Constraints on supersymmetric soft phases from renormalization group relations

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(Received 2 October 1996)

By using relations derived from renormalization group equations (RGEs), we find that strong indirect constraints can be placed on the top squark mixing phase in A_t from the electric dipole moment of the neutron (d_n) . Since m_t is large, any GUT-scale phase in A_t feeds into other weak scale phases through RGEs, which in turn contribute to d_n . Thus *CP*-violating effects due to a weak-scale A_t are strongly constrained. We find that $|\text{Im } A_t^{\text{EW}}|$ must be smaller than or of order $|\text{Im } B^{\text{EW}}|$, making the electric dipole moment of the top quark unobservably small in most models. Quantitative estimates of the contributions to d_n from A_u , A_d , and *B* show that substantial fine-tuning is still required to satisfy the experimental bound on d_n . While the low energy phases of the *A*'s are not as strongly constrained as the phase of B^{EW} , we note that the phase of a universal \widehat{A}^{GUT} induces large contributions in the phase of B^{EW} through RGEs, and is thus still strongly constrained in most models with squark masses below a TeV. $[$ S0556-2821(97)03703-X $]$

PACS number(s): 12.60.Jv, 11.10.Hi, 11.30.Er, 12.10.Kt

I. INTRODUCTION

Supersymmetry $(SUSY)$ [1] is one of the most compelling extensions of the standard model (SM) . It is the only known perturbative solution to the naturalness problem $[2]$, it unifies the gauge coupling constants for the observed value of $\sin^2 \theta_W$, it allows radiative electroweak (EW) symmetry breaking, and the lightest SUSY partner provides a good dark matter candidate. SUSY models with such features are generally in excellent agreement with experiment, and there is even the possibility that a recent Collider Detector at Fermilab (CDF) event $[3]$ is of supersymmetric origin $[4]$.

One of the few phenomenological problems associated with SUSY models is their generically large predictions for the electric dipole moment (EDM) of the neutron, d_n . Supersymmetric models with universal soft breaking parameters have two physical phases, beyond the Cabibbo-Kobayashi-Maskawa (CKM) and strong phases of the SM, which can be taken to be the triscalar and biscalar soft breaking parameters *A* and *B*. These phases give a large contribution to d_n , of order 10^{-22} (100 GeV/ M_{SUSY})²e cm, where M_{SUSY} is a characteristic superpartner mass. The experimental upper bound on d_n is of order 10^{-25} *e* cm [5], so that if superpartner masses are near the weak scale, the phases of these complex soft parameters must be fine-tuned to be less than or of order 10^{-2} – 10^{-3} since there is no *a priori* reason for them to be small $[6]$. If one wants to avoid such a finetuning, there are two approaches: suppress d_n with very large squark masses (greater than a TeV) $[7]$, or construct models in which the new SUSY phases naturally vanish [8]. Models with very heavy squarks are unappealing because in such models lightest supersymmetric particle (LSP) annihilation is usually suppressed enough so that the relic density is unacceptably large [9]. They also lead to a fine-tuning problem of their own in getting the *Z* boson mass to come out right in EW symmetry breaking.

It is natural to consider solutions of the second type, and

demand that the soft phases be zero by some symmetry. While that would leave only a small CKM contribution to d_n $[10-13]$, and thus avoid any fine-tuning in meeting the experimental bound on d_n , it would also mean that there is no non-SM *CP* violation, which is needed by most schemes for electroweak baryogenesis [14]. Also, such models do not generate signals of non-SM *CP* violation, such as those involving top squark mixing. There are ways of naturally obtaining small nonzero soft phases which leave sufficient *CP* violation for baryogenesis $[15–18]$, but these phases would still have to meet the bounds from d_n and would probably be unobservably small in most EW processes—unless the soft terms are not universal.

Recently it has been pointed out that large non-SM *CP*violating top quark couplings could be probed at high energy colliders [19]. A measurement of a large top quark EDM, for example, would indicate physics beyond the SM, and it is interesting to ask whether SUSY models can yield an observable effect. Several references have attempted to use *CP* violation from top squark mixing due to the complex parameter A_t to yield large CP -violating effects in collider processes involving top quarks [20]. Such papers either explicitly or implicitly assume nonuniversal soft couplings A_q at the grand unified theory (GUT) scale, otherwise, the phase of A_t would be trivially constrained by d_n . We consider whether it is possible to obtain large effects due to the phase of *At* at the EW scale by relaxing the universality of *A*. We will show that due to renormalization-group-induced effects on other low energy phases, the phase of A_t is strongly constrained by d_n , and it is not possible, for most areas of parameter space, to have large *CP*-violating effects due to the imaginary part of A_t .

We will assume that no parameters are fine-tuned and thus we will require the phases at the GUT scale to be either identically zero (presumably through some symmetry) or no less than 1/10. If one permits an arbitrary degree of finetuning, the whole SUSY *CP* violation issue becomes moot,

and one can derive no constraints on the phase of A_t . While one can construct models which give small universal phases, as we said above, the fine-tuning needed to evade the constraints we derive is unlikely to be explained naturally. Our approach in this paper is to assume the reasonable finetuning criterion we have just outlined, and ask what it implies about low energy SUSY *CP*-violating phenomenology.

In Sec. II, we review the basics of SUSY *CP* violation. We present our results derived from renormalization group equations (RGEs) in Sec. III, and impose the neutron EDM constraints on ImA_t using those results in Sec. IV. In Sec. V we discuss top-squark mixing-induced *CP*-violating observables in more detail in light of our constraints on the phase of A_t , and we give some concluding remarks in Sec. VI. The details from Sec. III are written up in Appendix A, and the full one-loop calculation for the SUSY contribution to the neutron EDM is given in Appendix B.

II. SUSY *CP***-VIOLATING PHASES**

The soft breaking potential in the MSSM is

$$
-\mathcal{L}_{soft} = \frac{1}{2} |m_i|^2 |\varphi_i|^2 + \frac{1}{2} \sum_{\lambda} M_{\lambda} \lambda \lambda + \epsilon_{ij} [A_U \widetilde{U}_R^* Y_U \widetilde{Q}_L^i] H_u^j
$$

+
$$
\epsilon_{ij} [A_D \widetilde{D}_R^* Y_D^{\dagger} \widetilde{Q}_L^j] H_d^i + \epsilon_{ij} [A_E \widetilde{E}_R^* Y_E^{\dagger} \widetilde{L}_L^j] H_d^i
$$

+
$$
\epsilon_{ij} B_{\mu} H_u^i H_d^j + \text{H.c.},
$$

(1)

where we take $A_U = \text{diag}\{A_u, A_c, A_t\}, A_D = \text{diag}\{A_d, A_s, A_b\},$ and $A_E = \text{diag}\{A_e, A_\mu, A_\tau\}; Y_U, Y_D, \text{ and } Y_E$ are the Yukawa coupling matrices; \overline{Q} , \overline{L} , \overline{U}_R , \overline{D}_R , and \overline{E}_R are the squark and slepton fields; λ are the gauginos; and φ_i are the scalars in the theory.

A common simplifying assumption is that this soft Lagrangian arises as the result of a GUT-scale supergravity (SUGRA) model with universal soft triscalar coupling A, gaugino mass $M_{\lambda} = M_{1/2}$, and scalar mass $m_i = m_0$. This provides an explanation for the absence of flavor-changing neutral currents which arise from loops with squarks of nondegenerate mass [21]. Such supersymmetric models have only two independent physical *CP*-violating phases beyond the CKM and strong phases of the SM $[10]$ although these phases appear in several different linear combinations in low energy phenomenology $[17,22]$. We will take the two physical phases to be Arg *A* and Arg *B*.

It turns out that all *CP*-violating vertices in this model arise through the diagonalization of complex mass matrices [15]. The complex quantities which appear in these matrices are $A_q + \mu^* R_q$ and μ , where R_q is tan β (the ratio of Higgs vacuum expectation values) for $q=d$, *s*, *b* and cot β for $q=u,c,t$, and where the phase of μ is simply equal to the phase of B^* by a redefinition of fields. Thus for d_n , which involves only *u* and *d* quarks, there are only contributions from three low energy combinations of the two SUSY GUT phases: Arg $(A_d - \mu \tan \beta)$, Arg $(A_u - \mu \cot \beta)$, and Arg μ . (In Appendix B, a complete expression of d_n is given which includes suppressed contributions from phases of the other squark mixings.)

