

Energy-momentum tensor in theories with scalar fields and two coupling constants. II. Yukawa theory

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We examine the question of renormalization of the energy-momentum tensor in Yukawa theory following our earlier work on scalar QED and non-Abelian gauge theories with scalars. As in those cases, we consider two kinds of forms for the improvement term: (i) one in which the improvement coefficient is a finite function of bare quantities of the theory (so that the energy-momentum tensor can be derived from an action that is a finite function of bare quantities); (ii) one in which the improvement coefficient is a finite quantity. As in earlier cases discussed we show that neither form leads to a finite energy-momentum tensor to $O(g^2\lambda^n)$.

I. INTRODUCTION

Energy-momentum tensors and their finiteness have received a good deal of attention on account of their relevance in physics.¹⁻¹¹ The finiteness of energy-momentum tensors in theories with scalar fields is a non-trivial question on account of a need for an improvement term and has been studied in great detail by various authors.^{2-4,6-9} Until recently the question had been studied in detail in the context of $\lambda\phi^4$ theory (with one scalar field). Collins^{6,7} has shown that an improvement term of the form

$$H_0(\epsilon)(\partial_\mu\partial_\nu - g_{\mu\nu}\partial^2)\phi^2 \quad (1.1)$$

[where $H_0(\epsilon)$ is a unique power series in *non-negative* powers of $\epsilon=4-n$] leads to a finite energy-momentum tensor to all orders.

In Refs. 10 and 11 (henceforth referred to as I and II) we discussed two special cases of theories containing scalar fields and having more than one coupling constant: viz., scalar QED and non-Abelian theories with scalar fields. As explained in detail in II, the crucial question in such theories is whether the improvement term can be chosen in such a way that the root-mean-square mass radius of the scalar particle is a *prediction* of the theory or whether this piece of information is needed as an independent experimental input to fix an independent renormalization constant of the $\frac{1}{2}R\phi^2$ term in the action.⁴ This depends on whether a finite energy-momentum tensor can be constructed so that the improvement coefficient is either a finite function of bare quantities, or a finite function of renormalized quantities (and hence a finite number), i.e., whether the "finite improvement program"⁴ works in such theories.

To this end, we have considered energy-momentum tensors in theories with scalar fields and having *two* coupling constants (for simplicity). There are four such renormalizable models: (i) scalar QED; (ii) non-Abelian gauge theories with scalars; (iii) Yukawa theory; (iv) a

model with two interacting scalar fields. The first two cases were analyzed in I and II. We analyze the third case in this paper. As in I and II we reach a negative conclusion for either kind of improvement coefficients in Yukawa theory also. The case (iv) is analyzed elsewhere with a similar conclusion.

In theories without scalar fields finite energy-momentum tensors which are finite functions of bare quantities exist.^{3,4} As shown by Collins^{6,7} the forms for the improvement coefficient of both kinds [the same one mentioned in Eq. (1.1)] work in $\lambda\phi^4$ theory. In pure $\lambda\phi^4$ theory, the root-mean-square radius of the scalar particle is (or rather can be) an experimental prediction of the theory and the coefficient of the $R\phi^2$ term in the action need not be independently renormalized.⁸ This proves to be an *exception* rather than a rule. In theories with scalar fields and more than one coupling constant the root-mean-square mass radius of the scalar field is needed as an additional experimental input to renormalize the coefficient of the $R\phi^2$ -like term in all the four cases analyzed.

II. PRELIMINARIES

We shall be dealing with the Yukawa theory of the scalar-fermion interaction. The Lagrangian density is

$$\begin{aligned} \mathcal{L} = & \frac{1}{2}(\partial_\mu\phi)^2 - \frac{1}{2}M_0^2\phi^2 - \frac{\lambda_0}{4!}\phi^4 + \bar{\psi}(i\partial - m_0)\psi \\ & + ig_0\bar{\psi}\gamma_5\psi\phi, \\ S = & \int d^4x \mathcal{L}. \end{aligned} \quad (2.1)$$

We shall work with dimensionally regularized quantities and use the minimal subtraction (MS) scheme.^{12,13} The unrenormalized but dimensionally regularized Green's functions, connected Green's functions, and proper vertices are generated, respectively, by $W[J, \eta, \bar{\eta}]$, $Z[J, \eta, \bar{\eta}]$, and $\Gamma[\phi, \psi, \bar{\psi}]$ with

$$\begin{aligned}
W[J, \eta, \bar{\eta}] &= \frac{1}{N} \int \mathcal{D}\phi \mathcal{D}\psi \mathcal{D}\bar{\psi} \\
&\times \exp \left[i \int d^n x \left[\frac{\mathcal{L}}{a} + J\phi + \bar{\psi}\eta + \bar{\eta}\psi \right] \right], \quad (2.2)
\end{aligned}$$

where $W[0]=1$ and a is the loop expansion parameter:

$$Z[J, \eta, \bar{\eta}] = -i \ln W[J, \eta, \bar{\eta}], \quad (2.3)$$

$$\phi(x) = \frac{\delta Z}{\delta J(x)}, \quad \psi(x) = \frac{\delta Z}{\delta \bar{\eta}(x)}, \quad \bar{\psi}(x) = -\frac{\delta Z}{\delta \eta(x)}, \quad (2.4)$$

and

$$\Gamma[\phi, \psi, \bar{\psi}] = Z - \int d^n x (J\phi + \bar{\eta}\psi + \bar{\psi}\eta). \quad (2.5)$$

