## Implications of precision electroweak experiments for $m_t$ , $\rho_0$ , $\sin^2\theta_W$ , and grand unification

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The implications of precision Z-pole, W-mass, and weak-neutral-current data for  $SU(2) \times U(1)$  models are described. Within the minimal model one finds  $\sin^2 \hat{\theta}_W(M_Z) = 0.2334 \pm 0.0008$  in the modified minimal subtraction scheme or  $\sin^2 \theta_W \equiv 1 - M_W^2 / M_Z^2 = 0.2291 \pm 0.0034$  in the on-shell scheme, where the uncertainties include the  $m_t$  and  $M_H$  dependence. The top-quark mass is predicted to be  $124^{+28+20}_{-34-15}$  GeV, where the second uncertainty is from  $M_H$ , with  $m_t < 174$  (182) GeV at 90 (95)% C.L. For the first time subleading effects and vertex corrections allow a significant separation of  $m_t$  and  $\rho_0$  in models with a nonminimal Higgs structure. Allowing arbitrary  $m_t$  and Higgs representations one obtains  $\sin^2 \hat{\theta}_W(M_Z) = 0.2333 \pm 0.0008$ ,  $\rho_0 = 0.992 \pm 0.011$ , and  $m_t < 294$  (310) GeV. The implications of these results for ordinary and supersymmetric grand unified theories are considered. Supersymmetric theories with a grand desert between the supersymmetry and unification scales are in striking agreement with data for  $M_{SUSY}$  in the  $M_Z$ -1 TeV range. Ordinary grand unified theories breaking to the standard model in more than one step are also discussed.

Weak-neutral-current data and W and Z properties have been a major quantitative test of the standard  $SU(2) \times U(1)$  electroweak model [1-7]. In 1987 a systematic analysis of the implications of all existing data was carried out [2], which has been updated regularly [5,7]. There is now a considerable amount of highprecision data on the mass, total and partial widths, and asymmetries of the Z from CERN LEP [8-10], as well as important new results on the W mass [11,12], atomic parity violation in cesium [13,14], and  $\binom{(-)}{v}_{\mu}e$  scattering [15]. There are also new direct 95%-C.L. lower limits  $m_t > 89$  GeV [16] and  $M_H > 48$  GeV [17] on the topquark and Higgs boson masses, and new experimental constraints on the charm-quark threshold [18] relevant to the interpretation of deep-inelastic neutrino scattering.

It is therefore an appropriate time to reconsider the implications of all these results for testing the standard model, constraining  $m_t$ , and comparing the experimental value of  $\sin^2\theta_W$  with ordinary and supersymmetric grand unified theories [19–23]. In this paper we will consider the standard SU(2)×U(1) model and extensions involving higher-dimensional representations of Higgs field [24]. Other extensions of the standard model will be considered elsewhere [25,26].

The data used in the analysis are summarized in Table I. The LEP results are averages of the four LEP experiments ALEPH, DELPHI, L3, and OPAL [10], which include all of the 1989-1990 data, with a proper treatment of common systematic errors.  $\Gamma_Z$ ,  $\Gamma_{l\bar{l}}$ ,  $\Gamma_{had}$ , and  $\Gamma_{inv}$ refer, respectively, to the total, leptonic (average of  $(e,\mu,\tau)$ , hadronic, and invisible Z widths;  $N_{\nu} \equiv \Gamma_{in\nu} / \Gamma_{\nu\bar{\nu}}$  is the number of light neutrino flavors;  $A_{\rm FB}(\mu)$  is the asymmetry forward-backward for muons;  $\sigma_p^h$ = $12\pi\Gamma_{e\bar{e}}\Gamma_{had}/M_Z^2\Gamma_Z^2$  is the hadronic cross section on the pole; and  $\bar{g}_A^2, \bar{g}_V^2$  are effective couplings related to  $\Gamma_{\bar{l}\bar{l}}$ and  $A_{\rm FB}(\mu)$  by

$$\Gamma_{I\bar{I}} = \frac{G_F M_Z^3}{6\sqrt{2}\pi} (\bar{g}_A^2 + \bar{g}_V^2) ,$$

$$A_{FB}(\mu) = \frac{3\bar{g}_V^2 \bar{g}_A^2}{(\bar{g}_V^2 + \bar{g}_A^2)^2} .$$
(1)

