Exotic fermions in E_6 and the anomalous magnetic moments of leptons

Thomas G. Rizzo

Ames Laboratory and Department of Physics, Iowa State University, Ames, Iowa 50011

(Received 30 December 1985)

The mixing between the exotic fermions and the usual fermions (in E_6 theories) can lead to flavor-changing neutral currents. Such couplings can produce new and potentially large contributions to the anomalous magnetic moments of the electron and muon. We analyze these possibilities and comment on other processes (such as $\mu \rightarrow e\gamma$) which can be used to constrain E₆ theories.

The recent advent of superstring theories¹ has renewed interest in grand unifled theories (GUT's) based on the gauge group E_6 (Ref. 2). Superstrings have the promise of being complete and finite theories of all of the fundamental forces including gravity. Freedom from anomalies is also possible if the gauge group is chosen to be SO(32) or $E_8\times E'_8$ in ten dimensions.³ The latter gauge group has attracted the most attention since it leads to chiral fermions. In particular, upon compactification, E_8 breaks to E_6 with local $N=1$ supersymmetry below the Planck scale. The second group E'_8 remains unbroken and is associated with a new form of matter (shadow matter).⁴ The fermions then transform as (at least) three 27 representations of E_6 with the possibility that "mirror" fermions⁵ in the 27 representation could also be present. They are all singlets with respect to E'_{s} .

Under the SU(5) subgroup of E_6 the 27 decomposes as

$$
27 = (11 + \overline{5}1 + 10) + (5 + \overline{5}2) + 12, \t(1)
$$

where explicitly for the first generation we have

$$
10 = \begin{bmatrix} u \\ d \end{bmatrix}_L + u_L^c + e_L^c, \quad 5 = \begin{bmatrix} N \\ E \end{bmatrix}_L^c + D_L ,
$$

$$
\overline{5}_1 = \begin{bmatrix} v \\ e \end{bmatrix}_L + d_L^c, \quad \overline{5}_2 = \begin{bmatrix} N \\ E \end{bmatrix}_L + D_L ,
$$

$$
1_1 = v_L^c, \quad 1_2 = N_L^c .
$$
 (2)

Investigations⁶ of symmetry-breaking patterns in the context of E_6 have shown that the low-energy electroweak gauge group is larger than the standard model by (at least) an additional $U(1)$ factor [which we will call $U(1)_x$]. We have recently considered a large class of such theories with only an additional $U(1)_x$ factor⁷ and we limit ourselves to this class in our analysis although our results can be extended to a much larger set of theories.

An important point to notice is that the above fermions do not satisfy the Glashow-Weinberg-Paschos⁸ conditions for the natural absence of flavor-changing neutral currents (FCNC's}; i.e., any mixing among the states (e, E) , (d, D) , (v, N, N') leads to FCNC's.⁹ These FCNC interactions result not only from the usual Z (now called Z_1) exchange, but are also due to exchange of the gauge boson Z_2 (which may be light) associated with the second $U(1)_x$ factor. (The phenomenology of this possibility has recently been discussed by Robinett.⁹) The existence of nonchiral FCNC couplings can, as we will see, produce large contributions to the electron and muon anomalous magnetic moments. (We will neglect any mixing between the two Z bosons in this analysis since the relevant mixing angle has been shown to be quite small.⁷)

The neutral-current interactions for any fermion f can be described by

$$
L = \frac{g}{2c_w} \bar{f} \gamma_\mu [(T_{3L} + T_{3R} - 2x_w Q) - (T_{3L} - T_{3R}) \gamma_5] f Z_1^\mu
$$

