Constraints for $N=2$ superspace from extended supergravity in ordinary space

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The constraints on the supertorsion and Maxwell curvatures for $N = 2$ supergravity are derived from $N = 2$ extended supergravity in ordinary space with a closed gauge algebra. Nonlinear constraints are found.

I. INTRODUCTION

Supergravity can be formulated in ordinary spacetime or in superspace. In ordinary spacetime the theory obtains a closed gauge algebra if one adds auxiliary fields. For $N = 1$ supergravity, the minimal set consists of a scalar S , a pseudoscalar P, and an axial vector A_{m} ¹ For $N=2$, one set is known²⁻⁴ which is presumably minimal and consists of two spin- $\frac{1}{2}$ SO₂ doublets (λ^i, χ^i) , two symmetric traceless SO_2 tensors $(A_m^{i_j}, P^{i_j})$, four antisymmetric SO_2 tensors $(V_m^{ij}, t_{mn}^{ij}, M^{ij}, N^{ij}),$ and three SO_2 scalars (A_m, S, V_m) . The gauge algebra of spacetime symmetries, SO, rotations, and supersymmetry closes, and the structure constants depend on physical as well as on auxiliary fields.

At this point we emphasize that the SO, group we consider acts only on the photon B_{μ}^{ij} but not on the indices i of ψ^i_μ , etc. In other words, this local SO_2 gauges the central charge of $N=2$ supersymmetry, but not the outside charge which rotates the generators Q^i and Q^j into each other. For the action this simply means that the photon only appears in the Maxwell curl, so that the action is locally SO, invariant, but not, for example, as $\partial_{\mu}\psi_{\nu}^{i}+e\epsilon^{i}iB_{\mu}^{jk}\psi_{\nu}^{k}-(\mu\leftrightarrow\nu)$. As a result, local supersymmetry and SO, commute. The reason we make this choice is that in Ref. 3 only the results for this case have been given. The $N=2$ extended supergravity model was first found in Ref. 5 and an extension to a local SO, which acts on the indices for ψ^i_μ was given in Ref. 6 without auxiliary fields. This result agrees with a group-theoretical approach, in which one gauges $Sp(4) \times SO(2)$ (see Ref. 7). In this de Sitter algebra the photon charge appears of course on the lefthand and on the right-hand side. If one takes a Wigner-Inönü group contraction, one reverts to the $SO₂$ which is a central charge. Thus, we will consider a base manifold $R⁴$ in which general coordinate transformations act, and a tangent

manifold in which one has local Lorentz symmetry and local supersymmetry as well as SO, rotations of the photon alone. The latter still enable one to compute the parameter composition law for the $SO₂$ parameter. Also, in superspace we will restrict ourselves to this kind of $SO₂$ group.

In superspace one needs constraints on the supertorsions. For minimal $N=1$ supergravity, the constraints read $(a, b, c$ are four-component spinor indices, and r, s, t bosonic indices)

$$
T_{ab}^r = -\frac{1}{4} (C\gamma^r)_{ab}, \quad T_{rs}^t = T_{ab}^c = T_{ar}^s = 0.
$$
 (1)

They were found by a series of unrelated arguments,⁸ but no systematic method seems to exist. However, one can establish a bridge between ordinary space and superspace by making a choice of gauge in superspace and requiring compatiof gauge in superspace and requiring compati-
bility.⁹⁻¹¹ In this way it was found¹² that the set in Eq. (1) follows from the gauge algebra for $N=1$ theory with S, P, A_{m} . If one replaces in superspace at $\theta = 0$ the ordinary spin connection $\omega_{\mu}^{ab}(e, \psi)$ by the improved spin connection, one finds 12 the constraints which also appear in the approach in which superspace consists of two approach in which superspace consists of two
small chirally conjugated superspaces.¹³ Also for $N=1$ conformal supergravity, the constraints for $N=1$ conformal supergravity, the constrain
in Eq. (1) hold.¹⁴ Moreover, the bridge betwee ordinary space and superspace yields here also the local scale and chiral transformations in superspace, and they agree with the ones pre-
viously proposed.¹⁵ viously proposed.¹⁵

For $N=2$ superspace, Wess has proposed a set of constraints on the supertorsion similar to Eq. $(1).^{16}$ In this article we construct these Eq. $(1).$ ¹⁶ In this article we construct these constraints by using the bridge between ordinary space and superspace and assuming the results straints by using the bridge between ordinary
space and superspace and assuming the result
of the $N=2$ auxiliary fields.^{2,3} Our results are

$$
T_{ai, bj}^{r} = -\frac{1}{4} \delta_{ij} (C\gamma^{r})_{ab}, \quad T_{rs}^{t} = T_{ai,s}^{t} = 0,
$$

\n
$$
T_{Ai, Bj}^{c} = -\frac{3}{8} \epsilon_{AB} \epsilon_{ij} (\epsilon^{kl} \lambda_{i}^{c}) - \frac{1}{8} \delta_{iA}^{c} \epsilon_{B} \rho_{D} \delta_{iA}^{k} \lambda_{j}^{D},
$$

\n
$$
T_{Ai}^{ck} = 0, \quad T_{Ai, Bk}^{ck} = 0, \quad T_{Ai, Bj}^{ck} = \frac{1}{2} \delta_{A}^{c} T_{Di, Bj}^{jk},
$$
\n(2)

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where the symbol λ_i^C in our method is the $N=2$ auxiliary field, and as we shall discuss, must be replaced in Eq. (2) by

$$
\lambda_{BI} = 2T_{BIBR}^{\dot{D}k} \tag{3}
$$

The last two constraints in Eq. (2} are consequences of the following constraint which we will derive

$$
T^{\hat{C}k}_{\hat{A}i, Bj} = \frac{1}{2} \delta^{\hat{C}}_{\hat{A}} (\delta_{ij} \delta^{kl} - \frac{1}{2} \delta^k_i \delta^l_j) \lambda_{Bl} . \tag{4}
$$

