Revisiting models that enhance $B^+ \to K^+ \nu \bar{\nu}$ in light of the new Belle II measurement

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Belle II recently reported the new measurement $\mathcal{B}(B^+ \to K^+\nu\bar{\nu}) = (2.3 \pm 0.7) \times 10^{-5}$ [I. Adachi *et al.*) (Belle-II Collaboration), [arXiv:2311.14647](https://arXiv.org/abs/2311.14647)] which is two times larger than their previous result (although consistent within errors) and about 2.7σ higher than the SM prediction. We reexamine new physics scenarios we discussed previously, which can enhance this rate to determine if they can accommodate the higher value reported in the new measurement. We use consistency with existing bounds on $B \to K^* \nu \bar{\nu}$, $b \to s\ell^+\ell^-$, $B \to D^{(*)}\ell\bar{\nu}$, and B_s mixing to limit possible explanations for the excess. For the case of lepton flavor violating neutrino couplings, we find that only two leptoquarks remain viable requiring a large $C_{9'}^{\tau\tau} = -C_{10'}^{\tau\tau}$. For models with different types of light dark matter particle pairs (scalar, fermion, or vector), the preliminary q^2 distribution from Belle II, which shows that the excess appears mostly for bins with $3 \le q^2 \le 7$ GeV² [I. Adachi *et al.* (Belle-II Collaboration), [arXiv:2311.14647\]](https://arXiv.org/abs/2311.14647), implies only the vector current operators with scalar or vector dark matter particles with masses in the hundreds of MeV can match the anomaly.

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I. INTRODUCTION

The $B \to K^{(*)} \nu \bar{\nu}$ decays are amongst the cleanest modes to search for new physics due to their well-controlled theoretical uncertainty in the standard model (SM). A recent measurement by the Belle II Collaboration finds the branching ratio [[1\]](#page-10-0),

$$
\mathcal{B}(B^+ \to K^+ \nu \bar{\nu})_{\text{exp}} = (2.3 \pm 0.7) \times 10^{-5}, \qquad (1)
$$

which is higher than SM expectation by about 2.7σ . This result is also about twice as large, but consistent within errors, with a previous Belle II combination

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 $\mathcal{B}(B^+ \to K^+\nu\bar{\nu})^{2021}_{\text{exp}} = (1.1 \pm 0.4) \times 10^{-5}$ [[2\]](#page-10-1), and with the new average $\mathcal{B}(B^+ \to K^+ \nu \bar{\nu})_{\text{exp}}^{\text{ave}} = (1.3 \pm 0.4) \times 10^{-5}$ $\mathcal{B}(B^+ \to K^+ \nu \bar{\nu})_{\text{exp}}^{\text{ave}} = (1.3 \pm 0.4) \times 10^{-5}$ $\mathcal{B}(B^+ \to K^+ \nu \bar{\nu})_{\text{exp}}^{\text{ave}} = (1.3 \pm 0.4) \times 10^{-5}$ [1].
Currently, these numbers suggest consistency with the Currently, these numbers suggest consistency with the SM prediction (subtracting the so-called tree-level contribution) [\[3](#page-10-2)],

$$
\mathcal{B}(B^+ \to K^+ \nu \bar{\nu})_{\rm SM} = (4.43 \pm 0.31) \times 10^{-6}, \qquad (2)
$$

but the new measurement is sufficiently intriguing to entertain the possibility of new physics affecting this decay [\[4](#page-10-3)–[10](#page-10-4)]. Note that this number agrees with the one we use for our numerical study below, obtained from flavio, $\mathcal{B}(B^+ \to K^+\nu\bar{\nu})_{\text{SM}} = (4.4 \pm 0.6) \times 10^{-6}$. It has been noted that the Cabibbo-Kobayashi-Maskawa (CKM) dependence that the Cabibbo-Kobayashi-Maskawa (CKM) dependence of the SM prediction, which can hide new physics (NP), can be bypassed by considering certain ratios instead $[11]$ $[11]$ $[11]$.

