Neutrino beams from muon storage rings: Characteristics and physics potential

S. Geer

Fermi National Accelerator Laboratory, P.O. Box 500, Batavia, Illinois 60510 (Received 4 December 1997; published 13 April 1998)

High-intensity high-energy neutrino beams could be produced by exploiting a very intense future muon source, and allowing the muons to decay in a storage ring containing a long straight section. Taking the parameters of muon source designs that are currently under study, the characteristics of the neutrino beams that could be produced are discussed and some examples of their physics potential given. It is shown that the neutrino and antineutrino beam intensities may be sufficient to produce hundreds of charged current interactions per year in a detector on the far side of the Earth. $[$ S0556-2821(98)01011-X $]$

PACS number(s): 14.60.Pq, 07.77.Ka, $13.15.+g$, 13.35.Bv

I. INTRODUCTION

High-energy neutrino beams have played an important role in the development of particle physics. Experiments using neutrino and antineutrino beams produced from charged meson decays have, for example, demonstrated that muon neutrinos are different from electron neutrinos, discovered neutral currents, provided measurements of the structure of the nucleon via deep inelastic scattering, made precision tests of the standard model of electroweak interactions via measurements of charged and neutral current interactions, and provided increasingly sensitive searches for neutrino oscillations in short- and long-baseline experiments.

Recent results from atmospheric neutrino, solar neutrino, and short-baseline accelerator neutrino experiments indicate that neutrino oscillations may occur at rates which are within reach of the next generation of accelerator based experiments. It therefore seems certain that experiments utilizing neutrino and antineutrino beams will continue to make important contributions to particle physics, initially by clarifying whether the existing results are indeed indications that neutrinos oscillate, and perhaps eventually by making precise measurements of the oscillations. In addition, there continues to be interest in using neutrino beams to further probe the structure of the nucleon.

The long-term future of the experimental program at neutrino beam facilities will require the continued improvement of the intensity and/or quality of the beams. The present generation of high-energy neutrino beams are made by allowing charged pions and kaons to decay in-flight in a long decay channel. This paper describes how a new type of neutrino beam could be made by exploiting a very high intensity muon source of the type that is currently being designed as part of the effort to develop the technology for a highluminosity muon collider $[1,2]$. The muons would be stored and allowed to decay in a ring containing a long straight section that points in the desired direction. The advantages of producing a neutrino beam using muon decays are discussed in Sec. II. The method of producing a neutrino beam from decaying muons is described in Sec. III. The resulting calculated fluxes, which are sufficient to provide significant interaction rates in a detector on the far side of the Earth, are described in Sec. IV for short-, long-, and very-long-baseline neutrino experiments. We note that geophysical applications of a facility that can shoot neutrino beams through the Earth have been discussed in Ref. [3]. Some examples of how the neutrino beams from a high intensity muon storage ring might be used to search for or measure neutrino oscillations are given in Sec. V, and conclusions are summarized in Sec. VI.

II. MESON DECAYS VERSUS MUON DECAYS

Presently operating high-energy neutrino beams are made by allowing charged pions and kaons to decay in a long decay channel. Shielding downstream of the decay channel removes the undecayed mesons, but transmits the weakly interacting neutrinos to form a ''pure'' neutrino beam. If positively charged pions and kaons have been selected for the decay channel, the resulting beam downstream of the shielding will contain mostly muon neutrinos produced in the two-body decays $\pi^+ \rightarrow \mu^+ \nu_\mu$ and $K^+ \rightarrow \mu^+ \nu_\mu$. The neutrino beam will also contain a small component of electron neutrinos produced in the three-body decays K^+ $\rightarrow e^+ \pi^0 \nu_e$. In addition, if the primary proton energy is sufficiently high, the beam will contain a small ν_{τ} component coming predominantly from prompt tauonic decays of *Ds* mesons. Antineutrino beams can be made by using a negatively charged meson beam.

Thus, present neutrino (antineutrino) beams consist mostly of muon neutrinos (antineutrinos) with a small $O(1\%)$ mixture of electron neutrinos (antineutrinos), and if the beam energy is high enough, a small component of tau neutrinos (antineutrinos). In general, this beam composition is not ideal for neutrino experiments. In particular, (i) the finite precisions with which the v_e and v_μ fluxes can be determined are important sources of systematic uncertainty for many neutrino experiments, (ii) the small ν_e contamination in the otherwise pure ν_{μ} beam is a nuisance for experiments searching for v_e - v_μ oscillations, (iii) the smallness of the v_e flux makes v_e - v_τ oscillation searches difficult, and (iv) the small ν_{τ} contamination in the beam will eventually become a nuisance for experiments searching for v_{μ} - v_{τ} oscillations.

These difficulties can be overcome if the neutrino beam is produced by allowing muons to decay in the straight section of a storage ring. This would produce a beam with a precisely known mixture of neutrino types; namely, 50% muon

neutrinos and 50% electron antineutrinos if a μ^- beam is stored, and 50% muon antineutrinos and 50% electron neutrinos if a μ^+ beam is used.