+
$$
\frac{g_x}{2} \bar{f} \gamma_\mu [(x_L + x_R) - (x_L - x_R) \gamma_5] f Z_2^\mu , \qquad (3)
$$

where Q is the fermion charge, T_{3L} (T_{3R}) is the third component of weak isospin for the left-handed (righthanded) fermion. x_L (x_R) is the U(1)_x quantum number for the left- (right-) handed fermion and g_x is the associated coupling constant. We limit our discussion in what follows to the charged leptons e and E , but we can easily extend the analysis beyond the first generation. We now imagine a mixing between the e and E fields which, in general may be different for their left- and right-handed components:

$$
\begin{pmatrix} e \\ E \end{pmatrix}_{L,R} \rightarrow U_{L,R}^e \begin{pmatrix} e \\ E \end{pmatrix}_{L,R} , \qquad (4)
$$

where $U_{L,R}^e$ are some general 2×2 unitary matrices which we will take to be purely real and orthogonal for simplicity. Note that if the (e, E) mass matrix is Hermitian then $U_L^e = U_R^e$ so that $\theta_L^e = \theta_R^e$. Let us first examine the Z_L couplings after this rotation; we see that the combination $T_{3L} - 2x_w Q$ has the same value for both e and E. This implies that the resulting FCNC connecting e and E will be purely right handed and, thus, chiral. Writing

$$
L_{Z_1} = \frac{g}{2c_w} \overline{f} \, a \gamma_\mu (v_1 - a_1 \gamma_5) f_\beta Z_1^\mu \tag{5}
$$

we find (with $s_{L(R)}^e = \sin \theta_{L(R)}^e$, $c_{L(R)}^e = \cos \theta_{L(R)}^e$) the couplings

$$
f_{\alpha} = f_{\beta} = e, \quad v_1 = -\frac{1}{2} + 2x_w - \frac{1}{2}(s_R^{\epsilon})^2, \quad a_1 = -\frac{1}{2}(c_R^{\epsilon})^2, \nf_{\alpha} = f_{\beta} = E, \quad v_1 = -\frac{1}{2} + 2x_w - \frac{1}{2}(c_R^{\epsilon})^2, \quad a_1 = -\frac{1}{2}(s_R^{\epsilon})^2, \nf_{\alpha} = e, \quad f_{\beta} = E, \quad v_1 = -\frac{1}{2}s_R^{\epsilon}c_R^{\epsilon}, \quad a_1 = -\frac{1}{2}s_R^{\epsilon}c_R^{\epsilon}.
$$
\n(6)

A similar, but more complex, situation occurs for Z_2 couplings since $x_{L,R}^e$ and $x_{L,R}^F$ are not in general equal. Writing

$$
L_{Z_2} = \frac{g_x}{2} \overline{f} \, a \gamma_\mu (v_2 - a_2 \gamma_5) f_\beta Z_2^\mu \;, \tag{7}
$$

we obtain

$$
f_{\alpha} = f_{\beta} = e, \begin{pmatrix} v_2 \\ a_2 \end{pmatrix} = \left[x_L^e (c_L^e)^2 + x_L^E (s_L^e)^2 \right]
$$

\n
$$
\pm \left[x_R^e (c_R^e)^2 + x_R^E (s_R^e)^2 \right],
$$

\n
$$
f_{\alpha} = f_{\beta} = E, \begin{pmatrix} v_2 \\ a_2 \end{pmatrix} = \left[x_L^E (c_L^e)^2 + x_L^e (s_L^e)^2 \right]
$$

\n
$$
\pm \left[x_R^E (c_R^e)^2 + x_R^e (s_R^e)^2 \right],
$$

\n
$$
f_{\alpha} = e, f_{\beta} = E, \quad v_2
$$

\n
$$
f_{\alpha} = E, f_{\beta} = e, \quad a_2
$$

\n
$$
\pm (x_L^e - x_L^E) s_L^e c_L^e
$$

\n
$$
\pm (x_R^e - x_R^E) s_R^e c_R^e,
$$

\n(9)

so that unless either $(x_L^e - x_L^E)$ or $(x_R^e - x_R^E) = 0$ our FCNC's are nonchiral even if $\theta_L^E = \theta_R^e$. It should be noted that the mixing angles for the other fermions such as (d,D) could be different from θ_L^e and θ_R^e and different FCNC's will be present for those fermions. In the (e, E) sector we know the angle θ_R^e must be small from data on $e^+e^- \rightarrow \mu^+\mu^-$ (Ref. 9).