Just as in the $N=1$ nonminimal theory, one can introduce a spinor superfield T_{BI} such that its $\theta = 0$ component equals λ_{BI} . Capital letters denote two-component spinors (see the Appendix). We used in Eq. (2) four-component notations where it led to a simplification. In addition, we find for the $SO₂$ curvatures in superspace the constraints

$$
F_{aI, s}^{ij} = 0, \quad F_{aI, bI}^{ij} = -\frac{1}{\sqrt{2}} C_{a b} \delta_{i}^{i i} \delta_{I}^{j 1}.
$$

The symbols () and [] denote symmetrization and antisymmetrization; for example, $A^{(i}B^{j)}$ $=A^{i}B^{j}+A^{j}B^{i}$. For F^{ij}_{mn} we find the supercovaria photon curl, while T_{rs}^{ck} contains the supercovaria $\text{spin-}\frac{3}{2}$ curl and $T^{ck}_{ai,s}$ contains many auxiliar fields. Qur results agree with Wess's results on many points, except that he proposes (in our notation)

$$
\epsilon^{ii} D^A{}_i T_{Ai, \dot{B}j} \dot{c}^k = 0 \tag{5}
$$

which we do not find. (Note that Wess raises isospin indices with ϵ^{ij} , whereas we use δ^{ij}). Specifically Wess has proposed the constraints (in our notation) $17, 18$

$$
T_{ai, bj}^{r} + \frac{1}{4}\delta_{ij}(C\gamma^{r})_{ab} = T_{rs}^{t} = T_{ai,s}^{t} = 0,
$$
\n(6)

$$
T_{Ai,Bj}{}^{\dot{C}k} = T_{Ai,Bj}{}^{\dot{C}j} = \oint_{ijk} \epsilon_{i1} T_{Ai}{}^{k} \cdot \dot{B}^{i}{}^{k} = 0 \;, \tag{7}
$$

in addition to the constraint of Eq. (5). It can be seen that our results [first line of (2)] agree completely with Eq. (6) above. Similarly, the first two results in Eq. (7) agree. To see the equivalence of the remaining algebraic constraints, we note that the general solution of the "cyclicity" and traceless equations on T_{A} , \hat{h}^{C} is

$$
T_{A j, \dot{B}}^{I, \dot{C}}_{h} = (\delta_{j}^{I} \delta_{h}^{m} - \frac{1}{2} \delta_{k}^{I} \delta_{j}^{m}) f_{A \dot{B}}^{C}{}_{m}
$$
\n(8)

for some spinorial quantity f_{AB}^c . Next there is a dimension- $\frac{1}{2}$ superspace Bianchi identity,

$$
\begin{aligned} [[\mathfrak{D}_{Ai}\mathfrak{D}_{Bj}\}, \overline{\mathfrak{D}}_{\dot{C}k}\} + [[\overline{\mathfrak{D}}_{\dot{C}k}, \mathfrak{D}_{Bj}], \mathfrak{D}_{Ai}] \\ + [[\overline{\mathfrak{D}}_{\dot{C}k}, \mathfrak{D}_{Ai}\}, \mathfrak{D}_{Bj}] = 0 \;, \; (9) \end{aligned}
$$
 which on using the constraints on $T_{ai, bj}^r$ and $T_{ai, s}^t$

yields-

$$
T_{Ai, Bj}^{D}{}_{k}(\sigma^{r})_{\tilde{C}D} + T_{\tilde{C}k, Bj}^{D}{}_{i}(\sigma^{r})_{\tilde{D}A} + T_{\tilde{C}k, A i}^{D}{}_{j}(\sigma^{r})_{\tilde{D}B} = 0
$$
\n(10)

as the coefficient of D_r . This equation can be multiplied by $(\sigma_r)_{\vec{E}F}$. Utilizing the Fierz identity for two-component spinors (see Appendix) we obtain a result which contains two independent equations:

$$
f_{A(\dot{B}\dot{C})m} = 0 \to f_{AB} \dot{C}_m = \frac{1}{2} \delta_{\dot{B}}^{\dot{C}} T_{Am} \,, \tag{11}
$$

$$
T_{Ai, Bj}^{Ck} = -\frac{3}{8} \epsilon_{AB} \delta_{[i}^{k} T_{j]}^{C} - \frac{1}{8} \delta_{(A}^{C} \epsilon_{B)D} \delta_{i}^{k} T_{j}^{D}.
$$
 (12)

From Eq. (8) we thus obtain

$$
T_{Ai, \hat{B}j}{}^{\hat{C}k} = \frac{1}{2} \delta_{\hat{B}}^{\hat{C}} (\delta_{ij} \delta^{kl} - \frac{1}{2} \delta_{j}^{k} \delta_{i}^{l}) T_{AI}
$$
 (13)

and complete agreement has now been shown explicitly if one identifies $T_{Al} = \lambda_{Al}$.

The differential constraint, however, is in disagreement with our results. To show this, we use the fact that T_{Am} , the superfield defined in (11) and (8), is a good tensor. Hence, using the compatibility method foilowed in this paper, one finds

$$
\delta(\epsilon)\lambda^{am} = [(\epsilon^{\beta j}\partial/\partial \vartheta^{\beta j})T^{am}(x,\vartheta)]_{\vartheta=0}.
$$
 (14)

From the supersymmetry variation $\delta(\epsilon)$ of the auxiliary field λ^{Am} as given in Ref. 3, one finds that

$$
D_{bj}T^a{}_i = \left[\eta_{ij} - \frac{1}{2}(\dot{V} + i\gamma_5 \mathcal{A})_{ij} - \frac{1}{2}\dot{V}\delta_{ij}\right]
$$

$$
-\frac{1}{2}(M + i\gamma_5 N)_{ij}\Big]^{a}{}_{b} - \lambda^a{}_{i}\bar{\lambda}_{bj}.
$$
 (15)

(5) The constraint in (5) becomes with (13)

$$
D_{Am}\lambda_{R}^{A} + k \rightarrow m = 0.
$$
 (16)

Substituting (15) and using two-component notation [which means multiplying by $\frac{1}{2}(1+\gamma_5)^b{}_a = \delta^B_A$] we find

$$
2S\delta^{ij} + 2i p^{ij} - \lambda^{Ai}\lambda^{Bj}\epsilon_{AB} = 0.
$$
 (17)

This completes the proof that the formulation by Wess is *inequivalent* to that of de Wit and van 'Holten. We have not, however, shown that the formulation of Ref. 16 is inconsistent. This requires a detailed study of the superspace Bianchi identities. Perhaps it allows no action.

