

Nonuniversal neutral currents of  $\nu_e$  and  $\nu_\mu$

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The present status and the general consequences of a violation of  $\nu_e$ - $\nu_\mu$  universality in weak neutral currents are examined from a phenomenological point of view. If  $\xi$  is the strength of the  $\nu_e$  coupling relative to the  $\nu_\mu$  coupling, an analysis of the available data gives two possible values:  $\xi = 1.45 \pm 0.2$  and  $-0.6 \pm 0.2$ . A measurement of the total cross section  $\sigma(\nu_e e)$  would reveal the deviations of  $\xi$  from unity much more easily than the measurement of  $\sigma(\bar{\nu}_e e)$ . The effects of nonuniversality on neutrino cross sections when the  $\nu_e$  undergoes flavor oscillations in vacuum are studied. A sizable nonuniversality will facilitate the observation of matter effects on the neutrino oscillations.

I. INTRODUCTION

The experimental data using  $\nu_\mu$  and  $\bar{\nu}_\mu$  beams reveal that the structures of both the hadronic and electronic parts of the weak neutral current are in remarkable agreement with the minimal electroweak gauge model.<sup>1</sup> The corresponding data with  $\nu_e$  and  $\bar{\nu}_e$  beams are, however, not so extensive, because these beams are not available with sufficiently high intensity. Thus it remains an open question whether the coupling of the  $\nu_e$  to the neutral current is indeed equal to that of the  $\nu_\mu$ , as required by the standard model. Besides the need to verify the model further, it is obviously very important to check the data on various neutrino scattering processes for any violations of the neutrino flavor universality.

In the following we would like to examine the consequences of a neutral-current interaction which violates the hypothesis of  $\nu_e$ - $\nu_\mu$  universality in a minimal way, but otherwise conforms to the customary view. That is, we assume the validity of the conventional ( $V-A$ ) theory with  $e$ - $\mu$  universality for the charged-current interactions. For the neutral-current processes also we shall take the usual current  $\times$  current interaction with  $V$  and  $A$  currents, except that the  $\nu_e$  current has a coupling strength  $\xi$  relative to the  $\nu_\mu$  current;

$$H_W = \frac{G}{\sqrt{2}} (J_\lambda^\dagger J^\lambda + N_\lambda N^\lambda)$$

$$N_\lambda = N_\lambda^{(h)} + \bar{\nu}_\mu \gamma_\lambda (1 + \gamma_5) \nu_\mu + \bar{e} \gamma_\lambda (g_V + g_A \gamma_5) e$$

$$+ \xi \bar{\nu}_e \gamma_\lambda (1 + \gamma_5) \nu_e + \dots, \tag{1}$$

where  $G$  is the Fermi coupling constant,  $J_\lambda$  is the charged current,  $N_\lambda^{(h)}$  is the hadronic part of the neutral current, and the dots refer to the currents  $\bar{\mu}\mu$ ,  $\bar{\nu}_\tau\nu_\tau$ , etc., with which we will not be concerned here. The real parameter  $\xi$  will be equal to unity if the usually assumed  $\nu_e$ - $\nu_\mu$  universality holds.

In this paper, we present a phenomenological study of the consequences of  $\xi \neq 1$  for a variety of experiments that can be done using  $\nu_e$  and  $\bar{\nu}_e$  beams at reactors, meson factories, muon rings, and accelerators. Of special interest would be the effect of  $\nu_e$ - $\nu_\mu$  nonuniversality on the oscillations of neutrinos passing through vacuum and through a material medium, as relevant for the deep-mine experiments.

II. NEUTRINO-HADRON SCATTERING

From the definition of  $\xi$  in Eq. (1), it is obvious that the differential cross sections for a neutral-current process (either exclusive or inclusive) on a nuclear target induced by a  $\nu_e$  and by a  $\nu_\mu$  of the same energy are in the ratio of  $\xi^2$  to 1,

$$d\sigma(\nu_e N \rightarrow \nu_e + \dots) = \xi^2 d\sigma(\nu_\mu N \rightarrow \nu_\mu + \dots). \tag{2}$$

This relation, which may be taken as the definition of  $\xi$ , holds for the corresponding antineutrinos as well. A special case which is of experimental interest is the deuteron disintegration by reactor antineutrinos,

$$d\sigma(\bar{\nu}_e d \rightarrow \bar{\nu}_e np) = \xi^2 d\sigma(\bar{\nu}_\mu d \rightarrow \bar{\nu}_\mu np). \tag{3}$$

In order to estimate the value of  $\xi$  from the avail-

able information with reactor neutrinos, we shall substitute the calculated value for  $\bar{\nu}_\mu d$  using the Weinberg-Salam model (which has been tested in the GeV range), and use the available experimental information on the left-hand side for  $\bar{\nu}_e d$ . In the notation of Reines *et al.*,<sup>2</sup> the measured value of  $\sigma(\bar{\nu}_e d)$ , expressed as the ratio

$$\begin{aligned} [\sigma(\bar{\nu}_e d \rightarrow e^+ nn) / \sigma(\bar{\nu}_e d \rightarrow \bar{\nu}_e np)]_{\text{expt}} \\ = 0.167 \pm 0.093, \quad (4) \end{aligned}$$

should therefore be compared with the theoretical value<sup>2</sup> which accounts for a possible nonuniversal neutral-current coupling, namely,

$$(\sigma_{\text{CCd}} / \sigma_{\text{NCd}})_{\text{theory}} = (0.43 \pm 0.01) / \xi^2. \quad (5)$$