Only  $M_Z$ ,  $\Gamma_Z$ ,  $\Gamma_{l\bar{l}}$ ,  $R \equiv \Gamma_{had}/\Gamma_{l\bar{l}}$ , and  $A_{FB}(\mu)$  are used. The other LEP observables are not independent, but are displayed for completeness. Recent measurements of the W mass and weak-neutral-current data are also displayed in Table I. Older neutral-current results, included in the analysis, are described in [2,7]. The standard-model predictions for each quantity other than  $M_Z$  are also shown. These are computed using  $M_Z = 91.174 \pm 0.021$  GeV as input, using the range of  $m_t$  determined from the global fit, and 50 GeV  $< M_H < 1$  TeV.

In the standard model,

$$M_{Z}^{2} = \frac{A_{0}^{2}}{\hat{\rho}\hat{c}^{2}\hat{s}^{2}(1-\Delta\hat{r}_{W})} = \frac{A_{0}^{2}}{c^{2}s^{2}(1-\Delta r)} ,$$
  
$$M_{W}^{2} = \hat{\rho}\hat{c}^{2}M_{Z}^{2} = c^{2}M_{Z}^{2} ,$$
 (2)

where

$$A_0^2 = \pi \alpha / \sqrt{2} G_F = (37.2803 \text{ GeV})^2, \ \hat{s}^2 \equiv \sin^2 \hat{\theta}_W(M_Z)$$

refers to the weak angle in the modified minimalsubtraction scheme ( $\overline{MS}$ ) scheme [27],  $s^2 \equiv \sin^2 \theta_W$  $= 1 - M_W^2 / M_Z^2$  refers to the on-shell scheme [28],  $\hat{c}^2 \equiv 1 - \hat{s}^2$ , and  $c^2 \equiv 1 - s^2$ . The radiative correction parameters  $\Delta \hat{r}_W$ ,  $\hat{\rho} - 1$ , and  $\Delta r$  are taken from the calculation of Degrassi, Fanchiotti, and Sirlin [29]. As is well known,  $\hat{\rho} \sim 1 + \Delta \rho_t$ , where

$$\Delta \rho_t = \frac{3G_F m_t^2}{8\sqrt{2}\pi^2} \simeq 0.0031 \left[\frac{m_t}{100 \text{ GeV}}\right]^2$$
(3)

has a strong  $m_t$  dependence, while  $\Delta r \simeq \Delta r_0$ 

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TABLE I. Experimental values for LEP observables [10],  $M_W/M_Z$  [11],  $M_W$  [12], the weak charge in cesium  $Q_W$  [13,14], and the tree-level  $\sin^2\theta^0$  from  $\stackrel{(\nu)}{\nu}_{\mu}e \rightarrow \stackrel{(\nu)}{\nu}_{\mu}e$  [15], compared with the standardmodel predictions for  $M_Z = 91.174 \pm 0.021$  GeV,  $m_t = 124 + \frac{28}{34}$  GeV, and 50 GeV  $< M_H < 1$  TeV. Only the first five LEP observables are independent. The two errors for  $Q_W$  (Cs) are experimental and theoretical (in parentheses). The first uncertainty in the predictions is from the uncertainties in  $M_Z$  and  $\Delta r$ , the second is from  $m_t$  and  $M_H$ , and the third (in parentheses) is the theoretical QCD uncertainty. The older neutral-current quantities described in [2,7] are also used in the analysis.

