

**Energy-momentum tensor in theories with scalar fields and two coupling constants.
III. A model with two scalars**

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In this work, we continue our investigation of the improvement term needed for the energy-momentum tensor in field theories with scalar fields and two coupling constants. Here, we shall discuss a model having two scalar fields. As in earlier works, we shall consider improvement terms which are (i) a finite function of bare quantities, (ii) a finite function of renormalized quantities. We show that, unlike the case of $\lambda\phi^4$ theory with a single scalar field, in the present case neither form of improvement leads to a finite energy-momentum tensor.

I. INTRODUCTION

Energy-momentum tensors and their finiteness have received a good deal of attention on account of their relevance in physics.¹⁻¹² As is well known, finiteness of the energy-momentum tensor in theories with scalar fields is a nontrivial question on account of a need for an improvement term and has been studied in great detail by various authors.^{2-4,6-9} These investigations have been restricted to theories with scalar field(s) and a single coupling constant⁶⁻⁸

In Refs. 10-12 (henceforth referred to as I, II, III, respectively) we investigated the cases containing scalar field(s) and two coupling constants, viz., scalar electrodynamics, non-Abelian theories with scalar fields, and Yukawa theory. In this work we complete this investigation by considering a scalar field model with two fields and two coupling constants.

We have motivated this study in Ref. 11. The improvement term needed is generally of the form

$$\bar{g}(\partial_\mu\partial_\nu - \partial^2g_{\mu\nu})\phi^2.$$

As in earlier works, we shall consider two forms¹¹ for improvement terms: (i) one in which \bar{g} is a finite function of bare quantities (i.e., finite at $\epsilon=0$), (ii) one in which \bar{g} is a finite function of renormalized quantities. Unlike a $\lambda\phi^4$ theory (with one scalar field) but like the cases dealt with in I-III, neither form leads to a finite energy-momentum tensor. The physical significance of this result is explained in II. As the discussion in this case is very similar to that in I-III, we shall present our results with brevity.

Our conclusion is that in all (renormalizable) theories with scalar field(s) and two coupling constants neither kind of improvement term leads to a finite energy-momentum tensor, and that a new renormalization is needed to make $\theta_{\mu\nu}$ finite. This is unlike theories without scalar fields and also unlike $\lambda\phi^4$ theory.

II. PRELIMINARY

We consider a theory of two interacting scalar fields ϕ_1 and ϕ_2 . We would like to restrict ourselves to the case in which there are only two independent couplings. Hence

we shall assume that the couplings of quartic interactions ϕ_1^4 and ϕ_2^4 are the same. We shall also assume that their masses are equal: this simplifies the treatment without altering the final conclusion. We consider the Lagrange density

$$\mathcal{L} = \frac{1}{2}\partial_\mu\phi_1\partial^\mu\phi_1 - \frac{1}{2}m_0^2\phi_1^2 - \frac{\lambda_0}{4!}\phi_1^4 + \frac{1}{2}\partial_\mu\phi_2\partial^\mu\phi_2 - \frac{1}{2}m_0^2\phi_2^2 - \frac{\lambda_0}{4!}\phi_2^4 - \frac{1}{4}\kappa_0\phi_1^2\phi_2^2. \quad (2.1)$$

We shall use dimensional regularization and the minimal-subtraction scheme throughout.

We define the renormalization transformations:

$$\phi_i = Z^{1/2}\phi_i^R, \quad i=1,2, \quad m_0^2 = Z_m m^2, \quad (2.2)$$

$$\lambda_0 = \mu^\epsilon[\lambda Z_\lambda + \delta\lambda(\kappa_0)], \quad \kappa_0 = \mu^\epsilon\kappa Z_\kappa.$$

Here λ_0 is not multiplicatively renormalizable and $\delta\lambda(\kappa_0)$ starts as κ_0^2 .

