Particle spectra of general Ricci-type Palatini or metric-affine theories

W. Barker^{1,2,*} and C. Marzo^{3,†}

¹Astrophysics Group, Cavendish Laboratory, JJ Thomson Avenue, Cambridge CB3 0HE, United Kingdom
 ²Kavli Institute for Cosmology, Madingley Road, Cambridge CB3 0HA, United Kingdom
 ³Laboratory for High Energy and Computational Physics, NICPB, Rävala 10, Tallinn 10143, Estonia

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In the context of weak-field metric-affine (i.e. Palatini) gravity near Minkowski spacetime, we compute the particle spectra in the simultaneous presence of all independent contractions quadratic in Ricci-type tensors. Apart from the full metric-affine geometry, we study kinematic limits with vanishing torsion (i.e. a symmetric connection) and vanishing nonmetricity (i.e. a metric connection, which is physically indistinguishable from Poincaré gauge theory at the level of the particle spectrum). We present a detailed report on how spin-parity projection operators can be used to derive systematically and unambiguously the character of the propagating states. The unitarity constraints derived from the requirements of tachyon and ghost freedom are obtained. We show that, even in the presence of all Ricci-type operators, only a narrow selection of viable theories emerges by a tuning.

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I. INTRODUCTION

The success of the curvature-based geometrical formulation of gravity has stimulated a search for the dynamical interpretation of geometrical properties analogous to curvature. This program has been directed to solve, or mitigate, some of the main shortcomings of the current status quo, mainly the lack of a perturbatively renormalizable quantum theory of gravity, and the phenomenological need for a dark sector. While hopes to address the renormalization issue have, so far, had to rely on nontraditional routes [1-13], the possibility of interesting phenomenology from particles of geometrical origin is frequently used in cosmology and dark-sector model building [14–27]. Metric-affine gravity (MAG) [28–49] represents the principal realization of this program. It broadens the matter content by considering the affine connection as an independent three-index object, giving rise to torsion and nonmetricity (see Fig. 1) as dynamical fields. This results in a notable growth in computational complexity, both in the number of allowed operators and in the profiling of the multiple particle states carried in by the affine connection. Such a broad parameter space is expected, as is often the case, to narrow under the pressure of field-theoretical self-consistency

constraints, such as unitarity and elimination of tachyons [28,32,33,35,44,45,50–57]. The imposition of these requirements is a highly nontrivial task and has often necessitated severe simplifications to arrive at positive scenarios. Also, the analysis often has to rely on a very indirect route without directly accessing the pole structure of the propagator.

In this paper, we use the arena of Ricci-type MAG to illustrate, in a detailed step-by-step fashion, how the formalism of spin-parity projectors can unambiguously and straightforwardly reveal the nature of the (tree-level) particle spectrum. By *Ricci-type* we refer to all the operators that may be added to the Einstein-Hilbert Lagrangian which are quadratic in the rank-two traces of metric-affine curvature (in MAG there are nine such operators, whilst in standard GR there is only one). Building on early work (see e.g. [58–63]), this particular space of operators was first properly charted by Annala and Räsänen [35]. The reparametrization-based methods used in that work are particularly innovative, but they are



FIG. 1. The effects of metric-affine curvature $\mathcal{F}_{\mu\nu}{}^{\rho}{}_{\sigma}$ in Eq. (2) (vector rotation after parallel transport in a closed loop) torsion $\mathcal{T}_{\mu}{}^{a}{}_{\nu}{}_{\nu}$ in Eq. (1) [nonclosure of parallelograms formed from (infinitesimally) parallel-transported vectors] and nonmetricity $\mathcal{Q}_{\lambda\mu\nu}$ in Eq. (1) (change in vector length under parallel transport).

^{*}wb263@cam.ac.uk [†]carlo.marzo@kbfi.ee

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only able to access a "punctuated" bulk of the full parameter space due to certain degeneracy conditions which must be avoided. In the present work we build on these foundations by applying the spin-parity projection formalism: our method does not come with any restriction on the parameters of the model. Indeed, if the set of projector operators is known, the approach adopted can be applied to every tensor-valued Lagrangian admitting a Minkowskian background expansion. We have authored a new Wolfram Language implementation of this procedure for all such theories, including the MAG: Particle Spectrum for any Tensor Lagrangian (PSALTer). The PSALTer software will be properly presented in a dedicated paper [64],¹ but in this paper we use it to confirm all our results (see Appendices A to C). The body this paper is set out as follows. In Sec. II we develop the underlying theory by setting out the MAG conventions in Sec. II A and briefly presenting our spectral algorithm in Sec. II B and the computer implementation in Sec. II C. In Sec. III we present all of the results, respectively for the vanishing nonmetricity kinematic limit in Sec. III A and the vanishing torsion kinematic limit in Sec. III B, and for the full MAG in Sec. III C. Conclusions follow in Sec. IV. We will use the "West Coast" signature (+, -, -, -), other conventions will be introduced as needed.

II. THEORETICAL DEVELOPMENT

The requirements of locality and Lorentz invariance select tensor fields as the building blocks for most of the theoretical speculations about high-energy physics. The price to pay for using them is in the risk of uncontrolled unitarity violations. It is a profound realization that the consistent adjustments to recover unitarity restrict us to Maxwell's theory, for rank-one fields, as well as Einstein's theory of gravity, when applied to models of symmetric rank-two fields [50,68–71]. The continuation of this program within the intricate scenarios of higher-rank fields, as well as multifield quadratic Lagrangians, is technically prohibitive. Many indirect shortcuts have supported claims of healthy particle propagation, but the proliferation of indices often prevents a direct approach to the pole structure of the propagator. The spin-parity projection approach to spectral analysis has been most thoroughly expostulated by Lin, Hobson, and Lasenby in [55], with follow-up papers in [56,72]. Further information about the method can also be found in, e.g., [28,33,50-52,54,73]but it is [55] which provides the most concise introduction for our purposes.

A. Spin-parity kinematics of metric-affine gravity

To have full control over the particle spectrum we introduce operators which project the Lorentz index structure onto the labels J and P of spin and parity, enumerating the irreducible representation under the SU(2) little group. In general, further reduction to the U(1) little group of massless particles and helicity states can be done. This results in no practical improvement, making the decomposition into SU(2) representations generic enough. In the case of MAG, the needed set of projection operators has been recently completed [28,74], developing the past studies of [32,33,50–52,54,75]. We present here a summary of the main ideas regarding the use of projection operators for the computation of the poles and residues of the propagator. We refer to [28,51,55,76] for further details.

The independent MAG connection $A_{\mu \rho}^{\nu}$ carries in 64 new degrees of freedom (d.o.f.), which are not present in the derived Levi-Civita connection $\Gamma_{\mu \rho}^{\nu} \equiv \Gamma_{(\mu \rho)}^{\nu} \equiv g^{\nu\lambda} (\partial_{(\mu}g_{\rho)\lambda} - \frac{1}{2}\partial_{\lambda}g_{\mu\rho})$. These new d.o.f. are often partitioned into the torsion and nonmetricity tensors

$$\mathcal{T}_{\mu}{}^{\alpha}{}_{\nu} \equiv 2A_{[\mu]}{}^{\alpha}{}_{|\nu]}, \qquad \mathcal{Q}_{\lambda\mu\nu} \equiv -\partial_{\lambda}g_{\mu\nu} + 2A_{\lambda}{}^{\alpha}{}_{(\mu]}g_{\alpha|\nu)}. \tag{1}$$

The two tensors in Eq. (1) are geometric counterparts to the metric-affine (i.e. non-Riemannian) curvature

$$\mathcal{F}_{\mu\nu}{}^{\rho}{}_{\sigma} \equiv 2(\partial_{[\mu}A_{\nu]}{}^{\rho}{}_{\sigma} + A_{[\mu]}{}^{\rho}{}_{\alpha}A_{|\nu]}{}^{\alpha}{}_{\sigma}). \tag{2}$$

The influences of these three geometric properties on vectors which are being parallel transported through the spacetime are shown in Fig. 1. We will be particularly interested in the three *Ricci-type* contractions of the metric-affine curvature

$$\mathcal{F}_{\mu\nu} \equiv \mathcal{F}_{\mu\nu\alpha}{}^{\alpha}, \quad \mathcal{F}^{(14)}{}_{\mu\nu} \equiv \mathcal{F}_{\alpha\mu\nu}{}^{\alpha}, \quad \mathcal{F}^{(13)}{}_{\mu\nu} \equiv \mathcal{F}_{\alpha\mu}{}^{\alpha}{}_{\nu}, \quad (3)$$

by which we just mean the contractions with two free indices. The quantity $\mathcal{F}_{\mu\nu}$ is the *homothetic* curvature [77,78], while $\mathcal{F}^{(13)}_{\mu\nu}$ and $\mathcal{F}^{(14)}_{\mu\nu}$ are variously *pseudo*-Ricci tensors [28], or $\mathcal{F}^{(14)}_{\mu\nu}$ is the *co*-Ricci [77]. The Riemannian curvature of course yields only one Ricci-type contraction: the homothetic curvature vanishes identically in the absence of nonmetricity, and the (co)Ricci tensors coincide in the absence of torsion. We define $\mathcal{F} \equiv \mathcal{F}_{\mu\nu}^{\mu\nu}$ as the Ricci scalar: there is still only one such scalar in metric-affine geometry.

In the *first-order* or *Palatini* parametrization of MAG, we take the 10 d.o.f. in $g_{\mu\nu}$ and the 64 new d.o.f. in $A_{\mu}{}^{\nu}{}_{\rho}$ to be fundamental fields. An advantage of the first-order parametrization is that the MAG field strength tensor in Eq. (2) is free from second derivatives. In the *second-order* or *post-Riemannian* parametrization, we keep $g_{\mu\nu}$ but treat the tensor-valued difference $\Delta_{\mu}{}^{\nu}{}_{\rho} \equiv A_{\mu}{}^{\nu}{}_{\rho} - \Gamma_{\mu}{}^{\nu}{}_{\rho}$

¹See [65] for a recent application of *PSALTer* to theories proposed in [66,67].

(sometimes termed the *distortion*) as fundamental, effectively partitioning the 64 d.o.f. among the 24 d.o.f. in $\mathcal{T}_{\mu}{}^{\alpha}{}_{\nu} \equiv \mathcal{T}_{[\mu]}{}^{\alpha}{}_{|\nu]}$ plus the 40 d.o.f. in $\mathcal{Q}_{\lambda\mu\nu} \equiv \mathcal{Q}_{\lambda(\mu\nu)}$ according to

$$\mathcal{T}_{\mu}{}^{\alpha}_{\nu} \equiv 2\Delta_{[\mu]}{}^{\alpha}_{|\nu]}, \qquad \mathcal{Q}_{\lambda\mu\nu} \equiv 2\Delta_{\lambda(\mu\nu)}. \tag{4}$$

In the second-order parametrization Eq. (2) is expanded into the Riemannian curvature (which naturally has second derivatives in $g_{\mu\nu}$) and many other terms which are (Levi-Civita) covariant derivatives and second powers of $\mathcal{T}_{\mu}{}^{\alpha}{}_{\nu}$ and $\mathcal{Q}_{\lambda\mu\nu}$. The second-order parametrization has an advantage in revealing the true nature of all MAG-type theories: every MAG theory is a (non)minimal coupling of standard metric-based gravity to an asymmetric rankthree matter field $\Delta_{\mu}{}^{\rho}{}_{\nu}$. We are free to work with either set of variables due to reparametrization invariance of the physics.

Working in the first-order formulation, the weak-field regime near to Minkowski spacetime can be captured by an inherently perturbative $A_{\mu \ \rho}^{\ \nu}$ and a metric perturbation $g_{\mu\nu} \equiv \eta_{\mu\nu} + h_{\mu\nu}$. These perturbations carry multiple massive particle states

$$A_{\mu\nu\rho} \supset \left\{ 3_{1}^{-}, 2_{1}^{+}, 2_{2}^{+}, 2_{3}^{+}, 2_{1}^{-}, 2_{2}^{-}, 1_{1}^{+}, 1_{2}^{+}, 1_{3}^{+}, 1_{1}^{-}, 1_{2}^{-}, \right. \\ \left. 1_{3}^{-}, 1_{4}^{-}, 1_{5}^{-}, 1_{6}^{-}, 0_{1}^{+}, 0_{2}^{+}, 0_{3}^{+}, 0_{4}^{+}, 0_{1}^{-} \right\},$$

$$(5a)$$

$$h_{\mu\nu} \supset \{2_4^+, 1_7^-, 0_5^+, 0_6^+\},\tag{5b}$$

where we used the compact notation J_j^P in referring to the *j*th representation of spin *J* and parity *P*. Enumeration convention is adopted from [28]. From Eqs. (5a) and (5b) the 64 and 10 d.o.f. can be recovered, respectively, by summing the multiplicities 2J + 1 over all states. The *PSALTer* notation for these states differs from the subscript notation in Eqs. (5a) and (5b), and full definitions of these states are provided in Fig. 11.

B. Saturated propagator and particle spectra

Following [76], we use a synthetic notation to describe the quadratic (i.e. perturbative) action in momentum space

$$S[\Phi] = \frac{1}{2} \int d^4k \big(\Phi(-k), K(k) \Phi(k) + \mathcal{J}(-k) \Phi(k) + \mathcal{J}(k) \Phi(-k) \big), \qquad (6)$$

where K(k) is the Fourier-transformed kinetic term (wave operator) and we have introduced a linear coupling between the fields, collectively labeled as $\Phi(k)$, and a source $\mathcal{J}(k)$. Connecting this to the specific formulation in Sec. II A, we identify Φ as the collection of perturbative fields $h_{\mu\nu}$ and $A_{\mu}{}^{\rho}{}_{\nu}$. Within the quadratic approximation, the indices on these fields are raised and lowered using $\eta^{\mu\nu}$ and $\eta_{\mu\nu}$, which are nondynamical, and the greek indices refer to Cartesian coordinates on the Minkowski background.² Conjugate to Φ , the fields \mathcal{J} in MAG are the symmetric matter stressenergy tensor $T^{\mu\nu}$ and the rank-three current $W^{\mu}{}_{\rho}{}^{\nu}$. This latter current is sometimes called the *hypermomentum* [14,27,79–83]. In the second-order formulation, a separate current must be defined as conjugate to $\Delta_{\mu}{}^{\rho}{}_{\nu}$ —we do not bother to ascribe it another symbol.

The propagating states will appear in the form of isolated poles for the propagator $\mathcal{D}(k)$ defined through the equation $K(k) \cdot \mathcal{D}(k) \equiv \hat{1}$. The use of projection operators $P_{\{J,P\}}^{i,i}$ drastically simplifies the solution of this inversion problem. By exploiting the defining properties

$$\sum_{J,P,i} P^{i,i}_{\{J,P\}\mu_1\mu_2\cdots\mu_n} {}^{\nu_1\nu_2\cdots\nu_n} \equiv \hat{1}_{\mu_1\mu_2\cdots\mu_n} {}^{\nu_1\nu_2\cdots\nu_n}, \qquad (7a)$$

$$P_{\{J,P\}\mu_{1}\mu_{2}\cdots\mu_{n}}^{i,k}P_{\{J',P'\}\rho_{1}\rho_{2}\cdots\rho_{n}}^{j,w}P_{\{J',P'\}\rho_{1}\rho_{2}\cdots\rho_{n}}^{\nu_{1}\nu_{2}\cdots\nu_{n}}$$

$$\equiv \delta_{k,j}\delta_{J,J'}\delta_{P,P'}P_{\{J,P\}\mu_{1}\mu_{2}\cdots\mu_{n}}^{i,w},$$
(7b)

$$P_{\{J,P\}\mu_{1}\mu_{2}\cdots\mu_{n}}^{i,j} \equiv \left(P_{\{J,P\}}^{j,i} \mu_{1}\nu_{2}\cdots\nu_{n} \atop \mu_{1}\mu_{2}\cdots\mu_{n}\right)^{*}, \quad (7c)$$

it is possible to decompose the kinetic term as

$$\int d^{4}k \Phi(-k) K(k) \Phi(k) = \int d^{4}k \Phi(-k) \sum_{J,P,i,j} (a_{i,j}^{\{J,P\}} P_{\{J,P\}}^{i,j}) \Phi(k), \qquad (8)$$

where the tortuous fabric of the indices is reformulated in terms of the simpler spin-parity matrices $a_{i,j}^{\{J,P\}}$, obtained by tracing over the Lorentz indices

$$a_{i,j}^{\{J,P\}} \equiv \frac{1}{2J-1} \operatorname{Tr} P_{\{J,P\}}^{i,j} \mathcal{K}(k).$$
(9)

The orthonormality of the projection operators in Eq. (7b) reduces the computation of the propagator $\mathcal{D}(k)$ to an inversion problem for the matrices $a_{i,i}^{\{J,P\}}$

²Indeed, it should be explicitly stated that along with these raising and lowering rules $A_{\mu}{}^{\rho}{}_{\nu}$ is a dynamical *tensor* field in the quadratic approximation. Geometrically, $A_{\mu}{}^{\rho}{}_{\nu}$ is a connection, but the physics does not know or care about the geometric foundations of the theory: only the representations of the particle states are important. Of course, $\Delta_{\mu}{}^{\rho}{}_{\nu}{}_{\nu}$ is already geometrically a tensor in the second-order formulation of MAG. The reparametrization is shown explicitly in Eq. (B1), and the key point is that the linearized Levi-Civita connection is also tensorial at lowest order in perturbations because the partial derivative ∂_{μ} is covariant at that order.

$$\mathcal{D}(k) \equiv \sum_{J,P,i,j} b_{i,j}^{\{J,P\}} P_{\{J,P\}}^{i,j}, \qquad b_{i,j}^{\{J,P\}} \equiv (a_{i,j}^{\{J,P\}})^{-1}.$$
(10)

When the degeneracy of the spin-parity matrices hampers this inversion, the model displays gauge symmetry. It is a bonus of this formalism that full control over the gauge symmetries is provided by the inspection of such simple objects as the null vectors $X_i^{r=1,2,...,n}$ of $a_{i,j}^{\{J,P\}}$, so that manipulations of linear algebra replace tensor operations. In turn, this promotes a code-friendly implementation of the spectral problem. The matrix structure not only allows the identification of symmetries which are already present in the model, but it also allows parameter-tunings to be identified which lead to the emergence of new symmetries. Finally, the gauge-invariant *saturated* propagator is obtained by restricting the inversion to the nondegenerate subspace in the spin-parity matrices, and contracting the inverse matrix with constrained sources $\tilde{J}(k)$

$$\mathcal{D}_{\mathcal{S}}(k) \equiv \tilde{J}^{*}(k) \left(\sum_{J,P,i,j} \tilde{b}_{i,j}^{\{J,P\}} P_{\{J,P\}}^{i,j} \right) \tilde{J}(k), \quad (11a)$$

$$X_{j}^{*r}P_{\{J,P\}}^{i,j}\tilde{J}(k) \equiv 0, \qquad r = 1, 2, ..., n.$$
 (11b)

C. Computer implementation

The saturated propagator in Eq. (11a) represents the arrival point of our computation. Previous efforts in MAG, even within the projection operators approach, generally relied on an indirect determination for the signs of the residues of the poles in $\mathcal{D}_{S}(k)$. Such methods avoid the difficult computation of the source constraint equations Eq. (11b) in the massless limit. However, as detailed in [55,64,76], the constraint equations can be fully resolved by choosing a suitable reference frame (with $k^2 = 0$ as a limit of $k^2 > 0$). This technique removes the main theoretical challenges encountered in the computation of the spectrum leaving, as the only limitation, the technical bounds in manipulating large expressions. The simplifications afforded by the use of projection operators, as well as the unambiguous computational procedure, have encouraged the development of opportune tools to promptly and automatically tackle the particle spectrum. Some of the most recent results on this subject [57,65] and the core of this work's conclusions, are obtained with the use of PSALTer [64], a Wolfram Language implementation of these ideas. The PSALTer software can automatically return the spectrum of any tensor-valued field theory up to rank-three.

The *PSALTer* analyses of key theories to be considered in Sec. III are presented in Figs. 3, 5–10, 12, and 13. These are vector graphics which contain the following information;

- (1) The linearized (quadratic) action in Eq. (6), in the position-space representation. This expression is the only input to the *PSALTer* software, and it is not necessary to perform any kind of decomposition of fields into irreducible parts.
- (2) Automatically computed: the elements $a_{i,j}^{\{J,P\}}$ in Eq. (8) which encode the wave operator of the theory, provided as one matrix for each spin sector. Because the theories considered here are not parity violating, it is reassuring to see that the matrices are always block-diagonal across parity-even (red) and parity-odd (blue) subsectors: the mixed-parity (purple) blocks are empty.
- (3) Automatically computed: the inverse $b_{i,j}^{\{J,P\}}$ matrices in Eq. (10) which encode the saturated propagator of the theory. We believe our implementation to be the first which uses Moore-Penrose inversion [84,85] (i.e. a uniquely defined gauge fixing) to obtain these coefficients.
- (4) Automatically computed: the source constraints X_i^r in Eq. (11b), which are guaranteed to encode all the gauge symmetries present in the theory.
- (5) Automatically computed: the spectrum of all massive and massless particles present in the theory. This includes information about the particle spin *J*, parity *P*, pole residue and mass. In the case of massless particles, there is no physical notion of spin which survives, but the number of independent polarizations is given.
- (6) Automatically computed: the overall unitarity conditions which must be imposed on the Lagrangian coupling coefficients. These conditions are derived from the above pole residues and masses, so as to support the no-ghost and no-tachyon criteria. There is of course no guarantee that such conditions exist, so the calculation is time limited to ten seconds. In case of "timeout," the masses and residues provide all the relevant information for further tuning the theory anyway.

Because these various outputs may be extremely cumbersome and have uncertain dimensions after typesetting, *PSALTer* uses a rectangle-packing algorithm to find the most economical layout for each theory: consequently some of the formulas in Figs. 3, 5–10, 12, and 13 are rotated on the page. The various SO(3) irrep definitions are provided separately, in Figs. 2, 4, and 11. Each of these figures defines a "kinematic module" for the software: a declaration of the fundamental tensor fields and their conjugate sources which are present in a class of theories. Within each module, the spectral analysis of infinitely many distinct models can be performed, depending on the admixture of operators in the quadratic action.

Two steps of the analysis are computationally expensive: the Moore-Penrose inversion and the evaluation of massless residues. When the theory contains more than two or three independent Lagrangian couplings (parameters) and tensor fields of rank three or more, these calculations start to pose a highly nontrivial computer algebra problem. Consequently, many subroutines in *PSALTer* are automatically parallelized to take advantage of the available infrastructure. For expedience, analysis of each theory in this work was performed using a dedicated compute node consisting of 112 Intel[®] Sapphire Rapids CPUs, or 64 AMD[®] Ryzen Threadripper CPUs, depending on availability. The former setup is close to the current state of the art in highperformance computing. The resulting throughput is very fast, and in fact each theory would only have taken approximately 20 minutes to process on a modern PC with four CPU cores.

III. RESULTS

Using the building blocks in Eq. (3) the most general Ricci-type MAG in the first-order formulation is

$$S[g,A] = -\frac{1}{2} \int d^4x \sqrt{-g} \Big[-a_0 \mathcal{F} \\ + \mathcal{F}^{(13)\mu\nu} \Big(c_7 \mathcal{F}^{(13)}{}_{\mu\nu} + c_8 \mathcal{F}^{(13)}{}_{\nu\mu} \Big) \\ + \mathcal{F}^{(14)\mu\nu} \Big(c_9 \mathcal{F}^{(14)}{}_{\mu\nu} + c_{10} \mathcal{F}^{(14)}{}_{\nu\mu} \Big) \\ + \mathcal{F}^{(14)\mu\nu} \Big(c_{11} \mathcal{F}^{(13)}{}_{\mu\nu} + c_{12} \mathcal{F}^{(13)}{}_{\nu\mu} \Big) \\ + \mathcal{F}^{\mu\nu} \Big(c_{13} \mathcal{F}_{\mu\nu} + c_{14} \mathcal{F}^{(13)}{}_{\mu\nu} + c_{15} \mathcal{F}^{(14)}{}_{\mu\nu} \Big) \Big], \quad (12)$$

where we borrow the numbering of Lagrangian couplings directly from [28].³ As emphasized in Sec. II A, reparametrization invariance leads to the equivalence $S[g, A] \cong$ $S[g, \Delta]$ with the second-order formalism. The computational algorithm sketched in Sec. II B can obviously handle both representations of the dynamical fields and our tests have adopted both approaches as a further self-consistency check. While the final outcome does not change, the particular form of the intermediate steps does. In this regard, we find the second-order basis more convenient for enumerating the spin-parity states in kinematically restricted version of the MAG, due to the index symmetries in Eq. (4).