This same mixing results in a modification of the charged-current (CC) structure as well. Limiting ourselves to the electroweak group $SU(2)_L\times U(1)_Y\times U(1)_X$ we still have only a single W boson so that CC remains left handed. The two left-handed doublets (we ignore N' for now)

$$
\begin{bmatrix} v \\ e \end{bmatrix}_{L}, \quad \begin{bmatrix} N \\ E \end{bmatrix}_{L} \tag{10}
$$

are then mixed by the rotation U_L^e as well as the corresponding (v, N) rotation U_L^v . Thus the CC coupling to W takes the form

$$
\frac{g}{2\sqrt{2}} [\bar{v}\gamma_{\mu}(1-\gamma_{5})e + \bar{N}\gamma_{\mu}(1-\gamma_{5})E] \cos(\theta_{L}^{e} - \theta_{L}^{v})W^{\mu} + \frac{g}{\sqrt{2}} [\bar{v}\gamma_{\mu}(1-\gamma_{5})E + \bar{N}\gamma_{\mu}(1-\gamma_{5})e] \sin(\theta_{L}^{e} - \theta_{L}^{v})W^{\mu}.
$$
\n(11)

Clearly, universality constrains the difference in angles $\theta_L^e - \theta_L^v$ to be quite small.⁹ Note that these couplings are all still chiral.

Thus present CC and NC data tell us that θ_R^e is small and the difference $\theta_L^e - \theta_L^v$ is small; the individual values of $\theta_L^{e,v}$ may still be large.

Figure ¹ shows the two generic contributions of gauge bosons X and fermions F to the muon anomalous magnetic moment. The various contributions depend, of course, on the nature of the particles X and F . For the models under consideration here these are several distinct sets of identifications possible. For diagram (A) we have the possibilities

FIG. 1. Feynman diagrams contributing to a_{μ} in a general gauge theory.

*A*₀:
$$
X=Z_1
$$
, $F=\mu$,
\n*A*₁: $X=Z_1$, $F=M$,
\n*A*₂: $X=Z_2$, $F=\mu$,
\n*A*₃: $X=Z_2$, $F=M$,

where M is the new muon-type heavy charged lepton corresponding to E . For diagram (B) we find two possibilities

B₀:
$$
X = W
$$
, $F = v_{\mu}$,
\nB₁: $X = W$, $F = N_{\mu}$. (13)

The standard-model (SM) contributions are already well known¹⁰ and are given by the sum of A_0 and B_0 ; this leads to the SM prediction

$$
(a_{\mu})^{\rm SM} = 20 \times 10^{-10}
$$
 (14)

and thus¹¹

$$
(a_{\mu})^{\text{QED}} + (a_{\mu})^{\text{SM}} - (a_{\mu})^{\text{expt}} = (27 \pm 69) \times 10^{-10} , \quad (15)
$$

so that any additional contributions beyond the SM are quite constrained. We now turn to the new contributions A_{1-3} and B_1 .

The A_1 contribution¹⁰ can be written as

$$
A_1 = \frac{G_F m_\mu^2}{4\sqrt{2}\pi^2} (s_R^\mu c_R^\mu)^2 F(z)
$$
 (16)

with
$$
z = (M_M/M_1)^2
$$
, $r = (1-z)^{-1}$, and
\n
$$
F(z) = r \left[\frac{5}{6} - \frac{5}{2} + r^2 + r(r^2 - 3r + 2) \ln \frac{r-1}{r} \right] + \frac{r}{2} z \left[\frac{5}{6} + \frac{3r}{2} + r^2 + r^2(1+r) \ln \frac{r-1}{r} \right].
$$
\n(17)

Note that $F(z)$ is a slowly varying function of z. Figure 2

FIG. 2. A plot of the function $F(z)$ vs \sqrt{z} . FIG. 3. A plot of the function $G(z)$ vs \sqrt{z} .

$$
A_1 \leq (-0.50 \sim -0.73) \times 10^{-10} . \tag{18}
$$

The second contribution is A_2 ; for $m_\mu^2/M_2^2 \ll 1$ we find that

$$
A_2 = \frac{1}{4\pi^2} \frac{m_\mu^2}{M_2^2} \frac{g_x^2}{12} [(x_L^\mu + x_R^\mu)^2 - 5(x_L^\mu - x_R^\mu)^2]
$$