Quantity	Value Standard model	
$M_Z$ (GeV)	91.174±0.021	
$\Gamma_Z$ (GeV)	$2.487{\pm}0.009$	$2.485{\pm}0.0021{\pm}0.008(\pm0.011)$
$\Gamma_{II}$ (MeV)	83.3±0.4	$83.5 {\pm} 0.1 {\pm} 0.2$
$R = \Gamma_{\rm had} / \Gamma_{l\bar{l}}$	20.94±0.12	$20.78{\pm}0.003{\pm}0.016({\pm}0.13)$
$A_{\rm FB}(\mu)$	$0.0154 {\pm} 0.0048$	$0.0142{\pm}0.0002{\pm}0.0015$
$\Gamma_{\rm had}$ (MeV)	1744±10	$1735 \pm 1.6 \pm 5.6 (\pm 11)$
$\Gamma_{inv}$ (MeV)	493±10	$499.4{\pm}0.3{\pm}1.3$
N <sub>v</sub>	$2.97{\pm}0.04{\pm}0.04$	3
$\sigma_p^b$ (nb)	$41.44{\pm}0.28$	$41.44{\pm}0.02{\pm}0.02({\pm}0.21)$
$\overline{g}_{A}^{2}$	$0.250 {\pm} 0.001$	$0.251 \pm 0 \pm 0.001$
$\overline{g}_V^2$	$0.0013 \pm 0.0004$	$0.0012 \pm 0 \pm 0.0001$
$M_W$ (GeV)	79.91±0.39	$80.05 {\pm} 0.03 {\pm} 0.19$
$M_W/M_Z$	$0.8831 {\pm} 0.0055$	$0.8780 {\pm} 0.0001 {\pm} 0.0022$
$Q_W$ (Cs)	$-71.04\pm1.58(\pm0.88)$	$-73.12{\pm}0.08{\pm}0.05$
$\sin^2 \theta^0$	$0.240{\pm}0.009{\pm}0.008$	$0.232{\pm}0.0003{\pm}0.001$

 $-\Delta \rho_t / \tan^2 \theta_W$  is even more sensitive.  $\Delta \hat{r}_W \sim \Delta r_0 \sim 0.07$  has no quadratic  $m_t$  dependence. There is additional logarithmic dependence on  $m_t$  and  $M_H$  in  $\hat{\rho}$ ,  $\Delta \hat{r}_W$ , and  $\Delta r$ , as well as  $O(\alpha)$  effects associated with low-energy physics. These effects are important and are fully incorporated in the analysis, but will not be displayed here.

For the Z widths, we use the results of Hollik [30], which include the  $m_t$  dependence of the  $Z \rightarrow b\bar{b}$  vertex corrections [31]. For the hadronic widths we include QCD corrections to  $O(\alpha_s^2)$  using an effective  $\alpha_s(M_Z)=0.12\pm0.02$ . The uncertainty is chosen larger than the average of the LEP determination of  $\alpha_s$  (e.g.,  $\alpha_s=0.123\pm0.007$  [32]) in order to roughly incorporate other QCD uncertainties associated with higher-order effects. For  $A_{\rm FB}(\mu)$  we use the calculation of Degrassi and Sirlin [33], which agrees with Hollik [30].

The most precise measurement of atomic parity violation is in the cesium atom [14]. Cesium is a hydrogenlike atom, allowing a clean theoretical interpretation of the results. Recent calculations have reduced the theoretical uncertainty to the 1% level [13]. The current value of  $Q_W$  differs from the standard-model prediction by slightly over one standard deviation.

Deep-inelastic neutrino reactions are still an important constraint. The ratio of neutral-to-charged-current cross sections has been measured at the 1% level [2-4]. The most important theoretical uncertainty is the charmquark threshold in the charged-current denominator. In [2] the threshold was parametrized [34] in terms of an effective charm-quark mass  $m_c^{\text{eff}}=1.5\pm0.3$  GeV determined from the experimental cross section for neutrinoinduced dimuon production [35]. The  $m_c^{\text{eff}}$  range is important because a much lower value (e.g., 0.8 GeV [36]) would considerably weaken the upper limit on  $m_t$ . Recently, the Chicago-Columbia-Fermilab-Rochester (CCFR) group has measured the threshold precisely, obtaining [18]  $m_c^{\text{eff}}=1.34\pm0.21(\text{stat})^{+0.25}_{-0.04}(\text{syst})$  GeV, in agreement with the old result based on data from the CERN-Dortmund-Heidelberg-Saclay (CDHS) Collaboration [37]. The lower central value is offset by the smaller lower error bar in obtaining upper limits on  $m_t$ .

