

Quantum simulations for strong-field QED

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Quantum field theory in the presence of strong background fields contains interesting problems where quantum computers may someday provide a valuable computational resource. In the noisy intermediate-scale quantum era it is useful to consider simpler benchmark problems in order to develop feasible approaches, identify critical limitations of current hardware, and build new simulation tools. Here we perform quantum simulations of strong-field QED (SFQED) in $3 + 1$ dimensions, using real-time nonlinear Breit-Wheeler pair production as a prototypical process. The strong-field QED Hamiltonian is derived and truncated in the Furry-Volkov mode expansion, and the interactions relevant for Breit-Wheeler are transformed into a quantum circuit. Quantum simulations of a “null double slit” experiment are found to agree well with classical simulations following the application of various error mitigation strategies, including an asymmetric depolarization algorithm which we develop and adapt to the case of Trotterization with a time-dependent Hamiltonian. We also discuss longer-term goals for the quantum simulation of SFQED.

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I. INTRODUCTION

Quantum computing technology is rapidly developing. At present, real quantum devices do not offer useful advantage over classical computers, but in the not terribly distant future, quantum computing may mature into a valuable tool with diverse applications. In high energy physics (HEP), it is hoped that quantum computers will be able to simulate phenomena in models where classical simulation techniques are limited by sign problems or the need to explore a large Hilbert space [1–3]. In the near term, it is important to develop the tools to maximize the power of noisy intermediate-scale quantum (NISQ) devices on simulation problems, and to flesh out the space of target problems for which systematic improvements in quantum computing power would lead to clear advantage over classical simulations [4,5].

The Schwinger process, e^+e^- pair creation in $1 + 1$ dimensions in a strong background electric field [6], is a widely used benchmark for quantum simulations in HEP (see, e.g., [7–17].) As a nonperturbative, dynamical process in a confining gauge theory, it captures some of the essential physics of four-dimensional QCD-like theories

in a simpler setting more tractable for NISQ-era devices. The potential applications of quantum computing to QCD are well known [18–23]. Perhaps less widely appreciated are the potential applications of quantum simulation to QED, particularly in the presence of strong background electromagnetic fields (commonly referred to as strong-field quantum electrodynamics, or SFQED—for a recent review, see [24].) Indeed the Schwinger process [25–27] is a simple example of an SFQED phenomenon, but the richness and complexity of the theory arises in the presence of dynamical photons. Remarkably, although QED near the perturbative vacuum has been tested with unparalleled precision, there are still aspects of QED—most dramatically in the regime of ultrastrong fields—where the behavior is not fully understood (see, e.g., [24,28]). Qualitatively, ultrastrong fields should yield explosive particle production and backreaction, leading rapidly to a complicated quantum state. Thus, quantum computers may one day provide a unique probe of the most extreme regimes of QED. Similarly, quantum computers could provide a valuable tool to study phenomena in moderately intense fields. In this regime, which will be probed in upcoming experiments [29], analytic techniques are available for studying few-particle scattering amplitudes, at low loop order, in idealized background fields. As in QCD, quantum simulations could provide complementary access to real-time phenomena in more complex states.

These are future goals. Near-term quantum computers will be limited by size, connectivity, and noise. Therefore

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the present task is to design general simulation frameworks and test them on simple benchmark processes, tractable on present-day hardware.

In this paper we study real-time digital quantum simulations of $3 + 1$ D SFQED. Unlike the discretized-space approach generally used to study Hamiltonian lattice gauge theories [30,31], we use the discretized momentum-space SFQED Hamiltonian in the light-front formalism [32–34]. We perform quantum simulations by mapping Fock states onto qubits and ladder operators onto Pauli gates.¹ As a simple but nontrivial benchmark process, we examine nonlinear Breit-Wheeler pair production [38–45]. In this process, a photon “decays” into an electron-positron pair via an interaction with a strong background electromagnetic field. We study the real-time interference effects in pair-production probability due to two fields separated in time, a sort of “null double-slit” experiment [46–50].

Currently, quantum noise is unavoidable in simulations, and error mitigation is a powerful tool to “fit and subtract” various sources of noise. Fortunately, the past decade has seen rapid progress in the development of a suite of effective mitigation techniques [51–56]. Here we employ measurement mitigation, Pauli twirling [53], and depolarization mitigation. For depolarization mitigation, we improve on so-called “self-mitigation” [19,57,58] by relaxing the symmetric depolarization assumption. We find that at present these techniques are essential to obtain acceptable results. Without the combination of these mitigation techniques, our quantum simulations would yield unusable data.

This paper is organized as follows. In Sec. II we derive the momentum-space SFQED Hamiltonian and describe the truncation approach we use for quantum simulations. In Sec. III we translate the nonlinear Breit-Wheeler process into quantum circuits. In Sec. IV we describe the specific benchmark process of interest and develop a theoretical understanding of the relevant interference phenomena using time-dependent perturbation theory. We also perform real-time classical simulations in order to compare with quantum simulation results. In Sec. V we show raw quantum data and its comparison with Qiskit noisy simulations. In Sec. VI we describe the error mitigation strategies used to obtain the final mitigated results, which are found to be in good agreement with the classical simulations. Finally, in Sec. VII we describe some directions for future work.

A. Conventions

We work in natural units with $\hbar = c = 1$ and report values in MeV. $m = 0.511$ MeV is the mass of the electron and $e = 0.303$ is the electric charge.

It is convenient to work in light-front coordinates [33,59,60]. The coordinate four-vector is $x^\mu = (x^+, x^-, x^1, x^2) = (x^+, x^-, x^\perp)$, where $x^\pm \equiv x^0 \pm x^3$. Similar

notation holds for other four-vectors. The Minkowski metric and inverse are

$$g_{\mu\nu} = \begin{pmatrix} 0 & \frac{1}{2} & 0 & 0 \\ \frac{1}{2} & 0 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad g^{\mu\nu} = \begin{pmatrix} 0 & 2 & 0 & 0 \\ 2 & 0 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}. \quad (1)$$

To alleviate the subsequent bookkeeping of factors of two or one-half, we work mainly with raised indices, so that $x_\mp \rightarrow x_\mp^\pm$ and $x_i \rightarrow -x^i$ (for $i = 1, 2$). In other words, the covariant coordinate form is $x_\mu = (x_+, x_-, x_\perp) = (\frac{x^-}{2}, \frac{x^+}{2}, -x^\perp)$. An exception to this convention is with the four-gradient, which will be written as $\partial_\mu = (\partial_+, \partial_-, \partial_\perp)$.

We take x^+ to be the light-front time coordinate. The spatial three-vector and kinetic momentum are

$$\mathbf{x} = (x^-, x^1, x^2) \quad \mathbf{p} = (p^+, p^1, p^2). \quad (2)$$

We define their inner product as $\mathbf{p}\mathbf{x} = \frac{1}{2}p^+x^- - p^1x^1 - p^2x^2$. p^- is light-front energy, and free particles satisfy the dispersion relation

$$p^- = \frac{(p^\perp)^2 + m^2}{p^+}. \quad (3)$$

In the light front, p^+ and hence p^- are positive semi-definite. We will work in a Fock space constructed from a momentum lattice, and the zero mode $p^+ = 0$ does not propagate. Therefore we will take $p^+ > 0$.

Because of the conventions in Eq. (1), the integration measure on a slice of constant light-front time x^+ incurs a factor of one-half:

$$d^3\mathbf{x} = \frac{1}{2}dx^-dx^\perp. \quad (4)$$

We work with the chiral basis for the gamma matrices, so that with $\mu = 0, 1, 2, 3$

$$\gamma^\mu = \begin{pmatrix} 0 & \sigma^\mu \\ \bar{\sigma}^\mu & 0 \end{pmatrix}, \quad (5)$$

where $\sigma^\mu = (I, \vec{\sigma})$ and $\bar{\sigma}^\mu = (I, -\vec{\sigma})$. Light-front gamma matrices are defined as $\gamma^\pm = \gamma^0 \pm \gamma^3$. They obey the usual anticommutation relations with the light-front metric, $\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}$.

II. SFQED HAMILTONIAN

Given a quantum state $|\psi\rangle$, the generator of light-front time x^+ translations is $P_+, i\partial_+|\psi\rangle = P_+|\psi\rangle$. Equivalently, $\partial_+|\psi\rangle = -\frac{i}{2}P^-|\psi\rangle$ where we identify $P^- \equiv H$ as the

¹For additional discussion of quantum simulations in the light front, see Refs. [35–37].

Hamiltonian. The Hamiltonian density is given by the Legendre transform

$$\mathcal{H} = \frac{\partial \mathcal{L}}{\partial(\partial_+ \varphi_\alpha)} \partial_+ \varphi_\alpha - \mathcal{L}, \quad (6)$$

where \mathcal{L} is the Lagrangian density

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \bar{\psi} (i\not{D} - m)\psi. \quad (7)$$

Within the covariant derivative $D_\mu = \partial_\mu + ieA_\mu + ie\mathcal{A}_\mu$ we make explicit the electromagnetic gauge field A_μ and the classical² background field \mathcal{A}_μ . ψ is the Dirac field and the Dirac adjoint is defined as usual by $\bar{\psi} = \psi^\dagger \gamma^0$.

We work in the gauge $A^+ = 0$ for the photon field and assume the Lorenz gauge $\partial_\mu A^\mu = 0$ for the background field. For plane waves, \mathcal{A}_μ is a function of a light-front wave vector κ^μ : $\mathcal{A}_\mu = \mathcal{A}_\mu(\kappa_\mu x^\mu)$. If the background field is a null field propagating in the $-\hat{z}$ direction, $\kappa^\mu = (0, 2\omega, 0, 0)$, then the Lorenz gauge reduces to $\mathcal{A}^+ = 0$. Here ω is the frequency of the background wave.

Under these gauge conditions the Euler-Lagrange equations for A_μ and ψ reveal two constraint equations. For the electromagnetic field, we have

$$\partial_-^2 A^- = -\partial_- \partial_i A^i - \frac{1}{2} e \bar{\psi} \gamma^+ \psi. \quad (8)$$

For the Dirac field, we obtain the Dirac equation $(i\not{D} - m)\psi = 0$, but defining the projectors

$$\begin{aligned} \Lambda_+ &= \frac{1}{4} \gamma^+ \gamma^- = \text{diag}(1, 0, 0, 1) \\ \Lambda_- &= \frac{1}{4} \gamma^- \gamma^+ = \text{diag}(0, 1, 1, 0), \end{aligned} \quad (9)$$

and applying Λ_- from the left, we obtain the constraint

$$\partial_- \psi_+ = \frac{i}{4} \gamma^+ (i\gamma^i \partial_i + e\gamma^i A^i + e\gamma^i \mathcal{A}^i - m) \psi_-, \quad (10)$$

where $\psi_\pm = \Lambda_\pm \psi$ and $\psi_+ + \psi_- = \psi$. We will solve the constraints exactly for ψ_+ and A^- ,

$$\begin{aligned} \psi_+ &= \frac{i\gamma^+ (i\gamma^i \partial_i + e\gamma^i A^i + e\gamma^i \mathcal{A}^i - m) \psi_-}{4\partial_-} \\ A^- &= -\frac{\partial_- \partial_i A^i + e\psi_+^\dagger \psi_-}{\partial_-^2}. \end{aligned} \quad (11)$$

Note that $\gamma^0 \gamma^+ = 2\Lambda_-$ and $\Lambda_-^2 = \Lambda_-$, resulting in the appearance of ψ_- in the solution for A^- .

²Equation (7) lacks a kinetic term for the background field \mathcal{A}_μ because we consider cases where it obeys Maxwell's equations in vacuum.

In defining the Green functions appearing in Eq. (11), it is customary to take antisymmetric boundary conditions for fields at the longitudinal boundaries [32,34,61]. As a consequence the zero mode is omitted from the spectral decomposition.

Having fixed the gauge and solved the constraints exactly, the Hilbert space we construct is a physical one, with no additional non-gauge-invariant sectors. This approach minimizes the number of qubits needed to represent the degrees of freedom, at the expense of introducing additional interactions.

Inserting Eq. (11) into Eq. (6) we obtain a nonlocal expression for the Hamiltonian density in terms of the dynamical fields A^i and ψ_- . The full Hamiltonian density can be separated into several terms: the fermion energy

$$\mathcal{H}_\psi = -\frac{im^2}{2} \psi_-^\dagger \left(\frac{\psi_-}{\partial_-} \right) + \frac{i}{2} \psi_-^\dagger \left(\frac{\partial_i \partial_i \psi_-}{\partial_-} \right), \quad (12)$$

the photon energy

$$\mathcal{H}_A = \frac{1}{2} (\partial_1 A^2 - \partial_2 A^1)^2 - \frac{1}{2} (\partial_- \partial_i A^i) \left(\frac{\partial_i A^i}{\partial_-} \right), \quad (13)$$

the fermion-background energy

$$\mathcal{H}_{\psi\mathcal{A}} = e\mathcal{A}^- \psi_-^\dagger \psi_- + e\mathcal{A}^i \psi_-^\dagger \left(\frac{\partial_i \psi_-}{\partial_-} \right) - \frac{ie^2}{2} \psi_-^\dagger \mathcal{A}^i A^i \left(\frac{\psi_-}{\partial_-} \right), \quad (14)$$

the fermion-photon-background interaction

$$\mathcal{H}_{\psi A \mathcal{A}} = \frac{ie^2}{2} \psi_-^\dagger A^i \gamma^i \gamma^j \mathcal{A}^j \left(\frac{\psi_-}{\partial_-} \right) + \frac{ie^2}{2} \psi_-^\dagger \mathcal{A}^i \gamma^i \gamma^j \left(\frac{A^j \psi_-}{\partial_-} \right), \quad (15)$$

the fermion-photon interaction

$$\begin{aligned} \mathcal{H}_{\psi A} &= \frac{ime}{2} \psi_-^\dagger \gamma^i \left(\frac{A^i \psi_-}{\partial_-} \right) - \frac{ime}{2} \psi_-^\dagger \gamma^i A^i \left(\frac{\psi_-}{\partial_-} \right) \\ &\quad - \frac{e}{2} (\partial_- \partial_i A^i) \left(\frac{\psi_-^\dagger \psi_-}{\partial_-^2} \right) - \frac{e}{2} (\psi_-^\dagger \psi_-) \left(\frac{\partial_i A^i}{\partial_-} \right) \\ &\quad - \frac{e}{2} \psi_-^\dagger \gamma^i \gamma^j \left(\frac{\partial_i (A^j \psi_-)}{\partial_-} \right) - \frac{e}{2} \psi_-^\dagger A^i \gamma^i \gamma^j \left(\frac{\partial_j \psi_-}{\partial_-} \right), \end{aligned}$$

the four-fermion interaction

$$\mathcal{H}_{4\psi} = -\frac{e^2}{2} (\psi_-^\dagger \psi_-) \left(\frac{\psi_-^\dagger \psi_-}{\partial_-^2} \right), \quad (16)$$

and the double fermion-photon interaction

$$\mathcal{H}_{2\psi A} = \frac{ie^2}{2} \psi_-^\dagger A^i \gamma^i \gamma^j \left(\frac{A^j \psi_-}{\partial_-} \right). \quad (17)$$

A. Momentum-space Hamiltonian and discretization

Most simulations of quantum field theories on quantum computers use a real-space lattice discretization. We instead use a mode decomposition and construct the corresponding Fock space in the usual way. This approach has the advantage that it makes extracting physical quantities much easier with noisy, low-qubit-number devices. In terms of asymptotic scaling, the mode expansion results in a Hamiltonian with $O(n^4)$ terms, compared with $O(n^3)$ terms for a spatial lattice.