In the MS scheme the renormalization transformations are¹⁴

$$\begin{aligned}
\phi &= Z^{1/2} \phi^R, \quad \psi = \bar{Z}^{1/2} \psi^R, \\
\lambda_0 &= \mu^\epsilon [\lambda Z_\lambda + \delta\lambda(g)], \quad \bar{\psi} = \bar{Z}^{1/2} \bar{\psi}^R, \\
g_0 &= \mu^{\epsilon/2} g Z_g, \quad m_0 = Z_m m, \quad M_0^2 = Z_M M^2 + Z'_M m^2,
\end{aligned} \quad (2.6)$$

where μ is an arbitrary parameter of dimension of mass. $\delta\lambda$ in Eq. (2.6) starts with $O(g^4)$ and Z'_M starts with $O(g^2)$.

The renormalized Green's functions, connected Green's functions and property vertices are generated, respectively, by

$$W^R[J^R, \eta^R, \bar{\eta}^R], \quad Z^R[J^R, \eta^R, \bar{\eta}^R], \quad \text{and} \quad \Gamma^R[\phi^R, \psi^R, \bar{\psi}^R]$$

with

$$W^R[J^R, \eta^R, \bar{\eta}^R] = W[J, \eta, \bar{\eta}] \quad \text{and} \quad J^R = Z^{1/2} J, \text{ etc.}$$

Equations of motion imply that

$$\left\langle \phi \frac{\delta S}{\delta \phi} \right\rangle^R = \left\langle \phi \frac{\delta S}{\delta \phi} \right\rangle = \text{finite}, \quad (2.7)$$

$$\left\langle \bar{\psi} \frac{\delta S}{\delta \bar{\psi}} \right\rangle^R = \left\langle \bar{\psi} \frac{\delta S}{\delta \bar{\psi}} \right\rangle = \text{finite}, \quad (2.8)$$

$$\left\langle \frac{\delta S}{\delta \psi} \psi \right\rangle^R = \left\langle \frac{\delta S}{\delta \psi} \psi \right\rangle = \text{finite}. \quad (2.9)$$

Unlike the cases considered in I and II $M_0^2 \phi^2$ now is not a finite operator as seen from Eq. (2.6), using the mass independence of Z_m, Z_M, Z'_M :

$$\langle M_0^2 \phi^2 \rangle = -2M_0^2 \frac{\partial Z^{\text{UR}}}{\partial M_0^2} = -2 \left[1 + \frac{m^2 Z'_M}{M^2 Z_M} \right] M^2 \frac{\partial Z^R}{\partial M^2}. \quad (2.10)$$

Similarly, $m_0 \bar{\psi}\psi$ is also not a finite operator:

$$\begin{aligned}
\langle m_0 \bar{\psi}\psi \rangle &= -m_0 \frac{\partial Z^{\text{UR}}}{\partial m_0} \\
&= -m \frac{\partial Z^R}{\partial m} + \left[\frac{2m^2 Z'_M}{M^2 Z_M} \right] M^2 \frac{\partial Z^R}{\partial M^2}.
\end{aligned} \quad (2.11)$$

But the sum $M_0^2 \phi^2 + m_0 \bar{\psi}\psi$ is finite:

$$\langle M_0^2 \phi^2 + m_0 \bar{\psi}\psi \rangle = -2M^2 \frac{\partial Z^R}{\partial M^2} - m \frac{\partial Z^R}{\partial m} = \text{finite}. \quad (2.12)$$

Furthermore $\partial^2 \phi^2$ is a multiplicatively renormalizable operator:¹⁵

$$\{\partial^2 \phi^2\}^{\text{UR}} = Z_M^{-1} \{\partial^2 \phi^2\}^R. \quad (2.13)$$

