# Spectra of $\tau$ -leptonic atoms

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The low-lying atomic spectra are calculated nonrelativistically for a  $\tau$  lepton interacting Coulombically with various finite-sized nuclei. The proton is taken as a point charge in these calculations. The nonspherical charge distribution of the deuteron nucleus is included; its effects are small but not negligible. The magnetic-moment interaction is calculated in perturbation theory and is negligible. The transition lifetimes are small compared to the  $\tau$ -lepton lifetime. An experimental measurement of the  $\gamma$ -ray energies emitted by the  $\tau$ -leptonic atom while cascading to the ground state would lead to an independent measure for the  $\tau$ -lepton mass.

#### INTRODUCTION

 $\tau$ -leptonic atoms consist of a nucleus and a  $\tau$  lepton. The  $\tau$  lepton is here assumed to have a mass of twice the proton mass, consistent with the experimental value,<sup>1</sup> and we calculate the atomic levels of this system treating the  $\tau$  lepton as stable. Its experimental lifetime is  $4.6 \times 10^{-13}$  sec.<sup>2</sup> The Bohr radius of such a system is 28 fm or smaller, depending on the nucleus, so we ignore the presence of any electrons.  $\tau$ -leptonic hydrogen is treated as though the proton were a point charge.

 $\tau$ -leptonic deuterium is an intrinsically interesting three-body system to study. The  $\tau$  is supposed to be just like an electron, but heavy, with a mass of 1.8 GeV, and also unstable, with a lifetime of  $4.6 \times 10^{-13}$  sec. The three-body (bound) system of proton, neutron, and  $\tau$ should be well described nonrelativistically. Further, to some high degree of approximation, the neutron and the  $\tau$ do not interact with each other. The assumed Coulombic interaction betweeen the  $\tau$  and the proton and the presumed known neutron-proton force result in a possibly tractable three-body problem. However, here we neglect the presence of the  $\tau$  lepton in determining the neutronproton relative wave function, and instead determine the  $\tau$ -lepton atomic wave function with a given charge distribution for the nucleus. This approximation is least accurate for the ground state of the system, where the  $\tau$  is closest to the nucleus, but is still a reasonable approximation for that case. For other than s waves, the spectrum calculated is almost identical to that of a point nucleus charge. The nonspherical deuteron charge distribution splits the  $2p_{3/2}$  states from the  $2p_{1/2}$  state. This effect is larger than the relativistic point-charge splitting of the Dirac equation for the  $2p_{3/2}$  from the  $2p_{1/2}$  by three orders of magnitude. For s waves, the bound states have less binding energy than for a point nuclear charge, due to the assumed nuclear charge distribution. The Bohr radius of the  $\tau$  deuteron system is 28.8 fm, which is about 10 times the nuclear radius.<sup>3</sup> The nuclear binding energy of 2.226 MeV is about 100 times larger than the atomic binding energy of order 20 keV. A comparison of the order of magnitude of these sizes and energies tends to justify the neglect of the  $\tau$  lepton on the nuclear charge distribution.

For the  $A \ge 3$  nuclei, the nuclear charge distribution is assumed to be a uniform sphere of radius R, where R was determined from the rms radius from electron scattering.<sup>4</sup> Of the nuclei considered, only <sup>6</sup>Li has a spin as large as one  $\hbar$ , and as a consequence, a possible quadrupole moment, but its experimental quadrupole moment<sup>5</sup> is only  $-0.1 \text{ fm}^2$ , which we neglect. The largest nuclei considered is <sup>24</sup>Mg as the Bohr radius is exceeded by the size of this nucleus.

It is the purpose of this paper to ascertain the  $\gamma$ -ray energies that would result from a  $\tau$  lepton cascading down the hydrogenic levels in a  $\tau$ -leptonic atom. Estimates of the lifetime for these transitions are provided as well. The slowing-down time must be small compared to the  $\tau$ -lepton lifetime if the transitions to the ground state can be expected to be detected experimentally.