Fradkin and Vasiliev' have also given component results for SO_2 supergravity. Between their first and second works, they have made many field redefinitions. However, the net effect of these redefinitions is to bring their formulations into precise agreement with the form first presented by de Wit and van Holten. [The differences in notation used by these two sets of authors obscures this fact but, utilizing SU(2) Fierz identities, equivalence has been found.¹⁹

We have not constructed the constraints on Lorentz supercurvatures, since they follow according to a general theorem from the Bianchi identities, once the supertorsion constraints are known.²⁰

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Many other authors²¹ have studied constraints in superspace from points or view or by using methods which differ from ours. We intend to compare in the future their results to the results obtained by the method we follow, and to see whether proper choice of integration constants can simplify our or their results.

II. CONSTRAINTS IN FOUR-COMPONENT **FORMALISM**

We extend to $N=2$ supergravity a method used in Refs. 9-12 and 14 for finding superspace con-

strains. The SO₂ theory with auxiliary fields²⁻⁴ is invariant under local supersymmetry
$$
s
$$
, Maxwell m , and Lorentz transformations l , all of which act in the tangent space; and under general coordinate transformations g acting in the base manifold. The closed gauge algebra³ is given by

$$
\begin{aligned} \left[\delta(\epsilon_1, \varphi_1, \lambda_1, \xi_1), \delta(\epsilon_2, \varphi_2, \lambda_2, \xi_2) \right] \\ &= \delta_{\epsilon}(\epsilon_{12}) + \delta_m(\varphi_{12}) + \delta_l(\lambda_{12}) + \delta_{\epsilon}(\xi_{12}), \end{aligned} \tag{18}
$$

where

$$
\epsilon_{12}^{i} = \xi_{12}^{p} \partial_{\nu} \epsilon_{11}^{i} + \frac{1}{2} \lambda_{12}^{rs} \sigma_{rs} \epsilon_{11}^{i} + \frac{1}{2} \overline{\epsilon}_{1}^{i} \gamma^{\rho} \epsilon_{2}^{j} (\psi_{\rho}^{i} + \gamma_{\rho} \lambda^{i}) - \frac{3}{8} (\overline{\epsilon}_{1}^{i} \epsilon_{2}^{j} + \overline{\epsilon}_{1}^{i} \gamma_{5} \epsilon_{2}^{j} \gamma_{5}) \lambda^{j}
$$
\n
$$
- \frac{1}{8} (\overline{\epsilon}_{1}^{i} \gamma^{\rho} \epsilon_{2}^{j} \gamma_{\rho} + \overline{\epsilon}_{1}^{i} \gamma^{\rho} \gamma_{5} \epsilon_{2}^{j} \gamma_{\rho} \gamma_{5}) \lambda^{j} + \frac{1}{4} (\overline{\epsilon}_{1}^{i} \sigma^{rs} \epsilon_{2}^{j} \sigma_{rs}) \lambda^{j},
$$
\n
$$
\varphi_{12}^{i,j} = \xi_{12}^{v} \partial_{v} \varphi_{11}^{i,j} - \frac{1}{2} \overline{\epsilon}_{1}^{r} \gamma^{\rho} \epsilon_{2}^{k} b_{\rho}^{i,j} - \frac{1}{\sqrt{2}} \overline{\epsilon}_{1}^{r} \epsilon_{2}^{i} + \varphi_{12}^{ik} \varphi_{11}^{kj},
$$
\n
$$
\lambda_{12}^{rs} = \xi_{12} \partial_{v} \lambda_{11}^{rs} + \lambda_{12}^{rs} \lambda_{11}^{is} - \frac{1}{2} \overline{\epsilon}_{1}^{r} \gamma^{\rho} \epsilon_{2}^{k} \omega_{\rho}^{rs} + \frac{1}{4} \overline{\epsilon}_{1}^{i} (T + \gamma_{5} \overline{T})^{rsi} \delta_{\epsilon_{2}^{j}} + \frac{1}{2} \overline{\epsilon}_{11}^{i} (\sigma^{rs} \eta^{i,j} + \eta^{i,j} \sigma^{rs}) \epsilon_{2}^{j}
$$
\n
$$
\xi_{12}^{\mu} = \xi_{12}^{\nu} \partial_{v} \xi_{11}^{\mu} + \frac{1}{4} \overline{\epsilon}_{12}^{i} \gamma^{\mu} \epsilon_{11}^{i}
$$

and the supersymmetry variations of the physical fields are

$$
\delta e^{\mu}_{\mu} = \frac{1}{2} \overline{\epsilon}^i \gamma^r \psi^i_{\mu},
$$
\n
$$
\delta b^{ij}_{\mu} = -\frac{1}{\sqrt{2}} \overline{\epsilon}^{i} \psi^{j}_{\mu},
$$
\n
$$
\delta \psi^{i}_{\mu} = D_{\mu} \epsilon^{i} + \frac{i}{2} A_{\mu} \gamma_5 \epsilon^{i} + \frac{1}{2} (V^{ij}_{\mu} + \overline{\psi}^{i}_{\mu} \lambda^{j}) \epsilon^{j}
$$
\n
$$
- \frac{i}{2} (A^{ij}_{\mu} + i (\overline{\psi}^{i}_{\mu} \gamma_5 \lambda^{j})_s) \gamma_5 \epsilon^{j} - \frac{1}{4} \sigma \cdot T^{ij} \gamma_{\mu} \epsilon^{j}
$$
\n
$$
- \frac{1}{2} \gamma_{\mu} \eta^{ij} \epsilon^{j} - \frac{1}{2} \overline{\epsilon}^{i} \lambda^{j} \psi^{j}_{\mu} - \frac{1}{2} (\overline{\epsilon}^{i} \gamma_5 \lambda^{j})_s \gamma_5 \psi^{j}_{\mu}.
$$
\n(20)

In the case of $N=2$, the $\phi\phi$ term in (19) vanishes. The notations are as in Ref. 3, the gravitational coupling constant κ is taken equal to 1, and ϵ is normalized such that $\delta \psi_{\mu}^{i} = \partial_{\mu} \epsilon + \text{more. Under}$ Maxwell SO₂ rotations, $\delta b_\mu^{ij} = -\partial_\mu \phi^{ij}$ (thus our signs in ϕ_{12}^{ij} differ from Ref. 3).

