Variational approach to the spinless relativistic Coulomb problem

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By application of a straightforward variational procedure, me derive a simple, analytic upper bound on the ground-state energy eigenvalue of a semlrelativistic Hamiltonian for (one or two) spinless particles which experience some Coulomb-type interaction.

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I. INTRODUCTION

A standard (semi) relativistic description of bound states built up by particles of spin zero is provided by a Hamiltonian H which incorporates relativistic kinematics by involving the "square-root" operator of the relativistic kinetic energy, $\sqrt{p^2 + m^2}$ for a particle of mass m and momentum p, but describes the forces acting between the bound-state constituents by some coordinate-dependent static interaction potential $V(\mathbf{x})$.

For the case of bound states consisting of two particles of equal mass m , the generic Hamiltonian in the centerof-momentum frame of these constituents, expressed in terms of relative momentum \bf{p} and relative coordinate \bf{x} , reads

$$
H = 2\sqrt{p^2 + m^2} + V(\mathbf{x}).\tag{1}
$$

The equation of motion resulting from this type of Hamiltonian is the well-known "spinless Salpeter equation." It may be obtained from the so-called Salpeter equation [1] by ignoring all spin degrees of freedom and considering positive-energy solutions only. The Salpeter equation, in turn, approximates the Bethe—Salpeter equation [2] for bound states within a relativistic quantum field theory by eliminating, according to the spirit of an instantaneous interaction, any dependence on timelike variables. A central role in physics is played by the Coulomb potential, a spherically symmetric potential, i.e., one which depends only on the radial coordinate $r \equiv |\mathbf{x}|$, parametrized by some coupling strength κ :

$$
V(\mathbf{x}) = V_C(r) = -\frac{\kappa}{r}, \qquad \kappa > 0. \tag{2}
$$

The bound-state problem defined by the semirelativistic Hamiltonian (1) with the Coulomb potential (2) is what we call the "spinless relativistic Coulomb problem. "

Over the past years, this spinless relativistic Coulomb problem, in both the one- and two-particle formulations, has been the subject of intense study.¹ First of all, Herbst [3], in a rigorous mathematical discussion, developed the complete spectral theory of the one-particle counterpart of the operator (1) , (2) , from which one may directly deduce for the two-particle relativistic "Coulombic" Hamiltonian under consideration its essential self-adjointness for $\kappa \leq 1$ and the existence of its Friedrichs extension up to the critical value $\kappa_{cr} = 4/\pi$ of the coupling constant, and derived a strict lower bound on the energy E_0 of the ground state which translates in the two-particle case to

$$
E_0 \geq 2 m \sqrt{1-\left(\frac{\pi \kappa}{4}\right)^2} \quad \text{for} \quad \kappa < \frac{4}{\pi}.
$$
 (3)

Durand and Durand [4] obtained, in a certainly rather involved analysis of the spinless relativistic Coulomb problem for the particular case of vanishing relative orbital angular momentum of the bound-state constituents, as a by-product of the explicit construction of the corresponding wave function, a closed analytic expression for the exact energy eigenvalues. Castorina and co-workers [5] generalized, for small relative distances of the boundstate constituents, this wave function to arbitrary values of the orbital angular momentum. Hardekopf and Sucher [6] performed a comprehensive numerical analysis of oneand two-particle relativistic wave equations for both spin-0 as well as spin- $\frac{1}{2}$ particles, in the course of which they were able to show that the result reported in Ref. [4] for the energy eigenvalues must necessarily be wrong. Mar-

¹The semirelativistic Coulombic Hamiltonians for one- and two-particle problems, $H^{(1)} = \sqrt{P^2 + M^2} - \alpha/R$ and $H^{(2)} =$ $2\sqrt{p^2 + m^2} - \kappa/r$, respectively, may be easily equated. Relate their respective phase-space variables (X, P) and (x, p) by rescaling them by some scale factor λ according to $P = \lambda p$ and $X = x/\lambda$, which preserves their fundamental commutation relations: $[X, P] = [x, p]$. Let $\lambda = 2$ and identify both mass and Coulomb coupling strength parameters according to $M = 2m$ and $\alpha = \kappa/2$.

$$
E_0 \ge 2 m \sqrt{\frac{1 + \sqrt{1 - \kappa^2}}{2}} \quad \text{for} \quad \kappa < 1. \tag{4}
$$

Finally, Raynal and co-workers [8] succeeded in restricting numerically the ground-state energy eigenvalue of the semirelativistic Hamiltonian (1), (2), considered as a function of the coupling strength κ , to some remarkably narrow band.

The aim of the present paper is to find some simple and, in any case, analytically given upper bound on the ground-state energy level of the above semirelativistic "Coulombic" Hamiltonian.