We define creation and annihilation operators a^\dagger/a , b^\dagger/b , and c^\dagger/c for the electron, positron, and photon, respectively. These respect the (anti)commutation relations

$$\begin{aligned} \{a_p^s, a_{p'}^{\dagger s'}\} &= (2\pi)^3 \delta_{s,r} \delta^3(\mathbf{p} - \mathbf{p}') \\ \{b_p^s, b_{p'}^{\dagger s'}\} &= (2\pi)^3 \delta_{s,r} \delta^3(\mathbf{p} - \mathbf{p}') \\ [c_k^j, c_{k'}^{\dagger j'}] &= (2\pi)^3 \delta_{j,j'} \delta^3(\mathbf{k} - \mathbf{k}'). \end{aligned} \quad (18)$$

Here, $s, r = \pm \frac{1}{2}$ are helicity indices for the fermions and $\lambda = 1, 2$ are polarization indices for the photon. The Fock states are labeled by occupation numbers and require an ordering, particularly for the fermions. We write the states in the following order:

$$|\dots, e^{-\frac{1}{2}}_{p_n}, \dots, e^{+\frac{1}{2}}_{p_n}, \dots, \gamma_{p_n}^1, \dots, e^{-\frac{1}{2}}_{p_n}, \dots, e^{+\frac{1}{2}}_{p_n}, \dots, \gamma_{p_n}^2, \dots\rangle, \quad (19)$$

where the superscript is the helicity/polarization, the subscript is the four-momentum, and n is an index for all allowed four-momenta for a given particle and helicity/polarization. (For a given three-momentum, p^- is determined by the dispersion relation.) We note that the fermion creation and annihilation operators induce a phase factor of $(-1)^\zeta$, where ζ is the sum of the number of fermions to the left of the operand as written in Eq. (19). The fermion states can be at most singly occupied, while the photon states can be arbitrarily highly occupied.

The mode expansion of the fields also requires helicity bispinors for the Dirac field and polarization vectors for the electromagnetic field. Since ψ_- is projected with Λ_- , it is convenient to define basis bispinors for each helicity as

$$w^{+\frac{1}{2}} = \begin{pmatrix} 0 \\ 1 \\ 0 \\ 0 \end{pmatrix} \quad w^{-\frac{1}{2}} = \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix}. \quad (20)$$

We define linear polarization four-vectors

$$\epsilon_1^\mu = (0, \epsilon_1^-, 1, 0) \quad \epsilon_2^\mu = (0, \epsilon_2^-, 0, 1). \quad (21)$$

In the free limit $e = 0$ the constraint equation determines $\epsilon_j^- = \frac{2k^j}{k^+}$, and the photon is transversely polarized, $\epsilon_j^\mu k_\mu = 0$.

With these definitions, the Schrödinger picture mode expansions are

$$\begin{aligned} \psi_- &= \int \frac{dp^+ dp^\perp}{(2\pi)^3} \sum_{s=\pm\frac{1}{2}} e^{-ipx} w^s a_p^s + e^{ipx} w^{-s} b_p^{s\dagger} \\ A^j &= \int \frac{dk^+ dk^\perp}{(2\pi)^3 \sqrt{k^+}} (e^{-ikx} c_k^j + e^{ikx} c_k^{j\dagger}). \end{aligned} \quad (22)$$

In this context, s is the positron helicity, and the electron has helicity $-s$. These fields can then be substituted into the Hamiltonian densities to obtain the Schrödinger-picture Hamiltonian:

$$H = \frac{1}{2} \int \mathcal{H} d^3\mathbf{x}. \quad (23)$$

Subsequent integrals may be evaluated by using

$$\int d^3\mathbf{x} e^{i(\mathbf{p}-\mathbf{p}')\cdot\mathbf{x}} = 2(2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}'). \quad (24)$$

After constructing the Hamiltonian in this way, one takes the additional step of normal ordering the creation and annihilation operators. (See Appendix A for sample calculations.) This procedure renormalizes away the simplest “ear diagram” divergences, and it clarifies that the Fock vacuum is the exact ground state in the light front.

Simulating the theory requires discretizing the momenta, truncating the set of single-particle states, and truncating the photon state occupation numbers. We use a momentum lattice with lattice spacing $\frac{2\pi}{L}$. The creation and annihilation operators then obey discrete (anti)commutation relations:

$$\{a_p^s, a_{p'}^{\dagger s'}\} = \delta_{s,r} \delta_{\mathbf{p},\mathbf{p}'} \quad \{b_p^s, b_{p'}^{\dagger s'}\} = \delta_{s,r} \delta_{\mathbf{p},\mathbf{p}'} \quad [c_k^j, c_{k'}^{\dagger j'}] = \delta_{j,j'} \delta_{\mathbf{k},\mathbf{k}'}. \quad (25)$$

As a result, they are scaled by a factor of $\sqrt{L^3}$ relative to the infinite volume case. For example, $a_p^s \rightarrow \sqrt{L^3} a_p^s$ in passing from the continuum to discretized single-particle state space.

B. Interaction picture

The SFQED Schrödinger-picture Hamiltonian can be split into a “free” piece, quadratic in the fluctuating fields, and an “interaction” piece, accounting for interactions of order e and e^2 . The free Hamiltonian $H_0 = H_\psi + H_A + H_{\psi A}$ is composed of number operators, so it is diagonal in the Fock basis, whereas the interaction Hamiltonian $V = H_{\psi A A} + H_{\psi A} + H_{A\psi} + H_{2\psi A}$ is not. As a result it is convenient to work in an interaction picture. Since even the Schrödinger-picture Hamiltonian is time dependent in the present context, let us review the derivation of the interaction picture time evolution operator.

The Schrödinger-picture time evolution operator is

$$U = \mathcal{T} e^{-\frac{i}{2} \int_0^{x^+} H_s(y^+) dy^+}, \quad (26)$$

where $H_S(x^+)$ is the complete time-dependent Schrödinger-picture Hamiltonian. The lower limit of integration is set to zero but in general denotes some initial fiducial time at which the Schrödinger-, interaction-, and Heisenberg-picture states are the same. Schrödinger-picture states are evolved with U , $|\psi_S(x^+)\rangle = U|\psi(0)\rangle$, and satisfy the Schrödinger equation with H_S . Now define the free evolution operator

$$U_0 = \mathcal{T} e^{-\frac{i}{2} \int_0^{x^+} H_0(y^+) dy^+} \quad (27)$$

and the interaction picture states

$$\begin{aligned} |\psi_{\text{int}}(x^+)\rangle &= U_0^\dagger |\psi_S(x^+)\rangle \\ &= U_0^\dagger U |\psi(0)\rangle \\ &= U_{\text{int}}(x^+) |\psi(0)\rangle. \end{aligned} \quad (28)$$

In the last line we have defined the interaction-picture time evolution operator $U_{\text{int}} = U_0^\dagger U$. It can be shown to satisfy

$$\begin{aligned} i\partial_+ U_{\text{int}} &= \frac{1}{2} (U_0^\dagger V U_0) U_{\text{int}} \\ &= \frac{1}{2} H_{\text{int}} U_{\text{int}}, \end{aligned} \quad (29)$$

where the second line defines the interaction-picture interaction Hamiltonian H_{int} . (Essentially it amounts to inserting the Volkov mode solutions [62] into the Schrödinger-picture interaction Hamiltonian V —see Appendix A.) The solution to Eq. (29) is

$$U_{\text{int}} = \mathcal{T} e^{-\frac{i}{2} \int_0^{x^+} H_{\text{int}}(y^+) dy^+}. \quad (30)$$

The Schrödinger picture time evolution operator can be written as $U = U_0 U_{\text{int}}$, where U_0 is entirely diagonal and amounts to phases when acting on free-particle basis states, so the typical probability between these states can be calculated just using U_{int} .

III. QUANTUM CIRCUIT DESIGN

There are many ways to build the same unitary as a quantum circuit, but in the NISQ era we must be careful to minimize the use of noisy gates, particularly CNOTs and SWAPs. At present we cannot afford to simulate the most scientifically interesting SFQED processes with many degrees of freedom. However, as a first step toward the long-term goal and as a proof of principle, we will show that

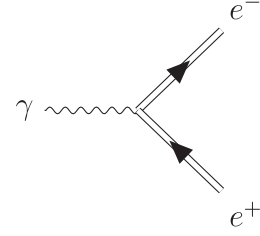


FIG. 1. Nonlinear Breit-Wheeler pair production. Double lines indicate e^+e^- propagation on the background field. This “decay” proceeds at tree level in SFQED in the Furry expansion [63] [see Eq. (A19)]. Applying the Jordan-Wigner transformation to represent three one-particle states by qubits, we translate the time evolution operator into a quantum circuit.

a simple example process, a tree-level three-body process with exclusive final states, is reliably simulated. In this section we build three-qubit circuits that optimally simulate the parts of the SFQED Hamiltonian responsible for the tree-level nonlinear Breit-Wheeler process, shown in Fig. 1.

We encode each particle in a qubit q . For the processes considered here it is sufficient to truncate boson occupation numbers at one. If q is in the $|0\rangle$ state, then the particle state is unoccupied; if q is in the $|1\rangle$ state, the particle state is occupied. For a truncated Hilbert space built from one single-particle state each for an electron, positron, and photon, the encoding is given by a bitstring of three qubits, such as $|001\rangle$, written as a Fock state according to Eq. (19). Thus the state $|001\rangle$ is a single photon and $|110\rangle$ is an electron-positron pair. We follow Qiskit’s little-endian notation, so that the qubits are numbered $|q_2, q_1, q_0\rangle$. That is, the photon is qubit zero, q_0 .

The terms in the Hamiltonian that describe the transition between these states are those with the operators $a^\dagger b^\dagger c$ and $c^\dagger b a$. We transcribe these two operator monomials into quantum gates with the Jordan-Wigner transformation. For example,

$$b = \frac{1}{2} (Z_2)(X_1 + iY_1) \quad c^\dagger = \frac{1}{2} (X_0 - iY_0), \quad (31)$$

where X_i , Y_i , and Z_i are the Pauli operators acting on the i th qubit. Z gates account for the fermion ζ factor. By using Pauli identities, we can write the monomials as a linear combination of Pauli strings, each written purely with X and Y gates. In the case of three qubits, there are $2^3 = 8$ different terms in the linear combination, since each qubit can be acted on with either an X gate or a Y gate. The resulting truncated Hamiltonian for three qubits is

$$\begin{aligned} H &= -\frac{2em}{8p^+ \sqrt{k^+ L^3}} e^{-ip^- x^+} e^{-2if(p)} [XXX - iXXY + iXYX + XYY + iYXX + YXY - YYX + iYYY] \\ &\quad -\frac{2em}{8p^+ \sqrt{k^+ L^3}} e^{ip^- x^+} e^{2if(p)} [XXX + iXXY - iXYX + XYY - iYXX + YXY - YYX - iYYY]. \end{aligned} \quad (32)$$

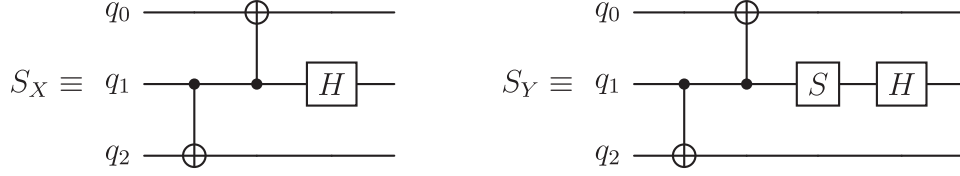


FIG. 2. GHZ diagonalization circuits for Pauli strings with an odd number of Xs (left) or odd number of Ys (right). Diagonalization facilitates implementation of the time evolution subcircuits with minimal gate counts.

Here, $f(p)$ is a function of light-front time defined in Eq. (A16), and p^μ and k^μ are the electron and photon momenta, respectively. See Appendix A for a derivation of the Hamiltonian. The upper set of operators corresponds to $c^\dagger ba$ while the lower set corresponds to $a^\dagger b^\dagger c$. In the following discussion we will write

$$a^\dagger b^\dagger c \sim c_{XXX}XXX + c_{XXY}XXY + c_{XYX}XYX + c_{XYY}XYY \\ + c_{YXX}YXX + c_{YYX}YYX + c_{YYY}YYY \quad (33)$$

and similar for $c^\dagger ba$.

We emphasize that because the interactions in the Hamiltonian are limited to three- and four-body, the Pauli string scaling is not exponential but rather polynomial in the number of qubits. (The nonlocality in momentum space increases the connectivity and therefore the power of the polynomial; however, by solving the constraints explicitly, even a position-space lattice implementation would have higher connectivity.)

Although the true time evolution operator is given in Eq. (30), this operator must be discretized as well to digitally simulate time evolution [64,65]. The typical prescription is to use the Lie-Trotter approximation, also referred to as first-order Trotterization. Although higher order Trotterizations such as the second-order Suzuki-Trotter formula lead to better accuracy for longer time steps, the circuit length is increased per time step, leading to more quantum error. It is a problem-dependent question whether it is worthwhile to implement the second-order formula, and in our case, we found that first order is preferable. Additionally, we wish to see fine-grained details not found using long time steps. Therefore we write

$$U_{\text{int}}(x^+) \approx e^{-\frac{i}{2}H_{\text{int}}(x_n^+)\Delta x^+} \dots e^{-\frac{i}{2}H_{\text{int}}(x_1^+)\Delta x^+}, \quad (34)$$

where n denotes the number of time steps used in the approximation, so $\frac{x^+}{n} = \Delta x^+$ and $x_n^+ = n\Delta x^+$.

As a result, we must build a quantum circuit representing the matrix exponential of Eq. (33). Here again we opt for a first-order as opposed to second-order Trotterization for the same reasons discussed above, so that

$$e^{-\frac{i}{2}\Delta x^+ a^\dagger b^\dagger c} \sim e^{-\frac{i}{2}\Delta x^+ c_{XXX}XXX} e^{-\frac{i}{2}\Delta x^+ c_{XXY}XXY} e^{-\frac{i}{2}\Delta x^+ c_{XYX}XYX} \\ \times e^{-\frac{i}{2}\Delta x^+ c_{XYY}XYY} e^{-\frac{i}{2}\Delta x^+ c_{YXX}YXX} e^{-\frac{i}{2}\Delta x^+ c_{YYX}YYX} \\ \times e^{-\frac{i}{2}\Delta x^+ c_{YYY}YYY}. \quad (35)$$

Each exponential factor now forms a subcircuit. Importantly, the order of these factors does not matter in the approximation, and it is very useful to shift certain subcircuits around and apply quantum gate identities to minimize the number of gates, particularly CNOTs.