The canonical energy-momentum tensor is

$$\theta_{\mu\nu}^c = -g_{\mu\nu}\mathcal{L} + \partial_\mu\phi_1\partial_\nu\phi_1 + \partial_\mu\phi_2\partial_\nu\phi_2. \quad (2.3)$$

We define an improved energy-momentum tensor

$$\theta_{\mu\nu}^{\text{imp}} = \theta_{\mu\nu}^c + \left[\frac{n-2}{4(1-n)} + \frac{\bar{g}}{1-n} \right] (\partial_\mu\partial_\nu - \partial^2g_{\mu\nu}) \times (\phi_1^2 + \phi_2^2), \quad (2.4)$$

where the second term on the right-hand side arises when one considers the conformally invariant action in curved space-time and obtains $\theta_{\mu\nu}$ from it via

$$\theta_{\mu\nu} = -\frac{2}{\sqrt{-g}} \frac{\delta S}{\delta g^{\mu\nu}} \Big|_{g^{\mu\nu}=\eta^{\mu\nu}},$$

and \bar{g} is an additional improvement to be determined. The improvement term is symmetric under $\phi_1 \leftrightarrow \phi_2$ because so is the Lagrange density. A term of the kind $(\partial_\mu\partial_\nu - \partial^2g_{\mu\nu})(\phi_1\phi_2)$ is not an allowed improvement because \mathcal{L} is invariant under $\phi_1 \rightarrow -\phi_1$ (and also under $\phi_2 \rightarrow -\phi_2$ separately).

A simple calculation shows that

$$\begin{aligned} \theta_{\mu}^{\text{imp } \mu} = & (n-4) \left[-\frac{\lambda_0 \phi_1^4}{4!} - \frac{\lambda_0 \phi_2^4}{4!} - \frac{1}{4} \kappa_0 \phi_1^2 \phi_2^2 \right] \\ & + m_0^2 (\phi_1^2 + \phi_2^2) + \frac{n-2}{2} \left[\frac{\delta S}{\delta \phi_1} \phi_1 + \frac{\delta S}{\delta \phi_2} \phi_2 \right] \\ & + \bar{g} \partial^2 (\phi_1^2 + \phi_2^2). \end{aligned} \quad (2.5)$$

To obtain $\langle \theta_{\mu}^{\text{imp } \mu} \rangle$ one needs to consider renormalization of operators appearing in the above equation. To this end we shall consider the following set:

$$\begin{aligned} O_1 = & - \left[\frac{\lambda \phi_1^4}{4!} + \frac{\lambda \phi_2^4}{4!} + \frac{1}{4} \kappa_0 \phi_1^2 \phi_2^2 \right], \\ O_2 = & m_0^2 (\phi_1^2 + \phi_2^2), \quad O_3 = \phi_1 \frac{\delta S}{\delta \phi_1}, \quad O_4 = \phi_2 \frac{\delta S}{\delta \phi_2}, \quad (2.6) \\ O_5 = & -\frac{1}{4} \kappa_0 \phi_1^2 \phi_2^2, \quad O_6 = \partial^2 (\phi_1^2 + \phi_2^2). \end{aligned}$$

The above set is closed under renormalization. A slight generalization of the argument in Ref. 10 shows that O_2 is a finite operator so that

$$O_2^{\text{UR}} = O_2^R. \quad (2.7)$$

Equations of motion further imply that¹⁰ O_3 and O_4 are also finite operators:

$$O_3^{\text{UR}} = O_3^R, \quad O_4^{\text{UR}} = O_4^R. \quad (2.8)$$

O_6 is a multiplicatively renormalizable operator¹⁰ with

$$O_6^{\text{UR}} = Z_m^{-1} O_6^R, \quad (2.9)$$

where Z_m^{-1} is defined below.

We note the following:

$$\begin{aligned} O_1 = & \lambda_0 \frac{\partial S}{\partial \lambda_0} + \kappa_0 \frac{\partial S}{\partial \kappa_0}, \\ O_2 = & -2m_0^2 \frac{\partial S}{\partial m_0^2}, \\ O_5 = & \kappa_0 \frac{\partial S}{\partial \kappa_0}, \end{aligned} \quad (2.10)$$

and thus define renormalized operators¹⁰ in terms of the generating functional for connected renormalized Green's function $Z^R[J_i^R, \lambda, \kappa]$ as

$$\begin{aligned} O_1^R = & \lambda \frac{\partial Z^R}{\partial \lambda} + \kappa \frac{\partial Z^R}{\partial \kappa}, \\ O_2^R = & -2m^2 \frac{\partial Z^R}{\partial m^2}, \\ O_3 = & -J_1^R \frac{\delta Z^R}{\delta J_1^R}, \\ O_4 = & -J_2^R \frac{\delta Z^R}{\delta J_2^R}, \\ O_5 = & \kappa \frac{\partial Z^R}{\partial \kappa}. \end{aligned} \quad (2.11)$$

