# Rational unification helps the seesaw mechanism

Karim Benakli<sup>\*</sup> and Goran Senjanović<sup>†</sup>

International Centre for Theoretical Physics, 34014 Trieste, Italy

(Received 2 May 1996)

In this paper we explore the possibility of intermediate scale physics in the context of superstring models with *higher Kac-Moody levels*, by focusing on left-right and Pati-Salam symmetries. We find that the left-right scale may lie in the range  $10^{10}-10^{12}$  GeV, which is favored by neutrino physics, while the Pati-Salam scale is at most two or three orders of magnitude below the unification scale  $M_X$ . We also show that the scale of B-L breaking can be as low as 1 TeV or so, providing protection against too rapid proton decay in supersymmetry. Our results allow a natural value for the scale  $M_X \sim 10^{18}$  GeV and the agreement with the experiment requires the value of  $\sin^2 \theta_w$  at  $M_X$  to be in general very different from the usually assumed 3/8. [S0556-2821(96)02419-8]

PACS number(s): 12.10.Dm, 11.25.Mj, 14.60.Pq

#### I. INTRODUCTION

One of the main reasons to study supersymmetric theories is that they could alleviate the problem of the gauge hierarchy. The minimal supersymmetric extension of the standard model (MSSM) leads to a remarkable prediction [1]: The gauge couplings  $\alpha_3$ ,  $\alpha_2$ , and  $\alpha_1$  of SU(3)<sub>c</sub>, SU(2)<sub>L</sub>, and U(1)<sub>Y</sub>, respectively, are unified at a scale  $M_{\rm GUT} \approx 2 \times 10^{16}$ GeV [2] through the relation

$$\alpha_3(M_{\rm GUT}) = \alpha_2(M_{\rm GUT}) = \frac{5}{3}\alpha_1(M_{\rm GUT}).$$
 (1)

While very exciting, this result rests, however, on the hypothesis of the "big desert scenario," which states that "nothing" happens between the scale of supersymmetry breaking (~ TeV) and the scale of unification ( $\sim M_{GUT}$ ). At higher energies the normalization factor  $\frac{5}{3}$  of U(1)<sub>Y</sub> allows one to embed SU(3)<sub>c</sub>×SU(2)<sub>L</sub>×U(1)<sub>Y</sub> in a grand unified theory (GUT) group SU(5), SO(10), ....

There are obvious reasons to look beyond this big desert scenario. The main one is the possibility of detecting some new particles at future colliders. This needs the intermediate scale physics to be of the order of TeV, which is usually hard to obtain naturally. This is the case, for example, of extra gauge bosons or more exotic such as an extra dimension.

Another important motivation for intermediate scales is the question of neutrino masses. In the MSSM neutrino masses are made to vanish by hand, through the requirement of the absence of their right-handed partners. If the latter are present, the small values of neutrino masses could naturally come from the seesaw mechanism [3,4]. If right-handed neutrinos get Majorana masses at some scale  $M_{\nu_R}$ , one expects a mass matrix of the form

$$\begin{pmatrix} 0 & m_D \\ m_D & M_{\nu_R} \end{pmatrix}, \tag{2}$$

where  $m_D$  are the Dirac masses of the neutrino. If  $m_D \ll M_{\nu_R}$ , then the masses of the right- and left-handed neutrinos are  $M_{\nu_R}$  and  $m_D^2/M_{\nu_R} \ll M_{\nu_R}$ , respectively. For the seesaw mechanism to give us the neutrinos masses, the large scale  $M_{\nu_R}$  should be predicted. Normally, one associates this scale with the breaking of some (gauge) symmetry. For example, this can be naturally implemented if at some intermediate scale  $M_I$  the symmetry is enhanced<sup>1</sup> to a left-right group [5].

In the standard model, left-handed quarks and leptons are doublets of SU(2)<sub>L</sub>, while their right partners are singlets. In left-right models, the explicit violation of parity is replaced with a spontaneous one, rendering our world more symmetric. In fact, right-handed quarks and leptons (and thus neutrinos too) appear now also in doublet representations, but under another gauge group SU(2)<sub>R</sub>. To relate these models to the MSSM, one has to introduce two scales: the scale of parity breaking,  $M_{parity}$ , and the scale of SU(2)<sub>R</sub> breaking,  $M_R$ , with obviously  $M_R \leq M_{parity}$ . It is then natural to relate these scales as  $M_{\nu_R} \sim M_R$  and  $M_{parity} \sim M_R$  or  $M_{parity} \sim M_X$ .

The simplest realization of this idea needs the introduction of a Higgs triplet of  $SU(2)_R$ , usually denoted  $\Delta_R$ , which gets a vacuum expectation value  $M_R$ . Anomaly cancellation imposes the presence of another field  $\overline{\Delta}_R$ . If one wants to have  $M_{\text{parity}}$  of the order of  $M_R$ , triplets  $\Delta_L + \overline{\Delta}_L$  of  $SU(2)_L$  must also be introduced. A natural question to ask is, what happens to the gauge coupling unification when these new states are present?

In all previous analyses, the unification of these models was supposed to appear inside an SO(10) gauge group, and the above triplets arise from **126** representations. It is easy to

<sup>\*</sup>Electronic address: benakli@ictp.trieste.it

<sup>&#</sup>x27;Electronic address: goran@ictp.trieste.it

<sup>&</sup>lt;sup>1</sup>Hereafter, for us  $M_I$  will denote any intermediate scale,  $M_W \ll M_I \ll M_X$ , where  $M_X$  is the unification scale.

see that unification constraints then imply that  $M_R \sim M_{GUT}$ , thus preventing the existence of such an intermediate scale. Thus these models were studied without referring to unification. Moreover, new *ad hoc* and more complicated states are often introduced with the only purpose being to lead to SO(10) unification.

In this paper, we want to study the possibility of string unification with arbitrary Kac-Moody levels for these models. In fact, these simple models with triplets of  $SU(2)_R$  are the first phenomenologically interesting ones, well motivated by the seesaw mechanism, which would *need stringy unification*. It is well known that this type of unification does not require the existence of a GUT group (see below). We will see below that, in contrast to all previous analyses, values of  $\sin^2 \theta_w$  at the unification scale very different from 3/8 could be in agreement with the experimental data. We will also use this opportunity to make some useful comments on the models.

In Sec. II we review the problem of gauge couplings in string theory. This will also allow us to define our notations and strategy. In Sec. III we will discuss the models and then determine some sets of values for Kac-Moody levels leading to gauge couplings unification. Comments of the results and conclusions are given in Sec. IV.

### **II. STRINGY UNIFICATION OF GAUGE COUPLINGS**

#### A. String unification and the MSSM

Superstring theory is considered today as a good candidate for the unification of all known interactions, as we hope that it contains a finite theory of quantum gravity. It is thus always important to readdress the question of unification of gauge couplings in its context.