A. Zero nonmetricity

The imposition of zero nonmetricity is easily realized in a formalism with explicit distortion. From Eq. (4) it is clear that a two-index-antisymmetric rank-three field $\Delta_{\lambda\mu\nu} \equiv$ $-\Delta_{\lambda\nu\mu}$ nullifies $Q_{\lambda\mu\nu} \equiv 0$. This achieves a reduction of the spin-parity sectors of Eq. (5a)

$$\Delta_{\mu\nu\rho} \supset \{2_3^+, 2_2^-, 1_2^+, 1_3^+, 1_3^-, 1_6^-, 0_3^+, 0_1^-\},$$
(13)

reflected by a further redundancy in the number of independent Ricci-type contractions

$$\mathcal{F}_{\mu\nu} \equiv 0, \qquad \mathcal{F}^{(14)}{}_{\mu\nu} \equiv -\mathcal{F}^{(13)}{}_{\mu\nu}.$$
 (14)

The action Eq. (12) is therefore simplified into

$$S[g,A] = -\frac{1}{2} \int d^4x \sqrt{-g} \Big[-a_0 \mathcal{F} + \mathcal{F}^{(13)\mu\nu} \Big(g_7 \mathcal{F}^{(13)}{}_{\mu\nu} + g_8 \mathcal{F}^{(13)}{}_{\nu\mu} \Big) \Big].$$
(15)

Note that we follow [28] in relabeling the dimensionless coefficients from c_i to q_i when passing from Eqs. (12)–(15) by kinematic restriction. The spectrum of this simple threeparameter model can be promptly, and unambiguously profiled within our formalism. Just within this section (but not within Secs. III B and III C) the accompanying PSALTer analysis in Appendix A will be made in the Poincaré gauge theory (PGT) formulation of zerononmetricity MAG. The kinematic structure of PGT is more extensive than its MAG counterpart due to extra antisymmetric parts of the tetrad fields (which are nullified by an extra Lorentz gauge symmetry). This kinematic structure is presented in Fig. 2, but is otherwise analogous to that in Eqs. (5b) and (13). For now, we proceed with our discussion as if we were working in the MAG formulation. To access the singular structure of the propagator we first identify the spin-parity sectors of the kinetic term. The PGT matrices are shown in Appendix A, in Fig. 3. We can immediately recognize in the absence of the 1_7^- and 0_6^+ sectors the hallmark of diffeomorphism invariance. All the information concerning the quadratic terms is encoded in such matrices and, from the arguments of Sec. II, a direct link exists between the zeroes of their determinants and the singularities of $\tilde{D}_{S}(k)$. The shape of the residue and the position of the singularity will determine the nature of the propagating particles. Once the degeneracies of $a_{i,j}^{\{1,-\}}$ and $a_{i,i}^{\{0,+\}}$ are removed, we immediately find that no massive poles are present. We can therefore extract the known result [35]:

No massive states propagate in linearized zerononmetricity Ricci-type theories.

The massless poles are present in the 2⁺ and 0⁺ sectors. Again, these are known traits of graviton propagation. To confirm that the graviton is present we explore the form of these massless poles in the final, gauge invariant, propagator. The constraints to be imposed are read off the null vectors of $a_{i,j}^{\{1,-\}}$ and $a_{i,j}^{\{0,+\}}$. By choosing the lightlike frame $k^{\mu} = (\mathcal{E}, 0, 0, \mathcal{E})$ the sources are constrained as

$$T^{00} = T^{03}, \quad T^{13} = T^{01}, \quad T^{23} = T^{02}, \quad T^{33} = T^{03}.$$
 (16)

³In that work, a more general MAG action is considered, in which a total of 28 invariant operators are present in the Lagrangian density.

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By substituting them into the propagator, we find the following structure:

$$\lim_{k^2 \to 0} \mathcal{D}_S(k) = \frac{1}{k^2} \left(\frac{2}{-a_0}\right) \sum_{i=1,2} |S_i|^2,$$
(17)

where the S_i are linearly independent combinations of the sources. Thus, Eq. (17) reveals the two (helicity) states with a residue $\propto -1/a_0$. We conclude by refining the previous statement:

Linearized zero-nonmetricity Ricci-type theories propagate only a healthy massless graviton with $a_0 < 0$.

B. Zero torsion

The zero-torsion case offers a stronger challenge to our methods, displaying a larger parameter space and nontrivial interplay among different operators. We will show that the spin-parity formalism, followed by direct access to the propagator, gives full access to the tree-level spectrum without imposing simplifying restrictions. First, in terms of the distortion tensor, the zero-torsion condition in Eq. (4) is achieved by working with a two-index-symmetric rank-three field $\Delta_{\lambda\mu\nu} \equiv \Delta_{\nu\mu\lambda}$. The available particle content in the symmetric distortion is then

$$\Delta_{\mu\nu\rho} \supset \left\{ 3_{1}^{-}, 2_{1}^{+}, 2_{2}^{+}, 2_{1}^{-}, 1_{1}^{+}, 1_{1}^{-}, 1_{2}^{-}, \\ 1_{4}^{-}, 1_{5}^{-}, 0_{1}^{+}, 0_{2}^{+}, 0_{4}^{+} \right\},$$
(18)

which causes $\mathcal{F}_{\mu\nu} \equiv -2\mathcal{F}^{(13)}{}_{[\mu\nu]}$, leaving seven independent combinations

$$S[g, A] = -\frac{1}{2} \int d^4 x \sqrt{-g} \Big[-a_0 \mathcal{F} \\ + \mathcal{F}^{(13)\mu\nu} \Big(h_7 \mathcal{F}^{(13)}{}_{\mu\nu} + h_8 \mathcal{F}^{(13)}{}_{\nu\mu} \Big) \\ + \mathcal{F}^{(14)\mu\nu} \Big(h_9 \mathcal{F}^{(14)}{}_{\mu\nu} + h_{10} \mathcal{F}^{(14)}{}_{\nu\mu} \Big) \\ + \mathcal{F}^{(14)\mu\nu} \Big(h_{11} \mathcal{F}^{(13)}{}_{\mu\nu} + h_{12} \mathcal{F}^{(13)}{}_{\nu\mu} \Big) \Big].$$
(19)

Note once again that we follow [28] in relabeling the dimensionless coefficients from c_i to h_i when passing from Eqs. (12)–(19) by kinematic restriction. The kinematic structures in Eqs. (5b) and (18) are reflected in the *PSALTer* definitions of the states, which are fully defined in Fig. 4. The computation of the spin-parity matrices produces the remaining output in Appendix B, with the general case Eq. (19) displayed in Fig. 5. After simple inspections of the determinants, we find that it is impossible to impose further gauge symmetries, unless a_0 , the only dimensionful parameter, is set to zero. This would nullify the graviton propagation and, therefore, we discard this option.

Taking into account the constraints of diffeomorphism invariance as illustrated in Sec. III A, we can compute the positions of the zeroes for all the determinants. Already for the 2^+ sector we find that a state of mass

$$m_2^{2+} = -\frac{a_0}{h_{10} + h_9} \tag{20}$$

is allowed to propagate whenever $h_{10} + h_9 \neq 0$. The computation of the limit $k^2 \mapsto m_{2^+}^2$ gives rise to the following pole residue, where we suppress the positive-definite quadratic form in the 2^+ sources

$$\lim_{k^{2}\mapsto m_{2^{+}}^{2}} \mathcal{D}_{S}(k) \propto 2 \left[4h_{10}^{2} + h_{11}^{2} + h_{12}^{2} + 2h_{11}(h_{12} - 2h_{9}) - 4h_{12}h_{9} + 4h_{9}^{2} - 2h_{9}a_{0} - 2h_{10}(2h_{11} + 2h_{12} - 4h_{9} + a_{0}) \right] \times \left[(h_{10} + h_{9})^{2}a_{0} \right]^{-1}.$$
(21)

The positivity of the spin-two mass in Eq. (20) and the residue in Eq. (21) select possible real values of the parameters involved. Among these, the requirement is seen for a positive a_0 . This shows that a massive spin-two is incompatible with the healthy propagation of the graviton: we must discard it. This, as can be seen from the determinants, can be accomplished by demanding $h_{10} = -h_9$ or the stronger $h_{10} = h_9 = 0$. The consequences of these different choices can be appreciated by observing that the theory also propagates a massive spin-one particle whose mass and residue is given by

$$m_{1^+}^2 = \frac{a_0}{h_9 - h_{10}}, \qquad \lim_{k^2 \mapsto m_{1^+}^2} \mathcal{D}_S(k) \propto \frac{4}{h_{10} - h_9}.$$
 (22)

We can therefore explore the possibility of keeping such a state by adopting $h_{10} = -h_9 \neq 0$ and studying the consequences for the rest of the spectrum. It is easy to show that this requires $h_{10} = -h_9 > 0$ and $a_0 < 0$, so that such propagation can indeed be afforded without spoiling the gravitational priorities of the model.

A more alarming scenario is presented by the degenerate spin 1⁻ sector, where the restricted determinant shows a quartic equation in the momentum. Imposing, for instance, $h_{10} = -h_9$ we would find

$$|a_{i,j}^{\{1,-\}}| = \frac{a_0^2}{32} \Big(5 \Big((h_{12} - h_{11})^2 + 8h_9(h_8 - h_7) \Big) k^4 - 4a_0 (3h_{11} - 3h_{12} - h_7 + h_8 + 12h_9) k^2 + 12a_0^2 \Big).$$
(23)

The determinant in Eq. (23) can be read off the denominators of the $b_{i,i}^{\{1,-\}}$ matrix elements.⁴ Bona fide unitarity demands the removal of the quartic term. The absence of a ghostly massive spin-two particle and the spin-one dipole motivates the defining constraint over the parameter space. Many (apparently) different solutions can be found by asking to solve such constraints in terms of different subsets of the couplings. It is a great advantage that the algorithmic disposition of the spin-parity approach allows a simple scan over the broad space of solutions. We proceed by considering the two separate branches obtained from $h_{10} =$ $-h_9 > 0$ and, for each of the two possibilities, gather the different solutions yielded by nullifying the quadratic coefficient of k^4 in Eq. (23). The theories associated with this scan are presented in [87]. Collecting all the masses and residues for, besides the graviton, the two massive spinone states of opposite parity we find, in all cases that

Linearized zero-torsion Ricci-type theories do not admit simultaneous propagation of massive spin-one states of opposite parity.

We can naturally continue by asking for the dismissal of the full 1⁻ propagation by setting to zero the coefficient of k^2 in Eq. (23). Again, we do this by gathering all the relevant equations and solving them for all the possible subsets of the free parameters, and again we refer to [87]. We find that healthy solutions are available, although not for all the given cases. In the healthy scenarios, we rediscover the mass/residue ratio of Eq. (22) rephrased in terms of the available couplings.

Finally, we can investigate the case $h_9 = h_{10} = 0$ which removes the propagation of the massive spin-one state of positive parity Eq. (22) and, simultaneously, of the massive spin-two. Under such circumstances, the cancellation of the dipole propagation simplifies to demanding $(h_{11} - h_{12})^2 =$ 0 in Eq. (23), namely $h_{11} = h_{12} \neq 0$ (see Figs. 9 and 10) and $h_{11} = h_{12} = 0$ (see Figs. 7 and 8). For the second, simpler scenario, we find in the second-order case Fig. 8

$$m_{1^-}^2 = \frac{-3a_0}{h_7 - h_8}, \qquad \lim_{k^2 \mapsto m_{1^-}^2} \mathcal{D}_S(k) \propto \frac{34}{h_7 - h_8}.$$
 (24)

Such a simple setup, which clearly presents a viable propagation, is only slightly modified when considering the branch $h_{11} = h_{12} \neq 0$, with only the residue's form being affected. Again, ghost- and tachyon-freedom can be accounted for. We conclude, accordingly, stating that, besides the graviton,

Either a healthy massive vector of negative parity or positive parity propagates in linearized zero-torsion Riccitype theories.

C. Generic case

1. General properties

The transition to the case of an unconstrained affine connection presents an obvious growth in computational complexity induced by the multiple components of Eqs. (5a) and (5b) and, consequently, by the independence of all three Ricci-type tensors. The challenges of the associated spectral problem are quite visible in the cumbersome spin-parity matrices of Appendix C. The general spectrum associated with Eq. (12) is shown in Figs. 12 and 13, respectively, for the first- and second-order formulations of the theory. The inclusion of all the components considerably changes the nature of the unconstrained spectrum. First, we notice how the massive state of spin-two is no longer present, the determinant having a simple proportionality to k^2 . Similarly, Eq. (23) is now of the form⁵

$$|a_{i,j}^{\{1,-\}}| = \frac{1}{4}a_0^4 k^2 (f_0 a_0 - 5f_1 k^2),$$
(25a)

$$f_0 \equiv (c_{10} - c_{11} + c_{12} - 16c_{13} + 4c_{14} + 4c_{15} - c_7 + c_8 - c_9),$$
(25b)

$$f_1 \equiv (c_{14} + c_{15})^2 + 4c_{13}(c_{10} - c_{11} + c_{12} - c_7 + c_8 - c_9),$$
(25c)

and the dangerous dipole of Eq. (23) leaves space for a massless vector. That this is indeed the case, and that the pole is not a spurious feature of the determinant, is demonstrated by the direct computation of the saturated propagator in the massless limit. For this computation we have to account for a further, associated peculiarity encountered in this scenario. The rank of the spin 0^+ sector is now reduced by 2, signaling the emergence of an Abelian symmetry. The presence of this symmetry was predicted by Iosifidis and Koivisto [88]-it appears whenever squares of the full metric-affine curvature are added to the Einstein-Hilbert term, and is a remnant of the full projective symmetry of that term. It is instructive to explicitly show what this entails in terms of source constraints in the lightlike frame $k = (\mathcal{E}, 0, 0, \mathcal{E})$. We find, together with Eq. (16), the following reduction:

$$W^{000} - W^{011} - W^{022} - W^{033}$$

= W^{300} - W^{311} - W^{322} - W^{333}. (26)

⁴Unfortunately, the $b_{i,j}^{\{1,-\}}$ matrices are very large expressions, so *PSALTer* frequently suppresses them when attempting to typeset the results for publication. Although Eq. (23) cannot therefore be confirmed from Appendix B, the full results are available in the Wolfram *Mathematica* notebook file from which *PSALTer* is run: this document, along with the source script, is made available in the Supplemental Material [87].

⁵Again, the $b_{i,j}^{\{1,-\}}$ matrices are suppressed in Appendix C, for full results see the Supplemental Material [87].

Accounting for Eq. (26) we recognize four independent states in the massless limit of the saturated propagator. Two of these are precisely the helicity states of the graviton, proportional to $-1/a_0$ and recognizable in Eq. (17). The residue of the other two states, while signaling unambiguously the propagation connected to the massless spin-one state, has an extremely convoluted form due to the concurrence of many different parameters in its definition. Nevertheless, the requirement for its positivity has the manageable structure

$$16c_{13} - 4(c_{14} + c_{15}) + c_{11} - c_{12} + c_7 - c_8 + c_9 - c_{10} > 0.$$
(27)

When searching for massive propagation, similarly to the torsionless case, both spin-one sectors source one state each, with masses

$$m_{1^{-}}^{2} = \left[a_{0}(c_{10} - c_{11} + c_{12} - 16c_{13} + 4c_{14} + 4c_{15} - c_{7} + c_{8} - c_{9})\right] \times \left[5(4c_{13}(c_{10} - c_{11} + c_{12} - c_{7} + c_{8} - c_{9}) + (c_{14} + c_{15})^{2})\right]^{-1},$$
(28a)

$$m_{1^+}^2 = -\frac{a_0}{c_{10} - c_{11} + c_{12} - c_7 + c_8 - c_9}.$$
 (28b)

The overall survey of the propagating states points, therefore, to three additional particles populating the spectrum besides the graviton. To solve the spectral problem we analyze the conditions for their simultaneous propagation.

2. Allowing the massless vector

The presence of a spin-one massless state can be included in our analysis. The related phenomenological concerns can then be seen as suggesting incompleteness, rather than an inconsistency: a mechanism to provide a mass gap is expected. Under such a hypothesis, we can investigate the coexistence of such a state with the others. The computation of the residues of the massive states is carried through the constrained propagator. The explicit effect of the different gauge symmetries on the sources is extracted in the rest-mass frame k = (m, 0, 0, 0) and gives

$$T^{00} = T^{01} = T^{02} = T^{03} = 0, (29)$$

for diffeomorphism invariance, and

$$W^{000} - W^{011} - W^{022} - W^{033} = 0, (30)$$

for the extra Abelian symmetry. When testing the sign of each residue, as well as the masses, with the requirement $a_0 < 0$, we immediately find an obstruction within the 1⁻ sector. Having committed to retaining the massless propagation, we must simplify the model by removing its massive counterpart. We proceed, therefore, by considering all the

11 solutions of $f_1 = 0$ in Eq. (25a) and recomputing the residues for the remaining propagating states. Once more, no viable solutions are found (see [87]). Finally, we kill the massive 1⁺ propagation in Eq. (28b) by adding the further condition

$$c_{10} - c_{11} + c_{12} - c_7 + c_8 - c_9 = 0.$$
(31)

To coherently include both constraints we consider pairs of parameters which are solutions of the corresponding equation system. Twelve solutions are found (see [87]), all of them with positive residues for the surviving massless sector. We can therefore draw the following:

A healthy massless vector of negative parity propagates in linearized generic Ricci-type theories.

Once again, we can make contact with the literature. We notice that Eq. (31) does not eliminate c_{13} , which controls the square of the homothetic curvature. It is known that when this operator is added to the Einstein-Hilbert term in full MAG geometry, the resulting theory cannot be distinguished from the vacuum Einstein-Maxwell theory [89,90]—the extra massless vector in this case is identified with our 1⁻ state.

3. Removing the massless vector

The only way to dispose of the massless vector is to introduce a further degeneracy in the 1⁻ sector. This can be enforced by solving for $f_1 = f_2 = 0$ in Eq. (25a). Once more, the spin-parity approach grants us the possibility to explore the results in a systematic way. The analysis is made more complicated by the peculiar challenges met in this scenario, where each solution of the $f_1 = f_2 = 0$ system affects the form of the gauge symmetry, thus necessitating, each time, a recomputation of all the main features of the theory. Despite the demanding computational task (see [87]), the outcome turned out to be the same for all the (20) different solutions defined in terms of pairs of independent parameters. We can, consequently, present the results for this scenario by focusing on a particular solution:

$$c_9 = c_{10} - c_{11} + c_{12} + 16c_{13} - c_7 + c_8, \quad (32a)$$

$$c_{15} = 8c_{13} - c_{14},\tag{32b}$$

producing the following propagator poles

$$m_{1^-}^2 = \frac{a_0}{20c_{13}}, \qquad m_{1^+}^2 = \frac{a_0}{16c_{13}}.$$
 (33)

The correlation among the masses of the two spin-one sectors is not an accident of the chosen solution but illustrates a common feature: the strict proportionality $m_{1^-}^2/m_{1^+}^2 = 4/5$. Such correlation signals the impossibility of removing the propagation of one state without interfering

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with the other. To assess the nature of the states we need to compute the saturated propagator in the presence of the augmented gauge constraints defined, for instance, by Eqs. (32a) and (32b). For the massive limit of k = (m, 0, 0, 0) for some *m* we must consider, on top of Eqs. (29) and (30) the extra degeneracy of the 1⁻ sector. Being generated by a three-dimensional vector, these source constraints take the form

$$W^{00i} - W^{0i0} = \frac{a_0}{m^2} \frac{W^{i00} - W^{i11} - W^{i22} - W^{i33}}{c_{11} - c_{12} - 4c_{14} + 2c_7 - 2c_8}.$$
 (34)

The odd appearance of couplings within the definitions of the symmetry is a curious feature of this analysis. This, however, does not present any theoretical downsides if we consider that such parameters will be normalized to pure numbers in the quadratic, final Lagrangian. The massive limit confirms the presence of the propagation of three states as expected for massive spin-one particles. For the two different sectors, and the representative selection of dependent parameters shown in Eqs. (32a) and (32b), we find the corresponding residue

$$\lim_{k^2 \mapsto m_{1^-}^2} \mathcal{D}_{S}(k) \propto \frac{1}{50c_{13}} \left[11 + \left(50(6c_{13} - c_{14})(c_{12} - c_{11} + 2(6c_{13} + c_{14} - c_7 + c_8)) \right) \times \left((c_{12} - c_{11} + 4c_{14} - 2c_7 + 2c_8)^2 \right)^{-1} \right], \quad (35a)$$

$$\lim_{k^2 \mapsto m_{1^+}^2} \mathcal{D}_S(k) \propto -\frac{1}{4096c_{13}^3} \left[c_{11}^2 - 2c_{11}(c_{12} + 2(8c_{13} - c_7 + c_8)) + c_{12}^2 + 4c_{12}(8c_{13} - c_7 + c_8) + 768c_{13}^2 + 64c_{13}(c_8 - c_7) + 4(c_7 - c_8)^2 \right].$$
(35b)

While the requirement of positivity of $\lim_{k^2 \mapsto m_{1^-}^2} \mathcal{D}_S(k)$ does not challenge the unitarity of the graviton sector, nor the tachyon-freedom conditions over the vector masses, this is not the case for $\lim_{k^2 \mapsto m_{1^+}^2} \mathcal{D}_S(k) > 0$, which calls for $c_{13} < 0$. Again, different choices in solving for $f_1 = f_2 = 0$ do affect the form of the residues, but no simultaneous solutions are found (see [87]). The correlation existing for the propagation of both massive vectors in Eq. (33) prohibits, therefore, both massive states from appearing in a healthy spectrum.

IV. CONCLUSIONS

It is difficult to overestimate the importance of accommodating the absence of ghosts and tachyons in quantum field theory. Control over unitarity is key to understanding the shape of possible new theories and future extensions of the current models. Spin-parity projectors provide a computational framework for fully controlling the propagation of quadratic Lagrangian, which lends itself well to computer implementation. Once the needed operators are collected, the spectral problem is basically solved [55,56]. The output is unambiguous, given the direct access to the propagator, and does not rely upon intricate field redefinitions or the introduction of spurious fields to achieve reductions to known cases. In this work, we have adopted the spin-parity formalism to illustrate its reach and the capacity to tackle a broader set of operators than previously possible. We have made a thorough survey of the Ricci-type MAG operator space, but our analysis is not intended to be exhaustive. The point we are making is that if further special cases turn out to be of interest in the future, then it will be economical to test them using our approach. Recently, some spectral analyses of the PGT and Weyl gauge theory have been made [55,56,72], which really *are* exhaustive. The trick to making exhaustive surveys is to recursively search over the root system of the wave operator determinant. This would make an appealing (and apparently straightforward) extension to our current PSALTer program, but we defer it to future work.

There are two key limitations to our approach. First, the authors of [35] are able to extend their analyses to particle spectra on Friedmann backgrounds: we cannot do this. There is some hope for the extension of the spectral algorithm to de Sitter backgrounds in the near future [91], but further applications to nonmaximally symmetric spacetimes are currently speculative [92]. Second, the theories in Eqs. (12), (15) and (19) may propagate more species in their full nonlinear dynamics than are revealed in the spectral analysis. This is already known to happen in the case of the theory in Eq. (19), for which the 1^+ and $1^$ torsion modes are strongly coupled near Minkowski spacetime [57]. When this happens, it means that the model is inherently nonperturbative around Minkowski spacetime, so the quadratic approximation in Eq. (6) is just a fictional model which has nothing to do with the actual physics. It is hard to see how this cannot be a pathology (with or without ghost-tachyon freedom of the strongly coupled modes), and the only sure way to diagnose it is via a nonlinear Hamiltonian analysis [93]. It is possible that the methods of [35] are also sensitive to strong coupling, if for example propagating d.o.f. are lost as the Friedmann background is deformed into the Minkowski background. However, it is not clear that such an approach would always detect the problem when it exists. Attempts at computer algebra Hamiltonian analysis were made in [94], but the implementation was not theory agnostic (restricted to PGT). The PSALTer software showcased here is theory agnostic by design.⁶

 $^{^{6}}$ It can be downloaded from github.com/wevbarker/PSALTer. In the longer term we hope that *PSALTer* will become an official contribution to the *xAct* project [86,95–98].