The slowing-down time has been calculated by Fermi and Teller<sup>6</sup> for muons in carbon or iron, and by Wightman<sup>7</sup> for muons in liquid hydrogen. The experimental problems of slowing down and capturing muons have been discussed by Devons and Duerdoth.<sup>8</sup> The slowing-down time for muons was calculated as  $10^{-13}$  to  $10^{-9}$  sec, depending on the initial muon energy assumed or moderator used, but in both cases,<sup>6,7</sup> the slowing-down time was small compared to the muon lifetime of about  $2 \times 10^{-6}$ sec. Both the Wightman and the Fermi-Teller calculations can be scaled via the mass to apply to the slowing down of  $\tau$  leptons. The slowing-down time for a 20-MeV au lepton in liquid hydrogen is found to be on the order of  $2 \times 10^{-10}$  sec, which is long compared to the  $\tau$  lifetime. Thus one must use  $\tau$  leptons produced nearly at rest in the laboratory frame and use a moderator of larger Z. The  $\tau$ slowing-down time from 20 keV until cascading into the ground state is estimated as  $9 \times 10^{-13}$  sec for a carbon moderator. This slowing-down time is still twice the  $\tau$ lifetime. Thus the 2P-to-1S  $\gamma$  yield would be reduced by a

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factor of order 10 compared to the number of  $\tau$  leptons produced.

## THEORY

For point charges the Dirac-equation energies are shown<sup>9</sup> in Table I. In Table I, *m* is the reduced mass of the atom, and is equal to  $2m_p/3$  for a proton nucleus. For a proton nucleus, the  $2P_{3/2}-2P_{1/2}$  energy splitting is 0.0554 eV. For a deuteron nucleus, this relativistic splitting is 0.0831 eV, for a point nuclear charge distribution.

For the deuteron nucleus, we treat the neutron and proton masses as equal. We introduce relative coordinates rfor the neutron-proton separation and x, and  $\tau$ lepton-deuteron-center-of-mass relative separation (see Fig. 1). We take the neutron and proton masses as equal. Then the  $\tau$ -lepton-proton separation y can be written as

$$y = |\vec{x} + \vec{r}/2|$$
 (1)

The nonrelativistic Schrödinger equation for the system can be written as

$$\left[\frac{-\hbar^2}{2M_r}\nabla_x^2 - \frac{\hbar^2}{m_p}\nabla_r^2 - \frac{e^2}{y} + V(\vec{r}) + B\right]\Psi(\vec{x}, \vec{y}) = 0, \quad (2)$$

where B is the positive binding energy for the three-body system,  $M_r$  is the reduced mass of the nucleus and the  $\tau$ lepton, and  $m_p$  is the proton mass. Neglecting the effect of the  $\tau$  on the nuclear charge distribution, we can write the wave function  $\Psi$  as

$$\Psi(\vec{\mathbf{x}}, \vec{\mathbf{y}}) \cong \sum_{mM_d} \psi_1^{M_d}(\vec{\mathbf{r}}) \frac{\chi_{lj}(\mathbf{x})}{\mathbf{x}} \mathscr{Y}_{l(1/2)j}^m(\hat{\mathbf{x}}) \begin{vmatrix} F & M_F \\ 1 & M_d \\ j & m \end{vmatrix}, \qquad (3)$$

where

$$\psi_1^{M_d}(\vec{\mathbf{r}}) = \sum_{l_d=0}^2 \phi_{l_d}(r) \mathscr{Y}_{l_d 1 1}^{M_d}(\hat{r})$$
(4)

and

$$\left[\frac{-\vec{n}^2}{m_p} \nabla_r^2 + V(\vec{r}) + B_0\right] \psi_1^{M_d}(\vec{r}) = 0.$$
 (5)

In Eq. (3), the quantity in square brackets is a Clebsch-Gordan coefficient and the  $\mathscr{Y}$  is an angular momentum eigenfunction with z component m.  $M_d$  is the z component of angular momentum for the deuteron nuclear wave function. The deuteron spin 1 is coupled to the lepton angular momenta j to form the atomic angular momenta F.

The sum over  $l_d$  in Eq. (4) is necessary as the neutronproton potential V(r) is not spherically symmetric. The radial eigenfunctions for each orbital state ( $l_d=0$  and  $l_d=2$ ) are labeled  $\phi_{l_d}(r)$ . The nuclear binding energy is  $B_0=2.226$  MeV.