In heterotic string theories, all the gauge (and the gravitational) couplings are given by the expectation value of a particular field: the dilaton. Moreover, the four-dimensional space-time gauge symmetries are associated with an appropriate Kac-Moody algebra on the two-dimensional string world sheet [6]. To a Kac-Moody algebra corresponds its level, a positive parameter  $k_i$  (integer for non-Abelian groups), which determines the corresponding tree-level gauge coupling constant  $\alpha_i$  in terms of the four-dimensional string coupling  $\alpha_{st}$ ,  $\alpha_i = \alpha_{st}/k_i$ . Since at the string scale  $\alpha_i/\alpha_j = k_j/k_i$ , which are rational numbers, we have coined this rational unification. The values of the levels also constrain the allowed unitary representations present in the chiral massless spectrum [7].

A natural question to ask then is, what are the values of the levels  $k_3$ ,  $k_2$ , and  $k_1$  associated with the standard model gauge groups SU(3)<sub>c</sub>, SU(2)<sub>L</sub>, and U(1)<sub>Y</sub>, respectively? In most of the models built up to now,  $k_3 = k_2 = 1$ , while  $k_1 \ge \frac{5}{3}$ . These level-1 constructions have the nice feature of explaining the presence of only singlet and fundamental chiral representations in the standard model. However, they generally suffer from the presence of fractionally charged particles in the massless spectrum at the string level. Moreover, the most popular value of  $k_1 = \frac{5}{3}$  does not allow one to embed the MSSM in a GUT group because of the absence of chiral adjoint representations.

In principle, arbitrary higher Kac-Moody levels  $(k_3, k_2>1)$  are allowed. However, the corresponding string mod-

els have been found to be very difficult to build. They also allow the presence of larger representations, leading to phenomenological problems [8]. For these reasons, they have been mainly disregarded. Recently, some (still unsuccessful) attempts have been made to build such theories, with  $k_3=k_2=\frac{3}{5}k_1=2$ , trying to embed the standard model in some GUT group and explain the *a priori arbitrary* normalization  $k_1$  of U(1)<sub>Y</sub> [9,10].

It is worth noticing that the (field theoretical) direct unification of gauge couplings, which is the remarkable prediction of the big desert scenario, takes place at a scale  $M_{\rm GUT} \approx 2 \times 10^{16}$  GeV. However, in contrast with the case in field theory, the unification scale  $M_{\rm SU}$  in string models can be predicted. Within some large class of models, it was found to be  $M_{\rm SU} \approx 2 \sqrt{\alpha_{\rm st}} \times 10^{18}$  GeV, nearly two orders of magnitude bigger than  $M_{\rm GUT}$  [11]. Some ideas have been presented to reconcile the two scales. They fall into two categories.

In the first category, one tries to push down  $M_{SU}$  toward  $M_{GUT}$  either by invoking large string threshold corrections [12] or by arguing on the possibility of a unified evolution of the gauge couplings between these two scales. These solutions have as good feature a small ratio  $M_{GUT}/M_{Planck}$ , which could be associated with some explanations of the fermions mass spectrum or with the strength of the fluctuations in the Cosmic Background Explorer (COBE) observations. Unfortunately, the needed large thresholds do not seem to appear naturally.

In the second category, one tries to push  $M_{GUT}$  toward  $M_{SU}$ . This involves either the modification of the hypercharge normalization (such as<sup>2</sup>  $k_1 \approx \frac{4}{3}$ ) [13] or the presence of some extra particles in some intermediate scale(s) [14]. These particles could be standard like or exotic fractionally charged ones [15]. These last possibilities are a clear abandon of the big desert scenario [16]. In fact, they seem natural solutions as string models usually contain more particles than the MSSM ones in the massless spectrum at the string level, and some of them could be lying somewhere between the TeV and the string scales.

### B. Stringy unification with one intermediate scale

Below, we would like to investigate the (next to minimal) situation where some intermediate scale appears corresponding to some symmetry breaking, with a minimal particle content motivated by some phenomenological reasons. We ask about the possible existence of string unification for these models. This corresponds to determining if there exists a set of levels  $k_i$ 's compatible with it. Most of such models contain some large representations that need some high-level Kac-Moody algebras [17]. With our actual knowledge of conformal field theories, building such compactifications is a challenging problem. In realistic models, we would also have to explain why and how only the wanted particles appear at the low and intermediate scales. In particular, some representations (of smaller conformal weight than the ones considered for example) could (probably would) appear and spoil

<sup>&</sup>lt;sup>2</sup>Such normalization does not appear in known level-1 constructions.

our analysis. In view of the above discussion, we will not try to answer these problems, but being less ambitious, we will constrain ourselves to the analysis of the gauge coupling unification in such hypothetical<sup>3</sup> cases. We plan to study the possible effects of extra matter states in the future.

We restrict our analysis to the one-loop unification of gauge couplings in some particular supersymmetric models. We mainly consider the possibility of one intermediate scale  $M_1$  lying in the region between the supersymmetry-breaking scale  $m_s$  and the unification scale  $M_X$ .

Below  $M_I$ , the gauge group is the standard model  $SU(3)_c \times SU(2)_L \times U(1)_Y$  with corresponding Kac-Moody levels  $k_3$ ,  $k_2$ , and  $k_1$  satisfying some constraints arising from the particle content of the model below the string scale [7]. More precisely, a representation  $(r_1, r_2, \ldots, r_n, q_1, \ldots, q_m)$  of  $SU(N_1) \times \cdots \times SU(N_n) \times U(1)_1 \times \cdots \times U(1)_m$ , of levels  $k_{N_1} \cdots k_{N_m}, k_{1_1} \cdots k_{1_n}$ , has a conformal weight

$$h = \sum_{i=1}^{n} \frac{C(r_i)}{k_{N_i} + C(N_i)} + \sum_{j=1}^{m} \frac{q_j^2}{k_{1_j}}.$$
 (3)

This state to be present in the string massless spectrum needs to have  $h \leq 1$ .

In the minimal case of MSSM content,  $k_3$  and  $k_2$  are positive integers while  $k_1 \ge 1$ . Notice that the case where all the levels go to infinity corresponds to the field theory limit as the string scale goes to infinity and all the representations are allowed. Unification is meaningless in this limit.

The associated effective couplings at the supersymmetry breaking scale  $m_s$  are given, at one loop, by

$$\frac{1}{\alpha_i(m_s)} = \frac{k_i}{\alpha_{\rm st}} + \frac{b_i}{2\pi} \ln\left(\frac{M_I}{m_s}\right) + \frac{b_i'}{2\pi} \ln\left(\frac{M_X}{M_I}\right),\tag{4}$$

where  $b_i$  and  $b'_i$  are the one-loop  $\beta$ -function coefficients in the corresponding energy domains. They are given by

$$b_i = -3C(G_i) + \sum_{\text{reps } R_i} T(R_i),$$
(5)

where the quadratic Casimir  $C(G_i)$  of the group  $G_i$  equals N for SU(N) and N-2 for SO(N). The index  $T(R_i)$  of the matter representation  $R_i$  is equal to  $\frac{1}{2}$  for chiral supermultiplets in the fundamental representation of SU(N), while it is given by the sum of the squares of charges in the case of U(1).