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APPENDIX A: ZERO NONMETRICITY WITH *PSALTER*

The full spectrum of the general theory in Eq. (15) is given in Fig. 3. Whilst the formulation in Sec. III A is consistent with the MAG conventions set out in Sec. II A, we take the unusual approach of reformulating the theory as a PGT [99] when presenting the *PSALTer* analysis. This allows us to recycle a preexisting PGT kinematic module within *PSALTer*, and meanwhile the spectral analysis of Fig. 3 is already well-known anyway [57]. The PGT kinematic module is displayed in Fig. 2.

Whilst our MAG conventions are designed to be identical to [28], our PGT conventions will be identical to those in [65,93,100–103]. To define the PGT, we introduce e^i_{μ} and e_i^{μ} as the cotetrad and tetrad components, which are associated with Roman Lorentz (i.e. anholonomic) indices, so that we can compare with the MAG metric $e^i_{\mu}e^j_{\nu}\eta_{ij} \cong$ $g_{\mu\nu}$ and inverse $e_i^{\mu}e_j^{\nu}\eta^{ij} \cong g^{\mu\nu}$ with identities $e^i_{\mu}e_i^{\nu} \equiv \delta^{\nu}_{\mu}$ and $e^i_{\mu}e_j^{\mu} \equiv \delta^i_j$ as kinematic restrictions. There is also a

PSA	LTer kinematic pa	nel	
Mome	ntum Norm Frame		
ν ^μ	$\mu^2 - \mu - \mu^{\mu} n^{\mu} - \frac{\mu^{\mu}}{2}$		
^	$ \kappa - \kappa_{\mu} \kappa n - \kappa_{k}$		
Fun	damental fields		
Fields	Symmetries	SO(3)	Sources
A ^{αβ} χ	StrongGenSet[{1,2},GenSet[-(1,2)]]	$\frac{4}{3} \frac{2}{3} \mathcal{A}^{\alpha\beta}_{\mu} + \frac{1}{2} \frac{1}{2} \mathcal{A}^{\beta}_{\mu} \delta^{\beta}_{\alpha} + \frac{1}{2} \frac{1}{3} \mathcal{A}^{\alpha}_{\mu} \delta^{\beta}_{\alpha} + \frac{1}{2} \mathcal{A}^{\beta}_{\alpha} n^{\alpha}_{\alpha} + \frac{1}{3} \mathcal{A}^{\beta}_{\alpha} \delta^{\beta}_{\alpha} n^{\alpha}_{\alpha} + \frac{1}{3} \mathcal{A}^{\beta}_{\alpha} n^{\alpha}_{\alpha} n^{\beta}_{\alpha} + \frac{1}{3} \mathcal{A}^{\beta}_{\alpha} n^{\beta}_{\alpha} n^{$	$\sigma_{ab\chi}$
		$2^{*}\mathcal{A}_{\mu}^{\mu}{}_{\chi}^{\mu}{}^{\mu}{}_{-\frac{1}{2}}^{\mu}\mathcal{A}_{\mu}^{\mu}{}$	
$f_{\alpha\beta}$	StrongGenSet[{},GenSet[]]	${}^{1^{*}}f^{l}{}_{\alpha\beta} + {}^{2^{*}}f^{l}{}_{\alpha\beta} + {}^{\frac{1}{3}} {}^{0^{*}}f^{l}{}_{\beta} \eta_{\alpha\beta} + {}^{1}f^{l}{}_{\beta} \eta_{\alpha} + {}^{1^{*}}f^{\mu}{}_{\alpha} \eta_{\beta} - {}^{\frac{1}{3}} {}^{0^{*}}f^{l}{}_{\beta} \eta_{\alpha} \eta_{\beta} + {}^{0^{*}}f^{\mu}{}_{\alpha} \eta_{\beta}$	$\tau (\Delta + \mathcal{H})_{\alpha\beta}$
600	2) innene		
50(3) irreps		
SO(3)	\$ymmetries	Expansion	Sources
⁰⁺ <i>A</i> l	StrongGenSet[{},GenSet[]]	$\mathcal{F}_{\alpha\beta}^{\ \ \beta} n^{lpha}$	0 ⁺ <i>σ</i>
${}^{0}\mathcal{R}^{I}$	StrongGenSet[{}, GenSet[]] -	$\mathcal{R}^{\partial \ell\delta} \epsilon \eta_{ab \ell\delta} n^a$	0.°01
1+ A al	StrongGenSet[{1, 2}, GenSet[-(1,2)]]	$\frac{1}{2} \mathcal{R}_{\alpha\chi\beta} \ n^{\chi} + \frac{1}{2} \mathcal{R}_{\beta\chi\alpha} \ n^{\chi} - \frac{1}{2} \mathcal{R}_{\beta\chi\delta} \ n_{\alpha} \ n^{\chi} \ n^{\delta} + \frac{1}{2} \mathcal{R}_{\alpha\chi\delta} \ n_{\beta} \ n^{\chi} \ n^{\delta}$	$1^+ \sigma^{\parallel}_{\alpha\beta}$
$^{1^+}\mathcal{H}^{\perp}_{al}$	StrongGenSet[{1,2},GenSet[-(1,2)]]	$\mathcal{R}_{\alpha\beta\chi} \ n^{\chi} + \mathcal{R}_{\beta\chi\delta} \ n_{\alpha} \ n^{\chi} \ n^{\delta} - \mathcal{R}_{\alpha\chi\delta} \ n_{\beta} \ n^{\chi} \ n^{\delta}$	$1^+ \sigma^+ a\beta$
${}^{1}\mathcal{R}^{l}{}_{\alpha}$	StrongGenSet[{}, GenSet[]] -	$\mathcal{R}_{\alpha\beta}^{\ \beta} + \mathcal{R}_{\beta\chi}^{\ \chi} \ n_{\alpha} \ n^{\beta} + \mathcal{R}_{\alpha\beta\chi} \ n^{\beta} \ n^{\chi}$	¹ .σ ¹ .α
${}^{1}\mathcal{A}^{\perp}{}_{\alpha}$	StrongGenSet[{},GenSet[]]	$\mathcal{F}_{\alpha\beta\chi} n^{\beta} n^{\chi}$	$^{1}\sigma^{+}\alpha$
2+ H ¹ al	StrongGenSet[{1,2}, GenSet[(1,2)]]	$\frac{1}{2} \mathcal{R}_{\alpha\chi\beta} \ n^{\chi} - \frac{1}{2} \mathcal{R}_{\beta\chi\alpha} \ n^{\chi} - \frac{1}{3} \mathcal{R}_{\chi}^{\ \delta} \ \eta_{\alpha\beta} \ n^{\chi} + \frac{1}{3} \mathcal{R}_{\chi\delta}^{\ \delta} \ \eta_{\alpha} \ n_{\beta} \ n^{\chi} + \frac{1}{2} \mathcal{R}_{\beta\chi\delta} \ n_{\alpha} \ n^{\chi} \ n^{\delta} + \frac{1}{2} \mathcal{R}_{\alpha\chi\delta} \ n_{\alpha} \ n^{\chi} \ n^{\delta} + \frac{1}{2} \mathcal{R}_{\alpha\chi\delta} \ n^{\delta} \ n^{\chi} \ n^{\delta} \ n^{\chi} \ n^{\delta} + \frac{1}{2} \mathcal{R}_{\alpha\chi\delta} \ n^{\chi} \ n^{\delta} + \frac{1}{2} \mathcal{R}_{\alpha\chi\delta} \ n^{\delta} \ n^{\chi} \ n^{\delta} \ n^{\chi} \ n^{\delta} \ n^{\chi} \ n^{\delta} \ n^{\chi} \ n^{\delta} \ n^{\delta} \ n^{\chi} \ n^{\chi} \ n^{\chi} \ n^{\delta} \ n^{\chi} $	²⁺ σ aβ
$^{Z}\mathcal{A}^{I}_{\alpha\beta}$	StrongGenSet[{1,2},GenSet[-(1,2)]]	$\frac{1}{2} \mathcal{R}_{\alpha\beta\chi} + \frac{1}{4} \mathcal{R}_{\alpha\chi\beta} - \frac{1}{4} \mathcal{R}_{\beta\chi\alpha} + \frac{3}{8} \mathcal{R}_{\beta\delta}^{\delta} \eta_{\alpha\chi} - \frac{3}{8} \mathcal{R}_{\alpha\delta}^{\delta} \eta_{\beta\chi} - \frac{3}{8} \mathcal{R}_{\beta\delta}^{\delta} n_{\alpha} n_{\chi} + \frac{3}{8} \mathcal{R}_{\alpha\delta}^{\delta} n_{\beta} n_{\chi} + \frac{1}{4} \mathcal{R}_{\beta\chi\delta} n_{\alpha} n^{\delta} + \frac{1}{2} \mathcal{R}_{\delta\chi\delta} n_{\alpha} n^{\delta} n^{\delta} n_{\alpha} n^{\delta} n^{\delta} n_{\alpha} n^{\delta} n^{\delta} n_{\alpha} n^{\delta} n^{\delta$	² σ αβχ
		$\frac{1}{2} \mathcal{R}_{\beta\delta\chi} \ n_{\alpha} \ n^{\delta} + \frac{1}{4} \mathcal{R}_{\chi\delta\beta} \ n_{\alpha} \ n^{\delta} + \frac{3}{8} \mathcal{R}_{\delta\epsilon}^{\ \epsilon} \ n_{\beta\chi} \ n_{\alpha} \ n^{\delta} - \frac{1}{4} \mathcal{R}_{\alpha\chi\delta} \ n_{\beta} \ n^{\delta} - \frac{1}{2} \mathcal{R}_{\alpha\delta\chi} \ n_{\beta} \ n^{\delta} - \frac{1}{4} \mathcal{R}_{\chi\delta\alpha} \ n_{\beta} \ n^{\delta} - \frac{3}{8} \mathcal{R}_{\delta\epsilon}^{\ \epsilon} \ \eta_{\alpha\chi} \ n_{\beta} \ n^{\delta} - \frac{1}{4} \mathcal{R}_{\alpha\delta\chi} \ n_{\beta} \ n^{\delta} - \frac{1}{4} \mathcal{R}_{\alpha\delta\chi} \ n_{\beta} \ n^{\delta} - \frac{1}{4} \mathcal{R}_{\alpha\delta\chi} \ n^{\delta} + \frac$	
		$\frac{1}{2} \mathcal{R}_{\alpha\beta\delta} \ n_{\chi} \ n^{\delta} - \frac{1}{4} \mathcal{R}_{\alpha\delta\beta} \ n_{\chi} \ n^{\delta} + \frac{1}{4} \mathcal{R}_{\beta\delta\alpha} \ n_{\chi} \ n^{\delta} - \frac{3}{8} \mathcal{R}_{\beta\delta\epsilon} \ \eta_{\alpha\chi} \ n^{\delta} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\delta\epsilon} \ \eta_{\beta\chi} \ n^{\delta} \ n^{\epsilon} - \frac{3}{8} \mathcal{R}_{\beta\delta\epsilon} \ n_{\alpha} \ n_{\chi} \ n^{\delta} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\delta\epsilon} \ n_{\beta\chi} \ n^{\delta} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\delta\epsilon} \ n_{\beta\chi} \ n^{\delta} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\delta\epsilon} \ n_{\beta\chi} \ n^{\delta} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\delta\epsilon} \ n_{\beta\chi} \ n^{\delta} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\delta\epsilon} \ n_{\beta\chi} \ n^{\delta} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\delta\epsilon} \ n_{\beta\chi} \ n^{\delta} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\delta\epsilon} \ n_{\beta\chi} \ n^{\delta} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\delta\epsilon} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\epsilon} \ n^{\epsilon} + \frac{3}{8} \mathcal{R}_{\alpha\epsilon}$	
0*f1	StrongGenSet[{},GenSet[]]	$f^{\alpha}_{\ \alpha} - f^{\alpha\beta} \ n_{\alpha} \ n_{\beta}$	0 ⁺ 1
0^+f^\perp	StrongGenSet[{},GenSet[]]	$f^{a\beta} n_a n_{\beta}$	0+ 7 [±]
$1^{+}f^{\dagger}_{\alpha\beta}$	StrongGenSet[{1, 2}, GenSet[-(1,2)]]	$\frac{\ell_{\alpha\beta}}{2} - \frac{\ell_{\beta\alpha}}{2} + \frac{1}{2} f_{\beta}^{\chi} n_{\alpha} n_{\chi} - \frac{1}{2} f_{\beta}^{\chi} n_{\alpha} n_{\chi} - \frac{1}{2} f_{\alpha}^{\chi} n_{\beta} n_{\chi} + \frac{1}{2} f_{\alpha}^{\chi} n_{\beta} n_{\chi}$	1 ⁺ τ [†] αβ
$1_{f_{\alpha}}$	StrongGenSet[{},GenSet[]]	$\int \beta_{\alpha}^{\beta} n_{\beta} - \int \beta_{\chi}^{\beta\chi} n_{\alpha} n_{\beta} n_{\chi}$	1 T ^I a
$1_{f_{\alpha}^{\perp}}$	StrongGenSet[{},GenSet[]]	$f_{\alpha}^{\ \beta} n_{\beta} f^{\beta \chi} n_{\alpha} n_{\beta} n_{\chi}$	1 t ¹ a
2 ⁺ f ¹ αβ	StrongGenSet[{1, 2}, GenSet[(1,2)]]	$\frac{t_{\alpha\beta}}{t_{\alpha\beta}} + \frac{t_{\beta\alpha}}{t_{\alpha}} + \frac{1}{3}f_{\chi}^{\chi} \eta_{\alpha\beta} + \frac{1}{3}f_{\chi}^{\chi} \eta_{\alpha} \eta_{\beta} - \frac{1}{2}f_{\beta}^{\chi} \eta_{\alpha} \eta_{\chi} - \frac{1}{2}f_{\alpha}^{\chi} \eta_{\alpha} \eta_{\chi} - \frac{1}{2}f_{\alpha}^{\chi} \eta_{\beta} \eta_{\chi} - \frac{1}{2}f_{\alpha}^{\chi} \eta_{\beta} \eta_{\chi} + \frac{1}{3}f^{\chi\delta} \eta_{\alpha\beta} \eta_{\chi} \eta_{\delta} + \frac{1}{3}f^{\chi\delta} \eta_{\delta} \eta_{\delta} \eta_{\chi} \eta_{\delta} + \frac{1}{3}f^{\chi\delta} \eta_{\delta} \eta_{\delta} \eta_{\delta} \eta_{\delta} + \frac{1}{3}f^{\chi\delta} \eta_{\delta} \eta_{$	2 ⁺ τ [†] αβ
			•

FIG. 2. Kinematic structure of PGT, as used in Fig. 3. We repurpose the PGT kinematic module in *PSALTer* for the study of zerononmetricity MAG in Sec. III A. Because of kinematic differences between the PGT and MAG (which are nullified by the extra Lorentz symmetry in PGT), the irreps displayed here do not completely map to those in Eq. (5b) and (13). The key point is that the spin connection $\mathscr{A}^{ij}{}_{\mu} \equiv \mathscr{A}^{[ij]}{}_{\mu}$ maps to the antisymmetric distortion $\Delta_{\lambda\mu\nu} \equiv \Delta_{\lambda[\mu\nu]}$, and the asymmetric tetrad perturbation $f_{i}{}^{\mu}$ contains at least the d.o.f. in the symmetric metric perturbation $h_{\mu\nu} \equiv h_{(\mu\nu)}$. Note that the 2⁻ state has a hidden multiterm cyclic symmetry on all its indices, which is not accommodated by the C language implementation of the Butler-Portugal algorithm [86].



FIG. 3. The full spectrum of the general theory in Eq. (15), but interpreted as a PGT in Eq. (A2). Kinematically, the 10 d.o.f. of the metric are replaced with the 16 d.o.f. of the tetrad field. Consequently however, the additional gauging of the Lorentz group results in six extra gauge generators on top of the diffeomorphism (translation) generators, so the formulations are not physically distinguishable. All the quantities in this output are defined in Fig. 2.

spin connection $\mathscr{A}^{ij}_{\ \mu} \equiv \mathscr{A}^{[ij]}_{\ \mu}$, so that the PGT torsion and PGT curvature are

$$\mathcal{T}^{k}{}_{ij} \equiv 2e_{i}{}^{\mu}e_{j}{}^{\nu}(\partial_{[\mu]}e^{k}{}_{|\nu]} + \mathscr{A}^{k}{}_{m[\mu]}e^{m}{}_{|\nu]}), \qquad (A1a)$$

$$\mathcal{R}^{kl}{}_{ij} \equiv 2e_i{}^{\mu}e_j{}^{\nu}(\partial_{[\mu]}\mathscr{A}^{kl}{}_{|\nu]} + \mathscr{A}^k{}_{m[\mu]}\mathscr{A}^{ml}{}_{|\nu]}). \tag{A1b}$$

The MAG torsion and curvature in Eqs. (1) and (2) are precisely analogous to the PGT counterparts in Eqs. (A1a) and (A1b), respectively, through the relations $\mathcal{T}_{\mu \nu}^{\ \alpha} \cong e^{i}_{\mu}e^{k}_{\ \nu}e^{j}_{\nu}\mathcal{T}^{k}_{\ ij}$ and $\mathcal{F}_{\mu\nu}^{\ \rho}_{\ \sigma} \cong e^{i}_{\ \mu}e^{j}_{\ \nu}e_{k}^{\ \rho}e_{l\sigma}\mathcal{R}^{k}_{\ lij}$, where we pay attention to the different ordering of the indices according to the two conventions. In terms of the PGT field strength tensors, the action in Eq. (15) corresponds to

$$S[e,\mathscr{A}] = \int \mathrm{d}^4 x e[\alpha_0 \mathcal{R} + \mathcal{R}^{ij}(\alpha_1 \mathcal{R}_{ij} + \alpha_2 \mathcal{R}_{ji})], \quad (A2)$$

where $S[e, \mathscr{A}] \cong S[g, A]$ and we adopt the dimensionful coupling α_0 and dimensionless couplings α_1 and α_2 in place of a_0, g_1 and g_2 . The contractions are defined $\mathcal{R}_{ij} \equiv \mathcal{R}^l_{ilj}$ and $\mathcal{R} \equiv \mathcal{R}^i_i$ with the measure $e \equiv \det(e^i_{\mu}) \cong \sqrt{-g}$. In the weak-field regime, we take \mathscr{A}^{ij}_{μ} to be inherently perturbative, and we define the exact tetrad perturbation $e_i^{\mu} \equiv \delta_i^{\mu} + f_i^{\mu}$, i.e. the "Kronecker" choice of Minkowski vacuum [65,104,105].

To lowest order in the quadratic action, the Greek and Roman indices are then interchangeable-indeed PSALTer only knows about one set of Lorentz indices on the Minkowski background, and these are strictly associated with Greek indices which represent Cartesian coordinates. There are 16 d.o.f. in f_i^{μ} and 24 d.o.f. in \mathscr{A}^{ij}_{μ} . The latter can clearly be accounted for, in the second-order formulation, by the d.o.f. in Eq. (13). Most of the former are accounted for by the metric d.o.f. in Eq. (5b), but there are six further d.o.f. in the antisymmetric part of the tetrad which do not appear in MAG. This is not a problem, because, these six d.o.f. are immediately eliminated by the six gauge generators of the Lorentz symmetry, part of the Poincaré symmetry, which also is not visible in the MAG. As a consequence, the spin-one matrices in Fig. 3 have two rows and two columns more than they would do in the MAG formulation, but the dimension of their null space also increases by 2. Kinematic extensions of the theory which are canceled by symmetries in this way do not alter the physics, and in this sense we understand the zero-nonmetricity MAG and the PGT to be equivalent theories.

Conjugate to the tetrad perturbation f_i^{μ} and the spin connection \mathscr{A}^{ij}_{μ} are the translational source (asymmetric

stress-energy tensor) $\tau^i_{\ \mu}$ and matter spin current $\sigma_{ij}^{\ \mu}$ [17,21,106]. The reduced-index SO(3) irreducible parts of these fields and sources label the rows and columns of the matrices in Fig. 3, and have spin-parity (J^P) labels to identify them. Duplicate J^P states are distinguished by additional parallel (||) and perpendicular (\perp) labels—but there is no significant meaning behind these auxiliary labels.

APPENDIX B: ZERO TORSION WITH PSALTER

Unlike in Appendix A, our *PSALTer* analysis corresponding to Sec. III B is fully grounded in the MAG formulation. The zero-torsion MAG kinematic module is displayed in Fig. 4. The first-order analyses in Figs. 5,7 and 9 share all our notational conventions above: the fields $h_{\mu\nu}$ and $A_{\mu\nu}^{\rho}$ are perturbative. To reach the second-order formulation, we only have to edit the quadratic action before substituting into the ParticleSpectrum function (which is the main function provided by the *PSALTer* package). The reparametrization used to transform the quadratic action is

$$A_{\mu}{}^{\rho}{}_{\nu} \mapsto A_{\mu}{}^{\rho}{}_{\nu} + \frac{1}{2} (2\partial_{(\mu|}h_{\lambda|\nu)} - \partial_{\lambda}h_{\mu\nu}). \tag{B1}$$

To lowest order in perturbative fields, Eq. (B1) captures the transition from $A_{\mu}{}^{\rho}{}_{\nu}$ to $\Delta_{\mu}{}^{\rho}{}_{\nu}$ set out in Sec. II A. The notation is slightly abusive, because $A_{\mu}{}^{\rho}{}_{\nu}$ on the rhs of Eq. (B1) is really $\Delta_{\mu}{}^{\rho}{}_{\nu}$. But since there is no advantage in defining a new kinematic module for *PSALTer* just to avoid the notational conflict, we therefore lazily recycle the first-order zero-torsion MAG module for all our second-order calculations.

Conjugate to the metric perturbation $h_{\mu\nu}$ and the affine connection $A_{\mu}{}^{\rho}{}_{\nu}$ are the (symmetric) stress-energy tensor $T^{\mu\nu}$ and the current $W^{\mu}{}_{\rho}{}^{\nu}$ which in MAG has become known as the *hypermomentum*. As with the PGT notation in Fig. 3, the J^P states are labelled as such. To distinguish the duplicate J^P states, apart from the (||) and (\perp) symbols, we use the letters (**s**), (h) and (t)—once again there is no significant meaning behind these labels. Different labels (numerical subscripts) are used in Eq. (18).

We show the general analysis of the theory in Eq. (19) in Figs. 5 and 6, respectively, for the first- and second-order formulations of the model. In Figs. 7–10 we consider tuned special cases of the model in which the massive 1^- state is allowed to propagate.

Further theories considered in Sec. III B, whose matrices are too cumbersome for the appendices, are presented in [87].

