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⁷P. Carruthers, Phys. Rev. D 2, 2265 (1970).

⁸P. Carruthers, Phys. Rev. Letters 26, 1346 (1971).

⁹Here the 3×3 matrix \mathfrak{M} transforms as $(3, \bar{3})$ and is related to the Hermitian nonet fields σ_i, ϕ_i by

$$\mathfrak{M} = \sum_{j=0}^8 (\sigma_j + i\phi_j) \lambda_j / \sqrt{2}.$$

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T-Matrix Perturbation Theory for the Faddeev Equations

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A *T*-matrix perturbation theory for the Faddeev equations is developed in two different ways giving slightly differing results for the first-order correction to the ground-state energy of the three-body system, and both expressions differ from the result recently obtained by Fuda. The application of the theory to the system of three identical spinless particles is discussed.

I. INTRODUCTION

When solving the Faddeev equations,¹ a separable two-body *T* matrix is often used as input, since this reduces the problem (after angular momentum decomposition) to one of solving one-dimensional (possibly coupled) integral equations instead of the more general two-dimensional equations. In many applications the potentials one wants to use are not separable, and consequently the corresponding two-body *T* matrices will not be separable, and thus some separable approximation, T_s , to the actual *T* matrix must be used. The difference between these two, $T_p = T - T_s$, the perturbing part of the *T* matrix, may not be negligibly "small," and it is often useful to be able to calculate the first-order correction to the ground-state energy of the three-particle system.

Fuda² has written down such a perturbation expansion, but he uses a separable *T* matrix that comes from a separable potential, while we take advantage of the considerable freedom in the particular breakup of the *T* matrix into separable and nonseparable parts. Let $V = V_1 + V_2$ with V_1 separable. Then

$$T = T_1 + T_2 + T_{12},$$

where T_1 is the solution of the Lippmann-Schwinger (LS) equation for V_1 , T_2 is the solution for V_2 , and T_{12} contains both V_1 and V_2 . Since V_1 is separable, so is T_1 , and it is this that Fuda calls T_s . Thus, for Fuda, $T_p = T_2 + T_{12}$. However, $T_1 + T_{12}$ is also separable,³⁻⁵ and thus it is possible to identify T_2 as T_p . The later breakup is advantageous, since T_s includes some of the effects of

V_2 , and T_p may be "small" even when V_2 is not. Yaes⁵ has also developed a perturbation expression, but he emphasizes expansion in terms of the potential.

In Sec. II we derive two relations for the first-order correction to the energy, both of which differ from Fuda's, and in Sec. III we discuss how the formulas may be applied in order to find the first-order Coulomb correction to the ground-state energy for a system consisting of three identical, spinless, charged particles.

II. THEORY

We first write the homogeneous Faddeev equations in the form

$$(E - H_0)|\psi\rangle = T(E)|\psi\rangle, \quad (2.1)$$

where

$$|\psi\rangle = \begin{pmatrix} |\psi^1\rangle \\ |\psi^2\rangle \\ |\psi^3\rangle \end{pmatrix} \quad (2.2)$$

and

$$T = \begin{pmatrix} 0 & T_1 & T_1 \\ T_2 & 0 & T_2 \\ T_3 & T_3 & 0 \end{pmatrix}, \quad (2.3)$$

with T_i given by

$$T_i = V_i + V_i G_0 T_i. \quad (2.4)$$

In the above equation V_i is the potential between particles j and k [throughout this paper $(ijk) = (123)$ or some cyclic permutation thereof], and G_0 is the free-particle Green's function given by

$$G_0(Z) = \frac{1}{Z - H_0}, \quad (2.5)$$

with H_0 the kinetic-energy operator. Thus T_i is just the T matrix, in three-body space, for particles j and k . We note that Eq. (2.1) looks like a coupled-channel Schrödinger equation except that the "potential" $T(E)$ is energy-dependent.

