

## Evidence for SU(3) octet mixing\*

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The strong decays of the two strange axial-vector mesons  $Q_1(1289)$  and  $Q_2(1404)$  are examined within the context of SU(3). It is found that the decays can be successfully explained by treating the  $Q$ 's as mixed states of two pure  $C=+1$  and  $C=-1$  SU(3) octets. The vector-axial-vector-pseudoscalar S-wave coupling constants are calculated to be approximately 2.8 GeV for the  $A_1$  multiplet and 4.2 GeV for the  $B$  multiplet, and the mixing angle approximately  $48^\circ$ .

Evidence has recently been presented<sup>1,2</sup> supporting the existence<sup>3</sup> of strangeness-one axial-vector mesons in the region 1300–1400 MeV. These particles, called  $Q_1$  and  $Q_2$ , are observed from partial-wave analysis<sup>4</sup> to have masses 1289 and 1404 MeV, respectively. It is tempting to assign these particles to two different SU(3) octets whose  $I=1$  members are the  $A_1(1100)$  and  $B(1235)$ . The charge-conjugation parity of the neutral and non-strange members of the  $Q_1, A_1$  multiplet would be even, while that of the  $Q_2, B$  multiplet would be odd. This assignment would be favored by the SU(6)  $\otimes$  O(3) quark model,<sup>5</sup> which predicts two axial-vector multiplets with even and odd  $C$  parity.

However, the decays of the  $Q$ 's are not easily incorporated into a scheme with approximate SU(3) symmetry of the coupling constants. For example, the ratio  $\Gamma(Q_1 \rightarrow \rho K)/\Gamma(Q_1 \rightarrow K^* \pi)$  should be about 1/3 according to SU(3) and phase-space considerations; however, it is observed<sup>4</sup> to be at least 10, even with the most liberal interpretation of errors. A possible way around this difficulty would be to construct a model with SU(3)-symmetry breaking. In a model of SU(3) breaking via a  $\lambda_8$  spurion,<sup>6</sup> one finds that in order to suppress the  $Q_1 \rightarrow K^* \pi$  decay one needs large SU(3) breaking, that is, the SU(3)-breaking parameters are as large as the SU(3)-preserving ones. This is certainly not in accord with our previous experience<sup>7</sup> with SU(3).

The near degeneracy of the mean mass of the two axial-vector multiplets suggests the possibility of mixing between them. This idea was proposed some time ago by several authors<sup>8</sup> when the characteristics of the  $Q$ 's were much less well defined. If, as has been suggested, the two multiplets considered here have different  $C$  parities, then invariance of the strong interactions under  $G$  parity would dictate that only the  $Q$ 's, which are eigenstates of strangeness and therefore not of  $G$  parity, would mix. Thus, the axial-vector-vector-pseudoscalar ( $AVP$  henceforth) vertex involving the  $Q$ 's will contain both  $f$ - and  $d$ -type coupling.

The Lorentz-covariant decay amplitude for  $A_i \rightarrow V_j P_k$  is given by

$$T = g_S \epsilon_A \cdot \epsilon_V + g_D \epsilon_A \cdot p_V \epsilon_V \cdot p_A, \quad (1)$$

where the  $\epsilon$ 's and the  $p$ 's are the polarizations and momenta of the vector and axial-vector mesons. We express the mixing of the strange members of the  $A_1$  and  $B$  octets via the angle  $\gamma$ :

$$\begin{aligned} g_S(Q_1) &= if_{ijk} g_A^S \cos \gamma + d_{ijk} g_B^S \sin \gamma, \\ g_S(Q_2) &= -if_{ijk} g_A^S \sin \gamma + d_{ijk} g_B^S \cos \gamma, \end{aligned} \quad (2)$$

with similar expressions for  $g_D$ .