PSALTer kinematic panel

 $\frac{\text{Momentum Norm Frame}}{k^{\mu}} k^2 == k_{\mu} k^{\mu} n^{\mu} == \frac{k^{\mu}}{k}$

Fundamental fields

Fields	Symmetries	SO(3)	Sources
h _{aβ}	StrongGenSet[{1,2},GenSet[(1,2)]]	$\frac{1}{2} \int_{\alpha_{\beta}} \sigma_{\beta}^{\alpha} h^{1} - \frac{1}{2} \int_{\alpha} \sigma_{\beta}^{\alpha} \sigma_{\beta}^{\alpha} h^{2} + \sigma_{\alpha} \sigma_{\beta}^{\alpha} \sigma_{\beta}^{\alpha} h^{\alpha} + \sigma_{\alpha}^{\alpha} \tau_{\beta}^{\alpha} h^{\alpha} + \sigma_{\alpha}^{\alpha} \tau_{\beta}^{\alpha} h^{\alpha} + \sigma_{\alpha}^{\alpha} \tau_{\beta}^{\alpha} h^{\alpha} + \sigma_{\alpha}^{\alpha} \tau_{\beta}^{\alpha} h^{\alpha} h^{\alpha$	$\mathcal{T}_{\alpha\beta}$
$\mathcal{A}_{\alpha\beta\chi}$	StrongGenSet[{1,3},GenSet[(1,3)]]	$\frac{1}{2} \eta_{\beta \chi} n_{\alpha} \circ \mathcal{I}_{\beta \delta}^{\beta} + \frac{1}{2} \eta_{\alpha \chi} n_{\beta} \circ \mathcal{I}_{\beta}^{\delta} + \frac{1}{2} \eta_{\alpha \beta} n_{\chi} \circ \mathcal{I}_{\beta}^{\delta} + \frac{1}{2} n_{\alpha} n_{\beta} n_{\chi} \circ \mathcal{I}_{\beta \delta}^{\delta} + \frac{1}{2} n_{\chi} : \mathcal{I}_{\beta \delta}^{\delta} + \frac{1}{2} n_{\alpha} : \mathcal{I}_{\beta \delta}^{\delta} + \mathcal{I}_{\beta \delta}^{$	$W_{\alpha\beta\chi}$
		$\frac{1}{3} n_{a} z^{2} \mathcal{R}_{a}^{\dagger} \rho_{\chi} + \tilde{z} \mathcal{R}_{a}^{\dagger} a_{\rho\chi} - \frac{1}{6} \eta_{\rho_{\chi}} z^{2} \mathcal{R}_{a}^{\dagger} \rho_{a} + \frac{1}{6} n_{\rho_{g}} n_{\chi} z^{2} \mathcal{R}_{a}^{\dagger} \rho_{a} + \frac{1}{3} \eta_{\alpha\chi} z^{2} \mathcal{R}_{a}^{\dagger} \rho_{\beta} - \frac{1}{3} \eta_{a} n_{\chi} z^{2} \mathcal{R}_{a}^{\dagger} \rho_{\beta} - \frac{1}{6} \eta_{\alpha\beta} z^{2} \mathcal{R}_{a}^{\dagger} \rho_{\chi} + \frac{1}{3} \eta_{\alpha\chi} z^{2} \mathcal{R}_{a}^{\dagger} \rho_{\chi} - \frac{1}{2} \eta_{\alpha\beta} z^{2} \rho_{\alpha\beta} z^$	
		$\frac{1}{15}n_{\beta}n_{\chi}\left[\mathcal{A}_{5}\mathbf{k}_{\alpha}\right]^{n}+\frac{1}{15}\eta_{\alpha\chi}\left[\mathcal{A}_{5}\mathbf{k}_{\beta}\right]_{\beta}-\frac{1}{15}n_{\alpha}n_{\chi}\left[\mathcal{A}_{5}\mathbf{k}_{\beta}\right]+\frac{1}{15}\eta_{\alpha\beta}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{15}n_{\alpha}n_{\beta}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]^{2}-\frac{1}{3}n_{\chi}\left[\mathcal{A}_{5}^{*}\mathcal{A}_{5}\right]+\frac{1}{3}n_{\beta}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}+\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{\gamma}\right]_{\gamma}-\frac{1}{3}n_{\alpha}\left[\mathcal{A}_{5}\mathbf{k}_{$	
		$\frac{2}{9} n_{0x} n_{\rho} \tilde{S} S^{ab} - \frac{1}{9} n_{ga} n_{x} \tilde{S} S^{ab} - \frac{1}{3} n_{\rho} n_{x} \tilde{S} S^{ab} - \frac{1}{3} n_{\sigma} n_{x} \tilde{S} S^{ab} - \frac{1}{3} n_{\sigma} n_{\rho} \tilde{S} S^{ab} + \frac{1}{3} n_{\sigma} n_{\rho} \tilde{S} S^{ab} + \frac{1}{3} n_{\sigma} n_{x} \tilde{S} S^{ab} + \frac{1}{3} n_{\sigma} n_{\mu} \tilde{S} S^{ab} + \frac{1}{3} n_{\mu} n_{\mu} \tilde{S} S^{ab} + $	

SO(3) irreps

SO(3) 5	ymmetries	Expansion	Sources
0 ⁺ . <i>h</i> ⁺	StrongGenSet[{},GenSet[]]	$n^{\alpha} n^{\beta} h_{\alpha\beta}$	$^{0^{+}}\mathcal{T}^{\perp}$
0* <i>h</i> I	StrongGenSet[{}, GenSet[]] -	$n^{a} n^{\beta} h_{a\beta} + h^{a}_{a}$	0*7~II
$1_{h^{+}\alpha}$	StrongGenSet[{},GenSet[]]	$\pi^{\beta} h_{\alpha\beta} - n_{\alpha} \pi^{\beta} \pi^{\chi} h_{\beta\chi}$	${}^{1}\mathcal{T}_{\alpha}^{\perp}$
²⁺ μ ¹ αβ	StrongGenSet[{1,2},GenSet[(1,2)]]	$h_{\alpha\beta} - n_{\beta} n^{\chi} h_{\alpha\chi} - n_{\alpha} n^{\chi} h_{\beta\chi} + \frac{1}{3} \eta_{\alpha\beta} n^{\chi} n^{\delta} h_{\chi\delta} + \frac{2}{3} n_{\alpha} n_{\beta} n^{\chi} n^{\delta} h_{\chi\delta} - \frac{1}{3} \eta_{\alpha\beta} h^{\chi}_{\chi} + \frac{1}{3} n_{\alpha} n_{\beta} h^{\chi}_{\chi}$	$2^{+}T^{+}a\beta$
${}^{0^+}\mathcal{A}_{s}{}^{\perp t}$	StrongGenSet[{},GenSet[]]	$n^{lpha} n^{eta} n^{\kappa} \mathcal{A}_{abk}$	$^{0^{+}}W_{s}^{\perp t}$
⁰⁺ <i>P</i> I _S ^{II}	StrongGenSet[{}, GenSet[]] -3	$n^{\alpha} n^{\beta} n^{\chi} \mathcal{F}_{\alpha\beta\chi} + 2 n^{\alpha} \mathcal{F}_{\alpha\beta} + n^{\alpha} \mathcal{F}_{\alpha\beta}^{\beta} + n^{\alpha} \mathcal{F}_{\alpha\beta}^{\beta}$	⁰⁺ W _s I
⁰⁺ <i>A</i> s ^{1h}	StrongGenSet[{}, GenSet[]] -	$n^{a} \mathcal{A}^{\ \beta}_{a\ \beta} + n^{a} \mathcal{A}^{\ \beta}_{a\beta}$	$^{0^+}W_{s}^{\perp h}$
$1^+ \mathcal{R}_{s^+ \alpha \beta}$	StrongGenSet[{1,2},GenSet[-(1,2)]]	$\pi^{\chi} \ \mathcal{A}_{\alpha\beta\chi} - n_{\beta} \ \pi^{\chi} \ n^{\delta} \ \mathcal{A}_{\alpha\chi\delta} - \pi^{\chi} \ \mathcal{A}_{\beta\alpha\chi} + n_{\alpha} \ \pi^{\chi} \ n^{\delta} \ \mathcal{A}_{\beta\chi\delta} + n_{\beta} \ \pi^{\chi} \ n^{\delta} \ \mathcal{A}_{\chi\beta\delta}$	$^{1^+}W_{s^{\perp}a\beta}$
${}^1\mathcal{R}_{\mathrm{S}}{}^{\mathrm{it}}{}_{\alpha}$	StrongGenSet[{}, GenSet[]] 2	$n^{\beta} n^{\chi} \mathcal{R}_{\alpha\beta\chi} + n^{\beta} n^{\chi} \mathcal{R}_{\beta\beta\chi} \cdot 3 n_{\alpha} n^{\beta} n^{\chi} n^{\delta} \mathcal{R}_{\beta\chi\delta}$	${}^{1}W_{s}^{\mu}{}_{\alpha}$
${}^{1}\mathcal{R}_{\mathrm{s}}{}^{\mathrm{pt}}{}_{\alpha}$	StrongGenSet[{}, GenSet[]] -2	$n^{\beta} n^{\chi} \mathcal{R}_{\alpha\beta\chi} + 2 \mathcal{R}_{\alpha\beta}^{\beta} - n^{\beta} n^{\chi} \mathcal{R}_{\beta\alpha\chi} + 3 n_{\alpha} n^{\beta} n^{\chi} n^{\delta} \mathcal{R}_{\beta\chi\delta} - 2 n_{\alpha} n^{\beta} \mathcal{R}_{\beta\chi}^{\chi} + \mathcal{R}_{\beta\beta}^{\beta} - n_{\alpha} n^{\beta} \mathcal{R}_{\beta\chi}^{\chi}$	${}^{1}W_{s}{}^{\parallel}_{\alpha}$
${}^1\mathcal{R}_{\rm s}{}^{\rm {\scriptscriptstyle \bot}h}{}_{\alpha}$	StrongGenSet[{}, GenSet[]] -	$\mu^{\beta} n^{\chi} \mathcal{F}_{\alpha\beta\chi} + n^{\beta} n^{\chi} \mathcal{F}_{\beta\alpha\chi}$	${}^{1}W_{s}^{ih}a$
${}^1\mathcal{R}_{\rm S}{}^{\rm Jh}{}_{\alpha}$	StrongGenSet[{},GenSet[]]	$n^{\beta} n^{\chi} \mathcal{A}_{\alpha\beta\chi} - \mathcal{A}_{\alpha\beta}^{\ \beta} - n^{\beta} n^{\chi} \mathcal{A}_{\beta\alpha\chi} + n_{\alpha} n^{\beta} \mathcal{A}_{\beta\chi}^{\ \chi} + \mathcal{A}_{\alpha\beta}^{\beta} - n_{\alpha} n^{\beta} \mathcal{A}_{\beta\chi}^{\chi}$	${}^{1}W_{s}^{\parallel h}\alpha$
$2^{+}\mathcal{A}_{s}^{\parallel}{}_{\alpha\beta}$	StrongGenSet[{1,2},GenSet[(1,2)]]	$n^{\chi} \ \mathcal{R}_{\alpha\beta\chi} + n^{\chi} \ \mathcal{R}_{\alpha\chi\delta} - 2 \ n_{\rho} \ n^{\chi} \ n^{\delta} \ \mathcal{R}_{\alpha\chi\delta} + n^{\chi} \ \mathcal{R}_{\beta\alpha\chi} - 2 \ n_{a} \ n^{\chi} \ n^{\delta} \ \mathcal{R}_{\beta\gamma\delta} - n_{\rho} \ n^{\chi} \ n^{\delta} \ \mathcal{R}_{\chi\alpha\delta} - n_{a} \ n^{\chi} \ n^{\delta} \ \mathcal{R}_{\chi\beta\delta} + n^{\chi} \ \mathcal{R}_{\alpha\beta\delta} $	²⁺ ₩ _s I _{αβ}
		$\eta_{\alpha\beta} n^{\chi} n^{\delta} n^{\epsilon} \mathcal{R}_{\chi\delta\epsilon} + 2 n_{\alpha} n_{\beta} n^{\chi} n^{\delta} n^{\epsilon} \mathcal{R}_{\chi\delta\epsilon} - \frac{2}{3} \eta_{\alpha\beta} n^{\chi} \mathcal{R}_{\chi\delta}^{\delta} + \frac{4}{3} n_{\alpha} n_{\beta} n^{\chi} \mathcal{R}_{\chi\delta}^{\delta} - \frac{1}{3} \eta_{\alpha\beta} n^{\chi} \mathcal{R}_{\chi\delta}^{\delta} + \frac{1}{3} n_{\alpha} n_{\beta} n^{\chi} \mathcal{R}_{\chi\delta}^{\delta}$	
$^{2^{+}}\mathcal{A}_{S^{\perp}\alpha\beta}$	StrongGenSet[{1, 2}, GenSet[(1,2)]] -	$\frac{1}{2}n^{\chi}\mathcal{A}_{a\beta\chi} + n^{\chi}\mathcal{A}_{a\chi\delta} - \frac{1}{2}n_{\beta}n^{\chi}n^{\delta}\mathcal{A}_{a\chi\delta} - \frac{1}{2}n^{\chi}\mathcal{A}_{\beta\alpha\chi} - \frac{1}{2}n^{\chi}\mathcal{A}_{\beta\alpha\chi} + \frac{1}{2}n_{\alpha}n^{\chi}n^{\delta}\mathcal{A}_{\beta\chi\delta} +$	$^{2^+}W_{s^{\perp}\alpha\beta}$
		$\frac{1}{2} n_{\beta} n^{\chi} n^{\delta} \mathcal{R}_{\chi \alpha \delta} + \frac{1}{2} n_{\alpha} n^{\chi} n^{\delta} \mathcal{R}_{\chi \beta \delta} + \frac{1}{3} \eta_{\alpha \beta} n^{\chi} \mathcal{R}_{\chi \delta}^{\delta} - \frac{1}{3} n_{\alpha} n_{\beta} n^{\chi} \mathcal{R}_{\chi \delta}^{\delta} - \frac{1}{3} \eta_{\alpha \beta} n^{\chi} \mathcal{R}_{\chi \delta}^{\delta} + \frac{1}{3} \eta_{\alpha \beta} n^{\chi} \mathcal{R}_{\chi \delta}^{\delta} + \frac{1}{3} n_{\alpha} n_{\beta} n^{\chi} \mathcal{R}_{\chi \delta}^{\delta} + \frac{1}{3} n_{\alpha} n^{\chi} n^{\chi} n^{\chi} \mathcal{R}_{\chi \delta}^{\delta} + \frac{1}{3} n^{\chi} $	
${}^{2}\mathcal{R}_{s}{}^{\dagger}{}_{\alpha\beta\chi}$	StrongGenSet[{1, 2}, GenSet[(1,2)]] -	$\frac{1}{3} \mathcal{A}_{\alpha\beta\chi} + \frac{1}{3} n_{\chi} n^{\delta} \mathcal{A}_{\alpha\beta\delta} + \frac{2}{3} \mathcal{A}_{\alpha\chi\delta} - \frac{2}{3} n_{\mu} n^{\delta} \mathcal{A}_{\alpha\chi\delta} - \frac{2}{3} n_{\chi} n^{\delta} \mathcal{A}_{\alpha\delta\rho} + \frac{1}{3} n_{\rho} n^{\delta} \mathcal{A}_{\alpha\delta\chi} + \frac{1}{6} n_{\rho\chi} n^{\delta} n^{\epsilon} \mathcal{A}_{\alpha\delta\epsilon} + \frac{1}{6} n_{\rho} n_{\chi} n^{\delta} n^{\epsilon} \mathcal{A}_{\alpha\delta\epsilon} - \frac{1}{6} n_{\rho\chi} \mathcal{A}_{\alpha\delta} + \frac{1}{6} n_{\rho\chi} n^{\delta} n^{\epsilon} \mathcal{A}_{\alpha\delta\epsilon} + \frac{1}{6} n_{\rho\chi} n^{\delta} n^{\epsilon} n^{\epsilon} \mathcal{A}_{\alpha\delta\epsilon} + \frac{1}{6} n_{\rho\chi} n^{\delta} n^{\epsilon} \mathcal{A}_{\alpha\delta\epsilon} + \frac{1}{6} n_{\rho\chi} n^{\delta} n^{\epsilon} n^{\epsilon}$	² W _s [∎] _{aβχ}
		$\frac{1}{6}n_{\beta}n_{\chi}\mathcal{A}_{a}^{b}c^{-\frac{1}{3}}\mathcal{A}_{\beta\alpha\chi} + \frac{1}{3}n_{\chi}n^{\delta}\mathcal{A}_{\beta\alpha\alpha\delta} - \frac{2}{3}n_{a}n^{\delta}\mathcal{A}_{\beta\alpha\delta} + \frac{1}{3}n_{a}n^{\delta}\mathcal{A}_{\beta\delta\chi} + \frac{1}{6}n_{\alpha\chi}n^{\delta}n^{\epsilon}\mathcal{A}_{\beta\delta\kappa} + \frac{1}{6}n_{\alpha}n_{\chi}n^{\delta}n^{\epsilon}\mathcal{A}_{\beta\delta\kappa} + \frac{1}{6}n_{\alpha}n_{\chi}\mathcal{A}_{\beta}^{b}c^{-\frac{1}{6}}n_{\alpha}\mathcal{A}_{\beta}\mathcal{A}_{\beta} + \frac{1}{6}n_{\alpha}n_{\chi}\mathcal{A}_{\beta}^{b}c^{-\frac{1}{6}}n_{\alpha}\mathcal{A}_{\beta}\mathcal{A}_{\beta} + \frac{1}{6}n_{\alpha}n_{\chi}\mathcal{A}_{\beta}\mathcal{A}_{\beta} + \frac{1}{6}n_{\alpha}n_{\chi}\mathcal{A}_{\beta} + \frac{1}{6}n_{\alpha}n_{\chi}\mathcal{A}_{\beta} + \frac{1}{6}n_{\alpha}n_{\chi}\mathcal{A}_{\beta} + \frac{1}{6}n_{\alpha}n_{\chi}\mathcal{A}_{\beta}$	
		$\frac{1}{3}n_{\beta}n^{\delta}\mathcal{A}_{\chi\alpha\delta} + \frac{1}{3}n_{\alpha}n^{\delta}\mathcal{A}_{\chi\delta\delta} - \frac{1}{3}\eta_{\alpha\beta}n^{\delta}n^{\epsilon}\mathcal{A}_{\chi\delta\epsilon} - \frac{1}{3}n_{\alpha}n_{\beta}n^{\delta}n^{\epsilon}\mathcal{A}_{\chi\delta\epsilon} + \frac{1}{3}\eta_{\alpha\beta}\mathcal{A}_{\chi\delta}^{\delta} - \frac{1}{3}n_{\alpha}n_{\beta}\mathcal{A}_{\chi\delta}^{\delta} - \frac{1}{6}\eta_{\beta\chi}n^{\delta}n^{\epsilon}\mathcal{A}_{\delta\alpha\epsilon} - \frac{1}{6}n_{\beta}n_{\chi}n^{\delta}n^{\epsilon}\mathcal{A}_{\delta\alpha\epsilon} - \frac{1}{6}n_{\beta}n_{\chi}n^{\delta}\mathcal{A}_{\delta\alpha\epsilon} - \frac{1}{6}n_{\chi}n^{\delta}\mathcal{A}_{\delta\alpha\epsilon} - \frac{1}{6}n_{\chi}n^{$	
		$\frac{1}{6}\eta_{ax} n^{\delta} n^{\epsilon} \mathcal{A}_{\delta\beta\epsilon} - \frac{1}{6}n_{a} n_{\chi} n^{\delta} n^{\epsilon} \mathcal{A}_{\delta\delta\epsilon} + \frac{1}{3}\eta_{a\beta} n^{\delta} n^{\epsilon} \mathcal{A}_{\delta\epsilon\epsilon} + \frac{1}{3}\eta_{a} \eta_{\beta} n^{\delta} n^{\epsilon} \mathcal{A}_{\delta\epsilon\epsilon} + \frac{1}{6}\eta_{\beta\chi} n_{a} n^{\delta} \mathcal{A}_{\delta\epsilon} + \frac{1}{6}\eta_{a\chi} n_{\beta} n^{\delta} \mathcal{A}_{\delta\epsilon} - \frac{1}{3}\eta_{a\beta} n_{\chi} n^{\delta} \mathcal{A}_{\delta\epsilon} + \frac{1}{2}\eta_{a\beta} n_{\chi} n^{\delta} \mathcal{A}_{\delta\epsilon} + \frac{1}{2}\eta_{\alpha} n_{\chi} n^{\delta} n^{\delta} \mathcal{A}_{\delta\epsilon} + \frac{1}{2}\eta_{\alpha} n_{\chi} n^{\delta} n^$	
		$\frac{1}{6}\eta_{\beta\chi}\mathcal{A}^{\delta}_{\alpha\delta} - \frac{1}{6}n_{\beta}n_{\chi}\mathcal{A}^{\delta}_{\alpha\delta} + \frac{1}{6}\eta_{\alpha\chi}\mathcal{A}^{\delta}_{\beta\delta} - \frac{1}{6}n_{\alpha}n_{\chi}\mathcal{A}^{\delta}_{\beta\delta} - \frac{1}{3}\eta_{\alpha\beta}\mathcal{A}^{\delta}_{\chi\delta} + \frac{1}{3}n_{\alpha}n_{\beta}\mathcal{A}^{\delta}_{\chi\delta} - \frac{1}{6}\eta_{\beta\chi}n_{\alpha}n^{\delta}\mathcal{A}^{\delta}_{\delta\epsilon} - \frac{1}{6}\eta_{\alpha\chi}n_{\beta}n^{\delta}\mathcal{A}^{\delta}_{\delta\epsilon} + \frac{1}{3}\eta_{\alpha\beta}n_{\chi}n^{\delta}\mathcal{A}^{\delta}_{\delta\epsilon} - \frac{1}{6}\eta_{\alpha\chi}n_{\beta}n^{\delta}\mathcal{A}^{\delta}_{\delta\epsilon} + \frac{1}{3}\eta_{\alpha\beta}n_{\chi}n^{\delta}\mathcal{A}^{\delta}_{\delta\epsilon} - \frac{1}{6}\eta_{\alpha\chi}n_{\beta}n^{\delta}\mathcal{A}^{\delta}_{\delta\epsilon} + \frac{1}{3}\eta_{\alpha\beta}n_{\chi}n^{\delta}\mathcal{A}^{\delta}_{\delta\epsilon} + \frac{1}{3}\eta_{\alpha\beta}n_{\chi}n^{\delta}$	
${}^{3}\mathcal{R}_{5}{}^{\dagger}{}_{\alpha\beta\chi}$	StrongGenSet[{1, 2, 3}, GenSet[(1,2), (2,3)]]	$\frac{1}{3}\mathcal{F}_{\alpha\beta\gamma} + \frac{1}{3}\sigma_{\alpha\gamma} + \frac{1}{3}\mathcal{F}_{\alpha\gamma} + \frac{1}{3}\mathcal{F}_{\alpha\gamma} + \frac{1}{3}\sigma_{\alpha\gamma} + \frac{1}{3}\sigma_{\alpha\gamma} + \frac{1}{3}\sigma_{\alpha\gamma} + \frac{1}{3}\sigma_{\beta\gamma} + $	³ W _s [∥] _{αβχ}
		$\frac{1}{3}\mathcal{B}_{\beta\alpha_{1}}+\frac{1}{3}n_{\chi}n^{\delta}\mathcal{B}_{\beta\alpha_{5}}-\frac{1}{3}n_{a}n^{\delta}\mathcal{B}_{\beta\beta_{0}\beta}+\frac{1}{3}n_{a}n^{\delta}\mathcal{B}_{\beta\beta_{0}\gamma}+\frac{2}{15}\eta_{\alpha_{\chi}}n^{\delta}n^{\epsilon}\mathcal{B}_{\beta\beta_{6}}+\frac{2}{31}n_{a}n_{\chi}n^{\delta}n^{\epsilon}\mathcal{B}_{\beta\beta_{6}}+\frac{2}{15}\eta_{\alpha_{\chi}}\mathcal{B}_{\beta}^{\delta}+\frac{1}{31}n_{a}n_{\chi}\mathcal{B}_{\beta}^{\delta}\partial+\frac{1}{31}n_{\mu}n^{\delta}\mathcal{B}_{\gamma_{\alpha}\delta}-\frac{1}{3}n_{\mu}n^{\delta}\mathcal{B}_{\gamma_{\alpha}}+\frac{1}{3}n_{\mu$	
		$\frac{1}{2}n_n n^6 \mathcal{A}_{vab} + \frac{2}{16}n_{aab} n^6 n^6 \mathcal{A}_{vabc} + \frac{8}{16}n_n n_a n^\delta n^6 \mathcal{A}_{vabc} - \frac{2}{12}n_{ab} \mathcal{A}_{vb} + \frac{2}{16}n_n n_a \mathcal{A}_{vb} + \frac{1}{16}n_{ab} n_b n^\delta n^6 \mathcal{A}_{bac} + \frac{4}{16}n_a n_a n^\delta n^6 \mathcal{A}_{bac} + \frac{1}{16}n_{ab} n^{\delta} n^{\delta} n^{\delta} \mathcal{A}_{bac} + \frac{1}{16}n_{ab} n^{\delta} n^{\delta} n^{\delta} \mathcal{A}_{bac} + \frac{1}{16}n_{ab} n^{\delta} n^{\delta} n^{\delta} n^{\delta} \mathcal{A}_{bac} + \frac{1}{16}n_{ab} n^{\delta} n^{\delta}$	
		$\frac{1}{2}\eta_{n}\eta_{v}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{12}\eta_{n}\eta_{n}\theta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{4}{24}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta^{\delta}\eta^{\delta}\mathcal{A}_{hqs} + \frac{1}{2}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta_{n}\eta_{n$	
		$ \begin{bmatrix} 1 & \cdots &$	
		$ \begin{array}{c} \vdots \\ \eta_{ij} $	
		ע א אי כז אין איי אי אי אי אי גע איי איי אין איי אין איי אין א	1

FIG. 4. Kinematic structure of zero-torsion MAG, as studied throughout Sec. III B. The SO(3) irreps precisely correspond to those in Eqs. (5b) and (18), though the labelling of duplicate J^P states is different from that in [28]. The key point is that the connection field carries an extra symmetry restriction $A_{\mu}{}^{\rho}{}_{\nu} \equiv A_{(\mu}{}^{\rho}{}_{|\nu)}$, and by referring to Eq. (1) we see that this kills off the torsion in the first-order formulation. In moving to the second-order formulation, we make a slight notational abuse in Eq. (B1), but from Eq. (4) we see that the effect will still be as desired if we keep using this kinematic module. As with Fig. 2, the 2⁻ and 3⁻ states have extra cyclic symmetries which are hidden. These definitions are used in Figs. 5–10.

