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Bounds on $e^+e^- \rightarrow l^*\bar{l}$ and $lp \rightarrow l^*X$ (l^* = excited lepton), and prospects for visible l^* tracks in cosmic-ray emulsion events

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The small limits of discrepancy between experimental and theoretical values of $(g - 2)_{e,u}$ are shown to eliminate any chance of seeing single excited e*'s or µ*'s in colliding-beam or inelastic electro- or muoproduction experiments. On the other hand, visible l^* tracks in very-high-energy ($E > 10^{19}$ eV) cosmic-ray events are possible for $m_i^* \sim 1$ TeV. The results presuppose a new dynamics, disjoint from QED, underlying the l^* excitations.

The close agreement between the experimental values of the anomalous magnetic moments (a_e, a_μ) of the electron¹ and muon,² and the best theoretical estimates³ using "known" dynamics places severe constraints^{4,5} on the mass scales associated with any subconstituents of the electron and muon.⁶ As shown in Refs. 4 and 5, a variety of considerations lead to the result

$$\left|\delta a_{l}^{\text{non-QED}}\right| \sim O(m_{l}/M_{\text{const}}), \qquad (1)$$

where $\delta a_1^{\text{non-QED}}(l=e,\mu)$ is the contribution to a_1 $\equiv \frac{1}{2}(g_1 - 2)$ from a dynamical structure associated with a constituent of mass M_{const} , in the limit α_{EM} $\rightarrow 0$. The bounds placed on $\delta a_i^{\text{non-QED}}$ by the combined theoretical and experimental uncertainty in a_e (about 5×10⁻¹⁰) or a_{μ} (about 2×10⁻⁸) requires (in both cases) that

$$M_{\rm const} \ge 10^6 \,\,{\rm GeV}$$
 (2)

In this note we will implement a version of this analysis, using the Drell-Hearn-Gerasimov (DHG) sum rule,⁷ to extract severe bounds on the single l^* production processes $e^+e^- \rightarrow l^+\overline{l} + l\overline{l^*}$ and $lp \rightarrow l^+X$, where l^* is an excited state of l associated with compositeness. The result, in essence, is that such processes will never be observed, regardless of the energy of the interaction, as long as $m_{1}*$ $\leq 10^3$ TeV. We will also comment on the feasibility of obtaining visible emulsion tracks due to l^* 's produced in high-energy $(>10^{19} \text{ eV})$ cosmic-ray events.

Consider the matrix element for the photo- or electroexcitation of a lepton resonance of mass m_1 *. For the present, we consider only parityconserving dipole excitations of a spin- $\frac{1}{2}$ resonance, and choose⁸:

$$\langle l^* | e J_{\mu} | l \rangle = \frac{ie}{2M} \,\overline{U}(p+q,\Lambda) \sigma_{\mu\nu} q^{\nu} \binom{1}{\gamma_5} u(p,\lambda) \,,$$
(3)

where the mass M is a measure of the strength of the excitation. (In a composite model, M^{-1} may be very small, but it is formally of zero order in α .) From Eq. (3), the contribution of l^* to $\delta a_l^{\text{non-QED}}$, as obtained through the DHG sum rule, is easily calculated

$$(\delta a_{l}^{\text{non-QED}})^{2} = \frac{m_{l}^{2}}{2\pi^{2}\alpha} \int_{0}^{\infty} \frac{dv}{\nu} [\sigma_{3/2}(\nu) - \sigma_{1/2}(\nu)]_{l} *$$
$$= -(m_{l}/M)^{2}$$
(4)

and hence⁹

$$\left| \delta a_{l}^{\text{non-QED}} \right| \simeq \left(m_{l} / M \right) \,. \tag{5}$$

This is in agreement with the results of Refs. 4 and 5, if M is taken to be the mass of the constituent. However, we do not demand this identification of M. The lower limit on M is then fixed by the uncertainties in $a_{e,\mu}$, just as was the limit on M_{const}

$$M \ge 10^6 \text{ GeV}. \tag{2'}$$

However, we now have an additional piece of information since, independent of the interpretation of M in Eq. (2'), the matrix element (3) is now bounded. Since this matrix element controls the processes $e^*e^- \rightarrow l^*\overline{l} (+l\overline{l}^*)$ and $lp \rightarrow l^*X$, and the radiative decay rate $l^* \rightarrow l_{\gamma}$, we can apply our bound (2') directly to these processes.