There are numerous ways to simulate each subcircuit. We employ the GHZ diagonalization used in previous literature [19,66]. Subcircuits for Pauli strings with an odd number of X (Y) gates can be diagonalized with a GHZ transformation circuit. The transformation circuits are shown in Fig. 2 and the diagonalization relations are shown in Table I. We note that the qubit that is acted on by X- or Y-basis transformation gates is the qubit that will always have a Z gate upon diagonalization. Therefore, to reduce the use of long-ranged CNOT gates, it is ideal to place these transformation gates at a middle qubit, making the middle qubit the control qubit.

Since S_X and S_Y are unitary, we may write $e^{S_X^\dagger A S_X} = S_X^\dagger e^A S_X$. Then by rearranging the matrix exponentials in Eq. (35), we can write the Trotterization with one half diagonalized by S_X and the other half diagonalized by S_Y . This way Eq. (35) becomes

$$e^{-\frac{i}{2}\Delta x^+ a^\dagger b^\dagger c} \sim S_X^\dagger e^{-\frac{i}{2}\Delta x^+ c_{XXX}IZI} e^{\frac{i}{2}\Delta x^+ c_{YXX}ZZI} e^{\frac{i}{2}\Delta x^+ c_{YXY}ZZZ} \\ \times e^{\frac{i}{2}\Delta x^+ c_{XYY}IZZ} S_X S_Y^\dagger e^{-\frac{i}{2}\Delta x^+ c_{YYY}ZZZ} e^{\frac{i}{2}\Delta x^+ c_{XXY}IZZ} \\ \times e^{\frac{i}{2}\Delta x^+ c_{XYX}IZI} e^{\frac{i}{2}\Delta x^+ c_{YXX}ZZI} S_Y. \quad (36)$$

TABLE I. Diagonalizations of Pauli strings consisting of X and Y operators by the operators of Fig. 2.

Odd Xs	$S_X(\text{string})S_X^\dagger$	Odd Ys	$S_Y(\text{string})S_Y^\dagger$
XXX	IZI	YYY	ZZZ
YYX	-ZZI	XXY	-IZZ
YXY	-ZZZ	YXX	-IZI
XYY	-IZZ	YXX	-ZZI

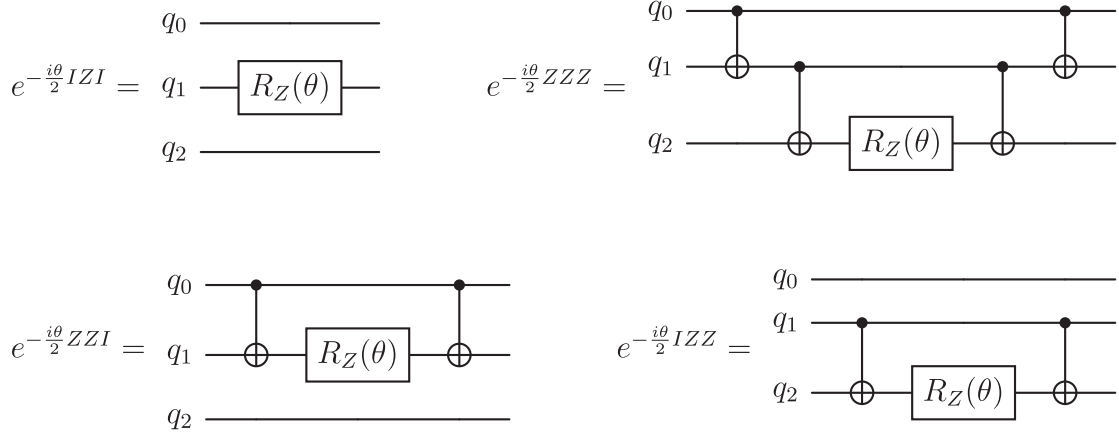


FIG. 3. Time evolution subcircuits after diagonalization. In Qiskit's convention $R_Z(\theta) \equiv e^{-\frac{i\theta}{2}Z}$. The CNOT gates in each subcircuit compute/uncompute the parity of the incoming bitstring.

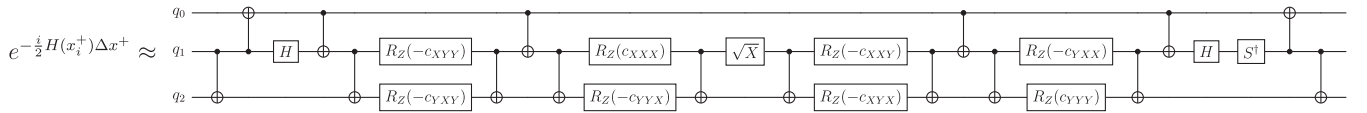


FIG. 4. First-order Trotterization circuit for a single time evolution step. By alternating the S_X and S_Y halves of the circuit in subsequent time steps, cancellations can be used to minimize the gate count. The first step has 16 CNOT gates, but only 12 CNOT gates are added with each succeeding step.

Each matrix exponential can now be simulated solely with CNOTs and R_Z gates. The four relevant circuits are shown in Fig. 3 and the final circuit is shown in Fig. 4.

In practice, IBM quantum computers use a limited set of native quantum gates: X , $R_Z(\theta)$, CNOT, I , and \sqrt{X} . Thus, the final step in circuit creation is to translate the circuit into these gates. For example, the Hadamard gate becomes $H = R_Z(\frac{\pi}{2})\sqrt{X}R_Z(\frac{\pi}{2})$ with an overall phase of $e^{i\frac{\pi}{4}}$, which falls out of the probability.

IV. NONLINEAR BREIT-WHEELER AND THE NULL DOUBLE SLIT EXPERIMENT

In the previous section we constructed a gate implementation for a truncation of the SFQED Hamiltonian. The truncation was drastic, but designed to capture the tree-level Breit-Wheeler interaction of Fig. 1 with fixed arbitrary initial and final states, and arbitrary null background fields. With this simple first-order interaction, it is possible to realize interesting quantum dynamics. An elegant example is a null version of the double-slit interference pattern [46,47]. We will use this process as a benchmark for quantum simulation of the truncated SFQED Hamiltonian, computing the real-time probabilities and comparing with classical simulation.

First, to build intuition, we describe the process using first order perturbation theory. Actually, although the truncated Hamiltonian operates on an eight dimensional Fock space, for the purposes of the Breit-Wheeler process, we need only consider a two-level subsystem corresponding to the basis

states $|e^-e^+\rangle = (1, 0)$ and $|\gamma\rangle = (0, 1)$. We are interested in the amplitude $\langle e^-e^+|U|\gamma\rangle$.

The simplest Hamiltonian truncation corresponds to the Hilbert subspace where the electron and positron have opposite helicities with equal and opposite transverse momenta. Let the positron have helicity $+\frac{1}{2}$ and the electron have helicity $-\frac{1}{2}$, and let the photon polarization be $\lambda = 1$. We take the photon momentum to be $K^\mu = (2p^+, 0, 0, 0)$, the electron momentum to be $P^\mu = (p^+, p^-, p^1, 0)$, and the positron momentum to be $Q^\mu = (p^+, p^-, -p^1, 0)$. In the latter two cases the mass shell condition fixes $p^- = \frac{(p^1)^2 + m^2}{p^+}$.

The background field we will consider corresponds to a linearly polarized plane wave $\mathcal{A}^\mu(\kappa_\mu x^\mu) = (0, 0, \mathcal{A}^1(\kappa_\mu x^\mu), 0)$. For a null field traveling in the $-\hat{z}$ -direction, we have $\kappa_\mu x^\mu = \omega x^+$. The electric field strength E can be nondimensionalized as $\xi = \frac{eE}{m\omega}$, where e is the gauge coupling [67]. A simple case to consider is a very short-duration pulse. An example is given by the gauge potential

$$e\mathcal{A}^1(\omega x^+) = m\xi(1 + \tanh(\omega x^+)). \quad (37)$$

Analytical treatment becomes straightforward if we take the limit of $\omega \rightarrow \infty$, creating a delta-function pulse and a Heaviside-function potential:

$$\lim_{\omega \rightarrow \infty} e\mathcal{A}^1(\omega x^+) = m\xi\theta(x^+). \quad (38)$$

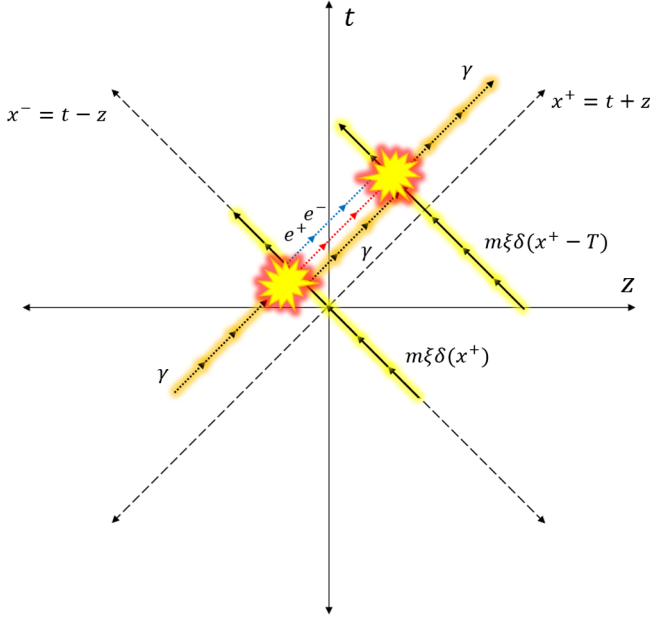


FIG. 5. Two delta-function laser pulses propagate in the $-\hat{z}$ direction, separated in time. A photon collides head-on with the first pulse, leading to a superposition of the $|e^-e^+\rangle$ and $|\gamma\rangle$ states. Collision with the second pulse may cause attenuation or enhancement of the $|e^-e^+\rangle$ state, depending on the pulse separation. Sketched is a case of total destructive interference in pair-production probability for a given electron-positron momentum mode. Light-front coordinates are convenient because all particles interact with the laser pulses at the same light-front time x^+ .

In the null double-slit process of [46], we allow the photon to collide with two such pulses separated in light-front time. Let one pulse appear at $x^+ = 0$ and the other appear at $x^+ = T$. Then our background field is

$$e\mathcal{A}^1(x^+) = m\xi[\theta(x^+) + \theta(x^+ - T)]. \quad (39)$$

Figure 5 shows a spacetime diagram of this nonlinear Breit-Wheeler process.

A. Perturbation theory

Since the process is relatively simple, we can describe the dynamics with first-order perturbation theory. This analysis is somewhat tangent to the main purpose of the paper, for which we could make do merely with a quantum simulation and a classical exact diagonalization numerical treatment for comparison. However, it is helpful to have an analytic description in order to understand the physics of the results and the limitations of perturbation theory.

To calculate the probability $|\langle e^-e^+|U(x^+)|\gamma\rangle|^2$ we expand the interaction-picture time evolution operator to first order,

$$U(x^+) \rightarrow 1 - \frac{i}{2} \int_0^{x^+} H_{\text{int}}(y^+) dy^+. \quad (40)$$

The interaction-picture Hamiltonian consists of

$$H_{\text{int}}(y^+) = -\frac{2me}{\sqrt{2p^{+3}L^3}} e^{ip^-y^+} e^{i(f(P)+g(Q))} a_P^{+\frac{1}{2}\dagger} b_Q^{+\frac{1}{2}\dagger} c_K^{+1}. \quad (41)$$

Here

$$f(P) + g(Q) = \int_0^{y^+} \frac{dz^+}{p^+} [-2p^1 e\mathcal{A}^1(z^+) + (e\mathcal{A}^1(z^+))^2]. \quad (42)$$

Using $\int_0^{y^+} \theta(z^+ - T)^n dz^+ = (y^+ - T)\theta(y^+ - T)$, we can evaluate the phase to be

$$f(P) + g(Q) = \frac{m\xi}{p^+} [(m\xi - 2p^1)y^+\theta(y^+) + (3m\xi - 2p^1)(y^+ - T)\theta(y^+ - T)]. \quad (43)$$

Now let us define the parameter combinations

$$\alpha = \frac{m\xi(m\xi - 2p^1)}{p^+}, \quad \beta = \frac{m\xi(3m\xi - 2p^1)}{p^+}. \quad (44)$$

Rewriting $e^{a\theta(y^+)} = (e^a - 1)\theta(y^+) + 1$, $e^{i(f(P)+g(Q))}$ is expressed compactly as

$$e^{i(f(P)+g(Q))} = e^{i\alpha y^+} (1 - \theta(y^+ - T)) + e^{i(\alpha+\beta)y^+} e^{-i\beta T} \theta(y^+ - T). \quad (45)$$

Note that $\theta(y^+) \rightarrow 1$ since we are considering $y^+ \geq 0$. Putting together the pieces we find

$$\langle e^-e^+|H_{\text{int}}(y^+)|\gamma\rangle = \frac{2me}{\sqrt{2p^{+3}L^3}} [e^{i(\alpha+p^-)y^+} (\theta(y^+ - T) - 1) - e^{i(\alpha+\beta+p^-)y^+} e^{-i\beta T} \theta(y^+ - T)]. \quad (46)$$

Integrating over light-front time, we have the probability amplitude

$$\begin{aligned} \langle e^-e^+|U(x^+)|\gamma\rangle \approx & \frac{me}{\sqrt{2p^{+3}L^3}} \left[\frac{e^{i(\alpha+p^-)x^+}}{\alpha + p^-} (1 - \theta(x^+ - T)) \right. \\ & - \frac{1}{\alpha + p^-} + \frac{e^{i(\alpha+p^-)T}}{\alpha + p^-} \theta(x^+ - T) \\ & + \frac{e^{i(\alpha+\beta+p^-)x^+}}{\alpha + \beta + p^-} e^{-i\beta T} \theta(x^+ - T) \\ & \left. - \frac{e^{i(\alpha+p^-)T}}{\alpha + \beta + p^-} \theta(x^+ - T) \right]. \quad (47) \end{aligned}$$

We can find separate probabilities for $x^+ \gtrsim T$. If $x^+ < T$, then the probability of pair production is

$$P_{\gamma \rightarrow e^- e^+}(x^+ < T) = \frac{2m^2 e^2}{(\alpha + p^-)^2 p^{+3} L^3} \sin^2 \left(\frac{\alpha + p^-}{2} x^+ \right). \quad (48)$$

If $x^+ > T$, then the probability is

$$\begin{aligned} P_{\gamma \rightarrow e^- e^+}(x^+ > T) &= \frac{m^2 e^2}{(\alpha + p^-)^2 (\alpha + \beta + p^-)^2 p^{+3} L^3} [\alpha^2 + \beta^2 + p^{-2} + \alpha(\beta + 2p^-) + \beta p^- \\ &\quad - \beta(\alpha + \beta + p^-) \cos[(\alpha + p^-)T] + \beta(\alpha + p^-) \cos[(\alpha + \beta + p^-)(x^+ - T)] \\ &\quad - (\alpha + p^-)(\alpha + \beta + p^-) \cos[(\alpha + \beta + p^-)x^+ - \beta T]]. \end{aligned} \quad (49)$$

Let us now inspect Eq. (48), the probability after the first pulse has arrived. Time averaging over a cycle, the probability is

$$\langle P \rangle_{\gamma \rightarrow e^- e^+}(x^+ < T) = \frac{m^2 e^2}{(\alpha + p^-)^2 p^{+3} L^3}, \quad (50)$$

which is maximized with respect to the background field strength when α is minimized. This corresponds to $m\xi = p^1$, or $\alpha = -\frac{(p^1)^2}{p^+}$ and $\beta = -\alpha$. Consequently, if we tune p^1 near to $m\xi$ we maximize the conversion probability. Furthermore, we note that since α is parabolic in $m\xi$, there will always be pairs of momenta modes that reach the same level of enhancement.

