# The $\delta$ expansion for stochastic quantization 

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#### Abstract

Using a recently proposed perturbation expansion called the $\delta$ expansion, we show how to solve the Langevin equation associated with a $g \phi^{4}$ field theory. We illustrate the technique in zero- and one-dimensional space-time, and then generalize this approach to $d$ dimensions.


## I. INTRODUCTION

Recently, a new perturbative technique, called the $\delta$ expansion, was proposed to solve nonlinear problems in physics. ${ }^{1}$ The technique involves replacing in a differential equation nonlinear terms such as $\phi^{3}$ by $\phi^{1+2 \delta}$ and expanding this term in powers of $\delta$ :

$$
\begin{equation*}
\phi^{1+2 \delta}=\phi \sum_{n=0}^{\infty} \frac{\delta^{n}}{n!}\left(\ln \phi^{2}\right)^{n} . \tag{1.1}
\end{equation*}
$$

We then obtain a solution to the differential equation as a perturbation series in powers of $\delta$. The perturbation parameter $\delta$ is a measure of the nonlinearity of the theory. When $\delta=0$ the theory is linear and typically can be solved in closed form. As $\delta$ increases from zero, the effects of the nonlinearity turn on smoothly. Thus, one would expect and we have indeed found in our research that the $\delta$-series representation of the solution has a finite radius of convergence. Furthermore, the $\delta$ expansion is nonperturbative in all physical parameters such as the coupling constant $g$.

As an example of a difficult nonlinear problem that we have successfully treated, consider the Blasius equation

$$
\begin{equation*}
y^{\prime \prime \prime}+y y^{\prime \prime}=0, \quad y(0)=y^{\prime}(0)=0, \quad y^{\prime}(\infty)=1 \tag{1.2}
\end{equation*}
$$

This problem cannot be solved analytically. However, we can introduce the perturbation parameter $\delta$ :

$$
\begin{equation*}
y^{\prime \prime \prime}+y^{\delta} y^{\prime \prime}=0 \tag{1.3}
\end{equation*}
$$

If we represent the solution to this equation as a series in powers of $\delta$,

$$
\begin{equation*}
y=\sum_{n=0}^{\infty} y_{n} \delta^{n}, \tag{1.4}
\end{equation*}
$$

we can easily calculate the coefficients $y_{n}$. Even a small number of terms in the $\delta$ series gives an accurate approximation ${ }^{1}$ to the exact solution to (1.2).

Our success in solving classical nonlinear differential equations using the $\delta$ expansion suggests that one could
apply these same methods to the Langevin equation, a nonlinear classical differential equation whose solution can be used to obtain the Green's functions of a quantum field theory. The Langevin equation for a quantum field theory is obtained by adding two terms to the classical equation of motion, a random source term $\eta$, and a derivative with respect to a fictitious time $\tau$. For example, suppose we want to solve a $g \phi^{4}$ field theory in $d$ dimensional space-time. The Euclidean action for this theory is given by

$$
\begin{equation*}
S[\phi]=\int\left[\frac{1}{2}\left(\partial_{\mu} \phi\right)^{2}+\frac{1}{2} m^{2} \phi^{2}+\frac{1}{4} g \phi^{4}\right] d^{d} x . \tag{1.5}
\end{equation*}
$$

We replace the classical field equation

$$
\begin{equation*}
\frac{\delta S}{\delta \phi}=\left(-\partial^{2}+m^{2}\right) \phi+g \phi^{3}=0 \tag{1.6}
\end{equation*}
$$

by the Langevin equation

$$
\begin{equation*}
\frac{\partial \phi}{\partial \tau}+\frac{\delta S}{\delta \phi}=\eta \tag{1.7a}
\end{equation*}
$$

or for $S[\phi]$ given by (1.5) we obtain

$$
\begin{equation*}
\frac{\partial}{\partial \tau} \phi(x, \tau)+\left(-\partial^{2}+m^{2}\right) \phi(x, \tau)+g \phi^{3}(x, \tau)=\eta(x, \tau) . \tag{1.7b}
\end{equation*}
$$

