Evidence for a four-dimensional $\mathcal{N} = 1$ integrable quiver in massive type IIA supergravity

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We are examining a newly discovered parametric family of backgrounds in massive type IIA supergravity that contains a warped AdS₅ factor. This family is the dual gravity description of four-dimensional quivers with $\mathcal{N} = 1$ supersymmetry. We are interested in the status of classical integrability in these theories, and we show that there exists a single choice of solutions that is special, while all other choices lead to nonintegrable quivers and chaotic string motion. By focusing on this special choice, we provide strong suggestive evidence for the integrability of the dual field theory based on analytic studies and extensive numerical analysis.

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I. INTRODUCTION

A. Motivation

Integrability is one of the cornerstones in our modern understanding of field theories. Integrable theories have a wealth of conserved quantities, and, more specifically, Liouville integrability is the statement that the motion of a dynamical system is confined to a submanifold of smaller dimensionality than its phase space. Not only that, but integrability is also a statement that the theory is exactly solvable for any value of the coupling constant. It should be obvious that these theories can be considered as beautiful toy models to provide us with intuition, since they admit exact and analytic solutions, on the structures of observables in the more interesting and challenging physics theories; see Ref. [1] for a more detailed exposition.

Using the holographic dictionary [2], we can relate a quantum field theory to the world-sheet description of string theory in an appropriately chosen background. This has a twofold importance in the framework of integrability. On the one hand, we can attempt to phrase the question of integrability in either side depending on which one appears to be the most tractable, and, on the other, integrable structures in string theory have a role on their own while also leading to new integrable gauge theories [3–5].

Even in the most basic and well-founded example of the AdS/CFT duality, the one that relates type IIB superstring theory in AdS₅ × S⁵ to the four-dimensional $\mathcal{N} = 4$ super Yang-Mills theory, the most prominent calculations rely on the integrability of the setup; see Ref. [3] for a detailed

account. Integrability of the classical string description, in this example, is manifest, since we can obtain the equations of motion that result from the Lagrangian description as a condition of flatness on the so-called Lax connection [6].

In spite of integrable theories having many coveted features, it is extremely cumbersome to declare a theory as being integrable, even at the classical level. The reason is that the proof of integrability relies on the existence of a Lax connection defined on the cotangent bundle of the theory. The sad reality is that there is no algorithmic and systematic way to construct that quantity. To make matters even worse, the precise statement is that there is no reason to determine, *a priori*, whether such a connection exists or not. Owing to this discussion, in order to generate integrable theories we resort, by and large, to appropriate deformations which preserve some structures of theories that are known to be integrable [7–9].

Even in two-dimensional models, which is some sense can be considered as the simplest ones, we lack a systematic approach to characterize the status of integrability. Indeed, hitherto we do not possess a classification of twodimensional integrable σ models, at least in all generality.

While this might seem to be discouraging at the very least, there has been a lot of progress in discovering new integrable theories with an explicitly known supergravity description that goes beyond the basic AdS_5/CFT_4 example mentioned above.

The $AdS_3 \times S^3 \times M^4$, with $M^4 = T^4$ or $M^4 = S^3 \times S^1$, enjoys integrability [10]. Furthermore, within the Gaiotto-Maldacena backgrounds [11], there exists a special choice, namely, the Sfetsos-Thompson solution [12,13], that is classically integrable [14]. Another example that enjoys integrability is coming from the AdS_7/CFT_6 paradigm,

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and the massive type IIA solution is of the form $AdS_7 \times S^2 \times I_z$ [15]. And the final explicitly known gauge-gravity dual pair for which we have a definite statement of classical integrability is the marginal Leigh-Strassler β deformation, when the deformation parameter is real. Its dual gravity description is given in terms of the Lunin-Maldacena solution [16], and the equivalence of the classical equations of motion of strings to the flatness of the Lax connection was derived in [17]. We should note, however, that there is a breakdown of integrability when the deformation parameter is complex [18–21].

From the list above, it is obvious that there is only one example of a classically integrable field theory in four dimensions and, with a minimum amount of supersymmetry, that of the real Leigh-Strassler deformation. This motivated us to look afresh for a new theory exhibiting these properties.

The story unfolds as follows: Recently, in [22] the authors derived twisted compactifications of the AdS₇ solutions developed in the series of papers [23–29] down to AdS₅. This is an example of a flow across dimensions that involves a twisted compactification. More specifically, the starting point was a family of six-dimensional $\mathcal{N} = (1,0)$ superconformal field theories parametrized by a function,¹ and, after their compactification on a two-dimensional manifold with constant curvature, at the end point of the flow there exists a new one-parameter family of strongly coupled four-dimensional $\mathcal{N} = 1$ superconformal field theories.

Regarding the statement of integrability: In the mother AdS_7 , we know that all, but a very specifically chosen, solutions are nonintegrable. For the special choice, we are certain on its classical integrability, since there exists an explicit Lax connection. While one might be tempted to argue that the same question is trivial in the daughter theories, this is far from being true. The twisted compactification that was performed in [22] was such that in the resulting theory, in the AdS₅ backgrounds, only $\mathcal{N} = 1$ supersymmetry is preserved. Since the twisted compactification of a given background is not a known method that preserves the integrability structures, whether we can discover signatures of integrability, let alone a proof of it, constitutes a highly nontrivial question.

B. The approach we undertake in this work

In light of the objective difficulties in spotting integrable structures, that we described above, and combining this with the fact that integrability has to be manifested in a universal manner for a given theory, reverse engineering the logic of the task at hand and, hence, searching for nonintegrable signatures appears as a ratiocination. To understand this dialectic approach to these matters, let us consider classical Hamiltonian systems. There, analytic nonintegrability is employing the use of Galois theory on differential equations, through which we can derive a conclusive statement on the structure of these systems. The statements of differential Galois theory on second-order, ordinary, linear differential equations were transformed to much simpler and straightforward algebraic form. Not only that, but also there exists an explicit algorithm producing the Liouvillian solutions of such equations, if any exist. This is the so-called Kovacic algorithm [30].

Now, we turn our attention to supergravity backgrounds. In order to perform the full-fledged analysis of the dynamics of classical strings, we would have to consider a generic string embedding and study the system of nonlinear partial differential equations that arise from the string σ model. A much more manageable approach is to study certain wrapped string embeddings, which are chosen in accord with the isometries of the background, and then analyze the equations of motion that result from that particular embeddings are chosen in such a way that they result in ordinary differential equations. This is a crucial requirement, in order to be able to apply the criteria of Kovacic's algorithm.

Since integrability is a property that has to be manifested universally in a theory, it means that each and every single string configuration has to result to integrable dynamics. Ergo, a single counterexample that is echoing signatures of nonintegrability is sufficient to allow us to declare the theory as being nonintegrable without any loss of generality. This is, in a nutshell, the method that has been dubbed *analytic nonintegrability*; see Ref. [31] for the first application in the critical string theory context.

We review the nuts and bolts of the above of the aforementioned method. To begin with, we write a string soliton that has \mathcal{Z} degrees of freedom and derive its equations of motion. Then, we proceed by finding simple solutions for $(\mathcal{Z} - 1)$ equations of motion. These simple solutions define the so-called invariant plane of solutions. We focus on the final equation of motion, and we replace the solutions found previously. Subsequently, we consider linearized fluctuations around this solution. In doing so, we arrive at a second-order, linear differential equation (NVE) and assumes the schematic form $\mathcal{A}_1 f'' + \mathcal{A}_2 f' + \mathcal{A}_3 f = 0$; see also [32].

In passing, and for completeness, we would like to mention that S-matrix factorization on the world-sheet theory of the string has been proven to be an invaluable tool to provide us with certain conditions of nonintegrability [33–36]. There is a nice connection between the analytic nonintegrability and S-matrix factorization approaches [37].

In the spirit of looking for signatures of nonintegrability, an equally important role is played by the standard arrows in our quivers when studying the dynamical evolution of

¹In principle, there are infinite choices for the function that parametrizes the field theory. In an abuse of language, we sometimes refer to quivers of this type as an "infinite family."

classical systems. That is, we can look for characteristic behaviors of a given system that are indicators of chaos. This can be implemented numerically in order to study quantities such as Lyapunov exponents, Poincaré sections, and power spectra.

At this point, it is worthwhile pausing for a moment to mention a subtle point. It has been noted in the past (see, for example, [32,38]), but we stress it for clarity. While it is not necessary that nonintegrability implies chaos, the converse is always true. More specifically, any chaotic system is nonintegrable by default. We would also like to point out that to the extent of our knowledge, in the context of the gauge-gravity duality, all systems that have been studied and found to be nonintegrable are chaotic as well. We are not certain what this is telling us about the dynamics of strings.

The ideas described above and dubbed analytic nonintegrability and studies of chaotic motion were first utilized, in a critical string theory setup, by Basu and Pando-Zayas in [31], building upon earlier studies of string solitons in the AdS-Schwarzschild [39]. Since then there have been many studies of various backgrounds using these ideas. We provide the readers a nonexhaustive but representative list of related works [32,38,40–59].²

C. What this method does and what it does not do

As we have said, the only way to prove integrability is to write the Lax connection [1]. An alternative way would be to connect directly to structures which are known to be integrable. However, since the latter structures are known to be integrable, effectively one obtains the Lax connection like so.

The methods that we use in this work, both the analytic and the numeric, can never prove integrability in a string background, not even in a single subsector of a given theory. As formal proof goes, these methods are well developed and very potent to prove nonintegrability; see also [32,38,42,44] for related explanations of the method.

In the case of absence of nonintegrability and/or chaos, the proper statement to be made is along the following lines.

After having carefully scrutinized all dynamical aspects of the system under examination, we find very strong suggestive evidence that the theory appears to be classically integrable.

The simple counterarguments against more powerful and robust statements are the following.

- (1) In the case of not spotting nonintegrable behavior analytically:
 - (a) The string soliton does not capture the breaking of the isometries, and, thus, it appears we are

describing a more symmetric situation than it actually is.

- (b) The string soliton is chosen only such that it leads to ordinary differential equations in order to apply Kovacic's algorithm. However, this is not the most general string embedding.
- (2) In the absence of chaos:
 - (a) We have already mentioned that nonintegrability does not necessarily imply chaos. Going to the other direction, this means that while a system is not chaotic, it could be nonintegrable. Hence, the absence of chaotic signatures is not enough to declare a theory integrable.

D. Comments on the field theory realization

We have already explained the main steps of which our method consists. It is clear that we need to consider a string soliton and study the resulting dynamics. Doing so, we end up obtaining an equation of motion for the time coordinate of the background spacetime. This has to be of appropriate form; namely, it has to give the energy of the string as its first integral. While this step might be characterized by most as nitpicking, it is highly important, in order to be able to argue, using the principles of the AdS/CFT duality, that a specific string configuration is a dual gravity description of a gauge invariant operator in the field theory [44]. This is necessary in order to be able to argue that the statement of (non)integrability is shared between the bulk and boundary pictures. With that in mind, we will always incorporate the time coordinate on the world sheet in the target spacetime coordinate by means of $t = t(\tau)$, and this is what enters the dynamics through the equations of motion derived from the σ model. Having said that, our string configuration possesses a well-defined holographic realization; see also [42,44] for related comments in different backgrounds.

To put it in simple terms, when resorting to the methods of analytic and numerical (non)integrability, we ought to be extremely careful about the consistency of our string embedding for any given (class of) supergravity background(s). It should be obvious from this discussion that the exact knowledge of the form of the operator is not strictly necessary. The same statement is true for the boundary superconformal field theory as well; see Ref. [42] for an application of the method without the knowledge of the specifics of the dual field theory. There is absolutely no need for the precise holographic dictionary to share the statement of nonintegrability so long as our string soliton is consistent in the way we described.³

Having already taken a small detour to discuss some aspects of the field theory interpretation, we also feel the need to add some commentary on the qualitative form of the boundary operators. We are using extended string

²It is worthwhile to also mention the very interesting approach of [60] that examined spectrum of anomalous dimensions and spotted signatures of chaos.

³This is up to the implicit assumption that the supergravity description has a well-defined global form.

configurations that wrapped in a nontrivial manner along some of the cyclic coordinates of the background. Using the standard holographic dictionary, the field theory operators that are the dual descriptions to these bulk states are long, unprotected operators with large quantum numbers; for example, they have large energy and/or angular momentum. In the case of the Klebanov-Witten model, this has been more thoroughly elaborated in [31], since we have a thorough understanding and control over the boundary superconformal field theory. In the limiting case where we take the wrapping of the string to be trivial, we have the low-energy description of the string which appears as a point particle moving along some geodesic paths.

E. Gaining intuition on discovering integrable theories

In the preceding discussion, we have clarified what the role and main steps of the analytic and numerical analyses of nonintegrability are in the context of examining the classical integrability of string motion. We have not pointed out, however, some intuition that has been built so far based on these studies that allow one to investigate more promising situations. In our opinion, this is the most important and exciting aspect in this kind of studies.