The fits of various data sets to the standard model are displayed in Fig. 1 and Table II. Figure 1 shows the  $\pm 1\sigma$ limits on  $\sin^2 \hat{\theta}_W$  for fixed  $m_t$  as a function of  $m_t$  for various combinations of data. It is apparent that the combination of  $M_Z$  with  $M_W$ ,  $\Gamma$ ,  $A_{FB}(\mu)$ , and deep-inelastic



FIG. 1. Values of  $\sin^2 \hat{\theta}_W(M_Z)$  with  $\pm 1\sigma$  errors for fixed  $m_t$  as a function of  $m_t$  for  $M_H = 250$  GeV. Also shown is the 90%-C.L. region allowed in a fit to all data. The direct lower limit [16]  $m_t > 89$  GeV is also shown.

TABLE II. Values of  $\sin^2 \hat{\theta}_W(M_Z)$  and  $m_t$  obtained for various data sets. The  $\sin^2 \hat{\theta}_W(M_Z)$  error includes  $m_t$  and  $M_H$ . The first error for  $m_t$  includes experimental and theoretical uncertainties for  $M_H = 250$  GeV. The second error is the variation for  $M_H \rightarrow 50$  GeV (-) and  $M_H \rightarrow 1000$  GeV (+). The last column lists the upper limits on  $m_t$  at 90% (95%) C.L. for  $M_H = 1000$  GeV, which gives the weakest upper limit. The direct CDF constraint  $m_t > 89$  GeV is included.

Data	$\sin^2 \widehat{\theta}_W(M_Z)$	$m_t$ (GeV)	$m_t^{\max}$ (GeV
$M_{Z}, M_{W}$	$0.2329 {\pm} 0.0014$	139±47±16	213 (225)
$M_{Z},\Gamma$	$0.2340^{+0.0003}_{-0.0015}$	$91^{+55+34}_{-2-1}$	176 (188)
$M_Z, A_{\rm FB}(\mu)$	$0.2323 \pm 0.0023$	$158 \pm 68 \pm 23$	250 (266)
$M_Z, \Gamma, A_{\rm FB}(\mu)$	$0.2337^{+0.0006}_{-0.0012}$	$110^{+43}_{-21}\pm22$	179 (190)
$M_Z, M_W, \Gamma, A_{\rm FB}(\mu)$	$0.2334 \pm 0.0009$	$124 \pm 33 \pm 17$	179 (188)
$M_Z, \nu N$	$0.2339^{+0.0005}_{-0.0016}$	$99^{+58+23}_{-10-9}$	184 (197)
All	$0.2334 \pm 0.0008$	$124_{-34-15}^{+28+20}$	174 (182)

vN scattering places upper limits of  $\sim 200$  GeV on  $m_{t}$ . This is quantified in Table II, where the values of  $\sin^2 \hat{\theta}_W$ and  $m_t$  are shown for simultaneous fits to various combinations of data. The  $\sin^2 \hat{\theta}_W$  errors are almost entirely due to the  $m_t$  uncertainty [for fixed  $m_t$ , the Z mass would yield  $\sin^2 \hat{\theta}_W(M_Z)$  to  $\pm 0.0001$  ( $\pm 0.0003$ ), where the first error is from  $\Delta M_Z$  and the second is theoretical, from  $\Delta r$ ]. The central values are for a Higgs-boson mass  $M_H = 250$  GeV. The uncertainties induced by allowing 50 GeV  $< M_H < 1$  TeV are folded into the  $\sin^2 \hat{\theta}_W$  errors and are shown explicitly for  $m_t$ . Also listed are the 90%and 95%-C.L. upper limits on  $m_t$  from each data set. The latter are for  $M_H = 1$  TeV, which gives the weakest upper limit. It is seen that the LEP results alone, as well as LEP combined with  $M_W$ , yield 90%-C.L. upper limits ~180 GeV on  $m_t$ , confirming previous limits which were dominated by  $M_Z$  and vN scattering. The vN scattering, atomic parity violation (which does not contribute significantly to the  $m_t$  limit), and other neutral-current observables are still very important, however, because of their role in constraining new physics beyond the standard model [25].