We shall use the renormalization group extensively. Below we give definitions and values of various renormalization-group quantities:¹⁶

$$\begin{aligned}
\beta^\lambda(\lambda, g, \epsilon) &\equiv \mu \frac{\partial \lambda}{\partial \mu} \Big|_{\lambda_0, g_0, M_0, m_0} \\
&= -\lambda \epsilon + \beta^\lambda(\lambda, g) \\
&= -\lambda \epsilon + \lambda^2 \frac{\partial Z_\lambda^{(1)}}{\partial \lambda} + \lambda \frac{\partial(\delta\lambda^{(1)})}{\partial \lambda} + \frac{g}{2} \frac{\partial(\delta\lambda^{(1)})}{\partial g} - \delta\lambda^{(1)} + \frac{1}{2} \lambda g \frac{\partial Z_\lambda^{(1)}}{\partial g} \\
&= -\lambda \epsilon + \beta_2 \lambda^2 + \dots, \\
\beta^g(\lambda, g, \epsilon) &\equiv \mu \frac{\partial g}{\partial \mu} \Big|_{\lambda_0, g_0, M_0, m_0} = -\frac{\epsilon g}{2} + \beta^g(\lambda, g) = -\frac{\epsilon g}{2} + \lambda g \frac{\partial Z_g^{(1)}}{\partial \lambda} + \frac{g^2}{2} \frac{\partial Z_g^{(1)}}{\partial g}, \\
\gamma_M(\lambda, g, \epsilon) &\equiv +\frac{1}{2} \mu \frac{\partial}{\partial \mu} \ln M^2 = \gamma_M(\lambda, g), \\
\gamma_m(\lambda, g, \epsilon) &\equiv +\frac{1}{2} \mu \frac{\partial}{\partial \mu} \ln m^2 = \gamma_m(\lambda, g), \\
\bar{\gamma}_M(\lambda, g, \epsilon) &= -\frac{1}{2} \mu \frac{\partial}{\partial \mu} \ln Z_M = \bar{\gamma}_M(\lambda, g) = \frac{\lambda}{2} \frac{\partial}{\partial \lambda} Z_M^{(1)} + \frac{g}{4} \frac{\partial}{\partial g} Z_M^{(1)} = \gamma_M^{(1)} \lambda + \dots, \\
\gamma(\lambda, g, \epsilon) &\equiv +\frac{1}{2} \mu \frac{\partial}{\partial \mu} \ln Z = \gamma(\lambda, g), \\
\bar{\gamma}(\lambda, g, \epsilon) &= +\frac{1}{2} \mu \frac{\partial}{\partial \mu} \ln \bar{Z} = \bar{\gamma}(\lambda, g).
\end{aligned} \quad (2.14)$$

We shall work with the following set of operators:

$$\begin{aligned} O_1 &= -\frac{\lambda_0 \phi^4}{4!} + \frac{i}{2} g_0 \bar{\psi} \gamma_5 \psi \phi, \quad O_2 = M_0^2 \phi^2, \\ O_3 &= m_0 \bar{\psi} \psi, \quad O_4 = \phi \frac{\delta S}{\delta \phi}, \quad O_5 = \bar{\psi} \frac{\delta S}{\delta \psi}, \\ O_6 &= \frac{\delta S}{\delta \psi} \psi, \quad O_7 = \frac{1}{2} i g_0 \bar{\psi} \gamma_5 \psi \phi, \quad O_8 = \partial^2 \phi^2. \end{aligned} \quad (2.15)$$

We note that, at zero momentum,

$$\begin{aligned} \int d^n x O_1 &= \lambda_0 \frac{\delta S}{\delta \lambda_0} + \frac{1}{2} g_0 \frac{\delta S}{\delta g_0}, \\ \int d^n x O_2 &= -2M_0^2 \frac{\delta S}{\delta M_0^2}, \\ \int d^n x O_3 &= -m_0 \frac{\delta S}{\delta m_0}, \\ \int d^n x O_7 &= \frac{1}{2} g_0 \frac{\delta S}{\delta g_0}. \end{aligned}$$

We shall define another set of renormalized operators:

$$\begin{aligned} \left\langle \int d^n x X_1^R \right\rangle^R &= \lambda \frac{\partial Z^R}{\partial \lambda} + \frac{1}{2} g \frac{\partial Z^R}{\partial g}, \\ \left\langle \int d^n x X_2^R \right\rangle^R &= -M^2 \frac{\partial Z^R}{\partial M^2}, \\ \left\langle \int d^n x X_3^R \right\rangle^R &= -m \frac{\partial Z^R}{\partial m}, \\ \left\langle \int d^n x X_4^R \right\rangle^R &= -\int J^R(x) \frac{\delta Z^R}{\delta J^R(x)} = \left\langle \int d^n x O_4^R \right\rangle, \\ \left\langle \int d^n x X_5^R \right\rangle^R &= -\int \bar{\eta}^R(x) \frac{\delta Z^R}{\delta \bar{\eta}^R(x)} = \left\langle \int d^n x O_5^R \right\rangle, \\ \left\langle \int d^n x X_6^R \right\rangle^R &= -\int \frac{\delta Z^R}{\delta \eta^R(x)} \eta^R(x) = \left\langle \int d^n x O_6^R \right\rangle, \\ \left\langle \int d^n x X_7^R \right\rangle^R &= \frac{1}{2} g \frac{\partial Z^R}{\partial g}, \quad X_8 = O_8. \end{aligned} \quad (2.16)$$