The perturbative unification at the scale  $M_X$  imposes a strong constraint  $\alpha_{st}/k_i < 1$ , which rules out the possibility of a low  $M_I$  scale in most of our models.

We would like to have some reasonable constraints on the allowed values of  $k_i$ 's. We first notice that the string unification scale is predicted to be of the order of

$$M_{SU} \approx 2\sqrt{\alpha_{st}} \times 10^{18} \text{ GeV}$$
$$= 2\sqrt{\alpha_i k_i} \times 10^{18} \text{ GeV}$$
$$= 2\sqrt{\alpha_3 k_3} \times 10^{18} \text{ GeV}. \tag{6}$$

We would like to keep  $M_{SU} \approx 10^{18} \text{ GeV} < M_{Planck}$ . As in all our cases,  $\alpha_3 \ge 1/25$ , this means that  $k_3$  should not exceed a number of order 100. A stronger constraint could come if we assume the existence of an extra non-Abelian group with a smaller level  $k < k_3$  than  $k_3 \le 25k$ . Moreover, another condition that could be imposed on the levels is

$$\frac{1}{3}k_3 + \frac{1}{4}k_2 + \frac{1}{4}k_1 =$$
integer, (7)

which is required in order to avoid the appearance of fractionally electrically charged particles in the massless spectrum [7]. If this condition is not satisfied, these undesired particles could, however, still get masses of the order<sup>4</sup>  $M_X$ . Notice that what we mean by charge quantization is that all color singlet states have a charge which is integer multiple of the electron charge. In orbifold compactifications, for instance, it has been shown that a weaker charge quantization, where the elementary charge is a fraction 1/N (N < 12) of the electron charge, can be imposed [18].

In Eq. (4) we have absorbed the unknown string threshold corrections (usually denoted  $\Delta_i$ ) in the definition of  $M_X$ , which can then be different from the computed value  $M_{SU}$ . While the *natural value of*  $M_X$  *is*  $\simeq 10^{18}$  GeV, for practical computations, we allow it to take values between  $10^{16}$  GeV and (more natural value)  $10^{18}$  GeV. The former value has the advantage of introducing naturally a small ratio in the theory and thus it is often considered as a good value in the literature.

In addition to  $M_X$ , our other inputs are of two kinds: As experimental inputs without our one-loop approximations, the values of the strong, electromagnetic coupling constants at  $m_Z$ ,  $\alpha_s \equiv \alpha_3$ ,  $\alpha_{\rm em} = \alpha_1 \alpha_2 / (\alpha_1 + \alpha_2)$ , respectively, and the weak angle  $s \equiv \sin^2 \theta_w = \alpha_1 / (\alpha_1 + \alpha_2)$  will be taken in the range

$$\alpha_{\rm em} = 1/128 \quad 0.230 \le s \le 0.233, \text{ and } 0.11 \le \alpha_s \le 0.13;$$
(8)

as theoretical inputs, we take  $m_S = m_Z$ , which is usually a good approximation at one loop. Below  $M_I$ , we will make the assumption that the massless spectrum is the one of the MSSM with three generations  $(n_g=3)$  and two Higgs doublets  $(n_H=2)$ . The coefficients  $b_i$  take the values

<sup>&</sup>lt;sup>3</sup>Neither the MSSM nor its phenomenological viable extensions (even some versions with extra chiral matter with the *appropriate spectrum* needed to raise the unification scale) have been by now derived from strings.

<sup>&</sup>lt;sup>4</sup>If they are confined at very high scale by some extra gauge factor [15,19] or get a mass through one of the mechanisms discussed in [17].

$$b_3 = -9 + 2n_g = -3, \quad b_2 = -6 + 2n_g + n_H/2 = 1, \quad b_1 = (10/3)n_g + n_H/2 = 11.$$
 (9)

Our strategy is to solve Eq. (4) in order to get the ratios  $k_1/k_2$  and  $k_2/k_3$  as a function of  $\ln(M_X/M_l)$ :

$$\frac{k_1}{k_2} = \frac{1 - s - (\alpha_{\rm em}/2\pi)[b_1 \ln(M_I/m_Z) + (b_1' - b_1)\ln(M_X/M_I)]}{s - (\alpha_{\rm em}/2\pi)[b_2 \ln(M_I/m_Z) + (b_2' - b_2)\ln(M_X/M_I)]},$$
(10)

$$k_{2} = \frac{(s/\alpha_{\rm em}) - (1/2\pi) [b_{2} \ln(M_{I}/m_{Z}) + (b_{2}' - b_{2}) \ln(M_{X}/M_{I})]}{(1/\alpha_{s}) - (1/2\pi) [b_{3} \ln(M_{I}/m_{Z}) + (b_{3}' - b_{3}) \ln(M_{X}/M_{I})]}.$$
(11)

By plotting these functions as well as the values of  $\alpha_{st}/k_i$ , which have to remain small, one can read the allowed intermediate scale and the corresponding ratios of levels.

### **III. MODELS WITH INTERMEDIATE MASS SCALES**

#### A. Models

We want to focus on a single intermediate scale, although we shall also discuss a case with two such scales. Our analysis continues the analysis of [17] and it is parallel to the ones for SO(10) unification. There are four possible rank-5 gauge groups at the intermediate scale, with their respective levels. (a)  $SU(3)_c \times SU(2)_L \times U(1)_R \times U(1)_{B-L}$  with levels  $k_3$ ,  $k_2$ ,  $k_R$ , and  $k_{B-L}$ . This model has two nice features if the scale  $M_I$  lies in the TeV region. On the one hand, it predicts the observation of a new vector boson at future colliders, and on the other hand, the B-L gauge symmetry forbids the appearance of (nonrenormalizable) operators leading to fast proton decay.

(b)  $SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$  with levels  $k_3, k_2, k_{2R}$ , and  $k_{B-L}$ . In addition to the aesthetic beauty of parity as a symmetry [5], these models can explain small neutrino masses through the seesaw mechanism [4]. Two cases are *a priori* allowed and will be analyzed below. The first case corresponds to the direct breaking to the standard model at the scale  $M_I$ . If this happens at the TeV scale, then B-L symmetry protects the proton from decay [21]. The model predicts then the observation of extra neutral and charged gauge bosons. However, in the natural approximation that the neutrino Dirac masses are of the same order as the corresponding charged lepton masses, the seesaw mechanism leads to a spectrum of too heavy left-handed neutrino masses overclosing the universe.