[The definition of O_1^R is valid only up to $O(\kappa^0)$. In this

definition, we have equated $\langle \lambda_0 (\partial S / \partial \lambda_0) \rangle^R = \lambda \partial Z^R / \partial \lambda$. But this cannot be true in $O(\kappa_0^2)$ because the right-hand side vanishes at $\lambda=0$ whereas the left-hand side does not vanish at $\lambda=0$; as then $\lambda_0 = \mu^\epsilon \delta \lambda(\kappa_0) \neq 0$ in $O(\kappa_0^2)$. However, this definition is valid in $O(\kappa_0)$ because $\delta \lambda_0$ vanishes to this order. Our treatment needs only $O(\kappa)$ quantities and hence the above definition of O_1^R is correct for our purpose.]

We introduce the following renormalization-group definitions and indicate the nontrivial leading term whenever they are needed in the future:

$$\begin{aligned} \gamma(\lambda, \kappa, \epsilon) = & \mu \frac{\partial}{\partial \mu} \ln Z \Big|_{\lambda_0, \kappa_0, \epsilon} = \gamma(\lambda, \kappa), \\ \gamma_m(\lambda, \kappa, \epsilon) = & -\frac{1}{2} \mu \frac{\partial}{\partial \mu} \ln Z_m = \gamma_m(\lambda, \kappa) \\ & = \gamma_{m(1)} \lambda + \gamma_{m(1)\kappa} + \dots, \end{aligned} \quad (2.12)$$

$$\begin{aligned} \beta^\lambda(\lambda, \kappa, \epsilon) = & \mu \frac{\partial \lambda}{\partial \mu} \\ & = \beta^\lambda(\lambda, \kappa) - \lambda \epsilon \\ & = \lambda^2 \frac{\partial}{\partial \lambda} Z_\lambda^{(1)} + \kappa \lambda \frac{\partial Z_\lambda^{(1)}}{\partial \kappa} + \kappa \frac{\partial}{\partial \kappa} \delta \lambda^{(1)} - \delta \lambda^{(1)} \\ & = -\lambda \epsilon + \beta_2 \lambda^2 + \dots, \\ \beta^\kappa(\lambda, \kappa, \epsilon) = & \mu \frac{\partial \kappa}{\partial \mu} = \beta^\kappa(\lambda, \kappa) - \kappa \epsilon \\ & = \kappa^2 \frac{\partial Z_\kappa^{(1)}}{\partial \kappa} + \lambda \kappa \frac{\partial Z_\kappa^{(1)}}{\partial \lambda} \\ & = -\kappa \epsilon + \frac{2}{16\pi^2} \lambda \kappa + \dots. \end{aligned}$$

For the six operators of Eq. (2.6), the renormalization matrix is defined by

$$O_i^{\text{UR}} = Z_{ij} O_j^R. \quad (2.13)$$

It can be shown that

$$Z_{ik}^{-1} \mu \frac{\partial}{\partial \mu} Z_{kj} = \gamma_{ij} = \text{finite at } \epsilon=0. \quad (2.14)$$

Equations (2.6)–(2.9) imply the following structure for Z_{ij} :

$$Z = \begin{pmatrix} Z_{11} & Z_{12} & Z_{13} & Z_{14} & Z_{15} & Z_{16} \\ 0 & 1 & 0 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1 & 0 & 0 \\ Z_{51} & Z_{52} & Z_{53} & Z_{54} & Z_{55} & Z_{56} \\ 0 & 0 & 0 & 0 & 0 & Z_m^{-1} \end{pmatrix}. \quad (2.15)$$