PSALTer results panel Wave operator and propagator Multiplicities Spin-parityform Covariantform a.) <u>.</u> $k^{0^+} W_{s^{\perp}t} + 2 i^{0^+} T^{\perp} = 0$ 2 $\partial_{\beta}\partial_{\alpha}\mathcal{T}^{\alpha\beta} = \partial_{\lambda}\partial_{\beta}\partial_{\alpha}\mathcal{W}^{\alpha\beta\chi}$ $\sqrt{5}(-k^2(n_{11}-2,n_7+4,n_9))$ $\frac{1}{12} \left(k^2 \left(4 \, h \right)_{11} - 2 \left(h \right)_{7} + 4 \, h \right) \right)$ $\frac{1}{12}(k^2(4h_{10}+h_{12}-2h_8)+$ 3 $2\,\partial_{\chi}\partial^{\chi}\partial_{\beta}\mathcal{T}^{\alpha\beta}+\partial_{\delta}\partial_{\chi}\partial_{\beta}\partial^{\alpha}\mathcal{W}^{\beta\chi\ \delta}$ 1 Gsth a 0 1 k a $\frac{+h}{12}$ Total expected gauge generators x² (4 A 2+.A.I. 2. Ast a 2+ hl 2 Acl Ika 2.+ hl † 0 0 -I¤ 4 √3 a. 1 k a 4 √3 $k^{2}(2h, -h, -h, -4(h, +h))+2h$ $\frac{1}{6} \left(-k^2 \left(h_{10} + h_{11} + h_{12} + h_{1} \right) - 3 a_{0}$ 0 ^{2,*}*A*s¹† (.° + (°,4 + $\frac{k^2 (2 h_{10} h_{11} h_{12} - 4(h_{7} + h_{8}) + 2 h_{9})}{12 \sqrt{2}}$ $\frac{1}{12}(k^2(h_{11}+h_7+h_9)+$ $\frac{1}{12} \left(-k^2 \left(h_{10} - 2 h_{11} - 2 h_{12} + 4 \left(h_{7} + h_{1} \right) + h_{1} \right) + 3 a_{0}$ k^{2} (4 h + h 2 h) - 4 a0 ?.⁺*F*is⁺† 0 $\frac{1}{12}\sqrt{\frac{5}{2}}\left(-2\,k^{2}\left(h_{10}+h_{12}\right)\right)$ $2k^{2}(k_{11}+k_{7}+k_{9})$ r Se[⊥] μ *i k* a 8 √3 4-0 -0 2. As¹ † 0 0 0 $^{0^+}W_{s}^{\perp l}$ $0^+_{\tau} T^{+}$ $0^+ T^ ^{0^{+}}\mathcal{W}_{s}^{\perp t}$ 0⁺W_sI $8k^2(2h-h-h+2h)$ $\frac{4i\sqrt{2}k}{3(4+k^2)}$ $4k^2(2k^2)$ $\frac{10 i k}{3(4+k^2)a}$ $0^+ T^ 3(4+k^2)^2a$ $\sqrt{3}(4+k^2)a$ $\frac{8k^2(2h,h,h+2h)}{\sqrt{3}(4+k^2)a_0^2}$ $2 k^{2} (h, -h, -h, +h, +h, +h,)+a$ $\frac{\frac{161k(2k-k-k+2k)}{10}}{\sqrt{3}(4+k^2)a_0^2}$ $\frac{4i\sqrt{\frac{2}{3}}}{ka_{0}}$ $\frac{2i}{\sqrt{3}ka}$ $0^{+}\mathcal{T}^{\parallel}$ a.) a.) 0 (;e $\frac{1}{6} \left(-5 \, k^2 \, (h_{11} + h_7 + h_0) + 2 \, a_0 \right)$ $\sqrt{\frac{5}{2}} \left(-2k^2 \left(h_{10} + h_{12} + h_{0}\right) +$ $\frac{1}{12} \sqrt{5} (k^2 (h_{11} - 2h_7 + 4h_0) +$ $\frac{16ik(2k-k-k+2k)}{\sqrt{3}(4+k^2)a_0^2}$ $16(2 k^{2} (7 h, +3(h, +h, +h, +h, +h, +7 h)) = 0$ $16ik^{3}(7h_{10}+3(h_{11}+h_{12}+h_{7}+h_{8})+7h_{9})+8ik$ $\frac{20}{3(4+k^2)a}$ $8\sqrt{2}$ $8(4+k^2)$ $\frac{1}{6}\sqrt{5}(k^2(h_{10}+h_{12}+h_{8})+$)⁺W_s⊥t $\frac{11}{3(4+k^2)^2}a$ $3(4+k^2)^2 a_1$ 1 A_S ^{∥t} $\frac{1}{4}i\sqrt{\frac{5}{6}}k$ 101 k 12a,+3k²a $\frac{2i}{\sqrt{3} k a}$ 20 12a,+3k²a, 0 $^{0^+}W_{\rm S}$ 0 0 $-\frac{4i\sqrt{\frac{2}{3}}}{ka_{0}}$ $\frac{4i\sqrt{2}k}{12a+3k^2a}$ $\frac{8\sqrt{2}}{12a_1+3k^2a_1}$ 0 0 We¹ 3 W. ³ *A*_s[∥]_{ab} -<u>2</u> a. - 0 - 0 \mathcal{R}^{1} 리워 2 Ws 2*W5 $\sqrt{5}(k^2(h_{10}+h_1+h_8)+a_0)$ a.) $2^{+}_{\cdot} \mathcal{T}^{\parallel}_{al}$ 2+ W. $a_{n}h^{a}_{a}\partial_{x}\mathcal{M}^{\beta,\chi}_{\beta}+a_{n}h^{a}_{a}\partial_{x}\mathcal{M}^{\beta,\chi}_{\beta}=2h_{\mu}\partial_{\beta}\mathcal{M}^{\delta}_{\chi\delta}\partial^{\lambda}\mathcal{M}^{a}$ $\frac{1}{5}(k^{2}(h_{11}+h_{7}+h_{0}) - 2a_{0})$ $h_{\alpha\beta}$ +2 $a_{\alpha}h^{\alpha\beta}\partial_{\beta}\mathcal{R}_{\alpha\chi}^{\chi}$ 2 $a_{\alpha}h^{\alpha\beta}\partial_{\chi}\mathcal{R}_{\alpha\beta}^{\chi}$ $1h_{i} \partial^{\chi} \mathcal{R}^{\alpha\beta}_{\ a} \partial_{\delta} \mathcal{R}^{\ b}_{\ \chi}^{\ b})](t, x, y, z] d z d y d x d i$ $\frac{k^4 \left((k_{\frac{1}{21}}+k_{\frac{1}{22}})^2 - 4(k_{\frac{1}{2}}+k_{\frac{1}{2}})(k_{\frac{1}{2}}+k_{\frac{1}{2}}+k_{\frac{1}{2}}) + 8k^2 \left(\frac{k_{\frac{1}{2}}}{k_{\frac{1}{2}}+k_{\frac{1}{2}}+k_{\frac{1}{2}}+k_{\frac{1}{2}}+k_{\frac{1}{2}}\right)}{k^2 a_0^{-2} \left(k_2^2 \left(k_{\frac{1}{2}}+k_{\frac{1}{2}}+k_{\frac{1}{2}}\right) + a_0^2\right)}$ $i(k^2(h, -h, -h, +h,)-a, 10, 11, 12, 9, 0)$ $2i\sqrt{\frac{2}{3}}(k^2(2k_1+k_1+k_2+2k_2)+4a_1)$ $\frac{1}{12}(k^2(4h_{10}+h_{12}-2h_8)+$ $H_{9} = \partial_{\alpha} \mathcal{R}^{\alpha\beta\chi} \partial_{\delta} \mathcal{R}_{\chi \beta}^{\ \delta} + 2 \ h \cdot \partial^{\chi} \mathcal{R}^{\alpha \beta}_{\ \alpha} \partial_{\delta} \mathcal{R}_{\chi \beta}^{\ \delta} +$ 0 2⁺T⁻¹† $2h_{11} \partial^{\chi} \mathcal{A}^{\alpha\beta} \partial_{\delta} \mathcal{A}^{\delta}_{\beta\chi} + 2h_{12} \partial^{\chi} \mathcal{A}^{\alpha\beta} \partial_{\delta} \mathcal{A}^{\delta}_{\beta\chi}$ √3 k a. (Waby 4 h, d^x A^a^B d₅ A⁵₈ + 4 h, d^x A^a^B d₅ A⁵₈ - $2h_{11}\partial_{x}\mathcal{A}^{\delta}_{\ \ Be}\partial^{x}\mathcal{A}^{\alpha}_{\ \ Be}\partial^{x}\mathcal{A}^{\alpha}_{\ \ \sigma} = 2h_{0}\partial^{\beta}\mathcal{A}^{\delta}_{\ \ x}\partial^{x}\mathcal{A}^{\alpha\beta}_{\ \ \sigma}$ $2h_{11} \partial_a \mathcal{R}^{ab \chi} \partial_b \mathcal{R}^{\delta}_{\beta \chi} 2h_{22} h_{12} \partial_a \mathcal{R}^{ab \chi} \partial_b \mathcal{R}^{\delta}_{\beta \chi}$ $k = (k^2 (h_1 + h_2) + a$ $^{\delta}$ +4 h , $\partial^{\chi} \mathcal{R}^{a\beta}_{\ a} \partial_{\delta} \mathcal{R}_{\beta}$ $2h_{12}\partial_{\beta}\mathcal{R}^{\delta}_{\ \ \chi} \partial^{\rho}\mathcal{R}^{\alpha}_{\ \ \alpha}^{\ \ \beta} 2h_{\eta}\partial_{\chi}\mathcal{R}^{\beta}_{\ \ \delta}\partial^{\nu}\mathcal{R}^{\alpha}_{\ \ \alpha}^{\ \ \beta}$ $2h_{a}^{}\partial_{x}\mathcal{B}^{}_{\ \ \beta\delta}\partial^{x}\mathcal{B}^{a\beta}_{\ \ \alpha} = 2h_{a}^{}\partial_{\beta}\mathcal{B}^{a\beta}_{\ \ \alpha}\partial_{\delta}\mathcal{B}^{a}_{\ \ \alpha}^{\delta}$ $2h_{B}\partial_{\beta}\mathcal{R}^{a\beta\chi}\partial_{\delta}\mathcal{R}^{\delta}_{a\chi}^{\delta}$ - $2h_{10}\partial_{a}\mathcal{R}^{a\beta\chi}\partial_{\delta}\mathcal{R}_{\beta\chi}$ 1 grs^⊥t 1 k a 4 √6 41(k²(-h,+h,+h,-h)+a 10 11 12 9 0 $\frac{4\sqrt{2}}{3(k^2(h_1+h_2)+a_1)}$ 0 ${}^{2^+}W_{s}^{\parallel}$ † $3(k^2(h_1+h_2)+a_1)$ $\mathcal{F}^{X}_{\beta\chi}$ +4 $\mathcal{F}^{\alpha\beta\chi}$ $\sqrt{3} k a_{h} (k^{2} (h_{10} + h_{10}) + a_{10})$ 2 k² (k + $\frac{2i\sqrt{\frac{2}{3}}(k^2(2\lambda_{10}+\lambda_{11}+\lambda_{12}+2\lambda_{10})+4a_{0})}{ka_{0}(k^2(\lambda_{10}+\lambda_{10})+a_{0})}$ $\frac{4 \sqrt{2}}{3(k^2 (h_1 + h_2) + a_1)}$ 0 2+ Ws+ t $\frac{1}{3(k^2(h_1+h_2)+a_1)}$ 4 a. +2 a, $\mathcal{R}^{\alpha\beta}_{\alpha}$ 0 0 0 $2h_{12} \partial^{\chi} \mathcal{A}^{\alpha \beta}{}^{\beta} \partial_{\delta} \mathcal{A}_{\beta \chi}$ 2 W. 1 k a $-\frac{1}{4}i\sqrt{\frac{5}{6}k}$ *i k a* 4 √6 1 k 0 8 \3 $\frac{i k a}{0}$ $1 h^{\perp}_{a}$ ${}^{0^+}h^{\parallel} {}^{0^+}\mathcal{F}_{r}^{\perp}$ 0+.A.1h 0 0⁺h⁺ 0+ A. $\frac{i k a}{4}$ $\frac{i k \frac{a}{0}}{8 \sqrt{2}}$ Xga B . 0,+ 0 0 0 1 k a 4 √3 0,+ h 0 5 8 √6 $1^+ \mathcal{A}_{s^+a_l}$ $\frac{1}{2}(k^2(h_{10}-h$ 0 0 0 0 0 $S == \iiint \left[\left(\frac{1}{4} \left(-2 a \right) \right) \right]$ 0 $\frac{a_{0}}{4\sqrt{2}}$ $0^+\mathcal{H}_{S}^{\perp}$ $2k^{2}(5k_{10}+2k_{11}+2k_{12}-k_{7}-k_{1}+5k_{9})+3a_{0}$ 1 k g 4 √3 4. 2 -<u>1</u> i k a $\frac{2}{3}k^{2}(h_{10} + h_{11} + h_{12} + h_{1} + h_{8} + h_{9})$ 0. AS $^{1}\mathcal{A}_{s}{}^{^{\mathrm{a}t}}\dagger^{a}$ $^1 \mathcal{R}_{\rm S}{}^{\rm lf} \dagger^a$ $^{1}\mathcal{R}_{s}^{,\mathrm{th}}+^{a}$ 1 A_sth†^a $^{1}H^{1}$ 51 k 4 8 √6 $\frac{a_{\dot{0}}}{4\sqrt{2}}$ +2 k +2 k -k -k +5 h)+3 a $\frac{1}{12} \left(-k^2 \left(13 h_{10} - 8 h_{11} - 8 h_{12} + 7 \left(h_7 + h_6 \right) + 13 h_6 \right) - 3 a_6 \right) \\ 0$ [⊥]s € $0^+ \mathcal{R}_{S}^{\perp \dagger}$ Massive and massless spectra Massive particle Massive particle Massless particle $\frac{2(4k_{10}^2+k_{11}^2+k_{11}^2+2k_{11}(k_{12}-2k_{11})-4k_{11}k_{12}+4k_{11}^2-2k_{10}k_{10}+4k_{10}^2-2k_{10}k_{10}-2k_{10}(2k_{11}+2k_{11}-4k_{10}+a_{11}))}{2k_{10}^2+k_{10}^2+2k_{1$ Pole residue: $\frac{1}{a} > 0$ Poleresidue 4 >0 Poleresidue (h, +h,)2 a, $\frac{a_{.}}{b_{.}}_{10} > 0$ Polarisations: 2 Square mass $\frac{a_{0}}{h_{10}+h_{0}} > 0$ Square mass Spin: Spin Parity Even Parity Unitarity conditions (Timeout after 20 seconds)

FIG. 5. The full spectrum of the general theory in Eq. (19). No general unitarity conditions are obtained, without further tuning. All the quantities in this output are defined in Fig. 4.

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Wave operator and propagator																								
$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	ZWs ¹ ak	$\frac{1}{12} + 4(\lambda_{2} + h_{3}) + 2\lambda_{3}(\lambda_{2}) = 0$	$\frac{44(x_{j}+h_{j}^{2}-2x_{j})a_{j}}{2}$ 0	$(r_{7}+r_{8}+r_{9})(r_{6}+12r_{0}^{2}) = 0$	*12	2																		
$ 2 \alpha_{1} \beta_{2} \alpha_{1} \beta_{2} \beta_{2} \beta_{2} \alpha_{2} \beta_{3} \beta_{4} \beta_{2} \beta_{4} \beta$		$2^{(n_{1}, n_{2}, n_{3}) + 4(n_{1}, n_{2}^{2} n_{1}^{2} n_{2}^{2} n_{3}^{2})} = \frac{2^{1} \sqrt{\frac{2}{3}} s t^{(n_{1}, n_{1}, n_{2}^{2}) - 4(n_{2}^{2} n_{3}^{2})}{2^{2} n_{1}^{2} n_{2}^{2}} = \frac{1}{2^{2} n_{1}^{2} n_{1}^{2}} = \frac{1}{2^{2} n_{1}^{2} n_{2}^{2}} = \frac{1}{2^{2} n_{1}^{2} n_{1}^{2}} = \frac{1}{2^{2} n_{1}^{2} n_$	$\frac{1^{2} \left(2 \lambda_{2} + \lambda_{1} + \lambda_{2} + \lambda_{2} + \lambda_{2} + \lambda_{2}^{2} + \lambda_{2}^{2} \right) \left(2 \lambda_{1} + \lambda_{2}^{2} + \lambda_{1}^{2} + \lambda_{2}^{2} +$	$\frac{1}{(h_1 + h_2)h_2} = \frac{1}{(h_2 + h_2)h_2} + \frac{1}{(h_1 + h_2)h_2} = \frac{1}{(h_1 + h_2)h_2} + \frac{1}{(h_1 + h_2)h_2}$		1.34 ₆ ¹⁰ .34 ₆ ¹⁰ .0	0 2 ² (a. 4 + 4 + 4 + 4 - 4 - 4 - 4 - 4 - 4 - 4 -	$\frac{1}{12}\sqrt{\frac{5}{2}} \left(2k^2(k_{12}^2 + k_{12}^2 + k_{21}^2) + a_2^2\right)\frac{1}{12}\sqrt{5}\left(k^2(k_{12}^2 + k_{12}^2 + k_{21}^2) + a_2^2\right)}$	$\frac{1}{12} \left(\frac{1}{12} \left(\frac{1}{12} + \frac{1}{7} + \frac{1}{9} \right) + \frac{1}{9} = \frac{1}{9} \left(\frac{1}{12} + \frac{1}{12} + \frac{1}{12} + \frac{1}{9} + \frac{1}{9} + \frac{1}{9} \right) = \frac{1}{12} \left(\frac{1}{12} + \frac{1}{1$	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	0 0	$\sum_{n=1}^{2} \frac{8! \sqrt{3} + (a_{2n} + a_{2n} + a_{$	$\frac{1}{10^{-1}} = \frac{2\sqrt{2}(2x)e^{-\frac{1}{2}x} + x_{1}\cdot 2x_{2})e^{-\frac{1}{2}x_{2}}}{3e^{-\frac{1}{2}x_{1}}}$	$\frac{2\sqrt{2}(\sqrt{2}\sqrt{2})}{3a_{2}} \frac{\sqrt{2}(\sqrt{2}\sqrt{2}\sqrt{2})}{3a_{2}} \frac{\sqrt{2}(\sqrt{2}\sqrt{2}\sqrt{2})}{3a_{2}} \frac{\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}}{3a_{2}} \frac{\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}\sqrt{2}2$	$\frac{1}{2} \frac{1}{2} + \frac{1}{2} + \frac{1}{2} $	¹¹⁻²⁰ م	$\frac{I P(T_{33} - 6_{31}, 6_{31}, 6_{32} + 3(x_{3} + x_{3}) + 7x_{3})}{4 + 4}$	4 $\frac{0}{2}$	$\frac{1}{2} \frac{2^{2/3} (5_{20} + 2^{2}_{20} + 2$	$\frac{1}{12} \left(k^2 \left(13 \frac{h}{10} - 8 \frac{h}{11} - 8 \frac{h}{12} + 7 \left(\frac{h}{7} + \frac{h}{8} \right) + 13 \frac{h}{9} - 3 \frac{h}{0} \right)$	$2^{2}\mathcal{B}_{5,40}^{4,0}$ $2^{2}\mathcal{B}_{5,40}^{4,0}$ $2^{2}\mathcal{B}_{5,40}^{4,0}$ 0	$\frac{e^{2}(2A_{23}+1)}{12A_{23}+1}\frac{e^{4}(b_{2}+1)+2A_{2}}{12\sqrt{2}}$ 0	$(4^{2}(h_{10}^{2}-h_{11}^{2}-2,h_{11}^{2}-2,h_{12}^{2}+4,(h_{1}^{2}+h_{3}^{2})+3,h_{3}^{2})=0$	0
الم من الأسان المحالية المح محالية محالية المحالية المحالي محالية محالية المحالية المحال	2.	$\frac{1}{2} e^{\frac{2}{3}} c^{\frac{2}{3}} t^{\frac{2}{3}} t^{\frac{2}{$	$\mu^{6}((s_{23}^{-1}+A_{12}^{-1})^{2}-4(A_{2}^{-1}+b_{0}^{-1})(A_{23}^{-1}+b_{0}^{-1})^{-4}$	$\frac{1}{4} + 2x_p^2 \frac{1}{60} \frac{\sqrt{2}}{\sqrt{2}} \frac{\sqrt{2}}{2} \frac{x^2}{2} \frac{(x^2}{(h_{11} + h_{12})^2 + (h_{21} + h_{22})^2 + (h_{22} + h_{22})^2}{3a^2} \frac{1}{4b^2}$	•	1.34 t ^{tt} 0	0	$\begin{array}{c c} & 1 & 0 \\ & 0 & 0 \\ & 0 & 0 \\ & 0 & 0 \\ & 0 & 0$	$\frac{1}{12}\sqrt{\frac{5}{2}}\left(-2k^{2}\left(h_{10}+h_{12}+h_{0}\right)+a_{0}\right)$	$\binom{1}{2}$, $\frac{1}{12}$, $\sqrt{5}$ (- k^2 (h_{11} - 2 h_{7} + 4 h_{1}) + a_{0})	0	$\frac{+2\alpha_{3}}{\sqrt{5}\alpha_{3}^{2}} = \frac{41(\alpha_{3}+\beta$	$\frac{e^2(4x_0^2+x_0^2)+e_0}{2} = \frac{e^2(4x_0^2-2(x_0+x_0^2-2x_0^2))+5}{3e_0^2}$	$(x_1) + 5a_0$ $(x_2) + 5a_0$ $(x_1) + 5a_0$ $(x_2) + (x_1 + x_2 + x_3 + x_3)$ $(x_2) + (x_2 + x_3 + x_3 + x_3)$ $(x_2) + (x_2 + x_3 + x_3 + x_3)$	$\frac{1}{32}$ $\frac{2}{3}$ $\frac{1}{3}$ $\frac{2}{3}$ $\frac{\sqrt{2}(-2)}{3}$ $\frac{1}{(2)}$ $\frac{1}{($	st 0 ⁺ 3 ¹	1 A ² (s ₂ + s ₄ + s ₂) 2 43	2 0	$= \frac{2}{3}k^2(h_{10} + h_{11} + h_{12} + h_{7} + h_{8} + h_{8}$	$\frac{2x^{2}(5x, +2x, -2x)}{2} \frac{2x^{2}(5x, +2x, -2x)}{12\sqrt{2}} + \frac{2x^{2}(5x, +2x)}{12\sqrt{2}} + \frac{2x^{2}}{3} + 2x$	$\frac{2^{2}}{3} \mathcal{G}_{g_{1}}^{g_{1}}$	$(x^{4})^{4}$ $(x^{4})^{4}$ $(x^{4})^{4}$ $(x^{4})^{5}$ $(x^{4})^{5}$ $(x^{4})^{6}$	$\frac{x^2(2s_{y_0}s_{y_1}+s_{y_1}+(s_{y_1}+s_{y_1}+2s_{y_1})}{12\sqrt{2}}$ $\frac{1}{12}$	0
$\begin{split} \prod_{i=1}^{n-1} \left[\left(\frac{1}{2} + \right) \right) \right) \right) \right) \right) \right) \right) \\ & + \left(\frac{1}{2} + \right) \right) \right) \right) \right) \right) \right) \right) \right) \\ & + \left(\frac{1}{2} + \right) + \left(\frac{1}{2} + \left(\frac{1}{2} + \left(\frac{1}$	$2^* \mathcal{T}^{\dagger}_{\alpha\beta}$	$= 2 x^4 (h_{11} + h_{12} + h_{13} + h$	$\ell (k^2 (t_{12}^{-1} + t_{12}^{-1})^2 - 4(t_{12}^{-1} + t_{22}^{-1})(t_{12}^{-1} + t_{22}^{-1})) + 4 k (t_{12}^{-1} + t_{22}^{-1} + t_{22}^{-1}) = 0$ $\ell (3 u_{22}^{-1} + t_{22}^{-1}) (t_{12}^{-1} + t_{22}^{-1}) + 4 k (t_{12}^{-1} + t_{22}^{-1} + t_{22}^{-1}) = 0$	$\frac{2I\sqrt{\frac{2}{2}}(\mu^2((e_{11}+h_{12})^2,4(h_{2}+h_{3})(h_{3}+h_{3}))x((2)_{10},2h_{12},2h_{12}+4(h_{2}+h_{3}))x((2)_{10},2h_{12}+4(h_{2}+h_{3}))x((2)_{10},2h_{12}+4(h_{2}+h_{3}))x((2)_{10},2h_{12}+2h_{12})x((2)_{10},2h_{12}$	0 4 01 0	$\frac{1^{2}\mathcal{G}_{2}^{-1}}{2} \frac{1^{2}\mathcal{G}_{2}^{-1}}{2} \frac{1^{2}\mathcal{H}_{1}}{2} \frac{1^{2}\mathcal{H}_{2}}{2} \frac{1^{2}\mathcal{G}_{2}^{-1}}{2} 1^$	0 0 0	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$0 \qquad 0 \qquad \frac{2x^2(y_{11} + y_{12} + y_{13} + y_{14} - y_{1$	$0 \qquad 0 \qquad \frac{1}{12} \left(k^2 \left(4 \dot{h}_{10} + \dot{h}_{12} - 2 \dot{h}_{1} \right) + c \right) = 0$	·γ	$0 \frac{4(2 e^{2} h_{0} $	$0 \frac{4r\lambda(2_{A_0} + 1, b_1 + 2, b_2)}{\sqrt{3} a_0^2} \frac{2r^2(7_{A_0} + 3(t_1 + b_1 + t_2))}{3 a_0^2}$	$0 \frac{4^{4} X[a_{0} + b_{12} + b_{12}$	$0 \frac{8! \sqrt{\frac{4}{3}} x \left(\frac{1}{2} x \left(\frac{1}{2} x \right)^{\frac{1}{3}} + \frac{1}{3} \frac{1}{2} x \left(\frac{1}{2} x \left(\frac{1}{2} x \right)^{\frac{1}{3}} + \frac{1}{3} \frac{1}{2} \right)^{\frac{1}{3}}}{a_{0}^{\frac{1}{3}}} \frac{2 \sqrt{2}(2 x^{\frac{1}{3}} (2 x_{0} + \frac{1}{3} +$	14,0 14,0 -4	$0 \frac{1}{4} (-2k^4 (h_{10} - h_{11} - h_{12} + h_{7} + h_{8} + h_{1}) + k^2 a_{0}) 0$	0	$\frac{t^{2}(h_{0}^{-1}, a_{0}^{-1}, a_{0}^{-1}, b_{0}^{-1})}{2\sqrt{3}} \xrightarrow{\frac{1}{2}}{2}$	$\frac{i k'(T_{00} + k_{0} + 5k_{0} + 5k_{0} + 5k_{0} + 7k_{0})}{4 \sqrt{6}} \frac{4}{4 \sqrt{5}}$	$\frac{2^{2}h_{1,0}}{8} = \frac{2^{2}h_{1,0}}{h_{1,1}} + \frac{2^{2}h_{1,0}}{h_{1,2}} + \frac{2^{2}h_{1,1}}{h_{1,2}} + \frac{2^{2}h_{1,1}}{h_{1,2}} + \frac{2^{2}h_{1,1}}{h_{1,2}} + \frac{2^{2}h_{1,1}}{h_{1,2}} + \frac{2^{2}h_{1,2}}{h_{1,2}} $	$\frac{i e^{2} (y_{10}^{-1}) + i e^{2y_{10}^{-1}}}{4 \sqrt{3}} = \frac{1}{6} (A^{2} (h) + i e^{2y_{10}^{-1}}) + i e^{2y_{10}^{-1}} (A^{2} (h) + i e^{2y_{10}^{-1}}) + i$	$i \frac{i}{2} \left(2 h_{0} - 3 h_{0} + 4 h_{0} + 2 h_{0} + 2$	0
о Ш		2*71 t ^{as}	W2 ¹⁺⁰⁸	w ¹⁺ t ^{as}	W _s ¹ † ^{all}	8 8 8	$1 h^{\pm} +^{\alpha}$	$\mathcal{B}_{s}^{,t,\dagger}$	98°14°	a,h+" کو	0+7-4	0 ⁺ 7 ⁻¹ †	rWs ^{±t} †	†1°wr‡	w₅ th †	0, ** +	0^41	°, t+	',,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,	ع _{د 1} 4 +	2*hl † 08	<i>S</i> r_1 + ∞	St, † ⁴⁰	H1-15
Massive and massless spectra			-2-	2*.	2	11			₩ ²	1			÷.	0.	.0			÷.	0.	-0- -		2+	N I	3
$\frac{1}{2} \frac{1}{\frac{1}{2}} \frac{1}{\frac$	1+ particle	? ? e	Polarisations: 2	Massless particle		J ⁰ = [₃₀ , 0, 0, 30]																		
Poleresidue: $\frac{21n_{10}h_{10}^{-1}h$	S:	->0 ->0 >0 																						
Unitarity conditions																								