$$\frac{d\sigma}{dy}(\nu_\mu e) = \hat{\sigma} \left[ (g_V + g_A)^2 + (g_V - g_A)^2 (1-y)^2 + \frac{my}{E} (g_A^2 - g_V^2) \right], \quad (7)$$

$$\frac{d\sigma}{dy}(\nu_e e) = \hat{\sigma} \left\{ [2 + \xi(g_V + g_A)]^2 + \xi^2 (g_V - g_A)^2 (1-y)^2 + \frac{my}{E} [(1 + \xi g_A)^2 - (1 + \xi g_V)^2] \right\}, \quad (8)$$

$$\hat{\sigma} \equiv \frac{G^2 m E}{2\pi}. \quad (9)$$

The corresponding formulas for  $\bar{\nu}_e$  are obtained simply by interchanging the coefficients of the first two terms [the coefficient of the first ( $y$ -independent) term is exchanged with that of the  $(1-y)^2$  term].

At high enough energies, where we can neglect the  $(m/E)$  terms, we obtain

$$\xi^2 = \frac{(1-y)^2 \frac{d\sigma}{dy}(\nu_e e) - \frac{d\sigma}{dy}(\bar{\nu}_e e)}{(1-y)^2 \frac{d\sigma}{dy}(\nu_\mu e) - \frac{d\sigma}{dy}(\bar{\nu}_\mu e)}. \quad (10)$$

This relation presumably will be more useful if we consider only those events in which the recoil electron energy is maximal ( $y \simeq 1$ ), since

$$\xi^2 = \left[ \frac{d\sigma}{dy}(\bar{\nu}_e e) / \frac{d\sigma}{dy}(\bar{\nu}_\mu e) \right]_{y=1}. \quad (11)$$

On the other hand, we may also integrate Eqs. (7) and (8) over  $y$  from 0 to 1 and obtain

$$\xi^2 = \frac{\sigma(\nu_e e) - 3\sigma(\bar{\nu}_e e)}{\sigma(\nu_\mu e) - 3\sigma(\bar{\nu}_\mu e)}. \quad (12)$$

This relation for the case of  $\xi=1$  was derived first by Sehgal,<sup>3</sup> and was also found to be valid in a more general context<sup>4-6</sup> such as  $S, P, T$  couplings and the presence of nondiagonal  $V, A, S, P, T$  couplings.

We thus obtain the estimate

$$|\xi| = 1.6 \pm 0.5. \quad (6)$$

The error on  $\xi$  may have to be increased to reflect the uncertainties inherent in the calculated reactor neutrino spectrum that enters Eq. (5), and also the complete neglect of any oscillations of reactor antineutrinos within a few meters.

### III. ELASTIC $\nu_e$ SCATTERING

Denoting the mass of the electron by  $m$ , the laboratory energy of the incident neutrino by  $E$ , and the laboratory kinetic energy of the recoil electron by  $(yE)$ , the  $y$  distributions for  $\nu_\mu e$  and  $\nu_e e$  scattering according to Eq. (1) are

For fixing the sign of  $\xi$  we look at the sign of the interference term in Eq. (8) between the charged-current contribution 2 and the neutral-current contribution  $\xi(g_V + g_A)$ . If we set  $\xi=1$  by assuming  $\nu_e$ - $\nu_\mu$  universality, then experiments can test for the destructive interference between the charged and neutral currents expected in the Weinberg-Salam model, a point well emphasized by Kayser *et al.*<sup>6</sup> On the other hand, treating  $\xi$  as unknown but assuming the validity of the Weinberg-Salam model for  $\nu_\mu e$  scattering, we can fix the sign of  $\xi$ . In any case it is obvious that the sign of the interference term gives only the overall sign of the combination  $\xi(g_V + g_A)$ .

In order to see the sensitivity of the total cross sections to the parameter  $\xi$ , we shall integrate the expressions (7) and (8) over  $y$  from 0 to 1, and substitute the couplings of the Weinberg-Salam model,

$$\begin{aligned} g_V &= -\frac{1}{2} + 2 \sin^2 \theta_W, \\ g_A &= -\frac{1}{2}. \end{aligned} \quad (13)$$

For purposes of numerical work we shall substitute the typical value<sup>1</sup>

$$\sin^2 \theta_W = 0.23, \quad (14)$$

and obtain the total cross sections of high-energy

neutrinos on electron target in units of  $\hat{\sigma}$  defined by Eq. (9),

$$\begin{aligned}\sigma(\nu_\mu e) &= 0.362\hat{\sigma}, \\ \sigma(\bar{\nu}_\mu e) &= 0.309\hat{\sigma}, \\ \sigma(\nu_e e) &= [2.20 + (1-\xi)(1.80 - 0.362\xi)]\hat{\sigma}, \\ \sigma(\bar{\nu}_e e) &= [0.922 + (1-\xi)(0.411 - 0.309\xi)]\hat{\sigma}.\end{aligned}\quad (15)$$

These cross sections are plotted in Fig. 1 as a function of  $\xi$ . The dots on the curves at  $\xi=1$  are the expectations on the basis of  $\nu_\mu$ - $\nu_e$  universality. The  $\sigma(\bar{\nu}_e e)$  curve shows that the universality point  $\xi=1$  is located inside a broad minimum (exact minimum is at  $\xi=1.17$  for  $\sin^2\theta_W=0.23$ ), and hence the existence of even significant departures from universality,  $\xi \simeq (\frac{1}{2}-2)$ , will not easily be noticeable in the measurements of  $\sigma(\bar{\nu}_e e)$ . For this reason data on  $\sigma(\nu_e e)$ , say from meson factories, will be very valuable. Secondly, for a given absolute value of  $\xi$ , the negative value leads to a larger cross section than the positive value—an effect due to the constructive interference between the charged-current and neutral-current interactions.