Even with universal boundary conditions, the elements of the matrices A_U , A_D , and A_E have distinct phases at the EW scale because of renormalization group evolution. We will also relax, in some places, the assumption that their phases started the same at the GUT scale. We assume (for simplicity) that these matrices are diagonal. One possible consequence of this approach is that one could have $d_n \approx 0$ because Im A_d and Im $A_u \approx 0$, but other A_q , notably A_t , could have large phases which lead to observable effects. These include angular correlations and polarizations $[20]$, including effects attributable to the electric dipole moment of the top quark, d_t . As discussed in the Introduction, this scenario is strongly constrained by RGE running.

III. RENORMALIZATION GROUP FLOW OF COMPLEX SOFT TERMS

The goal of this section is to demonstrate how a large phase in A_t can feed into other parameters in the theory through renormalization group running. The imaginary part through renormalization group running. The imaginary part
of A_t at the weak scale, $\overline{A}_t^{\text{EW}}$, is determined by running of A_t at the weak scale, A_t^{cut} , is determined by running $\overline{A}_t^{\text{GUT}}$ (and, for large tan β , $\overline{A}_b^{\text{GUT}}$) down to the weak scale via the renormalization group equations. (For compactness of the renormalization group equations. (For compactness of notation, we will define \bar{x} =Im *x* in the following sections.) notation, we will define $x=Im x$ in the following sections.)
We will show that large $\overline{A}^{\text{EW}}_r$ induces potentially large values *t* We will show that large A_t^{av} induces potentially large values
of B^{EW} and $\overline{A}_{u,d}^{\text{EW}}$, which give an unacceptably large neutron electric dipole moment.

Rather than write RGEs for the whole effective theory, we need only consider a complete subset of them which includes A_q and *B*. The running of these soft terms depends upon the gaugino masses, the top and bottom Yukawas (we ignore tiny effects from the other Yukawa couplings), and the gauge coupling constants $\alpha_a = \lambda_a^2/4\pi$ $(a=1,2,3)$. We define $t = 1/4\pi \ln(Q/M)_{\text{GUT}}$ and write

$$
\frac{dM_a}{dt} = 2b_a \alpha_a M_a \,, \tag{2}
$$

$$
\frac{dA_t}{dt} = 2c_a \alpha_a M_a + 12\alpha_t A_t + 2\alpha_b A_b, \qquad (3)
$$

$$
\frac{dA_{u,c}}{dt} = 2c_a \alpha_a M_a + 6 \alpha_t A_t, \qquad (4)
$$

$$
\frac{dA_b}{dt} = 2c'_a \alpha_a M_a + 2\alpha_t A_t + 12\alpha_b A_b, \qquad (5)
$$

$$
\frac{dA_{d,s}}{dt} = 2c'_a \alpha_a M_a + 6\alpha_b A_b, \qquad (6)
$$

$$
\frac{dB}{dt} = 2c''_a \alpha_a M_a + 6\alpha_t A_t + 6\alpha_b A_b, \qquad (7)
$$

$$
\frac{d\alpha_t}{dt} = 2\alpha_t(-c_a\alpha_a + 6\alpha_t + \alpha_b),\tag{8}
$$

$$
\frac{d\alpha_b}{dt} = 2\alpha_b(-c_a'\alpha_a + \alpha_t + 6\alpha_b),\tag{9}
$$

$$
\frac{d\alpha_a}{dt} = 2b_a \alpha_a^2,\tag{10}
$$

where *a* is summed from 1 to 3, and

$$
b_a = (\frac{33}{5}, 1, -3), \tag{11}
$$

$$
c_a = \left(\frac{13}{15}, 3, \frac{16}{3}\right),\tag{12}
$$

$$
c'_a = (\frac{7}{15}, 3, \frac{16}{3}), \tag{13}
$$

$$
c''_a = (\frac{3}{5}, 3, 0), \tag{14}
$$

and the Yukawa coupling constants $\alpha_{t,b} = \lambda_{t,b}^2 / 4\pi$ are related to the masses by

$$
\lambda_t = \frac{g_2}{\sqrt{2}} \frac{m_t}{m_W} \frac{1}{\sin \beta}, \quad \lambda_b = \frac{g_2}{\sqrt{2}} \frac{m_b}{m_W} \frac{1}{\cos \beta}.
$$
 (15)

We note that some references [23] list the α_iA_i coefficient in Eq. (4) as 2, but we have confidence that the coefficient is actually 6 $[13,24]$. Nevertheless, our conclusions do not depend qualitatively on this coefficient.

In detainary on this coefficient.
We are mainly interested in the evolution of \overline{A}_q and \overline{B} . We can set the phase of the common gaugino mass to zero at We can set the phase of the common gaugino mass to zero at the GUT scale by a phase rotation and then \overline{M}_i =0 at all scales. Therefore the RGE for the imaginary parts of the A_q and *B* can be written without the M_a terms:

$$
\frac{d\overline{A}_t}{dt} = 12\alpha_t \overline{A}_t + 2\alpha_b \overline{A}_b, \qquad (16)
$$

$$
\frac{d\overline{A}_b}{dt} = 2\,\alpha_t \overline{A}_t + 12\,\alpha_b \overline{A}_b,\tag{17}
$$

$$
\frac{d\overline{A}_u}{dt} = 6\,\alpha_t \overline{A}_t,\tag{18}
$$

$$
\frac{d\overline{A}_d}{dt} = 6 \alpha_b \overline{A}_b, \qquad (19)
$$

$$
\frac{d\overline{B}}{dt} = 6\,\alpha_t \overline{A}_t + 6\,\alpha_b \overline{A}_b. \tag{20}
$$

Using the above RGEs, we can derive the useful relations

$$
\Delta \overline{B} = \Delta \overline{A}_{u,c} + \Delta \overline{A}_{d,s} = \frac{6}{14} (\Delta \overline{A}_t + \Delta \overline{A}_b),
$$
 (21)

$$
\Delta \overline{A}_{u,c} = \frac{3}{35} (6\Delta \overline{A}_t - \Delta \overline{A}_b),
$$
 (22)

$$
\Delta \overline{A}_{d,s} = \frac{3}{35} (6 \Delta \overline{A}_b - \Delta \overline{A}_t),
$$
 (23)

where $\Delta \overline{B} = \overline{B}^{\text{GUT}} - \overline{B}^{\text{EW}}$, etc. For small tan β , we can neglect m_b so that these relations simplify to

$$
\Delta \overline{B} = \Delta \overline{A}_{u,c} = 3 \Delta \overline{A}_b = \frac{1}{2} \Delta \overline{A}_t,
$$

$$
\Delta \overline{A}_{d,s} = 0.
$$
 (24)

Thus, given the GUT values, to obtain the low energy values for the imaginary parts of all the soft terms, one only needs for the imaginary parts of all the soft terms, one only needs
to find $\overline{A}^{\text{EW}}_t$ and $\overline{A}^{\text{EW}}_b$, and for small tan β , we only need the former.

In the small tan β limit $(\alpha_b \approx 0)$, we can use Eq. (16) to obtain the ratio of EW- to GUT-scale values of the imaginary part of A_t :

$$
r_t = \overline{A}_t^{\text{EW}} / \overline{A}_t^{\text{GUT}} = \exp\bigg[-\int_{t_{\text{EW}}}^{t_{\text{GUT}}} 12\alpha_t(t)dt\bigg].\tag{25}
$$

If the top quark were light, the integral in Eq. (25) would be small and r_t would be close to 1, but since the top quark is heavy, we find that r_t is well below 1. We can use the relations in Eq. (24) and the definition for r_t in Eq. (25) to relate the low energy values for the imaginary parts of A_t to B and A_u (for small tan β):

$$
\overline{A}_t^{\text{EW}} = \frac{-2r_t}{1 - r_t} \left(\overline{B}^{\text{EW}} - \overline{B}^{\text{GUT}} \right) = \frac{-2r_t}{1 - r_t} \left(\overline{A}_u^{\text{EW}} - \overline{A}_u^{\text{GUT}} \right). \tag{26}
$$

We will make the simplifying assumption that $\overline{A}_{\mu}^{\text{GUT}}$ and We will make the simplifying assumption that A_u^{UU} and B_{GUT} are zero. As we will see in the next section, this is reasonably well justified by our fine-tuning criterion, at least for the phase of *B*.