Namely, with  $m_D \simeq m_l$  and  $M_{\nu_p} \simeq 1$  TeV, we predict

$$m_{\nu_e} \simeq 1 \text{ eV}, \quad m_{\nu_{\mu}} \simeq 10 \text{ keV}, \quad m_{\nu_{\tau}} \simeq 1 - 10 \text{ MeV}.$$
(12)

Now,  $\nu_{\tau}$  can in principle decay through the weak currents,  $\nu_{\tau} \rightarrow e^+ e^- \nu_e$ , by assuming the Cabibbo-Kobayashi-Maskawa- (CKM-) like matrix in the leptonic sector. The case of  $\nu_{\mu}$  is more problematic, and it requires the presence of  $\Delta_L$ , a left-handed analogue of  $\Delta_R$  [22].

On the other hand, if all  $\nu$ 's are lighter than 10–100 eV, then we have no problem with the overclosure of the universe, and  $\Delta_L$ 's are not necessarily present. In this case we have a constraint  $M_{\nu_R} \ge 10^8$  GeV (assuming the relation  $m_D = m_l$  as in the above). If we wish to have the MSW explanation of the solar neutrino puzzle (through  $\nu_e$ - $\nu_{\mu}$  oscillations), a preferred value for the intermediate scale becomes  $M_I \approx 10^{10}$  GeV, with

$$m_{\nu_e} \simeq 10^{-7} \text{ eV}, \quad m_{\nu_{\mu}} \simeq 10^{-3} \text{ eV}, \quad m_{\nu_{\tau}} \simeq 1 \text{ eV},$$
(13)

in which case  $\nu_{\tau}$  can play a role of dark matter (or some fraction of it). Because of the uncertainties in  $m_D$ , we quote this as  $M_I \approx 10^8 - 10^{12}$  GeV. In this case, B - L does not protect the proton from decaying too fast.

A spontaneous breaking of *R* parity giving a vacuum expectation value (VEV) to the sneutrino  $\langle \tilde{\nu}_R \rangle$  could lead to proton decay through a dimension-4 operator in the superpotential. One then could introduce a discrete symmetry or look for models where such operators are forbidden by some string selection rules.

Another possibility to get rid of this problem is to break the group in two steps: first, to  $SU(3)_c \times SU(2)_L$  $\times U(1)_R \times U(1)_{B-L}$  and then to break it again at the TeV scale to the standard model group, protecting the proton from decaying too fast. We will investigate a minimal version of this scenario too.

(c)  $SU(4)_c \times SU(2)_L \times U(1)_R$  with levels  $k_4$ ,  $k_2$ , and  $k_R$ . This partially unified model provides a symmetry between quarks and leptons.

(d)  $SU(4)_c \times SU(2)_L \times SU(2)_R$  with levels  $k_4$ ,  $k_2$ , and  $k_{2R}$ . This Pati-Salam partial unification is the minimal unification based on a simple group of the standard group [23] and offers both left-right symmetry and quark-lepton unification.

#### **B.** Unification constraints

# 1. $SU(3)_c \times SU(2)_L \times U(1)_R \times U(1)_{B-L}$

The minimal content of matter is just the MSSM spectrum plus three chiral superfields with quantum numbers  $\nu_R = (1,1,-1/2,1)$  under the  $SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$  group, which can then be identified as right-handed neutrinos. The condition imposed by the presence of these states on the Kac-Moody levels is

$$\frac{1}{4k_R} + \frac{1}{k_{B-L}} \le 1.$$
 (14)

The level  $k_1$  given by

$$k_1 = k_R + \frac{k_B - L}{4} \tag{15}$$

can take the standard value  $k_1 = 5/3$ . When one of the new fields  $\langle \tilde{\nu}_R \rangle$  gets a VEV, it breaks one combination U(1)' of U(1)<sub>R</sub>×U(1)<sub>B-L</sub>, leading to an extra Z' massive vector boson at the scale  $M_I$ , while it leaves the hypercharge U(1)<sub>Y</sub> with  $Y = Q_R + (B-L)/2$  unbroken, where  $Q_R$  is the generator of U(1)<sub>R</sub>. How would the neutrinos get a mass in this case? One possibility is then the mechanism suggested in [20]. A seesaw mechanism is obtained through the mass matrix between the left and right neutrinos and the gaugino partner of Z'. This, however, gives a mass only to the neutrino whose partner got a VEV. The other neutrinos presumably get masses through some nonrenormalizable operators.

Another possibility is to explain neutrino masses by the usual seesaw mechanism. The gauge symmetry breaking is achieved by giving a VEV to some extra state with gauge numbers (1,1,1,-2) [one also introduces a (1,1,-1,2) representation to cancel the U(1) anomalies]. Then one has the condition

$$\frac{1}{k_R} + \frac{4}{k_{B-L}} \leq 1, \tag{16}$$

which implies that  $k_R > 1$  and  $k_{B-L} > 4$ , and so  $k_1 > 2$ . As the new states couple equally to all the neutrinos, they generate small masses to all of them through the seesaw mechanism. An important question to raise is, what are the expected values for the Dirac masses? In the general case we are considering, it is not possible to make a model-independent statement. As the right-handed neutrinos and electrons are *a priori* independent and could come from different sectors of the string compactification, having different moduli dependence, the relative Yukawa couplings could be very different. Having smaller values for  $m_D$  corresponding to  $\nu_{\mu}$ would allow a low scale  $M_I$ .

In any case the gauge couplings of  $SU(3)_c \times SU(2)_L \times U(1)_Y$  are not affected and evolve in the same way as in the MSSM, with the new appropriate normalizations  $k_i$ . The scale  $M_I$  is only constrained by collider experiments, and the corresponding gauge coupling of U(1)' can be now computed, because at  $M_X$  it is equal to the one of  $U(1)_Y$  ( $k' = k_R + k_{B-L}/4 = k_1$ ) and the associated  $\beta$ -function coefficient is known:  $b_1$  or  $b_1 + 2 = 13$  in the first and second examples described above, respectively. At low energies, the corresponding coupling is equal or smaller than the hypercharge one.

# 2. $SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$

This case is more interesting because it is more restricting. The minimal matter content of the model is  $n_g$  generations of matter representations Q = (3,2,1,1/3),  $Q^c = (3,1,2,-1/3)$ , L = (1,2,1,-1), and  $L^c = (1,1,2,1)$ , which correspond to the quarks and leptons. There is also a Higgs sector consisting in  $n_{22}$  bidoublets (1,2,2,0) and  $n_{\Delta_L}$  pairs of  $SU(2)_L$  triplets { $\Delta_L = (1,3,1,2) + \overline{\Delta}_L = (1,3,1,-2)$ } as well as  $n_{\Delta_R}$  pairs of  $SU(2)_R$  triplets { $\Delta_R = (1,1,3,-2)$ }  $+ \overline{\Delta}_R = (1,1,3,2)$ } and a possible parity odd singlet to make

TABLE I. Examples of values of Kac-Moody levels  $(k_3, k_2, k_1)$  of  $SU(3)_c \times SU(2)_L \times U(1)_Y$ , allowing for  $SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$  at an intermediate scale of order  $M_I \approx 10^{12}$  GeV.