= $\frac{\alpha}{\pi} \frac{m_\mu^2}{M_1^2} \left[\frac{\alpha_x}{\alpha} \right] \left[\frac{M_1}{M_2} \right]^2 \Delta$, (19)

where $\alpha_x = g_x^2/4\pi$ and $\Delta = x_L^\mu x_R^\mu - (x_L^{\mu 2} + x_R^{\mu 2})/3$. Note $x_{L(R)}^{\mu} = x_{L(R)}^{\epsilon}$ $x_{L(R)}^M = x_{L(R)}^E$; these values can be found in Table I for the four models discussed in Ref. 7. In this same reference we saw that, roughly speaking, $(\alpha_x/\alpha)(M_1/M_2)^2$ can be in the range $(6-50)\times 10^{-3}$ and the value of Δ is given by Table I. Thus A_2 is quite small:

$$
|A_2| = (1.6 - 268) \times 10^{-12}
$$
 (20)

with the sign depending on the sign of Δ ; three out of the four cases lead to $A_2<0$. This contribution is quite small.

We now turn to the potentially large contribution A_3 ; in this case we have $(z = m_M^2/M_2^2)$

$$
S = \frac{\alpha_x}{4\pi} \frac{m_{\mu}^2}{M_2^2} \left[(v_2^2 + a_2^2) F(z) \right]
$$

\nFor l and l is the second contribution to a^{μ} and $a^{\mu} = \frac{a_x}{4\pi} \frac{m_{\mu}^2}{M_2^2} \left[(v_2^2 + a_2^2) F(z) \right]$
\n
$$
A_1 \leq (-0.50 \sim -0.73) \times 10^{-10}.
$$

\n
$$
A_2 \leq (-0.50 \sim -0.73) \times 10^{-10}.
$$

\n
$$
B = \frac{(18)}{4\pi} \frac{m_{\mu}^2}{M_2^2} \left[(v_2^2 + a_2^2) F(z) \right]
$$

\n
$$
+ (v_2^2 - a_2^2) G(z) \left[\frac{M_2}{m_{\mu}} \right] , \quad (21)
$$

 v_2 and a_2 are given by Eq. (9) and $F(z)$ is given by Eq. (17). $G(z)$ is $[r=(1-z)^{-1}]$

$$
G(z) = rz^{1/2} \left[\left[2r - 1 + 2r(r-1) \ln \frac{r-1}{r} \right] - \frac{1}{2} z \left[\frac{1}{2} + r + r^2 \ln \frac{r-1}{r} \right] \right].
$$
 (22)

Note that the term proportional to $(v_2^2 - a_2^2)$ is enhanced by an extremely large factor $M_2/m_\mu \ge 2000$. This term vanishes identically, however, in the limit of chiral couplings $v_2^2 - a_2^2 \rightarrow 0$. The function $G(z)$ is shown in Fig. 3. Equation (21) can be recast in the form

$$
A_3 = -\frac{\alpha_x}{2\pi} \frac{m_\mu^2}{M_2^2} \left[[(x_L^\mu - x_L^M)^2 (s_L^\mu c_L^\mu)^2 + (x_R^\mu - x_R^M)^2 (s_R^\mu c_R^\mu)^2] F(z) + 2(x_L^\mu - x_L^M) (x_R^\mu - x_R^M) s_L^\mu s_R^\mu c_L^\mu c_R^\mu \right. \\ \left. \times G(z) \left[\frac{M_2}{m_\mu} \right] \right]. \tag{23}
$$

TASI.E I. Values of the various parameters for the four models discussed in Ref. 7.

Model	хĭ	хå	x_L	×Ř			ρ
A			— z			-9	4
в			-2	— z		15	4
C	$\overline{}$	-2				-12	
			$\overline{}$		$\overline{12}$		$\overline{4}$