We now turn to an analysis of the reactor data of Reines *et al.*<sup>7</sup> on the difficult reaction  $\bar{\nu}_e e \rightarrow \bar{\nu}_e e$  for estimating the value of  $\xi$ . The experimental results are available in terms of the charged-current cross section  $\sigma_{\nu-A}$  averaged over a calculated antineutrino spectrum and integrated over two ranges of the recoil electron's kinetic energy  $T_e$ . Avignone and Greenwood<sup>8</sup> have listed the numerical values of the energy integrals (called  $G_i$ 's by them) which occur in the  $\bar{\nu}_e e$  cross-section formula, due to the folding of the incident spectrum as a function of the electron energy cuts. (The  $G_i$ 's appropriate for the recoil-

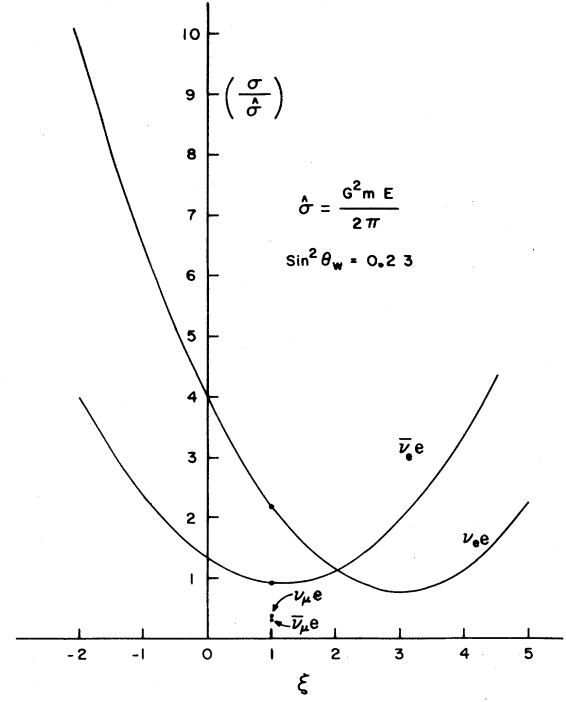


FIG. 1. The total cross sections  $\sigma(\nu e)$  as a function of the nonuniversality parameter  $\xi$ . The values of  $\sigma(\nu_\mu e)$  and  $\sigma(\bar{\nu}_\mu e)$ , and the two points marked on the curves of  $\sigma(\nu_e e)$  and  $\sigma(\bar{\nu}_e e)$  at  $\xi=1$  are obtained from the Weinberg-Salam model when  $\sin^2\theta_W=0.23$ .

electron-energy bins can also be readily extracted from Table VI of Ref. 6.) We use these integrals which incorporate the updated neutrino reactor spectrum, substitute  $\sin^2\theta_W=0.23$ , and obtain the following relations for  $\xi$  using the data of Reines *et al.*:

$$1.5 < T_e < 3.0 \text{ MeV: } 9.68\xi^2 - 14.86\xi + 21.2 = 21.2(0.87 \pm 0.25), \quad (16)$$

$$3.0 < T_e < 4.5 \text{ MeV: } 2.53\xi^2 - 2.28\xi + 2.9 = 2.9(1.7 \pm 0.44). \quad (17)$$

The numbers 21.2 and 2.9 appearing in the  $\xi$ -independent terms are the values of  $\sigma_{\nu-A}$  (in arbitrary units). Each of the quadratic equations implies a pair of alternative values of  $\xi$ ,

$$\xi = \begin{cases} 1.32 \pm 0.50 \\ 0.22 \pm 0.50 \end{cases} \quad \text{for } 1.5 < T_e < 3.0 \text{ MeV}, \quad (18)$$

$$\xi = \begin{cases} 1.45 \pm 0.25 \\ -0.55 \pm 0.25 \end{cases} \quad \text{for } 3.0 < T_e < 4.5 \text{ MeV}. \quad (19)$$

These values as well as those of Eq. (6) coming from  $\bar{\nu}_e d$  are schematically shown in Fig. 2, and compared with the  $\nu_e$ - $\nu_\mu$  universality value expected in the Weinberg-Salam model. Here neutrino oscillations are ignored. The dashed line denotes the alternative solution of  $\xi$  which is less than unity. The higher energy bin data of  $\bar{\nu}_e d$  lead to the determination of  $\xi$  with the smallest error. It should however be noted that the value of  $\xi$  from  $\bar{\nu}_e d$  data, although poorly determined, arises from the ratio of ratios and hence is less sensitive to the assumed reactor neutrino spectrum.

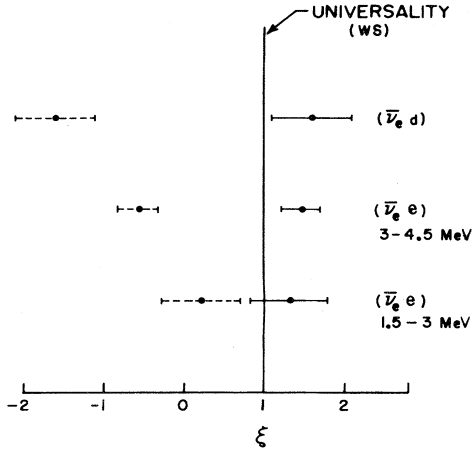


FIG. 2. Values of  $\xi$  determined from the reactor neutrino experiments  $\bar{\nu}_e d$  and  $\bar{\nu}_e e$ . Dashed line denotes a possible value of  $\xi$  arising from the quadratic ambiguity, but from the same experimental input.