III. EXPRESSIONS FOR Z_{1j}

From Eqs. (2.5)–(2.9) the expression for the improved trace can be written as

$$\begin{aligned}
\langle \theta_{\mu}^{\text{imp } \mu} \rangle &= -\epsilon \langle O_1 \rangle^{\text{UR}} + \langle O_2 \rangle^R + \frac{n-2}{2} \langle O_3 + O_4 \rangle^R + \bar{g} Z_m^{-1} \langle O_G \rangle^R \\
&= -\epsilon \langle O_1 \rangle^{\text{UR}} + \bar{g} Z_m^{-1} \langle O_G \rangle^R + \text{finite} \\
&= -\epsilon \sum_{j=1}^5 Z_{1j} \langle O_j \rangle^R + (-\epsilon Z_{16} + \bar{g} Z_m^{-1}) \langle O_6 \rangle^R + \text{finite} .
\end{aligned} \tag{3.1}$$

In the future discussion, we shall need Z_{1j} ($j=1, 2, \dots, 5$). These are obtained by noting the identity

$$\int d^n x O_1 = -S + \frac{1}{2} \int d^n x (O_3 + O_4) . \tag{3.2}$$

An expression for $\langle S \rangle$ can be obtained along the lines of Ref. 10. Here we shall give the final result:

$$\begin{aligned}
\left\langle \int d^n x O_1 \right\rangle^{\text{UR}} &= -\langle S \rangle + \frac{1}{2} \left\langle \int d^n x O_3 + \int d^n x O_4 \right\rangle \\
&= \left[1 - \frac{\beta^\lambda}{\lambda \epsilon} \right] \left\langle \int d^n x O_1^R \right\rangle + \frac{\gamma_m}{\epsilon} \left\langle \int d^n x O_2 \right\rangle + \frac{\gamma}{\epsilon} \left\langle \int d^n x O_3 + \int d^n x O_4 \right\rangle^R + \left[\frac{\beta^\lambda}{\lambda \epsilon} - \frac{\beta^\kappa}{\kappa \epsilon} \right] \left\langle \int d^n x O_5 \right\rangle^R .
\end{aligned}$$

This yields

$$\begin{aligned}
Z_{11} &= 1 - \frac{\beta^\lambda}{\lambda \epsilon}, \quad Z_{12} = \frac{\gamma_m}{\epsilon}, \\
Z_{13} &= \frac{\gamma}{\epsilon} = Z_{14}, \quad Z_{15} = \frac{\beta^\lambda}{\lambda \epsilon} - \frac{\beta^\kappa}{\kappa \epsilon} .
\end{aligned} \tag{3.3}$$

As Z_{1j} ($j=1, \dots, 5$) have only simple poles in ϵ , it follows from Eq. (3.1) that at zero momentum (where O_6 does not contribute), $\langle \theta_{\mu}^{\text{imp } \mu} \rangle$ is finite.

IV. IMPROVEMENT-TERM DEPENDENCE OF THE FORM $\bar{g}(\epsilon, \kappa_0 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon})$

In this section, we shall consider an improvement coefficient $\bar{g}(\epsilon, \kappa_0 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon})$ which is a finite function of bare couplings κ_0 and λ_0 , at $\epsilon=0$. Using the renormalization-group equation satisfied by Z_{16} we shall show that such an improvement term cannot yield a finite energy-momentum tensor $\theta_{\mu\nu}^{\text{imp}}$ for any choice of $\bar{g}(\epsilon, \kappa_0 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon})$ even to first order in κ_0 .

We rewrite Eq. (3.1) (noting that Z_{1j} are finite for

$j=1, \dots, 5$) as

$$\langle \theta_{\mu}^{\text{imp } \mu} \rangle = -\epsilon X \langle O_6 \rangle^R + \text{finite} , \tag{4.1}$$

where

$$X = Z_{16} - \frac{\bar{g}}{\epsilon} Z_m^{-1} = Z_{16} + g Z_m^{-1} . \tag{4.2}$$

Here

$$g \equiv - \frac{\bar{g}(\epsilon, \kappa_0 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon})}{\epsilon}$$

can have $1/\epsilon$ terms when expanded in a series in powers of $(\kappa_0 \mu^{-\epsilon})$ and $(\lambda_0 \mu^{-\epsilon})$.

