Phase transition in the σ model at finite temperature*

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We study the phase transition through which the spontaneously broken symmetry of the σ model is restored at finite temperature. The methods of nonrelativistic many-body theory, in which the equations of motion are approximated in a self-consistent manner, are applied to the σ model in 1 time and d space dimensions for $2 < d \leq 3$. We consider several different approximations of this type and discuss difficulties associated with their renormalization. The Hartree approximation predicts a second-order transition for all d , but breaks down at high temperatures when $d = 3$. The "modified Hartree approximation," a variant of Hartree theory which incorporates more of the effects of thermal fluctuations, predicts a first-order transition for all d. This result is shown to be an artifact of the approximation. The σ model with N fields [the O(N) model] is studied in the limit of large N. For $2 < d < 3$ this model undergoes a second-order transition whose critical exponents are computed to $O(1/N)$. When $d = 3$, however, the large-N approximation breaks down at high temperatures.

I. INTRODUCTION

The concept of spontaneously broken symmetry has come to play an increasingly important role in relativistic field theory' and descriptions of high-density matter.² As Kirzhnits and Linde³ first proposed, one might expect spontaneously broken symmetries in field theories to be restored at sufficiently high temperatures, analogous to the way in which nonrelativistic systems, such as superfluids, superconductors, and ferromagnets, in broken symmetry or "condensed" states, become symmetric or "normal" above a critical temperature, T_c . Such possible restoration of symmetry in the hot early universe has interesting cosmological consequences. ⁴ Finite-temperature phase transitions from broken symmetry to normal states also have importance for the behavior of hot high-density matter. For example, the pion-condensed state of neutron star matter should be destroyed at high enough temperature; an understanding of the condensation near the critical temperature is needed to assess the effects of pion condensation on the cooling of neutron stars.⁵ Also of interest are possible finite-temperature phase transitions from normal to abnormal nuclear matter.⁶

Subsequent to the suggestion of Kirzhnits and Linde, several authors⁷ have given descriptions of how finite-temperature field fluctuations can restore broken global as well as gauge symmetries; approximate calculations of critical temperatures approximate calculations of critical temperature
have also been presented.⁷ In this paper we further explore phase transitions in finite-temperature field theory, focusing on the transition expected in the σ model. This model is particularly interesting in the study of high-density matter

since it forms the basis of the description both of the pion-condensed state of neutron star matter and of abnormal nuclear matter. Limiting our considerations here only to states with no nucleons present, we shall study the extent to which various approximate calculational schemes reproduce the finite-temperature behavior physically expected in the model and the predictions they make about the order of the phase transition from broken symmetry to the normal state.

The σ model is described by the Lagrangian density

$$
L = -\frac{1}{2} \left[(\partial_{\mu} \sigma)^2 + (\partial_{\mu} \pi)^2 \right] + \frac{1}{2} m_0^2 (\sigma^2 + \pi^2)
$$

- $\frac{1}{4} \lambda_0 (\sigma^2 + \pi^2)^2$. (1.1)

Here σ and π are real fields, ${m_{\mathfrak 0}}^2$ is positive, and for the moment we take the π field to have only one component; the extension to three or more π fields is trivial. Because of the negative squared-mass term, $-m_0^2$, the complex field $\phi \equiv (\sigma + i\pi)/\sqrt{2}$ has a nonvanishing constant vacuum expectation value.

At finite temperature the behavior of the σ model is described in terms of thermal expectation values $\langle X \rangle$, defined for any operator X by

$$
\langle X \rangle = \mathrm{Tr}(e^{-H/T}X) / \mathrm{Tr}(e^{-H/T}), \qquad (1.2)
$$

where H is the Hamiltonian derived from (1.1) and temperature units are chosen so as to make Boltzmann's constant unity. The trace is over all states with the internal quantum numbers of the vacuum. At sufficiently low temperatures we expect the broken-symmetry state with thermal expectation value $\langle \phi \rangle$ nonzero to persist. We choose the phases of the fields so that $\langle \phi \rangle = \langle \sigma \rangle / \sqrt{2}$ is real; $\langle \pi \rangle$ is thus zero at all temperatures. The effect of thermal fluctuations is to decrease $\langle \sigma \rangle$ monotoni-

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FIG. 1. (a) Qualitative behavior of $\langle \sigma \rangle$ as a function of temperature. T_c is the system transition temperature. (b) Qualitative behavior of the masses m_{σ} and m_{π} below T_c , and the common mass $m_a = m_{\pi} = m$ above T_c .

cally with temperature from its vacuum value. The critical temperature, T_c , is defined by the vanishing of $\langle \sigma \rangle$, as in Fig. 1(a); the continuous decrease of $\langle \sigma \rangle$ to zero at T_c shown in that figure is characteristic of a second-order phase transition.⁸ Defining the masses m_{σ} and m_{π} at any temperature as the poles at wave vector \vec{k} = 0 of the thermal Green's functions for the σ and π excitations, respectively, we expect m_{σ} and m_{π} to behave as in Fig. 1(b). Below T_c , the π field is the Goldstone boson⁹ in the condensed state and one has $m_{\pi} = 0$. The mass m_{σ} is nonzero below T_{σ} . Since the symmetry between σ and π is restored at the critical temperature, m_a must go to zero at T_c , and $m_q = m_\pi$ above T_c .

Several authors⁷ have pointed out that the σ model at finite temperature should behave qualitatively in this way. Such behavior is characteristic of a wide variety of familiar nonrelativistic systems (e.g., superfluids, superconductors, and ferromagnets) which undergo second-order phase transitions. Indeed, one can invoke the principle of "universality" (which, crudely stated, asserts that the asymptotic behavior of systems very close to a second-order phase transition is determined purely by the space dimensionality and by the symmetry of the order parameter and of the Hamilton-

ian and is independent of the detailed dynamics¹⁰) to conclude that the critical exponents characterizing the asymptotic behavior of the σ model with $(N-1)$ π fields are identical to those of the classical N-component Heisenberg ferromagnet. This correspondence is a fruitful one since a great deal is known about the critical behavior of classical is known about the critical behavior of classical
magnets. In particular, Polyakov,¹¹ Migdal,¹² and Brezin and Zinn-Justin¹³ have recently shown that for $N>2$ the nonlinear σ model [i.e., λ , $m_0^2 \rightarrow \infty$ with m_a^2/λ remaining finite in the notation of Eq. (1.1)] becomes asymptotically free in two dimensions. Brezin and Zinn-Justin¹³ show that in $(2+\epsilon)$ dimensions with ϵ small, this model, whose critical properties are identical to those of the N-component Heisenberg ferromagnet, has the qualitative behavior shown in Fig. 1 with T_c given (in units of m_0^2/λ) by $T_c = 2\pi\epsilon/(N-2)+O(\epsilon^2)$. They compute the critical exponents for the nonlinear σ model to $O(\epsilon^2)$; the results should apply equally well to the linear σ model of Eq. (1.1).

Our goal in this paper is to find simple approximate treatments of the linear σ model that reproduce the qualitative behavior of Fig. 1 and agree as closely as possible with the exact results available in $(2+\epsilon)$ dimensions. Such approximate treatments can form a useful basis for future calculations of spontaneous symmetry breaking in high-density matter at finite temperature; the emphasis in such calculations is on having an approximately correct description over a wide range of temperatures, rather than an exact description of the critical region.

The approach we follow is to generate self-consistent approximations to the equation of motion for $\langle \phi \rangle$ and the Green's function equations derived from (1.1). Schematically written, these equations are $G_0^{-1}(\phi) = \lambda_0 \langle \phi^{\dagger} \phi \phi \rangle$ and $G^{-1} = G_0^{-1} - \Sigma$, where $G_0^{-1} = \Box^2 + m_0^2$. We approximate $\langle \phi^\dagger \phi \phi \rangle$ and the self-energy Σ in terms of $\langle \phi \rangle$ and the complete Green's function, G, and then solve the resulting equations for G and $\langle \phi \rangle$, thus obtaining expressions for m_{σ} , m_{π} , and $\langle \sigma \rangle$ as functions of temperature. One variant of this scheme was used by Kirzhnits and Linde 14 in their recent analysis of the finite-temperature phase transition in the σ the finite-temperature phase transition in the σ model. A similar approach was used by Chang,¹⁵ who showed that one can induce a phase transition in the σ model at zero temperature by varying the coupling constant. Techniques for generating selfconsistent approximations directly from the equations of motion have been used extensively in tions of motion have been used extensively in
nonrelativistic many-body theory,¹⁶ notably in models of superfluid helium and superconductors.

Such self-consistent approximation methods are particularly useful in problems such as the present one, where there is no small parameter. Simple approximations to $\langle \phi^{\dagger} \phi \phi \rangle$ and Σ correspond to the summation of whole classes of diagrams written in terms of the bare propagator G_0 . Moreover, by choosing these approximations in accordance with certain prescriptions, discussed in detail in the following section, one is guaranteed that the self-consistent solutions automatically obey the Goldstone theorem or preserve the conservation laws (i.e., the Ward identities).

In this paper we examine a number of such selfconsistent approximations for the σ model. The natural starting point is the familiar semiclassinatural starting point is the familiar semiclassi-
cal or "tree" approximation.¹⁷ This scheme is inadequate in that thermal fluctuations are completely neglected; symmetry breaking persists to arbitrarily high temperatures. Next we examine bitrarily high temperatures. Next we examine
Hartree theory,¹⁸ an approximation which incorporates fluctuation effects in the simplest possible way. Because of its (albeit crude) treatment of fluctuations, Hartree theory does predict a finite transition temperature and a second-order phase transition. Unfortunately the structure of this theory is trivial below T_c ; both the σ and π fields are free and massless.

Hartree theory provides the first illustration of a general difficulty encountered in formulating self-consistent approximations in relativistic systems: The ultraviolet properties of the theory can give rise to anomalies in the self-consistently computed Green's functions. In particular, we find that in the Hartree scheme in four space-time dimensions it is impossible to define the mass of the σ and π fields self-consistently at temperatures well above T_c . This failure is readily seen to be associated with the ultraviolet behavior of the theory; it is (both in Hartree theory and more generally) purely technical and does not, we hope, affect the critical (i.e., infrared) properties with which we are primarily concerned.