This diffusion equation must be solved subject to the initial condition

$$
\begin{equation*}
\phi\left(x, \tau_{0}\right)=0, \tag{1.7c}
\end{equation*}
$$

where $\tau_{0}$ is the time at which the source term $\eta$ first turns on. Thus, $\phi$ is regarded as quiescent before the source term begins to operate. The source term $\eta$ represents white noise. This means that there is no correlation between the noise at two different points in ( $x, \tau$ ) space:

$$
\begin{equation*}
\langle\eta(x, \sigma) \eta(y, \tau)\rangle=2 \delta(x-y) \delta(\sigma-\tau) . \tag{1.8}
\end{equation*}
$$

We also assume that $\langle\eta(x, \sigma)\rangle=0$.
We can express the correlation function of a product of white-noise sources $\eta$ in terms of the functional integral:

$$
\begin{equation*}
\left\langle\eta\left(x_{1}, \tau_{1}\right) \eta\left(x_{2}, \tau_{2}\right) \cdots \eta\left(x_{n}, \tau_{n}\right)\right\rangle=\frac{\int \mathcal{D} \eta \exp \left[-\frac{1}{4} \int d^{d} x \int_{\tau_{0}}^{\infty} d \tau \eta^{2}(x, \tau)\right] \eta\left(x_{1}, \tau_{1}\right) \eta\left(x_{2}, \tau_{2}\right) \cdots \eta\left(x_{n}, \tau_{n}\right)}{\int \mathscr{D} \eta \exp \left[-\frac{1}{4} \int d^{d} x \int_{\tau_{0}}^{\infty} d \tau \eta^{2}(x, \tau)\right]} \tag{1.9}
\end{equation*}
$$

Evaluating this functional integral, we find that if $n$ is odd, the correlation function vanishes, and, if $n=2 m$,

$$
\begin{align*}
& \left\langle\eta\left(x_{1}, \tau_{1}\right) \eta\left(x_{2}, \tau_{2}\right) \cdots \eta\left(x_{2 m}, \tau_{2 m}\right)\right\rangle \\
& \quad=2^{m}\left[\delta\left(x_{1}-x_{2}\right) \delta\left(\tau_{1}-\tau_{2}\right) \delta\left(x_{3}-x_{4}\right) \delta\left(\tau_{3}-\tau_{4}\right) \cdots \delta\left(x_{2 m-1}-x_{2 m}\right) \delta\left(\tau_{2 m-1}-\tau_{2 m}\right)+\text { permutations }\right] \tag{1.10}
\end{align*}
$$

In all, there are $(2 m-1)!$ terms on the right-hand side of (1.10). To obtain the $N$-point Green's functions $G_{N}\left(x_{1}\right.$, $\ldots, x_{N}$ ) for the quantum field theory given by (1.5), which are conventionally expressed as a path integral,
$G_{N}\left(x_{1}, \ldots, x_{N}\right)=\frac{\int \mathcal{D} \phi \exp (-S[\phi]) \phi\left(x_{1}\right) \cdots \phi\left(x_{N}\right)}{\int \mathcal{D} \phi \exp (-S[\phi])}$,
we first solve (1.7b) for $\phi(x, \tau)$. We regard the classical field $\phi$ as a functional of the noise source $\eta$. Second, we calculate the equal $-\tau$ stochastic average using (1.10) and compute the Green's functions from the prescription ${ }^{2}$
$G_{N}\left(x_{1}, \ldots, x_{N}\right)=\lim _{\tau \rightarrow \infty}\left\langle\phi\left(x_{1}, \tau\right) \phi\left(x_{2}, \tau\right) \cdots \phi\left(x_{N}, \tau\right)\right\rangle$.

For the special case of supersymmetric quantum mechanics, there is another, and even simpler, procedure for obtaining the Green's functions. ${ }^{3}$ The rules are as follows. We consider the zero-dimensional Langevin equation

$$
\begin{equation*}
\dot{x}+W(x)=\eta(\tau) \tag{1.13}
\end{equation*}
$$

We compute the corresponding Green's function

$$
\begin{equation*}
G_{2}(\tau, \sigma)=\langle x(\tau) x(\sigma)\rangle \tag{1.14}
\end{equation*}
$$

using (1.10). We then take the limit $\tau+\sigma \rightarrow \infty$, $|\tau-\sigma|=T$ fixed. The result agrees with the two-point function of supersymmetric quantum mechanics whose Euclidean Lagrangian is

$$
\begin{equation*}
L=\frac{1}{4} \dot{x}^{2}+\frac{1}{4} W^{2}(x)-\frac{1}{2} W^{\prime}(x) \tag{1.15}
\end{equation*}
$$