(1) Consider the one-photon annihilation process

$${}^{+}e^{-} \rightarrow l^{*}\overline{l} (+ l\overline{l}^{*}) .$$

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It is a straightforward algebraic exercise to evaluate the contribution to R from the $l l \gamma$ final state, using the matrix element (4) (properly crossed). It is

$$\delta R(l \, \overline{l} \gamma) = \sigma_{e^+e^- l \overline{l} \gamma}^{\text{non-QED}} / \sigma_{e^+e^- \mu^+ \mu^-}$$

$$= \frac{1}{4} \left(\frac{m_1 *}{M} \right)^2 \left(\frac{s}{m_1 *^2} - \frac{3m_1 *^2}{s} + \frac{2m_1 *^4}{s^2} \right)$$

$$\times F(s) \left(\frac{\Gamma_l * I \gamma}{\Gamma_l t t} \right), \qquad (6)$$

where F(s) is a form factor $(|F(s)| \le 1)$. It is clear, in comparing Eqs. (6) and (2'), that

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$$\delta R(l\,\bar{l}\,\gamma) \leq \frac{1}{4} (m_{l}*/10^{\circ} \,\mathrm{GeV})^{2} , \qquad (7)$$

assuming that F decreases fast enough with s to tame the kinematic factor. Hence this channel will not be seen at any s if m_l^* is of the order of 10^3 TeV or less [since $R^{\text{QED}}(l\bar{l}\gamma) \simeq 0.2$].¹⁰

(2) An analogous exercise shows that

$$\frac{d\sigma}{d^3p'/E'} (lp \to l^*X) \sim \left(\frac{m_l^*}{10^6 \text{ GeV}}\right)^2 \frac{d\sigma}{d^3p'/E'} (lp \to lX) \quad (8)$$

and the same conclusions hold as in (1) above. These results also carry over to the case of spin- $\frac{3}{2}$ resonances. Hence, it is fair to conclude that lower-mass l^* 's (in the TeV range) are best sought in experiments designed for associated production (e.g.,

$$e^+e^- \rightarrow l^+\bar{l}^+$$
 or $pp \rightarrow l^+\bar{l}^+ + X$).

(3) Finally, there is an optimistic aspect following from the bound (2') established on M: From the matrix element (4) we may calculate the proper

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lifetime for the radiative decay of the l^* to be $\tau_{l^*} = 4M^2/\alpha m_l *^3$. A bit of arithmetic then shows that if the radiative mode dominates the decay channels of the l^* , then an l^* carrying energy E in the laboratory has a mean free path of

$$\lambda_{lab}(\text{in mm}) \ge \frac{E(\text{in eV}) \times 10^{-19}}{[m*(\text{in TeV})]^4} .$$
(9)

Hence if $m_{l} * \sim 1$ TeV, cosmic-ray events in the energy range >10¹⁹eV could give rise to visible l^* paths in emulsions.

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- ⁸The most general gauge-invariant transition matrix element in an orthogonal basis is

$$\langle l^{*}(p)|eJ_{\mu}|l(p)\rangle = e\,\overline{U}(P)\,G_{\text{Coul}}(q^{2})[(P \cdot q)\,q_{\mu} - q^{2}P_{\mu}]$$

$$+ i\,(m_{l} + m_{l}^{*})G_{T}(q^{2})\,\epsilon_{\mu\nu\lambda\sigma}P^{\nu}q^{\lambda}\gamma^{\sigma}\gamma_{5}$$

$$\times \begin{cases} 1 \\ \gamma_{5} \end{cases} u\,(p)$$

[J. D. Bjorken and J. D. Walecka, Ann. Phys. (N.Y.) 38, 35 (1966)]. The DHG sum rule places a restriction only on the transverse piece $G_T(0)$. The choice (3) in the text for the matrix element corresponds to the symmetric constraint $G_{\text{Coul}}(q^2) = G_T(q^2)$. The results in part (3) of the text (pertaining to the decay l^* $\rightarrow l\gamma$) are *independent* of the assumed form (3), since this parametrization is completely general for real photons. The conclusions of parts (1) and (2) depend

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only on $G_{\rm Coul}$ not being grossly larger than G_T . Since most reasonable models of transition form factors of composite objects will give (for $|q^2| \sim m_I \star^2$ or less) $G_{\rm Coul} \approx G_T$, each being proportional to an overlap integral of parton wave functions of l and l^* , it is simplest for an order-of-magnitude estimate, to use the form (3) which (as stated above) enforces $G_{\rm Coul} = G_T$.

⁹The minus sign in Eq. (4) is not troublesome, since we are obtaining the contribution of a single l^* to $(\delta a^{\text{non-QED}})^2$. (It is analogous to the negative contribu-

tion of the $P_{11}(1450)$ to $a_{nucleon}^2$). A complete theory would presumably give rise to enough helicity $\frac{3}{2}$ scattering to correctly give a positive $(\delta a_1^{non-QED})^2$. We do, however, take the absolute contribution in Eq. (4) as indicative of the magnitude of $(\delta a^{non-QED})^2$.

¹⁰The converse of this statement is that if an l^* of mass $\leq 10^3$ TeV is observed in $e^+e^- \rightarrow l^*\overline{l}$, then the separation $\delta a^{\text{QED}} + \delta a^{\text{non-QED}}$ is not valid, and the dynamics of the l^* involves a coupling related to α_{EM} .