Helicity amplitudes proportional to those introduced by Colglazier and Rosner<sup>9</sup> are easily constructed from  $g_S$  and  $g_D$ :

$$\begin{aligned} H_0 &= [g_D m_A q^2 + (m_V^2 + q^2)^{1/2} g_S] / m_V, \\ H_1 &= g_S, \end{aligned} \quad (3)$$

where  $m_V$  and  $m_A$  are the masses of the vector and axial-vector mesons, and  $q$  is the center-of-mass momentum of the decay process. In terms of  $H_0$  and  $H_1$  the decay rates are given simply by

$$\Gamma(A \rightarrow VP) = \frac{q}{24\pi m_A^2} (H_0^2 + 2H_1^2). \quad (4)$$

A compilation of  $B \rightarrow \omega \pi$  data<sup>9</sup> yields  $g_B^D/g_B^S = -2.90$  GeV<sup>-2</sup>. While definitive data on the  $A_1$  multiplet are lacking, one may estimate  $g_A^D/g_A^S$  via a quark-model sum rule<sup>10</sup> involving the  $H$ 's:

$$2 \left( \frac{H_1}{H_0} \right)_{A \rightarrow \rho \pi} = \left( \frac{H_0}{H_1} \right)_{B \rightarrow \omega \pi} - 1. \quad (5)$$

We obtain  $g_A^D/g_A^S = 2.58$  GeV<sup>-2</sup>.

The decay processes to which we apply these formulas are listed in Table I. The small values for  $\Gamma(Q_1 \rightarrow K^* \pi)$  and  $\Gamma(Q_2 \rightarrow \rho K)$  imply that  $g_A \approx g_B$  and  $\gamma \approx 45^\circ$ . (These would be equalities if the two rates vanished.) These conditions also imply that  $\Gamma(Q_1 \rightarrow \rho K)/\Gamma(Q_2 \rightarrow K^* \pi) \approx 0.4$ , which is plausible if one includes the large systematic error in the  $Q_1 \rightarrow \rho K$  rate. The first five decay rates in the table are used for a minimum- $\chi^2$  fit, while the re-

maining three are predictions. The errors chosen for the minimization in the decays of the  $Q$ 's are the systematic ones. The  $Q_1$  decay rates depend critically on the  $Q_1$  mass, because of the small phase space available. Because there is a 25-MeV systematic uncertainty associated with the  $Q_1$  mass, we have chosen a value of 1300 MeV for our calculations. Raising or lowering the mass by 10 MeV changes the  $\rho K$  and  $\omega K$  rates accordingly by about 20%. We show the solutions found for the fit with and without the  $D$ -wave contribution included. In the latter case, there are two solutions with roughly the same  $\chi^2$ , so both are given. We note that it is a good first approximation to ignore the  $D$ -wave contribution. We have listed here only those solutions with positive coupling constants and mixing angle in the first quadrant. Other simple ambiguities exist due to choice of quadrant for  $\gamma$  and sign of  $g_A^S/g_B^S$ , but they yield the same results for the processes listed in Table I. As more data become available, one will hopefully be able to distinguish between these solutions.

Some  $A \rightarrow SP$  decays of the  $Q$ 's have also been observed. Simple calculation shows that these rates would be given by

$$\Gamma(Q_{1i} \rightarrow P_j S_k) = \frac{2}{3} \frac{q^3}{m_{Q_1}^2} \frac{(h_A d_{ijk} \cos\gamma + i h_B f_{ijk} \sin\gamma)^2}{4\pi}, \quad (6)$$

$$\Gamma(Q_{2i} \rightarrow P_j S_k) = \frac{2}{3} \frac{q^3}{m_{Q_2}^2} \frac{(-h_A d_{ijk} \sin\gamma + i h_B f_{ijk} \cos\gamma)^2}{4\pi},$$

where  $h_A$  and  $h_B$  are dimensionless coupling con-

TABLE I. Predicted and observed decay widths of the  $Q$ 's. Solution I, which includes the  $D$ -wave contribution, has  $g_A^S=2.78$  GeV,  $g_B^S=4.20$  GeV, and  $\gamma=47.8^\circ$ . Solutions II and III, with no  $D$  wave, have  $g_A=3.26$  GeV,  $g_B=3.57$  GeV,  $\gamma=54.7^\circ$  and  $g_A=2.85$  GeV,  $g_B=3.64$  GeV,  $\gamma=45.1^\circ$ . The first error in the observed-width column is statistical, while the second (in parenthesis) is systematic. The vector mixing angle is taken to be  $37.3^\circ$ . [D. H. Boal and R. Torgerson, Phys. Rev. D (to be published); R. Torgerson, Phys. Rev. D 10, 2951 (1974).]