Next, we must find r_t . We obtain a pseudoanalytic solution to Eq. (25) in terms of EW- and GUT-scale quantities in Eq. $(A1)$ of Appendix A, but this is useful only if one has already obtained the GUT values for the α 's by numerical integration of the RGEs. While we cannot find a truly analytic solution to Eq. (25) , we can place an analytic upper bound on r_t which is sufficient to make our point. We note that the integral in Eq. (25) is simply the area under the curve of the top Yukawa α_t as it runs from the EW scale to the GUT scale. Thus we can place an upper bound on r_t simply by finding a lower bound to that area. In Appendix A, we do this by placing a lower bound on $\alpha_t(t)$ at each *t*, and we obtain

$$
r_t \le 1 - 12\alpha_t^{\text{EW}} / f_{\text{EW}},\tag{27}
$$

which is valid for small $\tan\beta$ as long as $12\alpha_t^{EW} < f_{EW}$. Here $f_{\text{EW}} = 2c_a \alpha_a^{\text{EW}} \approx 1.5 + 32/3(\alpha_s^{\text{EW}} - 0.12)$, and so, for example, Eq. (27) is valid for $m_t=175$ if $1.3<\tan\beta \leq m_t/m_b$ (for smaller tan β , r_t gets closer to zero, but does not actually reach it). Thus we have placed an analytic bound on the running of A_t completely in terms of EW quantities. For $\alpha_s(M_Z)$ =0.12, sin β →1 (moderate tan β), and m_t =175 $(m_t=160)$, we find that $r_t<0.43$ $(r_t<0.52)$, which, from Eq. $(m_t = 160)$, we find that $r_t < 0.43$ ($r_t < 0.52$), which, from Eq. (26), corresponds to $|\overline{A}^{\text{EW}}_t| < 1.5|\overline{B}^{\text{EW}}|$ ($|\overline{A}^{\text{EW}}_t| < 2.2|\overline{B}^{\text{EW}})$). For small tan β , the bound is even stronger, so that for tan β small enough to neglect m_b effects, we obtain

$$
|\overline{A}^{\text{EW}}_t| < 2.2 \text{ min}\{|\overline{B}^{\text{EW}}|, |\overline{A}^{\text{EW}}_u|\},
$$
 (28)

and in practice the coefficient is less than 2.

If in practice the coefficient is less than 2.
In Fig. 1, we plot r_t = $\overline{A}^{\text{EW}}_t$ / $\overline{A}^{\text{GUT}}_t$ as a function of the top Yukawa coupling for different values of $\alpha_s(M_Z)$ in the limit where effects proportional to m_b can be ignored. For m_t 160 GeV, λ_t is always greater than about 0.87 for all values of tan β , which means that r_t is always less than 0.45,

FIG. 1. Plot of the ratios (a) $r_t = \overline{A}^{\text{EW}}_t / \overline{A}^{\text{GUT}}_t$, (b) $(1 - r_t)/2 =$ $-\overline{B}^{\text{EW}}/\overline{A}^{\text{OUT}}_t$, and (c) $(1-r_t)/2r_t = -\overline{B}^{\text{EW}}/\overline{A}^{\text{EW}}_t$ versus the top quark Yukawa coupling for $\alpha_s(M_Z)=0.118\pm0.006$.

in agreement with our analytic bounds. Also plotted are in agreement with our analytic bounds. Also plotted are $-\overline{B}^{\text{EW}}/\overline{A}^{\text{GUT}}_t = (1 - r_t)/2$, and $-\overline{B}^{\text{EW}}/\overline{A}^{\text{EW}}_t = (1 - r_t)/2r_t$, which is greater than 1 (0.6) for m_t =175 (160) . Thus which is greater than 1 (0.6) for m_t =175 (160).
 $|\overline{A}_t^{\text{EW}}| \lesssim |\overline{B}^{\text{EW}}|$, in agreement with our analytic results.

Next we consider moderate tan β , where one must take Next we consider moderate tan β , where one must take into account the mixing of \overline{A}_t and A_b but where tan β is not into account the mixing of A_t and A_b but where $\tan\beta$ is not order m_t/m_b . For $\overline{A_b^{\text{GUT}}}$ / $\overline{A_t^{\text{GUT}}}$ > 0, these effects lower r_t , and one can simply use the $m_b=0$ upper bound on r_t derived $above.¹$

 $\int \text{For } \overline{A}^{\text{GUT}}_b / \overline{A}^{\text{GUT}}_t < 0$ (recall that with universal *A* this ratio would simply be $+1$), one simply maximizes the positive would simply be $+1$), one simply maximizes the po
contribution to r_t from \overline{A}_b to obtain (see Appendix A)

$$
r_t < 1 - 12\alpha_t^{\text{EW}} / f_{\text{EW}} - \frac{1}{6} (\overline{A}_b^{\text{GUT}} / \overline{A}_t^{\text{GUT}}) \frac{\alpha_b^{\text{EW}}}{\alpha_t^{\text{EW}} - \alpha_b^{\text{EW}}}.
$$
 (29)

Note that the last term raises the upper bound on r_t , but the effect is small until tan β gets quite close to m_t/m_b . For effect is small until $\tan\beta$ gets quite close to m_t/m_b . For m_t =175 GeV, $\overline{A}_b^{\text{GUT}} = -\overline{A}_t^{\text{GUT}}$, and $\tan \beta = 0.7 m_t/m_b$ ≈35 (recall that we are evaluating all quantities at the EW scale, and so m_b is somewhat lower than the value at $q^2 = m_b^2$, we find the bound $r_t < 0.6$.

Effects due to m_b are evident in Figs. 2–4, which show Effects due to m_b are evident in Figs. 2–4, which show $\overline{A}^{\text{EW}}_t$, \overline{B}^{EW} , $\overline{A}^{\text{EW}}_u$, and $\overline{A}^{\text{EW}}_d$, normalized to $\overline{A}^{\text{GUT}}_t$, as a function of $tan\beta$ for various GUT-scale boundary conditions. In Fig. 2, only the phase of A_t^{GUT} is nonzero, while in Figs. 3 Fig. 2, only the phase of A_t^{GUT} is nonzero, while in Figs. 3 and 4, $\overline{A}_b^{\text{GUT}}$ has values of $+\overline{A}_t^{\text{GUT}}$ and $-\overline{A}_t^{\text{GUT}}$, respectively. In all cases, r_t (the solid curve) remains below 0.35 and has In all cases, r_t (the solid curve) remains below 0.35 and has
its largest value just below tan $\beta = m_t/m_b$ for $\overline{A}_b^{\text{GUT}}/\overline{A}_t^{\text{GUT}}$ $<$ 0 (Fig. 4), in agreement with our analytic results. This means that the EW value for the phase of A_t is constrained to be less than about a third, independent of constraints from low energy *CP*-violating observables. The magnitude of the low energy *CP*-violating observables. The magnitude of the imaginary part induced into $\overline{B}^{\text{EW}}/\overline{A}_t^{\text{GUT}}$ by \overline{A}_t is greater than

FIG. 2. The ratios of imaginary parts to Im A_t^{GUT} versus tan β with $\text{Im } A_t^{\text{GUT}} \neq 0$ and $\text{Im } A_b^{\text{GUT}} = 0$. The solid line is Im $A_t^{\text{EW}}/ \text{Im } A_t^{\text{GUT}}$, the dashed line is Im $B_{t}^{\text{EW}}/ \text{Im } A_t^{\text{GUT}}$, and the upper (lower) dotted line is $\text{Im } A_d^{\text{EW}}/\text{Im } A_t^{\text{GUT}}$ ($\text{Im } A_u^{\text{EW}}/\text{Im } A_t^{\text{GUT}}$).

