

Mixmaster universe: A chaotic Farey tale

Neil J. Cornish

DAMTP, University of Cambridge, Silver Street, Cambridge CB3 9EW, England

Janna J. Levin

Center for Particle Astrophysics, UC Berkeley, 301 Le Conte Hall, Berkeley, California 94720-7304

(Received 21 August 1996)

When gravitational fields are at their strongest, the evolution of spacetime is thought to be highly erratic. Over the past decade debate has raged over whether this evolution can be classified as chaotic. The debate has centered on the homogeneous but anisotropic mixmaster universe. A definite resolution has been lacking as the techniques used to study the mixmaster dynamics yield observer-dependent answers. Here we resolve the conflict by using observer-independent fractal methods. We prove the mixmaster universe is chaotic by exposing the fractal strange repeller that characterizes the dynamics. The repeller is laid bare in both the six-dimensional minisuperspace of the full Einstein equations and in a two-dimensional discretization of the dynamics. The chaos is encoded in a special set of numbers that form the irrational Farey tree. We quantify the chaos by calculating the strange repeller's Lyapunov dimension, topological entropy, and multifractal dimensions. As all of these quantities are coordinate or gauge independent, there is no longer any ambiguity—the mixmaster universe is indeed chaotic. [S0556-2821(97)02710-0]

PACS number(s): 98.80.Hw, 05.45.+b, 95.10.Eg, 98.80.Cq

I. INTRODUCTION

When space is most strongly deformed, Einstein's nonlinear theory of gravity may be fundamentally chaotic. The singular cores of collapsing stars and the big bang are suspected to tend toward chaos [1,2]. Beyond conjecture,¹ attempts to conclusively identify chaos near singularities stirred debate [5,6]. The debate has centered on the mixmaster universe, an archetypal singularity. In the mixmaster model, the three spatial dimensions oscillate anisotropically out of the big bang and finally toward a big crunch [7].

Relativistic chaos has the unique difficulty of demanding observer-independent signatures. Many of the standard chaotic indicators such as the Lyapunov exponents and associated entropies are observer dependent. The Lyapunov exponent quantifies how quickly predictability is lost as a system evolves. The metric entropy quantifies the creation of information as time moves forward. They are both tied to the rate at which a given observer's clock ticks. The relativism of space and time rejects the notion of a preferred time direction. In a curved space, the Lyapunov exponent and metric entropy become relative as well. Observer-independent tools are needed to handle chaos in relativity.

The challenge of relativistic chaos is well demonstrated in the debate over the chaoticity of the mixmaster universe. Barrow studied the chaotic properties of a discrete approximation to the full dynamics known as the Gauss map [8,9]. Barrow found the Lyapunov exponent, metric entropy, and topological entropy of the map. The dispute began when nu-

merical experiments run in different coordinate systems found that the Lyapunov exponents vanished [10–14]. Furthermore, the Gauss map itself corresponds to a specific time slicing. The ambiguity of time was manifest.

To conclude that the mixmaster approach to a singularity is indeed chaotic, an observer-independent signature must be uncovered. As has been promoted elsewhere [15–17], fractals in phase space are observer-independent chaotic signatures. Their existence and properties do not depend on the tick of a clock or the world line of an observer. As well, fractals have an aesthetic appeal. The chaos is consequent of a lack of symmetries. The loss of symmetry in the dynamics is appeased by the emergent symmetry of the self-affine fractal [18] in phase space.

We exploit the observer independence of the fractal to show unambiguously that the mixmaster universe is chaotic. This result was announced in Ref. [17].

Fractals in the minisuperspace phase space of the full dynamics are uncovered. Probing the full dynamics is always numerically intensive. To open a window on the numerical results we also study the Farey map, a discrete approximation to the dynamics related to the Gauss map. We focus on the map analytically but the real power of our conclusions is in the full dynamics where no approximation is made. Just as with a geographical map, the Farey map is used as a guide to navigate through the full dynamical problem. We use the map to locate the skeleton of the chaos and then verify its existence with the numerics.

The bare bones of the chaos is revealed in an invariant subset of self-similar universes known collectively as a strange repeller.² The strange repeller [20] is a fractal, non-

¹Direct two- and three-dimensional numerical simulations of inhomogeneous cosmologies and gravitational collapse have so far failed to conclusively support or refute this conjecture [3]. It is possible that numerical discretization may suppress chaos as it causes a coarse graining of phase space similar to that found in quantum mechanics [4].

²Page appears to describe a strange repeller in the minisuperspace phase space of a scalar field cosmology [19]. Remarkably, this paper was written in the same year strange repellers were first being described in chaos theory [20].

attracting, invariant set (Sec. III). More familiar are fractal, *attracting*, invariant sets or strange attractors. Though the repellor is a tiny subset of all possible universes, it isolates the essential features of the system [21]. A typical universe will accumulate chaotic transient eras as it brushes past the strange repellor in phase space.

The repellor is multifractal as shown in Sec. III A. The Farey map divides the fractal into two complementary sets. A particularly elegant quality of the Farey map is the connection between these complementary sets and number theory. As elaborated in Sec. III B, the multifractal can be completely understood in terms of continued fraction expansions and Farey trees.

In addition to the usual multifractal dimensions, we stress the importance of the Lyapunov dimension (Sec. IV A). This dimension is built out of a coordinate-invariant combination of the Lyapunov exponent (Sec. IV C) and the metric entropy. Both the Lyapunov exponent and the metric entropy can then be reinstated as valid tools even in curved space. We introduce a method for handling the repellor as a Hamiltonian exit system, to facilitate calculation of the Lyapunov dimension (Sec. IV B).

Maps have been put to good use before. We devote Sec. V to a detailed comparison of the Gauss map and the Farey map. There are two main differences we draw out. One distinction is in the character of the maps themselves. They are not topologically equivalent and therefore have different topological features such as topological entropy. The other notable difference is the technique for handling the maps. Previous methods treat the set of all universes. We concentrate on the repellor subset. The previous techniques and ours are complementary.

Having isolated the repellor in the map, we use this insight as an x ray to expose the chaotic skeleton in the full, unapproximated dynamics (Sec. VII).

II. MIXMASTER DYNAMICS

The richness of Einstein’s theory is revealed in the spectrum of solutions to the dynamical equations. To solve these equations universe by universe, symmetries of the Hamiltonian are sought. However, strong gravity is nonlinearly entangled and we cannot expect the universe to typically offer up symmetries of the motion. Einstein’s equations are more likely than not to be fundamentally nonintegrable, and so chaotic.

The singularity structure of the equations can indicate if the system is integrable. It has been shown that the mixmaster equations fail the Painlevé test [22]. This suggests, but does not prove, that the system is nonintegrable.

The fractals we find in the minisuperspace phase space show conclusively the emergence of chaos. To begin, we use dynamical systems theory on sets of universes and consider the minisuperspace of all possible mixmaster cosmologies. The minisuperspace Hamiltonian is

$$H = (\ln a)'(\ln b)' + (\ln a)'(\ln c)' + (\ln b)'(\ln c)' + \frac{1}{4} [a^4 + b^4 + c^4 - 2(b^2c^2 + a^2c^2 + b^2a^2)], \quad (2.1)$$

where a, b, c are the scale factors for the three spatial axes. For a full description of the geometry see Refs. [7,23]. The

Hamiltonian constraint equation requires $H=0$. The equations of motion governing the behavior of the three spatial axes are

$$(\ln a)'' = \frac{1}{2} [(b^2 - c^2)^2 - a^4] \quad \text{et cyc.} \quad (a, b, c). \quad (2.2)$$

A prime denotes $d/d\tau$ where τ is related to the cosmic time t through $dt = (abc)d\tau$. The universe comes out of the big bang with two axes oscillating between expansion and collapse while the third grows monotonically. The overall volume of the universe expands to a maximum and then collapses. On approach to the big crunch, two spatial dimensions will oscillate in expansion and collapse while the third will decrease monotonically. Eventually a bounce occurs at which point the axes permute and interchange roles.

The scale and expansion factors can be reparametrized in terms of the four variables (u, v, ω, Ω) [9]. As was done in Ref. [9], the initial conditions can be fixed on the surface of section $a_0 = 1, (\ln a)'_0 \geq 0$, without loss of generality,

$$\begin{aligned} \ln a_0 = 0, \quad (\ln a)'_0 &= \frac{3(1+u_0)\omega_0}{(1+u_0+u_0^2)}, \\ \ln b_0 &= \frac{3\Omega_0}{(1+v_0+u_0v_0)}, \quad (\ln b)'_0 = \frac{-3\omega_0u_0}{(1+u_0+u_0^2)}, \\ \ln c_0 &= \frac{3(v_0+u_0v_0)\Omega_0}{(1+v_0+u_0v_0)}, \quad (\ln c)'_0 = \frac{3(u_0+u_0^2)\omega_0}{(1+u_0+u_0^2)}. \end{aligned} \quad (2.3)$$

The relative sizes and velocities of the axes, respectively, are accounted for by v and u . The overall scale factors $\Omega = [\ln(abc)]/3$ and $\omega = \Omega'$ grow monotonically away from the big bang to the maximum of expansion and then shrink monotonically toward the big crunch.

The evolution of a given universe can be viewed as a particle scattering off a potential in minisuperspace. The right-hand sides of Eqs. (2.2) play the role of a potential. When this potential is negligible, the evolution mimics a simple Kasner model. The Kasner metric is

$$ds^2 = -dt^2 + t^{2p_a}dx^2 + t^{2p_b}dy^2 + t^{2p_c}dz^2. \quad (2.4)$$

The Kasner indices can be parametrized as

$$\begin{aligned} p_a(u) &= \frac{1+u}{(1+u+u^2)}, \\ p_b(u) &= \frac{-u}{(1+u+u^2)}, \\ p_c(u) &= \frac{u+u^2}{(1+u+u^2)}, \end{aligned} \quad (2.5)$$

and satisfy $\sum_{i=1}^3 p_i = 1 = \sum_{i=1}^3 p_i^2$. Moreover, $p_a(1/u) = p_c(u)$, $p_b(1/u) = p_b(u)$, and $p_c(1/u) = p_a(u)$. As is customary, we take u to lie in the range $[1, \infty]$ to ensure a definite ordering of the Kasner exponents:

$$p_b < p_a < p_c. \quad (2.6)$$

In terms of conformal time,

$$(\ln a)' \simeq p_a \Rightarrow a \propto \exp(p_a \tau). \quad (2.7)$$

When the potential gains importance, the trajectory is scattered and enters another Kasner phase. To study the effect of scattering it is a good approximation to consider only the change on the Kasner indices of Eq. (2.5). The full dynamical problem can therefore be reduced to a map which evolves the parameter u forward in discrete intervals of time. Belinskii, Khalatnikov, and Lifshitz (BKL) reduced the dynamics to the one-dimensional Gauss map. The Gauss map evolves u from bounce to bounce as the trajectory strikes the potential. We focus instead on a two-dimensional map, the Farey map, which includes not only bounces but also oscillations [24,25]. The Farey map has the additional nice property of being invertible and so preserves the time reversibility of Einstein's equations [26].

III. FAREY MAP AND THE STRANGE REPELLOR

With the discrete time map we are able to isolate the repellor. The repellor is the chaotic subset of self-similar universes. It is the Hamiltonian analogue of the better known strange attractor. In dissipative systems, the volume of phase space shrinks and trajectories are drawn onto an attracting set as energy is lost. Since the universe is a self-contained Hamiltonian system, there can be no dissipation and so no attractors. Still, the notion of an invariant set can be an incisive characterization of the system.

Nonattracting invariant sets are referred to as repellors since all trajectories which constitute the set are unstable in at least one eigendirection, and in that sense repel any near neighbors. The phase space volume is conserved as it squeezes in one direction while it expands in the other as it careens off the repellor. The transient chaos experienced by a typical, aperiodic universe reflects the passage of that trajectory near the core periodic orbits.

Lifshitz and Khalatnikov [1] were the first to introduce the u parametrization and derive the evolution rules

$$\begin{aligned} u &\rightarrow u-1 & \text{if } u > 1, \\ u &\rightarrow 1/u & \text{if } u < 1. \end{aligned} \quad (3.1)$$

As it stands, this prescription is discontinuous at $u=1$. A continuous map can be derived by considering either of the two double transformations

$$u' = u-1 \quad \text{and} \quad u'' = 1/u'$$

or

$$u' = 1/u \quad \text{and} \quad u'' = u' - 1. \quad (3.2)$$

This leads to the two entirely equivalent continuous maps

$$u_{n+1} = \begin{cases} u_n - 1, & u \geq 2, \\ \frac{1}{u_n - 1}, & 1 \leq u < 2, \end{cases} \quad (3.3)$$

and

$$u_{n+1} = \begin{cases} u_n - 1, & u \geq 1, \\ \frac{1}{u_n} - 1, & 0 \leq u < 1. \end{cases} \quad (3.4)$$

Both choices have shown up in the literature over the years, but these days the first choice is preferred as the transitional step where $u < 1$ is hidden in the double transformation. By hiding this step, a definite Kasner ordering can be maintained. It is worth mentioning that the two maps use *exactly the same* u variable, even though the first can formally be recovered from the second by making the substitution $u = \tilde{u} - 1$.

By combining the u map with its inverse in terms of v , one arrives at the two-dimensional Farey map. We call the map a Farey map due to its connection to the Farey trees of number theory as discussed in Sec. III A. The Farey map evolves the two parameters forward by discrete intervals in time, $F(u_n, v_n) = (u_{n+1}, v_{n+1})$, according to the rule

$$F(u, v) = \begin{cases} u-1, & v+1, & u \geq 2 \text{ (oscillations)}, \\ \frac{1}{u-1}, & \frac{1}{v} + 1, & u < 2 \text{ (bounces)}. \end{cases} \quad (3.5)$$

The map describes the evolution of a universe through a series of Kasner epochs. If $u \geq 2$, then two axes oscillate in expansion and collapse while the third coasts. The axes will oscillate until $u < 2$, at which point a bounce occurs, the three axes interchange roles, and u is bounced about the number line. In some studies, what we call oscillations are called epochs and what we call bounces are called eras.

Running in parallel with the Farey map is the volume map $V(\Omega, \omega)$ described by

$$\Omega_{n+1} = \Omega_n \left[1 + \frac{1 + u_n + u_n^2}{u_n(1 + v_n + u_n v_n)} \right], \quad (3.6)$$

$$\omega_{n+1} = \omega_n \left[1 - \frac{2u_n}{1 + u_n + u_n^2} \right]. \quad (3.7)$$

Like the Farey map, the volume map is a good approximation to the exact evolution away from the maximum of expansion where ω flips sign. During the collapse phase $\tau \rightarrow -\infty$, $\omega > 0$ and $\Omega < 0$. A nice feature of the combined Farey-volume map, $M(u, v, \Omega, \omega)$, is the way it splits the dynamics into oscillatory and monotonic pieces.

Clearly, the monotonic behavior of the V map excludes the possibility of periodic trajectories for the M map. Since all chaotic behavior relies on the existence of unstable periodic trajectories, this would seem to rule out chaos in the mixmaster dynamics. Indeed, similar reasoning has been used to claim that the full mixmaster equations cannot be chaotic [27]. The flaw in this reasoning was to neglect the noncompact nature of the mixmaster phase space. For example, the noncompact motion along a spiral of fixed radius appears as circular motion in an appropriate comoving frame of reference. Although a spiral trajectory never returns to its starting point, the motion is periodic in a projected subset of phase space. Similarly, the mixmaster dynamics allows unstable periodic trajectories in terms of the (u, v) coordinates.