Z_{16} satisfies the differential equation (obtained along the lines of Appendix C of I)

$$\begin{aligned}
(-\lambda \epsilon + \beta^\lambda) \frac{\partial Z_{16}}{\partial \lambda} + (-\kappa \epsilon + \beta^\kappa) \frac{\partial Z_{16}}{\partial \kappa} - 2\gamma_m Z_{16} \\
= Z_{11} \gamma_{16} + Z_{15} \gamma_{56} .
\end{aligned} \tag{4.3}$$

Using the Eqs. (4.2) and (4.3), one can obtain an equation satisfied by X :

$$\begin{aligned}
(-\lambda \epsilon + \beta^\lambda) \frac{\partial X}{\partial \lambda} + (-\kappa \epsilon + \beta^\kappa) \frac{\partial X}{\partial \kappa} - 2\gamma_m X - Z_{11} \gamma_{16} - Z_{15} \gamma_{56} \\
= \left[\mu \frac{\partial}{\partial \mu} g(\epsilon, \kappa_0 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}) \right] Z_m^{-1} \\
= -\epsilon \sum_{n=0}^{\infty} (\kappa_0 \mu^{-\epsilon})^n \left[n g_n(\epsilon, \lambda_0 \mu^{-\epsilon}) + \frac{\partial g_n(\epsilon, \lambda_0 \mu^{-\epsilon})}{\partial (\lambda_0 \mu^{-\epsilon})} (\lambda_0 \mu^{-\epsilon}) \right] Z_m^{-1} ,
\end{aligned} \tag{4.4}$$

where we have expanded g in powers of $(\kappa_0 \mu^{-\epsilon})$:

$$g(\epsilon, \kappa_0 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}) = \sum_{n=0}^{\infty} (\kappa_0 \mu^{-\epsilon})^n g_n(\epsilon, \lambda_0 \mu^{-\epsilon}) . \tag{4.5}$$

Now, if it is possible to choose \bar{g} such that X has no worse than simple poles in ϵ , the left-hand side of Eq. (4.4) would have no worse than simple poles and hence

$$\epsilon^2 \sum_{n=0}^{\infty} (\kappa_0 \mu^{-\epsilon})^n \left[n g_n(\epsilon, \lambda_0 \mu^{-\epsilon}) + \frac{\partial g_n(\epsilon, \lambda_0 \mu^{-\epsilon})}{\partial (\lambda_0 \mu^{-\epsilon})} \lambda_0 \mu^{-\epsilon} \right] Z_m^{-1} = \text{finite} . \quad (4.6)$$

We have analyzed a similar equation in Sec. V of I and shown there that Eq. (4.6) in particular implies that

$$g_0(\epsilon, \lambda_0 \mu^{-\epsilon}) \equiv g_0(\epsilon), \quad g_1(\epsilon, \lambda_0 \mu^{-\epsilon}) = 0 \quad (4.7)$$

implying that the improvement coefficient obtained in $O(\kappa^0)$, viz., $g_0(\epsilon)$, should be sufficient to $O(\kappa)$ to make $\theta_{\mu\nu}^{\text{imp}}$ finite to $O(\kappa\lambda^n)$. But this contradicts the result in the Appendix that in $O(\kappa\lambda^3)$, $X = Z_{16} + g_0(\epsilon)Z_m^{-1}$ does have double poles as verified explicitly. Hence such an improvement term does not yield a finite $\theta_{\mu\nu}^{\text{imp}}$ even to $O(\kappa\lambda^n)$.

V. IMPROVEMENT-TERM DEPENDENCE OF THE FORM $\bar{g}(\epsilon, \kappa, \lambda)$

In this section we shall consider an improvement coefficient which is a finite function (at $\epsilon=0$) of the renormalized coupling constants and λ . As in the previous section,

$$\langle \theta_{\mu}^{\text{imp} \mu} \rangle = -\epsilon X \langle O_6 \rangle^R + \text{finite} , \quad (5.1)$$

where

$$X = Z_{16} - \frac{\bar{g}(\epsilon, \kappa, \lambda)}{\epsilon} Z_m^{-1} \equiv Z_{16} + g(\epsilon, \kappa, \lambda) Z_m^{-1} . \quad (5.2)$$