In order to circumvent this difficulty we consider the σ model in 1 time and d space dimensions, where *d* is taken to be a continuous parameter less than or equal to $3.^{19}$ So long as $d < 3$ the eter less than or equal to $3.^{19}$ So long as $d < 3$ the theory is superrenormalizable; once mass renormalization is performed it is ultraviolet convergent, in contrast to the situation in three dimensions where infinite wave-function and couplingconstant renormalizations are required as well. One finds, correspondingly, that no high-temperature breakdown of Hartree theory occurs when $d < 3$. The mass of the particles is well defined at all temperatures and the theory predicts a secondorder phase transition. Except for the vanishing of m_o below T_c , the behavior of the approximation agrees with that of Fig. 1.

In an attempt to obtain a finite m_o below T_c we next examine a slightly more complex variant of

Hartree theory, incorporating more fluctuation effects. We call this the "modified Hartree approximation"; aside from trivial numerical factors, it is the approximation of Kirzhnits and tors, it is the approximation of Kirzhnits and
Linde.¹⁴ We find that when $d < 3$ this scheme predicts a first-order, rather than a second-order, phase transition, that is, $\langle \sigma \rangle$ jumps discontinuously to zero at T_c . (Although Kirzhnits and Linde state that the transition emerging from their approximation is of second order, more detailed analysis indicates that their equations actually predict a first-order transition.)

When $d = 3$ we encounter a serious difficulty in the modified Hartree theory: Removal of divergences by conventional renormalizations at zero temperature does not leave the self-consistent calculation divergence-free at finite temperature. We are forced to relax the self-consistency requirement of the scheme somewhat in order to obtain Green's functions free of divergences. The resulting approximation predicts a first-order transition, in agreement with the $d < 3$ prediction.

Mean field calculations that predict first-order phase transitions are not uncommon in nonrelativistic theories. In order to gain more insight into the modified Hartree theory prediction of a first-order transition, we briefly review two such calculations: the treatment of a type-II superconductor including electromagnetic field fluctuations by Halperin, Lubensky, and Ma²⁰ (HLM), and the Bogoliubov approximation²¹ for the interacting Bose gas. The first-order transitions predicted by these calculations arise in a manner mathematically identical to that oi the modified Hartree approximation. Because the order parameter is coupled to a gauge field in the superconductor, one can argue that the prediction of a first-order transition in this system is quite plausible. However, in the Bose gas, which is very closely analogous to the σ model, such a prediction is an artifact of the approximation used and is certainly wrong.

We next try to attain a more precise understanding of the critical behavior of the σ model and to pinpoint the failure of the modified Hartree approximation by considering the $O(N)$ model, i.e., the σ model with $(N-1)$ π fields.²² In the limit N the σ model with $(N-1)$ π fields.²² In the limit N the σ model with $(N-1)$ π helds.⁻⁻ In the limit N
 $\rightarrow \infty$ the behavior of the O(N) model for $d < 3$ is ex-

actly calculable.²³ We exhibit the large-N soluactly calculable. 23 We exhibit the large–N solution; the O(N) symmetry is spontaneously broken at low temperatures and is restored at a critical temperature via a second-order phase transition. The critical exponents are identical to those obtained for the classical $O(N)$ model in the large- N limit²⁴ and are consistent with the exact results¹³ available in $2+\epsilon$ dimensions.

We then apply the modified Hartree approxima-

tion to the O(N) model for $d < 3$. In the $N \rightarrow \infty$ limit this approximation agrees qualitatively with the exact large-N results; in particular, the transition is predicted to be of second order. For finite N, however, the deficiencies of the modified Hartree scheme become clear. We first compute the $O(1/N)$ corrections to the exact $N = \infty$ results.²⁵ $O(1/N)$ corrections to the exact $N = \infty$ results.²⁵ As expected, the only effect of the $1/N$ terms is to modifythe critical exponents. The order of the transition remains unchanged. In contrast, the modified Hartree approximation is found to predict a first-order transition for *any* finite value of N. We infer that this prediction is indeed an artifact of the approximation.

Abbott, Kang, and Schnitzer²⁶ have argued that when $d = 3$ there is no spontaneous symmetry breaking at $T = 0$ in the large-N limit. Turning to finite temperature we find that for sufficiently small T the large- N theory contains no anomalies and, as one would expect, exhibits no symmetry breaking. For T sufficiently large, however, the large-N limit does develop anomalies; the common mass of the σ and π excitations becomes complex. Thus the large-N approximation, the most satisfactory of our simple mean field approximations to the critical behavior of the σ model when $d < 3$, fails at high temperatures when d =3.

The organization of this paper is as follows: In Sec. II we briefly review the self-consistent approximation methods of many-body theory, and examine the tree, Hartree, and modified Hartree approximations to the σ model for $d < 3$ and $d = 3$. Section III is a brief discussion of the HLM treatment of a superconductor in a fluctuating electromagnetic field and the Bogoliubov approximation for the interacting Bose gas. In Sec. IV we examine the $O(N)$ model at finite temperature; we study the large-N limit for $d < 3$ and $d = 3$ and include $O(1/N)$ corrections for $d < 3$.

II. MEAN FIELD THEORY

A. Basic definitions

In terms of the complex field $\phi = (\sigma + i\pi)/\sqrt{2}$ the Lagrangian density (1.1) becomes

$$
L = -\partial_{\mu} \phi^{\dagger} \partial^{\mu} \phi + m_{0}^{2} \phi^{\dagger} \phi - \lambda_{0} (\phi^{\dagger} \phi)^{2} . \qquad (2.1)
$$

It is convenient to introduce the shifted field $\tilde{\phi}(x)$, defined by

$$
\phi(x) = \tilde{\phi}(x) + \phi^c(x) , \qquad (2.2)
$$

where ϕ^c , the order parameter, is the expectation value of ϕ ; thus $\tilde{\phi}$ has vanishing expectation value at all temperatures. In the presence of a source term $(-h\phi^{\dagger} - h*\phi)$ in the Lagrangian, the equation $\hbox{term}~(-h\phi^\dagger-h^*\phi)$ in the
of motion for ϕ^c become:

$$
\Box^2 \phi^c = -m_0^2 \phi^c + \eta + h \quad , \tag{2.3}
$$

where

$$
\eta = 2\lambda_0 \langle \phi^\dagger \phi \phi \rangle \tag{2.4}
$$

For nonzero h the quantity $n(x)$ can be regarded as a functional of $\phi^c(x')$. Note that the functional $(\eta - m_0^2 \phi^c)$ is simply the functional derivative with respect to ϕ^{c*} of the effective potential, $V(\lbrace \phi^c \rbrace)$, frequently employed in the study of spontaneous frequently employed in the study of spontaneous
symmetry breaking.²⁷ In any state with $\phi^c(x)$ uniform, the equation of motion (2.3) implies that

$$
\eta / \phi^c = m_0^2 \tag{2.5}
$$

when $h = 0$. We shall always assume that under such circumstances the phases of the fields are chosen so that ϕ^c is real, i.e.,

$$
\langle \sigma \rangle = \sqrt{2} \phi^c, \quad \langle \pi \rangle = 0 \quad . \tag{2.6}
$$

B. 4-derivable and gapless approximations

As Kirzhnits and Linde¹⁴ proposed, one can study the phase transition of this model within the framework of a simple mean field theory. Starting from the equation of motion (2.3), one approximates the quantity η in terms of lower-order correlation functions. This procedure has been used extensively in the study of superfluids, where the various approximations that have been studied are typically classified as either " Φ -derivable" or "gapless."¹⁶

The " Φ -derivable" approximations are formulated in terms of the quantity Φ , a functional of $\phi^{c}(x)$, $\phi^{c*}(x)$, and of the (time-ordered) temperature Green's function matrix $G(xx')$, defined by

$$
G(xx') = -i \begin{pmatrix} \langle (\tilde{\phi}(x)\tilde{\phi}^{\dagger}(x'))_{+} \rangle & \langle (\tilde{\phi}(x)\tilde{\phi}(x'))_{+} \rangle \\ \langle (\tilde{\phi}^{\dagger}(x)\tilde{\phi}^{\dagger}(x'))_{+} \rangle & \langle (\tilde{\phi}^{\dagger}(x)\tilde{\phi}(x'))_{+} \rangle \end{pmatrix}
$$

$$
= \begin{pmatrix} \delta \phi^{c}(x)/\delta h(x') & \delta \phi^{c}(x)/\delta h^{*}(x') \\ \delta \phi^{c*}(x)/\delta h(x') & \delta \phi^{c*}(x)/\delta h^{*}(x') \end{pmatrix}.
$$
(2.7)

Given a particular approximation for Φ , one determines η and the self-energy matrix Σ through the equations

$$
\eta(x) = \frac{i}{2} \left(\frac{\delta \Phi}{\delta \phi^{\sigma*}(x)} \right)_{\phi^c, \underline{G}} , \qquad (2.8)
$$

$$
\Sigma_{\alpha\beta}(xx') = \left(\frac{\delta\Phi}{\delta G_{\beta\alpha}(xx')}\right)_{\phi^c, \phi^{c*}} \quad . \tag{2.9}
$$

The relations

$$
(G^{-1})_{\alpha\beta} = (G_0^{-1} + m_0^2)\delta_{\alpha\beta} - \Sigma_{\alpha\beta}
$$
 (2.10)

are then solved along with Eq. (2.3) to determine ϕ^c and G self-consistently. [Here $G_0(\overline{\mathbf{p}}\mathbf{z})$ repre-

sents the free, massless propagator, $(z^2 - \vec{p}^2)^{-1}$. The Φ -derivable approximations are extremely useful in many-body theory since the correlation functions derived from them preserve the conserthe contract in many local model phase are contracted
functions derived from them preserve the consequential points.²⁸ Moreover, having determined ϕ^c and G one can uniquely construct the effective poand G one can uniquely construct the effective po-
tential, $V(\lbrace \phi^c \rbrace)^{28}$ With approximations which are not Φ -derivable this is often impossible; typically one can construct two or more different effective potentials from ϕ^c and G. The absence of such ambiguity is a great virtue of the Φ -derivable approximations. Unfortunately, most of them violate proximations. Unfortunately, most of them violat
the Goldstone,⁹ or Hugenholtz-Pines,²⁹ theorem in the state of broken symmetry.