In superspace we assume three local symmetries¹¹: general coordinate transformations $\delta_{G}(\Xi^{\Pi})$ with $\Pi = (\mu, \alpha)$, local Lorentz rotations $\delta_L(\Lambda^{rs})$ and SO₂ Maxwell rotations $\delta_M(\Phi^{ij})$. The index convention is as in (12) and (14) . The parameters $\Xi^{\Pi}(x,\theta)$, $\Lambda^{rs}(x,\theta)$, $\Phi^{ij}(x,\theta)$ depend on x^{μ}
and $\theta^{\alpha i}$ with $\alpha = 1, 4$, $i = 1, 2$ since we consider $N=2$ superspace supergravity.¹¹ The superfields are the vielbein $E_{\pi}^{A}(x, \theta)$, the Lorentz connection $\Omega_{\overline{n}}^{mn}(x,\theta)$, and the Maxwell connection $B_{\overline{n}}^{ij}(x,\theta)$. Their variations under the whole symmetry group are

$$
\delta(\Xi,\Lambda,\Phi)E_{\Pi}^{A} = \Xi^{\Lambda}\partial_{\Lambda}E_{\Pi}^{A} + (\partial_{\Pi}\Xi^{\Lambda})E_{\Lambda}^{A}
$$

+
$$
\frac{1}{2}\Lambda^{rs}X_{rs}E_{\Pi}^{A} + \frac{1}{2}\Phi^{ij}Y_{i,j}E_{\Pi}^{A},
$$

$$
\delta(\Xi,\Lambda,\Phi)\Omega_{\Pi}^{rs} = \Xi^{\Lambda}\partial_{\Lambda}\Omega_{\Pi}^{rs} + (\partial_{\Pi}\Xi^{\Lambda})\Omega_{\Lambda}^{rs}
$$

$$
- \partial_{\Pi}\Lambda^{rs} - \Omega_{\Pi}^{rt}\Lambda^{ts} - \Omega_{\Pi}^{st}\Lambda^{rt},
$$

$$
\delta(\Xi,\Lambda,\Phi)B_{\Pi}^{ij} = \Xi^{\Lambda}\partial_{\Lambda}B_{\Pi}^{ij} + (\partial_{\Pi}\Xi^{\Lambda})B_{\Lambda}^{ij}
$$

$$
- \partial_{\Pi}\Phi^{ij} - B_{\Pi}^{is}\Phi^{kj} - B_{\Pi}^{js}\Phi^{is},
$$
 (21)

where the generators of Lorentz and Maxwell transformations X_{rs} and Y_{ij} act on tensors as follows:

$$
\frac{1}{2}\Lambda^{rs}X_{rs}E_{\Pi}^{m} = \Lambda^{m}{}_{n}E_{\Pi}^{n}, \quad \frac{1}{2}\Lambda^{rs}X_{rs}E_{\Pi}^{ak} = \frac{1}{2}(\Lambda \cdot \sigma)^{ab}E_{\Pi}^{bk},
$$
\n
$$
\frac{1}{2}\Phi^{ij}Y_{ij}E_{\Pi}^{A} = 0, \quad \frac{1}{2}\Phi^{ij}Y_{ij}B_{\Pi}^{kl} = \Phi^{kj}B_{\Pi}^{jl} + \Phi^{lj}B_{\Pi}^{kl}.
$$
\n(22)

In $N=2$ supergravity, the $B\Phi$ terms in (21) and (22) vanish. From (21) one can extract the superspace gauge algebra 11 :

$$
\begin{aligned}\n\left[\delta(\Xi_1, \Lambda_1, \Phi_1), \delta(\Xi_2, \Lambda_2, \Phi_2)\right] &= \delta_G(\Xi_{12}) + \delta_1(\Lambda_{12}) + \delta_M(\Phi_{12}) \\
\Xi_{12}^{\Pi} &= \Xi_{12}^{\Lambda} \delta_{\Lambda} \Xi_{11}^{\Pi} + \delta_{11} \Xi_{21}^{\Pi}, \\
\Lambda_{12}^{rs} &= \Lambda_{12}^{rt} \Lambda_{13}^{ts} + \Xi_{12}^{\Pi} \delta_{\Pi} \Lambda_{13}^{rs} + \delta_{11} \Lambda_{23}^{rs},\n\end{aligned} \tag{23}
$$

Finally, we define covariant derivatives, supertorsions, and supercurvatures:

$$
D_{\Pi} = \partial_{\Pi} + \frac{1}{2} \Omega_{\Pi}^{S} X_{rs} + \frac{1}{2} B_{\Pi}^{ij} Y_{ij} , D_A = E_A^A D_A ,
$$

\n
$$
[D_A, D_B] = -2T_{AB}^C D_C + \frac{1}{2} R_{AB}^{mn} X_{mn} + \frac{1}{2} F_{AB}^{ij} Y_{ij} ,
$$
 (24)

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 (19)

 $\overline{22}$

)

÷,

$$
T_{AB}^{C} = \frac{1}{2}(-)^{\Lambda^{C}B^{*H}} E_{A}^{\Lambda} E_{B}^{T} [D_{\Lambda} E_{\Pi}^{C} - (-)^{\Lambda^{T}} D_{\Pi} E_{\Lambda}^{C}],
$$

\n
$$
R_{AB}^{mn} = (-)^{\Lambda^{C}B^{*H}} E_{A}^{\Lambda} E_{B}^{T} [\partial_{\Lambda} \Omega_{\Pi}^{mn} - (-)^{\Lambda^{T}} \partial_{\Pi} \Omega_{\Lambda}^{mn} + \Omega_{AB}^{ms} \Omega_{\Pi}^{sn} - (-)^{\Lambda^{T}} \Omega_{\Pi}^{ms} \Omega_{\Lambda}^{sn}],
$$
\n
$$
F_{AB}^{ij} = -(-)^{\Lambda^{C}B^{*H}} E_{A}^{\Lambda} E_{B}^{T} [\partial_{\Lambda} B_{H}^{ij} - (-)^{\Lambda^{T}} \partial_{\Pi} B_{A}^{ij} + B_{A}^{ik} B_{\Pi}^{kj} - (-)^{\Lambda^{T}} B_{H}^{ik} B_{A}^{kj}].
$$
\n(25)

We have chosen the same sign for the ΩX and BY terms in D_{II} , so that our BY term has opposite sign to the one in Ref. 11.