#### IV. INFLUENCE ON NEUTRINO OSCILLATIONS IN VACUUM

To illustrate the effect of unequal neutral-current couplings of neutrinos on the experiments which look for neutrino flavor oscillations in vacuum, we

$$d\sigma^{(L)}(\nu_\mu N \rightarrow \nu_\mu + \dots) = [(1-P) + \xi^2 P] d\sigma^{(0)}(\nu_\mu N \rightarrow \nu_\mu + \dots), \quad (22)$$

$$d\sigma^{(L)}(\nu_e N \rightarrow \nu_e + \dots) = [\xi^2(1-P) + P] d\sigma^{(0)}(\nu_e N \rightarrow \nu_e + \dots). \quad (23)$$

Here the probability  $P$  is given by Eq. (21) and the superscripts  $L$  and  $0$  denote the distances traveled by the neutrino between the points of its production and interaction; thus the  $d\sigma^{(0)}$  corresponds to the case when the oscillations are altogether absent. For antineutrinos, we can simply replace  $\nu_e$  by  $\bar{\nu}_e$  and  $\nu_\mu$  by  $\bar{\nu}_\mu$  in the above formulas.

What are the combined effects of unsymmetric neutrino couplings and neutrino oscillations in vacuum on the experimental data of the elastic reaction  $\nu e \rightarrow \nu e$ ? This question is of great experimental interest, for in case the neutrino involved is a  $\nu_e$  or  $\bar{\nu}_e$ , both the charged current interactions (which obey universality) and the neutral-current interactions (whose universality is under study) will be present. We shall list the  $y$  distributions for  $\nu e$  scattering when the neutrino flight path is  $L$ , in terms of the usual  $\nu_\mu e$  and  $\bar{\nu}_\mu e$  cross sections obtained on the basis of no oscillations (superscript zero)

$$\frac{d\sigma^{(L)}}{dy}(\nu_\mu e) = [(1-P) + \xi^2 P] \frac{d\sigma^{(0)}}{dy}(\nu_\mu e) + P\Sigma_1, \quad (24)$$

confine our attention to a simple two-level model. Let  $\nu_1$  and  $\nu_2$  denote the mass eigenstates with masses  $m_1$  and  $m_2$ , and  $\theta$  the mixing angle such that the flavor eigenstates  $\nu_e$  and  $\nu_\mu$  are given by

$$\nu_e = \nu_1 \cos\theta + \nu_2 \sin\theta, \quad (20)$$

$$\nu_\mu = -\nu_1 \sin\theta + \nu_2 \cos\theta.$$

If we start with a specific neutrino beam (say  $\nu_e$ ) of energy  $E$  reaching a detector located at a distance  $L$  along the length of the beam, the "other" type of neutrino ( $\nu_\mu$ ) will be present in the beam with relative intensity<sup>9</sup>

$$P(L/E) = \sin^2(2\theta) \sin^2 \left[ \frac{(m_1^2 - m_2^2)L}{4E} \right]. \quad (21)$$

If neutrino flavor universality ( $\xi=1$ ) is valid, as is generally assumed, there will be no effect of oscillations on the measured neutral-current cross sections. If, on the other hand, the neutral-current couplings of the neutrino are unequal ( $\xi \neq 1$ ), we see that the differential cross sections of neutrino-hadron interaction by weak neutral currents are given in terms of the standard  $\nu_\mu N$  cross sections as

$$\frac{d\sigma^{(L)}}{dy}(\bar{\nu}_\mu e) = [(1-P) + \xi^2 P] \frac{d\sigma^{(0)}}{dy}(\bar{\nu}_\mu e) + P\Sigma_2, \quad (25)$$

$$\begin{aligned} \frac{d\sigma^{(L)}}{dy}(\nu_e e) &= [\xi^2(1-P) + P] \frac{d\sigma^{(0)}}{dy}(\nu_\mu e) \\ &+ (1-P)\Sigma_1, \end{aligned} \quad (26)$$

$$\begin{aligned} \frac{d\sigma^{(L)}}{dy}(\bar{\nu}_e e) &= [\xi^2(1-P) + P] \frac{d\sigma^{(0)}}{dy}(\bar{\nu}_\mu e) \\ &+ (1-P)\Sigma_2, \end{aligned} \quad (27)$$

$$\Sigma_1 \equiv [4(1 + \xi g_V + \xi g_A) - 2\xi(g_V - g_A)my/E] \hat{\sigma}, \quad (28)$$

$$\begin{aligned} \Sigma_2 \equiv & [4(1 + \xi g_V + \xi g_A)(1-y)^2 \\ & - 2\xi(g_V - g_A)my/E] \hat{\sigma}. \end{aligned} \quad (29)$$

Here  $d\sigma^{(0)}(\nu_\mu e)$ , for instance, is the same as  $d\sigma(\nu_\mu e)$  given in Eq. (7), while  $d\sigma^{(0)}(\bar{\nu}_\mu e)$  is obtained from it

TABLE I. Coefficients  $a$  and  $b$  in the cross-section formula  $\sigma(\nu e)/\hat{\sigma} = a + bP(L/E)$ .