The "gapless" approximations are explicitly con-The "gapless" approximations are explicitly constructed to satisfy this theorem.¹⁶ One starts with an approximation to η as a functional of ϕ^c and G. The self-energies are then computed as

$$
\Sigma(xx')
$$
\n
$$
= \begin{pmatrix}\n(\delta \eta(x)/\delta \phi^c(x'))_{\phi^{c*}} & (\delta \eta(x)/\delta \phi^{c*}(x'))_{\phi^c} \\
(\delta \eta^*(x)/\delta \phi^c(x'))_{\phi^{c*}} & (\delta \eta^*(x)/\delta \phi^{c*}(x'))_{\phi^c}\n\end{pmatrix};
$$
\n(2.11)

the functional dependence of G on ϕ^c must be taken into account in performing this functional differentiation. Equations (2.3) and (2.10) are then solved self-consistently; the resulting Qreen's functions satisfy the Goldstone theorem. To see this we note that under a uniform gauge transformation $\phi(x) - e^{i \Lambda} \phi(x)$ one has $\phi^c(x) - e^{i \Lambda} \phi^c(x)$ and $\eta(x)+e^{i\Lambda}\eta(x)$. Thus when (2.11) is obeyed, the k $=0$ Fourier components of Σ satisfy

$$
\Sigma_{11}(k=0)\phi^{c} - \Sigma_{12}(k=0)\phi^{c*} = \eta . \qquad (2.12)
$$

For ϕ^c uniform and real we also have

$$
G_{\pi} = G_{11} - G_{12}, \quad G_{\sigma} = G_{11} + G_{12} \tag{2.13a}
$$

since under these conditions $\langle \sigma \pi \rangle = 0$; also,

$$
G_{\sigma}^{-1} = G_0^{-1} + m_0^2 - \Sigma_{11} - \Sigma_{12} ,
$$

\n
$$
G_{\pi}^{-1} = G_0^{-1} + m_0^2 - \Sigma_{11} + \Sigma_{12} .
$$
\n(2.13b)

Using (2.3), (2.12), and (2.13) we see that at $k=0$, $G_{\pi}^{-1}=0$ and thus, as expected, the π field is the Goldstone boson of the theory. (Note though that at finite temperature, where one has a preferred frame, G_{π}^{-1} does not vanish for general k with $k_{\mu}k^{\mu} = 0.$) Also, $G_{\sigma}^{-1}(k=0) = -2\Sigma_{12}(0)$.

C. Tree approximation

The familiar semiclassical (or tree) approximation is a "gapless" approximation which corresponds to taking

$$
\eta = 2\lambda_0 |\phi^c|^2 \phi^c \ . \tag{2.14}
$$

The self-energies are then given, according to (2.11), by

$$
\underline{\Sigma}(xx') = 2\lambda_0 \delta(x - x') \begin{pmatrix} 2|\phi^c|^2 & (\phi^c)^2 \\ (\phi^{c*})^2 & 2|\phi^c|^2 \end{pmatrix}.
$$
 (2.15)

From (2.5) , (2.14) , and (2.15) we immediately obtain

$$
\langle \sigma \rangle^2 = m_0^2 / \lambda_0 \tag{2.16a}
$$

$$
G_{\sigma}^{-1} = G_0^{-1} - 2m_0^2 \t{2.16b}
$$

$$
G_{\sigma}^{-1} = G_0^{-1}.
$$
 (2.16c)

We see from (2.16a) that symmetry breaking persists to arbitrarily high temperatures in the tree approximation. In order to observe the restoration of symmetry at finite temperature one must include effects of fluctuations. Any venture beyond the semiclassical treatment requires renormalizations. Since the renormalization counterterms in the Lagrangian are temperature independent, carrying out the various renormalizations to remove the infinites at any one temperature fixes the values of these counterterms uniquely. $(T = 0$ is a particularly convenient temperature at which to renormalize.) Kislinger and Morley³⁰ have recently argued how the renormalized theory which results is free from divergences at all temperatures. It is not true, however, that all approximations preserve this feature. Indeed we shall exhibit a very simple self-consistent approximation which cannot be rendered free of infinities at all temperatures by the standard renormalizations. Since our prime goal is to acquire some feeling for the critical (infrared) properties of the σ model, we shall temporarily avoid this complication by working in d space dimensions, where d is a continuous parameter less than 3. For any fixed $d < 3$ the theory is superrenormalizable; all infinities at $T = 0$ can be removed by mass renormalization alone and, at least for the elementary approximations we shall consider, no new divergences occur at finite temperature. We shall later return to the difficulties that arise when $d = 3$ and coupling-constant renormalization is required.

O. The Hartree approximation

Hartree theory, the simplest self-consistent scheme beyond semiclassical theory, amounts to approximating η as a functional of ϕ^c and G by

$$
\eta = 2\lambda_0 \left[|\phi^c(x)|^2 + \langle \tilde{\phi}^\dagger(x) \tilde{\phi}(x) \rangle \right] \phi^c(x) \,. \tag{2.17}
$$

This approximation can be " Φ -derived" from the functional

$$
\Phi = \frac{\lambda_0}{2i} \int dx \left[i \operatorname{Tr} G(xx) + 2|\phi^c(x)|^2 \right]^2 \tag{2.18}
$$

and the general prescription (2.8) and (2.9) ; then and from (2.19_b) ,

$$
[-m_0^2 + 2\lambda_0 |\phi^c|^2 + i\lambda_0 \operatorname{Tr} G(xx)]\phi^c = 0 \qquad (2.19a)
$$

and

$$
\Sigma_{\alpha\beta}(xx') = \delta(x - x')\delta_{\alpha\beta}
$$

$$
\times \lambda_0[2|\phi^e(x)|^2 + i \operatorname{Tr} G(xx)]. \quad (2.19b)
$$

In the broken-symmetry state

$$
\Sigma_{\alpha\beta} = \delta_{\alpha\beta}\delta(x - x')m_0^2;
$$

thus G_{σ} and G_{π} are free and massless. Since TrG $=2G_0$ and

$$
G_0(xx) = iT \sum_n \int \frac{d^d k}{(2\pi)^d} \frac{1}{(i\omega_n)^2 - \vec{k}^2}
$$
 (2.20)

(where $\omega_n = 2\pi nT$, $n = 0, \pm 1, \pm 2, \ldots$), (2.19a) determines $\langle \sigma \rangle$ as a function of temperature:

$$
\langle \sigma \rangle^2 = m_0^2 / \lambda_0 - 2 \int \frac{d^d k}{(2\pi)^d} \frac{1}{k} \left(\frac{1}{e^{k/T} - 1} + \frac{1}{2} \right). \tag{2.21}
$$

This equation shows how the thermal fluctuations of the field decrease $\langle \sigma \rangle^2$ and act toward restoring the symmetry. A trivial mass renormalization removes the (temperature-independent) infinity, no coupling-constant renormalization is required for $d < 3$, and (2.21) becomes

$$
\langle \sigma \rangle^2 = \mu^2 / \lambda_0 - 2a(d) T^{d-1}, \qquad (2.22)
$$

where μ^2 is the renormalized value of m_0^2 and $a(d)$ is the positive constant

$$
a(d) = \int \frac{d^d x}{(2\pi)^d} \frac{1}{x(e^x - 1)} \quad . \tag{2.23}
$$

The expectation value $\langle \sigma \rangle$ decreases monotonically with temperature until a critical temperature, T_c , defined by

$$
a(d)T_c^{d-1} = \mu^2/2\lambda_0 , \qquad (2.24) \qquad c(d) = \frac{2}{1-\lambda} \int \frac{d^d x}{(2\lambda)^d} \frac{1}{(2\lambda)^d}
$$

at which $\langle \sigma \rangle$ vanishes; $\langle \sigma \rangle^2$ decreases linearly with temperature as T approaches T_c from below. The phase transition is of second order and the critical exponent β [the power of $(T_c - T)$ with which $\langle \sigma \rangle$ vanishes as $T \rightarrow T_c$ equals $\frac{1}{2}$.

Note that for $d \leq 2$ the contribution of the thermal fluctuations of the field in (2.22) is infrared divergent since the constant $a(d)$ is then infinite. Such behavior is a specific realization of the general theorems³¹ that forbid the spontaneous breaking of .a continuous symmetry above zero temperature when $d \leq 2$. We restrict ourselves henceforth to $d > 2.$

Above T_c , (2.22) has no solution and (2.3) implies that ϕ^c vanishes identically; still $G_{\sigma} = G_{\pi} = G$, where

$$
G^{-1} = G_0^{-1} - m^2 \t{,} \t(2.25)
$$

$$
Tr G(xx)]\phi^c = 0 \qquad (2.19a) \qquad m^2 = -m_0^2 + 2i\lambda_0 G(xx) . \qquad (2.26)
$$

Evaluating $G(xx)$ as in (2.21) we have

$$
m^{2} = -m_{0}^{2} + 2\lambda_{0} \left(\int \frac{d^{d}k}{(2\pi)^{d}} \frac{1}{2\omega_{k}} + I^{(d)}(m) \right), \quad (2.27a)
$$

where

$$
\omega_k = (\vec{k}^2 + m^2)^{1/2} \tag{2.27b}
$$

and

$$
I^{(d)}(m) = \int \frac{d^d k}{(2\pi)^d} \frac{1}{\omega_k (e^{\omega_k/T} - 1)} \quad . \tag{2.27c}
$$

Equation (2.27a) is the d -dimensional version of Dolan and Jackiw's³² "gap equation," that describes the behavior of the $O(N)$ model above T_c in the N $\rightarrow \infty$ limit. It is not surprising that this equation emerges from the Hartree approximation: it is well known that the large- N approximation is identical to Hartree theory above the transition temperature.

After one has performed the same mass renormalization required in (2.21) , Eq. $(2.27a)$ becomes

$$
m^{2} + \lambda_{0} b (d) |m|^{d-1} = 2 \lambda_{0} I^{(d)}(m) - \mu^{2} , \qquad (2.28a)
$$

where

$$
b(d) = \frac{1}{d-1} \int \frac{d^d x}{(2\pi)^d} \frac{1}{(x^2+1)^{3/2}} \quad . \tag{2.28b}
$$

It is trivial to verify that for all $T > T_c$, Eq. (2.28a) has a positive solution, $m = m(T)$, which increases monotonically with T and vanishes at $T = T_c$. Since for small m we have

$$
I^{(d)}(m) = a(d)T^{d-1} - T|m|^{d-2}c(d) + \cdots, \qquad (2.29a)
$$

where

$$
c(d) = \frac{2}{d-2} \int \frac{d^d x}{(2\pi)^d} \frac{1}{(x^2+1)^2} , \qquad (2.29b)
$$

it follows that

$$
[m(T)]^2 \sim (T - T_c)^{2/(d-2)} \equiv (T - T_c)^{\gamma} \tag{2.30}
$$

as T approaches T_c from above. Hence the critical exponent γ that describes the behavior of m^2 above T_c is simply $2/(d-2)$ in the Hartree approximation.³³ tion.³³

E. The modified Hartree approximation: $2 < d < 3$

Hartree theory provides a Simple picture of how the thermal fluctuations of the fields in the σ model lead to a second-order phase transition. The theory is unphysical, though, in that the spectrum of σ -like excitations, described by G_{σ} , is that of a free massless particle for all $T < T_c$. By contrast, the tree approximation predicts a σ mass,

 $(2m₀²)^{1/2}$, independent of T, and no phase transi-

tion. An interesting question is whether one can construct simple approximations that describe the phase transition while incorporating a more realistic description of the spectrum of the fields below T_c . Let us therefore consider a slightly more complex variant of the Hartree approximation. Aside from insignificant numerical factors, it is the mean field approximation suggested by Kirzthe mean field approximation suggested by Kirz
hnits and Linde.¹⁴ This scheme, which we shal

refer to as the "modified Hartree approximation, " is a hybrid between the Φ -derivable and gapless methods.