We proceed now as in Refs. 12 and 14 by first choosing a gauge through the identifications at order θ^0 :

$$
E_{\rho}^{r}(x, \vartheta=0) = e_{\rho}^{r}(x), \quad \Xi^{\rho}(x, \vartheta=0) = \xi^{\rho}(x),
$$

\n
$$
\Omega_{\rho}^{rs}(x, \vartheta=0) = \omega_{\rho}^{rs}(x), \quad \Lambda^{rs}(x, \vartheta=0) = \lambda^{rs}(x),
$$

\n
$$
E_{\rho}^{ai}(x, \vartheta=0) = \psi_{\rho}^{ai}(x), \quad \Xi^{\alpha}^{i}(x, \vartheta=0) = \epsilon^{\alpha}^{i}(x),
$$

\n
$$
B_{\rho}^{ij}(x, \vartheta=0) = b_{\rho}^{ij}(x), \quad \Phi^{ij}(x, \vartheta=0) = \varphi^{ij}(x).
$$
 (26)

Then the higher-order components follow by requiring consistency between the gauge algebras of SO_2 superspace supergravity and of SO_2 ordinary supergravity. As the purpose of this paper is to find constraints on gauge-invariant quantities, all the "integration constants" are set equal to zero. The results, to first order in θ , are

given in Tables I and II.

Using (25) and Tables I and II, it is now straightforward to derive $T_{AB}^C(x, \theta=0)$. We get the following tensor relations (therefore true to all orders in θ):

$$
T_{r,s}^{t} = T_{ai,s}^{t} = 0, \quad T_{ai,bj}^{t} = -\frac{1}{4} \delta_{ij} (C \gamma^{t})_{ab},
$$

\n
$$
T_{ai,bj}^{ck} = \frac{1}{4} \delta_{ij} (C \gamma^{\rho})_{ab} (\gamma_{\rho} \lambda^{k})^{c}
$$

\n
$$
- \frac{3}{16} [\delta_{ij}^{k} C_{ab} \lambda_{j1}^{c} + \delta_{li}^{k} (C \gamma_{5})_{ab} (\gamma_{5} \lambda_{j1})^{c}]
$$

\n
$$
- \frac{1}{16} [\delta_{ij}^{k} (C \gamma^{\rho})_{ab} (\gamma_{\rho} \lambda_{j})^{c} + \delta_{li}^{k} (C \gamma^{\rho} \gamma_{5})_{ab} (\gamma_{\rho} \gamma_{5} \lambda_{j1})^{c}]^{\prime}
$$

\n
$$
+ \frac{1}{8} \delta_{ij}^{k} (C \sigma^{\rho \sigma})_{ab} (\sigma_{\rho \sigma} \lambda_{j})^{c}, \qquad (27)
$$

$$
T_{r,s}^{ck} = \frac{1}{2} e^{\mu}_r e^{\nu}_s (\psi_{\mu\nu}^{\text{cov}})^{ck} - \psi_r^{\alpha t} \psi_s^{\beta j} (T_{\alpha t, \beta j}^{ck}),
$$

$$
T_{ai,s}^{\alpha t} = \psi_s^{\beta j} (T_{ai, \beta j}^{ck}) + \frac{1}{2} H_{s,ai}^{ck},
$$

where $(\psi_{\mu\nu}^{\text{cov}})^{ck}$ is the supercovariant spin- $\frac{3}{2}$ curl i.e. ,

$$
(\psi_{\mu\nu}^{\infty})^{ck} = \tilde{D}_{\Gamma\mu}\psi_{\nu 1}^{ck} \tag{28}
$$

and \tilde{D}_{μ} is defined by $\delta_s \psi_{\mu}^k = \tilde{D}_{\mu} \epsilon^k$ [see the supersymmetry variation of ψ_{μ}^{ck} in (20)]. $H_{s,ai}^{ck}$ is the tensor $e_s^{\nu} (D_{ai} E_{\nu}^{ck} - D_{\nu} E_{ai}^{ck})$, which can be writte (order θ^{δ}) as

$$
e_{s}^{\nu} \partial_{ai} E_{\nu}^{ck} - \frac{1}{2} (\omega_{s} \cdot \sigma)^{cd} \delta_{a}^{d} \delta_{i}^{k} \tag{29}
$$

and from the explicit form of $E_v^{ck}(x, \theta^1)$ we see that the ω terms cancel, leaving

$$
\begin{split} \Xi^{\mu} &= \xi^{\mu} - \frac{1}{4} \epsilon^{j} \gamma^{\mu} \mathcal{S}^{j} \\ \Xi^{i} &= \epsilon^{i} - \frac{1}{2} \lambda^{rs} \mathcal{S}^{i} + \frac{1}{4} \epsilon^{j} \gamma^{\rho} \mathcal{S}^{j} (\psi_{\rho}^{i} + \gamma_{\rho} \lambda^{i}) \\ &\quad - \frac{3}{16} (\epsilon^{i} \mathcal{S}^{j} + \epsilon^{i} \gamma_{5} \mathcal{S}^{j} \gamma_{5}) \lambda^{j} - \frac{1}{16} (\epsilon^{i} \gamma^{\rho} \mathcal{S}^{j} \gamma_{\rho} + \epsilon^{i} \mathcal{S}^{j} \gamma_{\rho} \gamma_{5}) \lambda^{j} + \frac{1}{8} (\epsilon^{i} \sigma^{\alpha} \mathcal{S}^{j}) \sigma_{\rho \sigma} \lambda^{j} \\ \Phi^{ij} &= \varphi^{ij} + \frac{1}{4} (\overline{\mathcal{S}}^{k} \gamma^{\rho} \epsilon^{k}) b_{\rho}^{ij} + \frac{1}{2 \sqrt{2}} \mathcal{S}^{i} i \epsilon^{j} \mathcal{I} \\ \Lambda^{rs} &= \lambda^{rs} - \frac{1}{4} (\epsilon^{i} \gamma^{\rho} \mathcal{S}^{i}) \omega_{\rho}^{rs} + \frac{1}{8} \epsilon^{i} (T + \gamma_{5} \tilde{T})^{rs} \mathcal{S}^{i} \mathcal{S}^{j} + \frac{1}{4} \epsilon^{i} (\sigma^{rs} \eta^{ij} + \eta^{ij} \sigma^{rs}) \mathcal{S}^{j} \end{split}
$$