As discussed in I and is also evident from Eqs. (2.6), (2.10), and (2.11), $X_i^R \neq O_i^R$ for $i=1, 2$, and 3 , but one can still define a renormalization matrix Z_{ij} by

$$\langle O_i \rangle^{\text{UR}} = Z_{ij} \langle X_j \rangle^R. \quad (2.17)$$

(Note that $X_j = O_j$ for $j=4, 5, \dots, 8$.) Then

$$Z_{ij}^{-1} \mu \frac{\partial}{\partial \mu} Z_{jk} = \gamma_{ik} = \text{finite at } \epsilon=0. \quad (2.18)$$

Information on the structure of renormalization matrix

is obtained from Eqs. (2.7)–(2.11) and Eq. (2.13):

$$Z = \begin{pmatrix} Z_{11} & Z_{12} & Z_{13} & Z_{14} & Z_{15} & Z_{16} & Z_{17} & Z_{18} \\ 0 & Z_{22} & 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & Z_{32} & 1 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 1 & 0 & 0 \\ Z_{71} & Z_{72} & Z_{73} & Z_{74} & Z_{75} & Z_{76} & Z_{77} & Z_{78} \\ 0 & 0 & 0 & 0 & 0 & 0 & 0 & Z_{88} \end{pmatrix}. \quad (2.19)$$

It should be noted that Z_{22} and Z_{32} so defined depend on m/M .

We expand Z_M^{-1} in powers of g :

$$Z_M^{-1} = Z_{M(0)}^{-1} + g^2 Z_{M(2)}^{-1} + g^4 Z_{M(4)}^{-1} + \dots$$

We now state two results needed in Secs. V and VI. As shown by Collins,⁷ if $H(\lambda, M, \epsilon)$ is a finite function of λ, M at $\epsilon=0$ and

$$H(\lambda, M, \epsilon) Z_{M(0)}^{-1}(\lambda, \epsilon) = \text{finite at } \epsilon=0$$

keeping λ and M finite and fixed, then

$$H(\lambda, M, \epsilon) \equiv 0. \quad (2.20)$$

Second, as shown in Ref. 9, if

$$F \left[\lambda_0 \mu^{-\epsilon}, \frac{M_0^2}{\mu^2}, \epsilon \right]$$

is a finite function, at $\epsilon=0$, of $\lambda_0 \mu^{-\epsilon}$ and M_0^2/μ^2 and if

$$F \left[\lambda_0 \mu^{-\epsilon}, \frac{M_0^2}{\mu^2}, \epsilon \right] Z_{M(0)}^{-1}(\lambda, \epsilon) = \text{finite at } \epsilon=0$$

keeping λ and M fixed and finite, then

$$F \left[\lambda_0 \mu^{-\epsilon}, \frac{M_0^2}{\mu^2}, \epsilon \right] \equiv 0. \quad (2.21)$$

III. IMPROVED TRACE

The energy-momentum tensor of Yukawa theory as obtained from the action (2.1) is

$$\begin{aligned} \theta_{\mu\nu} &= -\frac{2}{\sqrt{-g(y)}} \frac{\delta S}{\delta g^{\mu\nu}} \Big|_{g^{\mu\nu(y)} = \eta^{\mu\nu}} \\ &= -g_{\mu\nu} \mathcal{L} + \partial_\mu \phi \partial_\nu \phi + i \bar{\psi} \gamma_\mu \partial_\nu \psi. \end{aligned} \quad (3.1)$$

One may carry out the analysis of Ref. 3 here also to show that this energy-momentum tensor has finite matrix elements at $q=0$ and to first order in q , the external momentum. But in $O(q^2)$ a further improvement term is needed. The most general improvement one can add to $\theta_{\mu\nu}$ is parametrized as

$$\theta_{\mu\nu}^{\text{imp}} = \theta_{\mu\nu} + \left[\frac{n-2}{4(1-n)} + \frac{\tilde{G}}{(1-n)} \right] (\partial_\mu \partial_\nu - g_{\mu\nu} \partial^2) \phi^2. \quad (3.2)$$

As before, \tilde{G} is a free parameter for which we shall be trying specific forms in Secs. V and VI, just as in I and II. The trace of $\theta_{\mu\nu}^{\text{imp}}$ is

$$\theta_{\mu}^{\text{imp}\mu} = (n-4) \left[-\frac{\lambda\phi^4}{4!} + \frac{1}{2}ig\bar{\psi}\gamma_5\psi\phi \right] + (M^2\phi^2 + m\bar{\psi}\psi) - \left[\frac{n-2}{2} \right] \phi \frac{\delta S}{\delta\phi} - (n-1)\bar{\psi} \frac{\delta S}{\delta\bar{\psi}} + \tilde{G}\partial^2\phi^2. \quad (3.3)$$

Using Eqs. (2.7), (2.8), and (2.13) one obtains

$$\langle \theta_{\mu}^{\text{imp}\mu} \rangle = \text{finite} + (n-4)\langle O_1 \rangle^{\text{UR}} + \tilde{G}Z_M^{-1}\langle \partial^2\phi^2 \rangle^R. \quad (3.4)$$

In the following sections we shall investigate if it is possi-

ble to choose \tilde{G} in either of the two ways mentioned in the Introduction so that $\langle \theta_{\mu}^{\text{imp}\mu} \rangle$ is finite.