Generalizing this procedure to calculate the $2 n$-point Green's function is straightforward. The advantage of this technique is that it is only necessary to solve a Langevin equation in the single variable $\tau$. This result is surprising because quantum mechanics is a quantum field theory in one-dimensional space-time, and, therefore, the Langevin equation corresponding to such a theory would ordinarily require two variables, the fictitious time $\tau$ and the real time $t$. Supersymmetric quantum mechanics is remarkable because it obviates the necessity of introducing the time $t$. Supersymmetry allows the one variable $\tau$ to play the role of the fictitious time as well as the real time.

We organize this paper as follows. In Sec. II we obtain
the first two terms in the $\delta$-series solution of the zerodimensional Langevin equation. We show that the equal- $\tau$ correlation function at large $\tau$ corresponds to the two-point function of zero-dimensional field theory. In Sec. III we verify that the unequal- $\tau$ correlation function corresponds to the two-point function of supersymmetric quantum mechanics. In Sec. IV we show how to generalize our procedure to arbitrary dimensions.

## II. ZERO-DIMENSIONAL LANGEVIN EQUATION

Consider the massless quantum field theory in zerodimensional space-time defined by

$$
\begin{equation*}
S(\phi)=g \frac{\phi^{4}}{4} \tag{2.1}
\end{equation*}
$$

The vacuum-persistence amplitude for this field theory reduces to an ordinary Riemann integral:

$$
\begin{equation*}
Z[J]=\frac{\int_{-\infty}^{\infty} d \phi \exp \left(-\frac{1}{4} g \phi^{4}+J \phi\right)}{\int_{-\infty}^{\infty} d \phi \exp \left(-\frac{1}{4} g \phi^{4}\right)} \tag{2.2}
\end{equation*}
$$

To obtain the $\delta$ expansion we first replace 4 everywhere in (2.2) by $2+2 \delta$, and instead, we study

$$
\begin{equation*}
Z_{\delta}[J]=\frac{\int_{-\infty}^{\infty} d \phi \exp \left[-g\left(\phi^{2}\right)^{1+\delta} /(2+2 \delta)+J \phi\right]}{\int_{-\infty}^{\infty} d \phi \exp \left[-g\left(\phi^{2}\right)^{1+\delta} /(2+2 \delta)\right]} \tag{2.3}
\end{equation*}
$$

The formula for the two-point function in this theory can be obtained in closed form by evaluating the integral exactly:

$$
\begin{align*}
G_{2} & =\frac{\int_{-\infty}^{\infty} d \phi \exp \left[-g\left(\phi^{2}\right)^{1+\delta} /(2+2 \delta)\right] \phi^{2}}{\int_{-\infty}^{\infty} d \phi \exp \left[-g\left(\phi^{2}\right)^{1+\delta} /(2+2 \delta)\right]} \\
& =\left[\frac{2+2 \delta}{g}\right]^{1 /(1+\delta)} \frac{\Gamma(3 /(2+2 \delta))}{\Gamma(1 /(2+2 \delta))} \tag{2.4}
\end{align*}
$$

It is straightforward to Taylor expand (2.4) as a series in powers of $\delta$ :
$G_{2}=\frac{1}{g}\left\{1-\delta L+\delta^{2}\left[-1+L+\frac{1}{2} L^{2}+\psi^{\prime}\left(\frac{3}{2}\right)\right]+\cdots\right\}$,
where

$$
\begin{equation*}
L=\psi\left(\frac{3}{2}\right)+\ln \frac{2}{g} \tag{2.6}
\end{equation*}
$$

We will now show how to obtain this same expansion from the Langevin equation. The Langevin equation for the original $\phi^{4}$ theory is

$$
\begin{equation*}
\frac{\partial \phi}{\partial \tau}+g \phi^{3}=\eta \tag{2.7}
\end{equation*}
$$

We replace (2.7) by

$$
\begin{equation*}
\frac{\partial \phi}{\partial \tau}+g \phi^{1+2 \delta}=\eta \tag{2.8}
\end{equation*}
$$