We now write

$$T(E) = T_0(E) + \lambda T_p(E), \quad (2.6)$$

where we assume that the solution of the Faddeev equations for the T matrix T_0 is known; i.e., we know $|\psi_0\rangle$ and E_0 , where

$$(E_0 - H_0)|\psi_0\rangle = T_0(E_0)|\psi_0\rangle. \quad (2.7)$$

If we now expand $|\psi\rangle$, E , $T_0(E)$, and $T_p(E)$ in powers of λ about their unperturbed values, we obtain

$$|\psi\rangle = |\psi_0\rangle + \lambda |\psi_1\rangle + \lambda^2 |\psi_2\rangle + \dots, \quad (2.8)$$

$$E = E_0 + \lambda E_1 + \lambda^2 E_2 + \dots, \quad (2.9)$$

$$T_0(E) = T_0(E_0) + \lambda E_1 \left. \frac{\partial T_0(E)}{\partial E} \right|_{E=E_0} + \lambda^2 \left(E_2 \left. \frac{\partial T_0(E)}{\partial E} \right|_{E=E_0} + \frac{1}{2} E_1^2 \left. \frac{\partial^2 T_0(E)}{\partial E^2} \right|_{E=E_0} \right) + \dots, \quad (2.10)$$

with an analogous expansion for $T_p(E)$. Using these expansions in Eq. (2.1) and comparing coefficients of different powers of λ , we obtain

$$[E_0 - H_0 - T_0(E_0)]|\psi_0\rangle = 0, \quad (2.11)$$

$$[E_0 - H_0 - T_0(E_0)]|\psi_1\rangle + E_1 \left(1 - \left. \frac{\partial T_0}{\partial E} \right|_{E=E_0} \right) |\psi_0\rangle - T_p(E_0)|\psi_0\rangle = 0, \quad (2.12)$$

plus higher-order equations. We note that Eq. (2.11) is just the unperturbed Faddeev equation. After a little manipulation, Eq. (2.12) will give us E_1 , the first-order correction to the energy. We now define the "conjugate wave function," $\langle\bar{\psi}|$, by

$$\langle\bar{\psi}| = \langle\psi| M = (\langle\psi_2| + \langle\psi_3|, \langle\psi_1| + \langle\psi_3|, \langle\psi_1| + \langle\psi_2|), \quad (2.13)$$

where

$$M = \begin{pmatrix} 0 & 1 & 1 \\ 1 & 0 & 1 \\ 1 & 1 & 0 \end{pmatrix}. \quad (2.14)$$

If we now multiply Eq. (2.12) on the left by $\langle\bar{\psi}_0|$, we obtain

$$E_1 \left\langle \bar{\psi}_0 \left| 1 - \left. \frac{\partial T_0}{\partial E} \right|_{E=E_0} \right| \psi_0 \right\rangle - \langle \bar{\psi}_0 | T_p(E_0) | \psi_0 \rangle = 0, \quad (2.15)$$

since taking the Hermitian conjugate of Eq. (2.11) gives

$$\langle \bar{\psi}_0 | [E_0 - H_0 - T_0(E_0)] = 0. \quad (2.16)$$

Thus from Eq. (2.15) we have, for the first-order correction to the energy,

$$E_1 = \langle \bar{\psi}_0 | T_p(E_0) | \psi_0 \rangle / \left\langle \bar{\psi}_0 \left| 1 - \left. \frac{\partial T_0(E)}{\partial E} \right|_{E=E_0} \right| \psi_0 \right\rangle. \quad (2.17)$$

If $[\partial T_0(E)/\partial E]_{E=E_0}$ is small compared to 1, we reproduce Fuda's result.

We can derive another equation for the first-order correction to the energy from a variational principle.⁶ We define

$$J[\psi] = \langle \psi | L(E) | \psi \rangle, \quad (2.18)$$

where

$$L(E) = M[E - H_0 - T(E)], \quad (2.19)$$

with

$$T(E) = T_0(E) + T_p(E).$$

Then $\delta J = 0$ gives

$$L(E)|\psi\rangle = 0 \quad \text{and} \quad \langle \psi | L(E) = 0, \quad (2.20)$$

which are just the Faddeev equations.