$(\sigma/\hat{\sigma})$	$\xi=1$		$\xi=1.6$		$\xi=-1.6$	
	$a$	$b$	$a$	$b$	$a$	$b$
$\nu_\mu e$	0.36	1.84	0.36	1.11	0.36	8.02
$\bar{\nu}_\mu e$	0.31	0.61	0.31	0.66	0.31	2.97
$\nu_e e$	2.20	-1.84	1.47	-1.11	8.38	-8.02
$\bar{\nu}_e e$	0.92	-0.61	0.97	-0.66	3.28	-2.97

by replacing  $g_A$  by  $(-g_A)$ . Note that Eqs. (26) and (27) are obtained from Eqs. (24) and (25), respectively, by replacing  $P$  by  $(1-P)$ , as is to be expected in a two-level oscillation scheme.

To gain some insight into the above expressions, we shall ignore the  $(m/E)$  terms, integrate over  $y$  from 0 to 1, and substitute the Weinberg-Salam-model values for  $g_V$  and  $g_A$  from Eqs. (13) and (14). The resulting total-cross-section formulas can be cast into a form which isolates the effects due to neutrino oscillations,

$$\frac{\sigma^{(L)}(\nu e)}{\hat{\sigma}} = a(\xi) + b(\xi)P(L/E). \quad (30)$$

We have listed in Table I the dimensionless constants  $a$  and  $b$  for three values of  $\xi$  [see Eq. (6)].

Since the  $b$ 's for  $\nu_\mu e$  and  $\bar{\nu}_\mu e$  are positive, the measured cross sections of these reactions will be larger than what they would be in a "no-oscillation" theory; the cross sections for  $\nu_e e$  and  $\bar{\nu}_e e$  are correspondingly depressed.<sup>10</sup> The ratio  $|b/a|$  which controls the relative importance of the oscillation term, is largest for  $\nu_\mu e$  and smallest for the reactor neutrino experiment  $\bar{\nu}_e e$ . As a function of  $\xi$ , we observe that even relatively modest deviations from  $\nu_e$ - $\nu_\mu$  universality can lead to rather large differences between the measured cross sections and those estimated in the standard model—especially for negative values of  $\xi$ . The actual values of cross sections, however, depend on the assumed values of the neutrino-mixing parameter and the masses which enter the expression for  $P$  in Eq. (21); moreover, a comparison of data with the theoretical expectations in any case requires the folding in of the incident neutrino energy spectrum and the ubiquitous  $y$  cuts. We therefore defer such calculations for the present, as we only want to illustrate here the effects of  $\xi \neq 1$  on  $\nu e$  scattering when the incident neutrinos undergo flavor oscillations.

#### V. NONUNIVERSALITY AND MATTER EFFECTS ON NEUTRINO OSCILLATIONS

When a neutrino beam passes through a piece of matter the emerging forward beam includes the neu-

trinos that had been scattered coherently in the forward direction. As a result the oscillation pattern gets affected, as was first emphasized in a fascinating paper by Wolfenstein.<sup>11</sup> This coherent forward scattering of the neutrinos arises in two ways: (i) by the charged-current interaction of the neutrino with the atomic electrons, when the neutrino involved is a  $\nu_e$  or  $\bar{\nu}_e$ , and (ii) by the neutral-current interaction of the neutrino with the  $e$ ,  $p$ , and  $n$  of the medium. Thus while the charged-current interaction selectively affects a  $\nu_e$  or  $\bar{\nu}_e$  involved in the oscillation and modifies the oscillation pattern, the neutral current will not affect the oscillations when the neutrino types involved have equal vector coupling. What we wish to point out is that a violation of the  $\nu_e$ - $\nu_\mu$  universality results only in a minor but important change in the oscillation length, which may lead to interesting experimental consequences.

Focusing our attention once again on the  $\nu_e$ - $\nu_\mu$  oscillations, the time evolution of the two states  $\nu_1$  and  $\nu_2$  will be governed by the Schrödinger equation  $i\partial_t\psi = H\psi$ , where the  $H = H_0 + H_W$  now includes the interaction of neutrinos with the matter through which they are passing;  $H$  is a  $2 \times 2$  matrix with elements

$$H_{12} = \langle \nu_1 | H_0 + H_W | \nu_2 \rangle,$$

etc. If we start initially with a beam of definite type, say  $\nu_e$ , which is a mixture of mass eigenstates with mixing angle  $\theta$  according to Eq. (20), then at a subsequent time the eigenstates will be given by

$$\nu_{1m} = \nu_e \cos\theta_m - \nu_\mu \sin\theta_m, \quad (31)$$

$$\nu_{2m} = \nu_e \sin\theta_m + \nu_\mu \cos\theta_m,$$

$$\tan(2\theta_m) = \frac{\tan(2\theta)}{1 - \beta \sec(2\theta)}, \quad (32)$$

$$\beta \equiv \frac{\langle \nu_\mu | H_W | \nu_\mu \rangle - \langle \nu_e | H_W | \nu_e \rangle}{(m_1^2 - m_2^2)/2E}, \quad (33)$$

where the matrix elements in the last equation refer to elastic scattering of neutrinos on matter. The parameter  $\beta$  does not vanish because the coherent-forward-scattering amplitudes of  $\nu_\mu$  and  $\nu_e$  on the intervening matter are not equal irrespective of the universality question.<sup>12</sup> The relative intensity of the  $\nu_\mu$  component in matter at a distance  $L$  along a  $\nu_e$  beam is now given by<sup>11</sup>

$$P_m(L/E) = (l_m/l_v)^2 \sin^2(2\theta) \sin^2(\pi L/l_m), \quad (34)$$

where  $l_v$  is the conventional oscillation length in vacuum,

$$l_v = \frac{4\pi E}{m_1^2 - m_2^2}, \quad (35)$$

and  $l_m$  is the oscillation length in matter given by

$$l_m = \frac{l_v}{[1 + \beta^2 - 2\beta \cos(2\theta)]^{1/2}}. \quad (36)$$

It is important to realize<sup>13</sup> that the “lengths”  $l_v$  and  $l_m$  can be positive or negative. Although a change of sign of  $l_v$  does not affect the intensity formula Eq. (21), it does affect the corresponding expression Eq. (34) in matter because  $\beta$  depends on  $l_v$ .