Next, we assume that $\phi$ has the expansion

$$
\begin{equation*}
\phi=\phi_{0}+\delta \phi_{1}+\delta^{2} \phi_{2}+\cdots . \tag{2.9}
\end{equation*}
$$

Inserting (2.9) into (2.8), we obtain a sequence of linear equations:

$$
\begin{align*}
& \frac{\partial \phi_{0}}{\partial \tau}+g \phi_{0}=\eta  \tag{2.10a}\\
& \frac{\partial \phi_{1}}{\partial \tau}+g \phi_{1}=-g \phi_{0} \ln \phi_{0}^{2}  \tag{2.10b}\\
& \frac{\partial \phi_{2}}{\partial \tau}+g \phi_{2}=-2 g \phi_{1}-g \phi_{1} \ln \phi_{0}^{2}-\frac{g}{2} \phi_{0}\left(\ln \phi_{0}^{2}\right)^{2} \tag{2.10c}
\end{align*}
$$

$$
\begin{align*}
G_{2}(\sigma, \tau)=\langle\phi(\sigma) \phi(\tau)\rangle= & \left\langle\left[\phi_{0}(\sigma)+\delta \phi_{1}(\sigma)+\delta^{2} \phi_{2}(\sigma)+\cdots\right]\left[\phi_{0}(\tau)+\delta \phi_{1}(\tau)+\delta^{2} \phi_{2}(\tau)+\cdots\right]\right\rangle \\
= & \left\langle\phi_{0}(\sigma) \phi_{0}(\tau)\right\rangle+\delta\left[\left\langle\phi_{0}(\sigma) \phi_{1}(\tau)\right\rangle+\left\langle\phi_{1}(\sigma) \phi_{0}(\tau)\right\rangle\right] \\
& +\delta^{2}\left[\left\langle\phi_{0}(\sigma) \phi_{2}(\tau)\right\rangle+\left\langle\phi_{2}(\sigma) \phi_{0}(\tau)\right\rangle+\left\langle\phi_{1}(\sigma) \phi_{1}(\tau)\right\rangle\right]+\cdots \tag{2.13}
\end{align*}
$$

In this subsection we calculate the first term in this series:

$$
\begin{align*}
\left\langle\phi_{0}(\sigma) \phi_{0}(\tau)\right\rangle & =e^{-g(\sigma+\tau)} \int_{\tau_{0}}^{\tau} d t \int_{\tau_{0}}^{\sigma} d s e^{g(t+s)}\langle\eta(t) \eta(s)\rangle \\
& =\frac{e^{-g T}}{g}-\frac{1}{g} e^{g\left(2 \tau_{0}-\tau-\sigma\right)}, \tag{2.14}
\end{align*}
$$

where $T=|\tau-\sigma|$. In the limit as $\tau$ and $\sigma$ approach infinity, with the time difference $T$ held fixed, the result in (2.14) approaches

$$
\begin{equation*}
\left\langle\phi_{0}(\sigma) \phi_{0}(\tau)\right\rangle \rightarrow \frac{e^{-g T}}{g} \tag{2.15}
\end{equation*}
$$

This is the form of a free Green's function in onedimensional space-time, where $g$ plays the role of the mass. Note that, even though the theory described by (2.1) has no bare mass term, to zeroth order in $\delta$, a mass has been generated. We have already seen this effect in
previous studies of the $\delta$ expansion. ${ }^{4}$ It is surprising that this one-dimensional result arises from the zerodimensional Langevin equation. We return to this point in Sec. III. Here, we simply set $T=0$ and obtain the result

$$
\begin{equation*}
\left\langle\phi_{0}^{2}\right\rangle=\frac{1}{g}, \tag{2.16}
\end{equation*}
$$

which agrees with the first term in (2.5).