We substitute Eq. (3) into Eq. (2), and use Eq. (5) to eliminate the nuclear potential and the nuclear Laplacian in favor of the nuclear binding energy  $B_0$ . Multiplying by  $\psi_1^{*\overline{M}}$  and integrating we have

$$0 = \int d\vec{\mathbf{r}} \,\psi_1^* \overline{M}(\vec{\mathbf{r}}) \left[ B - B_0 - \frac{\hbar^2}{2M_r} \nabla_x^2 - \frac{Ze^2}{y} \right] \\ \times \sum_{Mm} \psi_1^M(\vec{\mathbf{r}}) \frac{\chi_{ij}(x)}{x} \mathcal{Y}_{l(1/2)j}^m(\hat{x}) \begin{bmatrix} F & M_F \\ 1 & M \\ j & m \end{bmatrix}.$$
(6)

Using Eq. (1) in Eq. (6), and expanding in spherical harmonics, we obtain the following equation for  $\chi_{li}(x)$ :

$$0 = \left| B - B_0 - \frac{\hbar^2}{2M_r} \left[ \frac{d^2}{dx^2} - \frac{l(l+1)}{x^2} \right] - Ze^2 \langle V \rangle \right| \chi_{lj}(x) , \qquad (7)$$

where

$$\langle V \rangle = \sum_{\lambda l_d \overline{l}_d m's} \begin{bmatrix} 1 & \overline{M}_d \\ 1 & m_2 \\ \overline{l}_d & m_{\overline{l}_d} \end{bmatrix} \begin{bmatrix} 1 & M_d \\ 1 & m_2 \\ l_d & m_{l_d} \end{bmatrix} \begin{bmatrix} \overline{l}_d & 0 \\ l_d & 0 \\ \lambda & 0 \end{bmatrix} \begin{bmatrix} \overline{l}_d & m_{\overline{l}_d} \\ l_d & m_{l_d} \end{bmatrix} \begin{bmatrix} l & m_l \\ \overline{l} & \overline{m}_l \\ \lambda & \lambda \end{bmatrix} \\ \times \begin{bmatrix} j & \overline{m} \\ \frac{1}{2} & m_s \\ l & \overline{m}_l \end{bmatrix} \begin{bmatrix} j & m \\ \frac{1}{2} & m_s \\ l & m_l \end{bmatrix} \begin{bmatrix} F & M_F \\ j & \overline{m} \\ 1 & \overline{M}_d \end{bmatrix} \begin{bmatrix} F & M_F \\ j & m \\ 1 & \overline{M}_d \end{bmatrix} \begin{bmatrix} 2l_d + 1 \\ 2\overline{l}_d + 1 \end{bmatrix}^{1/2} \int_0^\infty r^2 dr \, \phi_{\overline{l}_d}(r) \phi_{l_d}(r) \frac{r_{<\lambda}}{r_{>}^{\lambda+1}} \,.$$

$$(8)$$

Here  $r_{<}$   $(r_{>})$  is the lesser (greater) of x and r/2, coming from the expansion of 1/y. For the case l=0,  $\langle V \rangle$  simplifies to

$$\langle V \rangle = \sum_{l_d=0}^{2} \int_{0}^{\infty} \frac{r^2 dr \, \phi_{l_d}{}^2(r)}{r_{>}} \,.$$
 (9)

In Eq. (7), the difference  $B - B_0$  is the atomic binding energy. This atomic binding energy varies for different

states because Eq. (8) for the potential  $\langle V \rangle$  depends on F, l, and j, in addition to the angular momentum barrier appearing in Eq. (7) of  $l(l+1)\hbar^2/2Mx^2$ .  $M_r$  is the reduced mass of the  $\tau$ -lepton—nucleus system in Eq. (7). We assume an analytic Hulthén form for the deuteron wave function

$$\phi_0 = N \cos w \left( e^{-\alpha r} - e^{-\beta r} \right) / r , \qquad (10)$$
  
$$\phi_2 = N \sin w \left( e^{-\alpha r} - e^{-\beta r} \right) / r ,$$

TABLE I. Point Coulombic energy levels.

Level	Dirac energy	Nonrelativistic binding energy
1S <sub>1/2</sub>	$m(1-Z^2\alpha^2)^{1/2}$	$mZ^2\alpha^2/2$
2 <i>S</i> <sub>1/2</sub>	$m\left(\frac{1\!+\!(1\!-\!Z^2\alpha^2)^{1/2}}{2}\right)^{1/2}$	$mZ^2\alpha^{2}(\frac{1}{8}+5Z^2\alpha^2/128)$
2 <i>P</i> <sub>1/2</sub>	$m\left[\frac{1\!+\!(1\!-\!Z^2\alpha^2)^{1/2}}{2}\right]^{1/2}$	$mZ^2\alpha^{2}(\frac{1}{8}+5Z^2\alpha^2/128)$
2P <sub>3/2</sub>	$\frac{m}{2}(4-Z^2\alpha^2)^{1/2}$	$mZ^2\alpha^2(\frac{1}{8}+Z^2\alpha^2/128)$

where  $\alpha$  reproduces the deuteron binding energy,  $\beta = 7\alpha$ , and  $\sin w = 0.15$ . These parameters will reproduce the deuteron quadrupole moment, and the analytic wave function chosen permits the integrals in Eq. (9) to be evaluated analytically.