In the muon rest frame the distribution of muon antineutrinos (neutrinos) from the decay $\mu^{\pm} \rightarrow e^{\pm} + \nu_e(\bar{\nu}_e)$ $+\bar{\nu}_\mu(\nu_\mu)$ is given by the expression [4]

$$
\frac{d^2N_{\nu_{\mu}}}{dx d\Omega} = \frac{2x^2}{4\pi} [(3-2x) \mp (1-2x)\cos \theta],
$$
 (1)

where $x=2E_\nu/m_\mu$, θ is the angle between the neutrino momentum vector and the muon spin direction, and m_{μ} is the muon rest mass. The corresponding expression describing the distribution of electron neutrinos (antineutrinos) is

$$
\frac{d^2N_{\nu_e}}{dx d\Omega} = \frac{12x^2}{4\pi} \left[(1-x) \mp (1-x) \cos \theta \right].
$$
 (2)

Thus, the neutrino and antineutrino energy and angular distributions depend upon the parent muon energy, the decay angle, and the direction of the muon spin vector. For an ensemble of muons we must average over the polarization of the initial state muons, and the distributions become

$$
\frac{d^2N_{\nu_\mu}}{dx d\Omega} \propto \frac{2x^2}{4\pi} \left[(3-2x) \mp (1-2x) P_\mu \cos \theta \right],\tag{3}
$$

and

$$
\frac{d^2N_{\nu_e}}{dx d\Omega} \propto \frac{12x^2}{4\pi} \left[(1-x) \mp (1-x) P_\mu \cos \theta \right],\tag{4}
$$

where P_{μ} is the average muon polarization along the chosen quantization axis, which in this case is the beam direction.

The advantages of producing a neutrino beam using muon decays rather than meson decays are, therefore, (a) the absolute neutrino fluxes can be easily and precisely calculated, provided the stored muon current, momentum, and polarization are carefully measured, and (b) only one type of neutrino and one type of antineutrino are present in the beam, and these types can be chosen by selecting the charge of the stored muons. Thus precise v_e , v_u , \bar{v}_e , and \bar{v}_u measurements can be made.

In addition, the muons can be polarized and the time dependence of the precessing muon spin vectors monitored, enabling measurements to be made as the differential spectra of the neutrinos and antineutrinos in the beam vary in a precisely known way.

III. USING A MUON STORAGE RING

The muon lifetime is about 100 times longer than the corresponding charged pion lifetime. For example, 20 GeV/*c* muons have a decay length $\gamma c \tau = 126$ km. Thus, a linear decay channel of the type used to produce conventional neutrino beams would in practice be too short to use efficiently as a muon decay channel. This problem can be overcome by using a muon storage ring with a straight section pointing towards the desired experimental area. For simplicity we will consider a storage ring that consists of two parallel straight sections connected together by two arcs. If the straight sections are long compared to the arc lengths, the circulating muons spend approaching 50% of their time traveling in the straight section pointing towards the experiment. In practice there is an advantage in keeping the size of the storage ring small. Therefore, in the following we will assume that the straight sections are equal in length to the arcs, and that 25% of the injected muons decay as they circulate in the ring whilst they are in the straight section pointing at the experiment.

To understand how a muon storage ring designed for a given muon momentum might be used as a neutrino source, and calculate the parameters of the resulting neutrino beams, we must understand some of the basic parameters of the muon source and storage ring, namely, the divergence of the beam in the straight sections, the size of the ring, and the number of muons available from the source.

If the ring lattice is properly designed, a beam divergence $\theta_b \leq O(10^{-4})$ should be achievable [5] in the straight section. Thus, if the circulating muons have momentum p/m_{μ} $\leq 10^4$ (corresponding to $p \leq 1000$ GeV/*c*) the angular divergence of the neutrino beam produced from decays in the straight sections will be dominated by the decay kinematics.

In the scheme we are considering, the size of the storage ring is determined by the length of the arcs. The circumference of the ring in meters (*L*) designed to store muons of momentum $p(GeV/c)$ is given by

$$
L = \frac{4\,\pi p}{0.3fB},\tag{5}
$$

where B is the field of the arc dipole magnets (T) and the ''packing fraction'' *f* is the fraction of the arc lengths occupied by the dipoles. As an example, choosing the reasonable values $f=0.7$ and $B=8$ T to store 20 GeV/*c* muons we obtain an estimate of 37 m for the arc lengths and 150 m for the ring circumference. The estimate we are using for the arc lengths has been confirmed by a study $[6]$ of storage ring lattices for rings designed to store muons with momenta from 10 up to 250 GeV/*c*.

The direction of the muon beam must be carefully monitored within the straight section to avoid significant systematic uncertainties on the calculated neutrino fluxes at the experiment. This can be done by placing beam position monitors (for example, wire chambers) within the muon beam at either ends of the straight section. The angular precision that would be achieved (σ_{θ}) would depend upon the spatial resolution of the beam position monitors (σ_x) , and the distance between the monitors $(L/4)$. Using Eq. (5) to relate *L* to the stored muon momentum we obtain

$$
\sigma_{\theta} \sim \frac{0.3fB}{\pi p} \sqrt{2} \sigma_x, \qquad (6)
$$

where σ_x is in meters. Hence, for *B*=4 T, *f*=0.8, and choosing $\sigma_x = 100 \mu \text{m}$ we obtain for $p = 20 \text{ GeV}/c$ the estimate $\sigma_{\theta} = 2 \mu r$, which is small compared to the anticipated divergence of the muon beam within the straight section. Thus, for muons in this momentum range, it would appear that the required storage ring would be sufficiently compact, and the beam direction could be monitored with sufficient precision, to contemplate building the ring in a plane tilted at a large angle with respect to the horizon. This would enable the neutrino beams to be directed through the Earth for a very-long-baseline neutrino oscillation experiment.