A projected strange repellor is all we require for chaos to occur. Consequently, we can drop the overall scaling of the V map and concentrate on the Farey map. Similarly, when we turn our attention to the full dynamics we will look for periodic behavior in scale-invariant quantities such as $(\ln b)/(\ln c)$. In this way we effectively project out the regular evolution of the system.

An important subset of universes are formed by the orbits periodic in (u, v) . The existence of a repelling, denumerably infinite, everywhere dense set of periodic orbits was first recognized by Bogoyavlenski and Novikov [28]. Physically, the periodic orbits correspond to discretely self-similar universes. After one orbital period, the proportionality of the scale and expansion factors recurs, while the volume of the universe has contracted overall. The collection of periodic orbits in (u, v) form a multifractal strange repellor.

To isolate the strange repellor we search for a *chaotic, nonattracting, invariant set*. In this section we justify each term in this definition.

Invariant set. The set of points which are invariant in time are located by the condition

$$F^k(\bar{u}_i, \bar{v}_i) = (\bar{u}_i, \bar{v}_i) \tag{3.8}$$

for $k = 1, \dots, \infty$. These fixed points lie on periodic orbits of period $p \leq k$. For Hamiltonian systems it is sufficient to consider the future invariant set as time reversal invariance can be used to find the complete strange repellor. For the Farey map the future invariant set is defined by the condition

$$F^k(\bar{u}_i) = \bar{u}_i \quad (v \text{ arbitrary}). \tag{3.9}$$

Consider for illustration the period-1 orbit. The invariant point satisfies $F(\bar{u}_1) = \bar{u}_1$. The map can be split into an oscillation map O and a bounce map B :

$$O(u_n) = u_n - 1, \quad u_n \geq 2, \tag{3.10}$$

$$B(u_n) = \frac{1}{u_n - 1}, \quad u_n < 2. \tag{3.11}$$

The equation $F(\bar{u}_1) = \bar{u}_1$ yields two possibilities: $O(\bar{u}_1) = \bar{u}_1$ or $B(\bar{u}_1) = \bar{u}_1$. Only the bounce fixed point, which generates the equation

$$\bar{u}_1 = \frac{1}{\bar{u}_1 - 1}, \tag{3.12}$$

has a solution. The solution is the golden mean

$$\bar{u}_1 = \frac{1 + \sqrt{5}}{2}. \tag{3.13}$$

The occurrence of the golden mean is no accident. As we discuss in the next section, it is the first leaf on the irrational Farey tree.

Another particularly simple set of fixed points can be located. These are the maximal value of u along a period k orbit. The maximal value of u corresponds to the largest number of oscillations before a bounce. The equation $O^{k-1}B(\bar{u}) = \bar{u}$ generates

$$\bar{u}_k(\bar{u}_k - k) = 1, \tag{3.14}$$

whose solution is

$$\bar{u}_k = \frac{k + \sqrt{k^2 + 4}}{2}. \tag{3.15}$$

These are the silver means (and can be found among the k th order leaves in the Farey tree). As elaborated in the next section, *all* of the fixed points comprise a countably infinite set of irrationals with periodic continued fraction expansions described by Farey trees. In this sense, we know all the fixed points.

Nonattracting. We can verify that all of the periodic orbits are unstable and so are not attracting but rather repelling. The repellor is the intersection in phase space of the unstable and stable manifolds. Since it is unstable in one direction, a near neighbor to a periodic orbit will deviate off the invariant set. Conservation of the Hamiltonian requires the other eigendirection to be attracting.

In the two-dimensional (2D) phase space defined by (u, v) , we show here that the unstable manifold is composed of fixed points along u . This is the future invariant set given by Eq. (3.9). Since time reversal corresponds to $(u, v) \rightarrow (v, u)$, the stable manifold is made up of the fixed points along v . The stable and unstable manifolds intersect at each point along a periodic orbit. The collection of such points forms the strange repellor. Since they are invariant, any point which is on both the unstable and stable manifolds will be mapped to points which by necessity inhabit the intersection of stable and unstable manifolds.

To demonstrate explicitly that the periodic orbits are unstable in the u direction but attracting in v consider an initial trajectory in the vicinity of a period p orbit. We look at the evolution in the u direction first. The aperiodic universe begins with $u_0 = \bar{u}_0 + \delta_0$ where \bar{u}_0 is the fixed point along the periodic universe. A period later, the deviation from the periodic orbit has grown to δ_p defined by [29]

$$F^p(u_0) = u_0 + \delta_p = F^p(\bar{u}_0 + \delta_0). \tag{3.16}$$

Expanding the right-hand side we can relate the evolved deviation to the initial deviation

$$\delta_p = c_p \delta_0, \tag{3.17}$$

where the stability coefficient is

$$c_p \equiv \left. \frac{dF^p(u_n)}{du_n} \right|_{u_n = \bar{u}_0} = F'(\bar{u}_0)F'(\bar{u}_1) \cdots F'(\bar{u}_{p-1}). \tag{3.18}$$

The elements in the final product are evaluated at the points along the unperturbed orbit. The periodic orbits are stable if the magnitude of δ_p shrinks and unstable if the magnitude grows. In terms of the stability coefficient the periodic orbits are stable if $|c_p| < 1$ and unstable if $|c_p| > 1$. Taking the derivative of the map

$$|F(u, v)'| = \begin{cases} 1, & 1 & (u \geq 2), \\ \frac{1}{(u-1)^2}, & \frac{1}{v^2} & (u < 2). \end{cases} \tag{3.19}$$

Along any periodic orbit, at least one point must fall between 1 and 2. Let \bar{u}_{\min} be the smallest value of \bar{u} along the orbit. It then follows from Eq. (3.18) that

$$|c_p| \geq (\bar{u}_{\min} - 1)^{-2} > 1. \quad (3.20)$$

Therefore all orbits on the repeller are unstable in the u eigendirection. Any universe which begins near an exact periodic orbit eventually deviates away from it. Since a typical universe is confined to bounce around the number line forever, it will eventually scathe past a periodic orbit before being repelled off only to stumble onto the repeller again. Thus, a typical universe will scatter around intermittently hitting chaotic episodes as it jumps on and off the repeller.

Repeating the stability analysis in v , we find $|c_k^v| = 1/v^2 < 1$. Near neighbors tend to follow periodic orbits in the v direction. The repeller along u is an attractor along v .

Chaotic. We have so far established that the periodic orbits comprise a nonattracting, invariant set. We can demonstrate that this set is chaotic by showing it has a positive topological entropy. The topological entropy, in analogy with the thermodynamic entropy, measures the number of accessible states on the repeller. The number of states on the repeller is equivalent to the number of fixed points, and so

$$H_T = \lim_{k \rightarrow \infty} \frac{1}{k} \ln N(k), \quad (3.21)$$

where $N(k)$ is the number of fixed points at order k . For a nonchaotic set, the number of fixed points is either finite or grows as a finite power of k , and so $H_T = 0$.

The fixed points at order k are found by solving Eq. (3.9): $F^k(\bar{u}_i) = \bar{u}_i$. To count the number of fixed points we can count the number of such possible equations. For a given period $p \leq k$, F^k can be broken into O^m oscillations and B^{k-m} bounces. Since O and B do not commute, the order in which they occur leads to different possible solutions for the \bar{u}_i . The number of ways to combine m O 's and $k-m$ B 's is

$$\binom{k}{m} = \frac{m!}{m!(k-m)!}. \quad (3.22)$$

For example, after $k=4$ iterations of the map and $m=2$ oscillations and $k-m=2$ bounces there are 6 possible permutations:

$$\begin{aligned} O O B B(\bar{u}_i) &= \bar{u}_i, \\ B B O O(\bar{u}_i) &= \bar{u}_i, \\ B O O B(\bar{u}_i) &= \bar{u}_i, \\ O B B O(\bar{u}_i) &= \bar{u}_i, \\ O B O B(\bar{u}_i) &= \bar{u}_i, \\ B O B O(\bar{u}_i) &= \bar{u}_i. \end{aligned} \quad (3.23)$$

However, the first four are all cyclic permutations of $O O B B$. Cyclic permutations must lie along the same orbit. Therefore, the solutions to the first 4 equations yield the

points along the same period-four orbit. Similarly, the last two are cyclic permutations of each other. They represent the recurrence of the two points along the period-2 orbit.

The total number of points belonging to the future invariant set with period $p \leq k$ is given by summing over all possible combinations of oscillations and bounces:

$$N(k) = \sum_{m=0}^{k-1} \binom{k}{k-m} = 2^k - 1. \quad (3.24)$$

There are thus $2^k - 1$ words of length k that can be built out of a two-letter alphabet. The topological entropy is then

$$H_T^u = \ln 2. \quad (3.25)$$

This entropy is independent of phase space coordinates. The topological entropy for the full 2D map $F(u, v)$ will be twice this quantity as the strange repeller is formed by the intersection of $2^k - 1$ horizontal and vertical lines. Thus, there are $(2^k - 1)^2$ roots of Eq. (3.8) at order k and $H_T = 2 \ln 2$. The result $H_T^u = \ln 2$ was first found by Rugh [10] using symbolic dynamics to describe typical, aperiodic orbits in u .

The topological entropy will be the same for all maps which are topologically conjugate, that is, which can be related by a continuous, invertible, but not necessarily differentiable, transformation of coordinates. In particular, this tells us that the map $F(u)$ can be obtained from the shift map, horseshoe map, or generalized baker's map [29]. We mention that the closely related Gauss map has a much higher topological entropy $H_T = \pi^2 / (6 \ln 2)$. As discussed in Sec. V, the two maps are not topologically conjugate. We do not expect then for their entropies to be the same.

The bare bones of the chaotic scattering have been exposed in the repeller. Taking this skeleton, we can now show that the repeller is in fact strange, that is, fractal.

A. The repeller is a multifractal

We can create an illustrative picture of the fractal set of self-similar universes. Distributing the fixed points u along the number line, the collection of periodic points forms a fractal in phase space. A fractal is a nowhere differentiable, self-affine structure. It cannot be undone by a coordinate transformation. The emergence of a fractal distribution of invariant points is an observer-independent declaration of chaos.

For multifractals there is information in the way points are distributed. There is an underlying fractal structure with an architecture of distributed points built on top. The fractal in phase space is simply constructed by locating all of the points along period $p \leq k$ orbits numerically and plotting them in a histogram as is done in Fig. 1. The histogram was generated by solving for all roots of Eq. (3.9) up to and including $k=16$. The histogram reveals how points are distributed in the future invariant set, otherwise known as the unstable manifold of the strange repeller. The self-similarity of the distribution is clear.

The distribution of points can be understood in terms of the period of the orbit. As k increases, the lowest period orbit, namely, the $p=1$ orbit, is visited the maximal number of times, that is, k times. On the other hand, the maximum

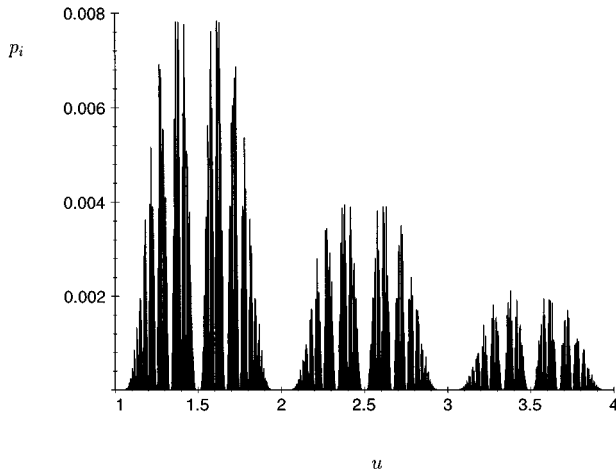


FIG. 1. A histogram of the future invariant set in the interval $1 < u < 4$.

value of u at order k lies on a $p = k$ orbit, and so is visited only once.

Points on the repeller are clustered in the interval $1 < u < 2$. The combinatorics of building words out of a two-letter alphabet (O, B) of oscillations and bounces favors the small u tower. Consider the integer interval $[n, n + 1]$. A root in this interval corresponds to the sequence, or word, $O^{(n-1)}B$ occurring somewhere along the orbit. The complete orbit is a sentence of $O^i B$ words, e.g., $(O^2 B)(B)(O^5 B)(OB)$, repeated in a cyclic fashion. Since the number of n -letter words that can be formed from a two-letter alphabet is 2^n , it follows that the fraction of roots in each interval is

$$\rho_n = 2^{-n}. \tag{3.26}$$

Note that the distribution is correctly normalized since

$$\sum_{n=1}^{\infty} \rho_n = 1. \tag{3.27}$$

The exponential falloff in the density of points on the repeller is clearly evident in Fig. 1.

The parametrization of the axes in terms of u in Eq. (2.5) shows that the small u orbits typical of the repeller are those with axes of similar scale and speed. These universes have axes which frequently switch from expansion to collapse. It follows that the strange repeller corresponds to the most isotropic mixmaster trajectories possible.

In the full 2D phase space of the Farey map, $F(u, v)$, the unstable manifold appears as a forest of vertical lines, while the stable manifold appears as a forest of horizontal lines. A portion of the future invariant set, the unstable manifold of the repeller, is displayed in Fig. 2. The collection of lines appears to form a multifractal cantor set. To confirm this, we calculate the fractal dimension of the set.

The fractal dimension can be thought of as a critical exponent. The dimension is a measure of the length of the collection of points. To find its box-counting dimension, cover the set with $N(\epsilon)$ boxes of size ϵ . As the size of the boxes is taken infinitesimally small, the number of boxes needed to cover the set grows. For fractals, the number of boxes needed grows faster than the scale shrinks. The struc-

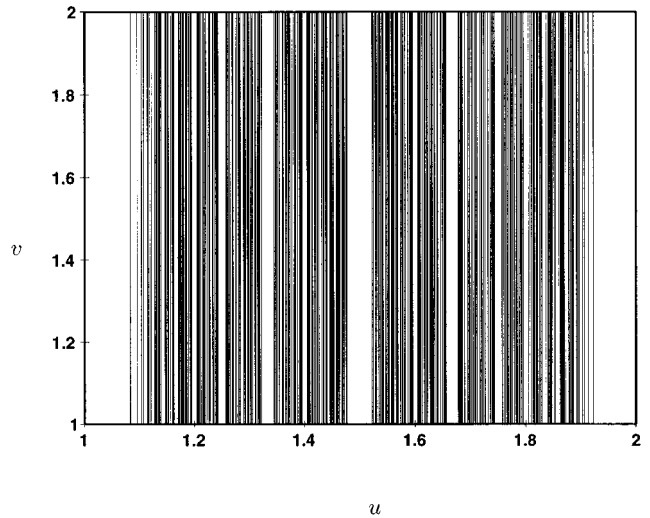


FIG. 2. The future invariant set in the interval $1 < u < 2$, $1 < v < 2$.

ture becomes more and more complex as smaller and smaller scales come into focus. The critical exponent can be defined as the number which keeps

$$\lim_{\epsilon \rightarrow 0} \epsilon^{D_0} N(\epsilon) \tag{3.28}$$

finite. Any exponent greater than D_0 would result in zero length, and an exponent smaller would result in infinite length. In this sense D_0 is a critical exponent.