An important property of variational principles is that they give second-order accuracy for some quantity when the trial function is good to first order. Let

$$|\psi\rangle = |\psi_0\rangle + |\delta\psi\rangle, \quad E = E_0 + \delta E, \quad (2.21)$$

where

$$(E_0 - H_0)|\psi_0\rangle = T_0(E_0)|\psi_0\rangle$$

and

$$(E - H_0)|\psi\rangle = [T_0(E) + T_p(E)]|\psi\rangle. \quad (2.22)$$

Then

$$\begin{aligned} J_0 &= \langle \psi_0 | L(E_0) | \psi_0 \rangle \\ &= (\langle \psi | - \langle \delta\psi |) L(E - \delta E) (|\psi\rangle - |\delta\psi\rangle) \\ &= -\delta E \left\langle \psi \left| \frac{\partial L}{\partial E} \right| \psi \right\rangle + (\text{higher-order term}) \\ &= -\delta E \left\langle \psi_0 \left| \frac{\partial L}{\partial E} \right|_{E=E_0} \right| \psi_0 \right\rangle + (\text{higher-order term}). \end{aligned}$$

Thus, since

$$J_0 = \langle \psi_0 | L(E_0) | \psi_0 \rangle = \langle \bar{\psi}_0 | T_p(E_0) | \psi_0 \rangle,$$

we have to second order

$$\delta E = \langle \bar{\psi}_0 | T_p(E_0) | \psi_0 \rangle / \left\langle \bar{\psi}_0 \left| 1 - \left. \frac{\partial T(E)}{\partial E} \right|_{E=E_0} \right| \psi_0 \right\rangle. \quad (2.23)$$

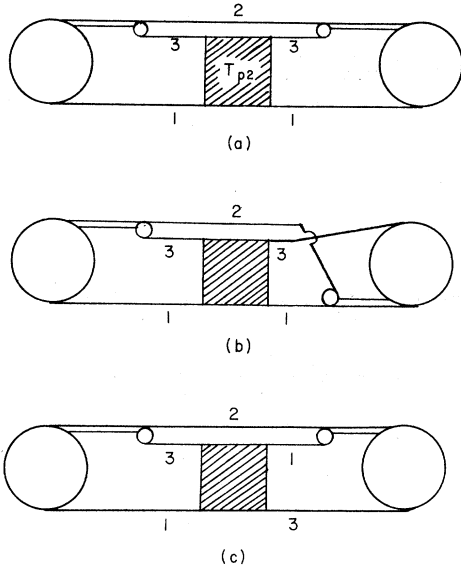


FIG. 1. (a) Graphical representation of $\langle \psi^1 | T_{p2} | \psi^1 \rangle$. (b), (c) Two different graphical representations of $\langle \psi^1 | T_{p2} | \psi^3 \rangle$.

We notice that (2.23) differs from (2.17) by the extra term $(\partial T_p / \partial E)_{E=E_0}$ in the denominator. The

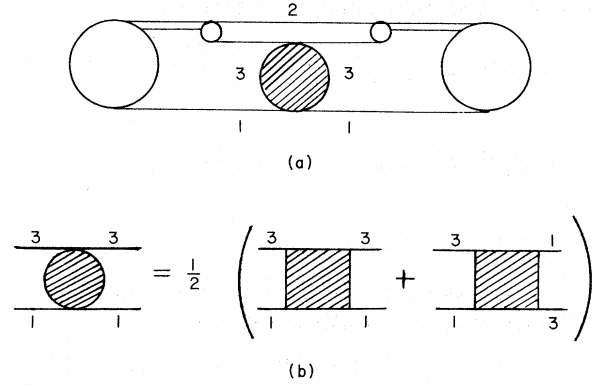


FIG. 2. (a) One half the sum of graphs (a) and (c) of Fig. 1. (b) Symmetrized form of the perturbing T matrix.

reason for this is that in the derivation of (2.17) it was assumed that T_p was "small" compared to T_0 , where no such condition is required in the derivation of (2.23). Thus, if T_p is small compared to T_0 , we expect E_1 given by Eq. (2.17) to be small compared to E_0 . On the other hand, we expect that Eq. (2.23) will be useful when δE is small compared to E_0 , whether or not T_p is small compared to T_0 .