0.35 except for large $\tan\beta$ and $\overline{A}_b^{\text{GUT}}/\overline{A}_t^{\text{GUT}}$ < 0. For 0.35 except for large $\tan\beta$ and $A_b^{\text{C}}/A_t^{\text{C}}$ < 0. For $\tan\beta = m_t/m_b$ and $\overline{A_b^{\text{GUT}}}$ / $\overline{A_t^{\text{GUT}}}$ < 0, \overline{B}^{EW} actually goes tan $\beta = m_r/m_b$ and $A_b^{\text{OUT}}/A_t^{\text{OUT}} < 0$, B_{EW} actually goes through zero, because ΔB gets equal and opposite contributhrough zero, because ΔB gets equal and opposite contributions from ΔA_t and ΔA_b there. At that point the *t* and *b* RGE coefficients are almost exactly the same at each t [because the Yukawa coupling runnings differ only in a small $U(1)$ coefficient], and the boundary conditions have opposite coefficient], and the boundary conditions have opposite
signs, so that $\overline{A}_t(t) \approx -\overline{A}_b(t)$ for all *t*. Of course $\overline{A}_a^{\text{EW}}$ and signs, so that $A_t(t) \approx -A_b(t)$ for all t. Of course A_a^{EW} and $\overline{A}_d^{\text{EW}}$ are nonzero because they involve different linear com- A_d^{EW} are nonzero because they involve different linear combinations of $\Delta \overline{A}_t$ and $\Delta \overline{A}_b$, and so there is still a strong constraint on \overline{A}_t from d_n there. constraint on A_t from d_n there.

Finally we note that for large $tan \beta$, one can place con-Finally we note that for large tan β , one can place constraints on $\overline{A}^{\text{GUT}}_b$ as well, since it can then affect other low energy phases through renormalization group running. For $\tan\beta \sim m_t/m_b$, the constraints are of the same order as on \bar{A}^{GUT}_t , while for small tan β , \bar{A}^{GUT}_b is unconstrained (though \bar{A}^{GUT}_t A_t^{tot} , while for small tan β , A_b^{tot} is unconstrained (though the \overline{A}_t contribution to $\overline{A}_b^{\text{EW}}$ for small tan β is constrained to be the A_t contribution to A_b^{av} for small tan β is constraine
small and $\overline{A_b^{\text{EW}}}$ can be large only if $|\overline{A_b^{\text{GUT}}}|\geqslant |\overline{A_t^{\text{GUT}}}|\$.

IV. BOUNDS FROM THE NEUTRON EDM

Now that we have placed an upper bound on the magni-Now that we have placed an upper bound on the magnitude of \overline{A}_t in terms of A_u , \overline{A}_d , and \overline{B} , we need to explore the

FIG. 3. Same as Fig. 2 with $\text{Im } A_b^{\text{GUT}} = \text{Im } A_t^{\text{GUT}}$.

¹There is a subtlety for the case of small positive $\overline{A}_{b}^{\text{GUT}}/\overline{A}_{t}^{\text{GUT}}$ for There is a subtlety for the case of small positive $A_L^{\text{bcl}}/A_L^{\text{dcl}}$ for which there can be a net positive contribution to r_t if A_b runs down below zero. However, the maximum effect on the bound is very small.

FIG. 4. Same as Fig. 2 with $\text{Im } A_b^{\text{GUT}} = -\text{Im } A_t^{\text{GUT}}$.

constraints on the latter three imaginary parts (in low energy observables, we will drop the label EW). As we mentioned in the Introduction, one of the strongest constraints on *CP*violating phases is the electric dipole moment (EDM) of the neutron, d_n . In Appendix B, we write expressions for the full supersymmetric contribution to d_n . One sees that all the supersymmetric contribution to d_n . One sees that all the pieces are proportional to \overline{A}_u , \overline{A}_d , or $\overline{\mu}$ (except for the negligibly small pieces proportional to A_q). We can redefine the *¯* Higgs fields so that the phase of μ is just the opposite of the phase of *B*, and thus

$$
\overline{\mu} = -\left|\frac{\mu}{B}\right| \overline{B} = \left|\frac{\mu}{B}\right| \left(\frac{1 - r_t}{2r_t} \overline{A}_t - \overline{B}^{\text{GUT}}\right),\tag{30}
$$

where the right-hand side (RHS) follows for small tan β .

In order to estimate the size of $\overline{\mu}$ we will need an estimate of $|\mu/B|$ in Eq. (30). We can find this ratio by considering the two equations which μ and *B* need to satisfy to ensure that EW symmetry breaking occurs and that the *Z* boson gets the right mass:

$$
2B\mu = -(m_{H_u}^2 + m_{H_d}^2 + 2\mu^2)\sin 2\beta,
$$
 (31)

$$
\mu^2 = -\frac{m_Z^2}{2} + \frac{m_{H_d}^2 - m_{H_u}^2 \tan^2 \beta}{\tan^2 \beta - 1}.
$$
 (32)

In the limit that tan $\beta \rightarrow \infty$, we see that the right-hand side of Eq. (32) goes to zero, and so $B\rightarrow 0$, whereas μ^2 is not forced to zero. For tan $\beta \rightarrow 1$, the right-hand side of Eq. (32) blows up, forcing μ to take on very large values. When μ^2 dominates Eq. (31) and $tan\beta=1$, then we are led to a value of $|B|=|\mu|$. So in both the tan $\beta\rightarrow\infty$ limit and the tan $\beta\rightarrow1$ limit we find that $|\mu|\geq |B|$. We have run thousands of models numerically $|25|$ which include the one-loop corrections to Eqs. (31) and (32) and found that $|\mu| \geq |B|$ is indeed a good relationship for most of the parameter space. As expected, it is violated most strongly for intermediate values of $tan \beta$. For example, for tan $\beta=10$ we have found a small region of parameter space where $|\mu|/|B|$ is as low as 0.4, although most solutions prefer $|\mu|/|B| > 1$. We will assume that $|\mu|/|B| \geq 1$, and thus the fine-tuning constraint on the phase of *B* is even stronger than on what we obtain below for the phase of μ .

From Appendix B, we see that d_n can be written in terms of the three imaginary parts:

$$
\frac{d_n}{10^{-25} e \text{ cm}} = k_n^{A_u} \frac{\overline{A}_u}{m_0} + k_n^{A_d} \frac{\overline{A}_d}{m_0} + k_n^{\mu} \frac{\overline{\mu}}{m_0}
$$

$$
= k_n^{A_u} \frac{\overline{A}_u}{m_0} + k_n^{A_d} \frac{\overline{A}_d}{m_0} - k_n^{\mu} \frac{\mu}{B} \frac{\overline{B}}{m_0}, \quad (33)
$$

where we have normalized the RHS by the SUSY mass scale m_0 and the LHS by the region of the experimental bound so that the coefficients *k* are dimensionless. We can rewrite the EW imaginary parts in Eq. (33) using Eq. (24) as

$$
\frac{d_n}{10^{-25} e \text{ cm}} = \frac{d_n^{\text{GUT}}}{10^{-25} e \text{ cm}} + \frac{1 - r_t}{2r_t} \left(-k_n^{A_u} + k_n^{\mu} \left| \frac{\mu}{B} \right| \right) \frac{\overline{A}_t}{m_0},\tag{34}
$$

where $d_n^{\text{GUT}}/10^{-25}$ *e* cm is just Eq. (33) with EW values of where $d_n^{\text{OUT}}/10^{-25}$ *e* cm is just Eq. (33) with EW values of $\overline{A}_{u,d}$ and \overline{B} replaced by GUT quantities. It vanishes if $A_{u,d}$ and *B* replaced by GUT quantities. It vanishes if $\overline{A}_{u,d}^{\text{GUT}}$ and $\overline{B}^{\text{GUT}}$ are zero. In supergravity models, $|A^{\text{GUT}}|$ and $|B^{\text{GUT}}|$ are of order m_0 , so that barring fine-tuned cancellations, the GUT-scale phases must be less than order $1/k_n$. If the *k*'s are greater than order 10, then our fine-tuning criterion dictates that we set the GUT phases to zero (presumably protected by some symmetry). Thus we need an estimate of the k_n 's.