Restoring the oscillating part of Eq. (48), we also see that the maximum probability is achieved when $x^+ = \frac{(2n+1)\pi}{\alpha + p^-}$ for $n \in \mathbb{Z}^+$.

Let us now inspect Eq. (49). Taking $\beta = -\alpha$ and time averaging, we find the following ratio with Eq. (48):

$$\frac{\langle P \rangle_{\gamma \rightarrow e^- e^+}(x^+ > T)}{\langle P \rangle_{\gamma \rightarrow e^- e^+}(x^+ < T)} = 1 + \frac{\alpha^2}{(p^-)^2} + \frac{2\alpha}{p^-} \cos^2 \left(\frac{\alpha + p^-}{2} T \right). \quad (51)$$

To maximize this ratio, we recognize that $\alpha < 0$, so we need $T = \frac{(2n+1)\pi}{\alpha + p^-}$. If we further take the high-energy limit $p^1 \gg m$, then $\alpha = -p^-$ and we obtain

$$\frac{\langle P \rangle_{\gamma \rightarrow e^- e^+}(x^+ > T)}{\langle P \rangle_{\gamma \rightarrow e^- e^+}(x^+ < T)} = 2. \quad (52)$$

This gives the factor of enhancement from the first time-averaged probability to the next.

B. Classical simulation

In the near-term the accuracy of quantum computations must be benchmarked against classical simulations. We begin by configuring the momentum lattice. The lattice

spacing is $\frac{2\pi}{L}$ and, to approximate the continuum limit, we need this spacing to be less than the typical momentum in the processes of interest. We will take a large $L = \frac{50\pi}{m}$.

Let us now investigate the pair production of electron-positron pairs with differing transverse momenta. As discussed above, the electron four-momentum is given by $P^\mu = (p^+, p^-, p^1, 0)$, the positron four-momentum is $Q^\mu = (q^+, q^-, q^1, 0)$, and the photon four-momentum $K^\mu = (k^+, 0, 0, 0)$. Here $p^+ = q^+$, so $k^+ = 2p^+$, and $q^1 = -p^1$. Because of equal and opposite transverse momenta, we can classify an electron-positron pair by the produced electron transverse momentum p^1 .

For concreteness we will take $\xi = 6$, so that the maximum possible enhancement is seen for $p^1 = 6m$. In addition to simulating this momentum mode, we simulate also $p^1 = 0$ and $p^1 = 12m$ modes, which should have equal enhancements to each other. p^+ is arbitrary and we fix it to $7.2m$. Correspondingly, the light-front energies are $k^- = 0$ for the photon, and $p^- + q^- \approx (0.3, 10, 40)m$ for $p^1 = (0, 6, 12)m$, respectively. Physically, the sign of the momentum results from a negative transverse electric field $E^\perp = -\partial_+ \mathcal{A}^\perp$. Positive particles will experience a force in the $-\hat{x}^1$ direction while negative particles will be accelerated in the $+\hat{x}^1$ direction.

In addition to momentum, the Fock states possess helicity and/or polarization quantum numbers. We let the initial photon be an equal superposition of both polarizations. We will simulate the production of electron-positron pairs with all four different helicity configurations, and sum their probabilities to obtain the final pair-production probability. Accordingly, we keep two photon polarization states for a single momentum mode and four electron-positron helicity states for three electron-positron momentum modes. This creates a 14-dimensional Hilbert space.

Lastly, let us discuss state preparation. The Fock states are not eigenstates of the interacting Hamiltonian. During time evolution, Fock states evolve into superpositions of the true eigenstates. We use adiabatic state preparation to map the initial and final states back and forth between these bases, adiabatically turning on/off the electric coupling e .

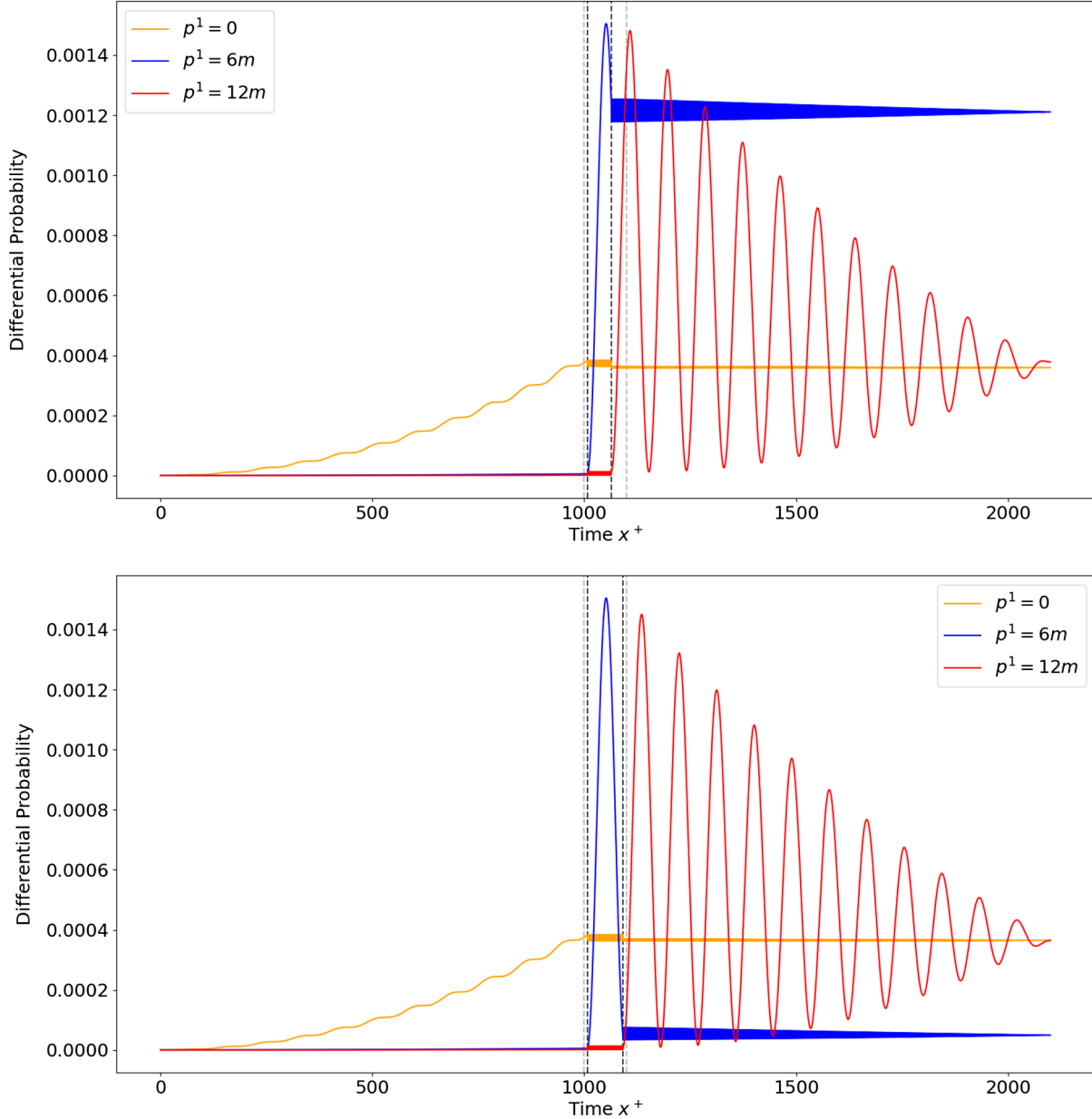


FIG. 6. Classical simulations of nonlinear Breit-Wheeler pair-production differential probabilities as a function of light-front time. The three curves represent three different electron-positron pairs, labeled by the p^1 electron transverse momentum. The vertical light gray lines enclose the physical evolution period when the coupling is held constant. Each dark gray line represents a collision of the incoming photon with the background laser pulse. Only the $p^1 = m\xi$ mode sees significant constructive or destructive interference. Top: two delta pulses with $\xi = 6$, spaced to show constructive interference in the production e^+e^- pairs with $p^1 = 6m$. Bottom: the same, but spaced to show destructive interference. (In the quantum simulations shown below, evolution is performed only between the dark gray lines.).

Adiabatic turn-on/off is performed linearly over 1000 time units, compared to the physical evolution time of 100 units during the two-pulse encounter, where the coupling is constant. However, the effects of adiabatic turn-on are negligible for the finite p^1 modes due to kinematic suppression (these modes are of high light-front energy). Since adding thousands of Trotter steps is prohibitive in the NISQ era, we will focus on the $p^1 = 6m$ mode in the quantum simulations, for which adiabatic turn-on can be

neglected.³ Furthermore, we measure probabilities just after the collision with the second pulse, neglecting adiabatic turn-off, which is also a good approximation for this mode.

Figure 6 shows the classical real-time Hamiltonian simulation. As expected, the pair-production probability

³This remains true even with the artificially inflated value of the coupling used in the quantum simulation, since mixing with the incoming photon state is still strongly kinematically suppressed.

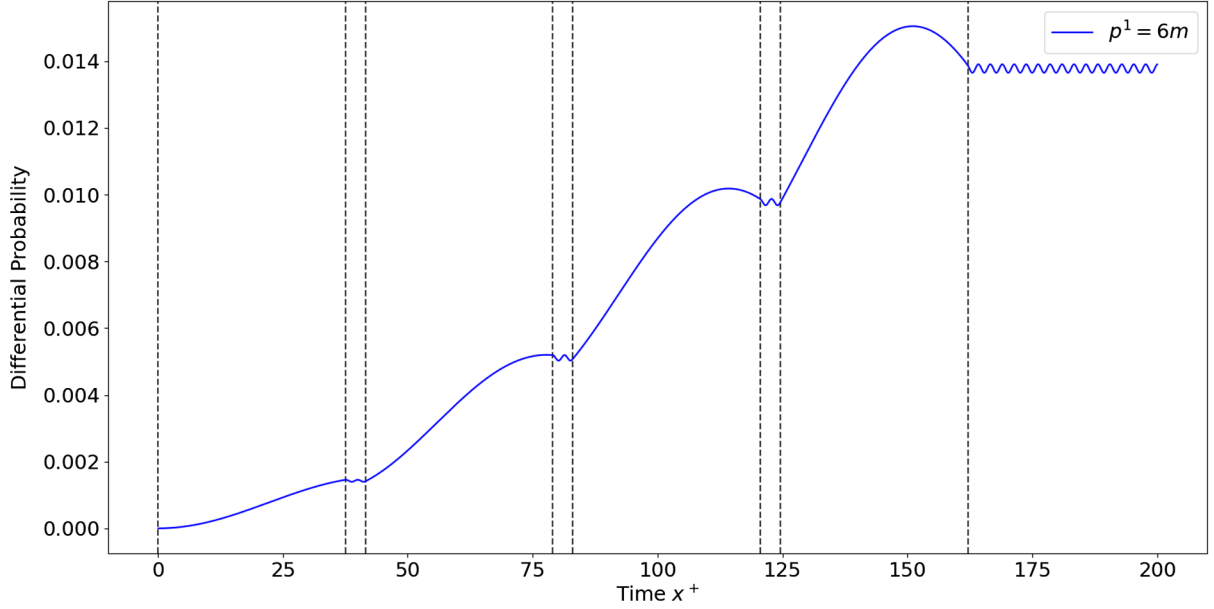


FIG. 7. Eight consecutive delta pulses with $\xi = 6$ are represented by dark gray vertical lines. The signs of the pulses appear in the order $++--++--$. A careful selection of pulse spacings can lead to substantial enhancement of the pair production probability in specific momentum modes.

for the $p^1 = 6m$ mode can experience constructive or destructive interference depending on the light-front time delay T separating the pulses. The $p^1 = 0$ and $p^1 = 12m$ modes experience equal enhancements asymptotically. The various oscillatory behaviors appearing before and after the second pulse are captured by the perturbative formulas. After the first pulse, Eq. (48) shows that both the frequency and amplitude of oscillations are sensitive to the combination $\alpha + p^-$, which is quadratic in p^1 . If $p^1 = m\xi$, then $\alpha + p^-$ is minimized and we see relatively shorter frequencies and larger amplitudes. If $p^1 \neq m\xi$, then we see longer frequency and shorter amplitude oscillations (as in the $p^1 = 0$ and $p^1 = 12m$ modes). From Eq. (49) we see that the oscillation frequencies after the second pulse depend on the combination $\alpha + \beta + p^-$. Although still quadratic in p^1 , this expression is minimized and equal to $\frac{m^2}{p^-}$, independent of ξ , when $p^1 = 2m\xi$. This is why the $p^1 = 12m$ mode has a lower frequency than the $p^1 = 0$ and $p^1 = 6m$ modes. The amplitudes of the modes now have a more complicated dependence in p^1 . Roughly speaking, a factor of $\frac{1}{(\alpha + \beta + p^-)^2}$ leads to large amplitudes for $p^1 \approx 2m\xi$. Away from this minimum, however, the amplitudes are suppressed by $\mathcal{O}(1/\xi^2)$. Correspondingly, for the $p^1 = 0$ and $p^1 = 6m$ modes we see much smaller amplitudes: the second pulse effectively “stops” the oscillations. For these modes far from $2m\xi$, adiabatic turn-off is therefore less significant.

The simulations allow us to visualize interference effects in real time, which are most important in the $p^1 = m\xi$ mode. The first pulse gives a kick to the $p^1 = m\xi$ mode and

its probability starts to oscillate. The second pulse stabilizes the oscillation of this mode near its maximum. At the same time, the application of two kicks also yields an amplitude in the $p^1 = 2m\xi$ mode.

As a side note, if we introduce a third field with strength $-m\xi$, we might expect to obtain further constructive interference in the $p^1 = m\xi$ mode. In other words, if we vary the net field between $m\xi$ with consecutive pulses, we expect to see multiple constructive interference enhancements in the $p^1 = m\xi$ mode pair-production probabilities. Figure 7 exhibits this qualitative behavior⁴ using a background field of the form

$$e\mathcal{A}^1(x^+) = m\xi[\theta(x^+) + \theta(x^+ - T_1) - \theta(x^+ - T_2) - \theta(x^+ - T_3) + \theta(x^+ - T_4) + \theta(x^+ - T_5) - \dots]. \quad (53)$$

The light-front time differences $T_{\text{odd}} - T_{\text{even}}$ and $T_{\text{even}} - T_{\text{odd}}$ were kept constant. Subsequent enhancements exhibit somewhat shorter frequency oscillations, so fine-tuning the pulse separations could lead to greater net constructive interference.