^a A term $-\Phi^{ij}\theta^j$ in Ξ^i present in Ref. 1 is absent here, because our SO₂ gauges the central charge.

$$
H_{r,a,i}^{ck}(x,\vartheta=0)=\frac{i}{2}A_r(\gamma_5)^c{}_a\delta_i^k+\frac{1}{2}(\gamma_r^{kj}+\overline{\psi}_r^{lk}\lambda^{j})\delta_a^c\delta_i^j
$$

$$
-\frac{i}{2}(A_r^{kj}+i(\overline{\psi}_r^k\gamma_5\lambda^j)_s)(\gamma_5)^c{}_a\delta_i^j-\frac{1}{4}(\sigma\cdot T^{kj}\gamma_r)^c{}_a\delta_i^j
$$

$$
-\frac{1}{2}(\gamma_r\eta^{kj})^c{}_a\delta_i^j-\frac{1}{2}\delta_i^{lk}\lambda_a^j\psi_r^{cj}-\frac{1}{2}(\delta_i^k(\gamma_5\lambda^j)_a)_s(\gamma_5\psi_r)^{cj}.
$$
 (30)

We turn now to the B-field strength F_{AB}^{ij} , and we derive at order θ^0 :

$$
F_{rs}^{ij} = e_r^{\mu} e_s^{\nu} (\partial_{\mu} b_{\nu j}^{ij}) + \frac{1}{\sqrt{2}} \bar{\psi}_r^{i} \psi_s^{j1},
$$

\n
$$
F_{ai,s}^{ij} = 0 , \quad F_{ak,be}^{ij} = -\frac{1}{\sqrt{2}} C_{ab} \delta_k^{i} \delta_e^{j1}.
$$
\n(31)

Notice that the right-hand side of the first relation is the supercovariant curl of δ [cf. the supersymmetry variation of b in (20).

III. CONSTRAINTS IN TWO-COMPONENT FORMALISM

We can now rewrite the constraints using the explicit representation of γ matrices given in the Appendix, and we immediately get for their dotted and undotted components

$$
T_{mn}^r = T_{Ai_1, Bj}^r = T_{Ai_1, Bj}^r = 0,
$$

\n
$$
T_{Ai_1, Bj}^r = -\frac{1}{4} (i\sigma^r)_{AB} \delta_{ij},
$$

\n
$$
T_{Ai_1, Bj}^{O_k} = -\frac{3}{8} \epsilon_{AB} \delta_{ij}^k \lambda_{j1}^C - \frac{1}{8} \delta_{(A}^C \epsilon_{B)D} \delta_{(i}^k \lambda_{j1}^D),
$$

\n
$$
T_{Ai_1, Bj}^{O_k} = \frac{1}{2} \delta_{B}^C (\delta_{ij} \delta^{k\ell} - \frac{1}{2} \delta_{j}^k \delta_{i}^{\ell}) \lambda_{\hat{A}i},
$$

\n
$$
T_{Ai_1, Bj}^{O_k} = \frac{1}{2} \delta_{\hat{A}}^{\hat{C}} (\delta_{ij} \delta^{k\ell} - \frac{1}{2} \delta_{i}^k \delta_{j}^{\ell}) \lambda_{\hat{B}l},
$$

\n
$$
T_{Ai_1, Bj}^{O_l} = 0, \quad T_{Ai_1, Bj}^{O_k} - \frac{1}{2} \delta_{B}^C T_{Ai_1, Dj}^{Dk} = 0,
$$

\n
$$
T_{Ai_1, Bj}^{O_l} = 0,
$$

\n
$$
T_{Ai_1, Bj}^{i\ell} = F_{Ai_1, Bl}^{i\ell} = F_{Ak_1, S}^{i\ell} = F_{Ak_1, S}^{i\ell} = 0,
$$

\n
$$
F_{Ai_1, Bl}^{i\ell} = -\frac{1}{\sqrt{2}} \epsilon_{AB} \delta_{k}^{[i} \delta_{k}^{j]},
$$

\n
$$
F_{Ai_1, Bk}^{i\ell} = -\frac{1}{\sqrt{2}} \epsilon_{AB} \delta_{k}^{[i} \delta_{k}^{j]},
$$

 $F_{m,n}^{ij}$ = supercovariant photon curl.

Note that $\lambda_{B\ell}$ can be replaced by $2T^{\dot{D}\dot{k}}_{\dot{B}B\dot{B}}$, as we see from the expression of $T_{Ai,Bj}^{\zeta_k}$. Hence, λ_{Bk} is the $\theta = 0$ component of a superfield T_{BE} .

Note added. Breitenlohner and Sohnius in Ref. 4 have also derived a set of constraints on supertorsions and supercurvatures in superspace with (gauged) SU_2 internal symmetry. The torsion constraints read

$$
T_{ai, bj}^{\prime} = \frac{1}{4} \epsilon_{ij} (C \gamma_{5} \gamma^r)_{ab} , T_{ai, m}^{\prime} = 0 ,
$$

 $T_{ai, bj}^{j}$ should not contain spin or isospin $\frac{3}{2}$.

$$
(\overline{\tau}\sigma_{mn}C^{-1}\gamma_5)^{ai.~bj}T_{bj.ai.}^{~ck}=0\ ,\quad T_{mn}^{\quad \ *}=0
$$

$$
F_{ai.~bj}=\epsilon_{ij}(C\gamma_5)_{ab}\exp(C)\ .
$$

They do not impose a differential constraint and state that their formulation is equivalent to that of Refs. ² and 3, the difference being a redefinition of fields and connections.