The fit to all data yields

$$\begin{aligned} \sin^2 \hat{\theta}_W(M_Z) &= 0.2334 \pm 0.0008 ,\\ \sin^2 \theta_W &= 0.2291 \pm 0.0034 , \end{aligned} \tag{4}$$
$$m_t &= 124^{+28}_{-34-15} \text{ GeV }, \end{aligned}$$

where the  $\sin^2 \theta_W$  errors include  $m_t$  and  $M_H$  (the on-shell  $\sin^2 \theta_W$  is about 4 times as sensitive to  $m_t$ ), and the second error in  $m_t$  is from  $M_H$ . One can also regard  $\Delta \hat{r}_W$  (or  $\Delta r$ ) as free parameters. From a fit to all data in which  $\sin^2 \theta_W$ ,  $m_t$ , and  $\Delta r$  are free, one obtains  $\Delta r = 0.056^{+0.006}_{-0.010}$ ,  $\Delta \hat{r}_W = 0.061^{+0.010}_{-0.005}$ , in good agreement with standard-model expectations [29].

The upper limit on  $m_t$  from all data is (174,182,197) GeV at (90,95,99)% C.L., respectively, for  $M_H = 1$  TeV. For  $M_H = 250$  GeV, the corresponding limits are (157,165,180) GeV, while  $m_t < (144, 154, 170)$  GeV for  $M_H = 50$  GeV. The  $\chi^2$  distributions as functions of  $m_t$ are shown in Fig. 2(a) for the three values of  $M_H$ . It is apparent from Table II and Fig. 2 that the central value and upper limit on  $m_t$  are rather sensitive to  $M_H$ . However, until  $m_t$  is known independently, no significant constraint can be obtained on  $M_H$  from these types of experiments provided one assumes  $M_H \leq 1$  TeV. (It is not clear that it makes any sense to discuss the minimal standard model perturbatively for larger  $M_H > 1$  TeV.) Even if  $m_t$  were known, present data would only constrain  $M_H$  weakly: For example, for  $m_t = 100$  GeV, present data would favor  $M_H \sim 48$  GeV, but would still allow  $M_H \sim 810$  (1150) GeV at 90% (95%) C.L.



FIG. 2. (a)  $\chi^2$  distribution as a function of  $m_t$  for 134 df in a fit to all data for  $M_H = 50$ , 250, and 1000 GeV. The  $\chi^2$  is minimized with respect to  $\sin^2 \hat{\theta}_W(M_Z)$  for each  $m_t$ . (b) Same as (a), except  $\rho_0$  is also a free parameter (133 df).

Now that  $\sin^2 \hat{\theta}_W$  is well determined, it is useful to compare with the predictions of grand unification [19-22]. Using as inputs  $\alpha^{-1}(M_Z) = 127.9 \pm 0.2$  [29],  $\alpha_s(M_Z) = 0.12 \pm 0.012$  [8], and  $\sin^2 \hat{\theta}_W(M_Z) = 0.2334 \pm 0.0008$ , one obtains

$$\alpha_1^{-1} \equiv \frac{3}{5} \alpha^{-1} \cos^2 \hat{\theta}_W(M_Z) = 58.83 \pm 0.11 ,$$
  

$$\alpha_2^{-1} \equiv \alpha^{-1} \sin^2 \hat{\theta}_W(M_Z) = 29.85 \pm 0.11 ,$$
 (5)  

$$\alpha_3^{-1} \equiv \alpha_s^{-1} = 8.33 \pm 0.83 ,$$

at  $M_Z$ . These can be propagated to high renormalization scales using the standard two-loop renormalization-group equations [20]. Of course, in simple grand unified theories breaking in one step to the standard model (SM) or the minimal supersymmetric standard model (MSSM), one expects the three couplings to meet at the unification scale  $M_X$ . (One actually expects a calculable discontinuity in the MS scheme [20]. This is included in the analysis but is too small to see in the figures.)

The running couplings in the standard model are shown in Fig. 3(a). It is clearly seen that they do not meet at a point, thus ruling out simple grand unified theories such as SU(5), SO(10), or  $E_6$  which break in a single step to the standard model [19]. Of course, such models are also excluded by the nonobservation of proton decay, but this independent evidence is welcome.