II. A VARIATIONAL UPPER BOUND

In order to derive an analytic upper bound on the energy eigenvalue of the ground state of the spinless relativistic Coulomb problem, we make use of a rather standard variational technique. The basic idea of this variational technique is (i) to calculate the expectation values of the Hamiltonian H under consideration with respect to a suitably chosen set of trial states $|\lambda\rangle$ distinguished from each other by some variational parameter λ , which yields the λ -dependent expression $E(\lambda) \equiv \langle \lambda | H | \lambda \rangle$, and (ii) to minimize $E(\lambda)$ with respect to λ in order to obtain the upper bound to the proper energy eigenvalue E of the Hamiltonian H in the Hilbert-space subsector of the employed trial states $|\lambda\rangle$ as the above λ -dependent expression $E(\lambda)$ evaluated at the point of the minimizin
value λ_{\min} of the variational parameter: $E \le E(\lambda_{\min})$.

For the Coulomb potential, the most reasonable choice of trial states is obviously the one for which the coordinate-space representation $\psi(\mathbf{x})$ of the states $|\lambda\rangle$ for vanishing radial and orbital angular momentum quantum numbers is given by the hydrogenlike trial functions $(\lambda > 0)$

$$
\psi(\mathbf{x}) = \sqrt{\frac{\lambda^3}{\pi}} \, \exp(-\lambda \, r).
$$

For this particular set of trial functions we obtain for the expectation values we shall be interested in, namely, the ones of the square of the momentum p and of the inverse of the radial coordinate r , respectively, with respect to the trial states $|\lambda\rangle$,

$$
\langle \lambda \left| \mathbf{p}^2 \right| \lambda \rangle = \lambda^2
$$

and

$$
\left\langle \lambda \left| \frac{1}{r} \right| \lambda \right\rangle = \lambda.
$$

Let us follow this line of argument in some detail. As an immediate consequence of the fundamental postulates of any quantum theory, the expectation value of a given Hamiltonian H taken with respect to any normalized Hilbert-space state and, therefore, in particular, taken with respect to any of the above trial states, must necessarily be larger than or equal to that eigenvalue E_0 of the Hamiltonian H which corresponds to its ground state:

$$
E_0 \leq E(\lambda) \equiv \langle \lambda | H | \lambda \rangle.
$$

The application to the semirelativistic Hamiltonian of Eq. (1) yields, for the right-hand side of this inequality,

$$
E(\lambda) = 2 \left\langle \lambda \left| \sqrt{\mathbf{p}^2 + m^2} \right| \lambda \right\rangle + \langle \lambda | V(\mathbf{x}) | \lambda \rangle.
$$

Here, the rather cumbersome although (for convenient trial states) not impossible evaluation of the expectation value of the square-root operator may be circumvented very easily by taking advantage of a trivial but nevertheless fundamental inequality. This inequality relates the expectation values, taken with respect to arbitrary Hilbert-space vectors \mathcal{N} normalized to unity, of both the first and second powers of a self-adjoint but otherwise arbitrary operator $\mathcal{O} = \mathcal{O}^{\dagger}$; it reads

$$
|\langle \mathcal{O} \rangle| \leq \sqrt{\langle \mathcal{O}^2 \rangle}.
$$

For the purposes of the present discussion it is sufhcient to replace, in turn, $E(\lambda)$ by its upper bound obtained by applying this inequality:

$$
E(\lambda) \leq 2\sqrt{\langle\lambda\left|\mathbf{p}^2\right|\lambda\rangle + m^2} + \langle\lambda\vert V(\mathbf{x})\vert\lambda\rangle.
$$

Identifying in this, as far as its evaluation is concerned, simplified, upper bound the until-now general potential $V(\mathbf{x})$ with the Coulomb potential (2) and inserting both of the λ -dependent expectation values given above implies

$$
E(\lambda) \le 2\sqrt{\lambda^2 + m^2} - \kappa \lambda. \tag{5}
$$

From this intermediate result, by inspection of the limit $\lambda \to \infty$, we may state already at this very early stage that, for the semirelativistic Hamiltonian (1), (2) to be bounded from below at all, the Coulombic coupling strength κ has to stay below a certain critical value: $\kappa\leq 2.$

The value of the variational parameter λ which minimizes the upper bound on the right-hand side of Eq. (5) may be determined from the derivative of this expression with respect to λ :

$$
\lambda_{\min} = \frac{m\,\kappa}{2\sqrt{1-\frac{\kappa^2}{4}}}
$$

For this value of λ , by shuffling together all our previous inequalities, we find that the energy eigenvalue corresponding to the ground state of the semirelativistic Hamiltonian (1) with the Coulomb potential (2) , E_0 , is bounded from above by

$$
E_0\leq 2\,m\sqrt{1-\frac{\kappa^2}{4}}.\tag{6}
$$

The reality of this upper bound requires again $\kappa \leq 2$.

to

(1951).

316 (1977) (addendum).

III. SUMMARY

Within the framework of a simple variational technique we derived an analytic upper bound, Eq. (6), to the ground-state energy level of the spinless relativistic Coulomb problem. However, for any nonvanishing value

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of the Coulombic coupling strength κ , this upper bound (6) turns out, in fact, to be violated by the lowest energy eigenvalue obtainable from the analytic expression given in Ref. [4]. This provides, of course, further confirmation of the corresponding findings of Ref. [6]; a similar observation has been made in Ref. [9].

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