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¹The SM value we use is a bit lower than the one quoted in $[1]$ due to the use of different CKM parameters; we consistently use $|V_{tb}V_{ts}^*| = (3.93 \pm 0.10) \times 10^{-2}$ whereas [\[12\]](#page-10-6), quoted in [[1\]](#page-10-0), uses $|V_{tb}V_{ts}^*|$ = 0.04185(93).

We previously studied three different new physics scenarios that can enhance the rate for this mode; lepton flavor violating (LFV) neutrino couplings (possibly induced by leptoquarks) [[13](#page-10-7)], a light sterile neutrino [[13](#page-10-7)], or other invisible light particles (dark matter) [\[14\]](#page-11-0). These three cases exploit the fact that the neutrinos (or their flavor) are not detected so they could be mimicked by other unseen particles.

In this paper, we revisit those scenarios with the new (high) measurement in mind. Specifically, we want to explore whether this large value of $\mathcal{B}(B^+ \to K^+ \nu \bar{\nu})$ (or indeed $B^+ \rightarrow K^+ + \cancel{E}$ can be accommodated in any of those scenarios while satisfying the existing upper bounds on $\mathcal{B}(B \to K^* \nu \bar{\nu})$ (again, more generally, $B \to K^* + \bar{E}$). All these NP scenarios introduce correlations with other modes that also need to be considered.

Specifically, we will base our study on the ratio,

$$
R_{\nu\nu}^{K} \equiv \frac{\mathcal{B}(B^{+} \to K^{+} \nu \bar{\nu})}{\mathcal{B}(B^{+} \to K^{+} \nu \bar{\nu})_{\text{SM}}},\tag{3}
$$

using the numerical values obtained from both the new Belle II measurement and from the new average reported in Fig. 23 of [\[1\]](#page-10-0)

$$
(R_{\nu\nu}^K)_{\text{new Belle II}} = 5.3 \pm 1.7,\tag{4a}
$$

$$
(R_{\nu\nu}^K)_{\text{newaverage}} = 3.0 \pm 1.0. \tag{4b}
$$

We will refer to the first number as the "new 1σ range". The rate predicted for the corresponding neutral and charged pseudoscalar meson modes is the same if isospin is conserved, and the measurement of the neutral mode is at present less restrictive. To discuss correlations with $R_{K^*}^{\nu\nu}$ we use two numbers,

$$
R_{K^*}^{\nu\nu} = \frac{\mathcal{B}(B \to K^* \nu \bar{\nu})}{\mathcal{B}(B \to K^* \nu \bar{\nu})_{\rm SM}} \le 2.7 \quad \text{or} \quad 1.9. \tag{5}
$$

The first number (2.7) arises from the combined charged and neutral modes as directly quoted by the Belle Collaboration [\[15\]](#page-11-1). The second number reflects the fact that the predictions for the decay rates of the charged and neutral modes are the same in all the models we consider. This suggests using the strongest experimental constraint, which in this case occurs for the neutral mode [\[15\]](#page-11-1),

$$
\mathcal{B}(B^0 \to K^{0*} \nu \bar{\nu}) \le 1.8 \times 10^{-5} \quad (90\% \text{ C.L.}). \quad (6)
$$

In combination with the corresponding SM prediction, this results in the second number (1.9) in Eq. [\(5\).](#page-1-0)

The overall picture suggested by these numbers is that the new Belle II measurement does not imply significant changes to the averages yet. However, it invites us to entertain the possibility of a larger $R_K^{\nu\nu}$ and to study its implications for phenomenology, in particular for the implied $R_K^{\nu\nu} > R_{K^*}^{\nu\nu}$. This is the subject of this paper.