In Figs. 5(a), 5(b), and 5(c), we plot the values for $k_n^{A_u}$, $k_n^{A_d}$, and k_n^{μ} , respectively, in many different models as a function of squark mass and as a function of $tan\beta$ in Figs. 5(d), 5(e), and 5(f). We see that $k_n^{A_u}$ and $k_n^{A_d}$ are fairly flat functions of $\tan\beta$, whereas $-k_n^{\mu}$ increases with $\tan\beta$ due to the μ tan β terms in the expression for d_n . We also see that most models give $k_n^{A_u} > 2(0.8)$, $k_n^{A_u} > 7(3)$, and $|k_n^{\mu}| > 100(40)$, for squark masses below 500 GeV (1 TeV), so that order-1 phases in all the SUSY complex quantities usually give a neutron EDM which is of order $100 (40)$ times the experimental bound. We note that these are substantially larger contributions (and thus stronger constraints) than claimed by the recent work of Falk and Olive $[26]$, though this is probably due to the fact that they use very heavy squark masses in an effort to find the smallest fine-tuning of phases consistent with cosmology. While one can argue whether or not the bounds on the phases of $A_{u,d}$ represent a whether or not the bounds on the phases of $A_{u,d}$ represent a fine-tuning, the bound on the phase of μ (and thus \overline{B}^{EW}), fine-tuning, the bound on the phase of μ (and thus B^{av} , which comes from $\overline{B}^{\text{GUT}}$ and $\overline{A}^{\text{GUT}}_t$) certainly does. Thus, by our fine-tuning criterion, the phases of B^{GUT} and A_t^{GUT} should be zero. We note that in the case of universal *A* it is irrelevant whether or not the low energy phases of A_u and A_d are strongly constrained, since the phase of the universal *A*GUT makes a large contribution to the low energy value of A^{GUT} makes a large con
 $\overline{\mu}$ (since $\overline{A}^{\text{GUT}}_t = \overline{A}^{\text{GUT}}$).

To give an idea of what level of neutron EDM one expects with different initial assumptions, we plot, in Fig. 6, $d_n/10^{-25}$ *e* cm with universal $|A^{GUT}|$ for three cases: (a),(d) $Arg A_{t}^{GUT} = Arg A_b^{GUT} = 0.1$ and all other phases zero, (b),(e) $ArgA_t^{GUT} = -ArgA_b^{GUT} = 0.1$ and all other phases zero, and (c) , (f) universal phases Arg A^{GUT} =Arg B^{GUT} =0.1. As one can see, even with phases of order 0.1, most models have an absolute value for $d_n/10^{-25}$ *e* cm greater than 1, inconsistent with the experimental bounds.

FIG. 5. Scatter plots of (a) and (d) $k_n^{A_u}$, (b) and (e) $k_n^{A_d}$, and (c) and (f) k_n^{μ} versus squark mass and versus tan β . Each point represents a solution of the supersymmetric parameter space with universal scalar and gaugino mass terms which is within other experimental limits.

As can be gathered by the spread of points in the scatter plots and the number of parameters involved, the results depend on one's model assumptions. For example, if one requires tan β to be small (say, because of *b*- τ unification), and the squarks are allowed to be very heavy, then there is very little fine-tuning needed for the current experimental bound on d_n . On the other hand, if SUSY is detected at the CERN e^+e^- collider LEP 2 or Fermilab TeV 33, then even the smallest tan β models would require fine-tuning.

In minimal supergravity models the natural scale for the *A* terms is m_0 . In Fig. 7 we have plotted $d_n/10^{-25}$ *e* cm versus Im A_t^{EW}/m_0 to succinctly demonstrate how quickly the EDM rises when $\text{Im} A_t^{\text{EW}} \neq 0$. To construct this plot we chose a random phase for A_t at the GUT scale, forced all other phases equal to zero at that scale, and then ran all the parameters down to the weak scale. A sharp drop in d_n occurs at Im $A_t^{\text{EW}}/m_0 \approx 0$ because Im A_t^{GUT} can be small there and thus induces only small phases into the other low energy soft parameters. Models with d_n around 10^{-25} *e* cm at $\lim_{t \to 0} A_t^{\text{EW}}/m_0 \approx 0$ occur for low $\tan \beta$ where $\lim_{t \to 0} A_t^{\text{GUT}}$ \gg Im A_t^{EW} but where d_n otherwise tends to be smaller. This means that most models with d_n below the experimental bound in Fig. 6 also have a small EW value for $\text{Im} A_t$, and thus from Fig. 7 we can place a stronger constraint on Im A_t^{EW} than we obtained on Im A_t^{GUT} : Im $A_t^{\text{EW}}/m_0 \le 1/20$.

Thus we conclude that models with universal GUT-scale phases of the soft parameters and models in which only A_t^{GUT} has a nonzero phase have difficulty meeting the bounds from d_n and our fine-tuning criterion. Models with nonzero from d_n and our fine-tuning criterion. Models with nonzero $\overline{A}_c^{\text{GUT}}$, $\overline{A}_s^{\text{GUT}}$, or $\overline{A}_l^{\text{GUT}}$ can meet the constraint from d_n with- A_c^{out} , A_s^{out} , or A_l^{out} can meet the constraint from d_n with-
out fine-tuning, as can those with nonzero $\overline{A}_b^{\text{GUT}}$ for small tan β . For the remainder of the paper, we will for simplicity set all the GUT-scale phases to zero except for that of A_t . set all the GUT-scale phases to zero except for that of A_t .
Even though our fine-tuning criterion implies that $\overline{A}^{\text{GUT}}_t$ should be zero, we find it useful to ask what effects one would have if one allows that fine-tuning.

V. TOP QUARK EDM

Now that the top quark has finally been discovered, one can envision some nice experiments which measure properties of this known particle. Future colliders, such as the NLC, can provide many precision measurements of the production cross-section and decay properties of the top quark. It is possible that signatures of new physics could arise out of such a study. One property of the top quark which has received much attention $[19]$ is the possibility of measuring its EDM by looking at the decay distributions of the $t\bar{t}$ pairs. (Other *CP*-violating observables are possible, such as those arising from $t \rightarrow bW$ decays, but we will make our point only with the top quark EDM.) It is generally estimated that the top quark EDM (d_t) can be measured to values as low as $\sim 10^{-18}$ *e* cm [19]. Given the constraints which we derived above, we ask if the minimal supersymmetric standard model can yield a value for d_t , this large.

In the context of supersymmetry, it has been proposed [20] that a large d_t is possible if the phase of A_t^{EW} is of order [20] that a large d_t is possible if the phase of A_t^{EW} is of order 1. But in Sec. III, we showed that $\overline{A}_t^{\text{EW}}$ is constrained to be smaller than or of order of the phases which contribute to d_n . The EDM of the top quark is thus constrained to be less than a constant times the neutron EDM:

$$
\frac{d_t}{d_n} \lesssim \xi \frac{m_t}{m_d} \frac{\det M_{\tilde{q}}^2}{\det M_{\tilde{t}}^2},
$$
\n(35)

FIG. 6. Scatter plots of $d_n/10^{-25}$ *e* cm versus squark mass and versus tan β for (a) and (d) Arg $A_t^{\text{GUT}} = \text{Arg } A_b^{\text{GUT}} = 0.1$ with all other phases zero, (b) and (e) $Arg A_t^{GUT} = -Arg A_b^{GUT} = 0.1$ with all other phases zero, and or (c) and (f) universal phases ArgA^{GUT}=Arg*B*^{GUT}=0.1. Each point represents a solution of the supersymmetric parameter space with universal scalar and gaugino mass terms which is within other experimental limits.

where det $M_{\tilde{q}}^2 = m_{\tilde{q}_1}^2 m_{\tilde{q}_2}^2$ is the determinant of the (down) squark mass-squared matrix, and the value of ξ depends upon many different SUSY parameters, but is generically of order 1. Normalizing d_n to the experimental bound, we see that

$$
d_t \lesssim \xi \frac{\det M_{\tilde{q}}^2}{\det M_{\tilde{t}}^2} \frac{d_n^{\text{expt}}}{10^{-25} e \text{ cm}} 2 \times 10^{-21} e \text{ cm.}
$$
 (36)

In addition to this constraint, we recall that the phase of A_t at the EW scale must be less than about 1/3, just from the RGE suppression factor r_t . Thus, as long as det $M_d \approx \det M_t$, we expect d_t to fall about three orders of magnitude below detectability at proposed future high energy colliders.

We can turn this analysis around. If a large top quark EDM is discovered, can it be explained in the MSSM? One possibility is that a conspiracy occurs between several large phases in the theory to render d_n below experimental limits, and yet produce a d_t detectable at high energy colliders. This is equivalent to saying that all the order 1 coefficients which we absorbed into the parameter ξ in Eq. (36) actually conspire to give $\xi \ge 10^3$. As we argued in the Introduction, we would not view this as a likely explanation.