FIG. 6. The results in Fig. 5 repeated in the second-order formulation. Note that the quadratic action in this case contains very many more operators than does Fig. 5. The matrix elements and the forms of the pole residues are expected to change, but the mass spectrum is the same. Once again, the unitarity conditions are not obtained for the full theory: despite the apparent changes to the residues, such conditions should be invariant under reparametrizations. All the quantities in this output are defined in Fig. 4.

PSA	LTe	er i	esu	lts p	ban	nel																						
Wa	ve o	pe	rato	or ai	nd	pro	pag	ato	r																			
	$^{1^+}W_{s}$	5 ¹ αβ	$_{\varphi}$ $^{1}\mathcal{T}_{a}^{\perp}$							${}^{1}\mathcal{W}_{s}{}^{it}{}_{a}$						${}^{1}W_{s}{}^{lt}{}_{a}$					${}^{1}\mathcal{W}_{s}{}^{h}{}_{\alpha}$							
${}^{1^+}\mathcal{W}_{s^{\perp}}$ †	β <u>4</u> a. 0				C)						0				0					0		0					
${}^{1}\mathcal{T}^{\perp}$ †	α 0		2 *2 (4 k ⁴ (hh.)) 7 8 (2+k ²	$\frac{(k_1+k_2)+k_1}{(k_1^2+k_2)^2} = \frac{2}{0} (k^2)$	$+8k^2(2k_1+\frac{7}{7})+3k_1^2(k_1-h_1)+3k_2^2(k_1-h_1)+3k_1^2(k$	* <u>;</u>)a,+3a ;,)	²)	21	$\sqrt{\frac{2}{3}} k (2 k^4) (h)$	$(2+k^2)$	$\frac{h_{7}+h_{8}-a_{0}}{a_{1}^{2}(k^{2}(h_{7}-k_{1}))}$	$3k^2(4(h_7+h_7)+3a_1)$.)-a.)a3	³ a, ²)	$\frac{i\sqrt{\frac{10}{3}}k(2k^2(h,-h)+3)}{(2+k^2)a_0(k^2(h,-h)+3)}$	<u>)</u>	2 <i>i k</i> (12	a,²+3 k² a √	$\frac{12}{3}(2+k^2)^2$	$\frac{4k_{+}+a_{+}+2k^{4}(h_{+}-h_{+})}{a_{0}^{2}(k^{2}(h_{+}-h_{+})+3a_{+})}$	$\frac{(4(k,+k,)+a,)}{7}$	(2	$\frac{4i\sqrt{\frac{2}{3}}k(k^2(h_1-h_2)-3a_1)}{(2+k^2)a_1(k^2(h_1-h_2)+3a_1)}$				
¹ We ^{⊥t} †	α 0	2	$1\sqrt{\frac{2}{3}}k(-2k^4)$	(h_7_8)(2(h.+h.)-a	.)+3a, ² +3	k ² a. (-4($k_{7}^{+k}_{8}^{+k}_{0}^{+a}_{0}^{+)$	-4(1	3 a, ² +2 k ² a,	(-4 h8)	h.+5a.)+k4	(-4 k ₇ ² +4 k ₇	a.+(-2 h.+	$+a_{0}^{(2)})$	$2\sqrt{5}(5 a) + k^2(2h) - 2h + 8$	a.)) 2	$\sqrt{2}(4 a_{0}^{2})$	+k ² a. (28	h,+20 h, +	$+a_{0}+k^{4}(8h_{7}^{2}-8h_{8}^{2})$	-2 h, a, +2 h, a, +a,	2))	$\frac{8(a.+k^2(kh.+2a.))}{\frac{0}{7}}$				
				(2+k2	$\frac{10}{2} k (2k^2)$	(h,-h,)+3a	·.)				3(2+ k) 2 √5(5	$(k^2)^{*} = (k^2)^{*} = (k^2$	-2 k. +a.))			$3(2+k^c)a_0(k^c(h_7,h_8)+3)$ 4 5	a.)			3($3(2+k^{-})a_{0}(k^{-}(h,-h_{0})+3a_{0})$ $4\sqrt{5}(k^{2}(h,-h_{0})+a_{0})$							
¹ W _s ^t †	α 0			(2+)	3 k ²) a. (k ²	(h,-h,)+3 a	<u>,</u>)				3(2+)	0 7 k ²) a. (k ² (h,	k.)+3a.)			$\overline{3a_0} = \overline{3k^2(h_7 - h_8) + 9a}$				3(2+ k ²)	-		$\frac{\frac{7}{78}}{\frac{3}{3}a_{0}\left(k^{2}\left(k,-h_{1}\right)+3}{\frac{7}{8}a_{0}}\right)}{\frac{7}{78}a_{0}}$					
¹ Ws ^{⊥h} †	α 0	-	2 <i>i k</i> (12 <i>a</i> . ² +	$\frac{-3 k^2 a_0}{\sqrt{3}}$ (12)	$\frac{h+4h}{7}$	$(k^2(h,-h,-))$	1#.)(4(/ 7 8 ·3a.)	7 8 0	2 √2 (4	$a_{0}^{2}+k^{2}a_{0}(2)$	$\frac{3k_1+20}{3(2+k^2)}$	$(k_{8}^{+}+a_{0}^{-})+k^{4}(k_{8}^{-})+k^{4}(k_{8}^{-})^{2}a_{0}^{-2}(k^{2}(k_{7}^{-}))$	8 h 2 8 h 2 2 h)+3 a)	h.a.+2h. 7 0 8	(a,+a, ²)	$\frac{\sqrt{10}(k^2(4k,-4k,-a))+4}{3(2+k^2)a(k^2(k,-k))+3}$	a.) a.)	2(32 a.	2+8k ² a.($\frac{10}{7} \frac{k}{2} + 2k}{8(2+k^2)^2}$	$(k^{2} (k^{2} (h, h)) + 3a_{0})^{2}$	$(h, a, -(4h, +a,)^2))$	- 8	$-\frac{8\sqrt{2}(5 a + k^{2}(-k + k + a))}{3(2 + k^{2})a_{0}(k^{2}(k - k) + 3 a_{0})}$				
¹ W _s ^{lh} †	α 0			41	$\sqrt{\frac{2}{3}} k (k^2)$	(hh.)-3 a.)				8(a	$k_{0} + k^{2} (h_{0} - h_{0} + h_{0}$	-2 a,))			$\frac{4\sqrt{5}(k^2(h,-h,)+a,)}{7}$				8 √2 (x	$\frac{2(h,-h,-a,)-5a,}{7800}$		4	$\frac{4}{3}\left(\frac{5}{a_1}-\frac{16}{b^2(b_1+b_1)+3a_1}\right)$				
1	1.5		i.	(2+)	k²) a. (k² 0	(A,-N,)+3a	.º	.0+		_	3(2+)	() a. (R= (A,-	*.)+3a.) 8 0			3 a. (k ⁻ (hh.)+3 a.) 0 7 8 0	2.9	.2+	2+	3(2+ k*) 4	1. (k" (hh.)+3 a.) 0 7 8 0	2 ⁺ 7 ⁻¹ -0 2 ⁺	Wela	2 ⁺ -2 ⁺ W ⁺ - 2 ⁺ W ⁺ -				
۳ _s lh ta	ras⊥h †α	99, ^{∥t} †″	.,, 74s±t+a	1, ⁵ , ⁴	l	₩s [™] t	₩s"+	t ^{sr} sw	0⁺7″†	0 ⁺ 7 ^{-⊥} †	S	==)))) (-)	14 (-2 a . F	$\mathcal{A}_{\alpha\chi\beta}\mathcal{A}^{\alpha}$	$a\beta \chi + 2 a$	$\mathcal{R}^{\alpha}{}^{\beta}{}_{\alpha}\mathcal{R}^{\chi}{}_{\beta\chi} +$	۲ ₅ " † ^{aßx}	77,-+~®	^{مه} +1 ⁵ لا	:+ hl + °°	2 ⁺ σ~1 + ^{αβ}	$8(h_{7} + h_{8} \frac{h_{1}}{h^{2}}) =$	4i	$\frac{8i\sqrt{\frac{2}{3}}}{\sqrt{\frac{2}{3}}}$	0			
0	0	0	0 0	ol 4 C	1+ 915	вİ	ы	8/ x()		4 x ² (6				4 34 2 a.	$h^{\alpha\beta} \partial_{\beta} \mathcal{R}$	$x + 4$ $y = h_{\alpha\beta} + \frac{x}{2} - 2a h^{\alpha\beta}$	0	1 k g	$\frac{ik_6}{4\sqrt{3}}$	0	2+ hla	a.2 0	(3 k a 0 8	k a.	0			
		414		5	αβ	41 \(\)2 1	101 k 2 a . + 3 k ²	$5k^2(n_7 + 4+k^2)^2$	0	$(h_{1}^{2})^{2}$	ì			∂_{χ}	$\mathcal{R}_{\alpha\beta}^{\chi} - a$	$h^{\alpha}_{\alpha} \partial_{\chi} \mathcal{R}^{\beta}_{\beta}^{\chi} +$, <u>1</u> (-1		[∞] ²⁺ W _s ¹ † ^{αβ}	- 1	3 a.	3a.0	0			
1 k a 4 1/6	1 4 6	615	1 * 9	0	$^{1}h^{2}a$	°.	0 ⁹	6 8 0 2 0		8 0 2				a. h'	$a_{\alpha} \partial_{\chi} \mathcal{R}^{\beta \chi}$	$_{\beta}$ -2 $h_{8} \partial_{\beta} \mathcal{A}_{\chi \delta}^{\delta}$		x ² (n	2 (h; +	4 2	$\mathcal{W}_{s}^{2^{+}}\mathcal{W}_{s}^{\perp}\dagger^{\alpha\beta}$	$\frac{8i\sqrt{\frac{2}{3}}}{ka_{0}}$	$\frac{4\sqrt{2}}{3a_{0}}$	$-\frac{4}{3a_{0}}$	0			
e la	-	4 6 1-1		0	-		5		4(2 k ² ()		2			∂ ^x .	$\mathcal{A}^{\alpha \beta}_{a} - 2$	$h_{\frac{1}{7}}\partial_{\chi}\mathcal{A}_{\beta}^{\delta}_{\delta}\partial^{\chi}\mathcal{A}_{\alpha}^{\alpha\beta} -$		v 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	. h.)-3 8	₩]o÷	$\mathcal{W}_{s}^{\alpha} = \frac{2}{2} \mathcal{W}_{s}^{\alpha} + \frac{\alpha \beta \chi}{2}$	0	0	0	$\frac{4}{a_{0}}$			
2 (-2 k	2 2 22	√5 (<i>k</i>	$\frac{1}{6}(-k^2)$	-	1.9		3 k a.	0	$\frac{7+h}{8}$ + a	0	<u>i</u>			2 h. i	$\partial_{\beta} \mathcal{A}^{ap \chi} \partial$	$\delta \mathcal{R}_{\alpha \chi}^{o} =$		h	<i>a</i> .)									
8 + • + °	17 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	² h. + 1	√6 7 -2 <i>a</i>	0 P	ls a			16(:01	81 8				4 h. i	∂ ^χ Я ^{α β} д	$\delta \mathcal{A}_{\alpha \chi}^{\delta} + \delta \mathcal{A}_{\alpha \omega}^{\delta} +$		$\frac{1}{3}k^{2}(h_{1})$.2+							
°.) □□□	- 12	°.)			-	8 √2	20 12 4, +3	$6 k^2 (n_7 + 3)$ $8(4 + k^2)^2$	0	$(6k^2(h_{7}^{2}))$ $(4+k^2)$	0+1			7 4 h.i	а д ^х Я ^{а в} д	$\delta \mathcal{R}^{\delta}_{\beta \chi}))[$	0	+ h.) 8	$\frac{(h_{7}+h_{8})}{3\sqrt{2}}$	1 * 5	Ars_ €							
² √5()	V 15	1 6 (-5 k	¹ / ₆ √5 (1		1 ² a	k ² α.	² а ² о		+h,)-a,) 2-a,2 0	-		t, x, y	, z] ď z	a y a	x d t		+ 9.014										
2 k ² h	-2 k ² h	2 h. +2	Y 6 ^	0	.α,s ^r β	0	0	20 3(4+ k	2	10/ 3(4+ k	₽ A	3.4	$W_{s_{\alpha\beta\chi}}$	3.74 I			4 ₀ !	0	0	0	" "FC.2							
+ a.)	»+ •	2 a.)	-a -	•				2) <u>a</u> . 3	40	²) ₄₀ 3	= 3°₩	s ^{II} † ^{αβχ}	- <u>2</u> a. 0	affx 3	ω.						*							
	1	<u>12</u>				0	0	$8\sqrt{2}$ $(4+k^2)a$	$\frac{4i\sqrt{2}}{ka_{0}}$	$\frac{W_{S}}{4t \sqrt{2} k}$ $\frac{4t \sqrt{2} k}{(4+k^{2})a}$	- -		r	⁴ ⁵ ⁶	-													
k ² π-2 6√2	(-k ² h	 5	2 k ² h	14.0	$1 \mathcal{R}_{s}$		0* <i>h</i> +	0* <i>h</i> I	^{0⁺} A _s ⊥t	^{0,+} <i>A</i> _s		0.	'As ^{⊥h}		Spin-par	ityform		Co	ovariant	form				м	ultiplicities			
1010	+ a.)	2 h. +	2 6	0.	а [.] Т	0 <u>.</u> + h+ +	0	0	0	$\frac{l k a}{4}$		ī	1 k a 8 √2	k	$k^{0^+}W_{s^{\pm t}}$	$+2i^{0^{+}}\mathcal{T}^{\perp} == 0$	2	" 02 .	$\partial_{\beta}\partial_{\alpha}T^{\alpha\beta}$	== ∂ _x ∂ _β ∂	$\partial_{\alpha}W^{\alpha b \chi}$			1				
	-	a.) 12			-	0.* <i>h</i> ∥†	0	0 0		$-\frac{i k a}{4 \sqrt{3}}$		-	5 <i>i k</i> a 8 √6	ĺ	2 A 3 W S	τ κ τ		02	$2 \partial_{\chi} \partial^{\chi} \partial_{\mu} \partial^{\chi} \partial_{\mu}$	$+ o_{\beta c}$ $_{\beta}T^{\alpha\beta} + \hat{c}$	$\partial_{\delta}\partial_{\chi}\partial_{\beta}\partial^{\alpha}W^{\beta\chi\ \delta}$							
1 12 (-2	- - 	√5(2	1 12 12		5	⁰⁺ <i>R</i> s ^{⊥t} †	0	0 0		4. 2		:	^a . 4 √2	1	Tota l exp	pected gauge gener	ators:							4				
k ² h ₇ -	* 2 <u>a</u>	k ² h.	4 16 (2 h. +	0 / k a	^π s ^h α	^{0.⁺} Яs [∥] †	$-\frac{1}{4}ik$	$\frac{a}{0} \frac{l k a}{4 \sqrt{3}}$	a. 2	$-\frac{2}{3}k^2(h_{-7})$	+ h.) 8	$\frac{2k^2(h)}{1}$	$(+h_1)+3a_1$ $(-2)\sqrt{2}$															
2.) 0	4	+ a.)	a.)			^{0,+} <i>Я</i> s ^{⊥h} †	$-\frac{ika}{2}$	$-\frac{5ika}{2\sqrt{6}}$	$\frac{a}{4\sqrt{2}}$	$\frac{2k^2(k_1+k_2)}{12k_1^2}$	+34.	$\frac{1}{12}$ (-7 k^2 ()	$h_{1} + h_{2} = \frac{1}{2}$	3 a.)														
							0.42	0 10		12 12																		
Mas	Siv	e a	nd r	nas	sle	SS S	pe	ctra																				
Parit	Squa	Pole																										
×.	re ma	-esidu	7		/			= (p, 0, 0, p)	/ ·																			
<u>0</u> H	:SSE	ë.	lassive	лн = (g	× -				$\langle \rangle$																			
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		د د ۷۵					Polaris	ations: 2	^a ; ⁻⁰ 2																			
Unit	Unitarity conditions																											
h.∈⊪ 7	$h_{.} \in \mathbb{R} \& \& a_{.} < 0 \& \& h_{.} < h_{.} \\ 7$																											

FIG. 7. The spectrum of the theory in Eq. (19) with the additional constraints $c_9 = c_{10} = c_{11} = c_{12} = 0$. These are sufficient to eliminate the 2⁺ and 1⁺ massive states, leaving only the massive 1⁻ state in Eq. (24). The overall theory is clearly unitary. All the quantities in this output are defined in Fig. 4.



FIG. 8. The results in Fig. 7 repeated in the second-order formulation. Note the much larger quadratic expansion, but the consistent mass spectrum and unitarity conditions. The results are used in Eq. (24). All the quantities in this output are defined in Fig. 4.