III. APPLICATION

In this section we will develop Eq. (2.17) or Eq. (2.23) for the case of three identical spinless charged particles. For this special case we have

$$\langle \bar{\psi} | T | \psi \rangle = 6(\langle \psi^1 | T_2 | \psi^1 \rangle + \langle \psi^1 | T_2 | \psi^3 \rangle) \quad (3.1)$$

and

$$\langle \psi | \psi \rangle = 6\langle \psi^1 | \psi^3 \rangle. \quad (3.2)$$

Then Eq. (2.23) becomes

$$\delta E = \frac{\langle \psi^1 | T_{p2} | \psi^1 \rangle + \langle \psi^1 | T_{p2} | \psi^3 \rangle}{\langle \psi^1 | \psi^3 \rangle - \langle \psi^1 | (\partial T_2 / \partial E)_{E=E_0} | \psi^1 \rangle - \langle \psi^1 | (\partial T_2 / \partial E)_{E=E_0} | \psi^3 \rangle}. \quad (3.3)$$

We give in Fig. 1 a graphical representation of the two terms in the numerator of Eq. (3.3). From the second graphical representation of $\langle \psi^1 | T_{p2} | \psi^3 \rangle$, we see that

$$\frac{1}{2}(\langle \psi^1 | T_{p2} | \psi^1 \rangle + \langle \psi^1 | T_{p2} | \psi^3 \rangle)$$

can be represented as in Fig. 2. We can thus write Eq. (3.3) as

$$\delta E = \frac{\langle \psi^1 | T_{p2} | \psi^1 \rangle}{\frac{1}{2}\langle \psi^1 | \psi^3 \rangle - \langle \psi^1 | [\partial T_2(E) / \partial E]_{E=E_0} | \psi^1 \rangle}, \quad (3.4)$$

where we must now use the symmetrized T matrices for T_p and T_0 .

For three identical spinless particles we have

$$\langle \vec{p}_1 \vec{p}_2 \vec{p}_3 | T_1 | \vec{p}_1 \vec{p}_2 \vec{p}_3 \rangle = (2\pi)^3 \delta(\vec{k} - \vec{k}') (2\pi)^3 \delta(\vec{q}_i - \vec{q}_i') t_i(\vec{k}_i, \vec{k}_i', E - k^2/12\mu - \frac{3}{4}q_i^2/2\mu), \quad (3.5)$$

where $t_i(\vec{k}, \vec{k}', E)$ is the two-body T matrix for particles j and k (i.e., the solution to the two-body Lippmann-Schwinger equation), E is the two-body internal-energy variable, $\mu = m/2$ ($m =$ mass of each particle), and as usual

$$\vec{k} = \vec{p}_1 + \vec{p}_2 + \vec{p}_3, \quad \vec{k}_i = \frac{1}{2}(\vec{p}_j - \vec{p}_k), \quad \vec{q}_i = \frac{1}{3}(\vec{p}_j + \vec{p}_k - 2\vec{p}_i). \quad (3.6)$$

We also note that the ψ^i 's all have the same functional form ψ , so that the Faddeev equations reduce to a

single homogeneous integral equation,

$$\psi(\vec{k}, \vec{q}) = \frac{2}{E - k^2/2\mu - \frac{3}{4}q^2/2\mu} \int \frac{d^3q'}{(2\pi)^3} t_s(\vec{k}, \vec{q}' + \frac{1}{2}\vec{q}, E - \frac{3}{4}q^2/2\mu)\psi(\vec{q} + \frac{1}{2}\vec{q}', \vec{q}'), \quad (3.7)$$

where

$$\psi(\vec{p}_1, \vec{p}_2, \vec{p}_3) = \psi(\vec{k}_1, \vec{q}_1) + \psi(\vec{k}_2, \vec{q}_2) + \psi(\vec{k}_3, \vec{q}_3) \quad (3.8)$$

and

$$t_s(\vec{k}, \vec{q}) = \frac{1}{2}[t(\vec{k}, \vec{q}) + t(\vec{k}, -\vec{q})] \quad (3.9)$$

is just the symmetrized two-particle T matrix (we will suppress the subscript s in what follows).