The summation over m values in Eq. (8) is over all values save  $M_F$ , where the limits of the m sums are from the respective minus L value to the plus L value. The summations over m in Eq. (8) can be reduced to a sum over 6-j symbols; the general result is not given here. The quantities in the square brackets in Eq. (8) are Clebsch-Gordan coefficients.

For nuclei with  $A \ge 3$ , we take

$$V = -Ze^{2}/r \text{ if } r > R ,$$
  
and  
$$V = \frac{-Ze^{2}}{2R} (3 - r^{2}/R^{2}) \text{ if } r < R .$$
 (11)

The parameter R is given in Table II for the various nuclei considered. We ignore any possible nonspherical charge distributions, and solve Eq. (7) for various l values. The difference  $B - B_0$  is the atomic binding energy, as before. It is always much smaller, 1% or less, than the nuclear binding energy, tending to justify the approximation of the  $\tau$ -lepton presence having no effect on the nuclear charge distributions.

# RESULTS

The spectra have been calculated for the levels corresponding to the hydrogenic 1S, 2P, 3D levels (see Table II).



FIG. 1. Geometry for  $\tau$ -leptonic deuterium. The nuclear charge distribution is somewhat cigar shaped.

For comparison, the levels for a point nucleus charge are shown. Only the S levels are noticeably shifted upwards in energy compared to the point nucleus results. For carbon and magnesium, the 2P levels are also shifted upwards from that of a point charge. The Bohr radius scales as  $n^2/mZ^2$ , so even the l=1 levels have Bohr radii comparable to the nuclear size for the carbon and magnesium nuclei. The 3D energies are all nearly the same as if the nucleus were a point charge.

The effect of the nonspherical charge distribution of the deuteron nucleus is shown in Fig. 2. Because of the deuteron quadrupole moment, the spectra associated with the  $2P_{3/2}$  leptonic state splits, depending on  $F = \vec{j} + \vec{l}$ . Energy splittings of 56, 14, and 70 eV from the  $2P_{1/2}$  state are seen for the  $F = \frac{3}{2}, \frac{5}{2}$ , and  $\frac{1}{2}$  states, respectively.

Including only the Coulomb interactions, all the  $j = \frac{1}{2}$  levels are degenerate since  $F = \frac{1}{2}$  or  $\frac{3}{2}$  both have the same energy. Including the fine structure from the magnetic field of the deuteron interacting with the assumed Dirac magnetic moment of the  $\tau$  lepton removes this degeneracy. The  $F = \frac{1}{2}$  state is lowest, shifted downward 0.669 eV in a perturbative calculation,<sup>10</sup> while the  $F = \frac{3}{2}$  level is shifted upwards by 0.515 eV. We now discuss transition life-times.

Using classical field theory, the probability<sup>11</sup> for spontaneous emission due to the dipole emission is

$$P = 4e^2 W^3 \langle r \rangle^2 / 3\hbar c^3 , \qquad (12)$$

where  $\langle r \rangle$  is the transition dipole moment and W is the transition energy. For the 2P-to-1S transition this leads to

Nucleus	Point-charge results		Nuclear	Extended-charge results		
	$E_{1S}$	$E_{2P}$	charge radius	$E_{1S}$	$E_{2P_{1/2}}$	$E_{3D}$
<sup>1</sup> <sub>1</sub> H	16.646	4.161	0	16.646	4.161	1.749
$^{2}_{1}\mathbf{H}$	24.969	6.242		22.592	6.238	2.774
<sup>4</sup> <sub>2</sub> He	133.168	33.292	2.08	122.051	33.300	14.796
<sup>8</sup> Li	337.082	84.270	3.59	279.915	83.974	37.453
<sup>7</sup> <sub>3</sub> Li	349.566	87.391	3.50	290.223	87.103	38.840
<sup>9</sup> ₄Be	653.733	163.433	3.92	495.604	162.114	72.637
${}^{12}_{6}C$	1540.944	385.236	3.04	1107.807	382.447	171.216
<sup>24</sup> <sub>12</sub> Mg	6637.914	1659.478	3.84	3089.414	1534.069	737.