As we have mentioned, in this work—as in many of the previous ones in the literature—we study a family of supergravity solutions that has nontrivial warp factors. Each of these warp factors can be written in terms of a function, which we call $\alpha(z)$ to connect with the rest of the work, of a single variable. That variable is one of the internal dimensions of the background metric.

It was observed in the past that, in situations where the warp factor in front of the AdS part of the geometry can be set to be equal to a constant consistently, we are led to obtain a classically integrable solution. This is hand-wavy and intuitive way to look for new integrable holographic theories and is by no means a golden hammer.⁴ Examples that follow this rule are [14,15].

So, why is the above working and on what premise? The answer, in simple terms, is that, by undoing the warping in front of AdS, one is hoping to end up with a background whose geometry is the direct product of integrable subspaces.

F. The structure of this paper

The structure of this work is the following: In Sec. II, we review and describe the basic aspects of the supergravity solutions that are of interest to us. We proceed to derive the equations of motion that govern the dynamics in this setup in Sec. III. In Sec. IV, we present the supergravity solution that appears to be special in the context of classical integrability. Section V contains the results of our numerical approaches. We conclude in Sec. VI.

We have included the Appendix in an effort to make this paper clear and self-contained. It contains the steps of Kovacic's algorithm. In the physics context, the basic steps of Kovacic's algorithm were presented most prominently in Appendix A in [43]. We, however, opt to present case III of Kovacic's algorithm, which has not been presented before in the physics literature.⁵ We feel that it is important to mention the excellent reference [41] that offers exposition to both the group theoretic and also differential equation approach as well.

II. PRELIMINARIES

We provide a brief overview of the main aspects of the new infinite family of massive type IIA backgrounds that is the focus in this work. This solution was obtained in [22] and is a flow between a family of six-dimensional $\mathcal{N} = (1,0)$ and four-dimensional $\mathcal{N} = 1$ superconformal field theories. Based on the existence of a monotonic quantity that interpolates between the conformal points at high and low energies, there is a proposal for an appropriate quiver capturing the low-energy dynamics. These quivers were proposed to hit a conformal fixed point at low energies. In essence, it is their strongly coupled dynamics that is being described by the new infinite family of massive type IIA vacua with an AdS₅ factor.

It is worthwhile to comment on the following subtlety that was observed in [22]: The β functions and the R-symmetry anomalies of the proposed quiver are canceled. In addition to that, the free-energy scaling with the parameters of the quiver matches the result obtained from the holographic calculation. However, the precise freeenergy coefficient in the field theory calculation does not match, exactly, the supergravity prediction.

The physical interpretation of the above comment is that the proposed quiver theory is a good first step to the end of having the appropriate field theory dual of the infinite family of massive type IIA vacua, but improvements need to be made for a precise identification. This, however, does not affect the integrability analysis from the supergravity side that we undertake in this work.

A. The massive type IIA backgrounds

We proceed to present the solution that was obtained in [22] and will be the focus of our analyses. We stress that we are not studying the flow of the six-dimensional quivers, but rather we focus on the IR fixed point of the flow, which was obtained in [25]. As we have already discussed, we are interested in the status of integrability in the fourdimensional theory.

⁴A simple counterexample is the $AdS_5 \times SE_5$ solution, with SE_5 being any Sasaki-Einstein space, that is nonintegrable [31,32,40] in spite of having a trivial warp factor in front of AdS.

⁵To the extent of our knowledge.

The NS-NS (Neveu-Schwarz) sector of the solutions consists of the string-frame metric, the dilaton, and a nontrivial 2-form field. The metric is

$$ds^{2} = f_{1}ds^{2}_{AdS_{5}} + f_{2}ds^{2}_{\mathbb{H}^{2}} + f_{3}dz^{2} + f_{4}\left(ds^{2}_{S^{2}} + \cosh\theta_{1}\sin^{2}\theta_{2}ds^{2}_{2}\right),$$
(2.1)

with the various submanifolds that appear in Eq. (2.1) being given by

$$ds_{AdS_{5}}^{2} = -\cosh^{2}\rho dt^{2} + d\rho^{2} + \sinh^{2}\rho ds_{S^{3}}^{2}, \qquad ds_{\mathbb{H}^{2}}^{2} = d\theta_{1}^{2} + \sinh^{2}\theta_{1}d\phi_{1}^{2}, ds_{S^{3}}^{2} = d\omega_{1}^{2} + \sin^{2}\omega_{1}(d\omega_{2}^{2} + \sin^{2}\omega_{2}d\omega_{3}^{2}), \qquad ds_{S^{2}}^{2} = d\theta_{2}^{2} + \sin^{2}\theta_{2}d\phi_{2}^{2}, ds_{2}^{2} = \cosh\theta_{1}d\phi_{1}^{2} + 2d\phi_{1}d\phi_{2}.$$
(2.2)

The dilaton and the NS-NS 2-form are given, respectively, by

$$e^{4\Psi} = f_5, \qquad B_2 = (f_6 \sin \theta_2 d\theta_2 - \pi \cos \theta_2 dz) \wedge (d\phi_2 + \cosh \theta_1 d\phi_1). \tag{2.3}$$

The RR sector of the theory is not relevant for our studies; however, we provide for completeness the relevant expressions below:

$$F_{0} = f_{7},$$

$$F_{2} = f_{6}f_{7}\mathrm{vol}_{S_{c}^{2}} + f_{8}(\cos\theta_{2}\mathrm{vol}_{\mathbb{H}^{2}} - \mathrm{vol}_{S_{c}^{2}}),$$

$$F_{4} = f_{9}\cos\theta_{2}\mathrm{vol}_{\mathbb{H}^{2}} \wedge \mathrm{vol}_{S^{2}} + f_{10}\mathrm{sin}^{2}\theta_{2}dz \wedge d\phi_{2} \wedge \mathrm{vol}_{\mathbb{H}^{2}}.$$
(2.4)

In all of the above expressions, the functions f_i for the various values of i are functions of z, and their explicit form is

$$f_{1} = 12\pi \sqrt{\frac{6}{m}} \sqrt{-\frac{\alpha}{\alpha''}}, \qquad f_{2} = \pi \sqrt{\frac{6}{m}} \sqrt{-\frac{\alpha}{\alpha''}}, \qquad f_{3} = \pi \sqrt{\frac{3m}{2}} \sqrt{-\frac{\alpha''}{\alpha}},$$

$$f_{4} = \pi \sqrt{6m} \frac{\sqrt{-\alpha''\alpha^{3}}}{3m\alpha\alpha'' - 2\alpha'^{2}}, \qquad f_{5} = \frac{3me^{4\Psi_{0}}}{(2\alpha'^{2} - 3m\alpha\alpha'')} \left(-\frac{\alpha}{\alpha''}\right)^{3},$$

$$f_{6} = 2\pi \frac{\alpha\alpha'}{2\alpha'^{2} - 3m\alpha\alpha''}, \qquad f_{7} = \frac{2^{1/4}e^{-\Psi_{0}}}{\sqrt{\pi}} \alpha''',$$

$$f_{8} = 2^{1/4} \sqrt{\pi}e^{-\Psi_{0}}\alpha'', \qquad f_{9} = 2^{5/4}\pi^{3/2}e^{-\Psi_{0}}\frac{\alpha\alpha'\alpha''}{2\alpha'^{2} - 3m\alpha\alpha''},$$

$$f_{10} = 2^{1/4}\pi^{3/2}e^{-\Psi_{0}}\alpha''. \qquad (2.5)$$

In the above and throughout this work, we used the abbreviation $\alpha \equiv \alpha(z)$. The canonical volume forms are given by

$$\operatorname{vol}_{S^{2}} = \sin \theta_{2} d\theta_{2} \wedge d\phi_{2}, \qquad \operatorname{vol}_{S^{2}_{c}} = \sin \theta_{2} d\theta_{2} \wedge (d\phi_{2} + \cosh \theta_{1} d\phi_{1}),$$
$$\operatorname{vol}_{\mathbb{H}^{2}} = \sinh \theta_{1} d\theta_{1} \wedge d\phi_{1}. \tag{2.6}$$

Consistency with the Bogomol'nyi-Prasad-Sommerfield (BPS) equations and the Bianchi identities that govern the fluxes requires that the mass parameter F_0 should be piecewise constant. Interpreting the solutions presented above in terms of fully localized sources, D8-branes, dictates that α''' must be piecewise constant. It is precisely the various consistent possibilities from which we can choose the function F_0 that provides us with a parametric family of solutions. It is should clear from the previous comments that $\alpha''' \propto F_0$.

We note that, in the above, we have used the abbreviation $' \equiv \partial_z$, and we keep this shorthand notation throughout our work.

B. The quiver picture

The authors of [22] put forth a phenomenological proposal in terms of which we can think the dual superconformal field theories. In order to have a better understanding of the systems under consideration in this work, we review some of their main characteristics. We begin by pointing out that in the UV regime the system above is dual to a six-dimensional superconformal field theory. More precisely, one can explicitly check that the solutions asymptote to a background spacetime that is of the form $AdS_7 \times I_z \times S^2$. Then, the six-dimensional field theory is understood to be defined on a spacetime of the form $\mathbb{R}^{1,3} \times \mathbb{H}_2$. This and the crucial fact that there exists a nontrivial fibration of the \mathbb{H}^2 over the S^2 of the internal manifold that does not vanish at infinity amounts to precisely provide the topological twisting in the field theory such that it partially preserves some amount of supersymmetry.

The above line of reasoning and specifically the parallelism to a six-dimensional superconformal field theory leads one to suggest naturally a geometric engineering in terms of magnetically charged 5-branes, NS5-branes, with D6- and D8-branes. Of course, this argument is rough around the edges.

The dual field theory has color (gauge) and flavor nodes. The associated numbers characterizing the nodes are $N_1, N_2, ..., N_P$ and $F_1, F_2, ..., F_P$, respectively. And now we have to turn to the six-dimensional picture in order to make a connection among those numbers explicitly. More specifically, anomaly cancellation in the six-dimensional quivers requires that [29]

$$F_i = 2N_i - N_{i-1} - N_{i+1}, \qquad (2.7)$$

a condition that is ensured when $\alpha(z)$ is chosen such that its second derivative is piecewise linear and continuous.

At this point, and since we know the special solution that we will present later and its physical interpretation, we wish to stress that Cremonesi and Tomasiello in [29] observed that there exists a possible scaling that involves taking the number of D8-branes to infinity, hence creating a continuous (smeared) distribution. However, under this scaling the AdS₇ backgrounds are still trustworthy descriptions of sixdimensional quivers with $\mathcal{N} = (1,0)$, and there are no issues with anomalies. As we will see below, this will play a crucial role to our suggestion for the integrable quiver.

Going back to the basic characteristics of the quiver in our setups, we have reached a UV conformal fixed point. These UV fixed points are deformed by some vacuum expectation values or by some relevant operators. These deformations, in close analogy to the presence of fibrations, topologically twist the six-dimensional superconformal field theory and trigger a renormalization group flow that ends in a four-dimensional superconformal field theory.

The Page charge quantization that was performed in [22] suggests the existence of a new, second set of NS5-branes. This is due to the twisting. So, we have exactly (g-1)P NS5-branes, on top of the previous one originating in the six-dimensional picture, that are orthogonal to the original NS5-branes, with g being a number. This generates additional color groups. So, with this extra stack of NS5-branes,

the D6-branes can stretch between the different stacks of the NS-branes.

After a careful counting of the degrees of freedom in the presence of the new ingredients, it was suggested that a quiver that consists of $P^2(g-1)$ copies of the six-dimensional mother theory should correctly capture the essential aspects in the four-dimensional daughter theory [22]. The quiver that was suggested in [22] is depicted in Fig. 1.

In the cartoon of the quiver, Fig. 1, we have used the very common notation for quiver diagrams. Namely, arrowed lines denote chiral multiplets, circles stand for vector multiplets, and boxes are flavors.