Now $(s_R^{\mu} c_R^{\mu})^2 \le 0.05$ suppresses the $(x_R^{\mu} - x_R^{\mu})^2$ term while $(s_L^{\mu} c_L^{\mu})^2$ can, in principle, be quite large with $\rho \equiv (x_L^{\mu} - x_L^{\mu})^2$ of order unity (see Table I for values of ρ in the four models under consideration). With both $F(z)$ and $G(z)$ of order unity, A_3 can be approximated by

$$
A_3 = -32.1 \times 10^{-10} \left[\frac{\alpha_x}{\alpha} \frac{M_1^2}{M_2^2} \right] \chi_s^{\mu} c_L^{\mu} s_R^{\mu} c_R^{\mu} G(z) \frac{M_2}{m_{\mu}} ,
$$
\n(24)

where $X = (x_L^{\mu} - x_L^M)(x_R^{\mu} - x_R^M)$. As a numerical example where $\chi = (x_L - x_L)(x_R - x_R)$. As a numerical example
we take $M_2 = 200$ GeV and $M_M = 180$ GeV with $(\alpha_x / \alpha) M_1^2 / M_2^2$ given by the above range. This yields

$$
A_3 = (-173 \text{ to } -1440) \times 10^{-10} \chi_s^{\mu} c_L^{\mu} s_R^{\mu} c_R^{\mu} \,. \tag{25}
$$

Since $|\chi| \sim 10$ in all but one model (where it vanishes identically) the value of A_3 can be quite large. For further demonstration purposes taking model 8 with $(s_R^{\mu} c_R^{\mu})^2 \approx 0.05$ we find

$$
A_3 = (-130 \text{ to } -1080) \times 10^{-10} (s_L^{\mu} c_L^{\mu} / s_R^{\mu} c_R^{\mu})
$$
 (26)

which can be very sizable. From this we conclude that either M_2 is much heavier than expected ($>$ 200 GeV or so), the mixing angles θ_L^{μ} , θ_R^{μ} are both very tiny, or that X is naturally small (or zero as in model D).

To see how (26) when combined with (15) can constrain Es model parameters let us again consider model B. With⁷ $\alpha_x \simeq \frac{1}{3000}$ we find in general that

$$
A_3 = -18\,083 \times 10^{-10} \frac{M_1}{M_2} (s_L^{\mu} c_L^{\mu}) (s_R^{\mu} c_R^{\mu}) G(z) \ . \tag{27}
$$

We take $G(z) \approx \frac{1}{2}$ and $\theta_L^{\mu} = \theta_R^{\mu}$ (the case of an Hermitian mass matrix) and find that at the $1\sigma (2\sigma)$ level

$$
\frac{M_1}{M_2} (s_L^{\mu} c_L^{\mu})^2 \le 4.65 \times 10^{-3} (1.23 \times 10^{-2}) \ . \tag{28}
$$

Figure 4 shows the allowed region in the $M_2/M_1-(s^{\mu}_L c^{\mu}_L)^2$
plane from the a^{μ} constraint as well as universality and the ρ neutral-current parameter.⁷ Similar results hold for the other models.

We now turn to the last contribution to a^{μ} , that coming from B_1 . We find $(z = M_N^2/M_W^2)$

> O.^l 8 O. ^l 5

strained by the ρ parameter, universality, and the value of a^{μ} . Both 1σ and 2σ limits are shown.

FIG. 5. A plot of the function $H(z)$ vs \sqrt{z} .

$$
B_1 = \frac{1}{4\pi^2} \frac{m_{\mu}^2}{M_{w}^2} \left[\frac{g}{2\sqrt{2}} \sin(\theta_L^{\mu} - \theta_L^{\nu}) \right]^2 H(z) .
$$
 (29)

With $r = (1-z)^{-1}$ and $t = rz$ we find

$$
H(z) \equiv 2r \left[\frac{5}{6} - \frac{3}{2}t + t^2 + t^2(1-t) \ln \frac{1+t}{t} \right]
$$

+ zr \left[\frac{5}{6} + \frac{5}{2}t + t^2 - t(2+3t + t^2) \ln \frac{1+t}{t} \right]. (30)

 $H(z)$ is shown in Fig. 5. B_1 can be rewritten as

$$
B_1 = \frac{G_F m_\mu^2}{4\sqrt{2}\pi^2} H(z) \sin^2(\theta_L^\mu - \theta_L^\nu)
$$

= 32.1 × 10⁻¹⁰H(z) sin²($\theta_L^\mu - \theta_L^\nu$). (31)

Note that the sign of B_1 is always positive unlike the model-dependent sign of the A_i . B_1 is constrained to be small by universality; if $\sin^2(\theta_L^{\mu} - \theta_L^{\nu}) \le 0.05$, then

$$
B_1 \leq 2.7 \times 10^{-10} \tag{32}
$$

which is a small but perhaps observable contribution.