We start with Φ given by (2.18); η , computed through Eq. (2.8) , has the same form (2.17) as in Hartree theory. At this point, however, we deviate from the Φ -derivable procedure and compute $\Sigma_{\alpha\beta}$ from (2.11), holding not only ϕ^c but also G constant during the functional differentiations to obtain

$$
\underline{\Sigma}(xx') = 2\lambda_0 \delta(x - x') \begin{pmatrix} 2|\phi^c|^2 + iG_{11}(xx) & (\phi^c)^2 \\ (\phi^{c*})^2 & 2|\phi^c|^2 + iG_{11}(xx) \end{pmatrix} .
$$
 (2.31)

This prescription differs from the gapless scheme where Σ is computed as the *total* functional derivative of η with respect to ϕ^c . Fortunately the Goldstone theorem is preserved 34 despite this heresy. When $\phi^{\texttt{c}}$ vanishes (above $T_{\texttt{c}}$) this approximation is identical to self-consistent Hartree theory, as is clear from Eqs. (2.19b) and (2.31). Below T_c , however, the structure is that of the tree approximation (2.15) plus the fluctuation term included in Hartree theory; choosing ϕ^c real we find

$$
G_{\sigma}^{-1} = G_{\sigma}^{-1} - m_{\sigma}^{2} , \qquad (2.32a)
$$

$$
G_{\pi}^{-1} = G_0^{-1} - m_{\pi}^2 \t\t(2.32b)
$$

where

$$
m_{\sigma}^{2} = 3\lambda_{0} \langle \sigma \rangle^{2} + \lambda_{0} i (G_{\sigma} + G_{\pi}) - m_{0}^{2} ,
$$
 (2.33a)

$$
m_{\pi}^{2} = \lambda_{0} \langle \sigma \rangle^{2} + \lambda_{0} i (G_{\sigma} + G_{\pi}) - m_{0}^{2}
$$

$$
= \eta / \phi^c - m_o^2 \tag{2.33b}
$$

Equation (2.5) implies that $m_\pi^{-2}\text{ = 0}$ in the state of broken symmetry and

 $m_{\sigma}^2 = 2\lambda_0 \langle \sigma \rangle^2$. (2.34)

Thus m_o is determined self-consistently from

$$
m_{\sigma}^{2} + 2\lambda_{0} i (G_{\sigma} + G_{\pi}) = 2m_{0}^{2} . \qquad (2.35)
$$

Evaluating G_{σ} and G_{π} as in (2.20) and (2.27a) and carrying out a single mass renormalization we obtain

$$
m_{\sigma}^{2} + 2\lambda_{0} \left[I^{(d)}(m_{\sigma}) - I^{(d)}(0) \right] - \lambda_{0} b(d) |m_{\sigma}|^{d-1}
$$

= $2\mu^{2} - 4\lambda_{0} a(d) T^{d-1}$ (2.36)

for $T < T_c$, where T_c , defined by the vanishing of m_{σ} , is again given by (2.24).

Now, however, a qualitatively new feature emerges: m_o does not decrease continuously to zero as T approaches T_c from below. To see this we note that $|m_{\sigma}|^{d-2}$, which occurs in the expansion (2.29a) of $I^{(d)}(m_q)$ for small m_q , is the dominant term on the left side of Eq. (2.36) as $m_a \rightarrow 0$. Thus

$$
c(d)|m_{\sigma}|^{d-2} = -2(d-1)T_c^{d-3}a(d)(T_c - T)
$$
 (2.37)

as T approaches T_c from below. The left and right sides of this equation have opposite sign; hence there is no solution below T_c of Eq. (2.36) for small, real m_{σ} .

The true behavior of m_{σ} near T_c in the modified Hartree approximation can be qualitatively understood by studying the effective potential, $V(\langle \sigma \rangle)$, whose derivative with respect to $\langle \sigma \rangle^2$ is simply $(\eta/\phi^c - m_o^2)/2$. In the state of equilibrium $V(\langle \sigma \rangle)$ is a minimum. Integrating (2.33b) with respect to $\langle \sigma \rangle^2$ using (2.34) we find for small $|\langle \sigma \rangle|$

$$
V(\langle \sigma \rangle) = V(0) + \lambda_0 a(d) (T^{d-1} - T_c^{d-1}) \langle \sigma \rangle^2
$$

-A | \langle \sigma \rangle |^d - B | \langle \sigma \rangle |^{d+1} + \lambda_0 \langle \sigma \rangle^4 / 4 , (2.38a)

where

$$
A = 2^{(d-2)/2} T c(d) \lambda_0^{d/2} / d ,
$$

\n
$$
B = 2^{(d-3)/2} b(d) \lambda_0^{(d+1)/2} / (d+1) .
$$
\n(2.38b)

A schematic plot of this function is shown in Fig. 2. It is clear from the figure that T_c is not the true transition temperature of the system. As expected, the absolute minimum of the effective potential occurs away from the origin for all temperatures less than T_c ; however, Fig. 2 shows that, owing to the presence of the $(-A|\langle \sigma \rangle|^d)$ term, this behavior persists even as T becomes greater than T_c . Only at a higher temperature, T_c^* , does the minimum at the origin become the absolute minimum of V. T_c^* is therefore the *true* transition temperature of the system. Furthermore, σ^* , the value of $\langle \sigma \rangle$ which minimizes V just below $T = T_c^*$, is nonzero; $\langle \sigma \rangle$ jumps discontinuously to zero at the transition temperature. Thus the modified Hartree approximation predicts a first-order, rather than a second-order, phase transition.

FIG. 2. Effective potential, $V(\langle \sigma \rangle)$, as a function of $\langle \sigma \rangle$ for the modified Hartree approximation with 2 < d < 3 and several different values of \boldsymbol{T} . From bottom to top these curves correspond to $T < T_c\,,\ T = T_c\,,\ T_c < T$ $\langle T_c^*, T = T_c^*$, and $T > T_c^*$, respectively. T_c^* is the true transition temperature of the system, and σ^* is the magnitude of the discontinuity in the equilibrium value of $\langle \sigma \rangle$ at T_c^* .

For all $T > T_c^*$, the gap equation (2.28a) still holds and has, as we have already noted, solutions m_{π} = m_{σ} = m . The qualitative behavior of m_{π} , m_{σ} , $\langle \sigma \rangle$, and *m*, as predicted by this approximation, is summarized in Fig. 3.

F. Hartree approximation in three dimensions

Before commenting further on the modified Hartree results let us consider the case $d=3$, where coupling-constant renormalization must be performed. We first examine the effects of this extra renormalization on the Hartree approximation. Above T_c the unrenormalized gap equation (2.27a) takes the form

$$
m^{2} = -m_{0}^{2} + 2\lambda_{0} I^{(3)}(m) + \frac{\lambda_{0} m^{2}}{8\pi^{2}} \left(1 + \ln\frac{m^{2}}{4\Lambda^{2}}\right) + \frac{\lambda_{0} \Lambda^{2}}{4\pi^{2}} ,
$$
\n(2.39)

where Λ is a high-momentum cutoff.

Defining the renormalized coupling constant, λ , and mass, μ^2 , by

$$
1/\lambda_0 = 1/\lambda - [\ln(4\Lambda^2/\mu^2) - 1]/8\pi^2
$$
 (2.40a)

and

$$
m_0^2/\lambda_0 = \mu^2/\lambda + \Lambda^2/4\pi^2 \,, \tag{2.40b}
$$

we are led to a gap equation free of infinities:

$$
m^{2}[1-\lambda \ln(m^{2}/\mu^{2})/8\pi^{2}] = 2\lambda I^{(3)}(m) - \mu^{2}. \quad (2.41)
$$

It is easy to verify that coupling-constant renormalization has no effect on Hartree theory below T_c , where the theory remains trivial (both the σ and π excitations are free and massless). From (2.21) and (2.40b) the condensate density varies with temperature according to

$$
\langle \sigma \rangle^2 = \mu^2 / \lambda - 2a(3)T^2 , \qquad (2.42a)
$$

and T_c is defined by

$$
T_c^2 = \mu^2/2\lambda a(3) \ . \tag{2.42b}
$$

Hartree theory in three dimensions presents the first example of an explicit breakdown of our approximation methods. To see this, note that the right side of (2.41) is positive at $m^2 = 0$ for $T > T_c$, and decreases monotonically with m^2 , approaching $(-\mu^2)$ as $m^2 \rightarrow \infty$; it also increases monotonically with T. The left side is zero at $m^2 = 0$, but, because of the $\ln(m^2/\mu^2)$ term, attains a maximum
at $m^2 = \mu^2 e^{-1 + 8\pi^2/\lambda}$ and goes to $-\infty$ for large m^2 thus for sufficiently large T Eq. (2.41) has no real solution. By contrast, the left side of the gap equation in Hartree theory for $2 < d < 3$ does not contain the $\ln(m^2/\mu^2)$ and increases monotonically with m , thereby ensuring that the equation has a solution for all T . The dominant behavior of the $\lambda \ln(m^2/\mu^2)$ for d=3 signals a breakdown of the approximation and the need to include higher-order correlations. This phenomenon is an interesting illustration of the interplay in relativistic theories of the infrared behavior, which determines critical properties, and ultraviolet effects.

FIG. 3. Qualitative behavior of m_{π} , m_{σ} , and $\langle \sigma \rangle$ as functions of temperature. T_c^* is the true transition temperature of the system. Above T_c , m_σ and m_π are identical.