IV. EXPRESSIONS FOR Z_{ij}

In this section we shall use the techniques of Ref. 10 to obtain expressions for Z_{1j} ($j=1,2,\dots,7$), which are needed in further investigations of Secs. V and VI. These expressions are obtained by considering renormalization of O_1 at zero momentum.

It is straightforward to show that

$$\int d^n x O_1 = -S + \frac{1}{2} \int \phi \frac{\delta S}{\delta\phi} d^n x + \int \bar{\psi} \frac{\delta S}{\delta\bar{\psi}} d^n x. \quad (4.1)$$

The last two terms on the right-hand side are finite operators. $\langle S \rangle^{\text{UR}}$ can be obtained by following the procedure of Refs. 10 and 17. The final result is

$$\begin{aligned} \langle \int d^n x O_1 \rangle^{\text{UR}} &= -\langle S \rangle^{\text{UR}} + \frac{1}{2} \left\langle \int d^n x \phi \frac{\delta S}{\delta\phi} \right\rangle^R + \left\langle \int d^n x \bar{\psi} \frac{\delta S}{\delta\bar{\psi}} \right\rangle^R \\ &= \left[1 - \frac{\beta^\lambda(\lambda, g)}{\epsilon} \right] \left\langle \int d^n x X_1 \right\rangle^R + \frac{\gamma_M(\lambda, g)}{\epsilon} \left\langle \int d^n x X_2 \right\rangle^R + \frac{\gamma_M(\lambda, g)}{\epsilon} \left\langle \int d^n x X_3 \right\rangle^R + \frac{\gamma(\lambda, g)}{\epsilon} \left\langle \int d^n x X_4 \right\rangle^R \\ &\quad + \frac{\tilde{\gamma}(\lambda, g)}{\epsilon} \left\langle \int d^n x (X_5 + X_6) \right\rangle^R + \left[\frac{\beta^\lambda(\lambda, g)}{\lambda\epsilon} - 2\frac{\beta^g(\lambda, g)}{g\epsilon} \right] \left\langle \int d^n x X_7 \right\rangle^R. \end{aligned} \quad (4.2)$$

Hence,

$$\begin{aligned} Z_{11} &= 1 - \frac{\beta^\lambda}{\lambda\epsilon}, \quad Z_{14} = \frac{\gamma}{\epsilon}, \\ Z_{12} &= \frac{\gamma_M}{\epsilon}, \quad Z_{15} + Z_{16} = 2\frac{\tilde{\gamma}}{\epsilon}, \\ Z_{13} &= \frac{\gamma_m}{\epsilon}, \quad Z_{17} = \frac{\beta^\lambda}{\lambda\epsilon} - \frac{2\beta^g}{g\epsilon}. \end{aligned} \quad (4.3)$$

From the fact that the right-hand side of Eq. (4.2) has only simple poles in ϵ , it follows that $\langle \theta_{\mu}^{\text{imp}\mu} \rangle$ is finite at zero momentum. Finiteness of $\langle \theta_{\mu}^{\text{imp}\mu} \rangle$ at zero momentum and to first order in q can be established along the lines of Ref. 3 using the conservation equation.

It should be noted that the above procedure does not yield Z_{15} and Z_{16} separately because $\int d^n x O_5 = \int d^n x O_6$. From Eq. (4.3) it only follows that $Z_{15} + Z_{16}$ has simple poles. However, the theory has a charge-conjugation invariance and the operator O_1 is charge-conjugation invariant. Hence the operator $O_5 - O_6 = i\partial^\mu(\bar{\psi}\gamma_\mu\psi)$ which is odd under charge conjugation cannot appear as a counterterm for O_1 . Hence only the combination $O_5 + O_6$ can appear in the expression for $\langle O_1 \rangle^{\text{UR}}$. This requires that $Z_{15} = Z_{16} = \tilde{\gamma}/\epsilon$.