Assuming  $\nu_\mu$ - $\nu_e$  universality ( $\xi=1$ ), the effect of matter arises only through the charged-current interaction of  $\nu_e$  with the  $e$ 's. In this case it can be shown that<sup>11,13,14</sup>

$$\beta = l_v / l_W, \quad (37)$$

where we define the Wolfenstein length  $l_W$  by

$$l_W \equiv \frac{\sqrt{2}\pi}{GN_e} \simeq \left[ \frac{N_A}{N_e} \right] 1.63 \times 10^9 \text{ cm}. \quad (38)$$

Here  $N_e$  denotes the number of electrons per unit volume of the target medium and  $N_A$  is the Avogadro number. Note that  $l_W$  does not depend on the energy  $E$ , in contrast to  $l_v$ . At small enough energies, when  $l_v \ll l_W$ , we have  $l_m \simeq l_v$  and the matter effects can be ignored. In the opposite extreme when  $l_v \gg l_W$ , we have  $l_m \simeq l_W$ , but the intensity becomes small due to Eq. (34); in fact,  $l_W$  then serves as a “cutoff” on  $l_m$  (since  $l_m$  no longer increases with  $E$  as  $l_v$  does) and the oscillations through matter would not be characterized by the difference ( $m_1^2 - m_2^2$ ).

With the nonuniversal neutral-current couplings the modification to be introduced is quite simple—instead of  $l_W$  being given by Eq. (38), it will be modified to  $l'_W$ , which is given by

$$\frac{\sqrt{2}\pi}{l'_W} = GN_e + (G_e N_e + G_p N_p + G_n N_n) - (G'_e N_e + G'_p N_p + G'_n N_n), \quad (39)$$

where  $G_x$  is the neutral-current vector coupling of  $\nu_e$  to the target particle  $x$  ( $=p, n, e$ ),  $G'_x$  refers to the corresponding vector couplings of the  $\nu_\mu$ , and  $N_x$  denotes the target particles of type  $x$  per unit volume in the medium. From Eq. (1) we see that  $G_e = G\xi g_V$ , etc. For a medium (such as the terrestrial medium) in which  $N_e = N_p = N_n$ , we obtain

$$\frac{\sqrt{2}\pi}{l'_W} = GN_e [1 + (\xi - 1)(g_V + g_V^p + g_V^n)]. \quad (40)$$

In the Weinberg-Salam model the sum of the vector-current couplings of  $\nu_\mu$  with  $e$ ,  $p$ , and  $n$  is  $g_V + g_V^p + g_V^n = -\frac{1}{2}$ , and this gives

$$l'_W = \frac{2\sqrt{2}\pi}{GN_e(3-\xi)}. \quad (41)$$

Thus the deviation from strict  $\nu_e$ - $\nu_\mu$  universality changes the Wolfenstein length  $l_W$  of Eq. (38) by the simple factor  $2/(3-\xi)$ .<sup>15</sup>

The oscillation length in matter is now given by

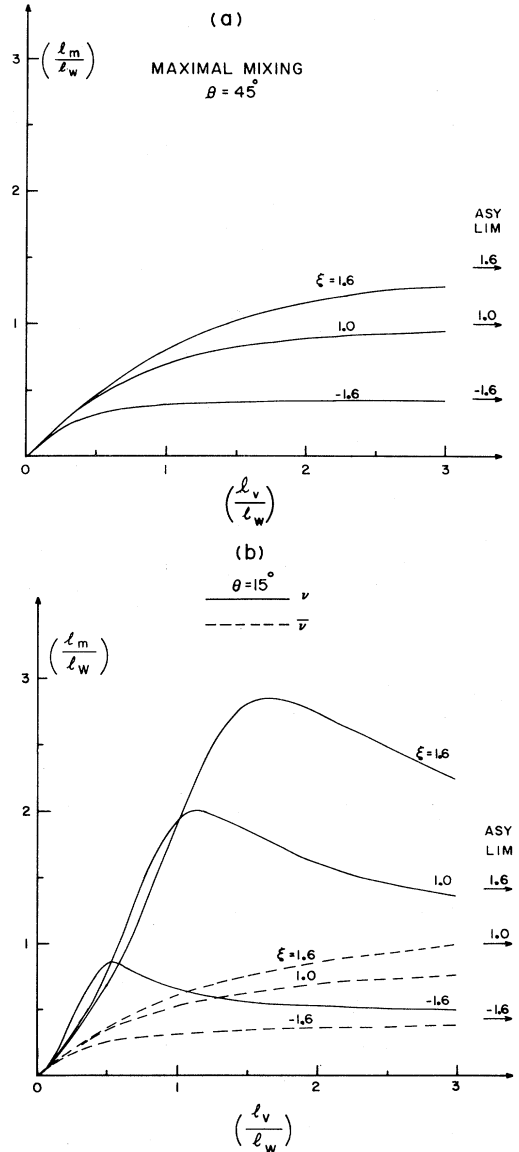


FIG. 3. Variation of the matter oscillation length  $l_m$  with respect to the vacuum oscillation length  $l_v$  (both measured in units of  $l_W$ ) for some values of the nonuniversality parameter  $\xi$ . The arrows denote the asymptotic limits of the curves at infinite neutrino energy. The neutrino mixing angle in vacuum  $\theta$  is taken to be (a)  $\theta=45^\circ$ , for which case the curves are identical for  $\nu$  and  $\bar{\nu}$ ; (b)  $\theta=15^\circ$ . In the latter case if the solid curve refers to  $\nu$  the dashed curve would refer to  $\bar{\nu}$  (or vice versa); the curves have a common asymptotic limit.