On the other hand, in the minimal supersymmetric extension of the standard model, the couplings do meet within the experimental uncertainties. This is illustrated in Fig. 3(b) for the case in which all of the new particles have a common mass  $M_{SUSY} = M_Z$ . Almost identical curves are obtained for larger  $M_{SUSY}$ , such as 1 TeV. The unification scale  $M_X$  is sufficiently large (>10<sup>16</sup> GeV) that proton decay by dimension-6 operators is adequately suppressed, although there may still be a problem with dimension-5 operators [23]. This success is encouraging for supersymmetric grand unified theories such as SUSY-SU(5) or SUSY-SO(10). However, some caution



FIG. 3. (a) Running couplings in the standard model. (b) Running couplings in the MSSM with two Higgs doublets for  $M_{SUSY} = M_Z$ . The corresponding figure for  $M_{SUSY} = 1$  TeV is almost identical.

is in order, because such models have many uncertainties, such as those involving possible splitting of the superheavy-Higgs-boson masses from the unification scale [19,22].

One can also use  $\alpha(M_Z)$  and  $\alpha_s(M_Z)$  to predict  $\sin^2 \hat{\theta}_W(M_Z)$ . The resulting prediction, the scale  $(M_X)$  at which  $(5\alpha_1^{-1}+3\alpha_2^{-1})/8$  and  $\alpha_3^{-1}$  unify, and the value of  $\alpha_3^{-1}(M_X)$  are listed in Table III for a number of models. It is seen in Table III and Fig. 4 that ordinary grand unified theories such as SU(5) with one or two Higgs doublets predict  $\sin^2 \hat{\theta}_W$  considerably below the experimental

TABLE III. Predictions of  $\sin^2 \hat{\theta}_W$  and the unification scale in the non-SUSY and SUSY SU(5) at two-loop order, and for non-SUSY SO(10) grand unification at one loop. (The same predictions hold for many larger groups.) For SO(10) we assume SO(10)  $\rightarrow$  SU(3)  $\times$  SU(2)<sub>L</sub>  $\times$  SU(2)<sub>R</sub>  $\times$  U<sub>B-L</sub> at the unification scale  $M_X$ , while SU(3)  $\times$  SU(2)<sub>L</sub>  $\times$  SU(2)<sub>L</sub>  $\times$  U<sub>B-L</sub>  $\rightarrow$  SU(3)  $\times$  SU(2)<sub>L</sub>  $\times$  U<sub>Y</sub> at  $M_R$ . The inputs are  $\alpha^{-1} = 127.9 \pm 0.2$ ,  $\alpha_S = 0.12 \pm 0.012$ , and  $\sin^2 \hat{\theta}_W (M_Z) = 0.2334 \pm 0.0008$ .  $N_H$  is the number of Higgs doublets.  $\delta$  and  $\Delta$  refer to two classes of SU(2)<sub>L</sub>  $\times$  SU(2)<sub>R</sub>  $\times$  U<sub>R-L</sub> models.