This dimension can be generalized to include not just the scaling property of the fractal but also the distribution of points on top of the underlying foundation. This leads to a continuous spectrum of critical exponents or dimensions. The spectrum of dimensions is more commonly expressed as

$$D_q = \frac{1}{q-1} \lim_{\epsilon \rightarrow 0} \frac{\ln \sum_{i=1}^{N(\epsilon)} (p_i)^q}{\ln \epsilon}, \tag{3.29}$$

where $N(\epsilon)$ are the number of hypercubes of side length ϵ needed to cover the fractal and p_i is the weight assigned to the i th hypercube. The p_i 's satisfy $\sum_{i=1}^{N(\epsilon)} p_i = 1$. The standard capacity dimension D_0 is recovered when $q = 0$, the information dimension when $q = 1$, the correlation dimension when $q = 2$, etc. For homogeneous fractals all the various dimensions yield the same result. The multifractal dimensions D_q are invariant under diffeomorphisms for all q , and D_1 is additionally invariant under coordinate transformations that are noninvertible at a finite number of points [30].

The quantitative importance of the fractal dimension is expressed in terms of the final state sensitivity $f(\delta)$ [31]. This quantity describes how the unavoidable uncertainty in specifying initial conditions gets amplified in chaotic systems, leading to a large final state uncertainty. By identifying a number of possible outcomes for a dynamical system, the space of initial conditions can be divided into regions corresponding to their outcome. If the system is chaotic, the boundaries between these outcome basins will be fractal. Points belonging to the basin boundary are none other than the chaotic future invariant set. The function $f(\delta)$ is the frac-

tion of phase space volume which has an uncertain outcome due to the initial conditions being uncertain within a hypersphere of radius δ . It can be shown [31] that

$$f(\delta) \sim \delta^\alpha, \quad \alpha = D - D_0, \quad (3.30)$$

where D is the phase space dimension and D_0 is the capacity dimension of the basin boundary. For nonchaotic systems $\alpha=1$ and there is no amplification of initial uncertainties, while for chaotic systems $0 < \alpha < 1$ and marked final state sensitivity can occur.

It can be argued that the capacity dimension must equal the phase space dimension for the mixmaster repeller. The periodic orbits are given by the periodic irrationals as explained in Sec. III B. The periodic irrationals are dense on the number line. Formally this means there is always another periodic irrational ϵ away from a neighbor for all $\epsilon > 0$. Since the periodic orbits are dense, there will always be an infinite number of fixed points in any box. It is sufficient to cover the set then with $N(\epsilon) \sim 1/\epsilon^2$ boxes. Therefore, the basic box-counting dimension is $D_0 = D = 2$ [32], though it converges slowly. In the infinite time limit this can be interpreted as an ultimate loss in predictability since $f(\delta) = 1$. The mixmaster is very mixed.

While the box-counting dimension saturates at the phase space dimension, the more heavily weighted dimensions of Eq. (3.29) do not. Taking the information dimension of the fractal in Fig. 2, we find

$$D_1^u = 1.87 \pm 0.01. \quad (3.31)$$

Figure 3 shows the fit³ used to determine D_1^u . The quality of the fit gives us confidence that the strange repeller is truly fractal.

By forming the intersection of the Farey map's stable and unstable manifolds, i.e., by solving Eq. (3.8) for all fixed points (\bar{u}, \bar{v}) , we uncover the strange repeller. A portion of the strange repeller is shown in Fig. 4 using all roots up to

³The fit is actually for the fractal formed by taking a horizontal cross section through Fig. 2. The final answer is obtained by adding 1 to this number.

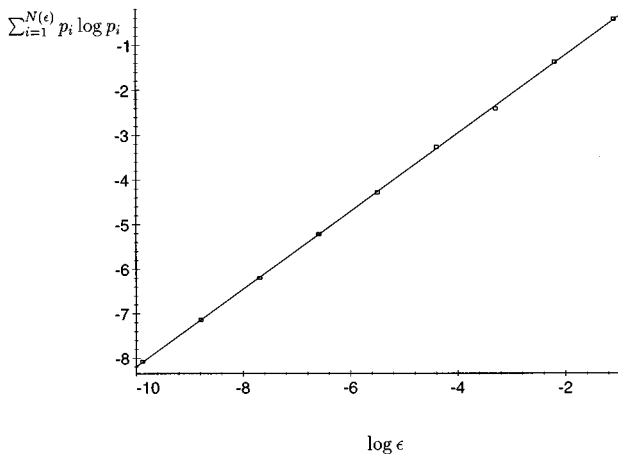


FIG. 3. Finding the information dimension of the future invariant set.

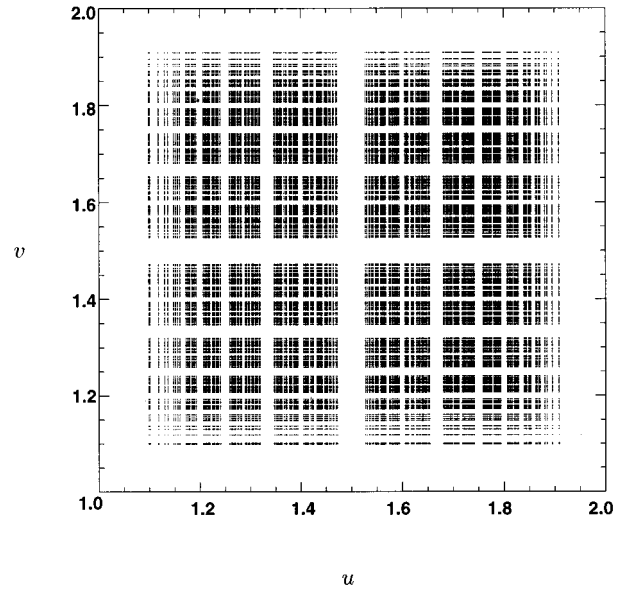


FIG. 4. The strange repeller in the region $1 < u < 2$, $1 < v < 2$.

$k = 12$. While it is possible to find the information dimension of the strange repeller directly from Fig. 4, the effort can be spared. Since the Farey map is Hamiltonian, we know the stable and unstable manifolds of the strange repeller share the same fractal dimensions. That is, $D_1^u = D_1^v = 1.87 \pm 0.01$. Now, the repeller is formed from the intersection of the stable and unstable manifolds, and so its information dimension is simply

$$D_1 = D_1^u + D_1^v - D = 1.74 \pm 0.02. \quad (3.32)$$

Since $D_0 \neq D_1$, we have confirmed the multifractal nature of the strange repeller. The dimension of Eq. (3.32) is calculated with all fixed points occurring at order $k = 16$. As k increases this number will continue to grow slowly as the rarified regions of the fractal continue to be populated. Thus, Eq. (3.32) actually represents a lower bound. However, we expect the ultimate value should only differ in the second decimal place.

B. Farey trees

The fractal of Fig. 4 reveals two complementary sets. The sequence of gaps corresponds to the rational numbers. In a complementary fashion, the strange repeller is made up from the periodic irrationals. Both sets are of Lebesgue measure zero. This division of the number line can be understood in terms of the properties of the continued fraction expansions (CFE's). We can write any number, rational or irrational, in terms of a CFE. While the connection to CFE's was known and explored by Belinskii, Khalatnikov, and Lifshitz, they explicitly ignored the periodic irrationals [2]. We focus on the periodic irrationals as they constitute the strange repeller.

Consider some initial condition u_0 not necessarily on a periodic orbit. We can decompose any number into its integer part and some left over:

$$u_0 = m_0 + x_0, \quad (3.33)$$

where $m_0 = [u_0]$ denotes the integer part of u_0 and x_0 , the fractional excess. According to the map, the next value of u following a bounce is $u_1 = 1/x_0$. Now decompose $1/x_0 = m_1 + x_1$, so that $m_1 \geq 1$. Solving this for x_0 we find $x_0 = 1/(m_1 + x_1)$. At the next bounce $u_2 = 1/x_1$ so that $x_1 = 1/(m_2 + x_2)$. In this way we generate the CFE for u_0 :

$$u_0 = m_0 + \frac{1}{m_1 + \frac{1}{m_2 + \frac{1}{\dots}}}, \tag{3.34}$$

where $m_i \geq 1 \forall i$. The integers $m_i - 1$ represent the number of oscillations between bounces. In shorthand form the CFE can be written as $u_0 = [m_0, m_1, m_2, m_3, m_4, \dots]$.

The map naturally distinguishes numbers on the basis of their CFE. A rational number can be written as the ratio of integers $x_0 = p/q$. Consequently, the CFE is finite. The fact that its finite means at some iteration $u_{n+1} = \infty$ and rationals are tossed out of the map. At the opposite extreme, the periodic orbits on the repellor have infinite CFE's which repeat. For example, an orbit of the form $\{O^{i-1}BO^{j-1}B\}$, where the curly brackets denote a repeated pattern, has the expansion

$$u_0 = i + \frac{1}{j + \frac{1}{i + \frac{1}{j + \frac{1}{i + \dots}}}}. \tag{3.35}$$

In shorthand form this reads $u_0 = [\{i, j\}]$. Lagrange showed that the necessary and sufficient condition for an irrational number to have a periodic CFE is for it to be the root of a quadratic equation with integer coefficients. It is easy to prove that the Farey map gives such equations for the periodic orbits at every order in k .

The map generates a Farey tree [33]. Consider the initial condition $\bar{u}_0 = 1$. With the first iteration of the map u is thrown to infinity. After two iterations, a universe with the initial condition $\bar{u}_0 = 2$ will be thrown to infinity. At the k th application of the map, after the universe has evolved forward in k jumps, another group of universes whose initial conditions were rational numbers are discarded. The pattern of discarded terms is a Farey tree [34]. Farey trees also arise in the quasiperiodic route to chaos described by the circle map [33].

In the interval $[1, 2]$ the Farey tree can be written as

$$\underbrace{\left(\frac{1}{1}\right)}_{k=0}, \underbrace{\left(\frac{2}{1}\right)}_{k=1}, \underbrace{\left(\frac{3}{2}\right)}_{k=2}, \underbrace{\left(\frac{4}{3}, \frac{5}{3}\right)}_{k=3}, \underbrace{\left(\frac{5}{4}, \frac{7}{5}, \frac{8}{5}, \frac{7}{4}\right)}_{k=4}, \dots \tag{3.36}$$

The bracketed terms each comprise a level of the Farey tree. For all $k \geq 2$, there are 2^{k-2} leaves at order k . Every rational number in the interval $[1, 2]$ occurs exactly once somewhere on the Farey tree. Each leaf on the Farey tree has a continued fraction expansion that satisfies

$$\sum_{i=1}^k m_i = k. \tag{3.37}$$

Now, we see that each m_i corresponds to a m_i -letter word $O^{(m_i-1)}B$. Each rational number corresponds to a sentence of these words. For example, at level $k=5$ we have the Farey number $10/7 = [1, 2, 3]$, which gives rise to the sentence *BOBOOB*. At the same order we also have $6/5 = [1, 5] = BOOOOB$, $9/5 = [1, 1, 4] = BBBOOB$, etc. In other words, a universe which began with $\bar{u}_0 = 10/7$ will follow the pattern of *BOBOOB* and then gets tossed to infinity. The universe then evolves with two axes fixed and the third expanding as $\propto t$. Such a universe has the form of a Rindler wedge $\times R^2$ [35].

At the opposite extreme from the rationals which escape the map are the periodic subset of irrationals. This explains the occurrence of the golden mean as the period-1 orbit with no oscillations. Its CFE is

$$\frac{1 + \sqrt{5}}{2} = 1 + \frac{1}{1 + \frac{1}{1 + \dots}} = [\{1\}]. \tag{3.38}$$

Numbers with a periodic CFE are a set of measure zero among the irrationals, though they are dense on the number line.

An irrational Farey tree for the repellor can be constructed using the rational Farey tree as a seed. Consider the construction in the interval $[1, 2]$. For each leaf on the rational Farey tree given by $[1, m_1, m_2, \dots, m_k]$, add a leaf to the irrational tree given by $[1, \{m_1, m_2, \dots, m_k\}]$. There are 2 times as many leaves on the irrational Farey tree than there are on the rational Farey tree. This follows since all rationals have two continued fraction expansions. The two CFE's differ in how they end. In general

$$[m_0, \dots, m_j + 1] = [m_0, \dots, m_j, 1]. \tag{3.39}$$

Consider $10/7 = [1, 2, 3]$. We see that $10/7$ also equals $[1, 2, 2, 1]$. Thus, $10/7 \approx 1.429$ has two corresponding leaves on the irrational Farey tree, $(\sqrt{15} - 1)/2 \approx 1.436$ and $(\sqrt{85} + 5)/10 \approx 1.422$. Consequently each rational generates two nearby irrationals.

A number theory result tells us that

$$\left| \alpha_{\pm} - \frac{p}{q} \right| < \frac{1}{2q^2} \leq \frac{1}{2k^2}, \tag{3.40}$$

where α_{\pm} stands for the two irrational numbers that lie on each side of their rational seed. As $k \rightarrow \infty$ the leaves on the rational and irrational Farey trees get closer and closer together. Another result from number theory tells us that $\alpha_+ > p/q$ and $\alpha_- < p/q$ (hence the names). The $1/q^2$ gaps around the rational Farey numbers seen in Figs. 1 and 2 are a consequence of this number theory result.

The asymptotic distribution of roots goes as follows. At level k there are 2^{k-2} leaves on the rational Farey tree in the interval $[1, 2]$. Each leaf produces 2 leaves on the irrational Farey tree. Thus, there are 2^{k-1} roots in the interval $[1, 2]$ at level k . This matches our earlier results that there are $\sim 2^k$ roots at order k , half of which lie in the interval $[1, 2]$.

Our brief excursion into number theory has produced a neat picture: The rational Farey tree gives us all the numbers that get mapped to $u = \infty$, while the irrational Farey tree

gives us all the periodic orbits. The leftovers are all the irrationals, save the rational and irrational Farey trees, both of which have measure zero. Thus, all typical trajectories are aperiodic, unbounded, and of infinite length. These are the trajectories typically studied in the literature. Our study is complementary.

C. Lyapunov exponents

In flat space, the extreme sensitivity of the dynamics can be quantified by Lyapunov exponents and the related metric entropy. The Lyapunov exponents determine how quickly in time trajectories diverge. The metric entropy measures the rate at which information is created. Since the exponent and the entropy are rates, they connect directly with the rate at which time pushes forward. Clearly, the observer dependence of the rate at which clocks tick make these tools suspect.

Although Lyapunov exponents are observer dependent and therefore ambiguous in general relativity, we see in Sec. IV A that they are related to an observer-independent quantity, namely, the Lyapunov dimension. Since there is still utility in them, we take the time to compute some Lyapunov exponents.

Since we know all the periodic orbits of the Farey map analytically, we are in a position to calculate the Lyapunov exponents for trajectories belonging to the Farey repellor. We calculate the Lyapunov exponents for the golden and silver mean orbits. Then we use the irrational Farey tree to write an analytic expression for the Lyapunov exponent for any periodic orbit on the repellor.

In close analogy to the manner in which the stability coefficient was found, the Lyapunov exponent for a given orbit is defined by

$$\lambda = \lim_{T \rightarrow \infty} \frac{1}{T} \sum_{n=0}^{T-1} \ln |F'(u_n)|, \quad (3.41)$$

where the u_n are the points along the orbit under scrutiny. The Lyapunov exponents along v are the negative of those along u as we now show. The points along a periodic orbit are the same whether time runs forward or backward. Time reversal corresponds to inverting the map. Note that $(F^{-1})'(u) = 1/F'(u)$ and therefore $\ln |F^{-1}'(u)| = -\ln |F'(u)|$. Since $F^{-1}(u, v) = F(v, u)$, it follows that $\ln |F'(v)| = -\ln |F'(u)|$. This identity in Eq. (3.41) shows that

$$\lambda^v = -\lambda^u, \quad (3.42)$$

as must be the case to conserve the phase space volume in the Hamiltonian system.