$(n_{22}, n_{\Delta_L}, n_{\Delta_R})$	$M_X$ (GeV)	$M_I$ (GeV)	$(k_1, k_2, k_3)$
(1,0,1)	10 <sup>18</sup>	10 <sup>11,5</sup>	(9,9,10)
(1,0,1)	1017	$10^{11}$	(6,5,5)
(1,0,1)	1016	$10^{11}$	(7,5,5)
(2,0,1)	$10^{18}$	$10^{11.5}$	(10,10,12)
(2,0,1)	$10^{17}$	$10^{12}$	(25/2,10,11)
(2,0,1)	$10^{16}$	$10^{12}$	(6,4,4)
(1,1,1)	$10^{18}$	$10^{12}$	(6,8,14)
(1,1,1)	$10^{17}$	$10^{11}$	(6,6,10)
(1,1,1)	$10^{16}$	$10^{10}$	(5,4,6)
(2,1,1)	$10^{18}$	$10^{12}$	(9/2,6,12)
(2,1,1)	$10^{17}$	$10^{11}$	(5,5,9)
(2,1,1)	$10^{16}$	$10^{10}$	(5,4,7)

 $n_{\Delta_L}=0$  [24]. The  $\Delta_R$  field, when it gets a VEV, breaking SU  $(2)_R \times U(1)_{B-L}$  to  $U(1)_Y$ , gives a Majorana mass to the right-handed neutrino.

The coefficients of the  $\beta$  functions of SU(3)<sub>c</sub>, SU(2)<sub>L</sub>, and  $U(1)_Y$  above  $M_I$  are

$$b'_{3} = -9 + 2n_{g} + n_{3}, \quad b_{2} = -6 + 2n_{g} + n_{22} + 4n_{\delta_{L}} + n_{2},$$
  
$$b'_{1} = -6 + (10/3)n_{g} + n_{22} + 6n_{\Delta L} + 10n_{\Delta \delta_{R}} + n_{1}, \quad (17)$$

respectively. In all of the discussions below, the numbers  $n_1$ ,  $n_2$ , and  $n_3$  will parametrize the (unknown) contributions of extra particles that could appear at this scale. Unless explicitly stated otherwise, we take  $n_g=3$ ,  $n_1=n_2=n_3=0$ . The normalization of U(1)<sub>Y</sub> is given by

$$k_1 = k_{2R} + \frac{k_{B-L}}{4}.$$
 (18)

The minimal values of levels are  $k_{2R} \ge 2$ ,  $k_{B-L} \ge 8$  (hence  $k_1 \ge 4$ ),  $k_2 \ge 2$  if  $n_{\Delta_L} \ne 0$ . One can define two possible cases: one with equal couplings for SU(2)<sub>L</sub> and SU(2)<sub>R</sub>, implying  $k_2 = k_{2R}$  and  $k_1 \ge k_2 + 2$ , and the other with different couplings  $k_2 \ne k_{2R}$ . In the more symmetric case, one has the constraint  $k_1/k_2 \ge 1$ , the equality corresponding to  $k_2 \rightarrow \infty$ . This constraint is too restrictive, and it leads usually to large intermediate scales. We will relax this constraint and allow for  $k_2 \ne k_{2R}$ . Hence  $k_1/k_2 \le 1$  is allowed, and the limits will come from the perturbative unification limit on  $\alpha_i$  at the unification scale.

We have made the analysis for different particles content and different unification scales, and we present our results in Tables I, II, and III and Figs. 1–13. In particular, we studied the following cases.

(i)  $n_{22}=1$ ,  $n_{\Delta_R}=1$ , and  $n_{\Delta_L}=0$ . The results are plotted in Figs. 1 and 2. This model could be made more symmetric if one explains  $n_{\Delta_L}=0$  through the introduction of some parity odd singlet [24]. While from the figures one can read that the intermediate scale as low as a few TeV is allowed by the running of the couplings, the neutrino spectrum forbids it. In fact, as discussed above, one gets a too heavy  $\nu_{\mu}$ , which is

TABLE II. Examples of values of Kac-Moody levels  $(k_3, k_2, k_1)$  of  $SU(3)_c \times SU(2)_L \times U(1)_Y$ , allowing for  $SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$  at an intermediate scale, leading to charge quantization.

$(n_{22}, n_{\Delta_L}, n_{\Delta_R})$	$M_X$ (GeV)	$M_I$ (GeV)	$(k_1,\!k_2,\!k_3)$
(1,0,1)	10 <sup>18</sup>	107	(22/3,10,11)
(1,0,1)	1017	$10^{8}$	(12,12,12)
(1,0,1)	1016	$10^{8.5}$	(19/3,5,5)
(2,0,1)	$10^{18}$	$10^{11.5}$	(10, 10, 12)
(2,0,1)	$10^{17}$	$10^{12}$	(44/5,8,9)
(2,0,1)	10 <sup>16</sup>	$10^{11}$	(44/3,12,13)
(1,1,1)	$10^{18}$	$10^{13}$	(32/3,12,19)
(1,1,1)	$10^{17}$	$10^{9}$	(8,12,24)
(1,1,1)	10 <sup>16</sup>	$10^{8.5}$	(12,12,21)
(2,1,1)	$10^{18}$	$10^{13}$	(32/3,12,22)
(2,1,1)	$10^{17}$	$10^{11.5}$	(40/3,12,20)
(2,1,1)	10 <sup>16</sup>	10 <sup>11.5</sup>	(52/3,12,17)

stable because of the absence of  $\Delta_L$ , thus overclosing the universe. We then have a constraint  $M_I \ge 10^7$  GeV as discussed above.

If one insists on the equality of left and right couplings, then for  $k_2 \approx 10$ , we get  $M_I \ge 10^{14.5}$ ,  $10^{11}$ , and  $10^8$  GeV for  $M_X = 10^{18}$ ,  $10^{17}$ , and  $10^{16}$  GeV, respectively. A possible set, for  $M_X = 10^{18}$  GeV, is  $k_3 = 11$ ,  $k_2 = k_{2R} = 10$ , and  $k_{B-L} = 8$ .

(ii)  $n_{22}=2$ ,  $n_{\Delta_R}=1$ , and  $n_{\Delta_L}=0$ . This case is quite similar to the first one. As we are interested in large  $M_I$  ( $\approx 10^{12}$  GeV), the ratio  $k_1/k_2$  is not sensible to the number of bid-oublets as shown in Fig. 3. The addition of the extra bidoublet is helpful to generate a correct Cabibbo angle. We present the corresponding results in Figs. 4 and 5.

(iii)  $n_{22}=1$ ,  $n_{\Delta_R}=1$ . This is the minimal fully left-right symmetric model allowing for seesaw "explanation" of the neutrino masses. For this model we present our results in Figs. 6 and 7. From these figures, we can see that now the scale  $M_I$  cannot be as low as 1 TeV; otherwise, the hyper-charge coupling will blow up before the scale  $M_X$ . In order that this does not happen, we get the lower value of  $M_I \approx 10^{10}$  GeV.