The rate at which muons are stored in the ring, and therefore the average neutrino beam intensity, will be determined by the performance of the muon source. In the following we will assume a muon source of the type being developed as part of an ongoing effort to determine the feasibility of building a high-luminosity muon collider $[1,2]$. The muon source consists of a proton accelerator, charged pion production target and collection system, pion decay channel, and muon cooling channel. Details of the proton accelerator design are given in Ref. $[7]$, and descriptions of the other components are given in Ref. $[1]$. For completeness, a brief overview of the muon source and its assumed performance is given in the following paragraph.

In the muon collider front-end design that we are taking as an example, the muon source receives protons from an accelerator complex that accelerates bunches containing 5 $\times 10^{13}$ particles to energies of 16 GeV. The protons interact in a target to produce approximately 3×10^{13} charged pions of each sign per proton bunch. These pions are produced with only a very limited component of momentum transverse to the incident proton direction. The charged pions can therefore be confined within a beam channel using, for example, a 20 T coaxial solenoid with an inner radius of 7.5 cm. To collect as many pions as possible within a useful energy interval, it is proposed to use rf cavities to accelerate the lower-energy particles and decelerate the higher-energy particles. Muons are produced by allowing the pions to decay. At the end of a 20 m long decay channel, consisting of a 7 T solenoid with a radius of 25 cm, on average 0.2 muons of each charge would be produced for each proton incident on the pion production target. With two proton bunches every accelerator cycle, the first used to make and collect positive muons and the second to make and collect negative muons, there would be about 1×10^{13} muons of each charge available at the end of the decay channel per accelerator cycle. If the proton accelerator is cycling at 15 Hz, in an operational year (10⁷ sec) about 1.5×10^{21} positive and negative muons would have been produced in the decay channel and collected. The muons exiting the decay channel populate a very diffuse phase space. The next step is to ''cool'' the muon bunch, i.e., to turn the diffuse muon cloud into a very bright bunch with small dimensions in six-dimensional phase space, suitable for accelerating and injecting into a muon storage ring. The proposed method of cooling the muons is to use ionization cooling $[8]$. At the end of the ionization cooling channel each muon bunch is expected to contain about 5×10^{12} muons with a momentum of order 100 MeV/*c*. We will assume that the losses in accelerating the muons to modest energies (up to a few \times 10 GeV or less) are small, and therefore that 7.5×10^{20} muons of the desired charge are injected into the storage ring each operational year. In the scheme we are considering, 25% of the muons will decay in the straight section pointing at the experimental area, and the resulting neutrino beam will contain 2×10^{20} neutrinos per year and 2×10^{20} antineutrinos per year, with energy and angular distributions described by Eqs. (1) and $(2).$

FIG. 1. Calculated neutrino and antineutrino fluxes at a far site located 10 000 km from a muon storage ring neutrino source. The parameters of the muon storage ring are described in the text. The fluxes are shown as a function of the energy of the stored muons for negative muons (top plots) and positive muons (bottom plots), and for three muon polarizations as indicated.

IV. FLUXES AND INTERACTION RATES

In the following we will consider three scenarios, namely, (i) a very-long-baseline experiment in which the neutrino beam passes through the Earth, (ii) a long-baseline experiment in which the neutrino beam dips down a few degrees with respect to the horizon, and (iii) a "near" experiment which is 1 km from the neutrino source.

A. Very-long-baseline experiment

Consider a geometry in which the plane of the storage ring dips at an angle of \sim 50° to the horizon, and the resulting neutrino beam exits the Earth at the ''far site'' after traversing 10 000 km. This geometry would correspond to a storage ring sited at the Fermi National Accelerator Laboratory in the United States with the far site in Japan.

The calculated neutrino and antineutrino fluxes at the far site are shown in Fig. 1 as a function of the energy and average polarization of the muons decaying in the straight section of the storage ring. The fluxes have been averaged over a 1 km radius ''spot'' at the far site. The electron neutrino and antineutrino fluxes are very sensitive to the muon spin direction. The reason for this can be understood by examining Eq. (2) which shows that for μ^+ (μ^-) decays the ν_e ($\bar{\nu}_e$) flux→0 for all neutrino energies as cos $\theta \rightarrow +1$ $(-1).$