II. ALTERING THE PROPERTIES OF THE NEUTRINOS

We parametrize possible new physics entering these decays through an effective Hamiltonian at the b mass scale with dimension-six operators responsible for $b \rightarrow s \nu \bar{\nu}$ in the low-energy effective field theory approach (LEFT) [\[16](#page-11-2)[,17\]](#page-11-3). The effective theory originates in extensions of the SM containing new particles at or above the electroweak scale that have been integrated out but also allowing for the possibility of light right-handed neutrinos. Specifically, we consider the effective Hamiltonian,

$$
\mathcal{H}_{\text{NP}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^{\star} \frac{e^2}{16\pi^2} \sum_{ij} \left(C_L^{ij} \mathcal{O}_L^{ij} + C_R^{ij} \mathcal{O}_R^{ij} + C_L^{ij} \mathcal{O}_L^{ij} + C_L^{ij} \mathcal{O}_L^{ij} + C_R^{ij} \mathcal{O}_R^{ij} \right) + \text{H.c.}, \tag{7}
$$

including the operators

$$
\mathcal{O}_L^{ij} = (\bar{s}_L \gamma_\mu b_L)(\bar{\nu}_i \gamma^\mu (1 - \gamma_5) \nu_j), \n\mathcal{O}_R^{ij} = (\bar{s}_R \gamma_\mu b_R)(\bar{\nu}_i \gamma^\mu (1 - \gamma_5) \nu_j),
$$
\n(8a)

$$
\mathcal{O}_L^{ij} = (\bar{s}_L \gamma_\mu b_L)(\bar{\nu}_i \gamma^\mu (1 + \gamma_5) \nu_j),
$$

\n
$$
\mathcal{O}_R^{ij} = (\bar{s}_R \gamma_\mu b_R)(\bar{\nu}_i \gamma^\mu (1 + \gamma_5) \nu_j).
$$
\n(8b)

The Wilson coefficients in Eq. [\(8\)](#page-1-1) are defined so that they only contain NP contributions. The SM contributes only to C_L^{ii} and is accounted for separately [[18](#page-11-4),[19](#page-11-5)],

$$
C_{L \text{SM}} = -\frac{X(x_t)}{s_W^2}, \qquad X(x_t) = 1.469 \pm 0.017. \quad (9)
$$

New interactions that conserve lepton number and have no new light particles can generate $O_{L,R}^{ij}$. In contrast, the operators $\mathcal{O}_{L,R}^{ij}$ are present when there are light righthanded neutrinos. Off diagonal (lepton flavor violating) operators occur for example, in models with leptoquarks. Other possibilities, such as scalar or tensor operators, are not discussed here.

Equation [\(7\)](#page-1-2) has contributions to $B \to K^{(*)} \nu \bar{\nu}$ that interfere with the SM, $\mathcal{O}_{L,R}^{ii}$; and others that do not, $\mathcal{O}_{L,R}^{i\neq j}$ and $\mathcal{O}_{L,R}^{ij}$. In $B \to K \nu \bar{\nu}$ only the vector current enters the hadronic matrix element making the contributions to the rate from \mathcal{O}_L^{ij} and \mathcal{O}_R^{ij} (or from \mathcal{O}_L^{ij} and \mathcal{O}_R^{ij}) the same. In $B \to K^* \nu \bar{\nu}$ both the vector and axial-vector currents enter the hadronic matrix element resulting in different contributions from \mathcal{O}_L^{ij} and \mathcal{O}_R^{ij} as well as from \mathcal{O}_L^{ij} and \mathcal{O}_R^{ij} .

The different neutrino chirality eliminates interference between the contributions from primed and unprimed operators for massless neutrinos. Expressions for $R^{\nu\nu}_{K^{(*)}}$ can be determined analytically [\[20\]](#page-11-6) and we find

$$
R_K^{\nu\nu} = 1 + \frac{2C_{L \text{SM}}}{3|C_{L \text{SM}}|^2} \sum_i \text{Re}\left(C_L^{ii} + C_R^{ii}\right) + \frac{1}{3|C_{L \text{SM}}|^2} \sum_{ij} \left(|C_L^{ij} + C_R^{ij}|^2 + |C_L^{ij} + C_R^{ij}|^2\right),\tag{10a}
$$