V. QUANTUM RESULTS

Near-term quantum devices are limited by quantum noise. This means we cannot precisely perform the same

⁴For the reasons discussed above, we did not include adiabatic turn-on or turn-off in making Fig. 7, and so we only display the dynamical time between pulses.

classical simulation on a quantum device. The equivalent number of qubits needed for the Breit-Wheeler simulation of Fig. 6 is $n = 14$.⁵ Based on the discussion below Eq. (33), we would then expect of order $3(2 \times 2)(2 \times 2)(2 \times 2) = 192$ distinct Pauli strings in the Hamiltonian. For a rough lower bound estimate, a single Trotter step could be implemented with $\gtrsim 600$ gates. Moreover, the majority of Trotter steps would require long-reaching CNOT gates which would introduce many SWAP gates and hence more noise. For this reason, we will return to the three qubit quantum simulation discussed in Sec. III, achieved by truncating to fewer helicities and polarizations and momenta.

Beyond gate errors, the probabilities we can actually measure are limited by shot noise. This is a problem particularly for real theories characterized by a weak coupling, like QED, where interesting processes may have quite low probabilities. To avoid requiring an infeasible number of shots, we artificially boost the coupling constant $e = 0.303 \rightarrow 60$ and decrease the lattice parameter $L \rightarrow \frac{6\pi}{m}$. For the electron momentum, we let $p^+ = p^1 = \frac{8m}{3}$. (It still is the case that the positron momentum has $q^+ = p^+$ and $q^1 = -p^1$. For the photon, $k^+ = 2p^+$ as before.) With these choices we are still effectively near the continuum limit, since the momentum lattice spacing $\frac{2\pi}{L} = \frac{m}{3}$ is an eighth of the typical momentum magnitude.

Although taking $e = 60$ appears decidedly nonperturbative, what our simulations compute are exclusive probabilities in small cells of Fock space, or effectively differential probabilities times small phase space factors. By retaining only a few states in the full Fock space, each with a small volume $(\frac{2\pi}{L})^3$, the truncated quantum mechanical model is still perturbative despite the large value of e . Tree-level results in the full QFT may then be obtained by rescaling. We emphasize however that this approach will break down if many more Fock states are kept in the simulation (increasing the amplitudes of second-order and higher transitions in the truncated model.) Expanding the truncated theory in this way is of course essential to match with continuum SFQED beyond the leading order and to see more interesting dynamics, so care will have to be taken to design simulations that can operate at the physical coupling without requiring prohibitively large shot counts. In any event, for our purposes an upscaling of the coupling is sufficient. Figure 8 shows that for $e = 60$ we still obtain qualitatively reasonable agreement between Trotterization and first-order perturbation theory in the truncated quantum mechanics (or the tree level prediction of full SFQED) in this case.

The single Trotter step circuit in Fig. 4, once expressed in native IBM gates, is 40 gates long. While the classical

⁵ $n = 14$ qubits = (2 photon momenta) \times (2 polarizations) + (3 electron momenta) \times (2 helicities) + (3 positron momenta) \times (2 helicities).

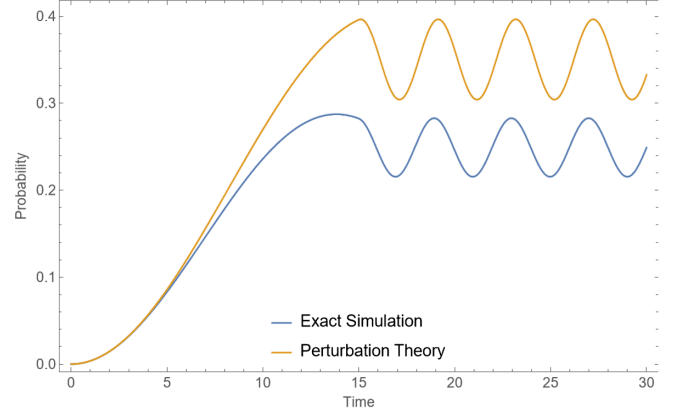
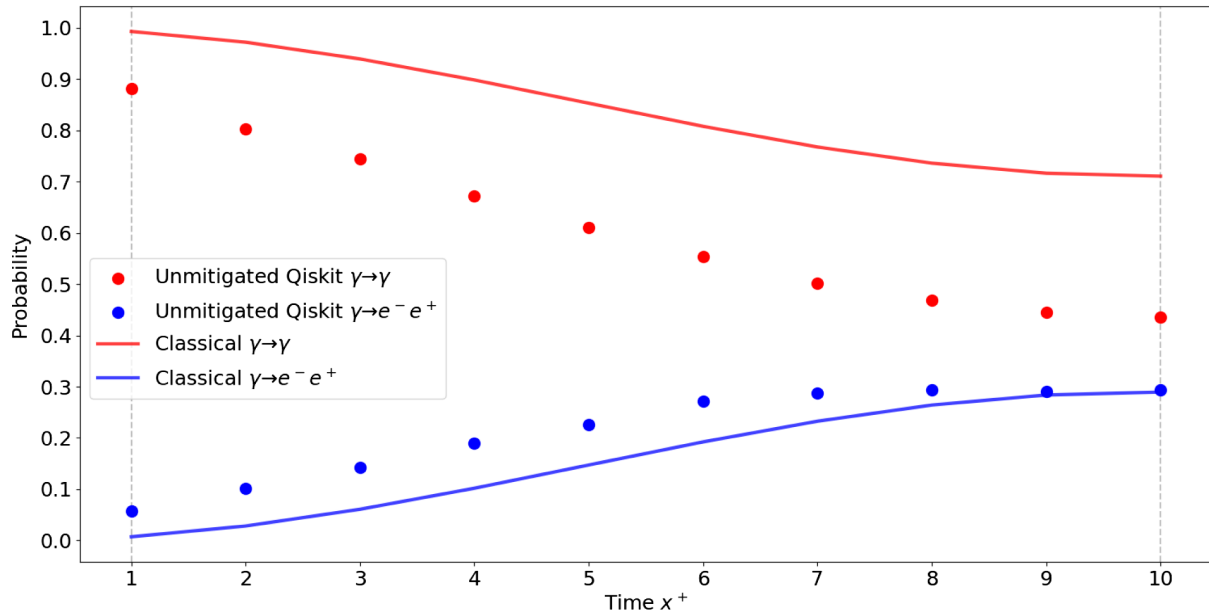


FIG. 8. Comparison of real-time pair-production probabilities computed with perturbation theory and “exact” time evolution (performed numerically via Trotterization with infinitesimal time step) for artificially large coupling. An initial delta pulse is encountered at time zero, and then the probability oscillation amplitudes change when the second pulse is encountered. At $e = 60$, perturbation theory still yields probabilities below one, and the two methods agree qualitatively. Perturbation theory provides a useful understanding of the process, despite the large coupling, because the phase space volume is small (both for the final state and the Hilbert space truncation as a whole, which prevents higher order effects from becoming large.).

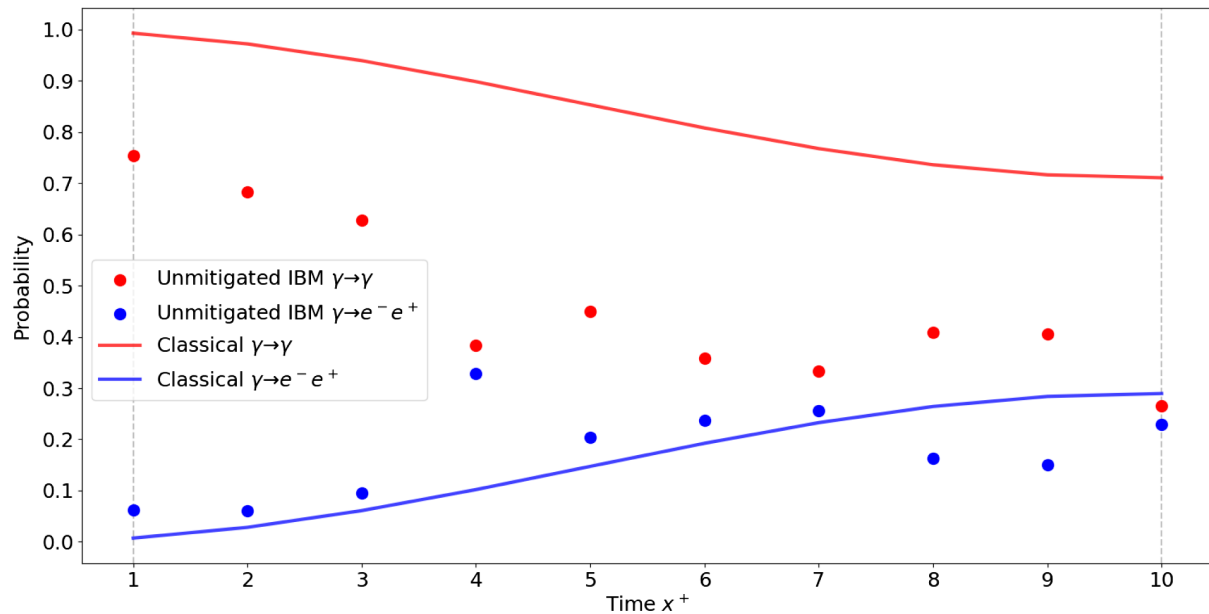
simulation discussed previously was performed over 21000 time steps, most of which were adiabatic turn-on/turn-off steps, the quantum simulation must be limited to a few hundred gates total in order to avoid total decoherence. To this end, we omit adiabatic turn-on/turn-off in the quantum simulations. Instead, we fix the initial state to the free photon state $|001\rangle$, which as we have discussed previously is a good approximation for the physics of interest. We also begin the time evolution at the (light-front) time of collision between the photon and the first pulse and end it just after the collision with the second pulse (e.g., the quantum simulation is performed only between the dark gray lines of Fig. 6). Within this region, the pair-production probability oscillates. Because of NISQ limitations, we can only afford to capture a small portion of these oscillations, such as half a period. Using ten Trotter steps, we show how the probability changes from minimum to maximum. To fully capture the noisy dynamics, we also record the $\gamma \rightarrow \gamma$ probabilities.

It is convenient to estimate the impact of noise first on a classical simulator, using a noise model based on historical measurements. Results using Qiskit’s FakeNairobi2 noisy simulator [68] are shown in Fig. 9(a). The data suggest that we should expect at least 10% errors already in the initial time step, increasing with time.

Figure 9(b) shows raw quantum data from IBM_NAIROBI [69]. The real noise level we observe is higher than seen in the noisy FakeNairobi2 simulator. At $x^+ = 4$ the error in the $\gamma \rightarrow \gamma$ probability is already 60%. Such results are not



(a) Results from Qiskit's noisy simulator reflect modest errors, more significant in $\gamma \rightarrow \gamma$ than $\gamma \rightarrow e^- + e^+$ probabilities.



(b) Data from IBM's quantum hardware reflects strong noise effects, particularly in $\gamma \rightarrow \gamma$ probabilities.

FIG. 9. Unmitigated quantum results for nonlinear Breit-Wheeler pair production. Pulses with $m\xi = \frac{8m}{3}$ are encountered at $x^+ = 1$ and $x^+ = 10$. The coupling is held fixed at $e = 60$. The electron transverse momentum mode is $p_1 = m\xi$.

uncommon from NISQ devices and can be substantially cleaned by mitigation algorithms.

VI. ERROR MITIGATION

Mitigating against specific sources of device noise is critical to extract useful information from near-term

quantum simulations. We employ measurement mitigation and Pauli twirling [53], as well as a modified version of self-mitigation [19,57,58] (mitigation of depolarization noise). The combination of these three error mitigation techniques significantly improved the quantum data as shown in Fig. 13(b). For completeness we discuss each mitigation algorithm in some detail.

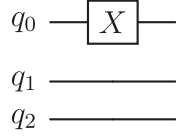


FIG. 10. The measurement mitigation circuit C_1 for preparing $|\psi_1\rangle = |001\rangle$ when $n = 3$ qubits. The initial state is $|\psi_0\rangle = |000\rangle$ so $C_1|000\rangle = |001\rangle$.

A. Measurement mitigation

With measurement mitigation we account for readout errors of bitstrings. For example, there is the probability that upon readout, the measured bitstring is the computational basis state $|000\rangle$ as opposed to $|001\rangle$. For n qubits, the mitigation strategy is to run 2^n circuits that prepare individual computational basis states and derive statistics for how many counts there are for the 2^n possible results.

Let $\{|\psi_i\rangle\}$ for $i = 0, 1, \dots, 2^n - 1$ be the computational basis. Prepare circuits $C_0, C_1, \dots, C_{2^n-1}$ that prepare each computational basis state as in Fig. 10. For each circuit C_j , measure $|\langle\psi_i|C_j|\psi_0\rangle|^2$. These probabilities are entries $C_{i,j} = |\langle\psi_i|C_j|\psi_0\rangle|^2$ of the calibration matrix \mathcal{C} . That is, the measurement probabilities from C_j form the j th column of \mathcal{C} .

For a general quantum circuit, let \vec{c}^{true} be the column vector of noiseless measurement counts of each basis state. The product $\mathcal{C}\vec{c}^{\text{true}}$ distributes the true measurement counts among other basis states such that the calibration matrix models the effects of readout noise. Thus, define $\vec{c}^{\text{meas}} = \mathcal{C}\vec{c}^{\text{true}}$.

In reality, noisy quantum measurements yield \vec{c}^{meas} , so in principle we can obtain the mitigated counts by matrix inversion: $\vec{c}^{\text{true}} = \mathcal{C}^{-1}\vec{c}^{\text{meas}}$. However, simply inverting \mathcal{C} can lead to negative counts and counts that do not sum to the number of shots N . Therefore, we utilize a least-squares

minimization protocol. This method takes advantage of the notion that $\vec{c}^{\text{meas}} - \mathcal{C}\vec{c}^{\text{true}} = 0$. Let \vec{x} be an estimate for \vec{c}^{true} . Then the problem becomes finding

$$\min_{\vec{x}} [(\vec{c}^{\text{meas}} - \mathcal{C}\vec{x})^T (\vec{c}^{\text{meas}} - \mathcal{C}\vec{x})] \quad (54)$$

subject to the constraints that $0 \leq x_i \leq N$ and $|\vec{x}| = N$. The minimization will yield the best estimate of \vec{c}^{true} as predicted by \mathcal{C} . The mitigated probabilities are then $\frac{1}{N}\vec{c}^{\text{true}}$.

B. Pauli twirling

CNOT gates may have different errors depending on the state that is fed to them. Pauli twirling implements gate identities around CNOTs in order to feed them different states, thus “symmetrizing” the noise. Figure 11 shows examples of CNOT identities. In total there are 16 ways to twirl a CNOT gate with extra Pauli gates. We randomly twirl all CNOTs with different identities in each run, then average over the twirled results in postprocessing.

C. Depolarization mitigation

The quantum channel often used to describe depolarizing noise [70] is

$$\tilde{\rho} = p \frac{I}{2^n} + (1-p)\rho. \quad (55)$$

Under this channel a state ρ may turn into the fully decohered state $\frac{I}{2^n}$ with probability p , and with probability $1-p$ the state remains unchanged. This type of depolarizing noise model is called symmetric depolarization and is the basis for many implementations of “self-mitigation” [19,57,58]. A more general noise model could take ρ not just to $\frac{I}{2^n}$, but to other general mixed states, which is what we will now consider.