III. DYNAMICS OF CLASSICAL CLOSED STRINGS

In this section, we are going to study analytically the motion of a closed string soliton in the backgrounds presented in Sec. II. Such dynamics are governed by the equations of motion derived from the extremization of the string- σ model action

$$S_{\sigma} = \frac{1}{4\pi\alpha'} \int d^2\sigma \left(h^{\alpha\beta}G_{MN} + \epsilon^{\alpha\beta}B_{MN}\right) \partial_{\alpha}X^M \partial_{\beta}X^N, \quad (3.1)$$

which has to be supplemented by the Virasoro constraint $T_{\alpha\beta} = 0$, with

$$T_{\alpha\beta} = \frac{1}{\alpha'} \left(G_{MN} \partial_{\alpha} X^{M} \partial_{\beta} X^{N} - \frac{1}{2} G_{MN} \eta_{\alpha\beta} \eta^{\gamma\delta} \partial_{\gamma} X^{M} \partial_{\delta} X^{N} \right),$$
(3.2)

where G_{MN} and B_{MN} are the target space metric and NS field, respectively, X^M are the target spacetime coordinates, and $d^2\sigma = d\tau d\sigma$, where (τ, σ) are the world-volume coordinates. In the equations above, we choose the convention $\eta_{\sigma\sigma} = -\eta_{\tau\tau} = 1, \eta_{\tau\sigma} = 0$, and $\epsilon^{\alpha\beta}$ is the Levi-Civita tensor with convention $\epsilon^{\tau\sigma} = 1$. Equation (3.1) makes explicit the fact that the string soliton couples to the target spacetime through a pullback to the world-volume coordinates. $T_{\alpha\beta}$ is the world-sheet energy-momentum tensor.

The starting point to study integrability analytically is to provide an ansatz for a suitable embedding $X^M(\tau, \sigma)$ capturing the coupling to the NS field and respecting the isometries of the background and construct the NVE. Accordingly, we place our string at three fixed angles on the three-sphere inside the AdS₅ and further suggest the following string embedding:

$$t = t(\tau), \qquad \rho = \rho(\tau), \qquad z = z(\tau), \qquad \theta_1 = \theta_1(\tau), \\ \theta_2 = \theta_2(\tau), \qquad \phi_1 = \alpha_1 \sigma, \qquad \phi_2 = \alpha_2 \sigma. \tag{3.3}$$

For the above embedding, Eq. (3.3), we evaluate the σ -model Lagrangian to get



FIG. 1. The quiver field theory. The Φ_i adjoint fields obtain a mass. We have used a double arrow and Q, \tilde{Q} for the bifundamentals fields while for the fundamentals are denoted by q, \tilde{q} . The vector multiplets are represented by the circles. We have used a fourdimensional, $\mathcal{N} = 1$ language. In this diagram, the dotted lines should be understood as continuing the quiver in the same pattern; e.g., bifundamentals connect N_2 to N_3 , N_3 to N_4 , and so on.

$$\mathcal{L} = f_1(\cosh^2 \rho \dot{t}^2 - \dot{\rho}^2) - f_3 \dot{z}^2 - f_2 \dot{\theta}_1^2 - f_4 \dot{\theta}_2^2 + (f_2 \sinh^2 \theta_1 + f_4 \cosh^2 \theta_1 \sin^2 \theta_2) \alpha_1^2 + f_4 \sin^2 \theta_2 \alpha_2^2 + 2f_4 \cosh \theta_1 \sin^2 \theta_2 \alpha_1 \alpha_2 + 2f_6 \sin \theta_2 (\cosh \theta_1 \alpha_1 + \alpha_2) \dot{\theta}_2 - 2\pi \cos \theta_2 (\cosh \theta_1 \alpha_1 + \alpha_2) \dot{z}.$$
(3.4)

The equations of motion that follow from the Lagrangian density given by Eq. (3.4) are

$$\begin{split} \dot{t} &= \frac{E}{f_1 \cosh^2 \rho}, \\ \ddot{\rho} &= -\frac{E^2}{f_1^2 \cosh^2 \rho} \tanh \rho - \frac{f_1'}{f_1} \dot{z} \dot{\rho}, \\ 2f_3 \ddot{z} &= f_1' \left(-\frac{E^2}{f_1^2 \cosh^2 \rho} + \dot{\rho}^2 \right) - f_3' \dot{z}^2 - 2\pi \alpha_1 \sinh \theta_1 \cos \theta_2 \dot{\theta}_1 + f_2' (\dot{\theta}_1^2 - \alpha_1^2 \sinh^2 \theta_1) \\ &\quad + 2(\alpha_1 \cosh \theta_1 + \alpha_2) \sin \theta_2 (\pi - f_0') \dot{\theta}_2 + f_4 (\dot{\theta}_2^2 - (\alpha_1 \cosh \theta_1 + \alpha_2)^2 \sin^2 \theta_2), \\ 2f_2 \ddot{\theta}_1 &= -2f_2 \alpha_1^2 \cosh \theta_1 \sinh \theta_1 - 2\alpha_1 (\alpha_1 \cosh \theta_1 + \alpha_2) f_4 \sinh \theta_1 \sin^2 \theta_2 + 2\pi \alpha_1 \sinh \theta_1 \cos \theta_2 \dot{z} \\ &\quad - 2f_2' \dot{z} \dot{\theta}_1 - 2\alpha_1 f_6 \sinh \theta_1 \sin \theta_2 \dot{\theta}_2, \\ 2f_4 \ddot{\theta}_2 &= -(\alpha_1 \cosh \theta_1 + \alpha_2)^2 f_4 \sin(2\theta_2) + 2\alpha_1 f_6 \sinh \theta_1 \sin \theta_2 \dot{\theta}_1 - 2\dot{z} [(\alpha_1 \cosh \theta_1 + \alpha_2) \sin \theta_2 (\pi - f_0') + f_4' \dot{\theta}_2], \end{split}$$
(3.5)

where *E* is a constant obtained by integrating the equation of motion for $t(\tau)$ and is associated with the energy of the string.

The equations of motion resulting from the Lagrangian, given by Eq. (3.5), are constrained by the Virasoro conditions, namely, the world-sheet equations of motion. We can use Eq. (3.2) to evaluate the different components of the stress-energy tensor, and, thus, the Virasoro conditions explicitly for the string embedding given by Eq. (3.3) read

$$T_{\tau\sigma} = T_{\sigma\tau} = 0,$$

$$2T_{\tau\tau} = 2T_{\sigma\sigma} = f_1(-\cosh^2\rho\dot{t}^2 + \dot{\rho}^2) + f_2\dot{\theta_1}^2 + f_4\dot{\theta_2}^2 + f_3\dot{z}^2 + (f_2\sinh^2\theta_1 + f_4\cosh^2\theta_1\sin^2\theta_2)\alpha_1^2 + f_4\sin^2\theta_2\alpha_2^2 + 2f_4\cosh\theta_1\sin^2\theta_2\alpha_1\alpha_2.$$
 (3.6)

The constraints above have to hold regardless of the Lagrangian equations of motion. It is easy to show that the energy-momentum tensor, Eq. (3.6), is conserved on shell, evaluated on the equations of motion given by Eq. (3.5). More explicitly, we have $\nabla^{\alpha}T_{\alpha\beta} = 0$ when we enforce the equations of motion on these expressions.

It is this agreement, precisely, between the Virasoro conditions and the equations of motion for the target spacetime coordinates, that indicates the consistency of the string soliton chosen in Eq. (3.3). This serves as a nice and quick check of the string embedding.

We, now, turn to the Hamiltonian formulation of the problem, in order to have a better grip on the physics and subsequent analysis. We begin by presenting the expressions for the conjugate momenta, which are

$$p_{t} = 2f_{1} \cosh^{2} \rho \dot{t},$$

$$p_{\rho} = -2f_{1}\dot{\rho},$$

$$p_{z} = -2f_{3}\dot{z} - 2\pi(\alpha_{1} \cosh\theta_{1} + \alpha_{2})\cos\theta_{2},$$

$$p_{\theta_{1}} = -2f_{2}\dot{\theta_{1}},$$

$$p_{\theta_{2}} = -f_{4}\dot{\theta_{2}} + 2f_{6}(\alpha_{1} \cosh\theta_{1} + \alpha_{2})\sin\theta_{2}.$$
(3.7)

We can, of course, use the expressions in Eq. (3.7) to derive an explicit expression for the Hamiltonian density. It is given by

$$\mathcal{H} = -f_1 \cosh^2 \rho \dot{t}^2 + f_1 \dot{\rho}^2 + f_3 \dot{z}^2 + f_2 \dot{\theta}_1^2 + f_4 \dot{\theta}_2^2 - (f_2 \sinh^2 \theta_1 + f_4 \cosh^2 \theta_1 \sin^2 \theta_2) \alpha_1^2 - f_4 \sin^2 \theta_2 \alpha_2^2 - 2f_4 \cosh \theta_1 \sin^2 \theta_2 \alpha_1 \alpha_2.$$
(3.8)

In the Hamiltonian formalism of the problem, the statement of the Virasoro conditions is equivalent to $\mathcal{H} = 0$. The Hamiltonian equations of motion that follow from Eq. (3.8) are equivalent to those that follow from the Lagrangian.

Owing to the above picture, it is very straightforward and instructive to provide a classical mechanics description of the system we have considered. The dynamics of the string embedding has effectively reduced to that of a particle that is moving in the presence of a nontrivial potential. It is precisely the nontrivial wrapping of the string soliton, the nontrivial winding of the string around the cyclic coordinates, that is responsible for generating the potential in the dynamics. The effective mass can be read off from the kinetic terms of the Hamiltonian and is due to the geometrical aspects of the setup we are examining. The dynamics of the string soliton in Eq. (3.3) is described by a quite involved system of differential equations; see Eq. (3.5). The major reason for the complication is that the equations are inherently coupled. Solving this system in all generality presents itself as a quite formidable task. In spite of this, there is an elegant alternative approach that manages to circumvent this complication and facilitates the needs for the forthcoming analysis. In order to make progress, we need to find a simple solution to the equations of motion and subsequently consider fluctuations around those. Such a fluctuation around the set of simple solutions is what is known as the NVE for a given coordinate.

Before we proceed, we stress that the Virasoro conditions constitute primary constraints; namely, the simple solutions to the equations of motion have to satisfy the Virasoro conditions as well.

In order to make progress, we wish to find an invariant plane of solutions and a simple equation for the $z(\tau)$. To that end, it is easy to check that⁶

$$\ddot{\rho}=\dot{\rho}=\rho=\ddot{\theta}_1=\dot{\theta}_1=\theta_1=\ddot{\theta}_2=\dot{\theta}_2=\theta_2=0 \quad (3.9)$$

solves all the equations in Eq. (3.5) except for the one for $z(\tau)$. We evaluate the *z* equation on the invariant plane of solution, and we obtain

$$2f_3\ddot{z} + f'_3\dot{z}^2 + E^2\frac{f'_1}{f_1^2} = 0.$$
 (3.10)

We insert the expressions for the f_1 and f_3 in the above, and we obtain

$$\ddot{z} - \left(\frac{\alpha \alpha''' - \alpha' \alpha''}{4\alpha^2}\right) \left(\frac{\alpha}{\alpha''}\right) \left(\frac{E^2}{36\pi^2} - \dot{z}^2\right) = 0, \quad (3.11)$$

which admits the simple solution

$$z_{\rm sol}(\tau) = \frac{E}{6\pi}\tau.$$
 (3.12)

It is a very straightforward calculation to evaluate the Virasoro conditions, Eq. (3.6), on the invariant plane of solutions given by Eq. (3.9) and, subsequently, enforce Eq. (3.12) to check the agreement of these expressions.

Therefore, we have shown that all the steps, so far, are self-consistent and allow for a natural interpretation of the string embedding as some operator with large R charge and large conformal dimension in the dual field theory.

We wish to make a comment about the invariant plane of solutions around which we fluctuate the string embedding.

⁶It should also be quite obvious that this plane of solutions is by no means unique. It is chosen for our convenience in this analysis and always taking into consideration the string theory constraints.

Naively, it would seem that the choice made in Eq. (3.9) reduces the string to a point particle, and, hence, the effect of the winding of the modes is neglected. However, this is not true as can be seen explicitly from Eqs. (3.20) and (3.23).

This sets up the scene for the derivation of the NVEs for the various dimensions. The NVEs play the protagonistic role in our analysis.

Since we have found a base solution for our system, we wish to study the NVE. This is achieved by allowing for small fluctuations in one of the equations of motion, working in linear order, and evaluating the resulting expression on the base solution defined by the invariant plane of solutions and the simple solution for t and z. All the NVEs can be brought into the schematic form

$$\ddot{y} + \mathcal{A}_1 \dot{y} + \mathcal{A}_2 y = 0, \qquad (3.13)$$

and under the change of variables given by $y = e^{\frac{1}{2}\int d\tau A_1} q$ we obtain the Schrödinger form

$$\ddot{q} = \mathcal{V}q, \qquad \mathcal{V} = \frac{1}{4} \left(2\mathcal{A}_1' + \mathcal{A}_1^2 - 4\mathcal{A}_2 \right), \qquad (3.14)$$

where in the above *y* denotes any generic fluctuation.

We will specialize the above picture to each one of the different coordinates and derive their associate NVEs.