## B. First-order calculation

To calculate the contribution to the two-point Green's function to first order in $\delta$, we must evaluate the correlation functions $\left\langle\phi_{0}(\sigma) \phi_{1}(\tau)\right\rangle$ and $\left\langle\phi_{1}(\sigma) \phi_{0}(\tau)\right\rangle$. This calculation is nontrivial because the source $\eta$ appears in the argument of a logarithm:

$$
\begin{equation*}
\left\langle\phi_{0}(\sigma) \phi_{1}(\tau)\right\rangle+\left\langle\phi_{1}(\sigma) \phi_{0}(\tau)\right\rangle=-g e^{-g(\sigma+\tau)}\left\langle\int_{\tau_{0}}^{\sigma} d t e^{g t} \eta(t) \int_{\tau_{0}}^{\tau} d r \int_{\tau_{0}}^{r} d s e^{g s} \eta(s) \ln \left[e^{-g r} \int_{\tau_{0}}^{r} d u e^{g u} \eta(u)\right]^{2}\right\rangle+(\sigma \leftrightarrow \tau) \tag{2.17}
\end{equation*}
$$

We encountered the analogous problem in our treatment of field theory in Ref. 4, and our approach to this problem here is similar; to wit, we use the identity

$$
\begin{equation*}
\left.\frac{d}{d \alpha} x^{\alpha}\right|_{\alpha=0}=\ln x \tag{2.18}
\end{equation*}
$$

to replace the logarithm in (2.17) by a power $\alpha$. In the subsequent calculation we regard $\alpha$ as an arbitrary integer, which allows us to use the identity (1.10). Specifically we have

$$
\begin{align*}
& \left\langle\phi_{0}(\sigma) \phi_{1}(\tau)\right\rangle+\left\langle\phi_{1}(\sigma) \phi_{0}(\tau)\right\rangle \\
& \quad=-g \frac{d}{d \alpha} e^{-g(\sigma+\tau)} 2^{\alpha+1}(2 \alpha+1)!!\int_{\tau_{0}}^{\sigma} d t e^{g t} \int_{\tau_{0}}^{\tau} d r e^{-2 \alpha g r} \prod_{n=1}^{2 \alpha+1} \int_{\tau_{0}}^{r} d z_{n} e^{g z_{n}} \delta\left(t-z_{1}\right) \delta\left(z_{2}-z_{3}\right) \\
&  \tag{2.19}\\
& \quad \times\left.\cdots \delta\left(z_{2 \alpha}-z_{2 \alpha+1}\right)\right|_{\alpha=0}+(\sigma \leftrightarrow \tau)
\end{align*}
$$

where the factor $(2 \alpha+1)!$ ! occurs because all permutations of (1.10) contribute equally. Performing the trivial integrations over the delta functions by integrating on $z_{1}, z_{3}, \ldots, z_{2 \alpha+1}$, we obtain

$$
\begin{align*}
\left\langle\phi_{0}(\sigma) \phi_{1}(\tau)\right\rangle & +\left\langle\phi_{1}(\sigma) \phi_{0}(\tau)\right\rangle \\
& =-\left.g \frac{d}{d \alpha} 2^{\alpha+1}(2 \alpha+1)!!e^{-g(\sigma+\tau)} \int_{\tau_{0}}^{\sigma} d t e^{2 g t} \int_{\tau_{0}}^{\tau} d r e^{-2 \alpha g r} \theta(r-t) \prod_{m=1}^{\alpha} \int_{\tau_{0}}^{r} d z_{2 m} e^{2 g z_{2 m}}\right|_{\alpha=0}+(\sigma \leftrightarrow \tau) \tag{2.20}
\end{align*}
$$

The integrals over $z_{2 m}, m=1,2, \ldots, \alpha$, are elementary:
$\left\langle\phi_{0}(\sigma) \phi_{1}(\tau)\right\rangle+\left\langle\phi_{1}(\sigma) \phi_{0}(\tau)\right\rangle=-\left.g \frac{d}{d \alpha} \frac{2^{\alpha+1} \Gamma\left(\alpha+\frac{3}{2}\right)}{g^{\alpha} \Gamma\left(\frac{3}{2}\right)} e^{-g(\sigma+\tau)} \int_{\tau_{0}}^{\tau} d r\left(1-e^{2 g\left(\tau_{0}-r\right)}\right)^{\alpha} \int_{\tau_{0}}^{\sigma} d t e^{2 g t} \theta(r-t)\right|_{\alpha=0}+(\sigma \leftrightarrow \tau)$.