Using Eq. (4.3) we obtain a differential equation satisfied by X of Eq. (5.2), viz.,

$$(-\lambda\epsilon + \beta^\lambda) \frac{\partial X}{\partial \lambda} + (-\kappa\epsilon + \beta^\kappa) \frac{\partial X}{\partial \kappa} - 2\gamma_m X - Z_{11}\gamma_{16} - Z_{15}\gamma_{56} = \left[(-\lambda\epsilon + \beta^\lambda) \frac{\partial g}{\partial \lambda} + (-\kappa\epsilon + \beta^\kappa) \frac{\partial g}{\partial \kappa} \right] Z_m^{-1} . \quad (5.3)$$

As in the previous section, if it is possible to choose $\bar{g}(\epsilon, \kappa, \lambda)$ such that X has no worse than simple poles, then this implies

$$\epsilon \left[(-\lambda\epsilon + \beta^\lambda) \frac{\partial g}{\partial \lambda} + (-\kappa\epsilon + \beta^\kappa) \frac{\partial g}{\partial \kappa} \right] Z_m^{-1} = \text{finite at } \epsilon=0 . \quad (5.4)$$

We expand

$$g(\epsilon, \kappa, \lambda) = \sum_{n=0}^{\infty} \kappa^n g_n(\epsilon, \lambda) \quad (5.5)$$

and consider Eq. (5.4) in various powers of κ . Following the same procedure as in Sec. VI of II (β^λ and β^κ have identical leading terms as β^λ and β^ρ of II) we obtain

$$g_0(\epsilon, \lambda) = g_0(\epsilon), \quad g_1(\epsilon, \lambda) = 0 \quad (5.6)$$

implying that the improvement coefficient obtained in $O(\kappa^0)$, viz., $g_0(\epsilon)$ should be sufficient to $O(\kappa)$. This is of course wrong as explained at the end of the last section. Hence this kind of an improvement function does not yield a finite $\theta_{\mu\nu}^{\text{imp}}$ even to $O(\kappa\lambda^n)$.

APPENDIX

In this appendix we shall explicitly show that $X = Z_{16} + g_0(\epsilon)Z_m^{-1}$ does have double poles in ϵ in $O(\kappa\lambda^3)$. This we shall deduce from a calculation of simple pole divergence in $O(\kappa\lambda^2)$ in Z_{16} , use of renormalization-group equation for Z_{16} , viz., Eq. (4.3) and the knowledge of $O(\epsilon^2)$ terms in $g_0(\epsilon)$.

In $O(\kappa\lambda^3)$, the double-pole terms in X come from those in Z_{16} and the $O(\kappa\lambda^3/\epsilon^4)$ terms in Z_m^{-1} multiplied by $O(\epsilon^2)$ term in $g_0(\epsilon)$.

A direct calculation shows that the order- $\kappa\lambda^2$ terms in Z_{16} are given by

$$Z_{16} = -\frac{1}{18(16\pi^2)\epsilon} \kappa\lambda^2 + \text{other terms} . \quad (A1)$$

Using the RG equation satisfied by Z_{16} [Eq. (4.3)] one can relate the $O(\kappa\lambda^3/\epsilon^2)$ terms in Z_{16} to those of $O(\kappa\lambda^2/\epsilon)$. The result is

$$Z_{16} = -\frac{1}{18(16\pi^2)^2\epsilon^2} \kappa\lambda^3 + \text{other terms} . \quad (A2)$$

We further use

$$g_0(\epsilon) = g_{02}\epsilon^2 + O(\epsilon^3) = -\frac{1}{480}\epsilon^2 + O(\epsilon^3) \quad (A3)$$

and

$$Z_m^{-1} = -\frac{1}{16\pi^2\epsilon} \lambda + \text{other terms} \quad (A4)$$

and use the renormalization-group equation for Z_m^{-1} to obtain

$$Z_m^{-1} = -\frac{1}{(16\pi^2)^4} \lambda^3 \kappa + \text{other terms} . \quad (\text{A5})$$

Using the results of Eqs. (A2)–(A5) we obtain

$$\begin{aligned} X &= Z_{16} + g_0(\epsilon) Z_m^{-1} \\ &= \frac{1}{(16\pi^2)^4} \left(-\frac{1}{18} + \frac{1}{480} \right) \kappa \lambda^3 + \text{other terms} . \end{aligned}$$

Thus X does have double poles in ϵ in order $\kappa \lambda^3$, a result used in Secs. IV and V.

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