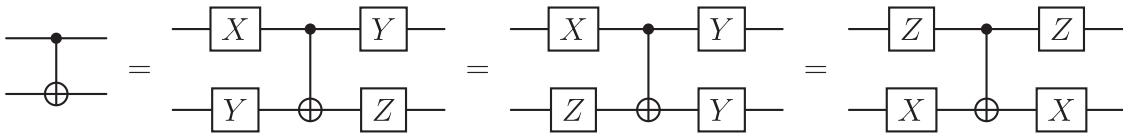


FIG. 11. Sample Pauli twirls of the CNOT gate. For error mitigation, each circuit was randomly compiled with every CNOT twirled using these equivalent gates.

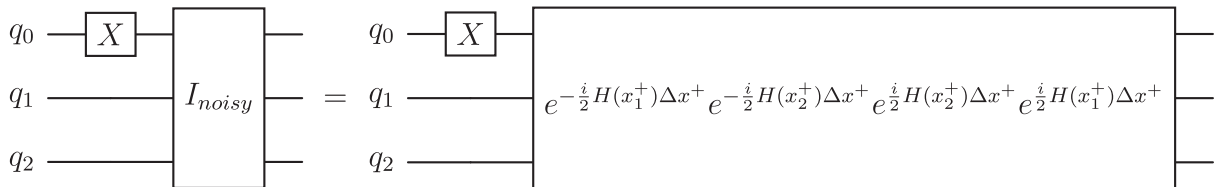
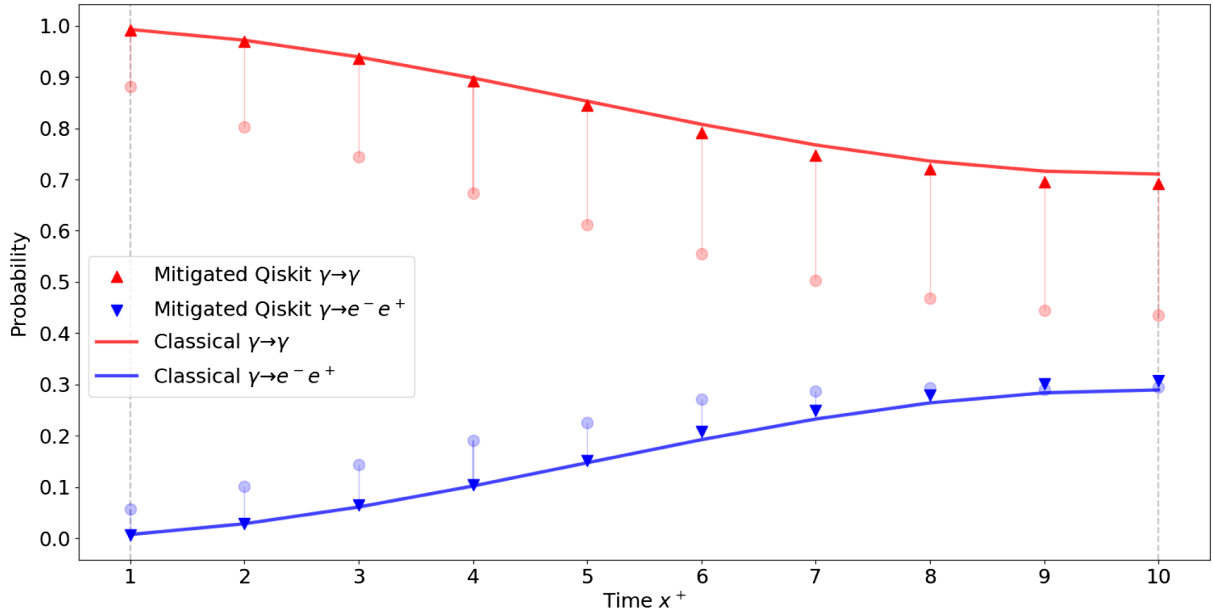
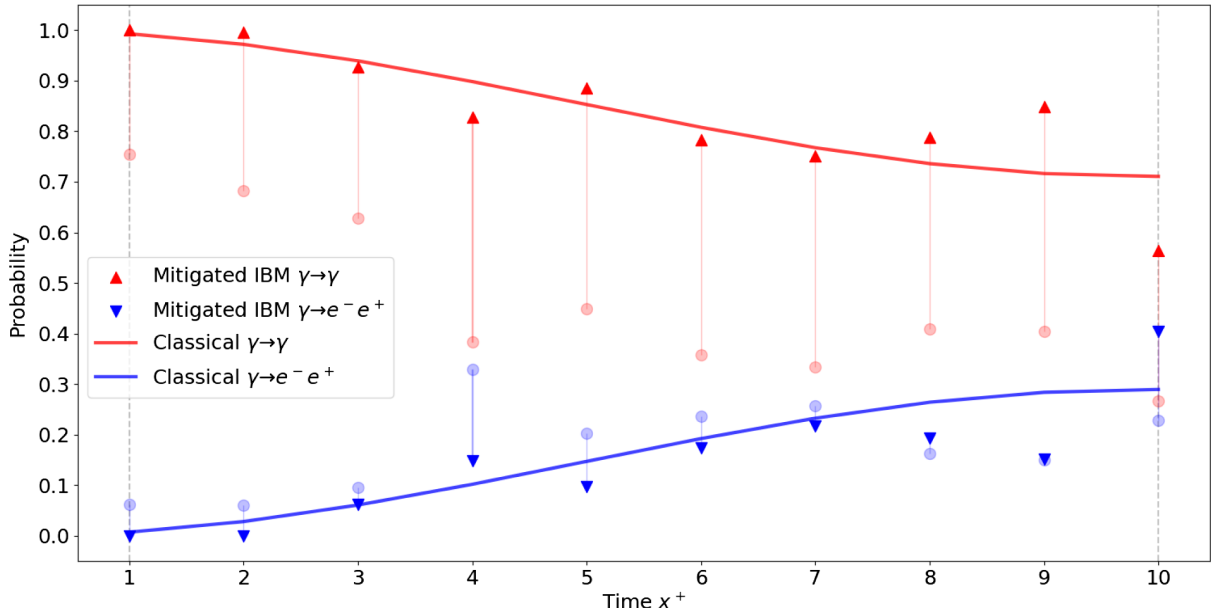


FIG. 12. The depolarizing mitigation circuit $C_1(x_2^+)$ for the second time step, preparing $|\psi_1\rangle = |001\rangle$ when $n = 3$ qubits. The noisy identity is the time evolution circuit used for the quantum simulation followed by its Hermitian conjugate.



(a) Mitigated data from Qiskit's fake simulator shows excellent agreement with classical simulation results.



(b) Mitigated data from IBM's quantum hardware agrees well with classical results until $x^+ = 9$, suppressing high levels of errors seen in raw data.

FIG. 13. Mitigated quantum results for nonlinear Breit-Wheeler pair production. Pulses with $m\xi = \frac{8m}{3}$ are encountered at $x^+ = 1$ and $x^+ = 10$. Robust measurement and depolarization mitigation techniques suppress errors at early time steps, but become less effective at late time steps. The unmitigated simulation is shown with the faded data for comparison.

We generalize the depolarizing noise model by introducing a transfer matrix M :

$$\sum_{j=0}^{2^n-1} M_{ij} P_j^{\text{true}} \equiv P_i^{\text{meas}}, \quad (56)$$

where P_i^{true} (P_i^{meas}) is the true (noisy measurement) probability of computational basis state $|\psi_i\rangle$. (Symmetric depolarizing noise is the case where M is a one-parameter matrix with entries $M_{ij} = \delta_{ij}(1 - 2^n \epsilon) + \epsilon$ for $\epsilon = \frac{p}{2^n}$. ϵ may be estimated by preparing circuits with similar noise elements as the circuit of interest.) Asymmetric

depolarizing noise corresponds, for example, to the following quantum channel:

$$\tilde{\rho} = \sum_{i=0}^{2^n-1} \sum_{j=0}^{2^n-1} M_{ij} |\psi_i\rangle \langle \psi_j| \rho |\psi_j\rangle \langle \psi_i|, \quad (57)$$

where the Kraus operators convert the j th computational basis state $|\psi_j\rangle$ to the i th state $|\psi_i\rangle$ with weight $\sqrt{M_{ij}}$. To estimate the elements of M_{ij} in our model, we construct circuits $C_0, C_1, \dots, C_{2^n-1}$, similar to those for measurement mitigation. However, in addition to preparing each computational basis state, the circuit applies an identity operator with the same gates and the physics circuit, as shown in Fig. 12. Each noisy identity is the current time evolution operator followed by its inverse. This means that the mitigation circuits C_i are time dependent $C_i(x^+)$, allowing us to mitigate against time-dependent noise. The transfer matrix elements are then estimated as $M_{ij}(x^+) = |\langle \psi_i | C_j(x^+) | \psi_0 \rangle|^2$.

We emphasize that in both the symmetric case and our case, Pauli twirling is essential to convert more general types of noise from the CNOTs into depolarization noise. Thus, depolarization mitigation is the last step of the error mitigation pipeline. Both the physics and mitigation circuits are twirled and the results are averaged before depolarization mitigation.

Furthermore, since the inserted identity operators are effectively “twice the physics circuit,” each mitigation circuit is practically a measure for the noise “squared.” This implies that we should be taking \sqrt{M} as the calibration matrix.⁶ That is,

$$\vec{c}^{\text{meas}} = \sqrt{M} \vec{c}^{\text{true}}. \quad (58)$$

As before, to determine \vec{c}^{true} , we solve an optimization problem

$$\min_{\vec{x}} [(\vec{c}^{\text{meas}} - \sqrt{M}\vec{x})^T W (\vec{c}^{\text{meas}} - \sqrt{M}\vec{x})] \quad (59)$$

subject to the same constraints on \vec{x} as before. The weights W are taken to be $W = \text{diag}(\frac{1}{p_i})$, where $p_i = |\langle \psi_i | C_i | \psi_0 \rangle|^2$.

We checked that symmetric depolarization noise mitigation following Pauli twirling did not perform well on our problem: the noise profile indicated some states “thermalize” with each other, but not all together. On the other hand, the construction of a complete calibration matrix for depolarization mitigation is computationally expensive and does not scale well with problem size. For the small size studied here it was not critical to

optimize it further, but it would be interesting to explore the performance restricted classes of calibration transformations constructed from fewer parameters and mitigation circuit measurements.

D. Mitigated results

For each time step, we performed ten Pauli twirls on the quantum simulation circuit as well as each of the $2^3 = 8$ mitigation circuits. Each time step was sent as a different set of circuits to IBM_NAIROBI at different times, so in addition we ran eight measurement mitigation circuits each time step. Therefore, 98 circuits were sent to IBM_NAIROBI at each time step.

As before we can compare with Qiskit’s noisy FakeNairobi2 simulator. Figure 13(a) shows the mitigated noisy simulator results. The mitigation is substantial and suggests that the real quantum simulation should also see improved results versus the raw data.

Figure 13(b) shows mitigated quantum data. As expected, the error mitigation protocol greatly improves the simulation results. The largest percent error we see is about 15% of the classical simulation. However, substantial errors do not occur until the last two time steps. This suggests that had we performed a longer or finer time evolution, decoherence may have been too strong to be well mitigated. (See Appendix B for a comparison between mitigation against symmetric versus asymmetric depolarizing noise.)

VII. CONCLUSION

Quantum field theory in the presence of strong background fields is a rich arena where quantum computers may provide valuable new access. We have seen that for a simple strong-field QED process—nonlinear Breit-Wheeler pair production in a sequence of laser pulses—the noisy quantum computers of today are already able to simulate the dynamics reliably with the help of error mitigation routines. Because of the polynomial scaling in circuit depth, the simulation of SFQED with many particles and strong backreaction presents an interesting, plausible long-term goal.

Capturing more complex electromagnetic cascade processes requires retaining a much larger Hilbert space. With a less drastic truncation, one could also study helicity and polarization effects in cascades, as well as begin to tackle the regime of ultrastrong fields, where few other tools exist. The light-front Fock space formulation used in this work can be applied directly to more complicated processes with bigger Hilbert spaces, and the operator-circuit map can be straightforwardly extended.

It is also possible to simulate O(30-50) qubit quantum computers with state-of-the-art GPUs and supercomputers [71]. Larger-scale SFQED quantum circuits could be designed on these simulators with adjustable noise. This capability could be useful to develop efficient circuits for

⁶This is also why the measurement mitigation and depolarization mitigation cannot be done at the same time in our problem, despite both being formulated in terms of a calibration matrix.

future quantum computers, and to establish noise targets for interesting physics processes. Investigations in this direction are ongoing.

ACKNOWLEDGMENTS

We thank Bryce Gadow for discussions and suggesting the possibility of greater enhancements from many pulses. This work was supported in part by the U.S. Department of Energy, Office of Science, Office of High Energy Physics under Award No. DE-SC0015655 and by its QuantISED program under an award for the Fermilab Theory Consortium ‘‘Intersections of QIS and Theoretical Particle Physics.’’ We acknowledge the use of IBM Quantum services for this work. The views expressed are those of the authors, and do not reflect the official policy or position of IBM or the IBM Quantum team.

APPENDIX A: QUANTIZED HAMILTONIAN

The SFQED Hamiltonian density was derived in Sec. II. Integrating the density over light-front space yields the Hamiltonian. We also expand the fields in momentum modes. The fields are given in Eq. (22) and the commutation relations are given in Eq. (18). In this Appendix, we

derive a few sample terms in the mode expansion of the Hamiltonian.