G. Modified Hartree approximation in three dimensionsrenormalization difficulties

The gapless and Φ -derivable approximations are not simple expansions in the coupling constant, but rather involve summing selected subsets of diagrams in a self-consistent manner. In general, however, when one must include a coupling-constant renormalization, such schemes are not consistently renormalizable. This difficulty is illustrated by the modified Hartree approximation in three dimensions, where the field equation for T T_c becomes

$$
m_0^2/\lambda_0 = \langle \sigma \rangle^2 + I^{(3)}(m_0) + I^{(3)}(0) + \Lambda^2 / 4\pi^2
$$

+
$$
m_0^2 [1 + \ln(m_0^2 / 4\Lambda^2)] / 16\pi^2 . \qquad (2.43)
$$

Since Eq. (2.34) implies that the mass m_o is temperature dependent, the divergent term $m_{\sigma}^{2} \ln \Lambda^{2}$ in (2.43) cannot be removed by simple temperature-independent coupling-constant and mass renormalizations.

One way to circumvent these difficulties is to relax the self-consistency of the approximation so as to allow the removal of all infinities by conventional renormalization.³⁵ Let us write $\lambda_0 = \lambda + \delta \lambda$, tional renormalization.³⁵ Let us write $\lambda_0 = \lambda + \delta \lambda$, where λ is the renormalized coupling constant and $\delta\lambda$ is the corresponding counterterm, and treat the $\delta \lambda \phi^4$ term in tree approximation to obtain

$$
2^{1/2}\eta = \lambda_0 \langle \sigma \rangle^3 + i\lambda (G_{\sigma} + G_{\pi}) \langle \sigma \rangle . \tag{2.44}
$$

At this point we abandon the self-consistency requirement used previously to determine the masses m_{σ} and m_{π} and simply demand that below T_c these masses satisfy the tree approximation conditions

$$
m_{\sigma}^{2} = 2\lambda \langle \sigma \rangle^{2}, \quad m_{\pi}^{2} = 0 \tag{2.45}
$$

The second of these equations builds in the Goldstone theorem. The first is the simplest assumption that allows the theory to be sensibly renormalized and is consistent with the tree approximation and the expectation that m_{σ} and $\langle \sigma \rangle$ should vanish simultaneously as T_c is approached from below. With (2.45) we have, at $T=0$,

$$
2^{1/2} \eta - m_0^2 \langle \sigma \rangle = \lambda_0 \langle \sigma \rangle^3
$$

+
$$
\frac{\lambda \langle \sigma \rangle}{4\pi^2} \left[\Lambda^2 + \frac{1}{2} \lambda \langle \sigma \rangle^2 \left(1 + \ln \frac{\lambda \langle \sigma \rangle^2}{2\Lambda^2} \right) \right]
$$

-
$$
m_0^2 \langle \sigma \rangle .
$$
 (2.46)

Convenient definitions of the renormalized mass and coupling constant are

 ${m_0}^2 = \mu^2 + \lambda \Lambda^2 / 4\pi^2$, (2.47a)

$$
\lambda_0 = \lambda - \lambda^2 [1 + \ln(\mu^2 / 2 \Lambda^2)] / 8\pi^2 , \qquad (2.47b)
$$

whereupon the expression for $(2^{1/2}\eta/\langle \sigma \rangle - m_0^2)$ at

finite temperature becomes

$$
2^{1/2} \eta / \langle \sigma \rangle - m_0^2 = -\mu^2 + \lambda \langle \sigma \rangle^2
$$

+
$$
[\lambda^2 \langle \sigma \rangle^2 \ln(\lambda \langle \sigma \rangle^2 / \mu^2)] / 8\pi^2
$$

+
$$
\lambda [I^{(3)}((2\lambda)^{1/2} \langle \sigma \rangle) + I^{(3)}(0)]. \quad (2.48)
$$

As in the case $2 < d < 3$, we now integrate (2.48) to arrive at the following expression for the effective potential, valid when $\langle \sigma \rangle$ is small:

$$
V(\langle \sigma \rangle) = V(0) + \lambda a(3) (T^2 - T_c^2) \langle \sigma \rangle^2
$$

–[(2\lambda^3)^{1/2} Tc(3)/3] | $\langle \sigma \rangle^3$]
+ $\frac{\lambda}{4} \left[1 + \frac{\lambda}{8\pi^2} \left(\ln \frac{\lambda \langle \sigma \rangle^2}{\mu^2} - \frac{1}{2} \right) \right] \langle \sigma \rangle^4,$
(2.49)

where $T_c^2 = \mu^2/2\lambda a(3)$. The qualitative behavior of this function is identical to that of the effective potential for the case $2 < d < 3$, plotted in Fig. 2. Again, owing to the presence of the negative $|\langle \sigma \rangle^3|$ term in V, $\langle \sigma \rangle$ approaches a finite value as the transition temperature is approached from below. Thus the modified Hartree approximation for $d = 3$ predicts a first-order transition as well.

Above T_c^* , $\langle \sigma \rangle = 0$ and the common mass of the σ and π fields is determined by the formula

corresponding counterterm, and treat
\nterm in tree approximation to obtain
\n
$$
m^2(T) = \frac{\partial^2 V}{\partial (\sigma)^2}|_{(\sigma)=0}
$$
\n
$$
= \lambda_0 \langle \sigma \rangle^3 + i \lambda (G_{\sigma} + G_{\pi}) \langle \sigma \rangle
$$
 (2.50)

Thus $m^2(T)$ is positive at T_c^* and increases monotonically with T thereafter. The behavior of $\langle \sigma \rangle$, m_{σ} , m_{π} , and m as functions of T is qualitatively identical to that shown in Fig. 3.

To what extent is this prediction of a first-order transition to be believed? Mean field approximations are at best marginally reliable in their predictions about critical phenomena, and the modified Hartree scheme makes approximations even beyond canonical mean field treatments. For 2 $d < 3$ we sacrificed some of the self-consistency of the usual gapless methods by ignoring the dependence of the Green's functions on ϕ^c in computing the self-energy. For $d=3$ even more selfconsistency was sacrificed in order to make the theory renormalizable. Thus one should view the first-order transition that emerges from the modified Hartree approximation with a healthy dose of skepticism.

III. ANALOGIES IN MANY-BODY THEORY

Mean field treatments that predict first-order phase transitions are familiar in nonrelativistic field theories. In trying to ascertain whether the first-order transition of the last section is qualitatively correct or whether it is simply an artifact of the approximations used, it is useful to review some of these examples briefly.

A. Type-I superconductors

We shall first consider the type-I superconductor and then the weakly interacting Bose gas. According to BCS theory 36 the superconducting phase transition is of second order. Furthermore, owing to the very large zero-temperature coherence length in type-I superconductors, the effects of fluctuations are negligible until the temperature is so close to T_c that³⁷ | $T - T_c$ | $/T_c \sim 10^{-15}$.

Recently, HLM²⁰ showed that the presence of a fluctuating electromagnetic field in such systems can change the order of the transition from second to first. The coupling of the superconductor to a, vector potential $\vec{A}(\vec{r})$ is described in their work by the classical Ginzburg-Landau "free-energy functional"

$$
V(\{\psi,\vec{A}\}) = \int d^3r \left[a\left|\psi\right|^2 + \frac{1}{2}b\left|\psi\right|^4 + \gamma \left(\vec{\nabla} - iq_0\vec{A}\right)\psi\right|^2
$$

$$
+ \frac{1}{8\pi\mu_0} \sum_{i>j} \left(\frac{\partial A_i}{\partial r_j} - \frac{\partial A_j}{\partial r_i}\right)^2\right], \quad (3.1)
$$

where ψ is the superconducting order parameter, $a = a'(T - T_c^0)/T_c^0$, $q_0 = 2e/\hbar c$, μ_0 is the magnetic permeability of the normal metal, and a' , b , and γ are temperature independent constants near T_c . Thermal expectation values, e.g., $\langle \psi(\vec{r}) \psi(\vec{r}') \rangle$, are defined by the functional integral (x) = $\int d^3r [a|\psi|^2 + \frac{1}{2}b|\psi|^4 + \gamma] (\nabla - i q_0 \mathbf{A}) \psi]^2$
+ $\frac{1}{8\pi \mu_0} \sum_{i > j} \left(\frac{\partial A_i}{\partial r_j} - \frac{\partial A_j}{\partial r_i} \right)^2$, (3.1)

re ψ is the superconducting order parameter,
 $\psi'(T - T_c^0)/T_c^0$, $q_0 = 2e/\hbar c$, μ_0 is the mag

$$
\langle X \rangle = \frac{\int \delta \{\psi\} \delta \{\vec{A}\} X e^{-\mathbf{v}(\{\psi, \vec{\Lambda}\})/T}}{\int \delta \{\psi\} \delta \{\vec{A}\} e^{-\mathbf{v}(\{\psi, \vec{\Lambda}\})/T}}, \qquad (3.2)
$$

where X is any functional of ψ and $\vec{\mathbf{\Lambda}}.$

We may cast the HLM calculation in a form which emphasizes its similarity to our treatment of the σ model by first writing the Ginzburg-Landau field equation which follows from (3.1):

$$
-\gamma\langle(\vec{\nabla}-iq_0\vec{A})^2\psi\rangle + a\langle\psi\rangle + b\langle|\psi|^2\psi\rangle = 0.
$$
 (3.3)

For translationally invariant systems the $\vec{\tau}$ terms in (3.3) vanish. The neglect of fluctuations in ψ is equivalent to the tree approximation and reduces (3.3) to

$$
(\gamma q_0^2 \langle \vec{\mathbf{A}}^2 \rangle + a + b |\langle \psi \rangle|^2) \langle \psi \rangle = 0.
$$
 (3.4)

This equation is an exact analog of Eq. $(2.19a)$.