Using the above relations, we have

$$\langle \psi^1 | T_2 | \psi^1 \rangle = \int d^3k_2 d^3q_2 d^3k'_2 \psi(-\frac{1}{2}\vec{k}_2 - \frac{3}{4}\vec{q}_2, \vec{k}_2 - \frac{1}{2}\vec{q}_2) t_2(\vec{k}_2, \vec{k}'_2, E_0 - \frac{3}{4}q_2^2/2\mu) \psi(-\frac{1}{2}\vec{k}_2 - \frac{3}{4}\vec{q}_2, \vec{k}'_2 - \frac{1}{2}\vec{q}_2). \quad (3.10)$$

If we let

$$\vec{x} = \frac{1}{2}\vec{q}_2, \quad \vec{y} = \vec{k}_2 - \vec{k}'_2, \quad \vec{z} = \vec{k}_2 + \vec{k}'_2, \quad (3.11)$$

and use the fact that

$$\psi(\vec{k}, \vec{q}) = \psi(-\vec{k}, \vec{q}), \quad (3.12)$$

we get

$$\begin{aligned} \langle \psi^1 | T_2 | \psi^1 \rangle = \int d^3x d^3y d^3z \psi(\frac{3}{2}\vec{x} + \frac{1}{4}\vec{y} + \frac{1}{4}\vec{z}, -\vec{x} + \frac{1}{2}\vec{y} + \frac{1}{2}\vec{z}) \\ \times t_p(\frac{1}{2}\vec{y} + \frac{1}{2}\vec{z}, -\frac{1}{2}\vec{y} + \frac{1}{2}\vec{z}, E_0 - \frac{3}{2}x^2/\mu) \psi(\frac{3}{2}\vec{x} - \frac{1}{4}\vec{y} - \frac{1}{4}\vec{z}, -\vec{x} - \frac{1}{2}\vec{y} + \frac{1}{2}\vec{z}), \end{aligned} \quad (3.13)$$

with an analogous expansion for $\langle \psi^1 | (\partial T_2 / \partial E)_{E=E_0} | \psi^1 \rangle$. Similarly we have

$$\langle \psi^1 | \psi^3 \rangle = (2\pi)^3 \int d^3y d^3z \psi(\vec{z} + \frac{1}{2}\vec{y}, \vec{y}) \psi(\vec{y} + \frac{1}{2}\vec{z}, \vec{z}). \quad (3.14)$$

The nine-dimensional integral can be reduced to six dimensions by noting that the integrand depends only on x , y , z , $\vec{x} \cdot \vec{y}$, $\vec{x} \cdot \vec{z}$, and $\vec{y} \cdot \vec{z}$. Similarly the six-dimensional integral can be reduced to three dimensions as the integrand depends only on y , z , and $\vec{y} \cdot \vec{z}$. Note that we are not restricting the perturbing part of the T matrix to act in only the s wave, and thus the integral will necessarily be more cumbersome than Fuda's.

We now let $t_p = t_c$, the Coulomb T matrix. For t_c and $\partial t_c / \partial E$ we could use an integral representation of the Coulomb T matrix⁷ given by

$$t_c(\vec{k}_1, \vec{k}_2, E) = 4\pi V_0 [1 + I(x)] / (\vec{k}_2 - \vec{k}_1)^2, \quad (3.15)$$

where

$$I(x) = -4a \int_0^1 \frac{t^a dt}{(1+t)^2 - x(1-t)^2}, \quad (3.16)$$

$$a = \frac{\mu V_0}{(-2\mu E)^{1/2}}, \quad x = 1 - \frac{(k_2^2 - 2\mu E)(k_1^2 - 2\mu E)}{-2\mu E(\vec{k}_2 - \vec{k}_1)^2}.$$

For identical particles of charge q , $V_0 = q^2$, and the Coulomb T matrix has no poles as a function of E . There are, however, singularities at $\vec{k}_1 = \vec{k}_2$ and at $E = k_1^2/2\mu$ and $E = k_2^2/2\mu$. The latter singularities do not occur in Eq. (3.13), since $E = E_0 - \frac{3}{2}x^2/\mu$ is always below threshold and the singularities at $\vec{k}_1 = \vec{k}_2$ are smoothed over in the integration.⁸

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