Recently, J. %ess replaced his differential constraint in Eq. (5) by

$$
D_i^C T_{Ci, \dot{B}j}{}^{\dot{B}k} = D_i^C T_{\dot{C}i, \dot{B}j}{}^{Bk}
$$

With Eqs. (13) and (15) one may verify that this constraint is not in agreement with our results. However, we can propose an alternative to Eq. (5) which is satisfied. We find

$$
\epsilon^{i\ell} [(D_{\ell}^{A} - 2T_{m,\dot{B}l,\dot{B}m})T_{Ai,\dot{C}j,\dot{C}_{k}}] + \epsilon^{i\ell} [(\overline{D}_{\ell}^{\dot{A}} - 2T_{m,\dot{B}l,\dot{B}m})T_{\dot{A}i,\dot{C}k,\dot{C}_{j}}] = 0
$$

To lowest order in θ this equation merely implies the reality of S and P_{ij} .

Stelle and West (Ref. 21) applied an algebraic method of determining constraints to $N=2$ super-

space supergravity. All constraints were deduced. with the exception of the differential constraint.

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APPENDIX: TWO-COMPONENT FORMALISM

We give here our conventions for two-component spinors, and the explicit representation of γ matrices used to translate from four-component to two-component formalism:

$$
\gamma_{k} = \begin{pmatrix} -i\sigma_{k} \\ i\sigma_{k} \end{pmatrix}, \quad \sigma_{k} \text{ are Pauli matrices}
$$

$$
\gamma_{4} = \begin{pmatrix} 1 \\ 1 \end{pmatrix}, \quad \gamma_{5} = \begin{pmatrix} 1 \\ 1 \end{pmatrix},
$$

$$
X^{A} = \left(\frac{1+\gamma_{5}}{2}\right)X, \quad X_{\tilde{A}} = \left(\frac{1-\gamma_{5}}{2}\right)X,
$$

$$
X^{\tilde{A}} = (X^{A})^{*}, \quad X_{A} = (X_{\tilde{A}})^{*}.
$$

Lorentz transformations on spinors

$$
X'_{a} = (e^{(1/2)\omega \cdot \sigma})_{a}^{b} X_{b} ,
$$

\n
$$
X'^{A} = \{ \exp[(\vec{\omega} + i\vec{\nu})_{i}\vec{\sigma}] \} {^{A}}_{B} X^{E}
$$

\n
$$
X'_{A} = \{ \exp[(\vec{\omega} - i\vec{\nu})_{i}\vec{\sigma}] \} {^{A}}_{A}^{B} X_{B}
$$

\n
$$
\omega_{k} = \frac{1}{4} \epsilon_{ijk} \omega^{ij} , \quad v^{k} = \frac{i\omega^{k4}}{2} .
$$

Invariants:

$$
\epsilon^{AB}, \epsilon_{AB} (\epsilon_{12} = 1),
$$

\n
$$
\epsilon^{\dot{A}\dot{B}}, \epsilon_{\dot{A}\dot{B}} (\epsilon_{\dot{1}\dot{2}} = -1),
$$

\n
$$
\delta^A_B, \delta^{\dot{A}}_{\dot{B}}.
$$

Contraction rule:

$$
X_A \equiv X^B \epsilon_{BA}, \quad X^A \equiv \epsilon^{AB} X_B
$$

and similar for dotted indices. From the definitions

$$
(\sigma_{\mu})^{A\dot{B}} \equiv (\vec{\sigma}, i1)^{A\dot{B}},
$$

$$
(\sigma_{\mu})_{\dot{A}B} \equiv (\vec{\sigma}, -i1)_{\dot{A}B},
$$

it follows that $(\sigma_\mu)_{A\dot{B}} = (\sigma_\mu)_{\dot{B}A}$. Conversely, given $(\sigma_{\mu})_A^B$ and defining $(\sigma_{\mu})_{AB}$. Conversely, gives
 $(\sigma_{\mu})^A$ and defining $(\sigma_{\mu})_{AB} = (\sigma_{\mu})_{BA}$, where $(\sigma_{\mu})_{BA}$
 $=(\sigma_{\mu})^{C\delta} \epsilon_{CB} \epsilon_{DA}$ one finds the quoted result for
 $(\sigma_{\mu})_{AB}$. Since the $(\sigma_{\mu})_{AB}$ are suggest.

 γ matrices in two-component notation:

$$
\gamma_{\mu} = \begin{pmatrix} -i(\sigma_{\mu})^{A\hat{B}} \\ i(\sigma_{\mu})_{\hat{A}B} \end{pmatrix}, \quad C = \gamma_{4}\gamma_{2} = \begin{pmatrix} \epsilon_{AB} \\ \epsilon_{\hat{A}\hat{B}} \end{pmatrix},
$$

$$
\sigma_{\mu\nu} = \frac{1}{4} [\gamma_{\mu}, \gamma_{\nu}] = \begin{pmatrix} \frac{1}{4} [\sigma_{\mu}, \sigma_{\nu}]^{A}{}_{B} \\ \frac{1}{4} [\sigma_{\mu}, \sigma_{\nu}]_{\hat{A}}^{\hat{B}} \end{pmatrix}.
$$

Note the useful properties (due to $\epsilon_{12} = -1$)

$$
(\sigma_{\mu})_{\hat{A}B}(\sigma^{\nu})^{BA} = 2\delta^{\nu}_{\mu},
$$

\n
$$
(\sigma_{\mu})_{\hat{A}B}(\sigma^{\mu})_{\hat{C}D} = 2\epsilon_{\hat{A}\hat{C}}\epsilon_{BD},
$$

\n
$$
(\sigma_{\mu})_{\hat{A}B}(\sigma^{\mu})^{C\hat{D}} = 2\delta^{\hat{D}}_{\hat{A}\hat{B}}\delta^C_{B},
$$

\n
$$
(\sigma^{\nu})^{AB}(\sigma_{\mu})_{\hat{A}C} = \delta^{\nu}_{\mu}\delta^A_C + 2(\sigma^{\nu}_{\mu})^A_C
$$

Our conventions are further that $C\gamma_{\mu}C^{-1}=-\gamma_{\mu}^{T}$ and $\bar{\theta}_{\alpha} = \theta^{\beta} C_{\beta \alpha}$ for a Majorana spinor. [Since in a general representation $C^{-1} = -C$ does not always hold, one should not define $\theta^{\alpha} = C^{\alpha\beta} \overline{\theta}_{\beta}^{T}$. This is already obvious from the indices of $(\gamma_{\mu})^{\alpha}{}_{\beta}$ because then C has indices $C_{\alpha\beta}$ and not $C^{\alpha\beta}$.] Re-
quiring that $C_{\sigma\mu\nu}C^{-1}=-\sigma_{\mu\nu}^T$ (i.e., that θ transforms as $\theta^{\dagger} \gamma_4$) leads to the general solution $C = A(1 + \gamma_5)\sigma_2 + B(1 - \gamma_5)\sigma_2$ since X^A and X_A transform intependently, but invariance of the Dirac equation equates A to B .