A. NVE for ρ

Here, we derive the NVE for the ρ coordinate, by considering $\rho = 0 + \epsilon r$ and expanding in the $\epsilon \rightarrow 0$ limit working to linear order in the small parameter. This results in

$$r''(\tau) + A_1 r'(\tau) + A_2 r(\tau) = 0$$
 (3.15)

with

$$\mathcal{A}_{1} = \frac{E}{12\pi} \left(\frac{\alpha'}{\alpha} - \frac{\alpha'''}{\alpha''} \right) \Big|_{z=z_{\text{sol}}},$$
$$\mathcal{A}_{2} = -\frac{E^{2}m}{864\pi^{2}} \frac{\alpha''}{\alpha} \Big|_{z=z_{\text{sol}}}.$$
(3.16)

We perform the appropriate change of variables as described above to bring it in the form

$$\ddot{\varrho} = \mathcal{V}_{\varrho}\varrho, \qquad \mathcal{V}_{\varrho} = -\frac{E^2}{f_1^2} \left(1 + \frac{1}{144\pi^2} f_1'^2 - \frac{1}{72\pi^2} f_1 f_1'' \right).$$
(3.17)

B. NVE for θ_1

Likewise, we consider $\theta_1 = 0 + \epsilon v_1$ and work as before. We have

$$v_1''(\tau) + \mathcal{A}_1 v_1'(\tau) + \mathcal{A}_2 v_1(\tau) = 0$$
 (3.18)

with

$$\mathcal{A}_{1} = \frac{E}{12\pi} \left(\frac{\alpha'}{\alpha} - \frac{\alpha''}{\alpha''} \right) \Big|_{z=z_{\text{sol}}},$$
$$\mathcal{A}_{2} = \alpha_{1}^{2} - \frac{\alpha_{1}E\sqrt{m}}{6\sqrt{6}\pi} \sqrt{-\frac{\alpha''}{\alpha}} \Big|_{z=z_{\text{sol}}}.$$
(3.19)

In this case as well, we wish to manipulate the NVE and bring it into a Schrödinger form as follows:

$$\ddot{\vartheta}_{1} = \mathcal{V}_{\vartheta_{1}}\vartheta_{1},$$

$$\mathcal{V}_{\vartheta_{1}} = -\frac{1}{f_{2}^{2}} \left(\alpha_{1}^{2}f_{2}^{2} - \frac{\alpha_{1}E}{6}f_{2} + \frac{E^{2}}{144\pi^{2}}f_{2}^{\prime 2} - \frac{E^{2}}{72\pi^{2}}f_{2}f_{2}^{\prime \prime} \right).$$

(3.20)

C. NVE for the θ_2

Finally, we derive the NVE associated with the θ_2 dimension. To do so, we allow $\theta_2 = 0 + \epsilon v_2$ and proceed as in the previous cases. This yields

$$v_2''(\tau) + \mathcal{A}_1 v_2'(\tau) + \mathcal{A}_2 v_2(\tau) = 0, \qquad (3.21)$$

where in the above we have

$$\begin{aligned} \mathcal{A}_{1} &= \frac{E}{12\pi(3m\alpha\alpha'' - 2\alpha'^{2})} \left(\frac{\alpha'}{\alpha} ((8+3m)\alpha\alpha'' - 6\alpha'^{2}) - \frac{\alpha'''}{\alpha''} (2\alpha'^{2} + 3m\alpha\alpha'') \right) \Big|_{z=z_{sol}}, \\ \mathcal{A}_{2} &= (\alpha_{1} + \alpha_{2})^{2} + \frac{(\alpha_{1} + \alpha_{2})E}{\sqrt{6m\pi(3m\alpha\alpha'' - 2\alpha'^{2})}} \cdot \\ &\left(\frac{2(1-3m)\alpha'^{2}\sqrt{-\alpha^{3}\alpha''}}{3\alpha^{2}} - \frac{m\alpha^{2}((2+3m)\alpha''^{2} - 2\alpha'\alpha''')}{2\sqrt{-\alpha^{3}\alpha''}} \right) \Big|_{z=z_{sol}}. \end{aligned}$$
(3.22)

As we did in the previous cases, we change variables appropriately to bring it to a Schrödinger form:

$$\begin{split} \dot{\vartheta_2} &= \mathcal{V}_{\vartheta_2} \vartheta_2, \\ \mathcal{V}_{\vartheta_2} &= -\frac{1}{f_4^2} \left((\alpha_1 + \alpha_2)^2 f_4^2 + \frac{E}{6\pi} (\alpha_1 + \alpha_2) (\pi - f_6') f_4 \right. \\ &\quad + \frac{E^2}{144\pi^2} f_4'^2 - \frac{E^2}{72\pi^2} f_4 f_4'' \right). \end{split}$$
(3.23)

We have ended up with second-order, linear, ordinary differential equations. These have all the necessary characteristics to be amenable to Kovacic's algorithm and, hence, examine whether or not they admit Liouvillian solutions.

IV. LIOUVILLIAN INTEGRABILITY

In this section, we discuss the status of integrability for closed strings in the background that we presented in Sec. II A. As we shall see, all choices for $\alpha(z)$ consistent with the BPS equations and the Bianchi identities lead to nonintegrable theories, except for the solution $\alpha(z) \propto \sin z$, where the proportionality is up to some numerical coefficients. These coefficients from the point of view of α are just the amplitude and the phase of the sin function. From the point of view of the quiver, they control the overall size and the number of flavors. All other solutions for $\alpha(z)$ that are consistent with the BPS equations and Bianchi identities lead to nonintegrability. *En passant*, we make some comments on the interpretation of the special solution.

A. The allowed solutions for $\alpha(z)$

The authors in [22] have thoroughly analyzed the BPS equations, Bianchi identities, and Page-charge quantization for the supergravity backgrounds presented in Sec. II. We review the main results of their analysis in order to understand what are the possible choices for the defining function $\alpha(z)$. We start by recalling that the authors of [22] showed that, in order for the solution to be given an interpretation as fully localized D8-branes, we must insist on $\alpha'''(z)$ being piecewise constant, since this is related to the Ramond-Ramond scalar. After performing a gauge

transformation, the study of quantization of the Page charges reveals that $\alpha''(z)$ must be a linear function with integer coefficients. Furthermore, they concluded that the fourth derivative $\alpha^{(4)}(z)$ should vanish. Finally, in order to avoid singular behavior, the defining function should satisfy $\alpha(0) = 0 = \alpha(P)$, with *P* being the upper end of the *z* interval; $\mathcal{I}_z \in [0, P]$.

Under these considerations, and following [22], the most general solution consistent with the supergravity equations is given by

$$\alpha(z) = \alpha_0 + \alpha_1 z + \alpha_2 z^2 + \alpha_3 z^3, \tag{4.1}$$

with $\alpha_0, \ldots, \alpha_3$ being some appropriately chosen constants.

The authors of [22] also gave a Hanany-Witten interpretation of the setup. The above picture can be realized in terms of a Hanany-Witten setup made of two sets of NS5branes, two sets of D6-branes, and one of D8-branes. From the six-dimensional picture, this can be roughly understood as follows. The AdS₇ backgrounds, from which the AdS₅ solutions originate, have an interpretation in terms of a Hanany-Witten setup with NS5-D6-D8 branes [29] and the twisted compactification induces a new set of NS5-branes and a new set of D6-branes.

Since we have already obtained the NVEs in all generality in the previous section (see Sec. III), and we have explained the basic aspects of $\alpha(z)$, we can use the most general and schematic form for the defining function given by Eq. (4.1) so that we check if we can make some robust statement on the integrability of closed strings in the background.

B. Some examples of linear quivers

To provide the interested reader a firm grip on the shapes of quivers and how they relate to the choice of $\alpha(z)$, we briefly review two examples and make some comments. These were presented thoroughly in [22].

1. Example 1

This first example is the dual originating from a sixdimensional liner quiver theory of rank $N_j = j$ and a flavor group given by SU(*PN*). The defining function $\alpha(z)$ is given by

$$\alpha(z) = \begin{cases} \frac{N}{6}(1-P^2)z + \frac{N}{6}z^3, & 0 \le z \le (P-1), \\ -\frac{1}{6}(2P^2 - 3P + 1)(P-z) + \frac{N}{6}(P-1)(P-z)^3, & (P-1) \le z \le P. \end{cases}$$
(4.2)

We note that $\alpha(z)$ vanishes at z = 0 and z = P and is continuous at z = (P - 1). Its derivative $\alpha'(z)$ is continuous at the same point.

2. Example 2

The second example, from a six-dimensional point of view, describes a linear quiver with (P-1) gauge nodes of rank $N_i = N$ and flavors $F_J = N(\delta_{J,1} + \delta_{J,P-1})$. The defining function $\alpha(z)$ in this example is given by

$$\alpha(z) = \begin{cases} \frac{N}{2}(1-P)z + \frac{N}{6}z^3, & 0 \le z \le 1, \\ \frac{N}{6} - \frac{PN}{2}z + \frac{N}{2}z^2, & 1 \le z \le (P-1), \\ -\frac{N}{2}(P-1)(P-z) + \frac{N}{6}(P-z)^3, & (P-1) \le z \le P. \end{cases}$$
(4.3)

The function $\alpha(z)$ vanishes at z = 0 and z = P and is continuous at z = 1 as well as at z = (P - 1), and its derivative is continuous at those two points.

We stress that, while we will present the details and plots of the numerical analysis for Eqs. (4.2) and (4.3), we have thoroughly checked many more examples and all lead to the same qualitative behavior.

Having described the main characteristics of the quiver theories and how they connect to the supergravity defining function, we proceed to analytically prove that these quivers are exactly nonintegrable. The way to do this follows from the simple argument that all quivers start with an SU(N) factor, and this translates to the $\alpha(z)$ assuming a very simple form, namely, $\alpha(z) = a_1 z + a_2 z^3$ in the first part of the interval. This can be explicitly seen from the two examples we presented above. We have obtained the NVEs in the most general case, without assuming anything on the form of the $\alpha(z)$. We will use the simplest choice $\alpha(z) =$ $a_1 z + a_3 z^3$ to demonstrate that the ordinary quivers are all necessarily nonintegrable.

We have explained Kovacic's algorithm sufficiently in the Appendix. Here, we just apply its logic on the NVEs.

With this election of the defining function, the potential appearing in the NVE for ρ adopts the form

$$\mathcal{V}_{\varrho}(\tau) = \frac{27a_1a_3E^2\pi^2}{(36a_1\pi^2 + a_3E^2\tau^2)^2} + \frac{a_3E^2(m-1)}{4(36a_1\pi^2 + a_3E^2\tau^2)}.$$
(4.4)

This potential fall into case 3 in the classification given in Kovacic's algorithm. That is, $V_{\varrho}(\tau)$ has order 2 at infinity, and the order of the poles does not exceed 2. So, in principle, this NVE could accept Liouvillian solutions.

The potential appearing in the NVE for θ_1 adopts the form

$$\mathcal{V}_{\theta_{1}} = \alpha_{1}^{2} + \frac{27a_{1}\pi^{2}}{a_{3}E^{2}} \frac{1}{\left(\frac{36a_{1}\pi^{2}}{a_{3}E^{2}} + \tau^{2}\right)^{2}} - \frac{1}{4\left(\frac{36a_{1}\pi^{2}}{a_{3}E^{2}} + \tau^{2}\right)} + \frac{-\alpha_{1}a_{3}\sqrt{m}E^{2}}{36a_{1}\pi^{2}}\sqrt{-\left(\frac{36a_{1}\pi^{2}}{a_{3}E^{2}} + \tau^{2}\right)} - \frac{a_{3}E^{2}\sqrt{m}\alpha_{1}\tau^{2}}{36a_{1}\pi^{2}} \frac{1}{\sqrt{-\left(\frac{36a_{1}\pi^{2}}{a_{3}E^{2}} + \tau^{2}\right)}}.$$

$$(4.5)$$

However, the above expression is not written in terms of rational functions, and, hence, it is not in an appropriate form for us to apply the criteria of Kovacic's algorithm. As we will see, however, in the later sections where we perform numerical analysis, this does not stop us from determining that the quivers of these type are not classically integrable. We note that the resulting potential for the θ_2 NVE also cannot be written in terms of rational functions, and, hence, it is not amenable to the algorithm by Kovacic.

C. A more exotic and special quiver

We claim that the situation above is drastically different when the quiver instead of being linear is of sinusoidal shape. This means that the defining function $\alpha(z)$ in the supergravity description is given by

$$\alpha(z) \propto \sin(z), \tag{4.6}$$

up to some unimportant numerical coefficients.