In illustration consider the golden mean period 1-orbit. All the $\bar{u}_n = \bar{u}_1$ and

$$F'(u_n) = -(\bar{u}_1 - 1)^{-2} = -(\bar{u}_1)^2. \quad (3.43)$$

From Eq. (3.41) it follows that

$$\lambda_1 = \lim_{T \rightarrow \infty} \frac{1}{T} (T-1) \ln \bar{u}_1^2, \quad (3.44)$$

which reduces to

$$\lambda_1 = 2 \ln \frac{1 + \sqrt{5}}{2} \approx 0.9624. \quad (3.45)$$

Similarly, for the silver means (3.15) we can find the exponent. We know the maximum value of u along the silver orbit is given by $u_{\max} = (\sqrt{k^2 + 4} + k)/2$, and so the sum of the ln's in Eq. (3.41) becomes the ln of the product,

$$F'(\bar{u}_0)F'(\bar{u}_1) \cdots F'(\bar{u}_{k-1}) = -(\bar{u}_{\max})^{-2}. \quad (3.46)$$

After k applications of the map, $F^k(\bar{u}) = \bar{u}$, so that after T application, a given value of \bar{u}_n has repeated $[T/k] = T/k$ times. Thus T must be a multiple of k . The sum in Eq. (3.41) can be written

$$\sum_{n=0}^{T-1} \ln |F'(u_n)| = \left[\frac{T-1}{k} \right] 2 \ln \bar{u}_{\max}, \quad (3.47)$$

which gives

$$\lambda_k = \frac{2}{k} \ln \frac{k + \sqrt{k^2 + 4}}{2}. \quad (3.48)$$

We have found the Lyapunov exponents for the two extreme cases of the period 1-orbit and the longest period- k orbit. We now use the irrational Farey tree to write down the general expression for Lyapunov exponents:

$$\lambda = 2 \left(\sum_{i=1} m_i \right)^{-1} \sum_{\text{cyc}} \ln(\{m_1; m_2, \dots\}), \quad (3.49)$$

where the second sum is taken over all terms in the cycle $\{m_1, m_2, \dots\}$ and $\sum m_i = k$ is the period of the orbit. Since $m_i > 1$, each term in the sum of logarithms is greater than zero. The Farey tree has given a nice compact form with which to express all the Lyapunov exponents. From Eq. (3.49) it is easy to recover the results for the golden and silver means:

$$\lambda_k = \frac{2}{k} \ln[\{k\}] = \frac{2}{k} \ln \frac{k + \sqrt{k^2 + 4}}{2}. \quad (3.50)$$

We can also calculate the average Lyapunov exponent for the periodic orbits:

$$\lambda_p = \lim_{l \rightarrow \infty} \frac{\sum_{m_1} \sum_{m_2} \cdots \sum_{m_l} \rho_l \lambda}{\sum_{m_1} \sum_{m_2} \cdots \sum_{m_l} \rho_l}. \quad (3.51)$$

Here l is the number of bounces along an orbit and ρ_l is the probability density. Upon summation over the m_i , the cyclic sum in Eq. (3.49) gives l identical contributions. Moreover, a long periodic orbit is composed of roughly equal numbers of oscillations and bounces so that $l \approx k/2$. These considerations allow us to replace the term $\lambda = (2/k) \sum_{\text{cyc}} \ln(\bar{u})$ in Eq. (3.51) by $\ln(\bar{u})$. Since we know the probability that each m_i equals a given integer n goes as 2^{-n} , it follows that

$$\lambda_p = \lim_{l \rightarrow \infty} \frac{\sum_{m_1} \sum_{m_2} \cdots \sum_{m_l} \rho_l \ln(\{m_1; m_2, \dots, m_l\})}{\sum_{m_1} \sum_{m_2} \cdots \sum_{m_l} \rho_l}, \quad (3.52)$$

where the density ρ_l is given by

$$\rho_l = \prod_{i=1, \dots, l} \rho_{m_i} = 2^{-(m_1 + m_2 + \dots + m_l)}. \quad (3.53)$$

The excellent convergence properties of the sums in Eq. (3.52) ensure that a finite truncation is able to provide a good estimate. Using $l=6$ and summing m_1 up to 30 and the $m_2 \cdots m_6$ up to 10 we find $\lambda_p = 0.793$. We can check this result by numerically evolving a large number of periodic orbits. Using MAPLE to find all $2^{15} - 1 = 32\,767$ periodic orbits at order $k=15$ and then evolving these orbits numerically, we find an average Lyapunov exponent of $\lambda_p = 0.792 \pm 0.002$. This is in excellent agreement with the finite truncation of the exact sum.

In contrast to the periodic orbits, typical aperiodic orbits have vanishing Lyapunov exponents. The average Lyapunov exponent for a typical aperiodic orbit can be approximated by

$$\langle \lambda \rangle = \frac{2}{u_{\text{av}}} \ln u_{\text{av}}. \quad (3.54)$$

From Fig. 5 we see that $u_{\text{av}} \rightarrow \infty$ as the number of iterations of the map grows large. Thus, as first noted by Berger [13], the average Lyapunov exponent for aperiodic trajectories tends to zero as $n \rightarrow \infty$. This behavior is characteristic of a chaotic scattering systems. Trajectories on the strange repeller have positive Lyapunov exponents while typical scattered orbits have Lyapunov exponents that tend to zero. The chaos in these systems is called transient as the brief chaotic encounter with the strange repeller is followed by regular asymptotic motion. Of course, all of these statements should be made with extreme care in general relativity as Lyapunov exponents are not gauge invariant.

IV. LYAPUNOV DIMENSION, TRANSIT TIMES, AND AVERAGE EXPONENTS

A. Lyapunov dimension

We have extolled the virtues of fractals in phase space as coordinate-independent signals of chaos. In this section we relate the fractal dimension of the invariant set to important dynamical quantities such as Lyapunov exponents and metric entropies. The Lyapunov dimension D_L combines these coordinate-dependent quantities into an invariant combination.

Remarkably, it has been shown that D_L equals the information dimension D_1 for typical chaotic invariant sets. The relation $D_1 = D_L$ has been rigorously established for certain dynamical systems [36] and has been numerically confirmed for many typical systems [37]. Specifically, it has been conjectured that [38,39]

$$D_1 = D_L = \Delta - \frac{(\lambda_1 + \lambda_2 \cdots + \lambda_\Delta) - 1 / \langle \tau \rangle}{\lambda_{\Delta+1}}. \quad (4.1)$$

Here $\langle \tau \rangle$ is the lifetime of typical chaotic transients, Δ is largest integer such that $\lambda_1 + \lambda_2 \cdots + \lambda_\Delta > 0$, and the Lyapunov exponents are ordered so that $\lambda_i \geq \lambda_{i+1}$. The equality of the information and Lyapunov dimension is be-

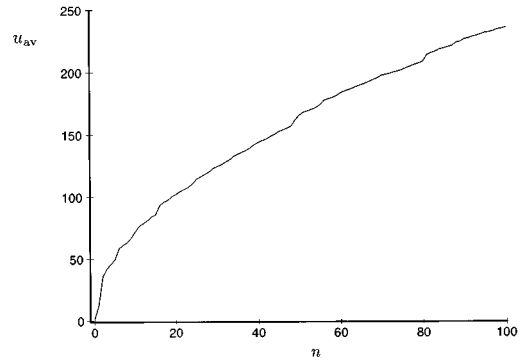


FIG. 5. The average value of u as a function of the number of iterations n . The average employs 10^6 randomly chosen trajectories initiated in the interval $u = [1, 2]$.

lieved to hold for most continuous dynamical systems, although a rigorous proof has only been given for discrete maps.

A heuristic derivation of this result can be found in the text of Ref. [29]. We sketch that reasoning here for the 2D Farey map. The repeller marks the intersection of the stable and the unstable manifolds. While it is repelling (unstable) in the direction of u , it is attracting (stable) in the direction of v . We can define a natural measure on the stable manifold, unstable manifold, and on the set itself. Ordinarily, an initial condition will produce an orbit which leaves the repeller never to return. Consequently, the number of trajectories near the repeller decays with time. The measure on a set is loosely related to the number of points which hang around the set. Consider some number $N(0)$ of random values for u scattered about the number line. These orbits are evolved by the map and eventually expelled. After a large number of iterations the only trajectories that remain belong to the invariant set, or are at least very close to a trajectories belonging to the invariant set. Thus, the measure of the repelling set can be defined as

$$\mu(t) \sim N(n)/N(0) \sim \exp(-n/\langle \tau \rangle). \quad (4.2)$$

The time scale $\langle \tau \rangle$ characterizes the decay time of typical trajectories leaving the repeller. In other words, $\langle \tau \rangle$ is the lifetime of typical chaotic transients. The measure can also be connected to the notion of the length of the set and consequently to the fractal dimension through

$$\epsilon^{2-D_1^u} \sim \mu, \quad (4.3)$$

where $\epsilon \sim \exp(-\langle \lambda \rangle n)$ and the Lyapunov exponent is defined by

$$\langle \lambda \rangle = \lim_{n \rightarrow \infty} \frac{1}{N(n)} \sum_{i=1}^{N(n)} \lambda_i. \quad (4.4)$$

The information dimension of the unstable manifold, D_1^u , makes an appearance in Eq. (4.3) since the measure accounts for the distribution of points as well as their location along the number line. Taking the natural logarithm of Eq. (4.3) yields

$$D_1^u = 2 - \frac{1}{\langle \lambda \rangle \langle \tau \rangle}. \quad (4.5)$$

Thus, we can relate the information dimension of the unstable manifold (the periodic irrationals in u) to the positive Lyapunov exponent and decay time of the Farey exit map. The dimension of the repeller is the sum of the dimensions of the stable and the unstable manifolds. For the conservative and hence invertible Farey exit map, the dimensions of the two manifolds are equal. Thus, the information dimension of the strange repeller is simply $D_1 = 2D_1^u - D$. Consequently, the Lyapunov dimension is

$$D_L = 2 \left(1 - \frac{1}{\langle \tau \rangle \langle \lambda \rangle} \right) = 2 \frac{h(\mu)}{\langle \lambda \rangle}, \quad (4.6)$$

where $h(\mu)$ is the the metric entropy:

$$h(\mu) = \langle \lambda \rangle - \frac{1}{\langle \tau \rangle}. \quad (4.7)$$

We see that D_L is given by the ratio of the metric entropy to the average Lyapunov exponent. Although neither Lyapunov exponents nor metric entropies are coordinate invariant, their ratio is. If a system is chaotic, that is, if $D_1 \neq 0$, a coordinate system can always be found in which h and $\langle \lambda \rangle$ are both finite and nonzero. From a dynamical systems perspective, such coordinate systems are preferable as they allow us to reinstate both Lyapunov exponents and metric entropies as useful chaotic measures.

In order to implement D_L , we derive $\langle \tau \rangle$ and $\langle \lambda \rangle$ for the mixmaster model in the following sections.

B. Hamiltonian exit systems

In our case, the system does not create an efficient repeller. Typical universes after being scattered off one periodic universe will eventually happen across another. The battle to toss trajectories off as they continually wash back ashore is constant. Because points which are discarded from the repelling set eventually return, the time $\langle \tau \rangle$ needed to discard typical trajectories from the repeller is infinitely long.

In principle, there is nothing wrong with a repeller that is revisited by previously scattered trajectories. In practice, it is easier to handle systems where scattered trajectories are discarded once and for all. We introduce a method for turning our thwarted repeller into a cleaner Hamiltonian exit system in this section.

The mixmaster dynamics presents some unique challenges as a dynamical system. It has some characteristics of a chaotic billiard [40], but this picture is upset by the non-compact nature of its phase space. The mixmaster also shares characteristics of a chaotic scattering system, but this picture is upset by the lack of absolute outcomes. Only a very special set of initial conditions leads to universes which terminate after a finite number of bounces.

There are other dynamical systems that combine features of chaotic billiards and chaotic scattering. These are known as Hamiltonian exit systems [41]. A standard example is a chaotic pool table. Trajectories can bounce chaotically around the table before falling into a particular pocket. When

the system is chaotic, the pocket a ball finishes up in depends sensitively on initial conditions. Each pocket has a basin of attraction in phase space, and the borders between the basins of attraction can be fractal. As discussed in Sec. III A, the fractal dimension of the basin boundaries provides a direct measure of the sensitive dependence on initial conditions.

In many ways, adding exits to a chaotic billiard is the best way to study the dynamics of the original closed system. By opening exits we can expose the underlying chaotic invariant set that encodes the chaotic behavior. The process would be familiar to an archeologist looking for bones. By running water through a sieve, dirt is washed away to expose the bones. Similarly, opening holes in a chaotic billiard allows most trajectories to escape, leaving behind the skeleton of unstable periodic orbits. As in archeology, some care has to be taken with the choice of sieve as bad choices can lead to the loss of bones along with the dirt.

For the mixmaster system, the placement of the pockets follows naturally. There are already three infinitesimally thin pockets. The three trajectories which lead out to these three pockets correspond to the Rindler universes with Kasner exponents $\{p_a=p_b=0, p_c=1\}$, $\{p_a=p_c=0, p_b=1\}$, or $\{p_b=p_c=0, p_a=1\}$. All we have to do is widen the pockets a little and the strange repeller will lie exposed. Since the three pockets correspond to large values of u , we know from our study of the Farey map that typical trajectories spend most of their time near a pocket. In addition, trajectories on the strange repeller are concentrated at small u values, and are thus far from the natural pockets. Only a small percentage of the strange repeller will be lost when we widen the pockets.

A nice visual representation of the mixmaster billiard is provided by the effective potential picture. The minisuper-space potential is given by

$$V = (a^4 + b^4 + c^4 - 2a^2b^2 - 2b^2c^2 - 2c^2a^2)/(abc). \quad (4.8)$$

The potential is best viewed by making the change of coordinates $(a, b, c) \rightarrow (\Omega, \beta_+, \beta_-)$:

$$\Omega = \frac{1}{3} \ln(abc), \quad (4.9)$$

$$\beta_+ = \frac{1}{3} \ln \left(\frac{bc}{a^2} \right), \quad (4.10)$$

$$\beta_- = \frac{1}{\sqrt{3}} \ln \left(\frac{b}{c} \right), \quad (4.11)$$

so that

$$V = e^{-8\beta_+} + 2e^{4\beta_+} (\cosh(4\sqrt{3}\beta_-) - 1) - 4e^{-2\beta_+} \cosh(2\sqrt{3}\beta_-). \quad (4.12)$$

Equipotentials of V are shown in Fig. 6. The natural pockets occur at the accumulation points $\theta_1=0$, $\theta_2=2\pi/3$, and $\theta_3=4\pi/3$. The angle θ is measured from the β_+ axis. Mixmaster trajectories can be thought of as a ball moving with unit velocity as it bounces between the triangular walls. The

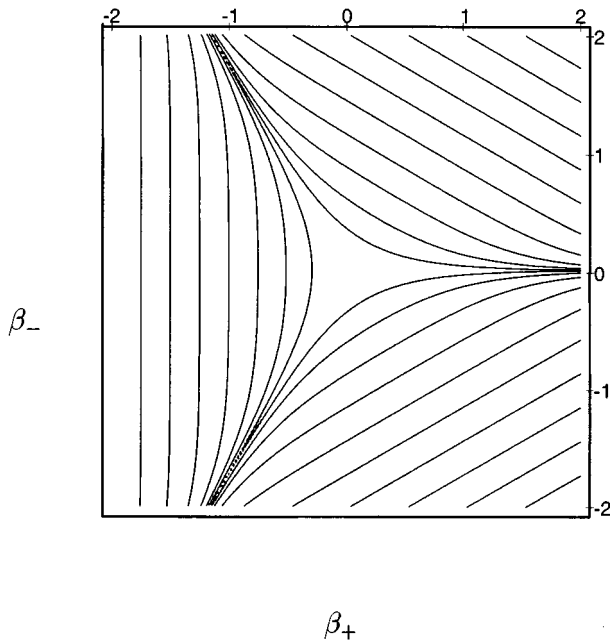


FIG. 6. Equipotentials of the minisuperspace potential.

walls are also moving, but at half the velocity of the ball. Equipotentials of V correspond to the position of the wall at different stages during the mixmaster collapse.