In the superconductor, the inverse, κ , of the penetration depth is defined by '

$$
\kappa^2 = 8\pi\mu_0 q_0^2 \gamma |\langle \psi \rangle|^2, \qquad (3.5)
$$

and plays the role of a photon mass; for small $|\langle \psi \rangle|$ the expectation value $\langle \vec{A}^2 \rangle$ is given by [cf. (2.27c) and (2.29a)]

$$
\langle \vec{A}^2 \rangle = 8\pi \mu_0 I^{(3)}(\kappa)
$$

\$\approx 8\pi \mu_0 T^2 a(3) - (32\pi \gamma q_0^2 \mu_0)^{1/2} \mu_0 T |\langle \psi \rangle| . (3.6)

Note that $\langle \vec{A}^2 \rangle$ is linear in $|\langle \psi | \rangle$, so that the field equation (3.4) has [as we found in our analogous Eq. (2.37)] no solution for arbitrarily small $|\psi\rangle$. As we have seen, such behavior indicates a firstorder phase transition. Again the effective potential, $V(\langle \psi \rangle)$, obtained by integrating the left side of the field equation has the form

$$
V(\langle \psi \rangle) = V(0) + a'(T - T_c) | \langle \psi \rangle |^2
$$

$$
- \frac{1}{6\pi} T(8\pi \mu_0 q_0^2 \gamma)^{3/2} | \langle \psi \rangle |^3
$$

$$
+ \frac{1}{2} b | \langle \psi \rangle |^4.
$$
 (3.7)

HLM are able to show that the transition in the superconductor occurs at a temperature, T_c^* , sufficiently above T_c [the temperature where a second-order transition would occur in the absence of the $|\langle \psi \rangle|^3$ term in (3.7)] that the fluctuations in ψ are completely negligible near T_c^* . Thus, for purposes of discussing the region near T_c^* , the initial neglect of the fluctuations in ψ is justifiable The prediction of a first-order phase transition in type-I superconductors is therefore on much firmer footing than is the analogous prediction for the σ model. The result for the superconductor derives from the electromagnetic field fluctuations, which are significant outside the extremely narrow region about T_c where the fluctuations in the order parameter are important. The σ model, on the other hand, has only one field: ϕ . When the loworder fluctuations in ϕ that we considered in the preceding section play a significant role, then fluctuation effects of arbitrarily high order become equally important. There is no justification for the omission of these higher-order fluctuations.

B. The weakly interacting Bose gas

The weakly interacting nonrelativistic Bose gas, which can be taken as a first model for superfluid helium, is more closely analogous to the σ model than is the superconductor. It is well known that the Bogoliubov approximation, 21 the analog in nonrelativistic Bose systems of the tree approximation in relativistic theories, provides a correct first description of the ground state and low-lying excitations of the weakly interacting Bose gas. However, when one tries to extend the approximation to describe the phase transition from the superfluid (broken symmetry) state to the normal state one finds, in a manner remarkably similar to that already observed for the σ model and the superconductor, a first-order transition. To see this, we start with the Hamiltonian

$$
\mathfrak{F} = \int d^3 r \left[\frac{1}{2m} (\vec{\nabla} \psi^{\dagger}) \cdot (\vec{\nabla} \psi) - \mu \psi^{\dagger} \psi + \frac{1}{2} \lambda \psi^{\dagger} \psi^{\dagger} \psi \psi \right],
$$
\n(3.8)

where *m* is the particle mass, μ the chemical potential, and ψ the field operator, and from it derive the equation of motion for the order parameter:

$$
(i\partial/\partial t + \vec{\nabla}^2/2m + \mu)\langle \psi \rangle = \lambda \langle \psi^{\dagger} \psi \psi \rangle . \qquad (3.9)
$$

The Bogoliubov approximation, like the tree approximation, starts with the factorization

$$
\langle \psi^{\dagger} \psi \psi \rangle = \langle \psi^{\dagger} \rangle \langle \psi \rangle^{2}. \tag{3.10}
$$

For a uniform, time-independent condensate, (3.9) then implies

$$
\mu = \lambda \, |\langle \psi \rangle|^2. \tag{3.11}
$$

The self-energy Σ [defined as in (2.11)] is computed by the gapless prescription (2.12), yielding

$$
\underline{\Sigma} = \lambda \delta(x - x') \begin{pmatrix} 2|\langle \psi \rangle|^2 & \langle \psi \rangle^2 \\ \langle \psi \rangle^{*2} & 2|\langle \psi \rangle|^2 \end{pmatrix} . \tag{3.12}
$$

Then

$$
G_{11}(\bar{\mathfrak{p}},z)=(z+\bar{\mathfrak{p}}^2/2m+\lambda\langle\psi\rangle^2)/(z^2-\omega_p^2),\quad (3.13)
$$

where $\omega_b^2 = \bar{p}^2 (\lambda \langle \psi \rangle^2 + \bar{p}^2 / 4m) / m$ and $\langle \psi \rangle$ is assumed real. The conservation law for the total number of particles in the system determines $\langle \psi \rangle$ as a function of T . Equation (3.13) implies

$$
\rho = \langle \psi \rangle^2 + \int \frac{d^d p}{(2\pi)^d} \frac{1}{\omega_p} \left[\frac{\bar{p}^2 / 2m + \lambda \langle \psi \rangle^2}{e^{\omega_p / T} - 1} + \frac{1}{2} \left(\frac{\bar{p}^2}{2m} + \lambda \langle \psi \rangle^2 - \omega_p \right) \right], \quad (3.14)
$$

where ρ is the total particle density and $\langle \psi \rangle^2$ represents the density of particles in the condensate. For small $\langle \psi \rangle$,

$$
\rho = \langle \psi \rangle^2 + f(d) (2m)^d / 2
$$

-g(d) m^{d/2} T λ ^{(d-2)/2} | $\langle \psi \rangle$ |^{d-2}
+... , (3.15)

where $f(d)$ and $g(d)$ are positive functions of the dimensionality. As we have seen before, the $|\braket{\psi}|^{d-2}$ term dominates for small $|\!\braket{\psi}|, \braket{\psi}$ does not vanish continuously as T approaches T_c [defined as the temperature where $\langle \psi \rangle = 0$ solves (3.14) from below, and the familiar first-order phase transition results. Again, the effective potential contains a negative $|\langle \psi \rangle|^d$ term, the harbinger of first-order transitions.

However, the conclusions to be drawn here are quite different from those in the superconductor since there is no temperature regime near T_c for superfluids where the neglect of fluctuations as in (3.10) is valid. There is no reason therefore to believe this mean field result. Indeed, the vast body of well-established theoretical and experimental evidence that the λ transition is of second

order indicates that the Bogoliubov prediction is simply wrong.

We believe the situation for the σ model to be rather similar to that for the imperfect Bose gas. The modified Hartree approximation has no apparent validity in the vicinity of the first-order transition it predicts. We have no general method for properly taking into account all of the important fluctuation effects. One can, however, acquire some feeling for the effect of higher order correlations in the σ model by considering a recorrelations in the σ model by considering a re-
lated theory, the O(N) model,²² which for large N does have a legitimate small parameter, viz. , $1/N$, near T_c . The existence of this parameter allows us systematically to include the effects of fluctuations and thus to see how a second-order transition results.

IV. THE O(N) MODEL

A. $2 < d < 3$

The O(N) model, a generalization of the σ model to N fields, is described by the Lagrangian density

$$
L = \frac{1}{2} \sum_{\alpha=1}^{N} \left[-(\partial_{\mu} \phi_{\alpha})^2 + m_{0}^2 \phi_{\alpha}^2 \right]
$$

$$
- \frac{1}{2} \frac{\lambda_{0}}{N} \left(\sum_{\alpha=1}^{N} \phi_{\alpha}^2 \right)^2 - \sum_{\alpha=1}^{N} h_{\alpha} \phi_{\alpha}.
$$
(4.1)

When m_0^2 > 0 and $h = 0$ we expect the O(N) symmetry to be spontaneously broken at low temperatures. As usual, the coupling constant is written as λ_0/N . so that the theory has a finite limit as $N \rightarrow \infty$. We consider dimension $2 < d < 3$, where difficulties arising from coupling-constant renormalization do not occur. It is trivial to generalize the methods of occur. It is trivial to generalize the methods of Coleman *et al.*³⁸ to finite temperature and so compute the effective potential in the $N \rightarrow \infty$ limit. We shall proceed slightly differently and compute the field equation and Green's functions to make contact with our earlier calculations.

We choose the "phase" of the condensate so that when $h = 0$ only $\langle \phi_1 \rangle$ is nonzero, and introduce the shifted fields $\tilde{\phi}_1 = \phi_1 - \langle \phi_1 \rangle$, $\tilde{\phi}_\beta = \phi_\beta$ for $\beta \neq 1$. In order to correspond with our previous notation we refer to the fields ϕ_{β} with $\beta \neq 1$ as π fields and (2/ $(N)^{1/2} \phi_1$ as the σ field. Then $\langle \phi_1 \rangle = (N/2)^{1/2} \langle \sigma \rangle$, where $\langle \sigma \rangle$ is of order unity. The equation of motion for $\langle \phi_{\alpha} \rangle$ is

$$
(\Box^2 + m_o^2) \langle \phi_\alpha \rangle
$$

= $h_\alpha + 2\lambda_o \sum_a (\langle \phi_\beta \rangle^2 \langle \phi_\alpha \rangle + \langle \tilde{\phi}_\beta^2 \rangle \langle \phi_\alpha \rangle)$

$$
+2\langle\bar{\phi}_{\beta}\bar{\phi}_{\alpha}\rangle\langle\phi_{\beta}\rangle+\langle\bar{\phi}_{\beta}^{2}\bar{\phi}_{\alpha}\rangle)/N. \quad (4.2)
$$

Setting $h = 0$ and taking $\alpha = 1$ in (4.2) we find the equation for $\langle \sigma \rangle$,

$$
+\left(\frac{2}{N}\right)^{3/2}\sum_{\beta}\left\langle \tilde{\phi}_{\beta} {}^{2}\tilde{\phi}_{1}\right\rangle, \qquad (4.3)
$$

since when $h = 0$ the Green's functions $G_{\alpha\beta}$ assume the form

$$
G_{\alpha\beta}(xx') = \delta_{\alpha\beta} [\delta_{\alpha 1} G_{\sigma}(xx') + (1 - \delta_{\alpha 1}) G_{\pi}(xx')].
$$
\n(4.4)

Let us now examine (4.3) in the $N \rightarrow \infty$ limit. The quantity $\langle \tilde{\phi}_B^2 \tilde{\phi}_1 \rangle$ in lowest order is $\gamma \chi_0 G_{\beta \beta}^2 G_{\alpha} \langle \phi_1 \rangle$ / quantity $\langle \varphi_{\beta} \varphi_{1'} \rangle$ in lowest of definition $\langle 4,3 \rangle$ thus be-
N and hence is $\gamma N^{-1/2} \langle \varphi \rangle$. Equation (4.3) thus becomes

$$
m_o^2 \langle \sigma \rangle = \lambda_o [\langle \sigma \rangle^2 + 2i G_\pi \langle xx \rangle] \langle \sigma \rangle \tag{4.5}
$$

for $N \rightarrow \infty$. This result, exact when N is strictly infinite, is identical in form to (2.19a) which was derived in the Hartree approximation for $N=2$.