V. IMPROVEMENT-TERM DEPENDENCE OF THE FORM $\tilde{G}(\epsilon, \lambda_0\mu^{-\epsilon}, g_0^2\mu^{-\epsilon})$

As shown in the previous section, the energy-momentum tensor

$$\theta_{\mu\nu}^{\text{imp}} = \theta_{\mu\nu} + \frac{n-2}{4(1-n)} (\partial_\mu\partial_\nu - g_{\mu\nu}\partial^2)\phi^2$$

is finite at zero momentum. Explicit calculation shows that it is finite only up to $O(\lambda^3)$ at $g=0$, up to $O(g^4)$ at $\lambda=0$ and also in $O(\lambda g^2)$, but a finite improvement is necessarily needed in $O(\lambda^4)$, $O(\lambda g^4)$, and $O(\lambda^2 g^2)$ [see, for example, the Appendix].

In this section, we shall consider a further improvement, where the improvement coefficient \tilde{G} is a *finite* function of *bare* coupling constants:

$$\begin{aligned} \theta_{\mu\nu}^{\text{imp}} &= \theta_{\mu\nu} + \left[\frac{n-2}{4(1-n)} + \frac{\tilde{G}(\epsilon, g_0^2\mu^{-\epsilon}, \lambda_0\mu^{-\epsilon})}{1-n} \right] \\ &\quad \times (\partial_\mu\partial_\nu - g_{\mu\nu}\partial^2)\phi^2, \end{aligned} \quad (5.1)$$

where

$$\tilde{G}(\epsilon, g_0^2\mu^{-\epsilon}, \lambda_0\mu^{-\epsilon}) = \sum_{r=0}^{\infty} \tilde{G}_{2r}(\epsilon, \lambda_0\mu^{-\epsilon})(g_0^2\mu^{-\epsilon})^r, \quad (5.2)$$

where $\tilde{G}_{2r}(\epsilon, \lambda_0\mu^{-\epsilon})$ are finite functions of λ_0 .

From Eq. (3.4) and the fact that Z_{1j} ($j=1,2,\dots,7$) involve simple poles, it follows that

$$\begin{aligned} \langle \theta_{\mu}^{\text{imp}\mu} \rangle &= \text{finite} + [-\epsilon Z_{18} + \tilde{G}(\epsilon, g_0^2\mu^{-\epsilon}, \lambda_0\mu^{-\epsilon})Z_M^{-1}] \\ &\quad \times \langle \partial^2\phi^2 \rangle^R. \end{aligned} \quad (5.3)$$

For future convenience, we shall reexpress this as follows:

$$\begin{aligned}\tilde{G}(\epsilon, g_0^2 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}) &= -\epsilon G(\epsilon, g_0^2 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}), \\ \tilde{G}_{2r}(\epsilon, \lambda_0 \mu^{-\epsilon}) &= -\epsilon G_{2r}(\epsilon, \lambda_0 \mu^{-\epsilon}),\end{aligned}\quad (5.4)$$

where now $G_{2r}(\epsilon, \lambda_0 \mu^{-\epsilon})$ may also contain $1/\epsilon$ terms when expanded in powers of ϵ and $(\lambda_0 \mu^{-\epsilon})$. We thus have

$$\langle \theta_{\mu}^{\text{imp } \mu} \rangle = \text{finite} - \epsilon X \langle \partial^2 \phi^2 \rangle^R, \quad (5.5)$$

where

$$X = Z_{18} + G(\epsilon, g_0^2 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}) Z_M^{-1}. \quad (5.6)$$

Thus to obtain a finite energy-momentum tensor, one

must find a $G(\epsilon, g_0^2 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon})$ such that X given above does not contain worse than simple poles. We shall show, in what follows, that it is not possible to do so consistently except at $g=0$.

For this purpose, we shall use the renormalization-group (RG) equation satisfied by Z_{18} , which can be derived straightforwardly from Eqs. (2.18) and (2.19) following the procedure of Appendix C of I. It is

$$\begin{aligned}(-\lambda\epsilon + \beta^\lambda) \frac{\partial Z_{18}}{\partial \lambda} + \left[-\frac{g\epsilon}{2} + \beta^g \right] \frac{\partial Z_{18}}{\partial g} - 2\tilde{\gamma}_M Z_{18} \\ = Z_{11}\gamma_{18} + Z_{17}\gamma_{78}.\end{aligned}\quad (5.7)$$

Substituting from Eq. (5.6) for Z_{18} and using

$$\begin{aligned}\mu \frac{\partial}{\partial \mu} [G(\epsilon, g_0^2 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}) Z_M^{-1}] &= 2\tilde{\gamma}_M G(\epsilon, g_0^2 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}) Z_M^{-1} + Z_M^{-1} \mu \frac{\partial}{\partial \mu} G(\epsilon, g_0^2 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}) \\ &= 2\tilde{\gamma}_M G(\epsilon, g_0^2 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}) Z_M^{-1} + Z_M^{-1} \mu \frac{\partial}{\partial \mu} \sum_{n=0}^{\infty} G_{2n}(\epsilon, \lambda_0 \mu^{-\epsilon}) (g_0^2 \mu^{-\epsilon})^n\end{aligned}\quad (5.8)$$

one obtains an equation satisfied by X :

$$\begin{aligned}(-\lambda\epsilon + \beta^\lambda) \frac{\partial X}{\partial \lambda} + \left[-\frac{g\epsilon}{2} + \beta^g \right] \frac{\partial X}{\partial g} - 2\tilde{\gamma}_M X - Z_{11}\gamma_{18} - Z_{17}\gamma_{78} \\ = -\epsilon \sum_{n=0}^{\infty} (g_0^2 \mu^{-\epsilon})^n \left[nG_{2n} + \frac{\partial G_{2n}}{\partial (\lambda_0 \mu^{-\epsilon})} \lambda_0 \mu^{-\epsilon} \right] Z_M^{-1}.\end{aligned}\quad (5.9)$$