As an example we will consider in more detail neutrino beams from unpolarized positive muons stored with momenta $p=20 \text{ GeV}/c$ (50 GeV/*c*) which, in the absence of neutrino oscillations, produce at the far site 2.2×10¹⁰(1.4×10¹¹) $\bar{\nu}_{\mu}$ m⁻² year⁻¹ and 2.2×10¹⁰ (1.4×10^{11}) ν_e m⁻² year⁻¹. These results assume that the neutrino beam is pointing exactly in the direction of the far site. The differential distributions dN_{ν}/dE_{ν} are shown in Fig. 2 as a function of the angle $\Delta\theta$ between the neutrino beam direction and the direction of the far site. As $\Delta\theta$ increases, both the maximum neutrino energy and the neutrino

FIG. 2. Calculated neutrino and antineutrino differential spectra at a far site located 10 000 km from a muon storage ring neutrino source. The parameters of the muon storage ring are described in the text. The spectra correspond to unpolarized positive muons circulating in the muon storage ring with momenta of $20 \text{ GeV}/c$ (left plots) and $50 \text{ GeV}/c$ (right plots). The solid curves are obtained by averaging the fluxes over a central "spot" with opening angle $\Delta \theta$ $=1$ mr. The dashed and dotted curves are obtained by averaging over annuli centered on the beam axis and covering the angular intervals $1<\Delta\theta<2$ mr and $2<\Delta\theta<3$ mr, respectively.

flux decrease. However, if, as expected, the beam direction can be monitored with a precision $\sigma_{\theta} \le 1$ mr, the systematic uncertainty on the predicted neutrino and antineutrino fluxes and differential distributions at the far site should be modest.

The charged current neutrino and antineutrino rates in a detector at the far site can be calculated using the approximate expressions $[10]$ for the cross sections:

$$
\sigma_{\nu N} \sim 0.67 \times 10^{-38} \text{ cm}^2 \times E_{\nu}(\text{GeV})
$$
 (7)

FIG. 3. Calculated neutrino and antineutrino charged current interaction rates in a detector located 10 000 km from a muon storage ring neutrino source. The parameters of the muon storage ring are described in the text. The rates are shown as a function of the energy of the stored muons for negative muons (top plots) and positive muons (bottom plots), and for three muon polarizations as indicated.

FIG. 4. Calculated lepton and antilepton differential spectra for particles produced in charged current interactions in a detector located 10 000 km from a muon storage ring neutrino source. The parameters of the muon storage ring are described in the text. The spectra correspond to unpolarized positive muons circulating in the muon storage ring with momenta of $20 \text{ GeV}/c$ (left plots) and 50 GeV/c (right plots). The solid curves are obtained by averaging the fluxes over a central "spot" with opening angle $\Delta \theta = 1$ mr. The dashed and dotted curves are obtained by averaging over annuli centered on the beam axis and covering the angular intervals 1 $<\Delta \theta$ < 2 mr and 2 < $\Delta \theta$ < 3 mr, respectively.

and

$$
\sigma_{\bar{\nu}N} \sim 0.34 \times 10^{-38} \text{ cm}^2 \times E_{\bar{\nu}} (\text{GeV}). \tag{8}
$$

The predicted charged current interaction rates are shown in Fig. 3 as a function of the energy and average polarization of the decaying muons, and the associated charged-lepton energy distributions are shown in Fig. 4 as a function of the angle $\Delta\theta$ between the beam direction and the direction of the far site. In the absence of neutrino oscillations, the number of charged current interactions in a 10 kT far site detector $(\Delta \theta = 0)$ when unpolarized 20 GeV/*c* (50 GeV/*c*) positive muons are stored in the ring are 610 (1.0×10⁴) \bar{v}_{μ} interactions per year and 1×10^3 (1.6×10⁴) v_e interactions per year. We conclude that, for these particular examples, interactions from neutrinos and antineutrinos should be readily detectable at the far site. Note that the predicted charged current interaction rates and the shapes of the associated lepton energy spectra are both sensitive to $\Delta \theta$. This could be exploited by locating one or more satellite detectors at angular distances $\Delta \theta = O(1 \text{ mr})$ from the main far site detector.

B. Long-baseline experiment

Consider a long-baseline geometry in which the plane of the muon storage rings tilts at just a few degrees to the horizon. To be explicit we will consider a far site that is 732

FIG. 5. Calculated fluxes and spectra in a detector 732 km downstream of a muon storage ring neutrino source in which 10 GeV/*c* unpolarized muons are circulating. The top plots show the neutrino and antineutrino spectra, and the bottom plots show the charged lepton spectra from charged current interactions when positive muons (bottom left) and negative muons (bottom right) are stored in the ring.

km from the neutrino source. This geometry would correspond $[11]$ to a storage ring sited at the Fermi National Accelerator Laboratory and a far site at the Soudan underground Laboratory in Minnesota.

The neutrino fluxes and charged current interaction rates at the far site can be obtained by scaling the results presented in Figs. 1 and 3 by a factor of 187. Thus, in the absence of neutrino oscillations, unpolarized positive muons stored with momenta $p=20 \text{ GeV}/c$ (50 GeV/*c*) will produce at the far site 4×10^{12} (2.6 \times 10¹³) $\bar{\nu}_{\mu}$ and ν_{e} m⁻² year⁻¹. In a 10 kT detector, these fluxes would result in 1.1×10^5 (1.8) $\times 10^6$) \overline{v}_μ charged current interactions per year and 1.8 $\times 10^5$ (2.9 $\times 10^6$) v_e charged current interactions per year.