Another possibility to consider is that the top squarks are much lighter than the other squarks. For d_t to be observable, we would need the determinants in Eq. (35) to have a ratio \approx 10³. This is possible, but it too would require some finetuning. The large top-quark-induced running of the \tilde{t}_R goes in the right direction—the lightest top squark mass eigenvalue tends to be smaller than the other quarks. However, \tilde{t}_2 value tends to be smaller than the other quarks. However, t_2 generally tracks fairly well with the other squarks, \tilde{q}_L , and thus, we estimate that

$$
\frac{\det M_d^2}{\det M_{\tilde{t}}^2} \lesssim \frac{m_{\tilde{d}}^2}{m_{\tilde{t}_1}^2},
$$
\n(37)

which means that we would need $m_{\tilde{t}_1} \leq m_{\tilde{d}} / \sqrt{1000}$ to yield an observable d_t . If experiment determines that $m_t \tilde{t}_1 > 80$ GeV, then this condition would imply that the superpartners

FIG. 7. Scatter plot of $d_n/10^{-25}$ *e* cm versus Im A_t^{EW}/m_0 . Each point represents a solution of the supersymmetric parameter space with universal scalar and gaugino mass terms which is within other experimental limits.

implied by the small ratio $m \tilde{t}_1/m \tilde{d}$. Finally, one could appeal to differences between d_t and d_d due to effects proportional to m_d^2/v^2 , which are negligible in d_d . To achieve $\dot{\xi}$ of order 10³, one again needs a finetuned conspiracy of couplings.

mentioned in the Introduction, with an additional fine-tuning

Thus we conclude that if a large d_t , were found, one would probably have to look beyond the MSSM for an explanation.

VI. CONCLUDING REMARKS

It has long been noted that the phases of soft supersymmetric parameters generically lead to an unacceptably large neutron EDM. This fine-tuning problem has slowly become less vexing as the theoretical expectations for the squark masses have risen faster than the experimental bound on the neutron EDM has fallen. Nevertheless, for squark masses below about a TeV, we showed in Sec. V that the phase of *B* and universal phase of *A* do not meet the fine-tuning criterion set forth in the Introduction [see Fig. $6(c)$]. Certainly, if supersymmetry is discovered at LEP 2 or TeV 33, a fundamental explanation for the absence of a neutron EDM would be needed, and any scheme for baryogenesis at the EW scale would require that mechanism to leave small effective low energy phases in the soft terms $[15–18]$.

From the phenomenological point of view, it is tempting to postulate that the soft phases are not universal, that the EW phase of A_t is large, while the other phases which directly contribute to the neutron EDM are small. This would allow interesting signatures of supersymmetric *CP* violation to be visible in top quark physics at future colliders. But we have demonstrated by using the renormalization group equations that the imaginary part of A_t must be less than twice the imaginary part of *B*, and A_t -induced *CP*-violating observables such as the top quark EDM are thus expected to be unobservably small in almost all minimal SUSY models.

These constraints are particularly important for models of EW baryogenesis which rely upon the phase of the top squark *LR* mixing parameter, $A_t + \mu^* \tan\beta$, to generate enough *CP* violation for baryogenesis. Such models must also have sufficiently small $A_t + \mu^*$ tan β to ensure that the phase transition is first order $[27]$. There has also been a recent attempt to explain the observed *CP* violation in the neutral kaon system with zero CKM phase and nonzero offdiagonal phases in the general *A* matrices [28]. If the universal diagonal *A* parameter has a large phase at the GUT scale, it will, as we noted above, give a large contribution to d_n through a renormalization-group-induced phase in μ , as well as from a direct contribution. One could evade such bounds by insisting that the off-diagonal components of the *A* matrices have a large phase, while the phases of the diagonal *A*'s and of *B* vanish. Although this hypothesis can probably be technically consistent with our fine-tuning criterion (phases either zero or large), this scenario strikes us as unnatural.

ACKNOWLEDGMENTS

We would like to thank G. Kane, S. P. Martin, D. Wyler, S. Thomas, A. Riotto, and A. Soni for helpful discussions. R. G. greatly appreciates the hospitality of the Brookhaven National Laboratory HEP Theory Group. The work of R. G. was supported in part by the Department of Energy, Contract No. DE-AC02-76CH00016. The work of J. W. was supported by the Department of Energy, Contract No. DE-AC03-76SF00515.

APPENDIX A

In this appendix, we provide the details related to our analytic results of Sec. III. It is interesting to note that we can use the RGEs for the top Yukawa and gauge coupling constants in Eqs. (8) and (10) to write a pseudoanalytical solution to $r_t = \text{Im} A_t^{\text{EW}}/\text{Im} A_t^{\text{GUT}}$. The integral in Eq. (25) can be rewritten as $\ln(\alpha_t^{\text{EW}}/\alpha_t^{\text{GUT}}) - \sum_a (c_a/b_a) \ln(\alpha_a^{\text{GUT}}/\alpha_a^{\text{EW}})$, which allows us to write a pseudoanalytic r_t in terms of EW and GUT-scale quantities (the latter of which cannot be found analytically):

$$
r_t = \frac{\alpha_t^{\text{EW}}}{\alpha_t^{\text{GUT}}} \prod_{a=1}^3 \left(\frac{\alpha_a^{\text{EW}}}{\alpha_a^{\text{GUT}}} \right)^{c_a/b_a}.
$$
 (A1)

To place an analytic upper bound on r_t , we must place a lower bound on the area $\int_{t_{\text{EW}}}^{t_{\text{GUT}}} 12\alpha_t(t)dt$. We will need the α_b =0 limit of the running of the top Yukawa coupling in Eq. $(8).$

$$
\frac{d\alpha_t}{dt} = -f(t)\alpha_t + 12\alpha_t^2,
$$
 (A2)

where $f(t)=2c_a\alpha_a$. While this cannot be solved analytically, we note that $\alpha_3(t)$ runs down with energy and one can show that $f(t)$ will be at its maximum value at the EW scale. Thus if we take $f(t)$ to the constant f_{EW} , we will minimize the running of α_t , and Eq. (A2) can be solved analytically to yield the bound

$$
\alpha_t(t) > \left[\frac{12}{f_{\rm EW}} + \left(\frac{1}{\alpha_t^{\rm EW}} - \frac{12}{f_{\rm EW}} \right) e^{f_{\rm EW}(t - t_{\rm EW})} \right]^{-1}, \quad \text{(A3)}
$$

which is valid for $12\alpha_t^{\text{EW}} < f_{\text{EW}}$ [larger α_t^{EW} allows the bound on $\alpha_t(t)$ to reach infinity for $t < t_{\text{GUT}}$ and thus makes the bound useless], which corresponds to tan β >1.3 for $m_t=175$. If we replace $\alpha_t(t)$ in the integral above by the RHS of Eq. $(A3)$, we can find an analytic solution for the lower bound on the area which, for the relevant range of f_{EW} and t_{EW} , can be approximated by

$$
-f_{\text{EW}}t_{\text{EW}} - \ln\left[\frac{12\alpha_t^{\text{EW}}}{f_{\text{EW}}} + \left(1 - \frac{12\alpha_t^{\text{EW}}}{f_{\text{EW}}}\right)e^{-f_{\text{EW}}t_{\text{EW}}}\right]
$$

$$
\approx -\ln\left[1 - \frac{12\alpha_t^{\text{EW}}}{f_{\text{EW}}}\right],\tag{A4}
$$

which yields Eq. (27) directly.