The Green's functions may be evaluated by taking the first variation of (4.2) with respect to h:

$$
(\square^2 + m_o^2) G_{\pi}(xx') = \delta(x - x') + \lambda_o [\langle \sigma \rangle^2 + 2i(1 + N^{-1}) G_{\pi}(xx) + 2iN^{-1} G_{\sigma}(xx)] G_{\pi}(xx')
$$

+
$$
4 \lambda_o \left[\left(\frac{2}{N} \right)^{1/2} \langle \sigma \rangle \frac{\delta}{\delta h_2(x')} \langle \tilde{\phi}_1(x) \tilde{\phi}_2(x) \rangle + \frac{1}{2N} \sum_{\delta} \frac{\delta}{\delta h_2(x')} \langle \tilde{\phi}_\beta^2(x) \tilde{\phi}_2(x) \rangle \right].
$$
 (4.6)

The last term in this equation is $\sim (\lambda_0^2/N) G_{\beta\beta}^2 G_{\pi}^2$ plus higher-order terms, while the penultimate term is \sim 1/N; these terms can thus be neglected in the N $\rightarrow \infty$ limit, as can the terms multiplied by explicit factors $N = \infty$. In this same limit G_{α} is given by

of
$$
N^{-1}
$$
. Using the field equation (4.5) we see that in the broken symmetry state G_{π} is free and massless for $N = \infty$. In this same limit G_{σ} is given by
\n
$$
(\Box^2 + m_0^2) G_{\sigma}(xx') = \delta(x - x') + \lambda_0 [3 \langle \sigma \rangle^2 + 2i G_{\pi}(xx)] G_{\sigma}(xx') + (\frac{2}{N})^{1/2} \lambda_0 \langle \sigma \rangle \sum_{\beta} \frac{\delta}{\delta h_1(x')} \langle \tilde{\phi}_{\beta}^2(x) \rangle.
$$
\n(4.7)

The N – 1 terms $\delta \langle \tilde{\phi}_B^2(x) \rangle / \delta h_y(x')$ with $\beta \geq 2$ are evaluated through use of $\delta G_\pi = -G_\pi \delta G_\pi$ ⁻¹ G_π and the $N = \infty$ form for G_{π} ; this produces the integral equation

$$
\frac{\delta G_{\pi}(xx)}{\delta h_1(x')} = 2\lambda_0 \int d\,\overline{x} \, G_{\pi}(x\overline{x}) \, G_{\pi}(\overline{x}x) \left[\left(\frac{2}{N} \right)^{1/2} \langle \sigma \rangle \, G_{\sigma}(\overline{x}x') + i \frac{\delta G_{\pi}(\overline{x}\,\overline{x})}{\delta h_1(x')} \right]. \tag{4.8}
$$

Solving (4.8) by Fourier transformation and using (4.5) we finally arrive at the exact $N \rightarrow \infty$ result for G_{σ} in the broken symmetry state:

$$
G_{\sigma}^{-1}(\bar{\mathbf{p}},z) = z^2 - \bar{\mathbf{p}}^2 - \frac{2\lambda_0 \langle \sigma \rangle^2}{1 + 2\lambda_0 S_T(\bar{\mathbf{p}},z)},
$$
 (4.9)

where S_T is the Fourier transform of $-iG_{\pi}(x\bar{x}) G_{\pi}(\bar{x}x)$, given explicitly by

$$
S_T(\vec{p}, i\omega_n)
$$

=
$$
T \sum_{m} \int \frac{d^d k}{(2\pi)^d} \frac{1}{(\omega_m^2 + \vec{k}^2) \left[(\omega_m + \omega_n)^2 + (\vec{p} + \vec{k})^2 \right]}.
$$

(4.10)

Since G_{π} is free and massless, we see from Eq. (4.5) that $\langle \sigma \rangle$ is given as a function of T by Eq. (2.22), which we write as

$$
\langle \sigma \rangle^2 = 2a(d) \left(T_c^{d-1} - T^{d-1} \right) \qquad (T < T_c), \tag{4.11}
$$

with T_c defined in terms of the renormalized mass μ by (2.24). The behavior of $\langle \sigma \rangle$ is thus identical to that found in Hartree theory. The transition is of second order with $\beta = \frac{1}{2}$.

In contrast to the Hartree theory result, the σ Green's function for large N is nontrivial below T_c . The "mass" of the σ excitations is defined by the

pole of $G_{\sigma}(\vec{0},z)$ that approaches the origin as T $-T_c$. The leading behavior of $S_r(\bar{0}, z)$ as $z\rightarrow 0$ is given by the $m = 0$ term of (4.10) and is $\sim (z^2)^{(d-4)/2}$. (The branch cut in the z^2 plane is along the positive real axis.) Thus as $z \rightarrow 0$,

$$
G_{\sigma}^{-1}(\bar{0}, z) \sim z^{2} + i l \langle d \rangle \langle \sigma \rangle^{2} (1 - e^{\pi i d}) (z^{2})^{(4-d)/2} / T + \cdots,
$$
\n(4.12)

where $l(d) = 2^{d-1} \pi^{d/2-1} \Gamma(d/2)$. Owing to the presence of the massless Goldstone bosons in the theory, the σ particle is unstable. Let us look for a zero of (4.12) of the form $z^2 = R^2 e^{i\theta}$. The first sheet of the function G_{σ} corresponds to $0 < \theta < 2\pi$. G_{σ}^{-1} then has zeros of the form

$$
R = | 2 \sin(\pi d/2) \langle \sigma \rangle^{2} l(d) / T |^{1/(d-2)}, \qquad (4.13a)
$$

$$
\theta = (d \pm 4\nu) \pi/(d-2) , \qquad (4.13b)
$$

where ν can assume any integral value; $sin(\pi d/2)$ is negative since $2 < d < 3$. It is clear that no value of ν corresponds to a θ on the first sheet, as is of course guaranteed by unitarity, but there are solutions on higher sheets. The "mass" of the σ is proportional to $\langle \sigma \rangle^{2/(d-2)}$ and so vanishes like (T_c) $(-T)^{1/(d-2)}$ as the critical temperature is approached from below.

Above the transition, $\langle \sigma \rangle$ vanishes and the $N = \infty$ limit is identical to Hartree theory. The common mass m of the σ and π excitations is given by the gap equation (2.28a). We recall that $m(T) \sim (T)$ gap equation (2.26a). We recall that $m(T) \sim (T - T_c)^{1/(d-2)}$ as T approaches T_c from above. The $\frac{1}{c}$, as 1 approaches $\frac{1}{c}$ from above. In critical index $(d-2)^{-1}$ is the same as that found below T_c .

The leading term of the $1/N$ expansion gives rise, as we have just seen, to a second-order phase transition with no anomalies when $2 < d < 3$. In Sec. II we observed that the modified Hartree calculation predicted a first-order transition for the $N = 2$ model. It is instructive to apply the modified Hartree approximation to the $O(N)$ model for large N in an attempt to pinpoint the source of the first-order result. For arbitrary N the modified Hartree approximation corresponds to writing the nartree approximation corresponds
self-consistency condition for m_{σ}^{-2} as

$$
m_{\sigma}^{2} = 2m_{0}^{2} + 4i\lambda_{0}N^{-1}[G_{\sigma} + (N-1)G_{\pi}].
$$
 (4.14)

Recall that for $N=2$, G_{σ} gives rise to the term $\mid m_{\alpha}\mid^{d-2}$, which is ultimately responsible for the first-order transition. In the $N \rightarrow \infty$ limit, however, the offending term vanishes, leaving us with

$$
m_{\sigma}^{2} = 2\lambda_{0} \langle \sigma \rangle^{2}
$$

= 4\lambda_{0} a(d)(T_{c}^{d-1} - T^{d-1}), (4.15)

with T_c defined by (2.24). This result is *not* the correct $N \rightarrow \infty$ answer. The order parameter $\langle \sigma \rangle$ is correct $N \rightarrow \infty$ answer. The order parameter $\langle \sigma \rangle$ is given correctly, but, owing to the failure of-the mean field theory to include the momentum-dependent renormalization of the coupling constant arising from the S $_T$ term in (4.9), the expression for ${m_{\mathfrak o}}^2$ and the analytic structure of the σ propagato are hopelessly wrong. Nonetheless modified Hartree does predict a second-order transition for $N = \infty$, in qualitative agreement with the proper result.

As in Sec. II, (4.14) predicts a first-order transition for any finite value of N . The discontinuity in the condensate magnitude at T_c becomes smaller

 \rightarrow

with increasing N , disappearing when N becomes strictly infinite. It is interesting to compare this prediction with that obtained by calculating the $1/N$ corrections to the true $N \rightarrow \infty$ result. The critical, i.e., infrared, properties of the quantum $\mathrm{O}(N)$ theory are expected to be identical to those of the classical O(N) model, whose critical behavior classical O(N) model, whose critical behavior
has been computed to $O(1/N)$.²⁵ To this order in the classical theory, Brézin and Wallace³⁹ have proved the exponent scaling law

$$
2\beta = \gamma \big[d(2-\eta)^{-1} - 1 \big], \tag{4.16}
$$

where η is the critical index that describes the spatial decay of the two-point Green's function at the critical temperature. Ma's classical calculation²⁵ of γ and η to $O(1/N)$, together with (4.16), yields

$$
m_{\sigma}^{2} = 2m_{0}^{2} + 4i\lambda_{0}N^{-1}[G_{\sigma} + (N-1)G_{\pi}]. \qquad (4.14) \qquad 2\beta = 1 - \frac{4S_{d}}{N} \frac{2d-5}{d-2} + O\left(\frac{1}{N^{2}}\right), \qquad (4.17a)
$$

where

$$
S_d \!=\! -2 \bigl[\sin(\pi d/2)/\pi (d-2) B \bigl(\tfrac{1}{2} d - 1 \,, \tfrac{1}{2} d - 1 \bigr) \bigr]
$$

(4.17b)

and $B(x, y)$ is the β function: $\int_0^1 d\alpha \, \alpha^{x-1}(1-\alpha)^{y-1}$.

It is easy to verify that to $O(\epsilon/N)$ Eq. (4.17) is in complete agreement with tbe exact calculations of Brézin and Zinn-Justin¹³ on the classical nonlinear σ model in $2+\epsilon$ dimensions. One can readily check, moreover, that the $N = \infty$ result for T_c [which is identical to the Hartree theory T_c and is given in (2.24) vanishes linearly with ϵ as d approaches 2 from above. Aside from a trivial normalization factor this T_c is identical to that found by Brézin and Zinn-Justin.