At this point, we set $\sigma=\tau$. (We return to the case where $\sigma \neq \tau$ in the next section.) Doing the $t$ integration we obtain

$$
\begin{equation*}
\left\langle\phi_{0}(\sigma) \phi_{1}(\sigma)\right\rangle+\left\langle\phi_{1}(\sigma) \phi_{0}(\sigma)\right\rangle=-\left.\frac{d}{d \alpha} \frac{2^{\alpha+1} \Gamma\left(\alpha+\frac{3}{2}\right)}{g^{\alpha} \Gamma\left(\frac{3}{2}\right)} e^{-2 g \sigma} \int_{\tau_{0}}^{\sigma} d r\left(e^{2 g r}-e^{2 g \tau_{0}}\right)\left(1-e^{2 g\left(\tau_{0}-r\right)}\right)^{\alpha}\right|_{\alpha=0} \tag{2.22}
\end{equation*}
$$

Now we take the limit $\sigma \rightarrow \infty$ :
$\left\langle\phi_{0}(\sigma) \phi_{1}(\sigma)\right\rangle+\left\langle\phi_{1}(\sigma) \phi_{0}(\sigma)\right\rangle$

$$
\begin{align*}
& \rightarrow-\left.\frac{1}{g \Gamma\left(\frac{3}{2}\right)} \frac{d}{d \alpha}\left[\left[\frac{2}{g}\right]^{\alpha} \Gamma\left(\alpha+\frac{3}{2}\right)\right]\right|_{\alpha=0} \\
& =-\frac{1}{g} L \tag{2.23}
\end{align*}
$$

This agrees precisely with the order- $\delta$ term in (2.5) and establishes the validity of our computational method.

It is clear now how to proceed to any order in $\delta$. In $n$th order we introduce $n$ exponential parameters, $\alpha_{1}, \ldots, \alpha_{n}$, and use the identity (2.18) to replace logarithms by powers. We then regard the parameters $\alpha_{k}$ as integers so that we can apply the identity in (1.10). Finally, using $\Gamma$ functions to analytically continue the combinatorial factors, we differentiate with respect to the parameters, and evaluate the resulting expressions at $\alpha_{k}=0$.

## III. SUPERSYMMETRIC QUANTUM MECHANICS

To obtain the two-point Green's function in supersymmetric quantum mechanics, it is sufficient to consider $G_{2}(\sigma, \tau)$ in (2.13) at $|\sigma-\tau|=T \neq 0$ in the limit $\sigma, \tau \rightarrow \infty$. We reevaluate the integral in (2.21) in this slightly more general case. We find that

$$
\left\langle\phi_{0}(\sigma) \phi_{1}(\tau)\right\rangle=\left\{\begin{array}{l}
-(L / 2 g) e^{-g T}, \quad \sigma>\tau,  \tag{3.1}\\
-(L / 2 g) e^{-2 T}(1+2 g T), \quad \sigma<\tau .
\end{array}\right.
$$

Combining this result with (2.15) we find, to first order in $\delta$, that

$$
\begin{equation*}
G_{2}(\sigma, \tau)=\frac{e^{-g T}}{g}[1-\delta L(1+g T)+\cdots] \tag{3.2}
\end{equation*}
$$

We can verify this result using the techniques described in Ref. 4. The supersymmetric quantummechanical theory corresponding to (1.13) is defined by the vacuum-persistence functional ${ }^{3}$

$$
\begin{equation*}
Z[J]=\frac{\int \mathcal{D} \phi \exp \left[-S[\phi]+\int J \phi d t\right]}{\int \mathcal{D} \phi \exp (-S[\phi])}, \tag{3.3}
\end{equation*}
$$

where

$$
\begin{equation*}
S[\phi]=\int d t\left[\frac{1}{4} \phi^{2}+\frac{1}{4} W(\phi)^{2}-\frac{1}{2} W^{\prime}(\phi)\right] \tag{3.4}
\end{equation*}
$$

The Langevin equation in (2.8) corresponds to the choice

$$
\begin{equation*}
W(\phi)=g \phi^{1+2 \delta} \tag{3.5}
\end{equation*}
$$

If we substitute (3.5) into (3.4) and keep terms of order $\delta$, we find the approximate action

$$
\begin{align*}
S[\phi]=\int d t & {\left[\frac{1}{4} \dot{\phi}^{2}+\frac{1}{4} g^{2} \phi^{2}+\frac{\delta}{4} g^{2} \phi^{2} \ln \phi^{4}\right.} \\
& \left.-\frac{\delta}{2} g \ln \phi^{2}+O\left(\delta^{2}\right)+\text { const }\right) \tag{3.6}
\end{align*}
$$

This is a nonpolynomial action. Following the procedure in Ref. 4, to this order we replace (3.6) by a provisional


FIG. 1. The Feynman rules for the provisional action $\widetilde{S}$ in (3.7).