Now, suppose it were possible to choose G_n 's such that X has no worse than simple poles (which would imply the existence of a finite energy-momentum tensor). Then as Z_{11} and Z_{17} have only simple poles, the left-hand side of Eq. (5.9) has at worst simple poles and hence so does the right-hand side. Hence

$$\begin{aligned}\epsilon^2 \left[\sum_{n=0}^{\infty} (g_0^2 \mu^{-\epsilon})^n \left[nG_{2n} + \frac{\partial G_{2n}}{\partial (\lambda_0 \mu^{-\epsilon})} \lambda_0 \mu^{-\epsilon} \right] \right] Z_M^{-1} \\ = \text{finite}.\end{aligned}\quad (5.10)$$

Then, following the same reasoning as in I, the above equation, after using Eq. (2.21), implies that

$$G_2(\epsilon, \lambda_0 \mu^{-\epsilon}) = 0. \quad (5.11)$$

This implies that the improvement term in Eq. (5.1) is consistent with the finiteness of $\langle \theta_{\mu}^{\text{imp } \mu} \rangle$ in $O(g^2 \lambda^n)$ only if

$$G(\epsilon, g_0^2 \mu^{-\epsilon}, \lambda_0 \mu^{-\epsilon}) = G_0(\epsilon) + O(g^4), \quad (5.12)$$

i.e., only if the improvement term obtained to $O(g^0)$ is sufficient even to $O(g^2)$. But this contradicts the result in the Appendix that an additional improvement term is necessarily needed to make $\theta_{\mu\nu}$ finite in $O(\lambda^2 g^2)$. Hence, we conclude that it is *not* possible to find an improved energy-momentum tensor of the form given in Eq. (5.1) which may be finite even to $O(g^2 \lambda^n)$.

VI. IMPROVEMENT-TERM DEPENDENCE OF THE FORM $\tilde{G}(\epsilon, g^2, \lambda)$

In this section we shall consider an improved energy-momentum tensor of the form

$$\theta_{\mu\nu}^{\text{imp}} = \theta_{\mu\nu} + \left[\frac{n-2}{4(1-n)} + \frac{\tilde{G}(\epsilon, g^2, \lambda)}{1-n} \right] (\partial_\mu \partial_\nu - \partial^2 g_{\mu\nu}) \phi^2, \quad (6.1)$$

where $\tilde{G}(\epsilon, g^2, \lambda)$ is finite for finite g^2 and λ at $\epsilon=0$. Re-tracing the steps of the previous section, we obtain

$$\langle \theta_{\mu}^{\text{imp } \mu} \rangle = \text{finite} - \epsilon X \langle \partial^2 \phi^2 \rangle^R, \quad (6.2)$$

where

$$X = Z_{18} + G(\epsilon, g^2, \lambda) Z_M^{-1}$$

and

$$G(\epsilon, g^2, \lambda) = -\frac{1}{\epsilon} \tilde{G}(\epsilon, g^2, \lambda) \equiv \sum_{n=0}^{\infty} g^{2n} G_{2n}(\epsilon, \lambda). \quad (6.3)$$

As in Sec. V the RG equation satisfied by X can be obtained and rewritten as [see II Eq. (6.4)]

$$\begin{aligned}(-\lambda\epsilon + \beta^\lambda) \frac{\partial X}{\partial \lambda} + \left[-\frac{\epsilon g}{2} + \beta^g \right] - 2\tilde{\gamma}_M X - Z_{11}\gamma_{18} - Z_{17}\gamma_{78} \\ = \left[(-\lambda\epsilon + \beta^\lambda) \frac{\partial G}{\partial \lambda} + \left[-\frac{g\epsilon}{2} + \beta^g \right] \frac{\partial G}{\partial g} \right] Z_M^{-1}.\end{aligned}\quad (6.4)$$

As before, the existence of $\langle \theta_{\mu}^{\text{imp } \mu} \rangle$ implies that

$$\epsilon \left[(-\lambda\epsilon + \beta^{\lambda}) \frac{\partial G}{\partial \lambda} + \left[-\frac{g\epsilon}{2} + \beta^g \right] \frac{\partial G}{\partial g} \right] Z_M^{-1} = \text{finite} . \quad (6.5)$$

Again, following the same procedure as in I and II and using the value of β^g from Eq. (2.14), Eq. (2.20) together with the above equation implies that

$$G(\epsilon, g^2, \lambda) = G_0(\epsilon) + O(g^4) , \quad (6.6)$$

and hence as argued at the end of Sec. V, it is not possible to find an improved energy-momentum tensor of the form given in Eq. (6.1), which may be finite to $O(g^2\lambda^n)$.