- We feel that some necessary comments are in order.
- (i) This is exactly the solution that led to an integrable six-dimensional theory in the mother AdS₇ backgrounds [15]. We have already, thoroughly, explained why this situation is by no means a trivial statement of our work. Instead, it presents itself as an interesting case scenario for further and deeper explorations between the sinusoidal-shaped quivers in massive type IIA supergravity and classical integrability.
- (ii) It is obvious that for the solution given by Eq. (4.6) we cannot have its derivative—of any order—being equal to zero. We know, however, that $\alpha'''(z)$ is related to the 0-form in the Ramond-Ramond sector, the Romans mass. Hence, the behavior of $\alpha'''(z)$ is associated with the distribution of D8-branes in the background. This resembles closely the interpretation that was given in [15] for the integrable quiver in the mother AdS₇ setup.
- (iii) This leads us to suggest the following physical interpretation for this special solution of the characteristic function. For $\alpha(z)$ given by Eq. (4.6), we have D8-brane sources that are smeared all along the range of the z dimension, instead of them being sharply localized as in the ordinary solutions—the ones presented in Sec. IVA.

Next, we proceed to examine the consequences of choosing Eq. (4.6) as the defining function for our supergravity solution on the various NVEs we have derived in Sec. III. We make our definition more precise by specifying $\alpha(z) = A \sin(\omega z)$, with ω being associated to the number of *D*6-branes. In the above, $A \in \mathbb{R}_{\geq 0}$ and $\omega \in \mathbb{R}_{\geq 0}$.

For this choice of the defining function $\alpha(z)$, the geometry in Eq. (2.1) becomes

$$ds^{2} = 12\sqrt{\frac{6}{m}}\frac{\pi}{\omega}ds^{2}_{AdS_{5}} + \sqrt{\frac{6}{m}}\frac{\pi}{\omega}ds^{2}_{\mathbb{H}^{2}} + \sqrt{\frac{3m}{2}}\pi\omega dz^{2} + \sqrt{6m}\frac{\pi}{\omega}\frac{\pi}{3m+2\cot^{2}(\omega z)} \times (ds^{2}_{S^{2}} + \cosh\theta_{1}\sin^{2}\theta_{2}ds^{2}_{2}).$$
(4.7)

It should be obvious from the form of Eq. (4.7) that up to the fibration, denoted by ds_2^2 , the geometry has reduced to that of integrable submanifolds. We note that by integrable submanifolds we mean that the Polyakov action on AdS_n, dS_n, Sⁿ, and Hⁿ is known to be integrable in the absence of a B_2 field; see Appendix B in [15] for the construction of the Lax connection in these cases.

Note that a similar, but indeed much easier to understand, situation arose in the study of the mother AdS_7 backgrounds [15]. Indeed, when studying the mother theory, there is, of course, no such complication as a nontrivial fibration. This is why in the UV, where the theory flows to the six-dimensional $\mathcal{N} = (1,0)$ superconformal field theory, it was more straightforward to obtain a Lax.

This is our first piece of evidence, although not strong and conclusive, and is entirely based on the isometries that are present in the setup. It is also once again quite obvious why the status of integrability in the AdS_5 does not immediately follow from the corresponding statement of the mother AdS_7 solution.

We proceed to quickly inspect the NVEs for the ρ and θ_1 in this case and offer some comments for the NVE of θ_2 that also connect to a similar peculiarity in the AdS₇ backgrounds.

1. NVE for ρ

In this case, Eq. (3.17) becomes

$$\varrho''(\tau) + \frac{E^2 m \omega^2}{864 \pi^2} \varrho(\tau) = 0, \qquad (4.8)$$

which has the simple solution

$$\varrho = \cos\left(\frac{E}{12}\sqrt{\frac{m}{6}}\frac{\omega}{\pi}\tau\right) + \sin\left(\frac{E}{12}\sqrt{\frac{m}{6}}\frac{\omega}{\pi}\tau\right), \quad (4.9)$$

that is clearly Liouville integrable.

2. NVE for
$$\theta_1$$

Equation (3.20) becomes

$$\vartheta_1''(\tau) + \frac{1}{6}\alpha_1(E - 6\alpha_1)\vartheta_1(\tau) = 0,$$
 (4.10)

which admits the Liouville integrable solution

$$\vartheta_1 = e^{\sqrt{-\alpha_1^2 + \frac{\alpha_1 \bar{E}}{6}\tau}} + e^{-\sqrt{-\alpha_1^2 + \frac{\alpha_1 \bar{E}}{6}\tau}}.$$
 (4.11)

3. NVE for θ_2

The NVE of θ_2 is much more cumbersome to analyze, even for the case of $\alpha(z) = A \sin(\omega z)$. This has to do with the increased complexity in the equations of motion of that particular coordinate. Similar difficulties appeared also in the analysis of the AdS₇ backgrounds, where we know the sinusoidal quiver is integrable, since we have an explicit Lax connection [15]. Truly, this was observed in their context for the NVE of the angular χ dimension, and we refer the reader to Eqs. (A.10) and (A.11) in [15] for the details in that setup. What we face here is entirely equivalent to that situation. From an inspection of the NVEs presented in Sec. III, we can see that in the NVE for the θ_2 coordinate, and more specifically its Schrödinger form, the warp factors f_4 and f_6 appear, unlike in the other two cases that only the f_1 and f_2 are present. This, in turn, results in obtaining a very complicated sum of large fractions, from which it is not possible to make any statements on the status of integrability for the string soliton. We will omit any further discussion here, and we will see from the numerical analysis that this is not a problem, as we are not led to chaotic behavior when studying this NVE. Borrowing intuition from the AdS₇ results (see Ref. [15]), we provide the hand-wavy argument that this is an unimportant complication that should not be discouraging at all.

V. NUMERICAL ANALYSIS OF STRINGS AND THEIR DYNAMICS

In this section, we proceed to carry out an exhaustive numerical analysis which will allow us to elaborate more on the analytic results obtained in the previous section. The numerical analysis also provides us with a more solid understanding of all four-dimensional $\mathcal{N} = 1$ quiver theories that are associated with these massive type IIA vacua. We will be able to compare the dynamics of any generic nonintegrable and chaotic quiver to those of the special ones that are integrable. As we will see, deforming the special solution that we have found exhibiting integrability even by some small amount results to signatures of chaos. This complements our previous findings which were based on the study of the NVEs.

We perform numerical computations by studying the dynamics of classical strings that are propagating in the backgrounds we consider in this work. These considerations can be used to demonstrate that the dynamics in the phase space of these classical strings exhibits characteristic signatures of chaos and are, thus, nonintegrable.

It is important at this point to pause for a brief moment and make a comment, once more, pertaining to the relation of chaos to the statement of a system's integrability. It is well understood that the nonintegrability of a system does not, necessarily, imply chaos. However, on the other hand, chaotic dynamics is indicative of the absence of integrability. It is also worthwhile mentioning that in the context of the gauge-gravity duality all known examples hitherto that have been argued to be nonintegrable were also shown to be chaotic; see also [38], where these ideas were explained.

We describe the general idea in order to explain our approach. We are following [61] in our discussion.⁷

Let us consider a dynamical system. The time evolution of this system is determined by a set of differential equations. This allows us to calculate the state of a system at a time, which we denote by t_{final} , knowing an earlier state of the system at some fixed point in time, which we denote by t_{earlier} . Following this logic, we can reach the beginning of the dynamical time evolution of the system at t_0 . Therefore, knowing the state of the system at the initial time t_0 , we can, in principle, solve the system at hand at any time.

Such a dynamical system is considered to be chaotic if it is sensitive to its initial conditions, namely, if the nearby points of the system depart from one another in an exponential way along the time flow. This is making it impossible to accurately predict its dynamical behavior based on the states of previous times. A simple, but illustrative, example of this is as follows: Let us assume that we have two adjacent initial conditions describing points that lie arbitrarily close to each other. We use $h_1(t_0)$ and $h_2(t_0) = h_1(t_0) + \epsilon$ to denote these points, and ϵ is a small parameter. We declare that the system exhibits chaotic dynamics when $|h_1(t) - h_2(t)| \sim e^{\lambda t}$. This is under the assumption that the system's trajectories in the phase space are bounded. This criterion on the boundedness of the trajectories is essential in the sense that it rules out a possible situation in which the trajectories are moving toward infinity. In that scenario, the exponential divergence is because they are moving apart [61]. This is the so-called trivial case [38].

The case of our interest is that of the motion of classical strings that are positioned at the center of the AdS_5 spacetime while moving and rotating in the internal space of the ten-dimensional solution. This is described by the system of equations that we derived in Eqs. (3.5) and (3.6) or their analogous ones; see Eqs. (3.7) and (3.8) for the Hamiltonian formulation. The coordinates of the internal space are bounded, and the same holds true for the associated conjugate momenta owing to the conservation of the Hamiltonian.

We will make a specific choice for the linear quiver to be used in the sections below as an illustrative example of our methods. We have already presented two concrete and interesting examples of quivers taken from [22]. We will make another concrete choice that will facilitate our analysis. This is given by

$$\alpha(z) = -81\pi^2 5\left(-6z + \frac{1}{6}z^3\right).$$
 (5.1)

Note that a similar solution was studied in [38] for the seed AdS_7 solutions. The interest in this is because it presents itself as the simplest, nontrivial solution with good behavior.

A. The dynamical evolution of the system and power spectra

We begin our numerical analysis section with the study of the classical string dynamics. This is the simple statement that we allow for time to evolve and we solve the differential equations resulting from the Lagrangian density numerically, for the choice of a linear quiver corresponding to Eq. (5.1). Subsequently, we plot the string motion along the various coordinates against time. This is shown in Fig. 2.

It should be quite obvious that, except for the ρ coordinate, the string motion along the rest is chaotic. It is, perhaps, quite instructive to examine the trajectories in the relative plane of coordinates for the string motion. We start by depicting the dynamical evolution in the (z(t), any coordinate) plane in Fig. 3.

A very quick inspection of Fig. 3 reveals that the string dynamics are highly disorganized and lack any integrability. More precisely, what we find out is that the motion is not periodic in any of those planes, and the path of the string does not close on itself. We will see a confirmation of this picture and intuition in the corresponding power spectrum, as well. We point out that one can choose one of those planes, e.g., the (z, ρ) plane, and allow the string to be energetically excited. And then proceed with this to the rest of the planes. The same qualitative conclusions can be reached, of course. We refrain from doing so, to avoid providing too many plots, and instead we opt to provide the other relative planes for a particularly frozen moment, for completeness. These are depicted for the $(\rho(t), any \text{ coordinate})$ plane in Fig. 4 and for $(\theta_1(t), \text{any coordinate})$ and $(\theta_2(t), \text{any coordinate})$ planes in Figs. 5 and 6, respectively.

We now move on to discuss the corresponding power spectra following [38,41,61]. We start by taking the Fourier transform of the numerical evolution of string motion, depicted in Figs. 3–6. Doing so, we can determine if the corresponding coordinates are periodic, quasiperiodic, or chaotic. If a signal has a well-defined period characterized by a frequency ω , the relevant spectrum will have a vertical line at the characteristic frequency of the system. Of course, one has to be careful with subtleties in the numerical errors

⁷See also the related numerical analysis in [38] for propagating strings in the mother AdS_7 backgrounds and also [41] for a similar analysis in the Gaiotto-Maldacena backgrounds.



FIG. 2. The dynamical evolution of the various string coordinates in time. The parameters chosen are $\alpha_1 = \alpha_2 = 1$, m = 1, and $\Psi_0 = 0$.



FIG. 3. The various frozen trajectories in the (z(t), any coordinate) plane. The parameters chosen are $\alpha_1 = \alpha_2 = 1$, m = 1, and $\Psi_0 = 0$.

and estimates when examining these behaviors [61]. As we see, however, in Fig. 7, the situation is very clear and suggestive.

As we see from the plot of the power spectra, which is presented in Fig. 7, for the various trajectories, there seems to be a fundamental frequency of value 0.01; however, the chaotic behavior observed in Fig. 2 is evident here as well. Truly, very quickly as we increase the frequency, we observe that all the higher harmonics are lost and what is happening is that there is a broad band of noise that is overtaking the spectra. By the end, the full picture is just dominated by pure noise. As we have mentioned previously in the examination of the dynamics in the relative planes and not in Fourier space, one can focus on a particular plane-or even choose all of them-and proceed to finetune the initial conditions of the system in such a way that the string is getting energetically excited. Then, proceed to examine the quantitative behavior of the system. While this is an interesting study, it presents itself as immaterial for our purposes, since we would reach the same qualitative behaviors. Since in Fig. 7 there is no sign to indicate that higher harmonics could, perhaps, survive and compete with the noise that is manifested, there is no need to further delineate this study, and we can make a conclusive argument even at this point.

Now, we turn our attention very quickly to the special quiver that corresponds to a sinusoidal shape. For the choice $\alpha(z) = A \sin(\omega z)$, we repeat the same analysis that we have presented above. We refrain from providing all the plots, and we present our findings for the power spectrum in this case since we can derive conclusive evidence from that. As we can observe by inspecting Fig. 8, the situation is drastically different compared to Fig. 7.