Euclidean action $\widetilde{S}[\phi]$ having polynomial interaction terms:
$\widetilde{S}[\phi]=\int d t\left[\frac{1}{4} \dot{\phi}^{2}+\frac{1}{4} g^{2} \phi^{2}+\frac{\delta g^{2}}{4} \phi^{4 \alpha+2}-\frac{\delta g}{2} \phi^{2 \alpha}\right]$.
Note that we recover the theory described by $S$ from the theory described by $\widetilde{S}$ by taking one derivative with respect to the parameter $\alpha$ and setting $\alpha=0$.

- Again, we treat $\alpha$ as an integer and read off the Feyn-

$$
\left[-\frac{\delta g^{2}}{4}(4 \alpha+2)!\left(\frac{1}{g}\right)^{2 \alpha} \frac{1}{2^{2 \alpha}(2 \alpha)!}+\frac{\delta g}{2}(2 \alpha)!\left(\frac{1}{g}\right)^{\alpha-1}\right]
$$

In (3.8) we have included the symmetry numbers shown in Fig. 2 for each graph. We have also included the amplitude for the loops; each loop has the value $1 / g$. Evaluating the integral in (3.8), taking the derivative with respect to $\alpha$, and setting $\alpha=0$, we obtain

$$
-\frac{e^{-g T}}{g} \delta L(1+g T)
$$

which agrees exactly with the order $\delta$ contribution in (3.2). We have thus verified that the $\delta$-expansion techniques when applied to the purely classical Langevin equation give, simply and directly, the correct fieldtheoretic Green's functions.

## IV. THE LANGEVIN EQUATION IN $\boldsymbol{d}$-DIMENSIONAL FIELD THEORY

In higher dimensions we must include the d'Alembertian in the Langevin equation (1.7). Again, we utilize the $\delta$ expansion by replacing (1.7b) by

$$
\begin{equation*}
\frac{\partial \phi}{\partial \tau}+\left(-\partial^{2}+m^{2}\right) \phi+g \phi^{1+2 \delta}=\eta . \tag{4.1}
\end{equation*}
$$

order $\delta^{0}$ :

order $\delta^{1}$ :

symmetry number $=\frac{1}{2^{2 \alpha}(2 \alpha)!}$

symmetry number $=\frac{1}{2^{\alpha-1}(\alpha-1)!}$
$\alpha-1$ loops

FIG. 2. The Feynman graphs in coordinate space that contribute to $G_{2}(\sigma, \tau)$ to order $\delta$.
man rules for $\widetilde{S}[\phi]$. The free propagator is $e^{-g\left|t_{1}-t_{2}\right|} / g$ in coordinate space and $2\left(p^{2}+g^{2}\right)^{-1}$ in momentum space. There are two vertices, a $(4 \alpha+2)$-point vertex, whose amplitude is $-(4 \alpha+2)!\delta g^{2} / 4$, and a $2 \alpha$-point vertex, whose amplitude is $(2 \alpha)!\delta g / 2$. These rules are illustrated in Fig. 1.
The three graphs contributing to $G_{2}(\sigma, \tau)$ to order $\delta$ are shown in Fig. 2. To order $\delta^{0}$ we have $e^{-g|\tau-\sigma|} / g$, which agrees with the first term in (3.2). To order $\delta^{1}$ there are two terms corresponding to two graphs shown in Fig. 2:

As before, we expand in powers of $\delta$ and assume the form (2.9). Thus, in place of the system (2.10) we have

$$
\begin{align*}
\frac{\partial \phi_{0}}{\partial \tau}+\left(-\partial^{2}+m^{2}+g\right) \phi_{0}= & \eta  \tag{4.2a}\\
\frac{\partial \phi_{1}}{\partial \tau}+\left(-\partial^{2}+m^{2}+g\right) \phi_{1}= & -g \phi_{0} \ln \phi_{0}^{2}  \tag{4.2b}\\
\frac{\partial \phi_{2}}{\partial \tau}+\left(-\partial^{2}+m^{2}+g\right) \phi_{2}= & -2 g \phi_{1}-g \phi_{1} \ln \phi_{0}^{2} \\
& -\frac{g}{2} \phi_{0}\left(\ln \phi_{0}^{2}\right)^{2} \tag{4.2c}
\end{align*}
$$