The four-component contraction $\bar{\chi} \lambda = \chi^{\beta} C_{\beta \alpha} \lambda^{\alpha}$ becomes in two-component notation $-\chi^A \lambda_A - \chi^A \lambda_A$ $=X_A\lambda^A+\chi_A\lambda^A$ since $C=(\epsilon_{AB}, \epsilon^{\lambda B})$ and one always contracts as $\lambda^A \epsilon_{AB}$ and $\epsilon^{AB} \chi_B$.

- ¹S. Ferrara and P. van Nieuwenhuizen, Phys. Lett. 74B, 333 (1978); K. Stelle and P. C. West, *ibid*. 74B. 303 (1978).
- ${}^{2}E$. S. Fradkin and M. A. Vassiliev, Lett. Nuovo Cimento 25, 79 (1979); Phys. Lett. 85B, 47 (1979).
- ³B. de Wit and J. W. van Holten, Nucl. Phys. B155. 530 (1979); B. de Wit, J. W. van Holten, and A. van Proeyen, Leuven report 1979 (unpublished), in Supergravity, proceedings of the Supergravity Workshop at Stony Brook, 1979, edited by P. van Nieuwen-
- huizen and D. Z. Freedman (North-Holland, Amsterdam, 1979).
- ${}^{4}P$. Breitenlohner and M. Sohnius, Nucl. Phys. B165, 483 (1980).
- ⁵S. Ferrara and P. van Nieuwenhuizen, Phys. Rev. Lett. 37, 1669 (1976).
- ⁶D. Z. Freedman and A. Das, Nucl. Phys. B120, 221 (1977); E. S. Fradkin and M. A. Vassiliev, Lebedev Institute report (unpublished).

 ${}^{7}P$. K. Townsend and P. van Nieuwenhuizen, Phys. Lett.

67B, 439 (1977).

- $3\overline{J}$. Wess and B. Zumino, Phys. Lett. 79B, 394 (1978).
- ^{9}R . Arnowitt and P. Nath, Phys. Lett. 65B, 73 (1976). 10 L. Brink, M. Gell-Mann, P. Ramond, and J. H. Schwarz, Phys. Lett. 74B, 336 (1978).
- 1iL. Brink, M. Gell Mann, P. Ramond, and J. H. Schwarz, Phys. Lett. 76B, 417 (1978).
- ¹²S. Ferrara and P. van Nieuwenhuizen, Ann. Phys. (N, Y) (to be published); in Supergravity, proceedings .of the Supergravity Workshop at Stony Brook, 1979, edited by P. van Nieuwemhuizen and D. Z. Freedman (North-Holland, Amsterdam, 1979).
- $13S$. J. Gates, Jr. and J. A. Shapiro, Phys. Rev. D 18, 2768 (1978);V. Ogievetsky and E. Sokatchev, Phys. Lett. 79B, 222 (1978); Dubna report, 1978 (unpublished); Dubna report, 1979 (unpublished).
- '4P. van Nieuwenhuizen and P. C. West, Nucl. Phys. B169, ⁵⁰¹ (1980); S. J. Gates, Jr. , Nucl. Phys. 8162, 79 (1980).
- ¹⁵P. Howe and P. Tucker, Nucl. Phys. **B142**, 301 (1978).
- $¹⁶J.$ Wess, in Proceedings of the International Confer-</sup> ence on High Energy Physics, Geneva, Switzerland, 1979 (CERN, Geneva, 1979).
- 17_{In} the notation utilized by Wess, one has as a definition $\mathbf{D}_{\alpha}^{i} = \mathbf{D}_{i}^{\alpha}$. Given this definition it is simple to
- transcribe the equations of Ref. 16 into our notation. We thank W. Siegel for bringing this to our attention. 18 J. Wess and B. Zumino, Phys. Lett. $79B$, 394 (1978). ¹⁹S. J. Gates, Jr. (unpublished).
- 20 N. Dragon, Z. Phys. C 2, 29 (1979).
- 21 J. Wess and B. Zumino, Phys. Lett. $74B$, 51 (1978); W. Siegel, Nucl. Phys. B142, 301 (1978); S. J. Gates and J. A. Shapiro, Phys. Rev. ^D 18, ²⁷⁶⁸ (1978); W. Siegel and S. J. Gates, Nucl. Phys. B147 (1979); Harvard Report No. HUTP-79/A034, 1979 (unpublished); J. G. Taylor, Phys. Lett. 78B, ⁵⁷⁷ (1978); 79B, 399 (1978); Kings College Report No. 79-0679, 1979 (unpublished); M. Roček and U. Lindstrom, Phys. Lett. 83B, 179 (1979);V. Akulov, D. Volkov, and V. Soroka, Zh. Eksp. Teor. Fiz. Pis'maRed. 22, 396 (1975) fJETP Lett. 22, 187 (1975)]; Theor. Math. Phys. (USSR) 31, 12 (1977); V. Ogievetsky and E. Sokatchev, JINR Report No. E2-12511, Dubna, 1979 (unpublished); Y. Ne'eman and T. Regge, Riv. Nuovo Cimento 1, No. ⁵ (1978); the on-shell (i.e. , making no distinction between constraints and equations of motion) has been discussed for $N=3$ and $N=8$; $(N=3)$ L. Brink, M. Gell-Mann, P. Ramond, and J. H. Schwarz, Phys. Lett. 76B, 417 (1978); $(N=8)$ L. Brink and P. Howe, $ibid.$ $88B.$ 268 (1979); K. S. Stelle and P. C. West, Phys. Lett. B (to be published).