Given these large interaction rates it is worthwhile considering using a lower-energy muon storage ring. Predicted fluxes and spectra corresponding to using 10 GeV/*c* stored muons are shown in Fig. 5. The calculated neutrino and antineutrino fluxes are both $\sim 1 \times 10^{12}$ m⁻² year⁻¹, and the corresponding charged current interaction yields are 3.1 $\times 10^{3} \mu^{-} kT^{-1}$ year⁻¹ and $1.4\times 10^{3} e^{+} kT^{-1}$ year⁻¹ when negative muons are stored in the ring, and 1.6 $\times 10^3 \mu^+$ kT⁻¹ year⁻¹ and 2.9×10³e⁻ kT⁻¹ year⁻¹ when positive muons are stored in the ring. The mean energies of the charged leptons and antileptons produced in these charged current interactions are, respectively, \sim 3.5 and \sim 2 GeV. Thus, neutrino and antineutrino charged current interactions should be readily detectable at the far site when the decaying muons in the storage ring have momenta as low as 10 GeV/*c*.

C. Near experiment

Next consider a short-baseline geometry in which the detector is 1 km from the neutrino source. This geometry is of

FIG. 6. Calculated fluxes and spectra in a detector 1 km downstream of a muon storage ring neutrino source in which 1.5 GeV/*c* unpolarized muons are circulating. The top plots show the neutrino and antineutrino spectra, and the bottom plots show the charged lepton spectra from charged current interactions when positive muons (bottom left) and negative muons (bottom right) are stored in the ring.

interest to short-baseline neutrino oscillation experiments if the neutrino energies are much lower than those considered so far, and of interest to deep inelastic scattering experiments if the neutrino energies are much higher than considered so far.

Consider first a 1.5 GeV/*c* muon storage ring. Averaging the fluxes at the detector over a ''spot'' with a radius of 5 m, the predicted neutrino and antineutrino fluxes resulting from unpolarized muon decays are both 1.2×10^{16} m⁻² per year. The corresponding charged current interaction rates in a 1 kT detector yield 2.9×10^6 μ^+ per year and 5.1×10^6 *e*⁻ per year if positive muons are stored in the ring, and 6.0 $\times 10^6$ μ^- per year and 2.6 $\times 10^6$ e⁺ per year if negative muons are stored in the ring. The calculated neutrino and antineutrino differential spectra are shown in Fig. 6 together with the charged lepton distributions from charged current interactions. A 1 kT detector would record millions of charged current interactions per year with mean chargedlepton energies of ~ 0.6 GeV and mean charged-antilepton energies of ~ 0.3 GeV.

Finally we consider a higher-energy storage ring suitable for deep inelastic scattering experiments. This might be the last recirculating linear accelerator ring in a future muon collider accelerator complex. We will assume the muons are unpolarized and choose 250 GeV/*c* for the stored muon momentum. After 1 km the neutrino beam has a radius of less than 1 m. Hence nearly all of the $O(10^{20})$ neutrinos and antineutrinos per year will pass through a reasonably compact detector. This, together with the very large neutrino flux, would make possible a small detector incorporating detailed tracking and particle identification. If, for example, the fiducial mass is 10 kg, then the estimated charged current interaction rate is 8×10^5 neutrino interactions per year and 5×10^5 antineutrino interactions per year.

TABLE I. Summary of the neutrino oscillation experimental configurations considered in the text. The number of ν_e charged current interactions per year and the mean energies of the interacting neutrinos are listed for a detector of mass m_{DET} a distance *L* from a storage ring in which 7.5×10^{20} unpolarized positive muons per year are injected with momenta p_{μ} , and 25% of the muons decay in the straight section pointing at the experiment.

p_μ (GeV/c)	m_{DET}	↵ (km)	$\langle E_{\nu}\rangle$ (GeV)	$L/\langle E_\nu \rangle$ (km/GeV)	$v_{e}CC$ interaction/yr
20	10 K	10 000	13	744	1×10^3
10	10 K	732	6.6	111	3×10^4
20	10 K	732	13	57	2×10^5
1.5	20 T				1×10^5

V. EXAMPLES

To illustrate the physics potential of the muon storage ring neutrino sources discussed in the previous section, consider the sensitivity of an experiment searching for v_e - v_u or ν_e - ν_τ oscillations performed by searching for charged current interactions producing ''wrong-sign'' muons. For example, if positive muons are stored at the neutrino source, unoscillated muon-antineutrino and electron-neutrino charged current interactions will produce μ^+ and e^- , respectively. However, if the v_e transforms itself into a v_u during its passage to the detector the charged current interaction will produce a μ ⁻. Similarly, if the ν_e transforms itself into a ν_{τ} and the neutrino energy is sufficiently large, the charged current interaction will produce a τ^- which, with a branching ratio of 17% [9], will decay to produce a μ^- .