and so on. To solve the first of these equations, (4.2a), we Fourier transform in all variables except in the artificial time variable $\tau$. The solution of the transformed equation is
$\widetilde{\phi}_{0}(k, \tau)=e^{-\left(k^{2}+m^{2}+g\right) \tau} \int_{\tau_{0}}^{\tau} d s e^{\left(k^{2}+m^{2}+g\right) s} \widetilde{\boldsymbol{\eta}}(k, s)$.
From (1.8) the transformed sources satisfy

$$
\begin{equation*}
\langle\widetilde{\eta}(k, \sigma) \widetilde{\eta}(p, \tau)\rangle=2(2 \pi)^{d} \delta(k+p) \delta(\sigma-\tau) . \tag{4.4}
\end{equation*}
$$

The zeroth-order two-point function is then obtained as
$\lim _{\sigma=\tau \rightarrow \infty}\left\langle\widetilde{\phi}_{0}(k, \sigma) \widetilde{\phi}_{0}(p, \tau)\right\rangle=\frac{1}{k^{2}+m^{2}+g}(2 \pi)^{d} \delta(p+k)$.

Again, we note a shift, to zeroth order in $\delta$, of the square of the bare mass by $g$.
To proceed further we use the Langevin Green's function

$$
\widetilde{D}(k, \tau, s)=e^{-\left(k^{2}+m^{2}+g\right)(\tau-s)} \theta(\tau-s)
$$

used in (4.3) to solve (4.2b) for $\phi_{1}$ in terms of $\phi_{0}$. As usual we introduce a parameter $\alpha$ and use the identity (2.18) in order to replace the logarithm in that equation by a power of $\phi_{0}$. This allows us to compute the average over the noise using (1.10). We are left with integrals over the Langevin Green's function $\widetilde{D}$. These integrals will diverge unless we introduce a regularization scheme. Thus, following Ref. 5, we modify the Langevin equation (1.7a) to read

$$
\begin{equation*}
\frac{\partial \phi}{\partial \tau}(x, \tau)+\frac{\delta S}{\delta \phi}(x, \tau)=\int d^{d} y R_{x y}\left(\partial^{2}\right) \eta(y, \tau) \tag{4.6}
\end{equation*}
$$

where, for example, the regulator $R$ may be taken to be

$$
\begin{equation*}
R=\left[1-\frac{\partial^{2}}{\Lambda^{2}}\right]^{-j} \tag{4.7}
\end{equation*}
$$

where $j$ is chosen large enough to make all integrals that occur finite. Introducing such a regulator modifies the zeroth-order two-point function (4.5):

$$
\begin{equation*}
G_{2}^{(0)}(k, p)=\frac{\widetilde{R}^{2}\left(k^{2}\right)}{k^{2}+m^{2}+g}(2 \pi)^{4} \delta(p+k) . \tag{4.8}
\end{equation*}
$$

We will not pursue this calculation any further because it is not the purpose of this paper to conduct a regulated Langevin-equation calculation. Rather, our purpose here was to establish that the $\delta$ expansion, which we have applied successfully to a wide variety of classical differential equations, is equally effective in solving the Langevin equation corresponding to a quantum field theory.

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${ }^{1}$ C. M. Bender, K. A. Milton, S. S. Pinsky, and L. M. Simmons, Jr., J. Math. Phys. (to be published).
${ }^{2}$ G. Parisi and Wu Yong-Shi, Sci. Sin. 24, 483 (1981).
${ }^{3}$ F. Cooper and B. Freedman, Ann. Phys. (N.Y.) 146, 262 (1983); G. Parisi and N. Sourlas, Nucl. Phys. B206, 321 (1982); E.

Gozzi, Phys. Rev. D 28, 1922 (1983).
${ }^{4}$ C. M. Bender, M. Moshe, K. A. Milton, S. S. Pinsky, and L. M. Simmons, Jr., Phys. Rev. Lett. 58, 2615 (1987).
${ }^{5}$ Z. Bern, M. B. Halpern, L. Sadun, and C. Taubes, Nucl. Phys. B284, 1 (1987).