$$\frac{l_m}{l_W} = \frac{l_\nu}{l_W} \left[ 1 + \frac{1}{4} \left( \frac{l_\nu}{l_W} \right)^2 (3 - \xi)^2 - \frac{l_\nu}{l_W} (3 - \xi) \cos(2\theta) \right]^{-1/2}. \quad (42)$$

Instead of considering  $l_m$  as a function of the incident neutrino energy  $E$ , it is convenient to look at it as a function of  $E/(m_1^2 - m_2^2)$  which is essentially  $l_\nu$ . In Fig. 3 we plotted  $(l_m/l_W)$  as a function of  $(l_\nu/l_W)$  for some typical values of  $\xi$  ( $= -1.6, 1.0,$  and  $1.6$ ). Since the only difference between the cases of  $\nu$  and  $\bar{\nu}$  amounts to changing the sign in front of the  $\cos(2\theta)$  term in Eq. (42), the curves for  $\nu$  and  $\bar{\nu}$  are the same for maximal mixing,  $\theta = \pi/4$ , shown in Fig. 3(a). For arbitrary mixing angles there will be a difference between the matter oscillation lengths of  $\nu$  and  $\bar{\nu}$ . This difference is depicted as a function of  $l_\nu$  in Fig. 3(b) for the case of  $\theta = 15^\circ$ . In this case we see that the lengths  $l_m(\nu)$  and  $l_m(\bar{\nu})$  could be different by more than a factor 3, but it would not be possible to decide which is the larger of the two since the intrinsic sign of  $[(3 - \xi)l_\nu \cos(2\theta)]$  is *a priori* unknown.

Negative values of  $\xi$  imply constructive interference between the two types of weak interactions and lead to smaller matter oscillation lengths than positive values of  $\xi$ , at sufficiently large energies of the  $\nu$  or  $\bar{\nu}$ . In this context, let us consider the following interesting, but somewhat extreme possibility. Suppose  $\nu_e$ 's undergo flavor oscillations but  $l_\nu \gg D_e$ , where  $D_e = 1.3 \times 10^9$  cm is the earth's diameter. Then vacuum oscillations cannot be observed in a terrestrial experiment of the type proposed by Mann and Primakoff.<sup>16</sup> However, can one detect the effects due to matter in such a case? Assuming a uniformly dense medium for the earth, we substitute  $N_e \simeq 2N_A$  in Eq. (38) and obtain  $l_W \simeq 0.6D_e$ . From Eq. (42), since  $l_m \simeq 2l_W/(3 - \xi)$ , negative values of  $\xi$  would be favorable for the observation of matter effects. For instance,  $\xi \simeq -2$  (allowed by Fig. 2) leads to  $l_m \simeq 0.25D_e$  and the effects due to matter may be barely observable.

## VI. COMMENTS AND SUMMARY

While there is considerable experimental evidence for the validity of the  $e$ - $\mu$  universality in the charged-current sector, it is still an open question whether the neutral-current couplings of  $\nu_\mu$  and  $\nu_e$  are in fact identical. It is possible that the neutral-current interaction of the neutrinos with matter is not universal, and if such a possibility is realized experimentally it would provide an interesting clue to the vexing problem of generations. As a first step,

therefore, it is worthwhile examining, purely from a phenomenological angle, how a simple deviation from the  $\nu_e$ - $\nu_\mu$  universality, as for instance the one characterized by the parameter  $\xi$  in Eq. (1), could affect the various experimental measurements.

As an example of a rather remote but interesting ramification due to the unequal neutral-current strengths of  $\nu_e$  and  $\nu_\mu$ , we shall briefly digress to consider the cosmic neutrino background. A suggestion originally due to Opher,<sup>17</sup> and subsequently elucidated by Lewis,<sup>14</sup> to detect the cosmic neutrino background ( $E \sim 10^{-2}$  eV) is to make use of the coherent scattering of neutrinos from the entire target of the detector system. According to this proposal, one measures the force exerted on a collector plate when neutrinos get totally reflected from its surface. If the collector plate has nuclei with atomic and mass numbers  $Z$  and  $A$ , the refractive indices will be given by

$$n(\nu_e) = 1 + \frac{GN_e}{\sqrt{2}E} [2Z - (A - Z)\xi], \quad (43)$$

$$n(\nu_\mu) = 1 - \frac{GN_e}{\sqrt{2}E} [A - Z], \quad (44)$$

the expressions for the corresponding antineutrinos being obtained by changing the sign in front of  $G$ . The sign of  $(n - 1)$  has important implications for the design of the Opher-Lewis detector. For  $\xi \gtrsim 2$ , the refractive index of  $\nu_e$  scattering on most materials of interest (Cu, Fe, . . .) is less than unity and thus cannot easily be discriminated from the case of  $\nu_\mu$ ; for  $\xi \lesssim 2$  (which includes the case of universality), however, the  $\nu_e$  exhibits the phenomena of total internal reflection ( $n > 1$ ), as the  $\bar{\nu}_\mu$ .