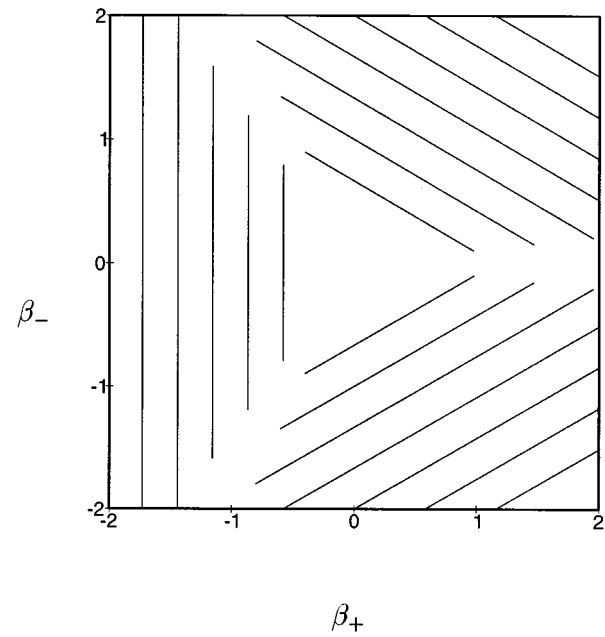
Typical trajectories oscillate around in one of the corners for a long time before bouncing out to a new corner where they oscillate around for a long time and so on *ad infinitum*. These are the typical, aperiodic mixmaster trajectories. Thus, the mixmaster dynamics leads to three accumulation points at the corners of the triangle. This behavior was first noted by Misner [7], and was later studied numerically by Creighton and Hobill [42].

In addition to the aperiodic trajectories, there are two special classes of mixmaster trajectories. One class corresponds to the rational numbers. After a finite number of oscillations and bounces they take the perfect bounce and head straight out one of the infinitesimally thin pockets. The other class of special trajectories forms the strange repellor and corresponds to the periodic irrationals. These trajectories regularly bounce from corner to corner and spend little time oscillating in each corner. Consequently, trajectories on the strange repellor are unlikely to exit the system when we widen the pockets.

By widening the natural pockets, we create exits. For example, the first pocket becomes the angular region $[-\Delta\theta, \Delta\theta]$. As the collapse proceeds, $\Omega \rightarrow -\infty$, and we can relate $\Delta\theta$ to the map parameter u via

$$\Delta\theta = 2\arctan\left(\sqrt{3}\left[\frac{u+1}{u-1}\right]\right) - \frac{2\pi}{3} \\ = \frac{\sqrt{3}}{u}\left(1 - \frac{1}{2u} + \frac{1}{4u^3} - \dots\right). \quad (4.13)$$

We see that choosing an exit value for u sets the angular width of the pockets. In the limit $u_{\text{exit}} \rightarrow \infty$ we recover the original mixmaster pockets with zero angular width. In Fig. 7 we display equipotentials of the minisuperspace potential

FIG. 7. The minisuperspace potential with exits set by $u_{\text{exit}} = 16$.

with pockets corresponding to the exit value $u_{\text{exit}} = 16$. While the pockets appear quite wide, we see from Eq. (3.26) that they have little effect on the strange repellor. Few bones will be lost from the chaotic skeleton. For $u_{\text{exit}} = 16$ less than 0.003% of the strange repellor will be lost out the corner pockets. The wider we make the pockets the faster the strange repellor is exposed. For numerical studies of the full dynamics we chose $u_{\text{exit}} = 8$, giving exits ≈ 1.7 times wider than those shown in Fig. 7, but still small enough to ensure that less than 1% of the chaotic skeleton is lost.

With the exits in place the mixmaster behaves like a classic chaotic scattering system. If we start N_0 randomly chosen trajectories, there will be

$$N(n) = N_0 \exp\left(-\frac{n}{\langle\tau\rangle}\right) \quad (4.14)$$

remaining after n iterations of the Farey map. A similar exponential decay can be seen for the full dynamics if we use the T time coordinate. Naturally, the rate of decay is coordinate dependent in general relativity. However, $\langle\tau\rangle$ is related to the Lyapunov dimension which is coordinate independent.

To build the pockets we modify the map to

$$u_{n+1} = \begin{cases} \infty, & u_n \geq u_{\text{exit}} \quad (\text{exit}), \\ u_n - 1, & 2 \leq u_n < u_{\text{exit}} \quad (\text{oscillation}), \\ \frac{1}{u_n - 1}, & 1 < u_n < 2 \quad (\text{bounce}). \end{cases} \quad (4.15)$$

We refer to this map as the Farey exit map.

In Fig. 8, we show a histogram of the fixed points of the exit map (4.15) by finding all roots up to order $k = 16$ with $u_{\text{exit}} = 8$. As promised, the strange repellor for the Farey exit map is little changed from the repellor of the full map shown in Fig. 1. The main difference is that the gaps around the rationals are now completely devoid of roots and are not just

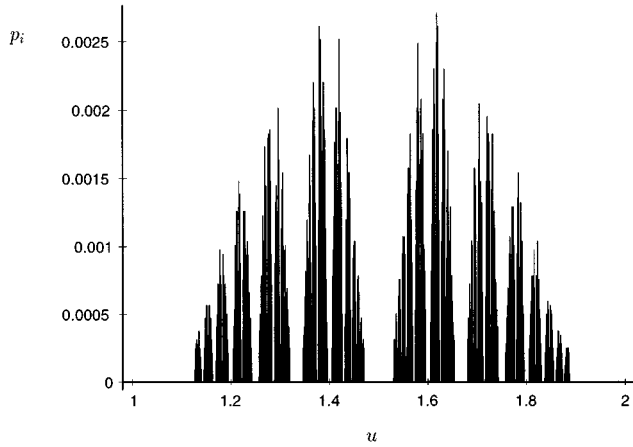


FIG. 8. A histogram of points belonging to the Farey exit map's future invariant set ($u_{\text{exit}}=8$).

sparingly populated regions. As a result, the capacity dimension of the fractal in Fig. 8 will not saturate at the phase space dimension. However, since the dense regions of the repeller have not been affected by the introduction of exits, we expect the information dimension to be little changed from that of the full map. We find $D_1^u = 1.87 \pm 0.01$, in agreement with the information dimension of the future invariant set shown in Fig. 2.

C. Measures and $\langle \lambda \rangle$

When dealing with maps, the transient lifetime and Lyapunov exponents for the aperiodic trajectories can be calculated from a knowledge of the periodic trajectories. In this way, the strange repeller provides a complete description of the dynamics even though the periodic trajectories are a set of measure zero.

Recall that the fixed points u_{jk} at order k are found by solving the $2^k - 1$ equations $F^k(u_{jk}) = u_{jk}$. Writing the stability coefficient of u_{jk} as c_{jk} , it can be shown [39] that

$$\mu(S) = \lim_{k \rightarrow \infty} \sum_{u_{ij} \in S} \frac{1}{c_{jk}}, \quad (4.16)$$

where the measure covers the region S . If S covers the entire invariant set, we find

$$\mu = \lim_{k \rightarrow \infty} \sum_j \frac{1}{c_{jk}} = \lim_{k \rightarrow \infty} \exp\left(-\frac{k}{\langle \tau \rangle}\right). \quad (4.17)$$

If $\langle \tau \rangle \neq \infty$, the measure decays as trajectories are lost from the system. The non-negative Lyapunov exponent is given by [39]

$$\langle \lambda \rangle = \lim_{k \rightarrow \infty} \frac{\sum_j (\ln c_{jk}) / c_{jk}}{\sum_j 1 / c_{jk}}. \quad (4.18)$$

To clarify, $\langle \lambda \rangle$ is the average Lyapunov exponent for a typical aperiodic trajectory. This is to be contrasted with the average Lyapunov exponent λ_p of the rare, periodic orbits described in Sec. III C.

Applying these relations to the Farey map without exits we find

$$\mu_k = \sum_j \frac{1}{c_{jk}} \simeq 1 - e^{-0.5k^{-0.268}}, \quad (4.19)$$

so that $\langle \tau \rangle = \infty$. As emphasized earlier, the Farey map without exits sees typical orbits return to the scattering region so $\mu = 1$.

Before we calculate D_L , we take a brief aside to demonstrate the effectiveness of the measure (4.16) by reproducing the topological entropy calculated analytically in Eq. (3.21). Since the measure does not decay, we can use the q -order entropy spectrum [43]

$$H_q = \frac{1}{1-q} \lim_{k \rightarrow \infty} \frac{1}{k} \ln \sum_j (c_{jk})^{-q} \quad (4.20)$$

as an independent calculation of the topological entropy. The topological entropy is recovered in the limit $q \rightarrow 0$, so that $H_T = H_0$. Using all roots up to $k=16$ and taking $q=0.01$ we find $H_{0.01} = 0.693 \simeq \ln 2$, in good agreement our analytic calculation of H_T . The fast convergence of $H_{0.01}$ as a function of k is shown in Fig. 9. In theory, by taking the limit $q \rightarrow 1$ we should recover the metric entropy $h(\mu) = H_1$. In practice, the expression for H_1 converges very slowly with k and we are unable to confirm that $h = H_1 = 0$.

Turning to the Farey exit map with an exit at $u_{\text{exit}}=8$ we find from Eqs. (4.17) and (4.18) that $\langle \tau \rangle = 12.6 \pm 0.1$ and $\langle \lambda \rangle = 0.724 \pm 0.005$. The convergence as $k \rightarrow \infty$ of Eqs. (4.17) and (4.18) is shown graphically in Figs. 10 and 11, respectively. Notice that the fits only start to converge once $k > u_{\text{exit}} + 1$. This makes sense as the the largest root at order k is $u_{\text{max}} \simeq k - 1$. The map is “unaware” it has exits until orbits start to escape. These results for $\langle \tau \rangle$ and $\langle \lambda \rangle$ combine to tell us that $D_L = 1.78 \pm 0.02$ and $h = 0.645 \pm 0.005$.

As an independent check we can use a direct Monte Carlo evaluation based on evolving a random collection of initial conditions. The Monte Carlo method allows us to estimate the properties of typical trajectories by measuring the properties of a large random sample and then inferring from this the properties of the collective.

By evolving $N(0) = 10^5$ randomly chosen initial points, we find from Eqs. (4.2) and (4.4) that $\langle \tau \rangle = 12.2 \pm 0.1$ and $\langle \lambda \rangle = 0.74 \pm 0.01$. The convergence of $\langle \tau \rangle$ and $\langle \lambda \rangle$ as a function of the number of iterations n is shown in Figs. 12 and 13. In each case the initial conditions in the first run were taken from the interval $u = [1, 2]$, and those in the second run from the larger interval $u = [1, 8]$. In this way, the asymptotic value was approached from above and below, leading to a more accurate estimate of the limit. After about $n = 60$ iterations statistical errors start to become large since the number of points remaining in the interval becomes small. This is apparent in Figs. 12 and 13. Using the Monte Carlo values for $\langle \tau \rangle$ and $\langle \lambda \rangle$ we find $D_L = 1.78 \pm 0.03$ and $h = 0.66 \pm 0.01$. The two methods agree.

In order to compare D_L to the information dimension of the strange repeller we need to numerically generate the repeller and find its fractal dimensions. Recall that in Fig. 8 we generated the unstable manifold of the Farey exit map by solving for all roots up to order $k=16$. With an exit at $u_{\text{exit}}=8$ we find $D_1^u = 1.87 \pm 0.01$. Consequently, the information dimension of the strange repeller is $D_1 = 1.74 \pm 0.02$,

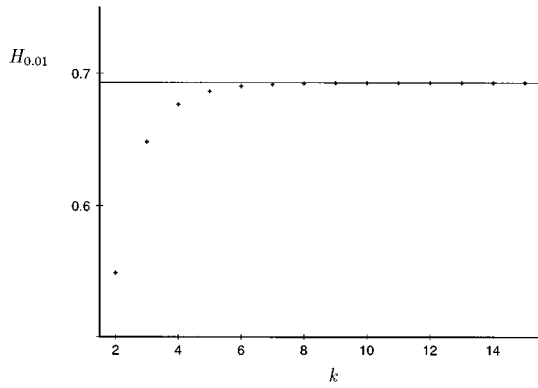


FIG. 9. Numerical convergence of the topological entropy, $H_{0.01} \approx H_T$, as a function of k using Eq. (3.20). The solid line indicates $\ln 2$.

which agrees with the result without exits [Eq. (4.32)]. As mentioned previously, the dimension found for $k=16$ is a lower bound for the true dimension. While there is reasonable agreement between D_1 and D_L at order $k=16$, we would like to test the relation at higher orders in k . However, this is difficult since the number of roots increases exponentially with k .

Alternatively, we can expose the future invariant set by the Monte Carlo method. By starting off $N(0) = 5 \times 10^6$ points in the interval $u = [1, 2]$ and iterating them $n = 60$ times, we are left with $N \approx 3 \times 10^4$ points. These remaining points must closely shadow the future invariant set. A histogram of the set generated in this way is shown in Fig. 14. The gap structure is identical to that seen in Fig. 8 using all roots at order $k=16$, but the bins are more evenly filled. As a result, the information dimension takes the slightly higher value $D_1^u = 1.89 \pm 0.01$. This method gives an information dimension of the strange repeller closer to $D_1 = 1.78 \pm 0.02$. Since the Monte Carlo method only finds orbits that shadow the repeller, and not the true periodic orbits, we expect the fractal dimension found in this way will be an upper bound.

As expected, the results for $\langle \tau \rangle$ and $\langle \lambda \rangle$ depend on how wide we make the pocket. In the limit $u_{\text{exit}} \rightarrow \infty$ we know $\langle \tau \rangle \rightarrow \infty$ and $\langle \lambda \rangle \rightarrow 0$. It would be interesting to see if the product $\langle \tau \rangle \langle \lambda \rangle$ remains finite as u_{exit} becomes large. Unfor-

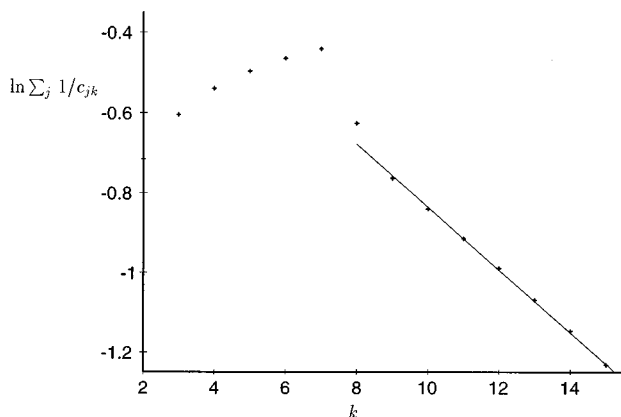


FIG. 10. Finding the decay time $\langle \tau \rangle$ for the Farey exit map using Eq. (3.17). The straight line fit yields a decay constant of $\langle \tau \rangle = 12.6 \pm 0.1$.