$$
R_{K^*}^{\nu\nu} = 1 + \frac{2C_{L \, \text{SM}}}{3|C_{L \, \text{SM}}|^2} \sum_{i} \text{Re}(C_L^{ii}) + \frac{1}{3|C_{L \, \text{SM}}|^2} \sum_{ij} \left(|C_L^{ij}|^2 + |C_R^{ij}|^2 + |C_L^{ij}|^2 + |C_R^{ij}|^2 \right) - 2\eta \left[\frac{C_{L \, \text{SM}}}{3|C_{L \, \text{SM}}|^2} \sum_{i} \text{Re}(C_R^{ii}) + \frac{1}{3|C_{L \, \text{SM}}|^2} \sum_{ij} \text{Re}(C_L^{ij} C_R^{*ij} + C_L^{'ij} C_R^{*ij}) \right],
$$
(10b)

where

$$
\eta \equiv \frac{F_{-}}{F_{+}}, \qquad F_{\pm} \equiv \int_{0}^{(1-\sqrt{x_K})^2} dx \lambda^{\frac{1}{2}}(1, x_K, x) \left[x A_1^2 + \frac{32x_K}{(1+\sqrt{x_K})^2} A_{12}^2 \pm \frac{x \lambda (1, x_K, x)}{(1+\sqrt{x_K})^4} V_0^2 \right], \tag{11}
$$

with $x \equiv q^2/m_B^2$ and $x_K \equiv m_{K^*}^2/m_B^2$. Numerically, $\eta =$ 0.63 ± 0.09 when using the recent light cone sum rule (LCSR) result for the form factors A_1 , A_{12} , V_0 [[21](#page-11-7)],
 $\frac{2C_{LSM}}{2|C|} = -(0.105 \pm 0.001)$ and $\frac{1}{2|C|} = 2.0083 \pm 0.0002$. $\frac{2C_{LSM}}{3|C_{LSM}|^2}$ = -(0.105 ± 0.001) and $\frac{1}{3|C_{LSM}|^2}$ = 0.0083 ± 0.0002. With these numbers, this expression agrees with the result presented in [\[13\]](#page-10-7) that was obtained by using flavio [[22](#page-11-8)]. It is interesting to note that the difference $R_K^{\nu\nu} - R_K^{\nu\nu}$ involves only the interference terms between the leftand right-handed currents,

$$
R_K^{\nu\nu} - R_{K^*}^{\nu\nu} = 2(1+\eta) \left[\frac{C_{L \text{ SM}}}{3|C_{L \text{ SM}}|^2} \sum_i \text{Re}(C_R^{ii}) + \frac{1}{3|C_{L \text{ SM}}|^2} \sum_{ij} \text{Re}(C_L^{ij} C_R^{*ij} + C_L^{'ij} C_R^{*ij}) \right].
$$
 (12)

We begin by examining the correlations between $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ implied by Eq. [\(10b\)](#page-2-0) in light of the new 1σ range. These are illustrated in Fig. [1](#page-3-0) for a scan of the 12 parameter space C_L^{ij} , C_R^{ij} in light gray in the left panel, and for the 12 parameter space C_L^{ij} , C_R^{ij} in darker gray in the left panel.² The primed operators do not interfere with the SM so the points in the darker shaded region satisfy $R^{\nu\bar{\nu}}_{K^{(*)}} \geq 1$. In addition, since there is no interference between the primed and unprimed coefficients (we ignore neutrino masses), it will suffice to look at the two cases separately. We will focus on groups that reproduce the new 1σ range of $R_K^{\nu\nu}$, $3.6 \le R_K^{\nu\nu} \le 7$, or the new average, $2 \le R_K^{\nu\nu} \le 4$ and these regions are marked by the vertical solid (dashed) black lines. We are also interested in groups that satisfy

 $R_{K^*}^{\nu\nu} \leq 2.7$ or $R_{K^*}^{\nu\nu} \leq 1.9$ and these two regions are bounded by the horizontal dashed red lines. For comparison we show the 3σ SM region in green (only theoretical parametric errors as estimated by flavio). The figure shows that it is possible to reproduce values of $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ in any of the desired regions with either the 12 unprimed or primed coefficients, albeit with increasing difficulty for the region where $R_K^{\nu\nu}$ matches the new measurement and $R_{K^*}^{\nu\nu} \leq 1.9$. The solid diagonal line, where $R_K^{\nu\nu} = R_{K^*}^{\nu\nu}$, will be of interest for certain cases that we discuss below [[13](#page-10-7)].