(iv)  $n_{22}=2$ ,  $n_{\Delta_R}=1$ , and  $n_{\Delta_L}=1$ . This is the most popular model. As the perturbative condition of the couplings does not allow small  $M_I$ , the contribution of the second bidoublet is small. The results are displayed in Figs. 8 and 9. The end result is again that the region  $10^{10}-10^{12}$  GeV is perfectly acceptable.

Thus one can have a realistic left-right model in the context of strings with the MSW mechanism and  $\nu_{\tau}$  as (some portion of) the dark matter of the universe.

TABLE III. Examples of values of Kac-Moody levels  $(k_3, k_2, k_1)$  of  $SU(3)_c \times SU(2)_L \times U(1)_Y$ , allowing for  $SU(4)_c \times SU(2)_L \times U(1)_R$  at an intermediate scale and charge quantization.

$\overline{M_X (\text{GeV})}$	$M_I$ (GeV)	$(k_1, k_2, k_3)$
10 <sup>18</sup>	10 <sup>13</sup>	(4,20,18)
$10^{17}$	$10^{12}$	(5,15,12)
10 <sup>16</sup>	$10^{11}$	(4,8,6)



FIG. 1. Values of  $1/\alpha_1$ ,  $1/\alpha_2$ , and  $1/\alpha_3$  at the unication scale  $M_X = 10^{18}$  GeV as a function of the position of the intermediate scale  $M_I$ , for model (b i) (the lines are doubled to represent minimal and maximal values).

Tables I–III give some possible values of the Kac-Moody levels consistent with  $M_I$  of the order of  $10^{10}-10^{12}$  GeV or electric charge quantization, respectively.

(v) We would like to investigate the possibility (from the gauge coupling unification point of view) to have a low-lying (B-L)-breaking scale around the TeV, protecting the proton from decaying, and a left-right breaking scale at an intermediate scale. These two intermediate breaking scales could be achieved, for example, with the following set of representations (in addition to the quarks, leptons, and Higgs bosons).

Between  $M_I$  and  $M_X$ , we have  $n_{\Delta L}$  pairs of  $SU(2)_L$  triplets  $\{\Delta_L = (1,3,1,2) + \overline{\Delta}_L = (1,3,1,-2)\}$ , as well as  $n_{\Delta_R}$  pairs of  $SU(2)_R$  triplets  $\{\Delta_R = (1,1,3,-2) + \overline{\Delta}_R = (1,1,3,2)\}$  and



FIG. 2. Values of  $k_1/k_2$  and  $k_2/k_3$ , as a function of the position of the intermediate scale  $M_I$ , for model (b i) with  $M_X = 10^{18}$  GeV (the lines are doubled to represent minimal and maximal values).



FIG. 3. Variation of the values of  $k_1/k_2$  and  $k_2/k_3$ , as a function of the position of the intermediate scale  $M_I$ , when the number of bidoublets is increased to two [model (b i) $\rightarrow$ (b ii) with  $M_X = 10^{18}$  GeV (the lines are doubled to represent minimal and maximal values)].

 $n_{T_L}$  SU(2)<sub>L</sub> triplets  $T_L = (1,3,1,0)$  and corresponding  $n_{T_R}$  SU(2)<sub>R</sub> triplets  $T_R = (1,1,3,0)$ . At  $M_I$  the neutral component of  $T_R$  gets a VEV and it breaks SU(2)<sub>R</sub> to U(1)<sub>R</sub>. The  $\beta$ -function coefficients are given by

$$b'_{3} = -9 + 2n_{g} + n_{3},$$

$$b_{2} = -6 + 2n_{g} + n_{22} + 4n_{\Delta_{L}} + 2n_{T_{L}} + n_{2},$$

$$b'_{1} = -6 + (10/3)n_{g} + n_{22} + 6n_{\Delta_{L}} + 10n_{\Delta_{R}} + 2n_{T_{R}} + n_{1}.$$
(19)



FIG. 4. Values of  $1/\alpha_1$ ,  $1/\alpha_2$ , and  $1/\alpha_3$  at the unication scale  $M_X = 10^{18}$  GeV as a function of the position of the intermediate scale  $M_I$ , for model (b ii) (the lines are doubled to represent minimal and maximal values).



FIG. 5. Values of  $k_1/k_2$  and  $k_2/k_3$ , as a function of the position of the intermediate scale  $M_I$ , for model (b ii) with  $M_X = 10^{18}$  GeV (the lines are doubled to represent minimal and maximal values).

For our analysis we take  $n_g = 3$ ,  $n_1 = n_2 = n_3 = 0$ ,  $n_{22} = 1$  or 2,  $n_{\Delta_R} = 1$ ,  $n_{\Delta_L} = 1$ , and  $n_{T_L} = n_{T_R} = 1$ .

At  $M_I$  where the new field  $T_R$  is supposed to get a VEV, some of the above particles become massive and decouple from the running of the coupling constants below  $M_I$ .

A complete analysis of the minimization of the full superpotential shows that below  $M_I$ , in addition to the particle content of the standard model, only the neutral component  $\Delta_R^0$  of  $\Delta_R$  and all of  $\Delta_R$  and  $\overline{\Delta}_L$  may remain massless. We found that the intermediate scale is then at most two orders of magnitude below  $M_X$ . An exhaustive study of the model, including the phenomenological implications, will be presented in a future publication [25].



FIG. 6. Values of  $1/\alpha_1$ ,  $1/\alpha_2$ , and  $1/\alpha_3$  at the unication scale  $M_X = 10^{18}$  GeV as function of the position of the intermediate scale  $M_I$ , for model (b iii) (the lines are doubled to represent minimal and maximal values).



FIG. 7. Values of  $k_1/k_2$  and  $k_2/k_3$ , as a function of the position of the intermediate scale  $M_1$ , for model (b iii) with  $M_x = 10^{18}$  GeV (the lines are doubled to represent minimal and maximal values).

### 3. $SU(4)_c \times SU(2)_L \times U(1)_R$

In this case, the quarks and leptons are unified in the same representations of  $SU(4)_c \times SU(2)_L \times U(1)_R$ , leading to predict the existence of the right neutrinos and the size of their expected Dirac masses. In addition to the quarks and leptons in the  $n_g$  representations,

$$(4,2,1) + (4,1,-1/2) + (4,1,1/2),$$
 (20)

and  $n_h$  electroweak Higgs representations

$$(1,2,-1/2) + (1,2,1/2),$$
 (21)



FIG. 8. Values of  $1/\alpha_1$ ,  $1/\alpha_2$ , and  $1/\alpha_3$  at the unication scale  $M_X = 10^{18}$  GeV as a function of the position of the intermediate scale  $M_I$ , for model (b iv) (the lines are doubled to represent minimal and maximal values).