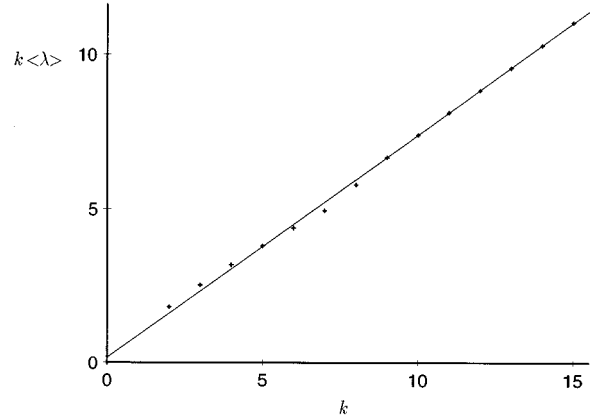


FIG. 11. Finding the Lyapunov exponent $\langle \lambda \rangle$ for the Farey exit map using Eq. (3.18). The straight line fit yields a Lyapunov exponent of $\langle \lambda \rangle = 0.724 \pm 0.005$.

tunately, the numerical values converge very slowly as the pockets become thinner, and so we were unable to test whether D_L remains fixed as $u_{\text{exit}} \rightarrow \infty$.

To conclude this section, we find from two separate methods that

$$D_L = 1.78 \pm 0.03. \tag{4.21}$$

Calculating the information dimension in two different ways gives the range

$$D_1 = (1.74 \rightarrow 1.78) \pm 0.02. \tag{4.22}$$

Systematic errors and slow convergence of the dimension can account for the discrepancy in the two values for D_1 . The Lyapunov dimension is in reasonable agreement with two different calculations of the information dimension. The Lyapunov dimension thus fares well as a coordinate-invariant combination of the otherwise ambiguous Lyapunov exponent and metric entropy.

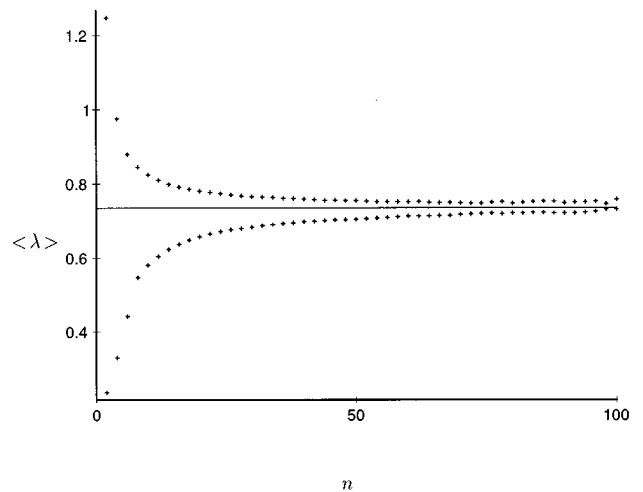


FIG. 12. Finding $\langle \lambda \rangle$ for the Farey exit map using the Monte Carlo method.

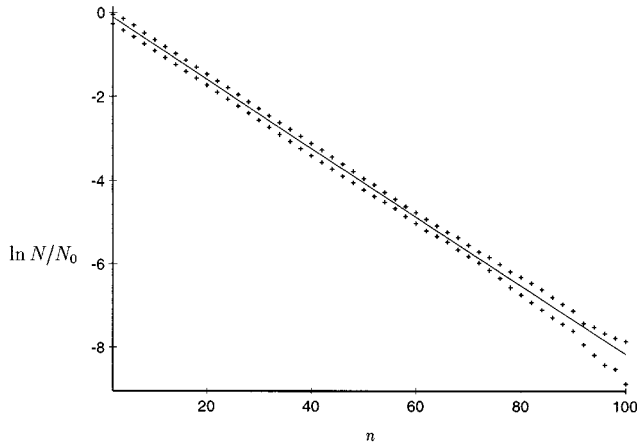


FIG. 13. Finding $\langle \tau \rangle$ for the Farey exit map using the Monte Carlo method. The decay time is found from the slope of the line of best fit.

V. COMPARING TO THE GAUSS MAP

Barrow used the Gauss map developed by BKL to derive dynamical systems properties such as metric and topological entropies. In this section we compare Barrow's results for the Gauss map to our results for the Farey map.

There are two critical features which distinguish the treatment of the maps. First, the Farey and Gauss maps are topologically inequivalent. Second, the techniques used to study the maps are different. We begin by isolating the strange repellor before using the repellor to reconstruct the properties of typical aperiodic trajectories. In contrast, Barrow worked directly with the aperiodic trajectories of the Gauss map and did not consider the repellor.

The Gauss map can be obtained from the Farey map by chopping out the oscillations. This was first done by BKL when they realized that the sensitive dependence on initial conditions was due to the bounces and not the oscillations. The Gauss map emerges from the Farey u map as follows: If $u_n \geq 2$, then oscillations will take place until $u_n < 2$. The bounce sequence that follows can then be expressed as

$$u_{N+1} = \frac{1}{u_N - [u_N]}, \quad (5.1)$$

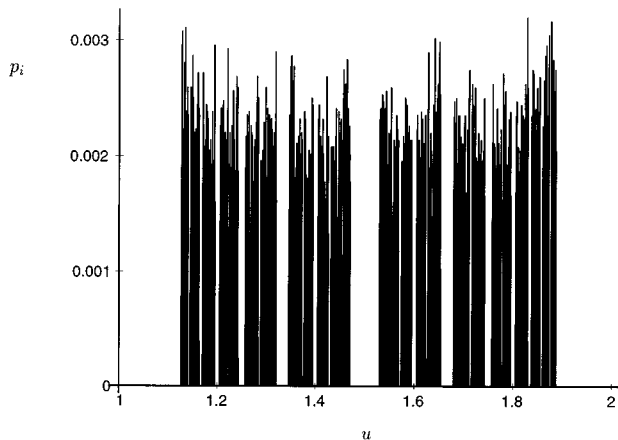


FIG. 14. A histogram of points shadowing the Farey exit map's future invariant set found by Monte Carlo methods ($u_{\text{exit}} = 8$).

where N is the number of bounces. Eq. (5.1) is just the Gauss map $G(u)$. The Gauss map is not invertible since it is derived from $F(u)$ and not the full $F(u, v)$.

The Farey and Gauss maps correspond to different discretizations of the full dynamics. The discrete ticks of the Farey clock correspond to periodic behavior in terms of the continuous time variable $T = \ln \ln(1/t)$. This is not true for the Gauss map. The time between bounces for the Gauss map can vary dramatically relative to T time. In compactifying the discrete time coordinate, $t_F = n$, of the Farey map to arrive at the discrete time coordinate, $t_G = N$, of the Gauss map, holes have had to be cut. There is no smooth coordinate transformation connecting the time variables t_F and t_G . As a result, the F and G maps are topologically inequivalent and have different topological entropies.

A. Repellor

Although topologically inequivalent, there are deep similarities between the two maps F and G . For one, they have identical fixed points. Still, some features of the repellor are altered.

Consider orbits with a Gauss period $P_G = 1$, so that the fixed points are obtained from

$$\bar{u}_1 = \frac{1}{\bar{u}_1 - [\bar{u}_1]}. \quad (5.2)$$

The solutions are

$$\bar{u}_1 = \frac{m_1 + \sqrt{m_1^2 + 4}}{2}, \quad (5.3)$$

where, again,

$$m_1 = [\bar{u}_1]. \quad (5.4)$$

Notice, that these \bar{u}_1 correspond to the silver means of Eq. (3.15). Hence, the period $P_G = 1$ orbit of the Gauss map generates all of the single bounce orbits of the Farey map, but with the longer period $P_F = m_1$. Notice also that there are already an infinite number of fixed points for the Gauss map at order $N = 1$.

We can calculate the Lyapunov exponents for the $P_G = 1$ orbits:

$$\lambda^G(\bar{u}) = \lim_{N \rightarrow \infty} \frac{1}{N} \ln(\bar{u})^{2N} \quad (5.5)$$

$$= 2 \ln \frac{m_1 + \sqrt{m_1^2 + 4}}{2}. \quad (5.6)$$

And so the silver mean exponents for the two maps are related by

$$\lambda^G(\bar{u}) = m_1 \lambda^F(\bar{u}). \quad (5.7)$$

The exponent for G is larger than that for F .

Continuing in the same vein, consider the period $P_G = 2$ fixed points. The condition

$$G^2(\bar{u}) = \bar{u} \quad (5.8)$$

generates the doubly infinite set of roots

$$\bar{u} = \begin{cases} \frac{m_1 + \sqrt{m_1^2 + 4m_1/m_2}}{2}, \\ 2 \\ \frac{2}{\sqrt{m_1^2 + 4m_1/m_2} - m_1}. \end{cases} \quad (5.9)$$

Notice that all the $P_G=1$ orbits are included in this solution when $m_1=m_2$. By considering the Gauss map we have been able to solve for all of the two-bounce fixed points of the Farey map at order $k=m_1+m_2$. These periodic orbits have two bounce epochs separated by an m_1-1 oscillation era and an m_2-1 oscillation era. The Lyapunov exponents for the period-2 orbits are given by

$$\lambda^G(\bar{u}) = \ln\left(\frac{m_0}{\sqrt{m_1^2 + 4m_1/m_2}} + 1\right). \quad (5.10)$$

These are related to the Lyapunov exponents for a corresponding periodic orbit of the Farey map by

$$\lambda^G(\bar{u}) = \left(\frac{m_1+m_2}{2}\right) \lambda^F(\bar{u}). \quad (5.11)$$

In terms of the CFE, the period of the Gauss map is the length of the generator for the periodic irrational. For example, the periodic irrational $[1, \{m_1, m_2, \dots, m_N\}]$ has $P_G=N$ and $P_F=\sum_{i=1}^N m_i$. From this we can deduce the general result

$$\frac{\lambda^G}{\lambda^F} = \frac{P_F}{P_G} = \frac{\sum_{i=1}^N m_i}{N}. \quad (5.12)$$

Interestingly enough, while the Lyapunov exponents and the period of the two maps differ, their product is invariant:

$$\lambda^G P_G = \lambda^F P_F. \quad (5.13)$$

Since typical periodic orbits of the Farey map have equal numbers of bounces and oscillations, it follows that the average Lyapunov exponents for the periodic orbits satisfy

$$\lambda_p^G = 2\lambda_p^F \approx 1.59. \quad (5.14)$$

A direct calculation of λ_p^G confirms this expectation.

B. Comparing entropies

In terms of continued fractions it is a simple matter to write down all the periodic orbits of the Gauss map. The period-1 orbits are given by $\bar{u}=[\{m_1\}]$, the period-2 orbits by $\bar{u}=[\{m_1, m_2\}]$, etc. From this it follows that the number of orbits with period $P_G=K$ scales as $N(K) \sim \infty^K$. Because the number of roots at each order is infinite, a regularization procedure has to be introduced before quantities such as H_T and $h(\mu)$ can be calculated.

After employing a suitable regulator [8], the measure (4.16) can be expressed in terms of an integral over $x=1/u$:

$$\mu(S) = \int_S \rho(x) dx, \quad (5.15)$$

where

$$\rho(x) = \frac{1}{(1+x)\ln 2}. \quad (5.16)$$

Taking S to cover the entire region $x=[0,1]$ ($u=[1,\infty]$), we have $\mu=1$ and $\langle \tau \rangle = \infty$. From the Gauss map in terms of the x coordinate, $G(x)=x^{-1}-[x^{-1}]$, we find $G'(x)=-x^{-2}$ and

$$\langle \lambda \rangle = \int_0^1 \ln|G'(x)|\rho(x) dx = \int_0^1 \frac{-2\ln x}{(1+x)\ln 2} dx = \frac{\pi^2}{6\ln 2}. \quad (5.17)$$

This expression is smaller by a factor of $\ln 2$ from the usual expression. The discrepancy can be traced to Barrow's use of \log_2 rather than natural logarithms. Similarly, the topological entropy of the Gauss map is given by

$$H_T^G = \frac{\pi^2}{6\ln 2}, \quad (5.18)$$

where again our expression differs from Barrow's oft-quoted result by a factor of $\ln 2$. As promised, the topological entropy of the Gauss map greatly exceeds that of the Farey map. Moreover, the metric entropy of the Gauss map, $h^G = \langle \lambda \rangle - \langle \tau \rangle^{-1} = \pi^2/(6\ln 2)$, is finite while the metric entropy of the Farey map vanishes.

C. Fractal dimension of the Gauss repeller

Although the Gauss map and the Farey map share common fixed points, they differ in how the fixed points are distributed as a function of period. It is impossible to produce a histogram in analogy to that of Fig. 1 since the Gauss map generates an infinite number of fixed points at every order. We can, however, argue that it is very likely that the fractal generated by the Gauss map is uniform and not multifractal. In other words, all the D_q 's are likely the same. Indeed, we can argue that the fractal dimensions saturate the phase space dimension, $D_q=D=1$, for all q , the reason being that the periodic irrationals are dense on the number line. We have already argued in Sec. III A that this ensures $D_0=2$ for the Farey map.

Let us consider this argument more closely. As P_G becomes large, the number of terms in the CFE cycle becomes large also. The longer the CFE cycle, the more uniform the coverage of the number line. For orbits of period P , the gaps between roots are always less than $1/f_p^2$, where

$$f_p = \frac{1}{\sqrt{5}} [\gamma^{(P+1)} + (-1)^P \gamma^{-(P+1)}]. \quad (5.19)$$

Here $\gamma = (\sqrt{5}+1)/2$ is the golden mean. For example, by the time $P=10$, $f_{10}=89$ and the largest possible gap is bounded by ≈ 0.0001 . In comparison, the Farey map at order $k=10$ has a finite number of roots and gaps as large as ≈ 0.2 . Moreover, the gaps close exponentially fast with P for the Gauss

map, while they close slower than $1/k$ for the Farey map. While the rate of closure makes no difference to D_0 , the multifractal dimensions with higher q 's take into account the density of points. The Gauss map lays down a very even covering of the number line, and so we expect $D_0 = D_1$, etc.

If the conjecture that $D_1 = D_L$ holds true, then we should find $D_L = 1$. This is in fact what we find. For the 1D map,

$$D_L = 1 - \frac{1}{\langle \lambda \rangle \langle \tau \rangle}. \quad (5.20)$$

The Gauss map is characterized by $\langle \tau \rangle = \infty$ and $\langle \lambda \rangle = \pi^2 / (6 \ln 2)$. Inserting these values into Eq. (5.20) we find $D_L = 1$, in accordance with our argument that $D_1 = 1$.