On the right panel of Fig. [1](#page-3-0) we consider only the six C_R^{ij} coefficients, as these are the ones induced by certain leptoquarks that can satisfy both the $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ constraints as discussed below. The red and blue regions of interest are the targets for our study and the color code will serve to map predictions for other modes with parameters producing values of $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ in the respective region.

To generate the operators with only left-handed neutrinos, $\overline{O_{L,R}^{ij}}$, we consider scalar S and vector V leptoquarks coupling to SM fermions as [[13](#page-10-7),[23](#page-11-9)–[26](#page-11-10)],

$$
\mathcal{L}_{S} = \lambda_{LS_0} \bar{q}_L^c i \tau_2 \ell_L S_0^{\dagger} + \lambda_{L \tilde{S}_{1/2}} \bar{d}_R \ell_L \tilde{S}_{1/2}^{\dagger} \n+ \lambda_{LS_1} \bar{q}_L^c i \tau_2 \vec{\tau} \cdot \vec{S}_1^{\dagger} \ell_L + \text{H.c.},
$$
\n(13a)

$$
\mathcal{L}_V = \lambda_{LV_{1/2}} \bar{d}_R^c \gamma_\mu \mathcal{C}_L V_{1/2}^{\dagger \mu} + \lambda_{LV_1} \bar{q}_L \gamma_\mu \vec{\tau} \cdot \vec{V}_1^{\dagger \mu} \mathcal{C}_L + \text{H.c.} \quad (13b)
$$

The leptoquark fields appearing in Eq. [\(13\)](#page-2-1), and their transformation properties under the SM group, are

$$
S_0^{\dagger} = S_0^{1/3} \colon (\bar{3}, 1, 1/3), \quad \tilde{S}_{1/2}^{\dagger} = (\tilde{S}_{1/2}^{-1/3}, \tilde{S}_{1/2}^{2/3}) \colon (3, 2, 1/6),
$$
\n(14a)

$$
\vec{\tau} \cdot \vec{S}_1^{\dagger} = \begin{pmatrix} S_1^{1/3} & \sqrt{2} S_1^{4/3} \\ \sqrt{2} S_1^{-2/3} & -S_1^{1/3} \end{pmatrix} : (\bar{3}, 3, 1/3), \tag{14b}
$$

²There are only six parameters for each of $C_{L,R}^{(l)}$ because the expressions in Eq. [\(10b\)](#page-2-0) are symmetric in $i \leftrightarrow j$. In addition, we assume CP is conserved and use real Wilson coefficients for our scans.

FIG. 1. The correlation between $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ scanning the 12 parameters C_L^{ij} , C_R^{ij} is shown in the left panel as the shaded light gray region. The corresponding 12 parameter scan in C_L^{ij} , C_R^{ij} is shown as the shaded dark gray region. The figure highlights in green the 3 σ SM prediction. The right panel shows the corresponding scan including only the six parameters C_R^{ij} . The points highlighted in blue or red fall within the new 1σ range of $R_K^{\nu\nu}$ (also marked by the vertical solid black lines). Additionally, the blue (red) points satisfy $R_{K^*}^{\nu\nu} \leq 2.7(1.9)$, respectively (values marked by the horizontal dashed red lines). The diagonal black line marks $R_K = R_{K^*}$.

$$
V_{1/2}^{\dagger} = (V_{1/2}^{1/3}, V_{1/2}^{4/3}) : (\bar{3}, 2, 5/6),
$$

$$
\vec{\tau} \cdot \vec{V}_{1}^{\dagger} = \begin{pmatrix} V_{1}^{2/3} & \sqrt{2}V_{1}^{5/3} \\ \sqrt{2}V_{1}^{-1/3} & -V_{1}^{2/3} \end{pmatrix} : (3, 3, 2/3). \quad (14c)
$$