We now briefly indicate how (4.17) follows directly from (4.3) and (4.6). The only term on the right side of Eq. (4.3) that remains to be computed to order $1/N$ is $G_r(xx)$. From (4.6) we obtain the $O(1/N)$ π self-energy,

$$
\Sigma_{\tau}(xx') = \lambda_0 \left\{ \delta(x - x') \left[\langle \sigma \rangle^2 + 2i(1 + N^{-1}) G_{\tau}^{(1)}(xx) + 2i N^{-1} G_{\sigma}(xx) \right] + 4(2N^{-1})^{1/2} \langle \sigma \rangle \delta \langle \tilde{\phi}_1(x) \tilde{\phi}_2(x) \rangle / \delta \langle \phi_2(x') \rangle \right. \\ \left. + 2\delta \langle \tilde{\phi}_3^2(x) \tilde{\phi}_2(x) \rangle / \delta \langle \phi_2(x') \rangle \right\} \tag{4.18}
$$

where $G_r^{(1)}$ denotes the π Green's function to $O(1/N)$. The penultimate term on the right side of (4.18) can be computed through an equation similar to (4.8); the $O(1/N)$ part of the final term is evaluated in the Appendix . The result is

$$
\Sigma_{\tau}(\vec{\mathbf{q}}, i\omega_{m}) = \lambda_{0} \left[\langle \sigma \rangle^{2} + 2i(1 - N^{-1})G_{\tau}^{(1)}(xx) + 2iN^{-1}G_{\sigma}(xx) \right]
$$

$$
-4N^{-1}T \sum_{n} \int \frac{d^{d}k}{(2\pi)^{d}} \frac{G_{\tau}(\vec{k} + \vec{\mathbf{q}}, i(\omega_{n} + \omega_{m}))G_{\tau}^{-1}(\vec{k}, i\omega_{n})G_{\sigma}(\vec{k}, i\omega_{n})}{1 + 2\lambda_{0}S_{T}(\vec{k}, i\omega_{n})} \right].
$$
(4.19)

Comparison of this result with the equation that follows from the substitution of (4.8) into (4.3) ,

$$
m_0^2 = \lambda_0 \left[\langle \sigma \rangle^2 + 2i(1 - N^{-1}) G_\pi^{(1)}(xx) + 2i N^{-1} G_\sigma(xx) - 4N^{-1} T \sum_n \int \frac{d^d k}{(2\pi)^d} \frac{G_\sigma(\vec{k}, i\omega_n)}{1 + 2\lambda_0 S_T(\vec{k}, i\omega_n)} \right],
$$
(4.20)

shows that

$$
\Sigma_{\tau}(\vec{\mathbf{q}}, i\omega_{m}) = m_{0}^{2} - 4\lambda_{0}N^{-1}T \sum_{n} \int \frac{d^{d}k}{(2\pi)^{d}} \left\{ \frac{G_{\sigma}(\vec{\mathbf{k}}, i\omega_{n})}{1 + 2\lambda_{0}S_{T}(\vec{\mathbf{k}}, i\omega_{n})} \left[G_{\tau}(\vec{\mathbf{k}} + \vec{\mathbf{q}}, i\omega_{n} + i\omega_{m}) G_{\tau}^{-1}(\vec{\mathbf{k}}, i\omega_{n}) - 1 \right] \right\}.
$$
 (4.21)

Since $\Sigma_{\tau}(\vec{0},0)$ is simply m_0^2 , this equation demonstrates to $O(1/N)$ that the π excitations are the Goldstone bosons of the O(N) model.

Expression (4.21) enables us to compute $G_{\epsilon}^{(1)}(xx)$ in terms of the known $N = \infty$ Green's functions G_{α} and G_{π} . Equation (4.20) then consists entirely of known quantities; after some straightforward algebra it becomes

$$
c(T_c - T) = \langle \sigma \rangle^2 \left(1 + \frac{8S_d}{N} - \frac{2d - 5}{d - 2} \ln \langle \sigma \rangle \right) \tag{4.22}
$$

for small $\langle \sigma \rangle$, where T_c is defined as the temperature where $\langle \sigma \rangle$ vanishes and c is a positive constant. To $O(1/N)$ this equation can be written as (T_c-T) $\sim \langle \sigma \rangle^{1/\beta}$, where β is the exponent defined in (4.17). Thus we arrive at the expected result: The inclusion of $1/N$ terms modifies the critical exponents but does not alter the order of the transition of the O(N) model.

Note that we were required to interpret Eq. (4.22) correctly to reach this conclusion. Had we retained in (4.22) only the $\langle \sigma \rangle^2 \ln \langle \sigma \rangle$ term, the dominant term on the right side for small $\langle \sigma \rangle$, we would have found that for $2d > 5$ the equation had no solution as $\langle \sigma \rangle$ approached zero; as in the modified Hartree approximation $cf. (4.14)$ we would have concluded that the transition was of first order. In (4.14) , the dominant term of $O(1/N)$ is actually a power of m_{σ} ; try as we might we cannot interpret it as a simple correction to the leading $(N = \infty)$ terms. In (4.22), on the other hand, the logarithmic $1/N$ term represents an obvious correction to the leading power. We expect inclusion of terms of higher order in $1/N$ merely to produce further corrections to the critical exponents.

The modified Hartree approximation completely neglects the $\langle \tilde{\phi}_1 \tilde{\phi}_2^2 \rangle$ term [the last term in (4.7)] of the σ self-energy, Σ_{σ} . In consequence Σ_{σ} is momentum independent and the logarithm which appears in the $1/N$ expansion is replaced by a power. A less drastic approximation to Σ_{σ} [such as the $N = \infty$ result of Eq. (4.9)] is evidently required if the order of the transition is to be given correctly.

B. $d=3$

In Sec. II we saw that simple Hartree theory, well-behaved when $2 < d < 3$, breaks down because of ultraviolet effects when $d = 3$., The breakdown was manifest in the absence of a solution of the gap equation (2.41) at sufficiently high temperature. Since Hartree theory and the $N \rightarrow \infty$ limit of the $O(N)$ model are intimately related, we might expect to find anomalies in the large-N approximation as well. Abbott, Kang, and Schnitzer²⁶ have argued that at zero temperature the $N \rightarrow \infty$ limit is free of anomalies and that spontaneous symmetry breaking does not occur. The ground state of the theory is $O(N)$ symmetric in the large-N approximation. It is straightforward to extend the work of these authors to finite temperature. The $O(N)$ symmetry is of course preserved for all T; the σ and π excitations are identical and are found to have a mass given by (2.41}. We conclude that the large-N approximation is consistent at low temperatures but does indeed break down at temperatures sufficiently large so that (2.41) has no solution.

Thus when $d = 3$ the thermodynamics of the $O(N)$ model cannot be reliably computed for all temperatures, even for large N. While we believe that spontaneous symmetry breaking does occur at low temperatures for small values of N and that the restoration of symmetry occurs via a second-order phase transition whose critical exponents are identical to those of the classical $O(N)$ model, we know of no simple approximation whereby this prejudice can be convincingly confirmed.

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APPENDIX: EVALUATION OF $\delta \langle \phi_3^2(x) \phi_2(x) \rangle / \delta \langle \phi_2(x') \rangle$

In this Appendix we compute to $O(1/N)$ the quantity

$$
I_{32}(xyx')\equiv\delta\langle\tilde{\phi}_3^{\ 2}(x)\tilde{\phi}_2(y)\rangle/\delta\langle\phi_2(x')\rangle
$$

which occurs in (4.18). Recalling that

$$
\langle \tilde{\phi}_3^2(x)\tilde{\phi}_2(y)\rangle = \frac{\delta G_3(xx)}{\delta h_2(y)}
$$

= $-\int dz \ d\overline{z} \sum_{\alpha\beta} G_{3\alpha}(xz) \frac{\delta \Sigma_{\alpha\beta}(z\overline{z})}{\delta h_2(y)} G_{\beta 3}(\overline{z}x)$ (A1)

and noting that to leading order in $1/N$,

$$
\Sigma_{\alpha\beta}(z\overline{z}) = \delta_{\alpha\beta}\delta(z-\overline{z})2\lambda_0 \sum_{\delta} \left[\langle \phi_{\delta}(z) \rangle^2 + iG_{\delta\delta}(zz) \right] / N \,, \tag{A2}
$$

we have \overline{a}

$$
\frac{\delta G_3(xx)}{\delta h_2(y)} = \int dz \sum_{\alpha} G_{3\alpha}(xz) G_{\alpha 3}(zx) \frac{2\lambda_0}{N} \sum_{\delta} \left[2 \langle \phi_{\delta}(z) \rangle G_{\delta 2}(zy) + i \frac{\delta G_{\delta \delta}(zz)}{\delta h_2(y)} \right], \tag{A3}
$$

whereupon it follows that

$$
I_{32}(xyx') = -\int dz G_{33}(xz)G_{33}(zx)\left[\frac{2\lambda_0}{N}\left(2\delta(z-x')G_{22}(zy) + (2N)^{1/2}\langle\sigma\rangle\frac{\delta G_{12}(zy)}{\delta\langle\phi_2(x')\rangle}\right) - 2i\lambda_0I_{32}(zyx')\right].
$$
 (A4)

The quantity $\delta G_{12}(zy)/\delta \langle \phi_2(x')\rangle$ is readily evaluated through an equation analogous to (4.8); to leading order we find that the Fourier transform, $\Gamma_3(\vec{k}, i\omega_n; \vec{p}, i\omega_m)$, of $\delta G_{12}(zy)/\delta \langle \phi_2(x') \rangle$ is

$$
\Gamma_{3}(\vec{k}, i\omega_{n}; \vec{p}, i\omega_{m}) = \frac{4\lambda_{0}}{(2N)^{1/2}} \langle \sigma \rangle \frac{G_{\sigma}(\vec{k}, i\omega_{n}) G_{\pi}(\vec{p}, i\omega_{m})}{1 + 2\lambda_{0} S_{T}(\vec{k}, i\omega_{n})}.
$$
\n(A5)

Equation (A4) then implies that the Fourier transform, $I_{32}(\vec{k}, i\omega_n; \vec{p}, i\omega_m)$, of $I_{32}(xyx')$ is simply,

$$
I_{32}(\vec{k}, i\omega_n; \vec{p}, i\omega_m) = \frac{4i\lambda_0}{N} \frac{S_T(\vec{k}, i\omega_n) G_\pi(\vec{p}, i\omega_m) G_\pi^{-1}(\vec{k}, i\omega_n) G_\sigma(\vec{k}, i\omega_n)}{1 + 2\lambda_0 S_T(\vec{k}, i\omega_n)}.
$$
(A6)

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