Within the framework of two-flavor vacuum oscillations, the probability that, whilst traversing a distance *L*, a neutrino of type 1 (mass m_1) oscillates into a neutrino of type 2 (mass m_2) is given by [4]

$$
P(\nu_1 \to \nu_2) = \sin^2(2\theta)\sin^2(1.27\Delta m^2 L/E),
$$
 (9)

where θ is the mixing angle, and $\Delta m^2 \equiv m_2^2 - m_1^2$ is measured in eV^2/c^4 , *L* in km, and the neutrino energy *E* is in GeV. Hence, in the absence of backgrounds or systematic uncertainties, a neutrino oscillation experiment can be characterized by the total number of neutrino interactions observed [and hence the minimum observable $P(\nu_1\rightarrow\nu_2)$] and the average *L*/*E* for the interacting neutrinos. These parameters are summarized in Table I for the experimental configurations discussed in the previous section. In the following we consider the very-long-, long-, and short-baseline configurations for a v_e - v_μ oscillation search, and then consider the long- and very-long-baseline configurations for a v_e - v_τ oscillation search.

A. Very-long-baseline experiment

Consider first the very-long-baseline experiment. The mean value of L/E for the interacting v_e downstream of a 20 GeV/*c* muon storage ring is 744 km/GeV, and the expected number of v_e charged current interactions per year in a 10 kT detector is $\sim 10^3$. In the absence of background, the minimum observable $P(\nu_e \rightarrow \nu_\mu)$ is therefore $O(10^{-3})$. To eliminate background from charge misidentification the experiment must be designed to provide a good measurement of the sign of the muon charge. Consider a detector design of the type being developed for the MINOS experiment $[11]$, namely, a detector constructed from 600 iron plates that are 4 cm thick and are magnetized toroidally with a field strength of 1.5 T. The iron plates provide the target for neutrino interactions, and tracking is provided by planes of streamer tubes placed between the iron plates with a 1 cm pitch. The curvature resolution for muons produced in neutrino interactions has been calculated for this detector, and is described by a Gaussian resolution function with an rms width given by $\lceil 11 \rceil$

$$
\sigma_{\kappa}/\kappa \sim 0.18/p_{\mu} + 0.06, \tag{10}
$$

for muons with momentum p_{μ} (GeV/*c*) in the range 1 $\langle p_\mu \langle 20 \text{ GeV}/c \rangle$. The non-Gaussian tails of the curvature resolution function resulting from hard interactions and from the non-Gaussian tails of the multiple scattering distribution have also been calculated, and are at the $4-5\times10^{-5}$ level. This non-Gaussian component dominates the calculated charge misidentification probability for $p_{\mu} > 1.3 \text{ GeV}/c$. However, measurements populating the tails of the resolution function can be suppressed by using the measured muon range to confirm the momentum determination, and hence reject large fluctuations in the curvature measurement. We anticipate that combined measurements of range and curvature should provide reliable charge determination with a misidentification probability $\leq 10^{-5}$ for muons with momenta >1 GeV/*c*. This would be adequate for the very-long baseline experiment provided most of the muons to be measured have momenta $p_\mu > 1$ GeV/*c*. This is seen to be the case in Fig. 7 which shows the predicted charged current $e^$ and μ^- spectra for sin²(2 θ)=1 and four choices of Δm^2 in the range $5 \times 10^{-4} < \Delta m^2 < 4 \times 10^{-3}$ eV²/ c^4 . The spectra are clearly sensitive to values of the oscillation parameters within this range.

Figure 8 compares the v_e - v_μ oscillation single-event sensitivity contour in the $[\Delta m^2, \sin^2(2\theta)]$ plane for the verylong-baseline configuration with the corresponding contours for the other configurations discussed in the following subsections. The contour has been calculated using the average value of *L*/*E* for the neutrinos producing charged current interactions in the detector. The very-long-baseline singleevent contour for one year of data taking extends down to $\Delta m^2 \sim 3 \times 10^{-5}$ eV²/ c^4 for sin²(2 θ)=1, with the "knee" in the contour at $\sin^2(2\theta) \sim 10^{-3}$. Since the probability of recording a background event due to charge misidentification would be small for a 10 kT yr data sample, in the absence of other significant sources of background, further data taking and/or a larger fiducial mass would improve the sensitivity of

FIG. 7. Predicted charged lepton energy distributions in a detector 10 000 km from a ring in which 20 GeV/*c* unpolarized positive muons are decaying. The predicted differential distributions for e^- (solid histograms) and μ^- (broken histograms) are shown assuming ν_e - ν_μ oscillations are occurring with sin² 2 θ =1 and Δm^2 as indicated on the four subplots.

the experiment. Note that the next generation of longbaseline neutrino oscillation experiments currently under design [11–20] are expected to probe down to $\Delta m^2 \sim 1$ $\times 10^{-3}$ eV²/ c^4 for sin²(2 θ)=1. Hence, the very-longbaseline experiment considered in this example would provide a significant improvement beyond the next generation of experiments.

FIG. 8. Contours of single-event sensitivity for v_e - v_μ oscillations for 1 year of running with the values of *L*/*E* specified on the figure, which correspond to the detector configurations summarized in Table I. The hatched and cross-hatched areas show the expected regions that will be explored by, respectively, the MINOS experiment [11] after 2 years of running and the MiniBooNe experiment $[12]$ after 1 year of running.