Exchange of S_0 , S_1 , or V_1 , generates the C_L coefficients only, whereas exchange of $\tilde{S}_{1/2}$ or $V_{1/2}$ generates only the C_R coefficients. We find that it is not possible to satisfy the bound on $R_{K^*}^{\nu\nu}$, Eq. [\(5\),](#page-1-0) with only $C_L^{i\bar{j}}$ coefficients because they predict $R_K^{\nu\nu} = R_K^{\nu\nu}$ and thus fall along the diagonal line shown in the left panel of Fig. [1](#page-3-0). Models with only S_0 , S_1 , or V_1 are thus incompatible with the new Belle II measurement. The same figure shows that they could satisfy the new average value of $R_K^{\nu\nu}$ but would only satisfy the weaker bound on $R_{K^*}^{\nu\nu}$.

These three leptoquarks are also constrained by modes with charged leptons. S_0 induces charged current operators and contributes to the "charged B anomalies" in $R_{D^{(*)}}$. If we choose its parameters to reach the largest point consistent with the new average value of $R_K^{\nu\nu}$ while satisfying the looser constraint on $R_{K^*}^{\nu\nu}$, $R_K^{\nu\nu} = R_{K^*}^{\nu\nu} \sim 2.7$, then we find that it can only enhance $R_{D^{(*)}}$ by about 6% over their SM value. This is insufficient to explain the current discrepancy as reported by HFLAV, $r_D = 1.19 \pm 0.0003$ 0.10, $r_{D^*} = 1.12 \pm 0.06$ [\[27](#page-11-11)], $r_{D^{(*)}} = R_{D^{(*)}} / (R_{D^{(*)}})_{\text{SM}}$. Conversely, tuning the parameters of S_0 to $r_{D^{(*)}} = 1.12$ results in $R_K^{\nu\nu} = R_{K*}^{\nu\nu} \gtrsim 14$.

The cases of S_1 and V_1 are also constrained by the "neutral B anomalies" as they generate the correlated coefficients [[13](#page-10-7)],

$$
S_1: C_9^{ij} = -C_{10}^{ij} = 2C_L^{ij}, \qquad (15a)
$$

$$
V_1: C_9^{ij} = -C_{10}^{ij} = \frac{1}{2} C_L^{ij}.
$$
 (15b)

Recent global fits suggest that the observables in $b \to s\ell^+\ell^-$ are best described with $C_9^{\mu\mu} \sim C_9^{ee} \sim -1$ implying $R_K^{\nu\nu} = R_{K^*}^{\nu\nu} \lesssim 1.1$ for S_1 and $R_{K^{(*)}}^{\nu} \lesssim 1.5$ for V_1 .

For the case where new physics appears only with C_R coefficients, on the other hand, there is a large region of parameter space that simultaneously satisfies the new 1σ range of $R_K^{\nu\nu}$ and the 90% C.L. upper bound on $R_{K^*}^{\nu\nu}$ as seen on the right panel of Fig. [1.](#page-3-0) However, if we restrict ourselves to models where only the off-diagonal (in lepton flavor) coefficients, $C_R^{i \neq j}$, are not zero, it is impossible to satisfy both constraints simultaneously because this scenario also results in $R_K^{\nu\nu} = R_{K^*}^{\nu\nu}$.

We are left with two leptoquarks that can satisfy both the new 1σ range of $R_K^{\nu\nu}$ and the 90% C.L. upper bounds on $R_{K^*}^{\nu\nu}$. They are $\tilde{S}_{1/2}$ and $V_{1/2}$ with both diagonal and off-diagonal flavor couplings. They generate the coefficients [\[13\]](#page-10-7),

$$
C_R^{ij} = C_{9'}^{ij} = -C_{10'}^{ij}
$$

=
$$
\frac{\pi}{\sqrt{2}\alpha G_F V_{tb} V_{ts}^*} \left(-\frac{\lambda_{L\bar{S}_{1/2}}^{2j} \lambda_{L\bar{S}_{1/2}}^{*3i}}{2m_{S_{1/2}}^2} + \frac{\lambda_{LV_{1/2}}^{3j} \lambda_{LV_{1/2}}^{*2i}}{m_{V_{1/2}}^2} \right),
$$
 (16)