FIG. 9. Values of  $k_1/k_2$  and  $k_2/k_3$ , as a function of the position of the intermediate scale  $M_I$ , for model (b iv) with  $M_X = 10^{18}$  GeV (the lines are doubled to represent minimal and maximal values).

we have  $n_{10}$  pairs of representations (10,1,-1) and  $(\overline{10},1,1)$  necessary to break  $SU(4)_c \times SU(2)_L \times U(1)_R$  to the  $SU(3)_c \times SU(2)_L \times U(1)_Y$ .

The standard model levels are given by

$$k_1 = \frac{2}{3}k_4 + k_R, \quad k_3 = k_4. \tag{22}$$

The associated  $\beta$ -function coefficients above the intermediate scale are

$$b'_{3} = -12 + 2n_{g} + 6n_{10} + n_{3}, \quad b_{2} = -6 + 2n_{g} + n_{h} + n_{2},$$
  
$$b'_{1} = -8 + (10/3)n_{g} + n_{h} + 24n_{10} + n_{1}. \tag{23}$$

We made the analysis for  $n_g=3$ ,  $n_1=n_2=n_3=0$ ,  $n_h=1$ , and  $n_{10}=1$ . We display in Figs. 10 and 11 the results



FIG. 10. Values of  $1/\alpha_1$ ,  $1/\alpha_2$ , and  $1/\alpha_3$  at the unication scale  $M_X = 10^{18}$  GeV as a function of the position of the intermediate scale  $M_I$ , for model (c) (the lines are doubled to represent minimal and maximal values).



FIG. 11. Values of  $k_1/k_2$  and  $k_2/k_3$ , as a function of the position of the intermediate scale  $M_I$ , for model (c) with  $M_X = 10^{18}$  GeV (the lines are doubled to represent minimal and maximal values).

for  $M_{\chi} \approx 10^{18}$  GeV. In Table III, we give some values of levels, allowing intermediate scale and electric charge quantization.

# 4. Pati-Salam group $SU(4)_c \times SU(2)_L \times SU(2)_R$

The Pati-Salam group needs large representations to get broken to the standard model one. The contribution of these particles to the running of  $U(1)_Y$  would make this coupling blow up very close to  $M_I$ . In fact,  $M_I$  is typically two to three orders of magnitude below  $M_X$  when the latter goes from  $10^{18}$  to  $10^{16}$  GeV. Hence it is not very appealing as an intermediate scale, since only for a low value of  $M_X$  does it become interesting for neutrino physics.

### IV. DISCUSSION AND CONCLUSION

Our results are mainly qualitative and need some comments. The maximal values of  $k_1/k_2$  are obtained in the absence of an intermediate scale. They vary between 1.4 for  $M_X \simeq 10^{18}$  GeV and  $\simeq 5/3$  for  $M_X \simeq 10^{16}$  GeV, confirming previous analyses for the MSSM. In the minimal models, we found that interesting values of the scale  $M_I$  are allowed. Furthermore, since renormalization group equations usually make the  $U(1)_{Y}$  coupling increase with energy faster than the SU(2)<sub>L</sub> one,  $\sin^2 \theta_w$  at  $M_X$  is typically larger than 3/8. Also, the necessary levels are often large, especially when one imposes the condition (7) for charge quantization. The cases of large values of  $k_3$  could be improved if one allows the presence of an octet of SU(3), for example. An increase of precision of our analysis, by improving the uncertainty on  $\alpha_s$  and by a serious two-loop analysis with the thresholds taken into account, could constrain the allowed values of the unification scale or intermediate scale for some of these simple models. For instance, in model  $(b \ i)$ ,  $k_3/k_2$  is constant, and if it takes a value such as 1.02, we cannot associate it with two small integers  $k_3$  and  $k_2$  ( $k_3 \leq 25$ ).

We would like to comment about the actual status of string model building of higher Kac-Moody level models.



FIG. 12. Variation of the values of  $1/\alpha_1$ ,  $1/\alpha_2$ , and  $1/\alpha_3$  at the unication scale  $M_X = 10^{18}$  GeV as a function of the position of the intermediate scale  $M_I$ , for model (b iv) when B-L charged singlets are introduced.

The actual known trick to get such models is to notice that if you take the "diagonal" part of the product of n factors of the same group G at level 1, you generate a gauge group G of level n. This method cannot generate the minimal spectrum considered above. For instance, a triplet of SU(2) comes from the product of two doublets and thus is always accompanied by a singlet. We have considered the effect of the additional singlets carrying B-L charges on our analysis. We found a notable effect. As an example, the deviation for model (b iv) is plotted in Figs. 12 and 13.

We would also like to compare with the case of usual supersymmetric SO(10) GUT. That case corresponds to  $k_1/k_2 = \frac{5}{3}$  and  $k_3 = k_2$ . From the analysis of our plots, it is obvious that there cannot be an intermediate scale with the



FIG. 13. Variation of values of  $k_1/k_2$  and  $k_2/k_3$ , as a function of the position of the intermediate scale  $M_I$ , for model (b iv) with  $M_X = 10^{18}$  GeV when B-L charged singlets are introduced.

content considered in this paper. The introduction of an intermediate scale would there necessitate an *ad hoc* introduction of a set of particles necessary for *fitting* the experimental data with a possible unification in an SO(10). From this point of view, the beauty of the prediction of the MSSM for a unification of couplings is totally lost.

Two kinds of string constructions of these GUT's have been investigated by now, using the method sketched above. The first uses fermionic constructions [26]. In this case, it has been obtained that possible Kac-Moody levels (normalization of any *non-Abelian* groups) are k=1,2,4,8. This allows only for ratios  $k_3/k_2$  equal to  $\frac{1}{8}, \frac{1}{4}, \frac{1}{2}, 1, 2, 4, 8$ , which is not satisfied in most of our cases.

The other analysis uses orbifold constructions [10], and it indicates that some extra states, the part of the adjoint that are not destroyed by the Higgs mechanism, must remain massless at the GUT scale. They are in representations

$$(8,1,0) + (6,1-2/3) + (6,1,2/3) + (1,3,0) + (1,3,1) + (1,3,-1)$$
  
(24)

or

(8,1,0) + (1,3,0) + (1,1,0) + (1,1,1) + (1,1,-1). (25)

They should then be lying somewhere, between the GUT and the electroweak scales. We found that the addition of these particles with a unique common mass (as one intermediate scale) always destroys the GUT unification prediction.

In conclusion, we have presented an analysis of possible string unification without GUT for simple and motivated extensions of the minimal supersymmetric standard model which, in contrast to the MSSM, necessarily need a departure from the more attractive level-1 string constructions. We have focused on left-right and Pati-Salam symmetries, and our analysis shows that  $M_R$  can be as low as  $10^5$  GeV or so, and furthermore the value  $M_R \approx 10^{10} - 10^{12}$  GeV, interesting for neutrino physics, is perfectly consistent with unification constraints. In the case of two-step breaking, we can have  $M_R$  as  $M_I$  in the range  $10^{14} - 10^{11}$  GeV, while allowing the B-L gauge symmetry to remain unbroken all the way down to 1 TeV. Among other effects, this can save the proton from decaying too fast, which is a generic problem in supersymmetric theories.