Decay	Predicted width (MeV)			Observed width (MeV)
	I	II	III	
$Q_1 \rightarrow \rho K$	62.7	59.3	54.2	$145 \pm 9$ ( $\pm 70$ ) <sup>a</sup>
$Q_1 \rightarrow K^* \pi$	6.9	6.3	1.9	$5 \pm 3$ ( $\pm 5$ ) <sup>a</sup>
$Q_2 \rightarrow \rho K$	4.5	1.7	1.4	$2 \pm 1$ ( $\pm 2$ ) <sup>a</sup>
$Q_2 \rightarrow K^* \pi$	139	144	136	$140 \pm 4$ ( $\pm 15$ ) <sup>a</sup>
$B \rightarrow \omega \pi$	123	123	128	$125 \pm 10$ <sup>b</sup>
$Q_1 \rightarrow \omega K$	16.1	15.1	13.9	...
$Q_2 \rightarrow \omega K$	1.2	1.0	0.2	...
$A_1 \rightarrow \rho \pi$	158	184	140	$\approx 300$ <sup>b</sup>

<sup>a</sup> See Ref. 4.

<sup>b</sup> See Particle Data Group, Ref. 3.

stants defined analogously to  $g_A$  and  $g_B$ . The fact that the  $\pi\pi$  channel is more strongly coupled to  $Q_1$  than  $Q_2$  also supports a nonzero value for the mixing angle, although a numerical analysis is not yet possible.

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<sup>1</sup>G. W. Brandenburg *et al.*, Phys. Rev. Lett. 36, 703 (1976); 36, 706 (1976).

<sup>2</sup>G. Otter *et al.*, Nucl. Phys. B106, 77 (1976).

<sup>3</sup>For a review of previous work, see Particle Data Group, Rev. Mod. Phys. 48, S1 (1976); Yu. Antipov *et al.*, Nucl. Phys. B86, 365 (1975); S. Tovey *et al.*, *ibid.* B95, 109 (1975); G. Otter *et al.*, *ibid.* B93, 365 (1975).

<sup>4</sup>R. K. Carnegie *et al.*, talk given at the International Conference on High Energy Physics, Tbilisi, U.S.S.R., 1976 (unpublished).

<sup>5</sup>R. H. Dalitz, in *Proceedings of the XIIIth International Conference on High Energy Physics* (Univ. of California Press, Berkeley, 1967), p. 215; B. T. Feld, *Models in*

*Elementary Particles* (Blaisdell, Waltham, 1969), p. 372.

<sup>6</sup>See B. J. Edwards and A. N. Kamal, Phys. Rev. Lett. 36, 241 (1976) and references contained therein.

<sup>7</sup>See Ref. 5 for a review. For a recent discussion of the role of SU(3) breaking in radiative decays, see Ref. 6 and D. H. Boal, R. H. Graham, and J. W. Moffat, Phys. Rev. Lett. 36, 714 (1976).

<sup>8</sup>See Ref. 5 and E. W. Colglazier and J. L. Rosner, Nucl. Phys. B27, 349 (1971). For multiplets with the same C parity, see also P. G. O. Freund, Phys. Rev. Lett. 12, 348 (1964); R. H. Graham and J. W. Moffat, Phys. Rev. 184, 1905 (1969).

<sup>9</sup>V. Chaloupka *et al.*, Phys. Lett. 51B, 407 (1974); S. U. Chung *et al.*, Phys. Rev. D 11, 2426 (1975).

<sup>10</sup>See Colglazier and Rosner, Ref. 8.