and are thus correlated with other modes with charged leptons [\[28,](#page-11-12)[29](#page-11-13)]. The flavor diagonal entries for the first two generations affect the $b \rightarrow s\ell^+\ell^-$ processes. Recent global fits³ to $b \rightarrow s\ell^+\ell^-$ observables, in addition to preferring $C_9^{\mu\mu} \sim C_9^{ee} \sim -1$, also allow $C_{9'}^{ii}$ and $C_{10'}^{ii}$ to be nonzero, but much smaller than one. This makes it obvious that $\tilde{S}_{1/2}$ and $V_{1/2}$ are not a preferred solution for the "neutral B anomalies". $\tilde{S}_{1/2}$ and $V_{1/2}$ can still

³After the corrected LHCb value of $R_{K(K^*)}$ [\[30\]](#page-11-14) (for a review with references see [\[31\]](#page-11-15)).

FIG. 2. Left panel: Scan of parameters C_R^{ij} where the only nonzero diagonal term is $C_R^{\tau\tau}$ with the same color code as in Fig. [1](#page-3-0). Center panel: A selected slice of a projection in parameter space illustrating the location of the red and blue clusters in the six-parameter space of C_R^{ij} . As this figure suggests, these clusters concentrate in regions where at least one of the diagonal entries $|C_R^{ij}|$ is near 10. Right panel: Parameter region allowed when only $C_R^{\tau\tau}$ and $C_R^{\mu\tau}$ are nonzero, we have added in this panel the hashed regions showing the corresponding solutions when the average value of $R_K^{\nu\nu}$ is used.

FIG. 3. Mapping of the parameter space selected in Fig. [2](#page-4-0) into predictions for other B meson decay modes. The solid black line for the lepton flavor conserving processes marks the SM prediction.

modify $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ but they are constrained to produce small values of $C_{9'}^{\mu\mu,ee}$ and $C_{10'}^{\mu\mu,ee}$ and additional NP would be needed to provide the preferred values of C_9 .

A scan of the six-parameter space for a symmetric C_R^{ij} is shown on the left panel of Fig. [2,](#page-4-0) marking the regions that satisfy Eq. [\(4\)](#page-1-3) and Eq. [\(5\)](#page-1-0) as before. With the aid of tools described in [[32](#page-11-16)–[34\]](#page-11-17) we can examine the parameter region responsible for the blue and red regions. We show in the center panel a slice of a projection mostly onto the diagonal elements C_R^{ii} (axes labeled $x_{1,2,3}$). The slice is thin in the orthogonal space, mostly the off-diagonal elements $C_R^{i \neq j}$. The inset represents the projection matrix corresponding to this view. The visualization indicates that the red and blue points concentrate in regions where at least one diagonal C_R^{ii} element is large, $\mathcal{O}(10)$. Given that the global fits $b \rightarrow s\ell^+\ell^-$ do not admit solutions with

large $C_{9',10'}^{ee,\mu\mu}$ we conclude that the case with large $C_R^{\tau\tau}$ is the only viable solution. 4 For further illustration, we show the viable two-parameter space $C_R^{\mu\tau} - C_R^{\tau\tau}$ region on the right panel of the same figure (the color code is the same throughout).

Quantitatively, in this case, we find a lower bound of $R_{K^*}^{\nu \nu} \geq 1.5$ for the red region. It corresponds to $C_R^{\mu \tau} \sim C_R^{\tau \mu} \sim 0$ and $C_R^{\tau\tau} \sim -13.2$. Similar solutions exist for small but nonzero C_R^{ee} , $C_R^{\mu\mu} \sim \mathcal{O}(0.1)$ in combination with a large $C_R^{\tau\tau} \sim \mathcal{O}(10)$.