B. Long-baseline experiment

Now consider the long-baseline (Fermilab→Soudan) experiment. Starting with neutrinos from a 10 GeV/*c* muon storage ring, the mean value of L/E for the interacting v_e is 111 km/GeV, and the expected number of v_e charged current interactions per year in a 10 kT detector is 3×10^4 . In the absence of background, the minimum observable $P(\nu_e \rightarrow \nu_\mu)$ is therefore $O(10^{-4})$ in a 10 kT yr exposure. Most of the muons to be measured have momenta p_u >1 GeV/*c* (Fig. 5). Hence, a detector of the type described in the previous section could be used.

The calculated long-baseline v_e - v_μ oscillation singleevent contour is shown in Fig. 8, and extends down to $\Delta m^2 \sim 3 \times 10^{-5}$ eV²/ c^4 for sin²(2 θ)=1, with the "knee" in the contour at $\sin^2(2\theta) \sim 3 \times 10^{-5}$. Hence, for a 10 kT yr exposure, the long- and very-long-baseline experiments would have comparable sensitivity at large mixing angles. However, with a significantly longer data taking period or a much larger fiducial mass the long-baseline experiment would begin to record background events unless the charge misidentification probability could be reduced well below the 10^{-5} level. In fact it is believed $[21]$ that a further large reduction in the charge misidentification could be obtained by applying reasonable goodness-of-fit criteria to the reconstructed tracks and/or using a detector design in which the tracking detectors are within a magnetic field and the curvature measurement is made in a region between the iron plates.

C. Near experiment

Next consider the short baseline (1 km) experiment at a 1.5 GeV/*c* muon storage ring. The mean value of *L*/*E* for the interacting v_e is 1 km/GeV. A detector of the type considered in the previous sections for the long- and very-longbaseline configurations is unsuitable for a very low-energy near experiment in which the muons to be measured have momenta $p_u < 1$ GeV/*c*. A smaller detector with an active target and an external momentum spectrometer would be more suitable. For example, a short-baseline experiment has been proposed $\left[13\right]$ at CERN with a detector consisting of a 2.4 ton emulsion target configured in six modules, with each module containing a tracker downstream of the emulsion. The modules are within a 0.7 T dipole field, and the expected curvature resolution provided by the tracking system within a single module has been estimated to be better than 10%. The finite range of very low-energy muons will impose an effective cutoff in p_{μ} . The detailed module design for a low-energy oscillation experiment would therefore need to be optimized to minimize this p_{μ} cutoff while maintaining a reasonably large fiducial mass and precise curvature measurement.

An explicit detector design is beyond the scope of this paper, but to illustrate the required design criteria for an experiment at a 1.5 GeV/*c* muon storage ring we will consider the sensitivity of a detector having a 20 ton fiducial mass, constructed from modules designed so that the curvatures of muons with momentum p_μ >0.2 GeV/*c* are measured with a precision σ_{k}/κ <0.2. In the absence of background there would be $\sim 1 \times 10^5 \nu_e$ charged current interactions within the fiducial mass per year, corresponding to a minimum observable $P(\nu_e \rightarrow \nu_\mu)$ of $\sim 10^{-5}$. Hence, it is desirable that the non-Gaussian tails of the curvature resolution function be at the level of $\leq 10^{-5}$. Assuming this is the case, the predicted single-event contour for the near experiment is shown in Fig. 8. The sensitivity of this experiment at large mixing angles is comparable to the expected sensitivities for the next generation of long-baseline experiments. At small mixing angles and "large" Δm^2 the experiment would be sensitive down to $\sin^2(2\theta) \sim 10^{-5}$, which is an order of magnitude lower than the expected reach of the next generation of neutrino oscillation experiments $[11–20]$.

We conclude that a very-low-energy near experiment would be of interest if a detector with a fiducial mass of at least 20 tons could be designed to measure muons (p_{μ}) >0.2 GeV/*c*) with a charge misidentification probability $\leq 10^{-5}$, and with adequate suppression of other potential sources of background from, for example, secondary production of neutrinos from interactions in the vicinity of the experiment, or a component of neutrinos produced from the decays of ''wrong-sign'' muons produced in the accelerator complex upstream of the storage ring.

D. v_e **-** v_τ oscillation search

Finally, we consider a search for v_e - v_τ oscillations, based on a search for wrong-sign muons. We will restrict ourselves to the long- and very-long-baseline configurations with the MINOS-type detector described previously. Since at low energies the v_{τ} charged current cross section is suppressed due to threshold effects, and since only 17% of the decaying τ leptons produce a muon, the sensitivity of the ν_e - ν_τ oscillation search will be less than the corresponding sensitivity for the v_e - v_μ oscillation search. Consider an experiment in which 20 GeV/*c* unpolarized positive muons are stored in the muon storage ring. The average energies of the interacting tau neutrinos will depend upon the oscillation parameters, but will be approximately 15 GeV, giving average values of *L*/*E* of 49 and 660 km/GeV for the long-and verylong-baseline configurations, respectively. Using these values of *L*/*E*, Fig. 9 shows the calculated single-event contours in the $[\Delta m^2, \sin^2(2\theta)]$ plane for ν_e - ν_τ oscillations. At large mixing angles the long- and very-long-baseline experiments would be sensitive to values of Δm^2 approaching 10^{-4} eV/ c^2 , an improvement beyond the sensitivities of past ν_e - ν_{τ} oscillation searches [22,23] of several orders of magnitude.