1. Fermion energy Hamiltonian

We begin with the free fermion energy H_ψ . We will need to compute expressions of the form $\frac{f(x^-)}{\partial_-}$. In momentum space the Green function contributes factors of $\frac{1}{ip^\pm}$. We have

$$\frac{\psi_-}{\partial_-} = \int \frac{dp^+ dp^\perp}{(2\pi)^3} \sum_{s=\pm\frac{1}{2}} \frac{2e^{-ipx}}{-ip^+} w^s a_p^s + \frac{2e^{ipx}}{ip^+} w^{-s} b_p^{s\dagger} \quad (\text{A1})$$

and

$$\begin{aligned} \frac{\partial_i \partial_i \psi_-}{\partial_-} &= \int \frac{dp^+ dp^\perp}{(2\pi)^3} \sum_{s=\pm\frac{1}{2}} \frac{2(ip^\perp)^2 e^{-ipx}}{-ip^+} w^s a_p^s \\ &+ \frac{2(-ip^\perp)^2 e^{ipx}}{ip^+} w^{-s} b_p^{s\dagger}. \end{aligned} \quad (\text{A2})$$

Here, s is the positron helicity, and the electron has helicity $-s$. Using orthogonality of the spinors,

$$w^{\pm s\dagger} w^{\pm s'} = \delta_{s,s'} \quad w^{\pm s\dagger} w^{\mp s'} = \delta_{-s,s'}, \quad (\text{A3})$$

and integrating over space, we find

$$\begin{aligned} H_\psi &= m^2 \int \frac{dp^+ dp^\perp}{(2\pi)^3} \int dp'^+ dp'^\perp \sum_{s=\pm\frac{1}{2}} \left(\frac{\delta^3(\mathbf{p}-\mathbf{p}')}{p'^+} a_p^{s\dagger} a_{p'}^s - \frac{\delta^3(\mathbf{p}-\mathbf{p}')}{p'^+} b_p^s b_{p'}^{s\dagger} - \frac{\delta^3(\mathbf{p}+\mathbf{p}')}{p'^+} a_p^{s\dagger} b_{p'}^{-s\dagger} + \frac{\delta^3(\mathbf{p}+\mathbf{p}')}{p'^+} b_p^s a_{p'}^{-s} \right) \\ &+ \int \frac{dp^+ dp^\perp}{(2\pi)^3} \int dp'^+ dp'^\perp \sum_{s=\pm\frac{1}{2}} (p'^\perp)^2 \left(\frac{\delta^3(\mathbf{p}-\mathbf{p}')}{p'^+} a_p^{s\dagger} a_{p'}^s - \frac{\delta^3(\mathbf{p}-\mathbf{p}')}{p'^+} b_p^s b_{p'}^{s\dagger} - \frac{\delta^3(\mathbf{p}+\mathbf{p}')}{p'^+} a_p^{s\dagger} b_{p'}^{-s\dagger} + \frac{\delta^3(\mathbf{p}+\mathbf{p}')}{p'^+} b_p^s a_{p'}^{-s} \right). \end{aligned} \quad (\text{A4})$$

The delta functions enforce momentum conservation, $p^+ = p'^+$ and $p^\perp = p'^\perp$. However, $\delta^3(\mathbf{p}+\mathbf{p}')$ and $p^+ > 0$ indicates that integrals with $p^+ \leq 0$ vanish. Only $\delta^3(\mathbf{p}-\mathbf{p}')$ terms contribute:

$$\begin{aligned} H_\psi &= m^2 \int \frac{dp^+ dp^\perp}{(2\pi)^3} \sum_{s=\pm\frac{1}{2}} \left(\frac{1}{p^+} a_p^{s\dagger} a_p^s - \frac{1}{p^+} b_p^s b_p^{s\dagger} \right) + \int \frac{dp^+ dp^\perp}{(2\pi)^3} \sum_{s=\pm\frac{1}{2}} (p^\perp)^2 \left(\frac{1}{p^+} a_p^{s\dagger} a_p^s - \frac{1}{p^+} b_p^s b_p^{s\dagger} \right) \\ &= \int \frac{dp^+ dp^\perp}{(2\pi)^3} \sum_{s=\pm\frac{1}{2}} \left(\frac{(p^\perp)^2 + m^2}{p^+} a_p^{s\dagger} a_p^s - \frac{(p^\perp)^2 + m^2}{p^+} b_p^s b_p^{s\dagger} \right) \\ &= \int \frac{dp^+ dp^\perp}{(2\pi)^3} \sum_{s=\pm\frac{1}{2}} p^- (a_p^{s\dagger} a_p^s - b_p^s b_p^{s\dagger}). \end{aligned} \quad (\text{A5})$$

Finally, we normal order by anticommuting the positron operators in the usual way. The final result for this term is

$$:H_\psi: = \int \frac{dp^+ dp^\perp}{(2\pi)^3} \sum_{s=\pm\frac{1}{2}} p^- (a_p^{s\dagger} a_p^s + b_p^{s\dagger} b_p^s) \quad (\text{A6})$$

which sums the light-front energies of the fermions.

2. Fermion-photon interaction Hamiltonian

The other terms for which we will derive the explicit mode expansion are the interaction Hamiltonian used in our simulations. These terms come from

$$\mathcal{H}_{\psi A} \supset \frac{ime}{2} \psi_-^\dagger \gamma^j \left(\frac{A^j \psi_-}{\partial_-} \right) - \frac{ime}{2} \psi_-^\dagger \gamma^j A^j \left(\frac{\psi_-}{\partial_-} \right). \quad (\text{A7})$$

(Here repeated indices, although all raised, are understood to be summed over.) It is convenient to define the following coefficients R :

$$R^{s,j} \equiv w^{s\dagger} \gamma^j w^{-s} \quad R^{+\frac{1}{2},j} = \begin{cases} 1 & j=1 \\ i & j=2 \end{cases} \quad R^{-\frac{1}{2},j} = \begin{cases} -1 & j=1 \\ i & j=2 \end{cases}. \quad (\text{A8})$$

Integrating over light-front space we obtain

$$\begin{aligned} \frac{1}{2} \int d^3 \mathbf{x} \psi_-^\dagger \gamma^j \left(\frac{A^j \psi_-}{\partial_-} \right) &= 2i \int \frac{dp^+ dp^\perp}{(2\pi)^3} \int \frac{dk^+ dk^\perp}{(2\pi)^3 \sqrt{k^+}} \int dp'^+ dp'^\perp \sum_{s=\pm\frac{1}{2}} \sum_{j=1,2} \\ &\times \left(\frac{\delta^3(-\mathbf{p} + \mathbf{k} + \mathbf{p}')}{k^+ + p'^+} R^{s,j} a_p^{s\dagger} c_k^j a_{p'}^{-s} - \frac{\delta^3(\mathbf{p} + \mathbf{k} + \mathbf{p}')}{k^+ + p'^+} R^{s,j} a_p^{s\dagger} c_k^{j\dagger} b_{p'}^{s\dagger} + \frac{\delta^3(-\mathbf{p} + \mathbf{k} - \mathbf{p}')}{k^+ - p'^+} R^{s,j} a_p^{s\dagger} c_k^j b_{p'}^{s\dagger} \right. \\ &- \frac{\delta^3(\mathbf{p} + \mathbf{k} - \mathbf{p}')}{k^+ - p'^+} R^{s,j} a_p^{s\dagger} c_k^{j\dagger} a_{p'}^{-s} + \frac{\delta^3(\mathbf{p} + \mathbf{k} + \mathbf{p}')}{k^+ + p'^+} R^{-s,j} b_p^s c_k^j a_{p'}^s - \frac{\delta^3(-\mathbf{p} + \mathbf{k} + \mathbf{p}')}{k^+ + p'^+} R^{-s,j} b_p^s c_k^{j\dagger} b_{p'}^{-s\dagger} \\ &\left. + \frac{\delta^3(\mathbf{p} + \mathbf{k} - \mathbf{p}')}{k^+ - p'^+} R^{-s,j} b_p^s c_k^j b_{p'}^{-s\dagger} - \frac{\delta^3(-\mathbf{p} + \mathbf{k} - \mathbf{p}')}{k^+ - p'^+} R^{-s,j} b_p^s c_k^{j\dagger} a_{p'}^s \right) \end{aligned} \quad (\text{A9})$$

and

$$\begin{aligned} \frac{1}{2} \int d^3 \mathbf{x} \psi_-^\dagger \gamma^j A^j \left(\frac{\psi_-}{\partial_-} \right) &= 2i \int \frac{dp^+ dp^\perp}{(2\pi)^3} \int \frac{dk^+ dk^\perp}{(2\pi)^3 \sqrt{k^+}} \int dp'^+ dp'^\perp \sum_{s=\pm\frac{1}{2}} \sum_{j=1,2} \\ &\times \left(\frac{\delta^3(-\mathbf{p} + \mathbf{k} + \mathbf{p}')}{p'^+} R^{s,j} a_p^{s\dagger} c_k^j a_{p'}^{-s} - \frac{\delta^3(\mathbf{p} + \mathbf{k} + \mathbf{p}')}{p'^+} R^{s,j} a_p^{s\dagger} c_k^{j\dagger} b_{p'}^{s\dagger} - \frac{\delta^3(-\mathbf{p} + \mathbf{k} - \mathbf{p}')}{p'^+} R^{s,j} a_p^{s\dagger} c_k^j b_{p'}^{s\dagger} \right. \\ &+ \frac{\delta^3(\mathbf{p} + \mathbf{k} - \mathbf{p}')}{p'^+} R^{s,j} a_p^{s\dagger} c_k^{j\dagger} a_{p'}^{-s} + \frac{\delta^3(\mathbf{p} + \mathbf{k} + \mathbf{p}')}{p'^+} R^{-s,j} b_p^s c_k^j a_{p'}^s - \frac{\delta^3(-\mathbf{p} + \mathbf{k} + \mathbf{p}')}{p'^+} R^{-s,j} b_p^s c_k^{j\dagger} b_{p'}^{-s\dagger} \\ &\left. - \frac{\delta^3(\mathbf{p} + \mathbf{k} - \mathbf{p}')}{p'^+} R^{-s,j} b_p^s c_k^j b_{p'}^{-s\dagger} + \frac{\delta^3(-\mathbf{p} + \mathbf{k} - \mathbf{p}')}{p'^+} R^{-s,j} b_p^s c_k^{j\dagger} a_{p'}^s \right). \end{aligned} \quad (\text{A10})$$

Now let us integrate \mathbf{p}' while keeping explicit the momentum conservation in each integrand. Once again integrals with $\delta^3(\mathbf{p} + \mathbf{k} + \mathbf{p}')$ vanish since they require a longitudinal momentum component to be nonpositive. We obtain

$$\begin{aligned} \frac{1}{2} \int d^3 \mathbf{x} \psi_-^\dagger \gamma^j \left(\frac{A^j \psi_-}{\partial_-} \right) &= 2i \int \frac{dp^+ dp^\perp}{(2\pi)^3} \int \frac{dk^+ dk^\perp}{(2\pi)^3 \sqrt{k^+}} \sum_{s=\pm\frac{1}{2}} \sum_{j=1,2} \\ &\times \left(\frac{1}{k^+ + p'^+} R^{s,j} a_p^{s\dagger} c_k^j a_{p'}^{-s} \Big|_{p'=p-k} + \frac{1}{k^+ - p'^+} R^{s,j} a_p^{s\dagger} c_k^j b_{p'}^{s\dagger} \Big|_{p'=k-p} \right. \\ &- \frac{1}{k^+ - p'^+} R^{s,j} a_p^{s\dagger} c_k^{j\dagger} a_{p'}^{-s} \Big|_{p'=p+k} - \frac{1}{k^+ + p'^+} R^{-s,j} b_p^s c_k^{j\dagger} b_{p'}^{-s\dagger} \Big|_{p'=p-k} \\ &\left. + \frac{1}{k^+ - p'^+} R^{-s,j} b_p^s c_k^j b_{p'}^{-s\dagger} \Big|_{p'=p+k} - \frac{1}{k^+ - p'^+} R^{-s,j} b_p^s c_k^{j\dagger} a_{p'}^s \Big|_{p'=k-p} \right) \end{aligned} \quad (\text{A11})$$

and

$$\begin{aligned}
 \frac{1}{2} \int d^3\mathbf{x} \psi_-^\dagger \gamma^j A^j \left(\frac{\psi_-}{\partial_-} \right) &= 2i \int \frac{dp^+ dp^\perp}{(2\pi)^3} \int \frac{dk^+ dk^\perp}{(2\pi)^3 \sqrt{k^+}} \sum_{s=\pm\frac{1}{2}} \sum_{j=1,2} \\
 &\times \left(\frac{1}{p'^+} R^{s,j} a_p^{s\dagger} c_k^j a_{p'}^{-s} \Big|_{p'=p-k} - \frac{1}{p'^+} R^{s,j} a_p^{s\dagger} c_k^j b_{p'}^{s\dagger} \Big|_{p'=k-p} + \frac{1}{p'^+} R^{s,j} a_p^{s\dagger} c_k^{j\dagger} a_{p'}^{-s} \Big|_{p'=p+k} \right. \\
 &\left. - \frac{1}{p'^+} R^{-s,j} b_p^s c_k^{j\dagger} b_{p'}^{-s\dagger} \Big|_{p'=p-k} - \frac{1}{p'^+} R^{-s,j} b_p^s c_k^j b_{p'}^{-s\dagger} \Big|_{p'=p+k} + \frac{1}{p'^+} R^{-s,j} b_p^s c_k^{j\dagger} a_{p'}^s \Big|_{p'=k-p} \right). \quad (\text{A12})
 \end{aligned}$$

As a cosmetic change, we write Eq. (A11) in terms of $\frac{1}{p^+}$. Combining terms we obtain

$$\begin{aligned}
 H_{\psi A} \supset &\int \frac{dp^+ dp^\perp}{(2\pi)^3} \int \frac{dk^+ dk^\perp}{(2\pi)^3 \sqrt{k^+}} \sum_{s=\pm\frac{1}{2}} \sum_{j=1,2} em \\
 &\times \left(\left(\frac{1}{p'^+} - \frac{1}{p^+} \right) R^{s,j} a_p^{s\dagger} c_k^j a_{p'}^{-s} \Big|_{p'=p-k} - \left(\frac{1}{p'^+} + \frac{1}{p^+} \right) R^{s,j} a_p^{s\dagger} c_k^j b_{p'}^{s\dagger} \Big|_{p'=k-p} \right. \\
 &+ \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) R^{s,j} a_p^{s\dagger} c_k^{j\dagger} a_{p'}^{-s} \Big|_{p'=p+k} - \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) R^{-s,j} b_p^s c_k^{j\dagger} b_{p'}^{-s\dagger} \Big|_{p'=p-k} \\
 &\left. - \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) R^{-s,j} b_p^s c_k^j b_{p'}^{-s\dagger} \Big|_{p'=p+k} + \left(\frac{1}{p'^+} + \frac{1}{p^+} \right) R^{-s,j} b_p^s c_k^{j\dagger} a_{p'}^s \Big|_{p'=k-p} \right). \quad (\text{A13})
 \end{aligned}$$

Once again we normal order. As an arbitrary convention we write the photon creation (annihilation) operators in the leftmost (rightmost) position. In this case, the delta functions in the normal ordering process produce terms that vanish. Rearranging terms to place Hermitian conjugates next to each other, the contribution to the normal-ordered Schrödinger-picture interaction Hamiltonian reads

$$\begin{aligned}
 :H_{\psi A}: \supset &\int \frac{dp^+ dp^\perp}{(2\pi)^3} \int \frac{dk^+ dk^\perp}{(2\pi)^3 \sqrt{k^+}} \sum_{s=\pm\frac{1}{2}} \sum_{j=1,2} em \\
 &\times \left(\left(\frac{1}{p'^+} - \frac{1}{p^+} \right) R^{s,j} c_k^{j\dagger} a_p^{s\dagger} a_{p'}^{-s} \Big|_{p'=p+k} + \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) R^{s,j} a_p^{s\dagger} a_{p'}^{-s} c_k^j \Big|_{p'=p-k} \right. \\
 &+ \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) R^{-s,j} c_k^{j\dagger} b_{p'}^{-s\dagger} b_p^s \Big|_{p'=p-k} + \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) R^{-s,j} b_{p'}^{-s\dagger} b_p^s c_k^j \Big|_{p'=p+k} \\
 &\left. + \left(\frac{1}{p'^+} + \frac{1}{p^+} \right) R^{-s,j} c_k^{j\dagger} b_p^s a_{p'}^s \Big|_{p'=k-p} - \left(\frac{1}{p'^+} + \frac{1}{p^+} \right) R^{s,j} a_p^{s\dagger} b_{p'}^{s\dagger} c_k^j \Big|_{p'=k-p} \right). \quad (\text{A14})
 \end{aligned}$$