For the Pati-Salam scale  $M_{\rm PS}$  we find that it has to be quite large, some two to three orders of magnitude below  $M_X$ . Only if we push  $M_X$  down to  $10^{16}$  GeV (not so appealing to us) can we keep  $M_{\rm PS}$  as low as  $10^{13}$  GeV to provide an interesting intermediate scale.

An optimistic point of view would be that these models can have interesting intermediate scales, allowing a correct string unification scale and charge quantization. A pessimistic point of view is that the predicted values of the associated Kac-Moody levels and the lack of knowledge of how to build such theories make it hopeless to construct these simple models in the near future, and they can only be studied from the effective field theory point of view, keeping in mind that gauge coupling unification could be achieved through the presence of only one tree level gauge coupling in heterotic string models. In any case further studies should include the possible effects of extra matter states which may exist in these theories.

### ACKNOWLEDGMENTS

We thank I. Antoniadis and C. Aulakh for useful comments.

- S. Dimopoulos, S. Raby, and F. Wilczek, Phys. Rev. D 24, 1681 (1981); N. Saki, Z. Phys. C 11, 153 (1981); L. E. Ibáñez and G. G. Ross, Phys. Lett. 105B, 439 (1981); M. B. Einhorn and D. R. T. Jones, Nucl. Phys. B196, 475 (1982); W. J. Marciano and G. Senjanović, Phys. Rev. D 25, 3092 (1982).
- [2] For LEP data analysis, see J. Ellis, S. Kelley, and D. V. Nanopoulos, Phys. Lett. B 249, 441 (1990); and 260, 131 (1991); U. Amaldi, W. de Boer, and H. Fürstenau, *ibid.* 260, 447 (1991); P. Langacker and M. Luo, Phys. Rev. D 44, 817 (1991). For more recent discussion, see P. Langacker and N. Polonsky, Phys. Rev. D 47, 4028 (1993); 49, 1454 (1994); L. J. Hall and U. Sarid, Phys. Rev. Lett. 70, 2673 (1993).
- [3] M. Gell-Mann, P. Ramond, and R. Slansky, in *Supergravity*, Proceedings of the Workshop, Stony Brook, New York, 1979, edited by P. van Niewenhuizen and D. Z. Freedman (North-Holland, Amsterdam, 1979); T. Yanagida, in *Proceedings of Workshop on Unified Theory and Baryon Number in the Universe*, Tsukuba, Japan, 1979, edited by O. Sawada and A. Sugamoto (KEK Report No. 79-18, Tsukuba, 1979).
- [4] R. N. Mohapatra and G. Senjanovíc, Phys. Rev. Lett. 44, 912 (1980); Phys. Rev. D 23, 165 (1981).
- [5] J. C. Pati and A. Salam, Phys. Rev. D 10, 275 (1974); R. N. Mohapatra and J. C. Pati, *ibid.* 11, 566 (1975); R. N. Mohap-

atra and G. Senjanović, ibid. 12, 1502 (1975).

- [6] P. Ginsparg, Phys. Lett. B 197, 139 (1987).
- [7] A. Schellekens, Phys. Lett. B 237, 363 (1990).
- [8] D. Lewellen, Nucl. Phys. B337, 61 (1990).
- [9] S. Chaudhuri, S.-W. Chung, and J. D. Lykken, Report No. Fermilab-Pub-94/137-T, hep-ph/9405374 (unpublished); S. Chaudhuri, S.-W. Chung, G. Hockney, and J. D. Lykken, in *The Albuquerque Meeting*, Proceedings of the Meeting of the Division of Particles and Fields of the APS, Albuquerque, New Mexico, 1994, edited by S. Seidel (World Scientific, Singapore, 1995), Report No. hep-th/9409151; B. Cleaver, *ibid.*, p. 1393.
- [10] G. Aldazabal, A. Font, L. E. Ibáñez, and A. M. Uranga, Nucl. Phys. B452, 3 (1995).
- [11] V. S. Kaplunovsky, Nucl. Phys. B307, 145 (1988); "Errata," Report No. STANFORD-ITP-838, 1992 (unpublished).
- [12] L. Dixon, V. Kaplunovsky, and J. Louis, Nucl. Phys. B355, 649 (1991); J.P. Derendinger, S. Ferrara, C. Kounnas, and F. Zwirner, *ibid.* 372, 145 (1992); I. Antoniadis, K. Narain, and T. Taylor, Phys. Lett. B 267, 37 (1991); I. Antoniadis, E. Gava, and K. Narain, *ibid.* 283, 209 (1992); Nucl. Phys. B383, 93 (1992); E. Kiritsis and C. Kounnas, *ibid.* 442, 472 (1995).
- [13] L. E. Ibáñez, Phys. Lett. B 318, 73 (1993).

- [14] I. Antoniadis, J. Ellis, S. Kelley, and D. V. Nanopoulos, Phys. Lett. B 272, 31 (1991).
- [15] I. Antoniadis and K. Benakli, Phys. Lett. B 295, 219 (1992).
- [16] For a recent work in this direction, see, for example, K. Dienes and A. Faraggi, Phys. Rev. Lett. **75**, 2646 (1995); Nucl. Phys. **B457**, 409 (1990); S. P. Martin and P. Ramond, Phys. Rev. D **51**, 6515 (1995).
- [17] A. Font, L. E. Ibánez, and F. Quevedo, Nucl. Phys. B345, 389 (1990).
- [18] I. Antoniadis, in 1990 Summer School in High Energy Physics and Cosmology, Proceedings, Trieste, Italy, edited by J. C. Pati et al. (World Scientific, Singapore, 1991), p. 677.

- [19] J. Ellis, J. L. Lopez, and D. V. Nanopoulos, Phys. Lett. B 245, 375 (1990); 247, 257 (1990).
- [20] R. N. Mohapatra, Phys. Rev. Lett. 56, 561 (1986).
- [21] R. N. Mohapatra, Phys. Rev. D 34, 3457 (1986); A. Font, L. E. Ibáñez, and F. Quevedo, Phys. Lett. B 228, 79 (1989); S. P. Martin, Phys. Rev. D 46, 2769 (1992).
- [22] M. Roncadelli and G. Senjanović, Phys. Lett. 107B, 59 (1981).
- [23] J. C. Pati and A. Salam, Phys. Rev. D 10, 275 (1974).
- [24] M. Cvetic, Phys. Lett. B 164, 55 (1985).
- [25] C. Aulakh, K. Benakli, and G. Senjanović (unpublished).
- [26] S. Chaudhuri, S.-W. Chung, G. Hockney, and J. D. Lykken, Nucl. Phys. B456, 89 (1995).