As for the current experimental information on  $\xi$ , the only available data are from the reactor experiments of the Irvine group. The information extracted from these data is displayed in Fig. 2, which may be summarized by noting the two *alternative* values of  $\xi$  (ensuing mainly from the  $\bar{\nu}_e e$  data):  $\xi = 1.45 \pm 0.20$ , the weighted average of the three values of  $\xi$  which are greater than unity, and  $\xi = -0.6 \pm 0.2$ , the weighted average of the three values less than unity. The first is about two standard deviations from the universality value  $\xi = 1$ . The errors on these values however should be much larger than what are quoted, to account for the many uncertainties, including that due to the assumed spectrum of reactor antineutrinos.

It should be mentioned that we have not considered here the possible violations of  $e$ - $\mu$  universality in neutral currents. From a phenomenological viewpoint, the  $e$ - $\mu$  universality need not be related to the  $\nu_e$ - $\nu_\mu$  universality discussed in this paper. For this purpose one may have to consider the reactions

$\nu_\mu$  and  $\nu_e$  scattering on  $\mu$ , and  $e^+e^- \rightarrow (e^+e^-, \mu^+\mu^-)$ .

In summary, starting with the reasonable assumption that the structure of the neutral-current interaction of  $\nu_e$  is the same as that established from the  $\nu_\mu$  reactions with matter, we have assumed the  $\nu_e$ - $\nu_\mu$  nonuniversality to be of a factorizable type characterized by the strength parameter  $\xi$ , and studied the consequences of it. Future scattering data using the  $\nu_e$  or  $\bar{\nu}_e$  beams of known fluxes and purity (as for example at the meson factories) on nuclear targets will be valuable in checking whether  $\xi=1$ . Secondly, even significant departures from universality, when  $\xi$  lies in the range 0.5–2, will only cause minor changes in  $\sigma(\bar{\nu}_e e)$  as in Fig. 1; in contrast, the corresponding changes in  $\sigma(\nu_e e)$  are appreciable.

Thirdly, the influence of (possible)  $\nu_e$ - $\nu_\mu$  oscillations on the data obtained on  $\sigma(\nu e)$  could be considerably modified by nonuniversality, as exemplified in Table I; here negative values of  $\xi$  could lead to spectacular increases in the measured  $\sigma(\nu_\mu e)$ , and to a less extent in  $\sigma(\bar{\nu}_\mu e)$ , as a result of oscillations. Lastly, the elusive matter effects on oscillations, in the context of experiments which seek to detect GeV neutrinos passing through the earth, may become observable provided sizable departures from universality ( $\xi \lesssim -3$ ) exist.

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<sup>1</sup>See the review articles, J. E. Kim, P. Langacker, M. Levine, and H. H. Williams, *Rev. Mod. Phys.* **53**, 211 (1981); P. Q. Hung and J. J. Sakurai, *Annu. Rev. Nucl. Part. Sci.* **31**, 375 (1981), for a comprehensive survey of the data and citation of the original references.

<sup>2</sup>F. Reines, H. W. Sobel, and E. Pasierb, *Phys. Rev. Lett.* **45**, 1307 (1980).

<sup>3</sup>L. M. Sehgal, *Phys. Lett.* **48B**, 60 (1974); *Nucl. Phys.* **70B**, 61 (1974).

<sup>4</sup>M. Gourdin, in *Proceedings of the 1976 International Neutrino Conference, Aachen*, edited by H. Faissner, H. Reithler, and P. Zerwas (Vieweg, Braunschweig, West Germany, 1977), p. 234.

<sup>5</sup>G. V. Dass and P. Ram Babu, *Ann. Phys. (N.Y.)* (to be published).

<sup>6</sup>B. Kayser, E. Fischbach, S. P. Rosen, and H. Spivack, *Phys. Rev. D* **20**, 87 (1979).

<sup>7</sup>F. Reines, H. S. Gurr, and H. W. Sobel, *Phys. Rev. Lett.* **37**, 315 (1976).

<sup>8</sup>F. T. Avignone III and Z. D. Greenwood, *Phys. Rev. D* **16**, 2383 (1977).

<sup>9</sup>See, for instance, S. M. Bilenky and B. Pontecorvo, *Phys. Rep.* **41C**, 225 (1978).

<sup>10</sup>B. V. Martemyanov, M. Yu. Khlopov, and M. G. Shchepkin, *Pis'ma Zh. Eksp. Teor. Fiz.* **32**, 484 (1980) [*JETP Lett.* **32**, 464 (1980)]; S. P. Rosen and B. Kayser, *Phys. Rev. D* **23**, 669 (1981).

<sup>11</sup>L. Wolfenstein, *Phys. Rev. D* **17**, 2369 (1978).

<sup>12</sup>For antineutrino, the coherent-forward-scattering amplitude is simply the negative of that for the corresponding neutrino; hence the value of  $\beta$  for  $\bar{\nu}$  will be equal and opposite to that for  $\nu$ .

<sup>13</sup>V. Barger, K. Whisnant, S. Pakvasa, and R. J. N. Phillips, *Phys. Rev. D* **22**, 2718 (1980).

<sup>14</sup>R. R. Lewis, *Phys. Rev. D* **21**, 663 (1980).

<sup>15</sup>An amusing possibility is obtained when  $\xi=3$ . In this case the charged-current effects will exactly cancel the neutral-current effects and  $\beta=0$ ; the medium will have no effect on the neutrino oscillation pattern irrespective of the matter density. Then  $l_m=l_\nu$  and  $\theta_m=\theta$ ; matter behaves as if it were vacuum, in so far as neutrino oscillations are concerned.

<sup>16</sup>A. K. Mann and H. Primakoff, *Phys. Rev. D* **15**, 655 (1977).

<sup>17</sup>R. Opher, *Astron. Astrophys.* **37**, 135 (1974).