PSALTer results panel

Wave operator and propagator a.) $\sqrt{5}(-k^2(h_{12}-2h_7) \frac{1}{12}(k^2(4h_{12}-2h_7)$ $\frac{1}{12}(k^2(h_{12}-2h_{8})+$ -2+.)+4a Multiplicities Spin-parityform Covariantform 12 12 1 Sg_lh 1 k 8 4 √6 0 $k^{0^+} \mathcal{W}_{s}^{\perp t} + 2 i^{0^+} \mathcal{T}^{\perp} = 0$ 2 $\partial_{\beta}\partial_{\alpha}\mathcal{T}^{\alpha\beta} == \partial_{\chi}\partial_{\beta}\partial_{\alpha}\mathcal{W}^{\alpha\beta\chi}$ ج (پ $2 k \, {}^{1} \mathcal{W}_{s}{}^{\perp h^{\alpha}} + k \, {}^{1} \mathcal{W}_{s}{}^{\perp t^{\alpha}} + 6 \, i \, {}^{1} \mathcal{T}^{\perp \alpha} = 0 \, 2 \, \partial_{\chi} \partial_{\beta} \partial^{\alpha} \mathcal{T}^{\beta \chi} + \partial_{\delta} \partial^{\delta} \partial_{\chi} \partial_{\beta} \mathcal{W}^{\beta \alpha \chi} = 0 \, \mathcal{V}_{s} \partial_{\mu} \partial$ $2\,\partial_{\chi}\partial^{\chi}\partial_{\beta}\mathcal{T}^{\alpha\beta}+\partial_{\delta}\partial_{\chi}\partial_{\beta}\partial^{\alpha}\mathcal{W}^{\beta\chi\ \delta}$ -12 Total expected gauge generators: (; $0^+, \mathcal{T}^+$ $^{0^+}W_{s}^{\perp h}$ ${}^{0^+}\mathcal{T}^{\parallel}$ 0+W.I (; 0 $^{0^+}W_{c}^{\perp t}$ $\sqrt{\frac{5}{2}} (-2k^2(h_{12} + h_{13}) +$ $4k^2(6k^2(2h_{12}+h_{12}+h_{13})a$ $\frac{3}{12} k(6k^2 (2h_{12} + h_{12} + h_{12}) a$ $\frac{1}{12}(-k^2(h_{12}+h_{7})+$ 16 k²/ $\frac{10 i k}{3(4 + k^2) d}$ $\frac{4i\sqrt{2}k}{(4+k^2)a}$ k² (h. -2 h.)+4 a. $0^+ \tau^{-1}$ $2^{+}h^{\parallel}a$ ${}^{2^+}\mathcal{R}_{s}{}^{I}{}_{a\beta}$ ${}^{2^+}\mathcal{A}_{S^{\perp}\alpha\beta}$ ${}^{2}\mathcal{A}_{S}{}^{\dagger}{}_{\alpha\beta\chi}$ $\frac{2k^2(h, +h)}{12}$ $3(4+k^2)^2a$ $\sqrt{3}(4+k^2)a$ $3(4+k^2)^2a$ 1 A_s⁺h 1 k a 8 $\sqrt{3}$ 0 $-\frac{ika_0}{4\sqrt{3}}$ $-\frac{ika_0}{2\sqrt{6}}$ 16 k² h $4(2 k^2 (-2 h + h + h + h)) + a$ 321 k h. 0 0 21 √3 k a $41\sqrt{\frac{2}{3}}$ $2^{+}h^{\parallel} + a^{\circ}$ $^{0^+}\mathcal{T}^{\parallel}$ $\sqrt{3}(4+k^2)a$ k² a.² $\sqrt{3}(4+k^2)a$ k a. 0 Ika $k^{2}(h_{12}+2(h_{7}+h_{8}))$ $8ik(6k^2(2h_{12}+h_{7}+h_{8}))$ ².⁺*F*Is[∥]† $\frac{1}{6}(-k^2(2h_1+h_1+h_2)-3a_1)$ 0 $16(6 k^2 (2 h + h + h)) a$ $\frac{32 i k h}{\sqrt{3} (4 + k^2) a_0^2}$ $8 \sqrt{2}$ $3(4+k^2)a$ $\frac{20}{3(4+k^2)a}$ 4 √3 리워 ${}^{0^+}W_{\rm S}^{\pm 1}$ $3(4+k^2)^2 a$ $3(4+k^2)^2 a$ $k^{2}(h_{12}+2(h_{7}+h_{8}))$ $\frac{1}{3}k^2(h_1-h_2-h_1)+$ 0 . a.) ?* As a.) $\frac{1}{6} \left(-5 \, k^2 \left(h_{12} + h_{7} \right) + 2 \, a_{0} \right)$ $-\frac{10 l k}{12 a + 3 k^2 a}$ a.) $\frac{2i}{\sqrt{3}ka}$ $\frac{20}{12a_1+3k^2a_2}$ 2 √6 0+Ws 0 0 $(h_{12} + h_{.}) + h_{12}$ $\sqrt{5} (-k^2 (h_{12} - 2 h_7) +$ a. _0 _4 $\sqrt{5}(k^{2}(h_{12} + h_{B}) +$ 0 0 0 $-\frac{4i\sqrt{\frac{2}{3}}}{ka_{0}}$ $-\frac{4i\sqrt{2}k}{12a+3k^2a}$ $\frac{8 \sqrt{2}}{12 a_0 + 3 k^2 a_0}$ W.I 0 0 1 A₅⊪ _°₽ ${}^{2^+}\mathcal{T}^{I}{}_{a\beta}$ ${}^{2^+}W_{s}{}^{\parallel}_{\alpha\beta}$ ${}^{2^+}W_{s^{\perp}\alpha\beta}$ $^{2}W_{s}^{\parallel}_{\alpha}$ 0 (-2 k² $k^{2} + k^{2} \left(-2 k_{1} + k_{1} + k_{1} + k_{2} \right) a_{0} a_{0}$ $k^{2} a_{0}^{3}$ $k(2k^2h_{12}+a_{12})$ 1 $\frac{41\sqrt{\frac{2}{3}}(k^2h_{12}+2a_{12})}{ka_{12}^2}$ 0 2⁺TI+ $\frac{1}{12}\sqrt{\frac{5}{2}}$ ($\mathcal{W}_{\alpha\beta\chi}$ $+2\ h_{12}\ \partial^{\chi}\mathcal{R}^{\alpha\beta}\ \partial_{\delta}\mathcal{R}^{\alpha\beta}\ \beta_{\delta}\mathcal{R}^{\delta}\ \beta_{\beta}))$ $2h_{8}^{}\partial_{\beta}\mathcal{R}^{\alpha\beta\chi}\partial_{\delta}\mathcal{R}^{\delta}_{a\chi}+2h_{12}^{}\partial^{\chi}\mathcal{R}^{\alpha\beta}_{a}\partial_{\delta}\mathcal{R}^{\gamma}_{\beta\chi}$ $2h_{12}\partial_{\beta}\mathcal{A}^{\delta}_{\chi}\partial^{\chi}\mathcal{A}^{\alpha}_{a}^{\beta}$ - $2h_{7}\partial_{\chi}\mathcal{A}^{\delta}_{\beta}\partial^{\delta}\partial^{\chi}\mathcal{A}^{\alpha}_{a}$ $\partial_{\delta} \mathcal{R}^{\delta}_{\beta\chi} + 4 h_{8} \partial^{\chi} \mathcal{R}^{\alpha\beta}_{\alpha} \partial_{\delta} \mathcal{R}^{\delta}_{\beta\chi} + 4 h_{12}$ $4i(2k^2h_{12}+a_0)$ ہ ا 11 -<u>8</u> 3a $\frac{4\sqrt{2}}{3a}$ 0 $h^{\alpha}_{a} \partial_{\chi} \mathcal{R}^{\beta\chi}{}_{\beta} - 2 h_{\alpha} \partial_{\beta} \mathcal{R}^{\delta}{}_{\chi\delta} \partial^{\chi} \mathcal{R}^{\alpha}{}^{\beta}$ 2+W51+ A^{abx} $4h \cdot \partial_{\alpha} \mathcal{A}^{\alpha\beta\chi} \partial_{\delta} \mathcal{A}_{\beta\chi}^{\delta} + 4h \cdot \partial^{\chi} \mathcal{A}_{\alpha}^{\alpha\beta}$ √3 k a a.) 0 $\frac{1}{6}(-k^2(h_1+h_2)-2a_0)$ (·) 1ºM 8 $2a_{0}h^{\alpha\beta}\partial_{\chi}\mathcal{A}_{\alpha\beta}^{\chi} - a_{0}h^{\alpha}_{\alpha}\partial_{\chi}\mathcal{A}_{\beta}^{\beta\chi}$ "S $\sqrt{5}(k^{2}(h_{12} + h_{8}) +$ $h_{\alpha\beta}$ +2 a, $h^{\alpha\beta}$ $\partial_{\beta} \mathcal{A}_{\alpha x}^{X}$ $4i\sqrt{\frac{2}{3}}(k^2h_{12}+2a_0)$ $\frac{4\sqrt{2}}{3a}$ $-\frac{4}{3a_{0}}$ + $\frac{1}{12}(k^2(h_{12}-2h_{8})+$ 0 2+ Ws+ + k a. $\frac{2k^{2}(h_{12}+h_{7})+}{12\sqrt{2}}$ αßχ 1 Grs^⊥t 1 k a Ť۶, 4 a. 3 NV $^{2}\mathcal{W}_{s}^{\parallel}\dagger^{\circ}$ 0 0 0 3 G 5 z]dzdydxdt $\mathcal{A}^{a\beta\chi}$ +2 a, $\mathcal{A}^{a\beta}^{a\beta}$ or Fas as Be So S 0+ h+ 0+.A.1 0+.A.1h 0+h 0+,91,1 0 0 0 0 $\frac{i k a}{4}$ 0.+ h+ 8 √2 1 k 6 8 \(3) *i k* a 4 √6 010 $-\frac{ika}{4\sqrt{3}}$ 5*i k a* 1,4,0 i k 8 0 0<u>*</u>hI 0 8 √6 B ť, ×, 4. -0 2 $\frac{a_{0}}{4\sqrt{2}}$ ${}^0^{\scriptscriptstyle +}\!\mathscr{F}\!\!{}_{\!S}{}^{\scriptscriptstyle \perp t}$ 0 0 (-2 a 0 1⁺ A_S[⊥] 0 0 0 0 $2k^2(-4k_{12}+h_{12}+h_{1}+h_{1})+3a_{0}$ $\frac{i k a}{4 \sqrt{3}}$ $-\frac{1}{4}ika_0$ ^a. 2 $\frac{2}{3}k^2(2h_{12}+h_7+h_8)$ $^{0^+}\mathcal{R}_{c}^{-1}$ -14 12 √2 ${}^1\mathcal{A}_{\mathsf{S}}{}^{\mathtt{i}\mathsf{t}}{}^{\dagger}^a$ $^{1}\mathcal{A}_{s}^{lt}\uparrow^{a}$ $^{1}\mathcal{A}_{s}^{\perp h} \uparrow^{a}$ s == IIII 1 /1-1 1 \mathcal{A}_{s}^{lh} $2k^{2}(-4k,+h,+h,+h)+3a$ 51) $\frac{1}{12}(k^2(16h_{12}-7(h_1+h_2))-3a_{12})$ $^+S_{S^+}$ 0.+ As $\frac{0}{4\sqrt{2}}$ 8 1/2 8 JF Massive and massless spectra $J^P = 1$ $k^{\mu} = (p, 0, 0, p)$ Massless particle Massive particle $\frac{\frac{1152k_{12}^{2}-288k_{12}(h_{7}-k_{1})+(h_{7}-k_{3})(68k_{7}-68k_{7}-75a_{0})}{(h_{7}-k_{3})^{2}(2k_{7}-2k_{3}-3a_{0})}>0$ Pole residue: $\frac{1}{a} > 0$ Pole residue: Polarisations: 2 $\frac{\frac{3a}{0}}{\frac{h_{\gamma}}{\gamma}\frac{h_{\beta}}{8}} > 0$ Square mass Spin: Parity: Odd Unitarity conditions

FIG. 9. The spectrum of the theory in Eq. (19) with the additional constraints $c_9 = c_{10} = c_{11} - c_{12} = 0$. As with Fig. 7 these are sufficient to eliminate the 2⁺ and 1⁺ massive states, leaving only the massive 1⁻ state in Eq. (24). The overall theory is clearly unitary. All the quantities in this output are defined in Fig. 4.



FIG. 10. The results in Fig. 9 repeated in the second-order formulation. Note the much larger quadratic expansion, but the consistent mass spectrum and unitarity conditions. All the quantities in this output are defined in Fig. 4.

APPENDIX C: GENERIC CASE WITH PSALTER

As with Appendix B, our *PSALTer* analysis corresponding to Sec. III C is fully grounded in the MAG formulation. The generic MAG kinematic module is displayed in Fig. 11. The generic first-order MAG kinematic module is of course larger