It is clear that in Fig. 8 none of the trajectories is losing its higher harmonics as we increase the frequency, and there is no sign of noise trying to overpower the spectrum. We have checked that, even if we energetically excite the string and repeat the same analysis, we can reach the same qualitative output. This is another strong indicator that the choice of a sinusoidal quiver is very special with regards to the status of integrability.

The observed behavior in Fig. 8 is also manifested in the dynamical evolution of the string motion in position space, as well as from the relative planes. Truly, if one plots the



FIG. 4. The various frozen trajectories in the ($\rho(t)$, any coordinate) plane. The parameters chosen are $\alpha_1 = \alpha_2 = 1$, m = 1, and $\Psi_0 = 0$.



FIG. 5. The various frozen trajectories in the ($\rho(t)$, any coordinate) plane. The parameters chosen are $\alpha_1 = \alpha_2 = 1$, m = 1, and $\Psi_0 = 0$.

dynamical evolution of the string along the various dimensions against time, then one sees clearly a periodic and canonical motion rather than the erratic and chaotic behavior observed in Fig. 2. The corresponding statement from the relative planes of motion is that the trajectories all close on themselves and form figures with clear patterns.

Before closing this subsection, we present the results from the computations of the power spectra for the two



FIG. 6. The various frozen trajectories in the ($\rho(t)$, any coordinate) plane. The parameters chosen are $\alpha_1 = \alpha_2 = 1$, m = 1, and $\Psi_0 = 0$.



FIG. 7. The power spectra for the $(z(t), \rho(t), \theta_1(t), \theta_1(t))$ trajectories. The parameters chosen are $\alpha_1 = \alpha_2 = 1$, m = 1, and $\Psi_0 = 0$.



FIG. 8. The power spectra for the $(z(t), \rho(t), \theta_1(t), \theta_1(t))$ trajectories. The parameters chosen are $\alpha_1 = \alpha_2 = 1$, m = 1, and $\Psi_0 = 0$.

examples that we described in Eqs. (4.2) and (4.3). These results are shown in Fig. 9.

B. The spectrum of Lyapunov exponents

The spectrum of Lyapunov exponents is considered to be a quantity determining that the dynamical information is lost when a system exhibits chaos. The loss of information is sourcing the dynamical Kolmogorov-Sinai entropy in a chaotic setup. When we consider a given dynamical system that is described by a nonzero Lyapunov exponent, the explanation that is associated with this phenomenon is that the time evolution of two trajectories that are close to each other in the phase space turns out to be extremely sensitive to an infinitesimal change in the initial conditions. This effect is amplified exponentially at sufficiently large times. More simply put, the existence of a nonzero Lyapunov exponent for a point Y = (q, p) in the phase space with initial condition $Y_0 = (q_0, p_0)$ at t = 0 is given by

$$\lambda = \lim_{t \to \infty \atop \Delta Y_{0} \to 0} \frac{1}{t} \log \frac{\Delta Y(Y_0, t)}{\Delta Y(Y_0, 0)}$$
(5.2)

and is intimately related to the degree of randomness associated with the dynamical phase space of a Hamiltonian system [32,38,41].

It measures the rate of separation between two trajectories that lie arbitrarily close in the phase space of the theory. The function $\Delta Y(Y_0, t)$ measures the separation between two such trajectories as a function of this initial location. For chaotic systems, we end up obtaining



(a) The power spectrum for equation (4.2).

(b) The power spectrum for equation (4.3).

FIG. 9. Power spectra for the two quivers discussed in Eqs. (4.2) and (4.3).

$$\Delta Y(Y_0, t) \sim \Delta Y(Y_0, 0) e^{\lambda t}.$$
(5.3)

We proceed by providing the details for the computation and implementation of the spectrum of Lyapunov exponents. In the description of the spectrum, we use λ_i to denote the *i*th Lyapunov characteristic exponent, henceforth abbreviated LCE, according to the standard nomenclature in the literature. We follow [62] in all the details and numerical techniques of our implementation. This has been used in many cases in the study of chaos in holographic setups; see, for example, [38,41]. We use the notational conventions of [38,41] for convenience.

For the computation of the Lyapunov exponents, we choose the initial conditions in such a way as to satisfy the Hamiltonian constraint, i.e., $\mathcal{H} = 0$. This, in turn, implies that, for a 2*N*-dimensional phase space,

$$\sum_{i=1}^{2N} \lambda_i = 0. \tag{5.4}$$

This is what measures the exponential growth associated with the *i*th direction in the phase space. For some systems, the sum of all positive Lyapunov exponents measures the Kolmogorov-Sinai entropy production during the dynamic evolution in the phase space.

We substitute appropriate initial conditions into the dynamical equations in order to generate a solution. Upon choosing these initial conditions for the phase-space variables and making sure that they satisfy the vanishing of the Hamiltonian, we find the corresponding Lyapunov spectrum for each of the quiver configurations. Notice that, in our analysis, we are eventually computing four LCEs (λ_i) that characterise a four-dimensional dynamical phase space.

The prescription of [62] is the following.

We consider a generic 2n-dimensional smooth dynamical canonical system, which can be written as

$$\dot{q} = \mathfrak{V}(q), \tag{5.5}$$

in all generality, and where q(t) is the 2*n*-dimensional state vector $q = (\vec{Y}(t), \vec{P}(t))$ at time $t, \dot{q} = \frac{dq}{dt}$, and \mathfrak{V} is a vector field defined on an open set U of the phase-space manifold. This vector field generates a flow f given by

$$\dot{f}^t(q) = \mathfrak{V}(f^t(q)) \text{ for all } q \in U, t \in \mathbb{R},$$
 (5.6)

with $f^{t}(q) = f(q, t)$.

Let us follow the evolution under the flow of two arbitrarily close points in phase space, q_0 and $q_0 + \delta_0$, where the subindex 0 denotes the initial time, with δ_0 being a small deviation of q_0 . After a time t, the deviation δ_t evolves to

$$\delta_t \equiv f^t(q_0 + \delta_0) - f^t(q_0) \approx D_{q_0} f^t(q_0) \cdot \delta_0, \quad (5.7)$$

where the difference in trajectories has been approximated with the gradient $D_{q_0} f^t(q_0)$. We can compute the averaged exponential growth of two trajectories as

$$\lambda(q_0, \delta_0) = \lim_{t \to \infty} \frac{1}{t} \log \frac{\|\delta_t\|}{\|\delta_0\|} = \lim_{t \to \infty} \frac{1}{t} \log \|D_{q_0} f^t(q_0) \cdot \delta_0\|,$$
(5.8)

where $||\delta||$ denotes the norm of the vector δ according to standard nomenclature, e.g., [62,63]. In the case that $\lambda > 0$, the nearby orbits will exponentially diverge. By making some mild assumptions on the vector field, the limit, given by Eq. (5.8), exists. In this case, the limit is finite and defines the largest LCE λ_1 .

We can introduce the LCEs of the order of p, with $1 \le p \le n$, as a means to describe the mean rate of growth of a p-dimensional volume in the tangent space. By constructing a solid of which each face is a parallelogram U_0 in the tangent space and with edges given by $\delta_1, \ldots, \delta_p$, we define LCEs of the order of p as



(a) The mass parameter here is: $m = \frac{1}{2}$ and the choice for the $\alpha(z) = \sin(z)$.



(b) The mass parameter here is: $m = \frac{1}{4}$ and the choice for the $\alpha(z) = \alpha(z) = -81\pi^2 5 \left(-6z + \frac{1}{6}z^3\right)$.

FIG. 10. Lyapunov for the integrable and nonintegrable quivers.

$$\lambda^{p}(q_{0}, U_{0}) = \lim_{t \to \infty} \frac{1}{t} \log \left[\operatorname{vol}^{p}(D_{q_{0}} f^{t}(U_{0})) \right], \quad (5.9)$$

where vol^p is the canonical volume form.

There exist *p* linearly independent vectors $\delta_1, ..., \delta_p$ satisfying [63]

$$\lambda^{p}(q_{0}, U_{0}) = \sum_{k=1}^{p} \lambda_{k}.$$
 (5.10)

 δ_t as defined in Eq. (5.7) dynamically evolves as

$$\dot{\Psi}_t(q_0) = D_q \mathfrak{B}(f^t(q_0)) \cdot \Psi_t(q_0), \quad \Psi_0(q_0) = \mathbb{1}, \quad (5.11)$$

where we have denoted $\Psi_t(q_0) = D_{q_0} f^t(q_0)$ and 1 denotes the identity. The trajectories in phase space are obtained by integrating the following system:

$$\begin{cases} \dot{q} \\ \dot{\Psi} \end{cases} = \begin{cases} \mathfrak{V}(q) \\ D_q \mathfrak{V}(q) \cdot \Psi \end{cases}, \qquad \begin{cases} q(t) \\ \Psi(t_0) \end{cases} = \begin{cases} q_0 \\ \mathbb{1} \end{cases}.$$
(5.12)

The spectrum of LCEs will be computed following the algorithmic process described in [64]. This procedure amounts to the computation of the order-*p* LCEs as defined in Eq. (5.10). The main upshot of the algorithm is a repeated application of the Gram-Schmidt procedure which simplifies the implementation [63]. We provide the main steps, very briefly, below.⁸

From the original set $\{\delta_i\}$, we can construct an orthonormal set $\{\hat{\delta}_i\}$ using the Gram-Schmidt recipe. The U_0 -shaped volume spanned by $\delta_1, \ldots, \delta_p$ is now constructed as

$$\operatorname{vol}\{\delta_{1},...,\delta_{p}\} = \|\hat{\delta}_{1}\|...\|\hat{\delta}_{p}\|.$$
(5.13)

The starting point of the procedure is to pick a set of initial conditions q_0 together with a matrix of $n \times n$ dimensionality $\Upsilon = \{\delta_1^0, ..., \delta_n^0\}$. Applying the Gram-Schmidt procedure, we obtain the basis $\hat{\Upsilon}_0 = \{\hat{\delta}_1^0, ..., \hat{\delta}_n^0\}$. Subsequently, we integrate Eq. (5.12) from $\{q_0, \Upsilon_0\}$ in a small neighborhood \mathfrak{T} , to obtain $q_1 = f^{\mathfrak{T}}(q_0)$ and

$$\Upsilon_1 \equiv [\delta_1^1, \dots, \delta_n^1] = D_{q_0} f^{\mathfrak{T}}(\Upsilon_0) = \Psi_{\mathfrak{T}}(q_0) \cdot [\delta_1^0, \dots, \delta_n^0].$$
(5.14)

The algorithmic process is such that it is repeating the process described above \mathcal{M} times. During the *m*th step, the *p*-dimensional volume vol^{*p*} defined in Eq. (5.9) is increased by a factor of $\|\mathbf{a}_1^m\| \dots \|\mathbf{a}_p^m\|$, with $\{\mathbf{a}_1^m, \dots, \mathbf{a}_p^m\}$ being the set of orthonormal vectors calculated from U_m . Then,

$$\lambda^p(q_0, \Upsilon_0) = \lim_{m \to \infty} \frac{1}{m\mathfrak{T}} \sum_{i=1}^m \log(\|\hat{\delta}_1^i\| \dots \|\hat{\delta}_p^i\|).$$
 (5.15)

We immediately obtain

$$\lambda_p = \lim_{k \to \infty} \frac{1}{m\mathfrak{T}} \sum_{i=1}^m \log \|\hat{\delta}_p^i\|.$$
 (5.16)

The full Lyapunov spectra is completed by computing the following:

$$\frac{1}{\mathcal{M}\mathfrak{T}}\sum_{i=1}^{\mathcal{M}}\log\|\hat{\delta}_{1}^{i}\|\approx\lambda_{1},\ldots,\frac{1}{\mathcal{M}\mathfrak{T}}\sum_{i=1}^{\mathcal{M}}\log\|\hat{\delta}_{1}^{n}\|\approx\lambda_{n},\quad(5.17)$$

for an appropriately chosen \mathfrak{T} , until we find them converging.

⁸See also [38] for an exposition on this approach.



FIG. 11. Some indicative Poincaré sections. (a) Poincaré sections for the quiver with defining function $\alpha(z) = \sin(\frac{\pi}{10}z)$ on top and for $\alpha(z) = \sin(\frac{\pi}{10}z) + 0.00071z$ on the bottom. The value of the parameter *m* is $m = \frac{2}{3}$. These plots strongly suggest that the classical string motion is integrable in the former case, while a small deformation away from that special solution results in chaotic motion. (b) Here, we have chosen $m = \frac{3}{4}$ and $\alpha(z) = \sin(\frac{\pi}{10}z)$ for the top, while $\alpha(z) = \sin(\frac{\pi}{10}z) + 0.00071z$ for the bottom. Regarding the status of (non) integrability, we reach the same conclusions.