Model	SUSY scale	$\sin^2 \hat{ heta}_W$	$M_X$	$\alpha_3^{-1}(M_X)$
$SU(5)(N_H = 1)$		$0.2102^{-0.0031}_{+0.0037}$	$4.5^{+3.4}_{-2.2} \times 10^{14}$	$41.2^{-0.1}_{+0.2}$
$SU(5)(N_{H}=2)$		$0.2142^{-0.0030}_{\pm 0.0036}$	$2.9^{+2.0}_{-1.4} \times 10^{14}$	$40.7_{\pm 0.2}^{+0.2}$
SUSY SU(5)( $N_H = 2$ )	$M_{Z}$	$0.2334_{\pm 0.0027}^{-0.0027}$	$1.9^{+1.7}_{-1.0} \times 10^{16}$	$25.2^{+0.2}_{+0.5}$
SUSY SU(5)( $N_H = 4$ )	$M_Z$	$0.2561^{-0.0022}_{\pm 0.0027}$	$9.2^{+7.0}_{-0.46} \times 10^{14}$	$23.9^{+0.4}_{+0.5}$
SUSY SU(5)( $N_H = 2$ )	$1 \ TeV$	$0.2315_{\pm 0.0032}^{-0.0026}$	$1.4^{+1.2}_{-0.7} \times 10^{16}$	$26.5_{\pm 0.5}^{+0.4}$
SUSY SU(5)( $N_H = 4$ )	1 TeV	$0.2525\substack{+0.0022\\+0.0027}$	$8.3^{+6.3}_{-4.2} \times 10^{14}$	$25.2^{+0.4}_{+0.5}$
	$M_R$	$\sin^2 \widehat{\theta}_W$	$M_{\chi}$	$\alpha_3^{-1}(M_X)$
$SO(10)(\delta_{I,R})$	1 TeV	$0.2811^{-0.0016}_{+0.0020}$	$2.0^{+2.1}_{-1.2} \times 10^{18}$	50.25-0.07
$SO(10)(\Delta_{L,R})$	1 TeV	$0.2765^{+0.0017}_{+0.0022}$	$5.6^{+6.6}_{-3.5} \times 10^{19}$	$53.97^{-0.18}_{+0.11}$
$SO(10)(\delta_{L,R})$	$4.5^{-2.0}_{+5.0} \times 10^{10}$	0.2334 (input)	$1.2^{+1.8}_{-0.8} \times 10^{16}$	$44.57_{\pm 0.31}^{-0.27}$
$SO(10)(\Delta_{L,R})$	$2.2^{-1.1}_{+3.2} \times 10^{10}$	0.2334 (input)	$5.3^{+10.0}_{-3.9} \times 10^{16}$	$46.22_{+0.56}^{+0.31}$

value. [These predictions actually hold for many grand unification groups larger than SU(5).] On the other hand, the predictions of the supersymmetric extension of SU(5) and similar groups predicts  $\sin^2 \hat{\theta}_W$  in agreement with observations for two Higgs doublets and  $M_{\rm SUSY}$  in the range  $M_Z - 1$  TeV. The predicted value is too high for four Higgs doublets.

We have seen that the observed low-energy couplings do not unify at a single scale if the running is due to the standard-model particles only, i.e., if there is a desert between  $M_Z$  and  $M_X$ . Adding supersymmetry in the desert is one way to achieve unification. Another possibility is to allow a group larger than SU(5) to break to the standard model in two or more stages. For example, an ordinary SO(10) model can break [38] first to left-right symmetric [39] SU(3)×SU(2)<sub>L</sub>×SU(2)<sub>R</sub>×U<sub>B-L</sub> at a scale  $M_X$  and then to the standard model at  $M_R$ .

The predictions of such models for a fixed  $M_R$  of 1 TeV are shown in Table III for two popular versions of the model: one in which  $SU(2)_R$  breaking is accomplished by introducing  $SU(2)_L$  and  $SU(2)_R$  doublets  $\delta_L$ and  $\delta_R$  [39], and one in which Higgs triplets [40]  $\Delta_{L,R}$ are introduced. The latter version can generate large Majorana masses for right-handed neutrinos, but are hard to incorporate into a superstring framework [24]. As is well known [41,19], the models generate much too high a  $\sin^2 \hat{\theta}_W$  for  $M_R \simeq 1$  TeV. However, one can obtain a viable model if  $M_R$  is left as a free parameter. In this case,  $\sin^2 \hat{\theta}_W$  is an input rather than a prediction. One obtains  $M_R \sim 10^{10} - 10^{11}$  GeV for the two versions, with a high enough  $M_X$  to avoid problems with proton decay. This scale of  $M_R$  is of relevance to seesaw models of neutrino mass that are suggested by the solar-neutrino problem [42].