As with the topological entropies, we find the fractal dimensions of the two maps are inequivalent. This does not invalidate the statement that fractal dimensions are coordinate-invariant signals of chaos. The two maps cannot be connected by a smooth coordinate transformation. They are, as we have already argued, topologically inequivalent. The discrete Farey time is topologically the same as the $T = \ln \ln(1/t)$ of the full dynamics. The discrete Gauss time, on the other hand, changes the topology of time by chopping out the oscillations.

VI. SYMBOLIC DYNAMICS

Before embarking on a numerical study of the continuum dynamics, we can gain additional analytic insight into the mixmaster universe by studying symbolic representations of the unstable periodic orbits. A chaotic system is characterized by the complexity of the representation required to describe its periodic orbits. The complexity is quantified by measuring the Rényi entropy spectrum [44] of the coding sequence. The most commonly measured Rényi entropies are the Shannon information entropy and the Hartley topological entropy.

There are several features that recommend symbolic or information dynamics as a method of studying chaos in general relativity: The description is coordinate independent, approximate maps are not required, and only the global behavior of trajectories need be known. The global structure of trajectories can be found by approximate methods so no computers are required.

In the case of the mixmaster we need to know how trajectories move around the anisotropy plane (β_+, β_-) . Here we are in luck as the work has been done for us by Bogoyavlenskii and Novikov [28]. From their work we see that a simple symbolic coding can be defined by recording crossings through three half-infinite lines in the anisotropy plane. The lines are

$$\begin{aligned} \text{(I)} \quad & \beta_- = 0, \quad \beta_+ < 0, \\ \text{(II)} \quad & \beta_- = \sqrt{3}\beta_+, \quad \beta_+ > 0, \\ \text{(III)} \quad & \beta_- = -\sqrt{3}\beta_+, \quad \beta_+ > 0. \end{aligned} \quad (6.1)$$

Whenever a trajectory crosses line (I) in a clockwise sense, the symbol X is recorded. An anticlockwise crossing of line (I) is recored with the symbol \tilde{X} . Similarly, crossings of lines (II) and (III) are recorded using Y and Z .

The symbolic length l of an orbit is defined to be the number of symbols required to describe the orbit. The shortest orbits are of the form $X\tilde{X}$. These describe a trajectory bouncing from one corner to another and then back again. There are six orbits of this type: $X\tilde{X}$, $\tilde{X}X$, $Y\tilde{Y}$, $\tilde{Y}Y$, $Z\tilde{Z}$, $\tilde{Z}Z$. In terms of the Farey coding, these orbits are of the form $BOBO$ or $BOOBO$, etc. The next shortest orbits are the two complete cycles XYZ and $\tilde{X}\tilde{Y}\tilde{Z}$. These correspond to Farey codings of the form BBB or $BOOBB$, etc. All longer orbits can be constructed by sewing together various combinations of order-2 and order-3 orbits.

Simple rules govern the sewing together of orbits. New primary orbits can be sewn onto the active sites of order- $(l-3)$ or $-(l-2)$ symbolic sequences to form the order- l coding sequences. By an active site we mean the segment of the coding string of length $l-3$ or $l-2$ that was added at the previous iteration. In other words, our coding tree only grows out from the tips of its branches. If the active site is of the form $\dots XYZ$, then new primary orbits can be added between the X and the Y , or the Y and the Z , or after the Z . Thus, a period-3 orbit provides three sites for adding a new orbit. Similarly, if the active site is of the form $\dots X\tilde{X}$, then new primary orbits can be inserted between the X and the \tilde{X} , or after the \tilde{X} , so a period-2 orbit only provides two sites for sewing on new orbits. Combining this information with our knowledge about how many new orbits may be added at each site, we are able to write down a recurrence relation that describes the growth in the number of orbits. Defining $P(l)$ to be the number of active period-2 orbits at order l and $Q(l)$ to be the number of active period-3 orbits at order l , then the total number of orbits at order l is $N(l) = P(l) + Q(l)$. The proliferation in the number of orbits follows from the coupled recurrence relations

$$\begin{aligned} P(l) &= 4P(l-2) + 6Q(l-2), \\ Q(l) &= 4P(l-3) + 6Q(l-3). \end{aligned} \quad (6.2)$$

The initial conditions for the recurrence relations are found by directly counting the number of orbits at order $l = 1, 2, 3$:

$$\begin{aligned} P(1) &= 0, \quad P(2) = 6, \quad P(3) = 0, \\ Q(1) &= 0, \quad Q(2) = 0, \quad Q(3) = 2, \\ N(1) &= 0, \quad N(2) = 6, \quad N(3) = 2. \end{aligned} \quad (6.3)$$

The solution to Eqs. (6.2) is given by

$$N(l) = c_1 a_1^l + c_2 a_2^l + c_3 a_3^l, \quad (6.4)$$

where the a_i are the three roots of the cubic equation

$$a^3 - 4a - 6 = 0, \quad (6.5)$$

and the c_i can be fixed using Eqs. (6.3). The Hartley entropy of the symbolic coding is then

$$H_H = \lim_{l \rightarrow \infty} \frac{1}{l} \ln N(l) = \ln(a_1), \quad (6.6)$$

where a_1 is the largest root of Eq. (6.5). Thus,

$$H_H = \ln \left(\frac{12 + (81 + 3\sqrt{537})^{2/3}}{3(81 + 3\sqrt{537})^{1/3}} \right) \approx 0.926. \quad (6.7)$$

The Hartley entropy of our symbolic coding differs from the topological entropy of the Farey or Gauss maps since their symbolic codings record different information about the orbits and assign different lengths to the orbits. There are many different ways we can assign a symbolic coding to the mixmaster orbits, and most have different symbolic entropies. However, since the rules governing a coding can be given in a coordinate-invariant way, the symbolic entropy of each coding is a coordinate-invariant quantity. As explained below, the dynamics is chaotic if the entropy of any (valid) coding is positive.

The symbolic coding we chose tended to throw out some information as oscillations that occur in the corner channels were not recorded. To record all the information needed to uniquely characterize a mixmaster orbit would require the introduction of three additional half-infinite lines extending into the corner channels. The six symbols associated with crossings of these lines, in addition to the six symbols used here, provide a complete alphabet for the mixmaster dynamics.

While our coding is not unique, it is onto; that is, a physical orbit belonging to the mixmaster's strange repeller has *only one* symbolic coding in terms of $\{X, Y, Z, \tilde{X}, \tilde{Y}, \tilde{Z}\}$. The converse is not true since many physical orbits share the same symbolic coding. Using the full 12-letter alphabet it is possible to arrive at a unique symbolic coding that is one to one. It would be nice to find the one-to-one coding as the Hartley entropy of this coding would provide a unique upper bound to the Hartley entropy of any coding of the mixmaster's strange repeller. In practice a one-to-one coding can be very hard to find so we often coarse-grain the available information and look for codings that are merely onto. These then provide lower bounds to the true Hartley entropy of the invariant set. Thus, we have shown that the continuum mixmaster dynamics is chaotic since the Hartley entropy of the repeller exceeds 0.926.

VII. NUMERICAL RESULTS FOR THE FULL DYNAMICS

To be certain the Farey map provides an accurate description, we now expose the strange repeller in the minisuperspace of the full continuum dynamics.

Continuous dynamical systems are much harder to solve than discrete dynamical systems. We cannot hope to find the periodic orbits analytically. Instead, we apply the same kind of Monte Carlo approach used to check the analytic results for the Farey map. Our ambition will be limited here to finding the fractal manifest in minisuperspace phase space and the fractal dimension.

Exposing the chaotic invariant set in a dissipative system is simple. By randomly choosing a collection of initial conditions and evolving them according to the dynamical equations, the trajectories will soon settle onto the various attractors in phase space. The decrease in available phase space volume forces all trajectories onto the chaotic attractors. For a Hamiltonian system this is not the case. Phase space volume is conserved and the chaotic invariant set remains hid-

den amongst the slurry of aperiodic trajectories. Nonetheless, the chaotic invariant set can be uncovered by a variety of methods. For compact Hamiltonian systems the proper interior maxima (PIM) procedure [45] efficiently reveals the strange repeller while for noncompact Hamiltonian systems, including Hamiltonian exit systems [41], fractal basin boundaries are the preferred method.

A. PIM procedure

The proper interior maxima procedure is able to isolate the strange repeller in most Hamiltonian systems. In particular, it could be applied to the mixmaster system with or without exits. Although we did not apply the PIM procedure to the mixmaster, we mention it here as an alternative to introducing exits.

The basic idea behind the PIM procedure is ingeniously simple. As discussed earlier, the stable and unstable manifolds of the strange repeller are interchanged under time reversal. As time moves forward trajectories near the repeller are repelled along the unstable manifold and attracted along the stable manifold. Evolving the system into the past reverses the attracting and repelling directions. The PIM procedure exploits this property in the following way: (a) Evolve a collection of trajectories near the strange repeller into the future and into the past by an equal amount of time. (b) Isolate subsets of the evolved bundles that remain nearby in both the past and the future and discard the remainder. (c) Zero in on the surviving trajectories and form a new, smaller bundle of trajectories around them and repeat the procedure. After a few iterations the surviving trajectories will closely shadow the invariant set as noninvariant trajectories have been discarded.

The PIM procedure is shown schematically in Fig. 15. To simplify the picture, the initial bundle of trajectories is shown centered on a point belonging to the strange repeller (black dot). The labels (a) and (b) correspond to the steps in the PIM procedure. The great feature of the PIM procedure is the way it uses the instability of the repeller's orbits to its advantage. Consider the original set of trajectories in the circle of radius R about the invariant point. Evolving into the future and into the past by an amount Δt leads to ellipses with axes of length $R \exp(-\langle \lambda \rangle \Delta t)$ and $R \exp(+\langle \lambda \rangle \Delta t)$. These ellipses share a Δt invariant set of radius $r \approx R \exp(-\langle \lambda \rangle \Delta t)$. Thus, after a few iterations, the repeller lies exposed. In effect, the PIM procedure turns the strange repeller into a strange attractor.

Using a variant of the PIM procedure entire trajectories belonging to the strange repeller can be reconstructed. While it would be valuable to apply these techniques to the mixmaster system, the numerical implementation of the procedure is difficult. A much simpler method of exposing the strange repeller is to introduce exits, and then chart the fractal coastline of the outcome basins.

B. Fractal basin boundaries

Rather than hunting for the chaotic skeleton among the aperiodic trajectories, the Hamiltonian exit method creates a sieve through which the slurry of aperiodic trajectories escape, leaving the strange repeller in clear view. Like the PIM procedure, the introduction of exits exploits the instability of

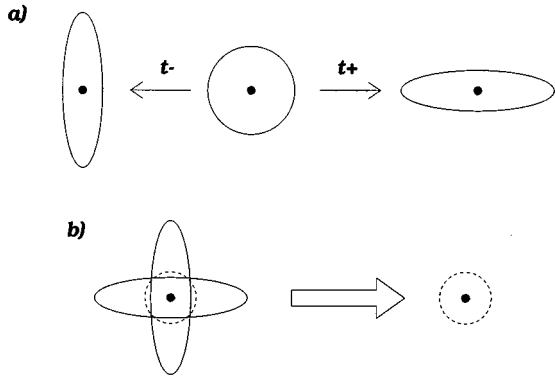


FIG. 15. Finding the invariant set using the PIM procedure.

the chaotic trajectories. The faster trajectories are expelled by the repeller, the faster the invariant set is exposed.

The repeller is manifest at the boundary in phase space which separates initial conditions on the basis of their outcomes. These basin boundaries often become fractal in chaotic scattering. The fractal basin boundary is composed of trajectories which spend substantial time on the repeller.

For the full mixmaster dynamics we employ the lessons of the Farey map. Indeed, it is a simple matter to directly implement exactly the same exit conditions we used for the map. By considering the ratios $(\ln a)' / (\ln b)'$ *et cyc.* (a, b, c) , we can test to see if a mixmaster universe is coasting in a Kasner phase. If it is, we can then use Eq. (2.5) to read off the value of u . When $u > u_{\text{exit}}$ we terminate the evolution and assign an outcome based on which axis is collapsing most quickly. By color coding the initial conditions near the maximum of expansion according to their outcome near the big crunch singularity, we produce plots of the outcome basins.

Using a fourth order Runge-Kutta integrator with adaptive step size, we evolved 300×300 grids of initial conditions and recorded their outcomes. The Hamiltonian constraint, Eq. (2.1) was monitored at all times, and the error tolerances were adaptively corrected to ensure the constraint was satisfied to within 1 part in 10^5 along each trajectory.⁴ In order to make contact with the Farey map, the initial conditions were chosen by fixing $\omega_0 = 1/3$ and taking (u_0, v_0) from a 300×300 grid. The value of Ω_0 was then found by solving the Hamiltonian constraint. Inserting these values of $(u_0, v_0, \Omega_0, \omega_0)$ into Eq. (2.3) yielded the initial conditions for the numerical integration of the equations of motion.

A portion of the basin boundaries in the (u, v) plane is displayed in Fig. 16. Depending on which axis is collapsing most quickly when the trajectory escapes, the initial grid point is colored black for a , grey for b , and white for c . The exit was set at $u_{\text{exit}} = 8$ in order to make comparison with the Farey exit map results. The basins form an intricately woven tapestry of roughly vertical threads. The basin boundaries appear to form a cantor set of vertical lines. By zooming in on a small portion of Fig. 16, we see from Fig. 17 that the dense weave persists on finer and finer scales.

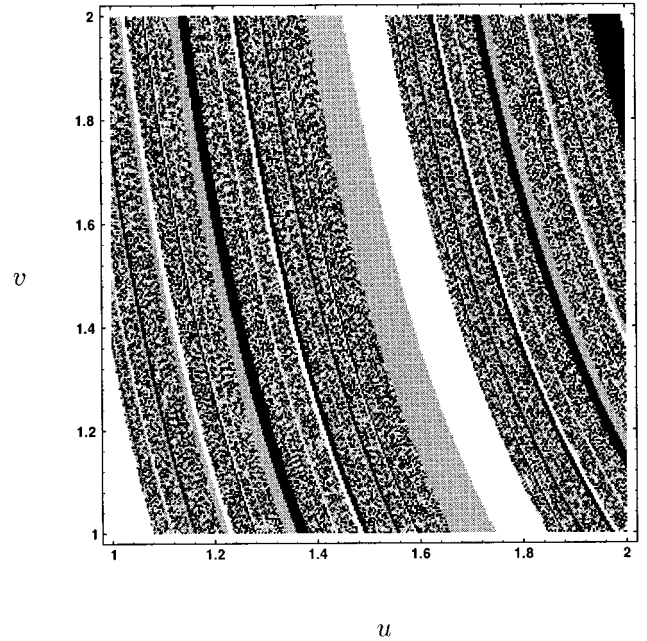


FIG. 16. Basin boundaries in the (u, v) plane for the full dynamics.

Points belonging to the fractal basin boundaries comprise a future invariant set. Trajectories belonging to this set never “decide” on a particular outcome, and so never escape through an exit. These are the trajectories that form the unstable manifold of the strange repeller. It is instructive to compare the future invariant set seen in Figs. 16 and 17 with the Farey map’s future invariant set shown in Fig. 2. The pattern of gaps and dense regions is strikingly similar. The warpage of the vertical stripes in Figs. 16 and 17 can be accounted for by our choice of starting point. Because our initial conditions start the universe near the maximum of its expansion, we are looking at the future invariant set in a region where the approximations used to derive the Farey map are very poor. Nonetheless, as the trajectories evolve toward the big crunch the continuum dynamics settles onto a pattern well described by the Farey map. Since the fine structure of the weave is laid down after many bounces and oscillations, the fractal structure produced by the full equations should be much the same as that produced by the Farey map.