We plot the LCEs for two examples. Following the discussion above, it should be clear that there no chaotic indicators for the choice $\alpha(z) = A \sin(\omega z)$, while for the simple quiver given by Eq. (5.1) the motion is chaotic, and, thus, there is no chance of Liouville integrability. We present our computations in Fig. 10.

C. Poincaré sections

Let us start by considering a ν -dimensional integrable system. The system has ν independent integrals of motion in involution, namely, that the Poisson bracket of any two of these conserved quantities is zero. This, in turn, implies that the corresponding phase-space trajectories are confined to the surface of an ν -dimensional Kolmogorov-Arnold-Moser (KAM) torus [61]. Upon a change of variables (q_i, p_i) with the so-called action-angle variables (ψ_i, G_i) , in such a way that the Hamiltonian describing the system depends on G_i only, the related classical paths on this KAM torus are specified entirely in terms of the ν frequencies, denoted by (ω_i) , which give the associated velocities along each of these angles. If there does not exist a set of integers η^i such that $\omega_i \eta^i = 0$, these trajectories will never close on themselves and, instead, will fill the surface of a KAM torus.⁹

The previous results imply that by simply taking cross sections of the phase-space trajectories we should be able to determine whether or not a system is integrable. As an example, if we plot the pair of data (ψ_1, G_1) each time $\psi_2 = \psi_{2,0}$ for $\psi_{2,0}$ an arbitrary number, we will obtain a set of embedded closed KAM trajectories associated to the cross sections of the embedded KAM tori. This cross

⁹See the application of these numerical methods in the context of $AdS_5 \times T^{1,1}$ [31] and much more directly related to our studies in [38] for the mother AdS_7 backgrounds.



FIG. 12. Poincaré sections. (a) Some indicative Poincaré sections for the quiver with defining function $\alpha(z) = \sin(\frac{\pi}{10}z)$ on top and for $\alpha(z) = \sin(\frac{\pi}{10}z) + 0.0953z$ on the bottom. The value of the parameter *m* is m = 1. These plots strongly suggest that the classical string motion is integrable in the former case, while a small deformation away from that special solution results in chaotic motion. (b) Here, we have chosen m = 1.79 and $\alpha(z) = \sin(\frac{\pi}{10}z)$ for the top, while $\alpha(z) = \sin(\frac{\pi}{10}z) + 0.835z$ for the bottom. Regarding the status of (non) integrability, we reach the same conclusions.

section is the so-called Poincaré section in the literature. The KAM theorem [61] is a statement about how these KAM trajectories change when we perturb an integrable Hamiltonian system \mathcal{H} by a small deformation $\epsilon \mathcal{H}_{new}$, with ϵ denoting a small number. Some of the KAM tori will no longer be closed-shaped curves upon the addition of the deformation. Let assume now that we increase the strength of the perturbation in the original integrable Hamiltonian system. This will result in the motion becoming completely random, and eventually all of the KAM tori will be random shapes.

To produce Poincaré sections for the supergravity backgrounds given by Eq. (2.1), we start by first picking a set of appropriate initial conditions. By appropriate, we mean that they have to be such that they are different but have the same energy *E*. We want to make sure that the Virasoro constraint is always satisfied for a given value of the energy, since the Virasoro constraint is a primary condition. We then allow the numerical evolution for the chosen initial points and plot conjugate pairs of coordinate and momentum at every time.

If the string soliton had an integrable motion, then the phase-space associated trajectories should be confined to some closed-shaped curves and, more specifically, tori. The Poincaré sections would then appear as circles generated by the various tori of different resonances. In the case of the absence of such closed, circular KAM tori, we conclude that the corresponding system is chaotic and, hence, nonintegrable.

We have already provided concrete evidence based on analytics and numerics that the regular, linear quivers, that can be interpreted as fully localized D8-branes in supergravity are chaotic and, thus, nonintegrable; here, we take a complementary approach.

We will exemplify in a very illustrative way just how special and properly chosen the solution $\alpha(z) = A \sin(\omega z)$ is.

To do so, we will present the Poincaré sections for this characteristic function. Having done so, we add to that choice of $\alpha(z)$ a very small deformation away from the perfectly chosen sinusoidal shape of the quiver.

We note that the choice

$$\alpha(z) = A\sin(\omega z) + \epsilon \cdot z, \qquad (5.18)$$

with ϵ being a small number, does not have a particular physical meaning. It is an interesting case study, though, mathematically. Indeed, as can be observed from the Poincaré sections presented in Figs. 11 and 12 that even the smallest deformation away from the sinusoidal quiver leads to signatures of chaos. In the Lyapunov exponent spectrum language, this would mean that one exponent is positive, and in the power spectrum language it means that the small deformation presents itself as the cause for the trajectories to lose their higher harmonics and the spectrum to be taken over by a band of noise.

Figures 11 and 12 present some examples that we provide for illustration and completeness. We have checked various other deformations away from the perfect sinusoidal quiver. The same qualitative answer is reached in every single study we performed.

This leads us to conclude that the choice for the characteristic function $\alpha(z) = A \sin(\omega z)$ is indeed the only choice that does not lead to chaotic string motion. Therefore, it is the only choice for which we can hope the dual field theory to be classically integrable.

Note, before ending this section, that these numerical findings are in accord with the brief analysis based on the isometries of the background that we presented in Sec. IV C. Our argument there was that, when $\alpha(z) = A \sin(\omega z)$, the warp factor in front of the AdS₅ submanifold is reduced to a constant, and, hence, our metric becomes the direct product of integrable submanifolds, up to the nontrivial fibration of the geometry. No matter how small a deformation way from that particular solution for $\alpha(z)$ is, the above argument can never hold true, since the warp factors will be nontrivial, even by a small amount.

VI. OUTLOOK AND CONCLUSIONS

In this work, we focused our attention on a recently discovered parametric family of supergravity backgrounds in the context of massive type IIA supergravity that are the dual descriptions of four-dimensional $\mathcal{N} = 1$ quiver theories developed in [22]. By performing a careful and detailed study of the dynamics of a string soliton, we provide very strong suggestive evidence that the dual field theory is classically integrable only for the choice $\alpha(z) = A \sin(\omega z)$, while all linear quivers are chaotic and, hence, not Liouville integrable. We have been careful with the choice of the string embedding such that it encapsulates the basic aspects of symmetries in the given class of solutions. Our studies have been based on both analytics and numerics.

In order to suggest that the sinusoidal quiver leads to an integrable theory, we used Kovacic's algorithm as our main tool with regards to analytics. It is only this choice that can pass all the conditions in all the respective cases. The polynomial rank quivers fail to satisfy said conditions.

For our numerical approach, we have provided evidence based on no signature of nonintegrability and no hallmark behavior of a chaotic system. To do so, we have studied the dynamical evolution of the string in time and examined how the dynamics are reflected on the plane by the other coordinates. Furthermore, we have computed the spectrum of Lyapunov exponents. Finally, we have studied the power spectra for these differently shaped quivers.

Below, we mention some exciting future directions.

- (1) We have already suggested that the physical interpretation of the special background, defined by $\alpha(z) = A \sin(\omega z)$, is in terms of smeared D8-branes. It would be very interesting to find the precise smearing form of this solution using the well-established framework and techniques that have been reviewed in [65].
- (2) Our focus here was on the \mathbb{H}_2 twisted compactification; however, the authors in [22] also studied T^2 and S^2 compactifications. In the latter two cases, however, there were some concerns raised regarding some pathologies of the resulting supergravity solutions. This is the reason that our focus here was on the \mathbb{H}_2 , since it is well under control. We want to point out, nevertheless, that T^2 and S^2 are maximally symmetric spaces as well, and, hence, an appropriate tuning of the defining function $\alpha(z)$ might lead to integrable string dynamics. For this reason, we believe that it would be very interesting to understand and systematically resolve the issues that were raised in [22].
- (3) It would be very interesting to understand aspects of probe D-branes in both the nonintegrable as well as the sinusoidal solution. There is a twofold motivation for these studies.
 - (a) Studying supersymmetric probe D-branes in a given background or class of backgrounds is relevant in extracting important information beyond the supergravity regime, namely, approaching the stringy regime. This was understood early on in the context of the $\mathcal{N} = 4$ SYM in [66] that the dual gravity picture must contain branes, such that it accommodates the Pfaffian operator for the SO(N) description or baryon vertices and domain walls in the SU(N) theory.
 - (b) Considering the dynamics of any given theory in the presence of boundaries is a very natural way to to probe it. The resulting defect theories are by themselves interesting, and, in the past, their studies have led to profound results and insight for the field theories [67,68].

- (4) It would also be very interesting to study and derive the spectrum of spin-2 operators in these backgrounds following the pioneering work of Bachas and Estes [69] and see how they fall into different representations of the superconformal algebra.
- (5) In addition to the above, a fascinating direction would be to carefully consider the Penrose limit, derive the corresponding pp-wave geometry [70] of the massive type IIA solutions, and perform the string quantization.
- (6) Another holographic observable of great interest that can be computed is the entanglement entropy [71–73].

And, of course, the three most important, and probably difficult, questions are the following.

- (i) What is the Lax connection, if one exists, for the sinusoidal solution in these new AdS₅ backgrounds?
- (ii) What is so special about these infinite linear quivers that arise in the context of massive type IIA theory that lead to integrability?
- (iii) Can we formulate a no-go theorem for integrability of a string background that has nontrivial warp factors? See Ref. [15] for related discussion.

We hope to be able to report on some of these in the near future.

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APPENDIX: THE KOVACIC ALGORITHM

In what follows, we summarize the basic notions of differential Galois theory, which formed the basis used by Kovacic [30] to construct an algorithm, pertaining to the existence of Liouville integrable solutions on second-order linear ordinary differential equations. By a Liouville integrable, closed-form solution, we are referring to one that is written in terms of algebraic, trigonometric, exponential functions or integrals of those.

The Kovacic algorithm has been presented in the physics literature in the past; see, for example, the appendixes in [38,42,43]. For notational convenience, we will follow [43]. We include case III of the algorithm, which to the best of our knowledge has not appeared before in the physics literature.

The theorem deals with second-order, linear, ordinary differential equations that assume the form

$$f''(x) + \mathcal{A}_1(x)f'(x) + \mathcal{A}_2(x)f(x) = 0,$$
 (A1)

where $x \in \mathbb{C}$ and $\mathcal{A}_{1,2}$ are rational complex functions. We can change variables using $f = e^{\frac{1}{2}\int dx \mathcal{A}_1 z}$ to eliminate the first-derivative term and obtain

$$z''(x) = \mathcal{V}(x)z(x), \qquad \mathcal{V} = \mathcal{A}_2 - \frac{1}{2}\mathcal{A}'_1 - \frac{1}{4}\mathcal{A}_1^2, \quad (A2)$$

where \mathcal{V} is called the potential. Obviously, f is Liouvillian integrable if and only if z is; there is no loss of generality under the change of variables.

The starting point of differential Galois theory for these type of equations, the Picard-Vessiot theory, is the group of automorphisms of its solutions, namely, $SL(2, \mathbb{C})$ and its possible subgroups. Let *X* be an algebraic subgroup of $SL(2, \mathbb{C})$; then one of the four cases can occur.

Case 1: X is triangulizable.

Case 2: Case 1 does not hold, and *X* is conjugate to a subgroup of

$$\left\{ \begin{pmatrix} a & 0 \\ 0 & a^{-1} \end{pmatrix} \middle| a \in \mathbb{C}, a \neq 0 \right\} \cup \left\{ \begin{pmatrix} 0 & c \\ -a^{-1} & 0 \end{pmatrix} \middle| a \in \mathbb{C}, a \neq 0 \right\}.$$
(A3)

Case 3: Case 1 does not hold, and the same is true for case 2 and X is finite.

Case 4: $X = SL(2, \mathbb{C})$.

If the differential equation satisfies one of the three first cases, it has Liouville-integrable solutions. Otherwise, in the case $X = SL(2, \mathbb{C})$, no such solutions exist.

Kovacic managed to translate the first three cases, cases 1-3, into algebraic arguments. These arguments are applied on the behavior of the potential we defined earlier, V, in Eq. (A2). These algebraic conditions make up the theorem that is stated as [30].

Theorem. The following necessary conditions have to hold for a particular case to be true.

Case 1: Every pole of \mathcal{V} must either have even order or else order 1. The order of \mathcal{V} at ∞ must be either even or else greater than 2.

Case 2: \mathcal{V} must have at least one pole that either has odd order greater than 2 or else has order exactly 2.

Case 3: The order of a pole of \mathcal{V} cannot exceed 2, and the order of \mathcal{V} at ∞ must be at least 2.