A simple extension of the standard model is to allow higher-dimensional Higgs representations. It is convenient to define  $\rho_0 \equiv M_W^2 / (\hat{\rho} \hat{c}^2 M_Z^2)$ , which is unity in the standard model or in extensions involving Higgs doublets or singlets only. At the tree level,

$$\rho_0 = \frac{\sum_i (t_i^2 - t_{3i}^2 + t_i) |\langle \varphi_i \rangle|^2}{\sum_i 2t_{3i}^2 |\langle \varphi_i \rangle|^2} , \qquad (6)$$

where  $t_i$  and  $t_{3i}$  are the total and third component of the weak isospin of  $\varphi_i$ . One expects  $\rho_0 = 1$  in most superstring models [24], while  $\rho_0$  may differ from unity in models involving compositeness [43].  $\rho_0$  modifies the SM expressions for observables by  $M_Z \rightarrow M_Z^{\text{SM}} / \sqrt{\rho_0}$ ,  $\Gamma_Z \rightarrow \rho_0 \Gamma_Z^{\text{SM}}$ , and  $L_{\text{NC}} \rightarrow \rho_0 L_{\text{NC}}^{\text{SM}}$ , where  $L_{\text{NC}}$  is an effective four-Fermi neutral-current operator. If  $m_t$  were known independently,  $\rho_0$  would be determined very precisely. For example, current data yield  $\rho_0 = 1.001 \pm 0.002$  $(0.992 \pm 0.002)$  for  $m_t = 100$  (200) GeV. However,  $\rho_0$ affects the observables in the same way as the quadratic terms in  $m_t$  (except for the  $Z \rightarrow b\bar{b}$  vertex [31]), and so it is hard to separate the two. To first approximation one determines the combination

$$\rho_{\rm eff} = \rho_0 (1 + \Delta \rho_t) , \qquad (7)$$

which takes the value 1.005 $\pm$ 0.0024. However,  $\rho_0$  and  $m_t$  can be separated [44] by the  $\ln m_t$  terms in the radia-



FIG. 4. Predictions for  $\sin^2 \hat{\theta}_W(M_Z)$  vs  $m_t$  in ordinary and supersymmetric grand unified theories, compared with the regions allowed by the data at 90% C.L. for  $M_H = 50$ , 250, and 1000 GeV.

tive corrections to  $M_W$  and  $M_Z$  and by the  $Z \rightarrow b\bar{b}$  vertex (which most sensitively affects the total Z width at present).

For the first time the data are sufficiently accurate to allow a reasonably precise separation. A fit to all data yields

$$\sin^2 \theta_W(M_Z) = 0.2333 \pm 0.0008 ,$$
  

$$\rho_0 = 0.992 \pm 0.011 ,$$
  

$$m_t = 203^{+73}_{-113} \text{ GeV },$$
(8)

where the errors in  $\hat{s}^2$  and  $\rho_0$  include  $m_t$  and  $M_{H_2}$  and those in  $m_t$  include the  $M_H$  effect. That is,  $\sin^2 \hat{\theta}_W$  is determined just as well with  $\rho_0$  free as in the standard model, while  $\rho_0$  is determined to  $\simeq 1\%$  even with  $m_t$  free. One also obtains the upper limit  $m_t < 294$  (310) GeV at 90% (95%) C.L. with  $\rho_0$  free. The  $\chi^2$  distribution as a function of  $m_t$  is shown in Figs. 2(b). From (8) one obtains  $\rho_0 < 1.004$  and  $\rho_0 > 0.979$ , both at 90% C.L. Using the tree-level expression (6), this implies  $|\langle \varphi^0 \rangle| < 25 \text{ GeV}$ for the triplet  $(\varphi^{++}\varphi^{+}\varphi^{0})$ , and  $|\langle \chi^{0}\rangle| < 7.9$  GeV for  $(\chi^+ \chi^0 \chi^-)$ , to be compared with the doublet expectation value  $\langle \varphi_{1/2} \rangle \sim 246$  GeV. As long as  $\rho_0 \simeq 1$ , it is safe to neglect the effects of nondoublet vacuum expectations values on radiative corrections, which are of order  $\alpha(\rho_0-1)$ . However, new scalar or fermion SU(2) multiplets with large mass splittings can lead to loop effects that are identical in form to  $\rho_0$  [43]. In the presence of such multiplets, one can interpret the results on  $\rho_0$  as applying to  $\rho_0/(1-\alpha T)$ , where T is the isospin-breaking parameter defined in [26].

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