For moderate tan β , we need to include m_b effects which \overline{A}_b , and α , with α_b . The coupled differential mix A_t with A_b , and α_t with α_b . The coupled differential equations (16) and (17) can be solved analytically only if the coefficients, which here are proportional to α_t and α_b , are coefficients, which here are proportional to α_t and α_b , are constants. To obtain bounds on the running of \overline{A}_t and \overline{A}_b , we can break up the range of energy from t_{EW} to t_{GUT} into small

regions where the coefficients are effectively constant, and iteratively evolve from the GUT scale down to the weak scale. At each energy t_i , the value for \overline{A}_i , is given by scale. At each energy t_j , the value for A_t is given by

$$
\overline{A}_t(t_{j+1}) \approx \overline{A}_t(t_j) \exp[-12\alpha_t(t_j)\delta t] - \frac{1}{6} \overline{A}_b(t_j)
$$

$$
\times \left\{ \frac{\alpha_b(t_j)}{\alpha_t(t_j) - \alpha_b(t_j)} \left\{ \exp[-12\alpha_b(t_j)\delta t] \right\} - \exp[-12\alpha_t(t_j)\delta t] \right\},
$$
(A5)

provided that $T \equiv \alpha_b/(\alpha_t - \alpha_b)$ is not large. Here $\delta t = t_i - t_{i+1}$, which is positive. Iterating Eq. (A5) gives a complicated expression with terms proportional to each of the $T(t_i)$'s. However, each of these terms is positive, so that taking $T(t_i)$ to its maximum value maximizes the size of the quantity in $\{ \}$'s in Eq. (A5), which is what we need for the quantity in $\frac{1}{2}$ is in Eq. (A5), which is what we need for the case $\overline{A}_{b}^{\text{GUT}}/\overline{A}_{t}^{\text{GUT}}$ < 0. Once we take $T(t_j) \rightarrow T_{\text{max}}$, many terms cancel, and we are left with (taking $\delta t \rightarrow 0$) the upper limit

$$
\bar{A}_t^{\text{EW}} < \bar{A}_t^{\text{GUT}} \exp\left(-\int_{t_{\text{EW}}}^{t^{\text{GUT}}} 12\alpha_t(t)dt\right)
$$

$$
-\frac{1}{6} \bar{A}_b^{\text{GUT}} \left(\frac{\alpha_b}{\alpha_t - \alpha_b}\right)_{\text{max}} \left[\exp\left(-\int_{t_{\text{EW}}}^{t^{\text{GUT}}} 12\alpha_b(t)dt\right)\right]
$$

$$
-\exp\left(-\int_{t_{\text{EW}}}^{t^{\text{GUT}}} 12\alpha_t(t)dt\right)\right].
$$
(A6)

One can show analytically that $T(t)$ reaches its maximum value at the lowest energy of the range, and thus we can replace T_{max} by $\alpha_b^{\text{EW}}/(\alpha_t^{\text{EW}} - \alpha_b^{\text{EW}})$. To obtain a simpler bound, one can reduce the $[$ $]$'s in Eq. (A6) to 1 by taking a lower bound on $\alpha_h(t)$ to be zero and an upper bound on $\alpha_t(t)$ to be infinity. Finally one uses the $m_b \approx 0$ bound on r_t obtained in Eq. (27) for the first term in Eq. $(A6)$ to obtain the upper bound on r_t in Eq. (29).

APPENDIX B

In this appendix we present analytic expressions for the full one-loop SUSY contribution to the neutron electric dipole moment d_n . The gluino [11] and chargino [7] contributions appear in the literature. While an expression for the neutralino contribution is given by Kizukuri and Oshimo $|7|$, it is written in terms of 4×4 complex unitary matrices which must be determined numerically. Below we give an expression for this neutralino contribution solely in terms of the mass matrices [and other minimal supersymmetric standard model (MSSM) parameters] and a useful approximation to that expression, which do not require calculating complex unitary matrices.

To find the neutron EDM, we first calculate the EDM of the up and down quarks (d_a) from one-loop diagrams with photons attached to either (*a*) an internal boson or (*b*) an internal fermion line. Then the neutron EDM is related to the quark EDM's in the naive quark model by $d_n = (4d_d - d_u)/3$, though recent work has argued that this expression overestimates d_n if the strange quark carries a large fraction of the neutron and proton spin [29]. The Feynman integrals associ-ated with (a) and (b) are $\lceil 30 \rceil$

$$
I^{a}(x) = \frac{1}{(1-x)^{2}} \left[-\frac{3}{2} + \frac{x}{2} - \frac{\ln x}{1-x} \right],
$$

$$
I^{b}(x) = \frac{1}{(1-x)^{2}} \left[\frac{1}{2} + \frac{x}{2} + \frac{x \ln x}{1-x} \right].
$$
 (B1)

As we mentioned earlier, all SUSY *CP*-violating effects arise from diagonalizing complex mass matrices $[15]$. Gluino loops contribute to the quark EDM d_a through the complex phase in the left-right mixing elements for up and down squarks:

$$
d_q(\tilde{g}) = \frac{-2}{3\pi} Q_q e \alpha_s \frac{m_q m_{\tilde{g}} \operatorname{Im}(A_q - \mu R_q)}{m_{\tilde{q}_0}^4} I^b \left(\frac{m_{\tilde{g}}^2}{m_{\tilde{q}_0}^2} \right).
$$
\n(B2)

We have averaged over the nearly degenerate squark mass eigenstates for simplicity: $m_{\tilde{q}_0}^2 = m_{\tilde{q}_1} m_{\tilde{q}_2}$ and $I(x_0)$ $\frac{1}{2}[I(x_1) + I(x_2)]/2$. Here Q_qe is the charge of quark *q*, $R_q = \tan\beta(\cot\beta)$ for $q = d(u)$, and $m_{\tilde{q}}$ is the gluino mass. (Note that we use Im z for the imaginary part of z in this (Note that we use Im *z* for the imaginary part of *z* in this appendix because it is clearer than \overline{z} in more complicated expressions.)

The chargino contribution is proportional to the imaginary part of products of elements of the matrices *U* and *V* which diagonalize the chargino mass matrix. It turns out that one can write those products directly in terms of the elements of the chargino mass matrix, so that the chargino contribution to d_q can be written

$$
d_q(\tilde{\chi}^+) = \frac{e}{(4\pi)^2} \left\{ g^2 R_q \frac{m_q m_{\tilde{W}} \text{Im } \mu}{m_{\tilde{q}_0'}^4} \frac{\left[\omega I^a(y_1) + Q_{q'} I^b(y_1)\right] - \left[\omega I^a(y_2) + Q_{q'} I^b(y_2)\right]}{y_1 - y_2} \right. \\ \left. + \frac{m_q \text{Re } \mu}{\sin \beta \cos \beta} \left[\omega I^a(y_0) + Q_{q'} I^b(y_0)\right] \sum_{r = q'_i}^{dsb \text{ or } uct} \frac{\text{Im}(A_r - \mu R_{q'})}{m_{\tilde{r}_0}^4} \frac{m_r^2 |V_{qr}|^2}{v^2} \right\}, \tag{B3}
$$

where $m\tilde{w}$ is the *W*-ino mass, $\omega = +1(-1)$ for $q = d(u)$, y_1 $=m_{\tilde{\chi}_1}^2/m_{\tilde{q}_0'}^2$, and $I(y_0)=[I(y_1)+I(y_2)]/2$. The primed quantities refer to the SU(2) partner, and so if $q=d$, then $q' = u$ and *r* is summed over the set $\{u, c, t\}$. Previous expressions and r is summed over the set $\{u, c, r\}$. Previous expressions
for $d_q(\tilde{\chi}^+)$ have neglected the squark mixing piece, which is the second term in Eq. $(B3)$. This piece is suppressed relative to the other contributions by $m_r^2 |\bar{V}_{qr}|^2/v^2$, which is less than 10^{-4} for $q=u$ or *d* (but it can affect the EDMs of other quarks), though it is interesting that there is a (tiny) *direct* contribution to d_n from Im A_t .