What about a_e ? Including the latest results from Kinoshita and co-workers¹² we find that

$$
(a_e)^{\text{th}} - (a_e)^{\text{expt}} = (109 \pm 111) \times 10^{-12} . \tag{33}
$$

The new E_6 contributions for a_e can be obtained from those for a_{μ} by letting $m_{\mu} \rightarrow m_e$ (with $\mu \rightarrow e$ in all the angles as well). As before, the only major contribution comes from A_3 ; we find

$$
A_3(e) = -29393 \times 10^{-12} G(z) \left[\frac{\alpha_x}{\alpha} \frac{M_1^2}{M_2^2} \right]
$$

$$
\times Xs_L^e c_L^e s_R^e c_R^e \left[\frac{M_2}{200 \text{ GeV}} \right]. \tag{34}
$$

With $\alpha_x \simeq \frac{1}{3000}$ and $\theta_L^e = \theta_R^e$ we find

$$
A_3(e) = -583\chi \times 10^{-12} \frac{M_1}{M_2} (s_L^e c_L^e)^2 G(z)
$$
 (35)

and for model B with $G(z) \approx \frac{1}{2}$ the constraint at the lo (2σ) level becomes

$$
\frac{M_1}{M_2} (s_L^{\epsilon} c_L^{\epsilon})^2 \le 2.29 \times 10^{-6} (2.58 \times 10^{-2}) \ . \tag{36}
$$

If the 1σ limit is taken seriously the improvement over (28) is better than a factor of 2000. The 2σ limit from (36) is about twice that from (28). Similar constraints hold for the other models as well.

As a last remark, we consider the radiative decays $\mu \rightarrow e\gamma$ and $\nu \rightarrow \nu' \gamma$ which are highly suppressed in the SM and slight extensions thereof. Since virtually the same kinds of diagrams as shown in Fig. ¹ are also responsible for these decays we might expect large enhancements in these radiative rates in E_6 theories (although further intergenerational mixing angles are also present). As is well known¹³ the existence of fermions with masses comparable or greater than that of the W or Z tend to negate the Glashow-Iliopoulos-Maiani mechanism responsible for suppressing these processes. It may be possible to strengthen our constraints on the M_1/M_2 ratio and the various mixing angles from further examination of these processes. An explicit calculation¹⁴ indicates no large contribution in radiative ν decay because of chiral cou-

- ¹J. H. Schwarz, Phys. Rep. 89, 223 (1982); G. T. Horowitz, University of California at Santa Barbara report, 1985 (unpublished); C. B. Thorn, lectures presented at the Workshop on Unified String Theories, Institute for Theoretical Physics, 1985 (unpublished).
- ²For recent work on E_6 not directly related to superstrings see, for example, R. W. Robinett and J. L. Rosner, Phys. Rev. D 26, 2356 (1982); R. W. Robinett, ibid. 26, 2388 (1982); P. K. Mohapatra, R. N. Mohapatra, and P. B. Pal, University of Maryland report, 1985 (unpublished); P. H. Frampton and T. W. Kephart, Phys. Rev. D 25, 1459 (1982); A. Sen, Phys. Rev. Lett. 55, 33 {1985).
- ³M. B. Green and J. H. Schwarz, Phys. Lett. 149B, 117 (1984); P. Candelas, G. T. Horowitz, A. Strominger, and E. Witten, Nucl. Phys. B258, 46 (1985); E. Witten, Phys. Lett. 149B, 351 (1984); D. J. Gross, J. Harvey, E. Martinec, and R. Rohm, Phys. Rev. Lett. 54, 502 (1985); Nucl. Phys. 8256, 253 (1985).
- ⁴E. W. Kolb, D. Seckel, and M. S. Turner, Report No. Fermilab-CONF-85 1114A, 1985 (unpublished); Report No. Fermilab-Pub-85 016-A, 1985 (unpublished).
- 5For some recent work on mirror fermions see, for example, J. Oliensis and C. Albright, Phys. Lett. 1608, 121 (1985); S. Mishra, S. P. Misra, S. Panda, and U. Sarkar, University of Calcutta Report No. IP/BBSR/84-12 (unpublished}; G. Senjanović, F. Wilczek, and A. Zee, Phys. Lett. 141B, 389 (1984); J. Maalampi and M. Ross, ibid. 146B, 333 (1984); S. Rajpoot and J. G. Taylor, ibid. 1428, 365 (1984); D. Forigon and M. Roos, ibid. 1478, 34 (1984); M. Roos, ibid. 135B, 4g7 (1984)
- 6B. A. Ovrut, University of Pennsylvania report, 1985 (unpublished); S. Cecotti, J.-P. Derendinger, S. Ferrara, and L.