To pass to the interaction picture, we evolve the creation and annihilation operators with U_0 [see Eq. (27)]:

$$\begin{aligned}
 U_0^\dagger a_p^s U_0 &= e^{-\frac{ip^- x^+}{2}} e^{-if(p)} a_p^s & U_0^\dagger a_p^{s\dagger} U_0 &= e^{\frac{ip^- x^+}{2}} e^{if(p)} a_p^{s\dagger} \\
 U_0^\dagger b_p^s U_0 &= e^{-\frac{ip^- x^+}{2}} e^{-ig(p)} b_p^s & U_0^\dagger b_p^{s\dagger} U_0 &= e^{\frac{ip^- x^+}{2}} e^{ig(p)} b_p^{s\dagger} \\
 U_0^\dagger c_k^j U_0 &= e^{-\frac{ik^- x^+}{2}} c_k^j & U_0^\dagger c_k^{j\dagger} U_0 &= e^{\frac{ik^- x^+}{2}} c_k^{j\dagger}.
 \end{aligned} \quad (\text{A15})$$

It turns out that $f(p)$ and $g(p)$ are the Volkov phases

$$f(p) = \int_0^{x^+} \frac{dy^+}{2} \left[e\mathcal{A}^- - \frac{2ep^i}{p^+} \mathcal{A}^i + \frac{e^2 \mathcal{A}^i \mathcal{A}^i}{p^+} \right] \quad g(p) = \int_0^{x^+} \frac{dy^+}{2} \left[-e\mathcal{A}^- + \frac{2ep^i}{p^+} \mathcal{A}^i + \frac{e^2 \mathcal{A}^i \mathcal{A}^i}{p^+} \right]. \quad (\text{A16})$$

These are part of the solutions to the Dirac equation in the presence of a background vector field \mathcal{A}_μ (specifically a null field, for which $\mathcal{F}\mathcal{F} = \mathcal{F}\tilde{\mathcal{F}} = 0$):

$$(i(\not{\partial} + ie\mathcal{A}) - m)\psi = 0. \quad (\text{A17})$$

Such solutions are called Volkov modes [62,72], and our interaction picture Hamiltonian can also be obtained by simply using (projected) expansions in Volkov modes. In any case, the interaction-picture interaction Hamiltonian is

$$\begin{aligned} U_0^\dagger : H_{\psi A} : U_0 \supset & \int \frac{dp^+ dp^\perp}{(2\pi)^3} \int \frac{dk^+ dk^\perp}{(2\pi)^3 \sqrt{k^+}} \sum_{s=\pm\frac{1}{2}} \sum_{j=1,2} em \\ & \times \left(\left(\frac{1}{p'^+} - \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(k^-+p^- - p'^-)} e^{i(f(p)-f(p'))} R^{s,j} c_k^{j\dagger} a_p^{s\dagger} a_{p'}^{-s} \Big|_{p'=p+k} \right. \\ & + \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(p^- - p'^- - k^-)} e^{i(f(p)-f(p'))} R^{s,j} a_p^{s\dagger} a_{p'}^{-s} c_k^j \Big|_{p'=p-k} \\ & + \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(k^-+p'^- - p^-)} e^{-i(g(p)-g(p'))} R^{-s,j} c_k^{j\dagger} b_{p'}^{-s\dagger} b_p^s \Big|_{p'=p-k} \\ & + \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(p'^- - p^- - k^-)} e^{-i(g(p)-g(p'))} R^{-s,j} b_{p'}^{-s\dagger} b_p^s c_k^j \Big|_{p'=p+k} \\ & + \left(\frac{1}{p'^+} + \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(k^- - p^- - p'^-)} e^{-i(g(p)+f(p'))} R^{-s,j} c_k^{j\dagger} b_p^{s\dagger} a_{p'}^s \Big|_{p'=k-p} \\ & \left. - \left(\frac{1}{p'^+} + \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(p^-+p'^- - k^-)} e^{i(f(p)+g(p'))} R^{s,j} a_p^{s\dagger} b_{p'}^{s\dagger} c_k^j \Big|_{p'=k-p} \right). \quad (\text{A18}) \end{aligned}$$

Finally, we discretize assuming a momentum lattice with spacing $\Delta p^i = \frac{2\pi}{L}$. Including factors associated with discretizing the ladder operators [see Eq. (25)], we obtain the final result:

$$\begin{aligned} U_0^\dagger : H_{\psi A} : U_0 \supset & \sum_{n_p^+, n_{p'}^+, n_p^2, n_{p'}^2} \sum_{n_k^+, n_k^1, n_k^2} \sum_{s=\pm\frac{1}{2}} \sum_{j=1,2} \frac{em}{\sqrt{k^+ L^3}} \\ & \times \left(\left(\frac{1}{p'^+} - \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(k^-+p^- - p'^-)} e^{i(f(p)-f(p'))} R^{s,j} c_k^{j\dagger} a_p^{s\dagger} a_{p'}^{-s} \Big|_{p'=p+k} \right. \\ & + \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(p^- - p'^- - k^-)} e^{i(f(p)-f(p'))} R^{s,j} a_p^{s\dagger} a_{p'}^{-s} c_k^j \Big|_{p'=p-k} \\ & + \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(k^-+p'^- - p^-)} e^{-i(g(p)-g(p'))} R^{-s,j} c_k^{j\dagger} b_{p'}^{-s\dagger} b_p^s \Big|_{p'=p-k} \\ & + \left(\frac{1}{p'^+} - \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(p'^- - p^- - k^-)} e^{-i(g(p)-g(p'))} R^{-s,j} b_{p'}^{-s\dagger} b_p^s c_k^j \Big|_{p'=p+k} \\ & + \left(\frac{1}{p'^+} + \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(k^- - p^- - p'^-)} e^{-i(g(p)+f(p'))} R^{-s,j} c_k^{j\dagger} b_p^{s\dagger} a_{p'}^s \Big|_{p'=k-p} \\ & \left. - \left(\frac{1}{p'^+} + \frac{1}{p^+} \right) e^{\frac{ix^+}{2}(p^-+p'^- - k^-)} e^{i(f(p)+g(p'))} R^{s,j} a_p^{s\dagger} b_{p'}^{s\dagger} c_k^j \Big|_{p'=k-p} \right). \quad (\text{A19}) \end{aligned}$$

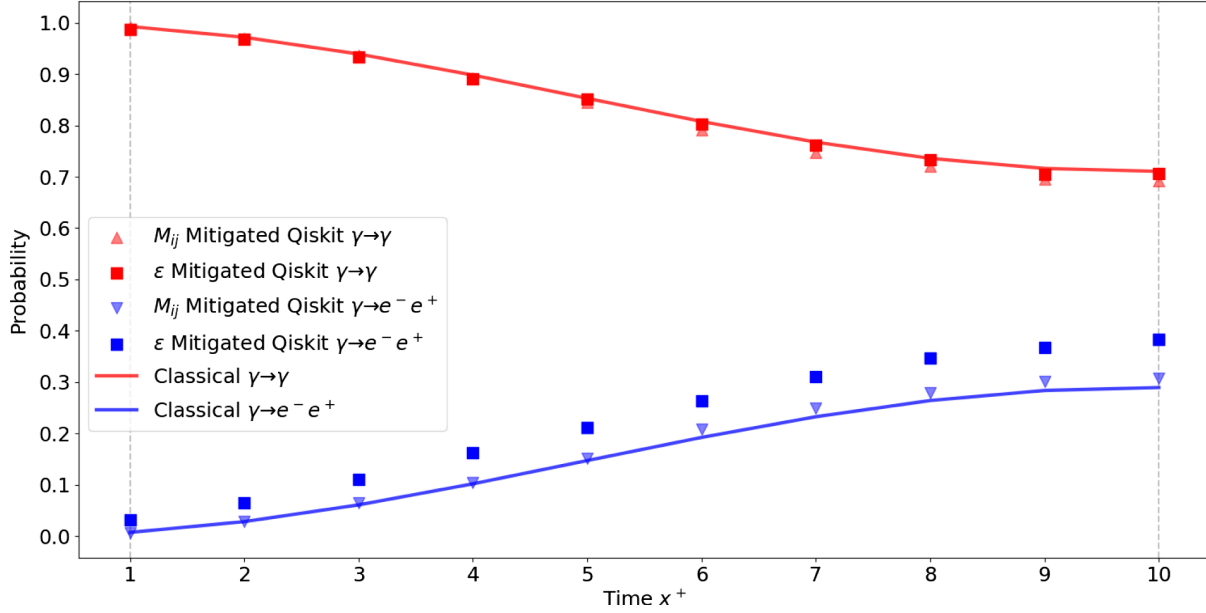
For general SFQED processes, the complete mode-expanded Hamiltonian can be derived similarly to the terms discussed in this Appendix. For the Breit-Wheeler simulation considered in the main body of the work, only the last two lines of Eq. (A19) were needed.

APPENDIX B: SYMMETRIC DEPOLARIZATION MITIGATION

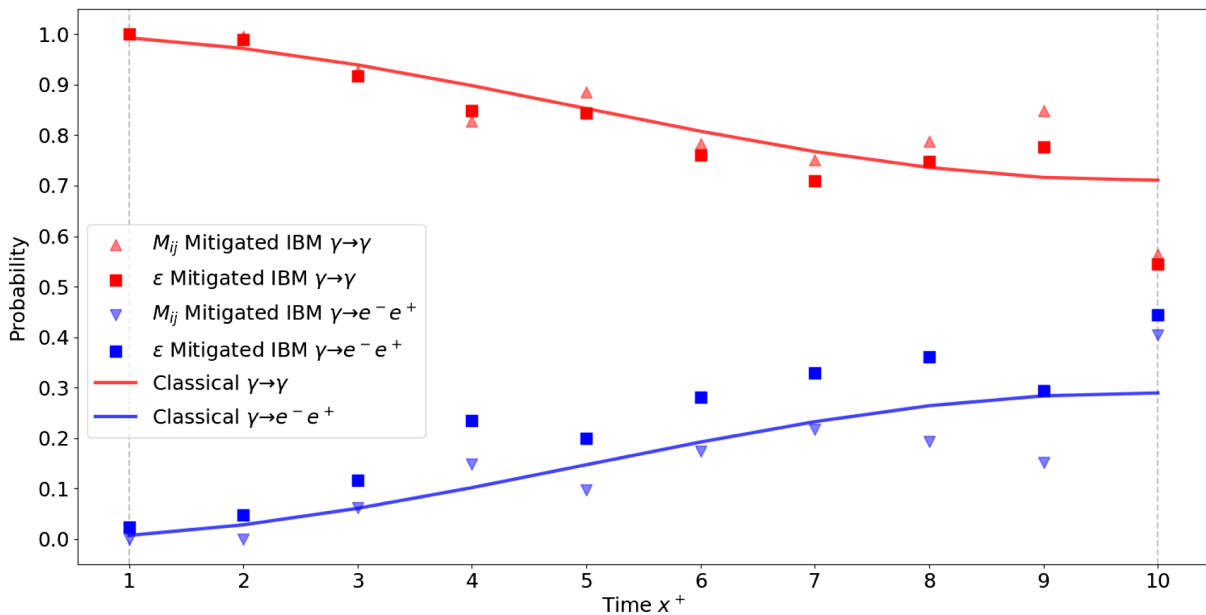
In this section we describe how our generalized depolarization mitigation improves upon the symmetric case. In the symmetric case, the transfer matrix $M_{ij} = \delta_{ij}(1 - 2^n \epsilon) + \epsilon$ depends on the single parameter ϵ . The noisy

probabilities P_i^{meas} for measuring the i th computational basis state are related to the ideal probabilities P_i^{true} by $P_i^{\text{meas}} = (1 - 2^n \epsilon)P_i^{\text{true}} + \epsilon$.

Previous implementations of symmetric depolarization mitigation use additional mitigation circuits to estimate the value of ϵ [19,57]. These circuits both mimic the noise of



(a) Mitigated data from Qiskit’s fake simulator shows an improvement of the transfer matrix (M_{ij}) mitigation over the ϵ mitigation.



(b) Mitigated data from IBM’s quantum hardware shows consistently better results from the use of a general transfer matrix in depolarization mitigation.

FIG. 14. Mitigated quantum results for nonlinear Breit-Wheeler pair production. Square data represent the use of single-parameter (ϵ) depolarization mitigation. Triangle data represent the use of a more general transfer matrix (M_{ij}) in depolarization mitigation.

the original physics circuit of interest and have a known noiseless final state. Call p_i^{meas} the noisy probabilities from the mitigation circuit and p_i^{true} the known ideal probabilities. One can then estimate ϵ via

$$\epsilon = \frac{1}{2^n} \frac{p_i^{\text{true}} - p_i^{\text{meas}}}{p_i^{\text{true}} - \frac{1}{2^n}}. \quad (\text{B1})$$

This allows us to calculate the ideal physics probability

$$P_i^{\text{true}} = P_i^{\text{meas}} \left(\frac{p_i^{\text{true}} - \frac{1}{2^n}}{p_i^{\text{meas}} - \frac{1}{2^n}} \right) - \frac{1}{2^n} \left(\frac{p_i^{\text{true}} - p_i^{\text{meas}}}{p_i^{\text{meas}} - \frac{1}{2^n}} \right). \quad (\text{B2})$$

It is possible to write this in the form used by Refs. [19,57]:

$$\frac{P_i^{\text{true}} - \frac{1}{2^n}}{P_i^{\text{meas}} - \frac{1}{2^n}} = \frac{p_i^{\text{true}} - \frac{1}{2^n}}{p_i^{\text{meas}} - \frac{1}{2^n}}. \quad (\text{B3})$$

In practice any computational basis state can be used to estimate ϵ in the mitigation circuit. In this study the

mitigation circuit will consist of the state $|000\rangle$ evolved by the identity operator in the form of forwards and backwards time evolution, similar to the setup shown in Fig. 12. Evidently, $p^{\text{true}} = 1$. However, the mitigation circuit measures the noise for “twice the physics circuit,” so we input $\sqrt{p^{\text{meas}}}$ into Eq. (B3).

Using Qiskit’s FakeNairobi2 noisy simulator, we can see a comparison between using an ϵ -mitigated simulation versus a general M_{ij} -mitigated simulation. Figure 14(a), shows near 30% improvements in M_{ij} mitigation over ϵ mitigation of $\gamma \rightarrow e^-e^+$ data. These results encouraged us to use the generalized depolarization mitigation in our final quantum simulation.

Figure 14(b) shows real quantum data from IBM_NAIROBI. The success of our asymmetric depolarization mitigation is more varied than the Qiskit simulation, but we still see near 30% improvements in asymmetric mitigation over symmetric mitigation of $\gamma \rightarrow e^-e^+$ probabilities in the middle of the simulation. Regardless of the mitigation strategy, the final time step with the largest gate count suffers the most error.

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