PSA	LTer kinematic panel			
Momont	um Norm Frame			
Moment	um Norm Frame			
k ^p	$k^2 == k_\mu k^\mu \left n^\mu \right = \frac{1}{k}$			
Func	amental fields			
Fields S	ymmetries SO(3)		Sources	
h _{aβ} S	trongGenSet[{1,2},GenSet[(1,2)]] $\frac{1}{3} \eta_{\alpha\beta} \stackrel{0^+}{}h$	$^{ij} + ^{2*}h^{i}_{\alpha\beta} + ^{i}_{b}h^{i}_{\beta} n_{\alpha} + ^{i}_{b}h^{i}_{\alpha} n_{\beta} - \frac{1}{3} ^{0*}h^{i}_{\beta} n_{\alpha} n_{\beta} + ^{0*}h^{i}_{\alpha} n_{\alpha} n_{\beta}$	$\mathcal{T}_{a\beta}$	
Я _{αβχ} S	trongGenSet[{},GenSet[]] $\frac{4}{3} \mathcal{R}_a _{\beta\chi}$	$_{a} + \frac{1}{2} \left[\left[\left[\mathcal{A}_{b} \right]_{ab\chi} + \frac{1}{2} \left[\left[\left[\mathcal{A}_{b} \right]_{\alpha} \right]_{\alpha} + \frac{1}{2} \left[\left[\left[\mathcal{A}_{b} \right]_{\alpha} \right]_{\alpha} - \frac{1}{6} \left[\left[\left[\mathcal{A}_{b} \right]_{\alpha} \right]_{\alpha} - \frac{1}{4} \right] \left[\left[\left[\mathcal{A}_{b} \right]_{\alpha} \right]_{\alpha\beta} + \frac{1}{15} \left[\left[\left[\mathcal{A}_{b} \right]_{\alpha} \right]_{\alpha\beta} - \frac{1}{2} \left[\left[\left[\left[\mathcal{A}_{b} \right]_{\alpha} \right]_{\alpha\beta} - \frac{1}{2} \left[\left[\left[\mathcal{A}_{b} \right]_{\alpha} \right]_{\alpha\beta} - \frac{1}{2} \left[\left[\left[\left[\mathcal{A}_{b} \right]_{\alpha\beta} \right]_{\alpha\beta} - \frac{1}{2} \left[\left[\left[\left[\mathcal{A}_{b} \right]_{\alpha\beta} \right]_{\alpha\beta} \right]_{\alpha\beta} - \frac{1}{2} \left[$	W _{aβχ}	
	$1^+ \mathcal{R}_{a}^+ \beta_{c}$	$\chi n_{a} + \frac{1}{3} \left\{ i \frac{2}{\beta_{b}} \left\{ n_{a} + \frac{4}{3} \left\{ i \frac{2}{\beta_{b}} \left\{ n_{a} + \frac{4}{3} \left(i \frac{2}{\beta_{b}} \right) n_{a} + \frac{4}{3} \left(i \frac{2}{$		
		$ \begin{array}{c} u_{\alpha\chi} n_{\beta} = \frac{2}{3} \cdot 3 r_{5}^{*} - u_{\alpha\chi} n_{\beta} - \frac{1}{2} \cdot 3 r_{5} + \alpha_{\alpha} n_{\beta} + \frac{1}{5} \cdot 3 r_{5} + \alpha_{\alpha} n_{\beta} + \frac{1}{3} \cdot 3 r_{5} + \alpha_{\alpha} n_{\beta} n_{\beta} + \frac{1}{3} \cdot 3 r_{5} + \alpha_{\alpha} n_{\beta} + \frac{1}{3$		
	2 .5% 1 1.9%	$a\beta \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \$		
	3 *	hav 2 shav 2 sahv 12 sahv 2 sahv 2 ahv ahv ahv ahv	1	
SO (3) irreps			
SO(3)	Symmetries	Expansion		Sources
0*.**	StrongGenSet[{},GenSet[]]	$h_{\alpha\beta} n^{-} n^{\prime\prime}$		·. 𝒯⊥ 0⁺==1
· /·	StrongGenSet[{},GenSet[]]	$h_a^n - h_{\alpha\beta} n^n n^n$		×. <i>Υ</i> *
1 h ⁺ a	StrongGenSet[{},GenSet[]]	$h_{\alpha\beta} r \cdot h_{\beta\chi} n_{\alpha} r r$		$^{1}\mathcal{T}_{\alpha}^{1}$
$\frac{2^{+}h^{+}\alpha\beta}{0^{+}c^{-}}$	strongGenSet[{1,2},GenSet[(1,2)]]	$ \begin{array}{c} h_{\alpha\beta} \cdot \overline{j} \cdot h_{\alpha\beta} \cdot n'_{\lambda} + \overline{j} \cdot n'_{\lambda} \cdot n_{\alpha} \cdot n_{\beta} \cdot h_{\beta\lambda} \cdot n_{\alpha} \cdot n'_{\beta} + n'_{\beta} \cdot \overline{j} \cdot h_{\alpha\beta} \cdot h_{\lambda\delta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n_{\alpha} \cdot n'_{\beta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n_{\alpha} \cdot n'_{\beta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n'_{\alpha} \cdot n'_{\beta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n'_{\alpha} \cdot n'_{\beta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n'_{\alpha} \cdot n'_{\beta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n'_{\alpha} \cdot n'_{\beta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n'_{\alpha} \cdot n'_{\beta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n'_{\alpha} \cdot n'_{\beta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n'_{\alpha} \cdot n'_{\beta} \cdot n'' + \overline{j} \cdot h_{\lambda\delta} \cdot n'_{\alpha} \cdot n''_{\beta} \cdot n''_{\alpha} \cdot n''$		$\mathcal{T}^{\dagger}_{\alpha\beta}$
° SRat	StrongGenSet[{}, GenSet[]] -	$\frac{1}{2}\mathcal{A}^{\mu}{}^{\mu}{}^{\rho}{}^{\rho}{}_{\rho}{}^{+}\frac{1}{2}\mathcal{A}^{\mu}{}^{\rho}{}_{a}{}^{\rho}{}_{\rho}$		"Wal
	StrongGenSet[{},GenSet[]]	\mathcal{A}^{ept} $n_a n_{\beta} n_{\chi}$		W _s ^{II}
°. As∎	StrongGenSet[{},GenSet[]]	$\mathcal{R}^{a_{\beta}} n_{a} + \mathcal{R}^{a_{\beta}} n_{\beta} + \mathcal{R}^{a_{\beta}} n_{\beta} - \mathcal{R}^{a_{\beta}} n_{\beta} - \mathcal{R}^{a_{\beta}} n_{\chi}$		° W _s I
⁰ A _s th	StrongGenSet[{},GenSet[]]	$\mathcal{R}^{ab}{}_{\rho}{}_{n}{}_{a}{}^{-\frac{1}{2}}\mathcal{A}^{a}{}_{\rho}{}^{n}{}_{\rho}{}^{-\frac{1}{2}}\mathcal{A}^{ab}{}_{\alpha}{}^{n}{}_{\rho}$		0"Ws1h
⁰ <i>R</i> _a	StrongGenSet[{},GenSet[]]	$\mathcal{R}^{(b)} \epsilon \eta_{\alpha\beta\gamma\delta} n^{\delta}$		⁰ Wa ^{II}
$\mathcal{A}_{a\beta}^{1^{+}}\mathcal{A}_{a\beta}^{\dagger}$	StrongGenSet[{1,2},GenSet[-(1,2)]]	$\frac{1}{4}\mathcal{A}_{\alpha\beta}^{\ \ \ }n_{\chi}^{\ \ -\frac{1}{4}}\mathcal{A}_{\alpha\beta}^{\ \ \ }n_{\chi}^{\ \ -\frac{1}{4}}\mathcal{A}_{\beta\alpha}^{\ \ \ }n_{\chi}^{\ \ -\frac{1}{4}}\mathcal{A}_{\beta\alpha}^{\ \ \ }n_{\chi}^{\ \ -\frac{1}{4}}\mathcal{A}_{\beta\alpha}^{\ \ \ }n_{\chi}^{\ \ \ }n_{\chi}^{\ \ -\frac{1}{4}}\mathcal{A}_{\beta}^{\ \ \ }n_{\alpha}^{\ \ \ }n_{\chi}^{\ \ \ }n_{\delta}^{\ \ -\frac{1}{4}}\mathcal{A}_{\alpha}^{\ \ \ \ \ }n_{\beta}^{\ \ \ \ }n_{\delta}^{\ \ \ \ }n_{\delta}^{\ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ $		^{1*} W _∂ ^I _{αβ}
$\mathcal{A}_{a^{\dagger}a\beta}^{1^{+}}$	StrongGenSet[{1,2},GenSet[-(1,2)]]	$\frac{1}{2} \mathcal{R}^{\chi}_{\alpha\beta} n_{\chi} - \frac{1}{2} \mathcal{R}^{\chi}_{\beta\alpha} n_{\chi} + \frac{1}{2} \mathcal{R}^{\chi}_{\beta} n_{\alpha} n_{\chi} n_{\delta} - \frac{1}{2} \mathcal{R}^{\chi\delta}_{\beta} n_{\alpha} n_{\chi} n_{\delta} - \frac{1}{2} \mathcal{R}^{\chi\delta}_{\alpha} n_{\beta} n_{\chi} n_{\delta} + \frac{1}{2} \mathcal{R}^{\chi\delta}_{\alpha} n_{\delta} n$		$^{1^+}W_{a^+\alpha\beta}$
$\mathcal{A}_{s^{\perp}\alpha\beta}^{1^{*}}$	StrongGenSet[{1,2}, GenSet[-(1,2)]] -	$\frac{1}{2}\mathcal{A}_{\alpha\beta}^{\ \chi} n_{\chi} - \frac{1}{2}\mathcal{A}_{\alpha\beta}^{\ \chi} n_{\chi} + \frac{1}{2}\mathcal{A}_{\beta\alpha}^{\ \chi} n_{\chi} + \frac{1}{2}\mathcal{A}_{\beta}^{\ \chi} n_{\chi} - \mathcal{A}_{\beta}^{\ \chi\delta} n_{\alpha} n_{\chi} n_{\delta} + \frac{1}{2}\mathcal{A}_{\beta}^{\ \chi\delta} n_{\chi} n_{\delta} + \frac{1}{2}\mathcal{A}_{\delta}^{\ \chi\delta} n_{\chi} n_{\delta} n_{\chi} n_{\delta} + \frac{1}{2}\mathcal{A}_{\delta}^{\ \chi\delta} n_{\chi} n_{\delta} n_{\chi} n_{\delta} n_{\chi} n_{\delta} n_{\chi} n_{\delta} n_{\chi} n_{\delta} n_{\chi} n_{\delta} n_{\chi} n_{\chi} n_{\delta} n_{\chi} n_{\chi} n_{\delta} n_{\chi}		$^{1^{+}}W_{s^{\perp}a\beta}$
${}^{1}\mathcal{R}_{a}{}^{\parallel}_{a}$	StrongGenSet[{}, GenSet[]] -	$\frac{1}{2} \mathcal{A}^{\mu}_{\ \alpha\beta} + \frac{1}{2} \mathcal{A}^{\mu}_{\ \beta\alpha} - \frac{1}{2} \mathcal{A}^{\beta}_{\ \beta} x^{\ \alpha}_{\ \alpha} n_{\chi} + \frac{1}{2} \mathcal{A}^{\beta}_{\ \beta} n_{\alpha} n_{\chi} + \frac{1}{2} \mathcal{A}^{\beta}_{\ \alpha} n_{\beta} n_{\chi} - \frac{1}{2} \mathcal{A}^{\beta}_{\ \alpha} $		${}^{1}W_{a}{}^{\parallel}_{\alpha}$
${}^1\mathcal{R}_{a^{\perp}\alpha}$	StrongGenSet[{},GenSet[]]	$\frac{1}{2}\mathcal{A}_{a}^{\beta}x^{\alpha}\beta_{\beta}n_{\chi}-\frac{1}{2}\mathcal{A}_{a}^{\beta}n_{\alpha}n_{\chi}$		${}^{1}\mathcal{W}_{a}{}^{\perp}{}_{\alpha}$
$\mathcal{R}_{s^{\perp t}\alpha}$	StrongGenSet[{},GenSet[]]	$\mathcal{R}_{a}^{\beta\chi} n_{\beta} n_{\chi} + \mathcal{R}_{a}^{\beta\chi} n_{\beta} n_{\chi} + \mathcal{R}^{\beta\chi} n_{\beta} n_{\chi} - 3 \mathcal{R}^{\beta\chi\delta} n_{a} n_{\beta} n_{\chi} n_{\delta}$		$^{1}W_{s}^{I}_{\alpha}$
${}^{1}\mathcal{R}_{s}{}^{\mu}{}_{\alpha}$	StrongGenSet[{},GenSet[]]	$\mathcal{R}_{\alpha\beta}^{\ \beta} + \mathcal{R}_{\alpha\beta}^{\beta} + \mathcal{R}_{\alpha\beta}^{\beta} + \mathcal{R}_{\beta\alpha}^{\beta} - \mathcal{R}_{\beta\alpha}^{\beta\chi} n_{\alpha} n_{\beta} - \mathcal{R}_{\beta}^{\beta\chi} n_{\alpha} n_{\chi} - \mathcal{R}_{\beta\beta}^{\beta\chi} n_{\alpha} n_{\chi} - \mathcal{R}_{\alpha}^{\beta\chi} n_{\beta} n_{\chi} - \mathcal{R}_{\alpha}^{\beta\chi} n_{\beta} n_{\chi} - \mathcal{R}_{\alpha}^{\beta\chi} n_{\beta} n_{\chi} + \mathcal{R}_{\alpha}^{\beta\chi} n_{\beta} n_{\chi} - \mathcal{R}_{\alpha}^{\beta\chi} n_{\chi} - \mathcal{R}_{\alpha}$		$W_{s}^{\dagger}a$
${}^1\mathcal{R}_{\mathrm{S}}{}^{\mathrm{ih}}{}_{\alpha}$	StrongGenSet[{},GenSet[]]	$\mathcal{R}_a^{\ \beta c} \ n_\beta \ n_x \cdot \frac{1}{2} \ \mathcal{R}^{\beta c} \ n_\beta \ n_x \cdot \frac{1}{2} \ \mathcal{R}^{\beta c} \ n_\beta \ n_x$		$^{1}W_{s}^{\perp h}{}_{\alpha}$
${}^1\mathcal{R}_{\rm S}{}^{\rm Jh}{}_a$	StrongGenSet[{},GenSet[]]	$\mathcal{R}^{\theta}_{\alpha\beta} - \frac{1}{2} \mathcal{R}^{\theta}_{\alpha\beta} - \frac{1}{2} \mathcal{R}^{\theta}_{\beta\alpha} - \mathcal{R}^{\theta\chi}_{\chi} n_{\alpha} n_{\beta} + \frac{1}{2} \mathcal{R}^{\theta\chi}_{\beta} n_{\alpha} n_{\chi} + \frac{1}{2} \mathcal{R}^{\theta\chi}_{\beta} n_{\alpha} n_{\chi} - \mathcal{R}^{\theta\chi}_{\alpha} n_{\beta} n_{\chi} + \frac{1}{2} \mathcal{R}^{\theta\chi}_{\alpha} n_{\gamma} + \frac{1}{2} \mathcal{R}^{\theta\chi}_{\alpha} n_{\gamma} + \frac{1}{2} \mathcal{R}^{\varphi\chi}_{\alpha} n_{\gamma} + \frac{1}$		${}^{1}\mathcal{W}_{s}{}^{\parallel h}{}_{\alpha}$
$^{2^{+}}\mathcal{R}_{a}^{\dagger}{}_{\alpha\beta}$	StrongGenSet[{1, 2}, GenSet[(1,2)]] -	$\frac{1}{4} \mathcal{A}_{\alpha\beta}^{\kappa} n_{\kappa} + \frac{1}{4} \mathcal{A}_{\alpha\beta}^{\kappa} n_{\kappa} - \frac{1}{4} \mathcal{A}_{\beta\alpha}^{\kappa} n_{\kappa} + \frac{1}{4} \mathcal{A}_{\alpha}^{\kappa} n_{\kappa} n_{\delta} - \frac{1}{6} \mathcal{A}_{\kappa}^{\kappa\delta} \eta_{\alpha\beta} n_{\delta} - \frac{1}{6} \mathcal{A}_{\kappa}^{\kappa\delta} \eta_{\delta} n_{\delta} n_{\delta} n_{\delta} - \frac{1}{6} \mathcal{A}_{\kappa}^{\kappa\delta} \eta_{\delta} n_{\delta} n_{\delta} n_{\delta} - \frac{1}{6} \mathcal{A}_{\kappa}^{\kappa\delta} \eta_{\delta} n_{\delta} n_{\delta} n_{\delta} n_{\delta} - \frac{1}{6} \mathcal{A}_{\kappa}^{\kappa\delta} \eta_{\delta} n_{\delta} n_$		²⁺ Wa ¹ αβ
$^{2^{+}}\mathcal{R}_{s}^{\parallel}{}_{\alpha\beta}$	StrongGenSet[{1, 2}, GenSet[(1,2)]]	$\frac{1}{2}\mathcal{R}_{ab}^{\ \ x}n_{\chi} + \frac{1}{2}\mathcal{R}_{ab}^{\ \ x}n_{\chi} + \frac{1}{2}\mathcal{R}_{ba}^{\ \ x}n_{\chi} + \frac{1}{2}\mathcal{R}_{ba}^{\ \ x}n_{\chi} + \frac{1}{2}\mathcal{R}_{ab}^{\ \ x}n_{\chi$		2+Wsl ag
2+ 01 +	StrongGanSet[[1 2] GanSet[[1 2]]	$\frac{1}{3}\mathcal{R}^{\delta_{\lambda}}n_{a}n_{b}n_{b}-\mathcal{R}^{\delta_{\lambda}}n_{a}n_{\chi}n_{b}-\mathcal{R}^{\lambda_{b}}n_{a}n_{\chi}n_{b}-\mathcal{R}^{\lambda_{b}}n_{a}n_{\chi}n_{b}-\mathcal{R}^{\lambda_{b}}n_{a}n_{\chi}n_{b}-\mathcal{R}^{\lambda_{b}}n_{a}n_{\chi}n_{b}-\mathcal{R}^{\lambda_{b}}n_{a}n_{\chi}n_{b}-\mathcal{R}^{\lambda_{b}}n_{a}n_{\chi}n_{b}-\mathcal{R}^{\lambda_{b}}n_{a}n_{\chi}n_{b}-\mathcal{R}^{\lambda_{b}}n_{\chi}n_{\mu}n_{\mu}n_{\mu}n_{\mu}n_{\mu}n_{\mu}n_{\mu}n_{\mu$		21011
*. 94 _{s⁺αβ}	strongGenseu{1, 2}, Genseu(1,2)] -	$\frac{1}{4} \mathcal{M}_{\alpha\beta}^{\alpha} n_{x}^{\alpha} - \frac{1}{4} \mathcal{M}_{\alpha}^{\delta} n_{\mu} n_{\tau}^{\alpha} - \frac{1}{4} \mathcal{M}_{\beta}^{\delta} n_{\alpha} n_{x} n_{\delta} - \frac{1}{4} \mathcal{M}_{\beta}^{\delta} n_{\alpha} n_{x} n_{\delta} + \frac{1}{4} \mathcal{M}_{\alpha}^{\delta} n_{\alpha} n_{\alpha} n_{\alpha} n_{\delta} + \frac{1}{4} \mathcal{M}_{\alpha}^{\delta} n_{\alpha} n_{\alpha} n_{\alpha} n_{\alpha} n_{\alpha} + \frac{1}{4} \mathcal{M}_{\alpha}^{\delta} n_{\alpha} $		* W _s * _{αβ}
$^{2}\mathcal{R}_{a}^{\parallel}{}_{\alpha\beta\chi}$	StrongGenSet[{1, 2}, GenSet[-(1,2)]] -	$\frac{1}{8}\mathcal{R}_{\alpha\beta\chi} + \frac{1}{8}\mathcal{R}_{\alpha\beta} + \frac{1}{8}\mathcal{R}_{\beta\alpha\chi} - \frac{1}{8}\mathcal{R}_{\beta\chi\alpha} + \frac{1}{4}\mathcal{R}_{\chi\alpha\beta} - \frac{1}{4}\mathcal{R}_{\chi\beta\alpha} + \frac{3}{16}\mathcal{R}_{\beta\delta}^{\delta} n_{\alpha\chi} - \frac{3}{16}\mathcal{R}_{\delta\beta}^{\delta} n_{\alpha\chi} - \frac{3}{16}\mathcal{R}_{\delta\beta}^{\delta} n_{\beta\chi} + \frac{3}{16}\mathcal{R}_{\delta\alpha}^{\delta} n_{\beta\chi} - \frac{3}{16}\mathcal{R}_{\delta\alpha}^{\delta} n_{\beta\chi} + \frac{3}{16}\mathcal{R}_{\delta\alpha}^{\delta} n_{\beta\chi} - \frac{3}{16}\mathcal{R}_{\delta\beta}^{\delta} n_{\alpha\chi} + \frac{3}{16}\mathcal{R}_{\delta\beta}^{\delta} n_{\alpha\chi} + \frac{3}{16}\mathcal{R}_{\delta\beta}^{\delta} n_{\alpha\chi} + \frac{3}{16}\mathcal{R}_{\delta\beta}^{\delta} n_{\alpha\chi} + \frac{3}{16}\mathcal{R}_{\delta\chi}^{\delta} + \frac{3}{16$		${}^{2}W_{a}{}^{\parallel}_{\alpha\beta\chi}$
		$\frac{3}{16} \mathcal{A}^{\delta}_{\ \delta\beta} \ n_{\alpha} \ n_{\chi} + \frac{3}{16} \mathcal{A}^{\delta}_{\ \alpha\delta} \ n_{\beta} \ n_{\chi} - \frac{3}{16} \mathcal{A}^{\delta}_{\ \delta\alpha} \ n_{\beta} \ n_{\chi} + \frac{1}{8} \mathcal{B}_{\beta\chi}^{\ \delta} \ n_{\alpha} \ n_{\delta} - \frac{1}{8} \mathcal{B}_{\beta\chi}^{\ \delta} \ n_{\alpha} \ n_{\delta} - \frac{1}{4} \mathcal{B}_{\chi\beta}^{\ \delta} \ n_{\alpha} \ n_{\delta} - \frac{1}{4} \mathcal{B}_{\chi\beta}^{\ \delta} \ n_{\alpha} \ n_{\delta} + \frac{1}{8} \mathcal{B}_{\beta\chi}^{\ \delta} \ n_{\alpha} \ n_{\delta} - \frac{1}{8} \mathcal{B}_{\lambda\beta}^{\ \delta} \ n_{\alpha} \ n_{\delta} + \frac{1}{8} \mathcal{B}_{\lambda\beta}^{\ \delta} \ n_{\alpha} \ n_{\delta} - \frac{1}{8} \mathcal{B}_{\lambda\beta}^{\ \delta} \ n_{\alpha} \ n_{\delta} + \frac{1}{8} \mathcal{B}_{\lambda\beta}^{\ \delta} \ n_{\delta} \$		
		$\frac{1}{8}\mathcal{R}_{\alpha\chi}^{\delta}n_{\beta}n_{\delta} + \frac{1}{8}\mathcal{R}_{\alpha\chi}^{\delta}n_{\beta}n_{\delta} - \frac{1}{4}\mathcal{R}_{\chi\alpha}^{\delta}n_{\beta}n_{\delta} + \frac{1}{4}\mathcal{R}_{\chi\alpha}^{\delta}n_{\beta}n_{\delta} - \frac{1}{8}\mathcal{R}_{\chi\alpha}^{\delta}n_{\gamma}n_{\delta} + \frac{1}{8}\mathcal{R}_{\alpha\chi}^{\delta}n_{\gamma}n_{\delta} + \frac{1}{8}\mathcal{R}_{\alpha\beta}^{\delta}n_{\chi}n_{\delta} - \frac{1}{8}\mathcal{R}_{\alpha\beta}^{\delta}n_{\chi}n_{\delta} + \frac{1}{8}\mathcal{R}_{\alpha\beta}^{\delta}n_{\chi}n_{\delta} - \frac{1}{8}\mathcal{R}_{\alpha\beta}^{\delta}n_{\chi}n_{\delta} + \frac{1}{8}\mathcal{R}_{\alpha\beta}^{\delta}n_{\chi}n_{\delta} - \frac{1}{8}\mathcal{R}_{\alpha\beta}^{\delta}n_{\chi}n_{\delta} - \frac{1}{8}\mathcal{R}_{\alpha\beta}^{\delta}n_{\chi}n_{\delta} + \frac{1}{8}\mathcal{R}_{\alpha\beta}^{\delta}n_{\chi}n_{\delta} - \frac{1}{8}$		
		$\frac{1}{8}\mathcal{R}_{\beta a}^{\delta} n_{\chi} n_{\delta} - \frac{1}{4}\mathcal{R}_{\alpha \beta}^{\delta} n_{\chi} n_{\delta} + \frac{1}{4}\mathcal{R}_{\beta a}^{\delta} n_{\chi} n_{\delta} - \frac{3}{16}\mathcal{R}_{\delta}^{\delta c} \eta_{\beta \chi} n_{a} n_{e} + \frac{3}{16}\mathcal{R}_{\delta}^{\delta c} \eta_{\beta \chi} n_{a} n_{e} + \frac{3}{16}\mathcal{R}_{\delta}^{\delta c} \eta_{\alpha \chi} n_{\beta} n_{e} - \frac{3}{16}\mathcal{R}_{\delta}^{\delta c} \eta_{\alpha \chi} n_{\beta} n_{e} + \frac{3}{16}\mathcal{R}_{\delta}^{\delta c} \eta_{\alpha \chi} n_{\delta} n_{e} + \frac$		
		$\frac{1}{16}\mathcal{R}^{ab}_{\ \rho}\eta_{\alpha}\eta_{\delta}\eta_{c}+\frac{1}{16}\mathcal{R}^{ab}_{\ \sigma}\eta_{\delta}\eta_{\delta}\eta_{\delta}\eta_{c}+\frac{1}{16}\mathcal{R}^{bb}_{\ \alpha}\eta_{\beta}\eta_{\delta}\eta_{\delta}\eta_{c}+\frac{1}{16}\mathcal{R}^{bb}_{\ \rho}\eta_{\alpha}\eta_{\delta}\eta_{c}+\frac{1}{16}\mathcal{R}^{bb}_{\ \rho}\eta_{\alpha}\eta_{\delta}\eta_{c}+\frac{1}{16}\mathcal{R}^{bb}_{\ \sigma}\eta_{\beta}\eta_{c}\eta_{\delta}\eta_{c}+\frac{1}{16}\mathcal{R}^{bb}_{\ \sigma}\eta_{\beta}\eta_{c}\eta_{c}\eta_{c}\eta_{c}\eta_{c}\eta_{c}\eta_{c}\eta_{c$		
${}^{2}\mathcal{R}_{S}{}^{\dagger}{}_{\alpha\beta\chi}$	StrongGenSet[{1,2},GenSet[-(1,2)]]	$\frac{1}{3}\mathcal{A}_{\alpha\beta\gamma} + \frac{1}{3}\mathcal{A}_{\alpha\beta\gamma} - \frac{1}{3}\mathcal{A}_{\beta\alpha\gamma} - \frac{1}{3}\mathcal{A}_{\beta\gamma\gamma} + \frac{1}{3}\mathcal{A}_{\beta} - n_{\alpha\gamma} - \frac{1}{3}\mathcal{A}_{\beta\beta}^{*} + n_{\alpha\gamma} - \frac{1}{3}\mathcal{A}_{\alpha\beta}^{*} - n_{\beta\gamma} + \frac{1}{6}\mathcal{A}_{\alpha\gamma}^{*} - \frac{1}{3}\mathcal{A}_{\beta\gamma}^{*} - n_{\beta\gamma} + \frac{1}{6}\mathcal{A}_{\alpha\gamma}^{*} - \frac{1}{3}\mathcal{A}_{\beta\gamma}^{*} - n_{\alpha\gamma} + \frac{1}{6}\mathcal{A}_{\beta\gamma}^{*} - n_{\alpha\gamma} + 1$		² Ws ^I _{αβχ}
		$\frac{1}{5}\mathcal{R}_{\delta\beta}^{\prime}n_{a}n_{x}+\frac{1}{3}\mathcal{R}_{a\delta}^{\prime}n_{\beta}n_{x}+\frac{1}{5}\mathcal{R}_{a\delta}^{\prime}n_{\beta}n_{x}+\frac{1}{3}\mathcal{R}_{\delta\alpha}^{\prime}n_{n}n_{x}+\frac{1}{3}\mathcal{R}_{\beta\chi}^{\prime}n_{a}n_{\delta}+\frac{1}{3}\mathcal{R}_{\beta\chi}^{\prime}n_{a}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{a}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{a}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{a}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}n_{\delta}+\frac{1}{3}\mathcal{R}_{\delta\chi}^{\prime}n_{\delta}+1$		
		$\frac{1}{3}\mathcal{M}_{a\chi}^{\prime}n_{\beta}n_{\delta}+\frac{1}{3}\mathcal{M}_{a\chi}^{\prime}n_{\beta}n_{\delta}+\frac{1}{3}\mathcal{M}_{\chi\alpha}^{\prime}n_{\beta}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha}^{\prime}n_{\beta}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha\beta}^{\prime}n_{\chi}n_{\delta}+\frac{1}{3}\mathcal{M}_{\beta\alpha}^{\prime}n_{\chi}n_{\delta}+\frac{1}{3}\mathcal{M}_{\beta\alpha}^{\prime}n_{\chi}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha}^{\prime}n_{\delta}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha\beta}^{\prime}n_{\chi}n_{\delta}+\frac{1}{3}\mathcal{M}_{\beta\alpha}^{\prime}n_{\chi}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha}^{\prime}n_{\delta}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha}^{\prime}n_{\delta}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha\beta}^{\prime}n_{\chi}n_{\delta}+\frac{1}{3}\mathcal{M}_{\beta\alpha}^{\prime}n_{\chi}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha}^{\prime}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha}^{\prime}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha\beta}^{\prime}n_{\chi}n_{\delta}+\frac{1}{3}\mathcal{M}_{\beta\alpha}^{\prime}n_{\chi}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha}^{\prime}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha\beta}^{\prime}n_{\chi}n_{\delta}-\frac{1}{3}\mathcal{M}_{\alpha\beta}^{\prime}$		
		$\frac{1}{6} \mathcal{M}_{\delta} \mathcal{M} \mathcal{M}_{\delta} \mathcal{M} \mathcal{M}_{\delta} \mathcal{M} \mathcal{M}_{\delta} \mathcal{M} \mathcal{M} \mathcal{M}$		
3 cm 1	StrongGonSot[(1, 2, 2), GonSot[(1, 2), (2, 2)]]	$6^{(i)}a^{ij}k_{i}'a^{ij}\epsilon_{3}'a^{j}a^{i$		Saul
: As' αβχ	5goenset[11, 2, 3}, Genset[(1,2), (2,3)]]	$ \begin{bmatrix} \delta^{-1}a\beta c & \delta^{-1}a\gamma b & \delta^{-1}b\gamma c & $.'Ws' _{αβχ}
		$\frac{15}{12} = \frac{10}{7} \frac{15}{12} = \frac{10}{7} \frac{15}{12} = \frac{10}{7} \frac$		
		$\frac{1}{2} \mathcal{A}_{m} \stackrel{\delta}{=} n_{\mu} n_{x} - \frac{1}{2} \mathcal{A}_{m} \stackrel{\delta}{=} n_{\mu} n_{\mu} n_{x} - \frac{1}{2} \mathcal{A}_{m} \stackrel{\delta}{=} n_{\mu} n_{\mu} n_{\mu} n_{\mu} n_{\mu} - \frac{1}{2} \mathcal{A}_{m} n_{\mu} n_{\mu} n_{\mu} n_{\mu} n_{\mu} n_$		
		$\frac{1}{4}\mathcal{A}_{\beta\alpha}^{\beta}n_{\alpha}n_{\gamma}n_{\delta}\cdot\frac{1}{4}\mathcal{A}_{\beta\alpha}^{\delta}n_{\gamma}n_{\delta}\cdot\frac{1}{4}\mathcal{A}_{\beta\alpha}^{\delta}n_{\gamma}n_{\delta}n_{\delta}+\frac{1}{14}\mathcal{A}_{\alpha}^{\delta\epsilon}n_{\alpha}n_{\delta}n_{\delta}n_{\delta}n_{\delta}n_{\delta}n_{\delta}n_{\delta}n_{\delta$		
		$\frac{1}{15}\mathcal{A}^{6c}_{\delta}\eta_{\alpha\chi}\eta_{\beta}\eta_{e} + \frac{1}{15}\mathcal{A}^{\delta}_{\delta}\eta_{\alpha\beta}\eta_{\chi}\eta_{e} + \frac{1}{15}\mathcal{A}^{6c}_{\delta}\eta_{\alpha\beta}\eta_{\chi}\eta_{e} - \frac{1}{2}\mathcal{A}^{\delta}_{\delta}\eta_{\alpha}\eta_{\beta}\eta_{\chi}\eta_{e} - \frac{1}{2}\mathcal{A}^{6c}_{\kappa}\eta_{\alpha}\eta_{\beta}\eta_{\chi}\eta_{e} + \frac{1}{15}\mathcal{A}^{6c}_{\kappa}\eta_{\alpha}\eta_{\beta}\eta_{\chi}\eta_{e} + \frac{1}{15}\mathcal{A}^{6c}_{\kappa}\eta_{\alpha}\eta_{\beta}\eta_{\mu}\eta_{\mu}\eta_{\mu}\eta_{\mu}\eta_{\mu}\eta_{\mu}\eta_{\mu}\eta_{\mu$		
		$\frac{1}{15}\mathcal{A}^{\delta\epsilon}_{\chi}\eta_{\alpha\beta}\eta_{\delta}\eta_{\epsilon} + \frac{1}{15}\mathcal{A}^{\delta\epsilon}_{\beta}\eta_{\alpha\chi}\eta_{\delta}\eta_{\epsilon} + \frac{1}{15}\mathcal{A}^{\delta}_{\beta}\eta_{\alpha\chi}\eta_{\delta}\eta_{\epsilon} + \frac{1}{15}\mathcal{A}^{\delta\epsilon}_{\beta}\eta_{\alpha\chi}\eta_{\delta}\eta_{\epsilon} + \frac{1}{15}\mathcal{A}^{\delta\epsilon}_{\alpha}\eta_{\beta\chi}\eta_{\delta}\eta_{\epsilon} + \frac{1}{15}\mathcal{A}^{\delta\epsilon}_{\alpha}\eta_{\delta}\eta_{\epsilon} + \frac{1}{15}\mathcal{A}^{\delta\epsilon}_{\alpha}\eta_{\epsilon}\eta_{\epsilon} + \frac{1}{15}\mathcal{A}^{\delta\epsilon}_{\alpha}\eta_{\epsilon} + \frac$		
		$\frac{4}{15}\mathcal{A}_{\chi}^{\delta\epsilon} n_{\alpha} n_{\beta} n_{\delta} n_{\epsilon} + \frac{4}{15}\mathcal{A}_{\chi}^{\delta\epsilon} n_{\alpha} n_{\beta} n_{\delta} n_{\epsilon} + \frac{4}{15}\mathcal{A}_{\chi}^{\delta\epsilon} n_{\alpha} n_{\beta} n_{\delta} n_{\epsilon} + \frac{4}{15}\mathcal{A}_{\beta}^{\delta\epsilon} n_{\alpha} n_{\chi} n_{\delta} n_{\epsilon} + \frac{4}{15}\mathcal{A}_{\delta}^{\delta\epsilon} n_{\alpha} n_{\chi} n_{\delta} n_{\epsilon} + \frac{4}{15}\mathcal{A}_{\delta}^{\delta\epsilon} n_{\alpha} n_{\chi} n_{\delta} n_{\epsilon} + \frac{4}{15}\mathcal{A}_{\delta}^{\delta\epsilon} n_{\epsilon} n_{\chi} n_{\delta} n_{\epsilon} + \frac{4}{15}\mathcal{A}_{\delta}^{\delta\epsilon} n_{\epsilon} n_{\kappa} n_{\epsilon} $		
		$\frac{4}{15}\mathcal{A}^{\delta\varepsilon}_{a} n_{\beta} n_{\chi} n_{\delta} n_{\varepsilon} + \frac{4}{15}\mathcal{A}^{\delta\varepsilon}_{a} n_{\beta} n_{\chi} n_{\delta} n_{\varepsilon} \frac{1}{5}\mathcal{A}^{\delta\varepsilon\phi} \eta_{\beta\chi} n_{a} n_{\delta} n_{\varepsilon} n_{\phi} \frac{1}{5}\mathcal{A}^{\delta\varepsilon\phi} \eta_{\alpha\chi} n_{\beta} n_{\delta} n_{\varepsilon} n_{\phi} \frac{1}{5}\mathcal{A}^{\delta\varepsilon\phi} \eta_{\alpha\beta} n_{\chi} n_{\delta} n_{\varepsilon} n_{\delta} $		

FIG. 11. Kinematic structure of the unrestricted MAG, as studied throughout Sec. III C. The SO(3) irreps precisely correspond to those in Eqs. (5a) and (5b), though the labeling of duplicate J^P states is different from that in [28]. As with Figs. 2 and 4, the 2⁻ and 3⁻ states have extra cyclic symmetries which are hidden. These definitions are used in Figs. 12 and 13.

than the zero-torsion counterpart in Fig. 4. By comparing Eq. (18) with Eq. (5a) we see that the spin-zero and spin-one matrices will have extra rows and columns. As with Appendix B, the transition to the second-order formulation is made using Eq. (B1). The *PSALTer* labeling of duplicate J^P states is again different from that shown in Eqs. (5a) and (5b), and a new label (a) is introduced.



FIG. 12. The full spectrum of the general theory in Eq. (12) using the first-order formulation. As with Figs. 5 and 6 completely general unitarity conditions cannot be found automatically within one minute, but we investigate the conditions in Sec. III C based on the masses and residues in these results. All the quantities in this output are defined in Fig. 11.

The full spectrum of the general theory in Eq. (12) using the first-order formulation is given in Fig. 12. The equivalent result in the second-order formulation is given in Fig. 13.

Further theories considered in Sec. III C, whose matrices are too cumbersome for the appendices, are presented in [87].



FIG. 13. The full spectrum of the general theory in Eq. (12) using the second-order formulation. Note that the matrix elements differ from Fig. 12, and the form of the gauge symmetries and pole residues seem to differ, but the mass spectrum is the same. Moreover, the number of gauge generators is the same, and the pole residues are expected to have the same implications for the unitarity. Note also the much larger quadratic expansion of the action. All the quantities in this output are defined in Fig. 11.

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