The neutralino contribution

$$
d_q(\tilde{\chi}^0) = \frac{-Q_q e}{(4\pi)^2} \frac{1}{m_{\tilde{q}_0}^2} \frac{m_q}{v_q} \left\{ \sum_{h=1}^{1,2} (a_{Lh}^q - a_{Rh}^q) \text{Im } \Phi_{h\hat{q}} \right.+ \frac{\text{Re}(A_q + \mu R_q) v_q}{m_{\tilde{q}_0}^2} \sum_{h,l}^{1,2} a_{Lh}^q a_{Rl}^q \text{Im } \Phi_{hl} - \frac{\text{Im}(A_q - \mu R_q) v_q}{m_{\tilde{q}_0}^2} \sum_{h,l}^{1,2} a_{Lh}^q a_{Rl}^q \text{Re } \Phi_{hl} \right\}
$$
(B4)

arises from the 4×4 complex neutralino mass matrix. The index $\hat{q} = 3(4)$ for $q = d(u)$. Recall [31] that the ''1'' and ''2'' weak eigenstates are gauginos, and the ''3'' and ''4'' weak eigenstates are Higgsinos which couple to down and up quarks respectively. Thus ''34'' and ''43'' terms are absent, which will allow us to simplify expressions involving the neutralino mass matrix, since that is the position of the complex coefficient μ . We have dropped terms of order m_f^2/v^2 relative to the others. The gauge coefficients a_{Li} are

$$
a_{L1}^q = \sqrt{2}g \tan \theta_w (Q_q - T_{3L}^q) = \sqrt{2}g \tan \theta_w / 6, \quad \text{(B5)}
$$

$$
a_{L2}^q = \sqrt{2}gT_{3L}^q, \qquad (B6)
$$

and the a_{Ri} are the same as the a_{Li} with $T_{3L} \rightarrow T_{3R} = 0$. The neutralino phases appear through a 4×4 matrix

$$
\Phi_{hl} = \sum_{i=1}^{4} U_{hi}^{T} \hat{M}_{ii} I_{ii}^{b} U_{il},
$$
\n(B7)

where *U* diagonalizes the neutralino mass matrix *M*, and *Mˆ* is the diagonal result. Here $I_{ij}^b = I^b(x_i) \delta_{ij}$ i.e., I_{ij}^b is the diagonal matrix of Feynman integrals for the corresponding mass eigenvalues in \hat{M} . In the limit that the $I^b(x_i)$ are equal, the real part of Φ_{hl} can simply be written

$$
\text{Re }\Phi_{hl} \simeq I^b(x_0) \text{Re } M_{hl},\tag{B8}
$$

where $I^b(x_0) = \sum_{i=1}^{4} I^b(x_i) / 4$. The imaginary part of Φ_{hl} is more difficult because it vanishes in the limit of degenerate neutralino masses (except for the irrelevant "34" and "43" terms). We know that Im μ is the only complex coefficient in the neutralino mass matrix M , and so we can write

$$
\text{Im } \Phi_{hl} \simeq \Omega_{hl} \text{ Im } \mu,\tag{B9}
$$

where Ω_{hl} is a real matrix to be determined. This allows us where Ω_{hl} is a real matrix to be determined. This allows us
to see that $d_g(\bar{\chi}^0)$ is proportional to Im μ and Im($A_q - \mu R_q$), just as the other contributions are. It turns out that Im Φ_{hl} is proportional to $\text{Im}(MM^*M)_{hl}$, $\text{Im}(MM^*MM^*M)_{hl}$, and $\text{Im}(MM^*MM^*MM^*M)_{hl}$ (except for the ''34'' and ''43'' pieces). To extract the Im μ dependence, we ignore all terms of higher order in Im $\mu/|\mu|$, which is a valid approximation for the phases allowed by the experimental bound on d_n . Then these products $[for (h,l) \neq (3,4), (4,3)]$ simplify as follows:

Im
$$
(MM^*M)_{hl}
$$
 \simeq Im $\mu \sum_{p=0}^{2} (-1)^p (M_R^p P M_R^{2-p})_{hl}$, (B10)

Im
$$
(MM^*MM^*M)_{hl}
$$
 = Im $\mu \sum_{p=0}^{4} (-1)^p (M_R^p M_R^{4-p})_{hl}$, (B11)

 $Im(MM*MM*MM*M)_{hl}$

$$
\simeq \text{Im } \mu \sum_{p=0}^{6} (-1)^{p} (M_R^p P M_R^{6-p})_{hl},
$$
\n(B12)

where M_R =Re *M* and *P* is a matrix with -1 in the 34 and 43 positions and 0 everywhere else [so that $\text{Im}(\mu P)$ \equiv Im *M* . After some calculation, we obtain an expression for the imaginary part of the complex matrix Φ :

Im
$$
\Phi_{hl} \simeq \Omega_{hl}
$$
 Im μ
\n
$$
\simeq \frac{3}{2} Im \mu \sum_{s=1}^{4} \left(I(x_s) - \sum_{j}^{4} I(x_j) \right) \left[\sum_{i,j,k}^{4} \epsilon_{ijk} \hat{M}_i^4 \hat{M}_j^6 \sum_{p=0}^{2} (-1)^p (M_R^p M_R^{2-p})_{hl} - \sum_{i,j,k}^{4} \epsilon_{ijk} \hat{M}_i^2 \hat{M}_j^6 \right] \times \sum_{p=0}^{4} (-1)^p (M_R^p M_R^{4-p})_{hl} + \sum_{i,j,k}^{4} \epsilon_{ijk} \hat{M}_i^2 \hat{M}_j^4 \sum_{p=0}^{6} (-1)^p (M_R^p M_R^{6-p})_{hl} \left[\sum_{i,j,k}^{4} \epsilon_{ijk} \hat{M}_i^4 \hat{M}_j^6 \right] 3 \hat{M}_s^2 - \sum_{n}^{4} \hat{M}_n^2 \right]
$$
\n
$$
- \sum_{i,j,k}^{4} \epsilon_{ijk} \hat{M}_i^2 \hat{M}_j^6 \left(3 \hat{M}_s^4 - \sum_{n}^{4} \hat{M}_n^4 \right) + \sum_{i,j,k}^{4} \epsilon_{ijk} \hat{M}_i^2 \hat{M}_j^4 \left(3 \hat{M}_s^6 - \sum_{n}^{4} \hat{M}_n^6 \right) \right]^{-1}, \tag{B13}
$$

where \sum_{j}^{f} means sum over the three members of the set $\{1,2,3,4\}$ ⁻ $\{s\}$. Here *M*_{*i*} are the mass eigenvalues of *M* (i.e., the four physical neutralino masses).

The expression above is completely analytic and exact except for the approximation we made in dropping higher order terms in Im $\mu/|\mu|$, but it has so many terms that it is not that useful. Let us find an approximation to this expression using the information about the neutralino mass eigenstates, namely, that they are fairly close together and the heaviest neutralino is lighter than the squarks $(x_4 \ll 1)$ in most SUSY models. This means that we can take a simple linear fit to the Feynman integral by evaluating $I^b(x)$ at the lowest and highest values of *x*:

$$
I^{b}(x) \approx K_0 + K_1 x \approx I^{b}(x_1) + S_{41}(x - x_1), \quad (B14)
$$

where S_{41} is the slope

$$
S_{41} = \frac{I^b(x_4) - I^b(x_1)}{x_4 - x_1},
$$
\n(B15)

and $x_j = m_{\tilde{x}_j^0} / m_{\tilde{q}_0^0}$. Thus $K_1 = S_{41}$ and $K_0 = I^b(x_1) - S_{41}x_1$. Note that this approximation gives exact values for x_1 and x_4 , and is only off for x_2 and x_3 —a rough estimate is that the approximation is correct to about 5%.

If we plug Eq. $(B14)$ into Im Φ_{hl} in Eq. $(B7)$, we see that the K_0 piece vanishes (except for the ''34'' and ''43'' pieces), and we obtain

Im
$$
\Phi_{hl} \approx S_{41} \frac{\text{Im}(MM^*M)_{hl}}{m_{\tilde{q}_0}^2} \approx \frac{S_{41}}{m_{\tilde{q}_0}^2}
$$

 $\times \text{Im } \mu \sum_{p=0}^{2} (-1)^p (M_R^p M_R^{2-p})_{hl}$. (B16)

The neutralino contribution to d_q is found by plugging Eq. (B8) for Re Φ_{hl} and Eq. (B13) or Eq. (B16) for Im Φ_{hl} into $(B4).$

Finally, we want to relate the expressions for the three SUSY contributions to the quark EDM in Eqs. $(B2)$, $(B3)$, and $(B4)$ in terms of the coefficients k_n from Sec. IV. Using the naive quark model, $d_n = 4/3d_d - 1/3d_u$, we can write $k_{n}^{x} = 4/3k_{d}^{x} - 1/3k_{u}^{x}$ and

$$
k_q^x = \frac{d_q^x}{10^{-25} \ e \text{ cm}} \frac{m_0}{\text{Im } x},
$$
 (B17)

where $x = A_u$, A_d , or μ , and d_q^x is the contribution to d_q from complex quantity *x*. We can see that if we neglect the tiny second term of $d_q(\tilde{\chi}^+)$ in Eq. (B3), then $k_a^{\hat{A}}(k_a^{\hat{A}d})$ gets contributions only from the gluino and neutralino contributions to $d_u(d_d)$, whereas k_n^{μ} gets contributions from all three of the SUSY contributions to d_u and d_d .

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