If $\mathcal{V} = \text{num/denom}$, then the poles of \mathcal{V} are the zeros of denom and the order of the pole is the multiplicity of the zero of denom. The order of \mathcal{V} at ∞ denotes the number deg denom—deg num, with deg denoting the degree of the polynomial.

These conditions are *necessary* for the relevant cases to be true, ergo, failing these conditions *sufficient* for case 4 to hold. Hence, we conclude that the failure of all three conditions is enough to declare the differential equation given by Eq. (A2) as Liouville nonintegrable.

However, if any of the conditions is satisfied, then the associated case might hold, and, assuming it does, then a Liouville integrable solution exists. So, when a condition is met we have to follow precisely the subalgorithm of the respective case to determine if such a solution exists. If a Liouvillian solutions exists, we can use the algorithm to find it.

Furthermore, Kovacic was able to produce the algorithmic steps for the three cases; see Ref. [30]. We present in subsections the algorithm for each case.

1. The case 1 algorithm

We assume that we are under the condition in which case 1 holds. We use Γ to denote the set of poles of \mathcal{V} .

Step 1.—For each $c \in \Gamma \cup \{\infty\}$, we define two complex numbers α_c^{\pm} and a rational function $[\sqrt{\mathcal{V}}]_c$ in the following way:

 (c_1) If $c \in \Gamma$ and c is an order 1 pole, then

$$[\sqrt{\mathcal{V}}]_c = 0, \qquad \alpha_c^{\pm} = 1.$$

 (c_2) If $c \in \Gamma$ and c is an order 2 pole, then

$$\left[\sqrt{\mathcal{V}}\right]_c = 0.$$

Let β_c be the coefficient of the term $1/(x-c)^2$ in the expansion in partial fractions for \mathcal{V} . Then,

$$\alpha_c^{\pm} = \frac{1}{2} \pm \frac{1}{2}\sqrt{1+4\beta_c}.$$

 (c_3) If $c \in \Gamma$ and c is an order $2\nu \ge 4$ pole, which in case 1 is necessarily an even number, then $[\sqrt{\mathcal{V}}]_c$ is the sum of all the terms that involve $1/(x-c)^i$ for $2 \le i \le \nu$ in the Laurent series expansion of $\sqrt{\mathcal{V}}$ at c. There are two possibilities for $[\sqrt{\mathcal{V}}]_c$ that are equal up to the overall sign, and we can choose either one of them. Hence,

$$[\sqrt{\mathcal{V}}]_c = \frac{a}{(x-c)^{\nu}} + \dots + \frac{d}{(x-c)^2}.$$

Let β_c be the coefficient of the term $1/(x-c)^{\nu+1}$ in \mathcal{V} minus the coefficient of the term $1/(x-c)^{\nu+1}$ in $[\sqrt{\mathcal{V}}]_c^2$. We have

$$\alpha_c^{\pm} = \frac{1}{2} \left(\pm \frac{\beta_c}{a} + \nu \right).$$

 (∞_1) If the order of \mathcal{V} at ∞ is bigger than 2, then

$$\left[\sqrt{\mathcal{V}}\right]_{\infty} = 0, \qquad \alpha_{\infty}^{+} = 0, \qquad \alpha_{\infty}^{-} = 1.$$

 (∞_2) If the order of \mathcal{V} at ∞ is equal to 2, then

$$[\sqrt{\mathcal{V}}]_{\infty} = 0.$$

Let b_{∞} be the coefficient of the term $1/x^2$ in the Laurent series expansion of \mathcal{V} at ∞ . (If $\mathcal{V} = \text{num/denom}$, where num and denom are coprime, then b_{∞} is the leading coefficient of num divided by the leading coefficient of denom.) Then,

$$\alpha_{\infty}^{\pm} = \frac{1}{2} \pm \frac{1}{2}\sqrt{1+4\beta_{\infty}}.$$

 (∞_3) If the order of \mathcal{V} at ∞ is $-2\nu \leq 0$, which in case 1 is necessarily an even number, then $[\sqrt{\mathcal{V}}]_{\infty}$ is the sum of all terms involving x^i for $0 \leq i \leq \nu$ in the Laurent series for $\sqrt{\mathcal{V}}$ at ∞ . We can choose either of the two possibilities. Ergo,

$$[\sqrt{\mathcal{V}}]_{\infty} = ax^{\nu} + \dots + d.$$

Let β_{∞} be the coefficient of $x^{\nu-1}$ in \mathcal{V} minus the coefficient of $x^{\nu-1}$ in $([\sqrt{\mathcal{V}}]_{\infty})^2$. Then,

$$\alpha_{\infty}^{\pm} = \frac{1}{2} \left(\pm \frac{\beta_{\infty}}{a} - \nu \right).$$

Step 2.—For each family $num = (num(c))_{c \in \Gamma \cup \{\infty\}}$, where num(c) is + or -, let

$$d = \alpha_{\infty}^{s(\infty)} - \sum_{c \in \Gamma} \alpha_c^{s(c)}.$$

If d is a non-negative integer, then

$$\omega = \sum_{c \in \Gamma} \left(s(c) [\sqrt{\mathcal{V}}]_c + \frac{\alpha_c^{s(c)}}{x - c} \right) + s(\infty) [\sqrt{\mathcal{V}}]_{\infty}$$

is a candidate for ω . If *d* is not a non-negative integer, then the family num can be removed from consideration.

Step 3.—We have to apply this step to each one of the families retained from step 2, until we succeed or all of the families have been exhausted. In the latter case, case 1 does not hold.

For each family, we search for a monic polynomial P of degree d (as defined in step 2) that satisfies the differential equation

$$P'' + 2\omega P' + (\omega' + \omega^2 - \mathcal{V})P = 0.$$

This we can achieve conveniently by using the undetermined coefficients and is a simple linear-algebra problem. Note that the problem may or may not have a solution. If such a polynomial exists, then $\eta = Pe^{\int \omega}$ is a solution of the differential equation (A2). If there is no such polynomial for any family retained from step 2, then case 1 does not hold.

2. The case 2 algorithm

In order for this algorithm to take effect, we assume that case 1 fails. We work similarly to case 1; namely, we first gather the data for each pole c of V and also for ∞ . The data form a set that is either E_c or E_{∞} and consists of some integer numbers, from one to three. Next, we consider the families of elements of these sets, discarding some and retaining others. If no families are retained, case 2 does not hold. However, for each family that is retained, we search for a monic polynomial that obeys a certain linear differential equation. If no such polynomial exists for any family, then case 2 does not hold. If such a polynomial exists, then we have obtained a solution to the differential equation given by Eq. (A2).

We use Γ to denote the set of poles of \mathcal{V} .

Step 1.—For each $c \in \Gamma$, we define E_c as follows:

(c₁) If c is an order 1 pole, then $E_c = \{4\}$.

 (c_2) If c is an order 2 pole and if β_c is the coefficient of $1/(x-c)^2$ in the expansion in partial fractions of \mathcal{V} , then

$$E_c = \{2 + k\sqrt{1 + 4\beta_c} | k = 0, \pm 2\} \cap \mathbb{Z}.$$

(c₃) If c is an order $\nu > 2$ pole, then $E_c = \{\nu\}$.

 (∞_1) If \mathcal{V} has order bigger than 2 at ∞ , then $E_{\infty} = \{0, 2, 4\}.$

 (∞_2) If \mathcal{V} has order exactly 2 at ∞ and β_{∞} is the coefficient of \mathcal{V} in the Laurent series expansion of \mathcal{V} at ∞ , then

$$E_{\infty} = \{2 + k\sqrt{1 + 4\beta_{\infty}} | k = 0, \pm 2\} \cap \mathbb{Z}$$

 (∞_3) If the order of \mathcal{V} at ∞ is $\nu < 2$, then $E_{\infty} = \{\nu\}$. Step 2.—We account for all families $(e_c)_{c \in \Gamma \cup \{\infty\}}$ with $e_c \in E_c$. Any family that has all its coordinates to be even can be discarded. Consider

$$d = \frac{1}{2} \left(e_{\infty} - \sum_{c \in \Gamma} e_c \right).$$

If d is a non-negative integer, then the family should be retained. If not, then the family is discarded. If there are no families retained, then case 2 does not hold.

Step 3.—For each of the families retained from step 2, we construct the rational function

$$\theta = \frac{1}{2} \sum_{c \in \Gamma} \frac{e_c}{x - c}.$$

We next search for a monic polynomial P of degree d, as defined in step 2, such that it satisfies

$$\begin{aligned} P^{\prime\prime\prime\prime} + 3\theta P^{\prime\prime} + (3\theta^2 + 3\theta^\prime - 4\mathcal{V})P^\prime \\ + (\theta^{\prime\prime} + 3\theta\theta^\prime + \theta^3 - 4\mathcal{V}\theta - 2\mathcal{V}^\prime)P &= 0. \end{aligned}$$

If we cannot find such a polynomial for any of the families retained from step 2, then case 2 does not hold.

Let us assume that we have found such a polynomial. Let us define $\varphi = \theta + P'/P$, and let ω be a solution of the equation

$$\omega^2 + \varphi \omega + \left(\frac{1}{2}\varphi' + \frac{1}{2}\varphi^2 - \mathcal{V}\right) = 0.$$

Then, $\eta = e^{\int \omega}$ is a solution of the differential equation given by Eq. (A2).

3. The case 3 algorithm

In case 3, we assume that cases 1 and 2 fail and the differential equation has only algebraic solutions. Let us assume η is a solution of the differential equation $z'' = \mathcal{V}z$, and let us define $\omega = \frac{\eta'}{\eta}$. In this case, ω is algebraic over $\mathbb{C}(x)$ of degree 4, 6, or 12, and it is the minimal polynomial for ω that we will have to find.

To do so, there are two possible ways. On the one hand, we can find a polynomial of degree 12 and then factor it. If ω is any solution of the 12th degree polynomial, then $\eta = e^{\int \omega}$ is a solution to the differential equation, and, hence, any of the irreducible factors can be used. On the other hand, we can first attempt to find a 4th degree equation for ω , subsequently a 6th degree equation, and finally a 12th degree equation. If an equation is found, it is guaranteed to be irreducible.

Let Γ be the set of poles of \mathcal{V} . Recall that, by the necessary conditions of case 2, \mathcal{V} cannot have a pole of order higher than 2.

Step 1.—For each $c \in \Gamma \cup \{\infty\}$, we define E_c as follows: (c₁) If c is an order 1 pole, then $E_c = \{12\}$.

 (c_2) If c is an order 2 pole and if α_c is the coefficient of the term $1/(x-c)^2$ in the expansion of \mathcal{V} in partial fractions, then

$$E_{c} = \left\{ 6 + \frac{12k}{n} \sqrt{1 + 4\alpha_{c}} | k = 0, \pm 1, \pm 2, \dots, \pm \frac{n}{2} \right\} \cap \mathbb{Z}.$$

 (∞) If the Laurent series for \mathcal{V} at ∞ has the form

$$\mathcal{V} = \gamma x^{-2} + \cdots (\gamma \in \mathbb{C}, \text{ possibly } 0),$$

then

$$E_{\infty} = \left\{ 6 + \frac{12k}{n} \sqrt{1 + 4\gamma} | k = 0, \pm 1, \pm 2, \dots, \pm \frac{n}{2} \right\} \cap \mathbb{Z}.$$

Step 2.—We consider all families $(e_c)_{c \in \Gamma \cup \{\infty\}}$ with $e_c \in E_c$. For each family, define

$$d = \frac{n}{12} \left(e_{\infty} - \sum_{c \in \Gamma} e_c \right).$$

If d is a non-negative integer, we should retain the family; in any other case, the family has to be discarded. If no families remain under consideration, then case 3 does not hold.

Step 3.—For each of the families from step 2, we construct the rational function

$$\theta = \frac{n}{2} \sum_{c \in \Gamma} \frac{e_c}{x - c}$$

and we further construct

$$S = \prod_{c \in \Gamma} (x - c).$$

Next, we look for a monic polynomial $P \in \mathbb{C}(x)$ that has degree *d* (as defined in step 2) in such a way that when we compute the polynomials $P_n, P_{n-1}, \dots, P_{-1}$ in a recursive manner by the formulas:

$$\begin{split} P_n &= -P, \\ P_{i-1} &= -SP'_i + ((n-i)S' - S\theta)P_i - (n-i)(i+1)S^2rP_{i+1}, \\ &\text{with } i = n, n-1, \dots, 0, \end{split}$$

then $P_{-1} = 0$ identically.

If we cannot find such a polynomial for any of the families retained from step 2, then case 3 does not hold.

Let us assume that we have found such a polynomial. Consider $\varphi = \theta + P'/P$ and let ω be a solution of the equation:

$$\sum_{i=0}^{n} \frac{S^{i} P_{i}}{(n-i)!} \omega^{i} = 0.$$

Then $\eta = e^{\int \omega}$ is a solution of Eq. (A2).

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