VI. CONCLUSIONS

A very intense muon source of the type currently being developed for a future high-luminosity muon collider would provide sufficient muons to make very intense neutrino and antineutrino beams. If $O(10^{20})$ muons per year were allowed

FIG. 9. Contours of single-event sensitivity for ν_e - ν_τ oscillations based on searching for ''wrong-sign'' muons with the verylong-baseline (solid contour) and long-baseline (dotted contour) configurations discussed in the text and corresponding to the average *L*/*E* values of, respectively, 660 and 49 km/GeV. The contours correspond to a 10 kT year exposure with 20 GeV/*c* unpolarized muons stored in the ring and the straight section pointing at detectors 10 000 and 732 km from the ring.

to decay within a 20 GeV/*c* storage ring with a straight section pointing in the desired direction, the resulting beams would produce hundreds of charged current neutrino interactions per year in a 10 kT detector on the other side of the Earth. High beam intensities, together with the purity of the initial flavor content of the neutrinos and antineutrinos within the beam, would provide unique opportunities for short-, long-, and very-long-baseline neutrino experiments. The physics program that can be pursued with muon storage ring neutrino sources could begin with a short baseline experiment using relatively low-energy muons $[O(1 \text{ GeV})]$ as soon as an intense muon source became operational, and be extended to include higher-energy muon storage rings and longer-baseline experiments as higher-energy muon beams became available.

ACKNOWLEDGMENTS

I would like to thank Ray Stefanski, Chris Quigg, and Alvin Tollestrup for valuable comments on the manuscript. This work was performed at the Fermi National Accelerator Laboratory, which is operated by Universities Research Association, under Contract No. DE-AC02-76CH03000 with the U.S. Department of Energy.

- [1] The Muon Collider Collaboration, Fermilab-Conf-96/092, 1996 (unpublished).
- [3] A. De Rujula, S. L. Glashow, R. R. Wilson, and G. Charpak, Phys. Rep. 99, 341 (1983).
- [2] R. Palmer, A. Tollestrup, and A. Sessler, Proceedings of the 1996 DPF/DPB Summer Study on New Directions for Highenergy Physics (Snowmass 96), Snowmass, CO, 1996 (unpublished).
- @4# See, for example, T. K. Gaisser, *Cosmic Rays and Particle Physics* (Cambridge University Press, Cambridge, England, 1990).
- $[5]$ C. Johnstone (private communication).
- [6] C. Johnstone, High Intensity Muon Storage Rings for Neutrino Production: Lattice Design, Proceedings of the Workshop on Physics at the First Muon Collider and Front-End of a Muon Collider, Fermilab, 1997 (unpublished).
- [7] S. Holmes *et al.*, Report No. FERMILAB-TM-2021, 1997 (unpublished).
- [8] A. N. Skrinsky and V. V. Parkhomchuk, Sov. J. Part. Nucl. 12, 223 (1981).
- @9# Particle Data Group, R. M. Barnett *et al.*, Phys. Rev. D **54**, 261 (1996).
- [10] See, for example, F. Boehm and P. Vogel, *Physics of Massive Neutrinos* (Cambridge University Press, Cambridge, England, 1987).
- $[11]$ E. Ables *et al.*, Report No. NUMI-L-63, 1995 (unpublished).
- [12] BooNE Collaboration, E. Church et al., report, 1997 (unpublished).
- [13] TOSCA Collaboration, A. S. Ayan et al., Report No. CERN-SPSC/97-5, 1997 (unpublished).
- [14] KEK-PS E362 Collaboration, K. Nishikawa et al., report, 1995 (unpublished); Report No. INS-924, 1992 (unpublished).
- [15] NOE Collaboration, M. Ambrosio et al., Nucl. Instrum. Methods Phys. Res. A 363, 604 (1995).
- [16] P. Ce Cennini et al., Report No. LAGS-94/99-I, 1994 (unpublished).
- [17] T. Ypsilantis, Nucl. Instrum. Methods Phys. Res. A 371, 330 $(1996).$
- [18] ICARUS Collaboration, A. Rubbia et al., Report No. CERN-SPSLC-96-58, 1996 (unpublished).
- [19] NESS Collaboration, G. Drexlin et al., report (unpublished).
- [20] ORLANDO Collaboration, B. Bugg et al., "Prospect for a Neutrino Oscillation Experiment at the National Spallation Neutron Source,'' Proceedings of the Intersections of Particle and Nuclear Physics Conference in Big Sky, MT, 1997 (unpublished).
- [21] A. Para (private communication).
- [22] The Fermilab E531 Collaboration, N. Ushida et al., Phys. Rev. Lett. **57**, 2897 (1986).
- [23] BEBC WA66 Collaboration, M. Talebzadeh et al., Nucl. Phys. **B291**, 503 (1987).