APPENDIX

It was shown in both Secs. V and VI that, in order that X has no worse than simple poles to $O(g^2)$, it was *necessary* (but not sufficient) that \bar{G} was a function of ϵ only. Now we show explicitly that even in this case X does have double poles to $O(\lambda^2 g^2)$. The proof proceeds exactly analogous to Appendix A of I.

The double poles in $X = Z_{18} - G_0(\epsilon)Z_M^{-1}$ arise entirely from those in Z_{18} in $O(\lambda^2 g^2)$, since $G_0(\epsilon)$ [which is the same as $g_0(\epsilon)$ of I and II] begins as $O(\epsilon^2)$ (Ref. 7). Direct calculation shows that Z_{18} has no worse than simple poles in $O(\lambda g^2)$ and the simple pole term is nonvanishing. The double-pole term in Z_{18} in $O(\lambda^2 g^2)$ is obtained using the RG equation satisfied by Z_{18} :

$$\begin{aligned} (-\lambda\epsilon + \beta^{\lambda}) \frac{\partial Z_{18}}{\partial \lambda} + \left[-\frac{g\epsilon}{2} + \beta^g \right] \frac{\partial Z_{18}}{\partial g} - 2\bar{\gamma}_M Z_{18} = & \left[1 - \frac{\beta^{\lambda}}{\lambda\epsilon} \right] \left[-\lambda \frac{\partial Z_{18}^{(1)}}{\partial \lambda} - \frac{g}{2} \frac{\partial Z_{18}^{(1)}}{\partial g} \right] \\ & + \left[\frac{\beta^{\lambda}}{\lambda\epsilon} - 2 \frac{\beta^g}{g\epsilon} \right] \left[-\lambda \frac{Z_{78}^{(1)}}{\partial \lambda} - \frac{g}{2} \frac{\partial Z_{78}^{(1)}}{\partial g} \right] . \end{aligned} \quad (A1)$$

Following the same procedure as in Appendix A of I and using (a) Z_{18} and Z_{78} vanish to $O(g^2)$, (b) Z_{78} vanishes at $g=0$, (c) at $g=0$, Z_{18} begins as λ^3 , (d) $Z_{18} - Z_{78}$ has no poles in $O(\lambda g^2)$, since

$$O_1 - O_7 = -\frac{\lambda_0 \phi^4}{4!}$$

(see II), and (e) β^g has no term of $O(\lambda g)$ as verified by direct calculation, one obtains

$$Z_{18(2,1)}^{(2)} = -\frac{1}{3}(\beta_2 - 2\gamma_M^{(1)})Z_{18(1,1)}^{(1)} \neq 0$$

as neither factor vanishes. (Note that $2\gamma_M^{(1)}$ is the same as $2\gamma_m^{(1)}$ of I and II.) Here $Z_{18(m,n)}^{(r)}$ is the coefficient of $\lambda^m g^{2n}/\epsilon^r$ in Z_{18} .

Hence, Z_{18} and therefore X does have double poles in $O(\lambda^2 g^2)$. Therefore, the improvement coefficient $G_0(\epsilon)$ obtained from $O(g^0)$ calculation does not suffice in $O(\lambda^2 g^2)$ to make $\theta_{\mu}^{\text{imp } \mu}$ finite to this order.

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- ¹⁴It has already been shown in I that λ is not multiplicatively renormalizable. The same argument holds here. In Yukawa theory, the scalar mass is also not multiplicatively renormalizable due to the presence of diagrams for the scalar propagator [such as the single fermion loop diagram in $O(g^2)$] which give contributions proportional to m^2 .
¹⁵ ϕ^2 is a multiplicatively renormalizable operator since there is no dimension-2 operator it can mix with. Its renormalization constant Z_M^{-1} can be obtained by dividing Eq. (2.10) by M_0^2 and putting $m=0$ as this renormalization constant is mass independent in the MS scheme.
¹⁶Here γ_M and $\bar{\gamma}_M$ are not the same in contrast with scalar QED and non-Abelian gauge theory (NAGT). The reason lies in the fact that M_0^2 is not multiplicatively renormalizable. One may verify that

$$\bar{\gamma}_M = \gamma_M + \frac{m^2}{M^2} \frac{Z'_M}{Z_M} \left[\gamma_m + \frac{1}{2}\mu \frac{\partial}{\partial \mu} \ln Z'_M \right] .$$

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