plings, whereas large contributions can occur in two out of the four models for $\mu \rightarrow e\gamma$.

In conclusion, we have examined the new fermion and gauge-boson contributions to the muon and electron anomalous magnetic moments present in E_6 theories. Because of the existence of nonchiral fiavor-changing neutral-current couplings present in such theories we can constrain the mixing between the ordinary and exotic fermions and/or the ratio M_1/M_2 . Both a_e and a_u give comparable limits at present, but new experiments could show evidence of non-SM contributions. The amplitude for $\mu \rightarrow e\gamma$ may also be enhanced in these models. (Our analysis has ignored Higgs-boson contributions which may also be potentially large in such theories but are very model dependent.) Much more work needs to be done in examining the phenomenology of E_6 theories.

This work was supported by the U.S. Department of Energy, Contract No. W-7405-Eng-82, Office of Basic Science (KA-01-01), Division of High Energy Physics and Nuclear Physics.

Girardello, Phys. Lett. 1568, 318 (1985); S. M. Barr, Phys. Rev. Lett. 55, 2278 (1985); R. Holman and D. Reiss, Report No. Fermilab Pub-851/130-A (unpublished); M. Dine, Institute for Advanced Studies report, 1985 (unpublished); M. Dine, V. Kaplunovsky, M. Mangano, C. Nappi, and N. Seiberg, Nucl. Phys. B259, 549 (1985); E. Witten, ibid. B258, 75 (1985); Y. Hosotani, Phys. Lett. 1268, 309 (1983); J. P. Derendinger, L. E. Ibanez, and H. P. Nilles, Nucl. Phys. 8267, 365 (1986};J. Rosner, Enrico Fermi Institute Report No. EFIL 85-34, 1985 (unpublished); V. Barger et al., Phys. Rev. Lett. 56, 30 (1986); 33, 1912 (1986).

- 7J. L. Hewett, T. G. Rizzo, and J. A. Robinson, Phys. Rev. D 33, 1476 (19S6).
- 8S. L. Glashow and S. Weinberg, Phys. Rev. D 15, 1958 (1977); E. A. Paschos, ibid. 15, 1966 (1977).
- ⁹R. W. Robinett, Phys. Rev. D 33, 1908 (1986). See, also, S. M. Barr, Phys. Rev. Lett. 55, 2278 (1985).
- ¹⁰See, for example, J. P. Leveille, Nucl. Phys. **B137**, 63 (1977).
- 11 For a recent comparison of theory and experiment see, S. R. Moore, K. Whisnant, and B-L. Young, Phys. Rev. D 31, 105 (1985), and references cited therein.
- ¹²T. Kinoshita and W. B. Lindquist, Phys. Rev. Lett. 47, 1573 (1981); Phys. Rev. D 27, S67 (1983); 27, 877 (1983); 27, 886 (1983). For a very recent comprehensive review, see, T. Kinoshita and J. Sapirstein, in Proceedings of the Ninth InternationaI Conference on Atomic Physics, edited by R. S. Van Dyck, Jr. and E. N. Fortson (World Scientific, Singapore, 1985).
- ¹³T. P. Cheng and L.-F. Li, Phys. Rev. Lett. 45, 1980 (1980); 38, 381 (1977); W. Marciano and A. Sanda, Phys. Lett. 67B, 303 (1977).
- 14 T. G. Rizzo (in preparation).