TABLE II.  $\tau$ -leptonic energy levels for various nuclei.





a transition lifetime of  $4.8 \times 10^{-14}$  sec in deuterium. Assuming the wave functions are Coulombic bound-state wave functions, the dipole transition matrix element is proportional to 1/MZ, where M is the reduced mass and Z is the nuclear charge. The energy of the transition varies as  $MZ^2$  so the probability of the transition due to spontaneous emission varies as  $MZ^4$ . This dependence certainly breaks down when the Bohr radius of the atom approaches the nuclear radius. This happens for a nuclear atomic weight of about 9 or 10 for atomic ground states.

TABLE III.  $\tau$ -leptonic atom transition energies and lifetimes.

Nucleus	$E_{2P-1S}$ (keV)	$E_{2D-2P}$ (keV)	$t_{2P-1S}$ (sec)
H	12.485	2.412	5.05×10 <sup>-14</sup>
$\frac{2}{1}$ <b>H</b>	16.354	3.464	4.80×10 <sup>-14</sup>
<sup>4</sup> <sub>2</sub> He	88.751	18.504	$2.15 \times 10^{-15}$
<sup>6</sup> <sub>3</sub> Li	195.941	46.521	5.69×10 <sup>-16</sup>
<sup>7</sup> <sub>3</sub> Li	203.120	48.263	5.49×10 <sup>-16</sup>
<sup>9</sup> <sub>4</sub> Be	333.490	89.477	$2.40 \times 10^{-16}$
6 <sup>12</sup> C	725.360	211.2	5.86×10 <sup>-17</sup>
<sup>24</sup> <sub>12</sub> Mg	1555.345	796.0	$2.75 \times 10^{-17}$

In Table III is shown the energy for the 2P-to-1S and for the 3D-to-2P transition, as well as the calculated transition lifetime for dipole radiation. These lifetimes are all small compared to the  $4.6 \times 10^{-13}$  sec measured<sup>2</sup> for the  $\tau$ -lepton lifetime. Providing the  $\tau$  could be captured directly into a low-lying atomic orbit, the measurement of the  $\gamma$  energy emitted as the orbit decays to the 1s ground state will provide an independent measure of the  $\tau$ -lepton mass. The transition energies are proportional to the reduced mass of the  $\tau$  lepton and the nucleus for a point charge distribution. The above transition energies are based on a  $\tau$ -lepton mass of twice the proton mass and on extended charge distributions. The slowing-down times for  $\tau$  leptons, as well as the cascade time must be less than or comparable to the  $\tau$  lifetime before experimental detection of the transition  $\gamma$ 's would be feasible.

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- <sup>1</sup>M. L. Perl et al., Phys. Lett. <u>70B</u>, 487 (1977).
- <sup>2</sup>G. J. Feldman, Phys. Rev. Lett. <u>48</u>, 66 (1982).
- <sup>3</sup>R. R. Roy and B. P. Nigam, *Nuclear Physics* (Wiley, New York, 1967).
- <sup>4</sup>M. A. Preston, *Physics of the Nucleus* (Addison-Wesley, Reading, Mass., 1962).
- <sup>5</sup>Harold A. Enge, *Nuclear Physics* (Addison-Wesley, Reading, Mass., 1966).
- <sup>6</sup>E. Fermi and E. Teller, Phys. Rev. <u>72</u>, 399 (1947).

- <sup>7</sup>A. S. Wightman, Phys. Rev. <u>77</u>, 521 (1949).
- <sup>8</sup>S. Devons and I. Duerdoth, Adv. Nucl. Phys. <u>2</u>, 295 (1969).
- <sup>9</sup>J. D. Bjorken and S. D. Drell, *Relativistic Quantum Mechanics* (McGraw-Hill, New York, 1964), p. 56.
- <sup>10</sup>J. D. Bjorken and S. D. Drell, *Relativistic Quantum Mechanics* (Ref. 9), p. 58.
- <sup>11</sup>L. I. Schiff, *Quantum Mechanics* (McGraw-Hill, New York, 1955), p. 261.