In order to test this proposition we numerically evaluate the fractal dimension of Fig. 17. To do this we employ the uncertainty exponent method described in Sec. III A. Randomly choosing 1000 points in the region covered by Fig. 17, we record how many have certain outcomes as a function of initial uncertainty δ . A plot of $\ln f(\delta)$ as a function of $\ln \delta$ is shown in Fig. 18. According to Eq. (3.30), the gradient of this graph yields the uncertainty exponent $\alpha = 0.14 \pm 0.01$ and a capacity dimension of $D_0^u = 1.86 \pm 0.01$.

The method described above delivers the capacity dimension of a sample of the repeller. The capacity dimension of a sample often is not equal to the capacity dimension of the full repeller. Rather, the capacity dimension of the sample equals the information dimension of the full repeller [46]. The reason is the following. The random sampling of the Monte Carlo approach favors the densest regions of the frac-

⁴We check the normalized constraint $\tilde{H} = H/[a^4 + b^4 + c^4 + 2(b^2c^2 + a^2c^2 + b^2a^2)]$.

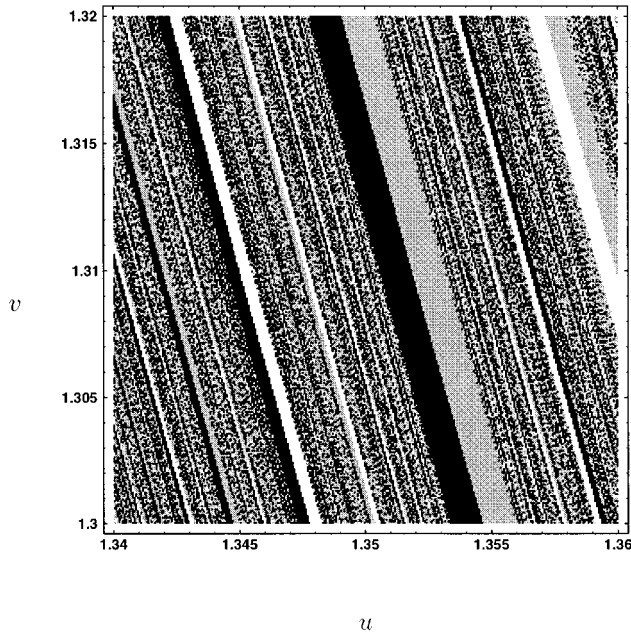


FIG. 17. A portion of Fig. 16 magnified 50 times.

tal. The method naturally weights therefore the densest regions. Since the weighted dimension is in fact the information dimension, it follows that D_0^u of the sample actually equals D_1^u . We have therefore really found the information dimension of the mixmaster’s future invariant set:

$$D_1^u = 1.86 \pm 0.01. \tag{7.1}$$

The information dimension of the repellor from these numerical experiments is then given by

$$D_1 = 2D_1^u - D = 1.72 \pm 0.02. \tag{7.2}$$

Within numerical uncertainty, this number agrees with what we found for the Farey map [Eqs. (3.32) and (4.22)]. The small discrepancy is probably due to the approximations used in deriving the map or systematic numerical errors.

It is worth noting that the fractal basin boundaries can be uncovered in any 2D slice through the 6D phase space. For example, in Fig. 19 we display the outcome basins in the anisotropy plane (β_+, β_-) . The initial conditions were chosen by setting $\Omega = -1$, $\dot{\alpha} = 10$, $\dot{\beta} = 10$, and selecting (β_+, β_-) from a 600×600 grid. The remaining coordinate $\dot{\gamma}$ was fixed by the Hamiltonian constraint. Again we see an interesting mixture of regular and fractal boundaries, this time forming a symmetrical patchwork about a central axis of symmetry. The symmetry of the basins reflects the symmetry of the MSS potential shown in Figs. 6 and 7.

By choosing different slices or different coordinate systems, many different views of the future invariant set can be uncovered. However, these are purely cosmetic changes. No matter what coordinates we choose there will always be fractal basin boundaries, and these boundaries will always have the same fractal dimension. No matter how you look at it, the mixmaster universe is chaotic.

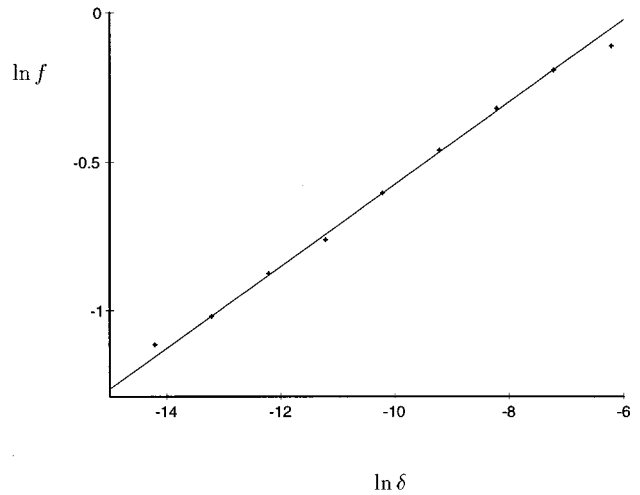


FIG. 18. Calculating the uncertainty exponent for Fig. 17.

VIII. SUMMARY OF RESULTS AND DISCUSSION

The power of the fractal in relativity is its observer independence. We isolate two related fractals. The strange repellor of the Farey map is shown to be a multifractal and analyzed in detail. The fractal repellor is then excavated numerically in the phase space of the full, unapproximated dynamics:

| Map | H_T | $\langle \lambda \rangle$ | λ_p | D_0 | D_1 | D_L |
|-------|-------------------------|---------------------------|-------------|-------|-----------|-------|
| Farey | $2 \ln 2$ | 0 | 0.793 | 2.0 | 1.74–1.78 | 1.78 |
| Gauss | $\frac{\pi^2}{6 \ln 2}$ | $\frac{\pi^2}{6 \ln 2}$ | 1.59 | 1.0 | 1.0 | 1.0 |

The full force of the conclusions comes with a comparison of the dimensions found in entirely different manners. We have collected our results for the discrete time maps in

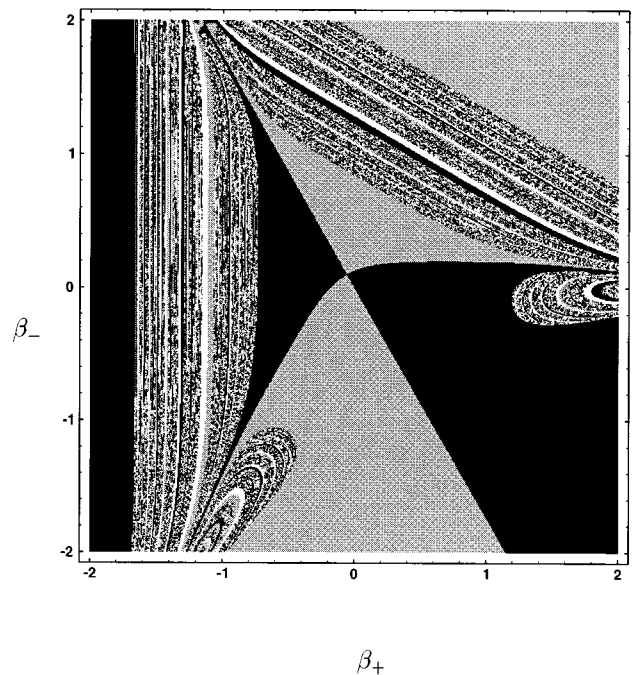


FIG. 19. The outcome basins in the anisotropy plane.

the chart above. For comparison with the chart, the information dimension of the fractal basin boundary in the full dynamics is $D_1 = 1.72 \pm 0.02$. Not only is a fractal a coordinate-invariant declaration of chaos, we have further found that the three information dimensions agree within errors: the information dimension of the repeller from the discrete Farey map, the Lyapunov dimension of the Farey map, and the information dimension of the fractal basin boundaries. As has long been suspected, the collective of mixmaster universes is certainly chaotic.

In a sense, chaos in the mixmaster universe makes it impossible to create one in the first place. The mixmaster is anisotropic but homogeneous. The extreme sensitivity to initial conditions induces an instability to inhomogeneities. Two different regions of spacetime with even slightly different initial conditions quickly evolve away from each other

[47]. The universe then becomes inhomogeneous as well as anisotropic, rendering the big crunch similar to a generic inhomogeneous collapse to a black hole. The churning dynamics on approach to a singularity may therefore be temporally and spatially chaotic, whether it be in a dying star or a dying cosmos.

ACKNOWLEDGMENTS

We would like to express our gratitude to J. Barrow, C. Dettmann, N. Frankel, P. Ferreira, and J. Silk for their comments and valuable discussions. We also thank C. Dettmann for letting us adapt his computer codes. N.C. thanks Case Western Reserve University for its hospitality and generous use of its computers.

-
- [1] E. M. Lifshitz and I. M. Khalatnikov, *Sov. Phys. Usp.* **6**, 495 (1963).
- [2] I. M. Khalatnikov and E. M. Lifshitz, *Phys. Rev. Lett.* **24**, 76 (1970); V. A. Belinskii, *Adv. Phys.* **19**, 525 (1970).
- [3] B. K. Berger, in *General Relativity and Gravitation, Proceedings of GR14*, Report No. gr-qc/9512003, 1995 (unpublished).
- [4] T. Dittrich and R. Graham, in *Information Dynamics*, edited by H. Atmanspacher and H. Scheingraber (Plenum, New York, 1991).
- [5] *Deterministic Chaos in General Relativity*, edited by D. Hobill, A. Burd, and A. Coley (Plenum, New York, 1994).
- [6] J. D. Barrow and F. Tipler, *Phys. Rep.* **56**, 372 (1979).
- [7] C. W. Misner, *Phys. Rev. Lett.* **22**, 1071 (1969).
- [8] J. D. Barrow, *Phys. Rev. Lett.* **46**, 963 (1981); *Phys. Rep.* **85**, 1 (1982).
- [9] D. F. Chernoff and J. D. Barrow, *Phys. Rev. Lett.* **50**, 134 (1983).
- [10] S. E. Rugh, *Cand. Scient. thesis*, The Niels Bohr Institute, 1990.
- [11] G. Francisco and G. E. A. Matsas, *Gen. Relativ. Gravit.* **20**, 1047 (1988).
- [12] J. Pullin, presented at VII Simposio Latinoamericano de Relatividad y Gravitacion, 1990 (unpublished).
- [13] B. K. Berger, *Gen. Relativ. Gravit.* **23**, 1385 (1991).
- [14] D. Hobill, D. Bernstein, D. Simkins, and M. Welge, *Class. Quantum Grav.* **8**, 1155 (1991).
- [15] C. P. Dettmann, N. E. Frankel, and N. J. Cornish, *Phys. Rev. D* **50**, R618 (1994); *Fractals* **3**, 161 (1995).
- [16] N. J. Cornish and J. J. Levin, *Phys. Rev. D* **53**, 3022 (1996).
- [17] N. J. Cornish and J. J. Levin, *Phys. Rev. Lett.* **78**, 998 (1997).
- [18] B. B. Mandelbrot, *Nature of Fractal Geometry* (Freeman, San Francisco, 1982).
- [19] D. N. Page, *Class. Quantum Grav.* **1**, 417 (1984).
- [20] M. Widom, D. Bensimon, L. P. Kadanoff, and S. J. Shenker, *J. Stat. Phys.* **32**, 443 (1983); L. P. Kadanoff and C. Tang, *Proc. Natl. Acad. Sci. USA* **81**, 443 (1984).
- [21] D. Auerbach, P. Cvitanović, J. P. Eckmann, G. Gunaratne, and I. Procaccia, *Phys. Rev. Lett.* **58**, 2387 (1987).
- [22] A. Latifi, M. Musette, and R. Conte, *Phys. Lett. A* **194**, 83 (1994); G. Contopoulos, B. Grammaticos, and A. Ramani, *J. Phys. A* **27**, 5357 (1994).
- [23] C. W. Misner, K. S. Thorne, and J. A. Wheeler, *Gravitation* (Freeman, New York, 1973).
- [24] D. Mayer, *Phys. Lett. A* **121**, 390 (1987).
- [25] A. B. Burd, N. Buric, and R. K. Tavakol, *Class. Quantum Grav.* **8**, 123 (1991).
- [26] B. K. Berger, *Phys. Rev. D* **47**, 3222 (1993).
- [27] R. Cushman and J. Sniatycki, *Rep. Math. Phys.* **36**, 75 (1995).
- [28] O. I. Bogoyavlenskii and S. P. Novikov, *Sov. Phys. JETP* **37**, 747 (1973).
- [29] E. Ott, *Chaos in Dynamical Systems* (Cambridge University Press, Cambridge, England, 1993).
- [30] E. Ott, W. D. Withers, and J. A. Yorke, *J. Stat. Phys.* **36**, 687 (1984).
- [31] S. W. McDonald, C. Grebogi, E. Ott, and J. A. Yorke, *Physica D* **17**, 125 (1985); C. Grebogi, E. Kostelich, E. Ott, and J. A. Yorke, in *Dynamical Systems*, edited by J. C. Alexander (Springer-Verlag, Berlin, 1988).
- [32] We thank Carl Dettmann for pointing this out to us.
- [33] See, e.g., M. Schröder, *Fractals, Chaos, Power Laws: Minutes from an Infinite Paradise* (Freeman, New York, 1991).
- [34] The connection to the *rational* Farey tree was noted by Rugh and can be found in his article in Ref. [5].
- [35] R. M. Wald, *General Relativity* (The University of Chicago Press, Chicago, 1984).
- [36] L. S. Young, *Ergod. Th. Dynam. Syst.* **2**, 109 (1982); A. Fathi, *Commun. Math. Phys.* **126**, 249 (1989).
- [37] H. Kantz and P. Grassberger, *Physica D* **17**, 75 (1985); T. Bohr and D. Rand, *ibid.* **25**, 387 (1987); G. H. Hsu, E. Ott, and C. Grebogi, *Phys. Lett. A* **127**, 199 (1988); Z. Kovacs and L. Wiesenfeld, *Phys. Rev. E* **51**, 5476 (1995).
- [38] J. L. Kaplan and J. A. Yorke, in *Functional Differential Equations and Approximations of Fixed Points*, edited by H. O. Peitgen and H. O. Walter (Springer, Berlin, 1979).

- [39] C. Grebogi, E. Ott, and J. A. Yorke, *Phys. Rev. A* **37**, 1711 (1988).
- [40] D. M. Chitre, Ph.D. thesis, University of Maryland, 1972.
- [41] S. Bleher, C. Grebogi, E. Ott, and R. Brown, *Phys. Rev. A* **38**, 930 (1988).
- [42] T. D. Creighton and D. W. Hobill, in *Deterministic Chaos in General Relativity* [5], p. 433.
- [43] H. Fujisaka, *Prog. Theor. Phys.* **70**, 1264 (1983).
- [44] A. Rényi, *Acta Math. Acad. Sci. Hung.* **10**, 193 (1959).
- [45] H. E. Nusse and J. Yorke, *Physica D* **36**, 137 (1989).
- [46] J. D. Farmer, E. Ott, and J. A. Yorke, *Physica D* **7**, 153 (1983).
- [47] G. Montani, *Class. Quantum Grav.* **12**, 2505 (1995).