The Wilson coefficients generated by either one of $\tilde{S}_{1/2}$ or $V_{1/2}$ [Eq. [\(16\)\]](#page-3-1), imply large rates for other B decay modes when chosen to produce values of $R^{\nu\nu}_{K^{(*)}}$ in the

⁴A short video illustrating the tools arriving at the center panel of Fig. [2](#page-4-0) can be found in the Supplemental Material [\[35\]](#page-11-18).

red/blue regions as we illustrate in Fig. [3](#page-4-1). We see that the CLFV branching fractions $\mathcal{B}(B^+ \to K^+ \mu^- \tau^+)$ and $\mathcal{B}(B_s \to \tau^+)$ $\mu^{-} \tau^{+}$) can reach values of 10⁻⁵, which are within factors of two below their current experimental upper limits; $2.8 \times 10^{-5} (90\%)$ and $4.2 \times 10^{-5} (95\%)$ [\[36\]](#page-11-19) respectively. Similarly, the branching ratios for the lepton flavor conserving modes $\mathcal{B}(B^+\to K^+\tau^+\tau^-)$ and $\mathcal{B}(B_s\to \tau^+\tau^-)$ can reach values of a few times 10^{-5} , whereas their current experimental upper limits are $\mathcal{B}(B^+ \to K^+\tau^+\tau^-)$ < 2.25 × $10^{-3}(90\%)$ and $\mathcal{B}(B_s \to \tau^+\tau^-) \leq 6.8 \times 10^{-3}(95\%)$ [[36](#page-11-19)]. These represent large enhancements over the SM (shown as a vertical black line in Fig. [3](#page-4-1) and could rule out this possibility in the future.

The case of leptoquarks was also discussed by $[5]$ $[5]$ $[5]$.⁵ They reach similar but not identical conclusions to us. For S_0 , S_1 or V_1 (called S_1 , S_3 and V_3 in that reference) we agree that only C_L coefficients are generated and this results in $R_K^{\nu\nu} = R_{K^*}^{\nu\nu}$. Our Fig. [1](#page-3-0) quantifies just how much this deviates from the current measurements. We also agree that global fits to $b \rightarrow s\ell^+\ell^-$ processes imply that significant deviations from the SM would correlate with τ -lepton flavor. We disagree with the conclusions reached for the case of $\tilde{S}_{1/2}$ or $V_{1/2}$ (called \tilde{S}_2 , V_2 in [\[5\]](#page-10-8)), which induce only C_R . We find that these two can explain the measurements of $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ as shown in Fig. [2](#page-4-0) for a certain region of parameter space, contrary to what is claimed in [[5\]](#page-10-8).

We briefly turn our attention to models with a light sterile neutrino (that couples to SM fields via a Z') [[37](#page-11-20),[38](#page-11-21)]. This model produces nonzero C_L^{ij} , C_R^{ij} as those in the left panel of Fig. [1.](#page-3-0) In this case, Eq. [\(10b\)](#page-2-0) implies that in the region where $3.6 \le R_K^{\nu\nu} \le 7$ there is a lower bound $R_{K^*}^{\nu\nu} \ge 1.47$. Both $C_L^{ij} \neq 0$ and $C_R^{ij} \neq 0$ are needed to deviate from $R_K^{\nu\nu} = R_{K^*}^{\nu\nu}$ and the parameter scan shows that there could be solutions of this type, the region satisfying both $R_K^{\nu\nu}$, $R_{K^*}^{\nu\nu}$ constraints is shown in dark gray in Fig. [1.](#page-3-0) However, the specific model described in [\[13\]](#page-10-7) is severely constrained by both $Z - Z'$ mixing and by B_s mixing [[39](#page-11-22),[40](#page-11-23)]. When we consider these two constraints, the model predicts $R_K^{\nu\nu} \approx R_{K^*}^{\nu\nu} \lesssim 2$, barely reaching the lower end of the new $R_K^{\mu\nu}$ average.

III. LIGHT DARK MATTER SCENARIO FOR $B \to K +$ *invisible*

A large $\mathcal{B}(B^+ \to K^+ \nu \bar{\nu})$ such as the one reported by Belle II can also be caused by invisible particles beyond the SM. In this case light dark matter (DM) pairs take the place of the neutrinos, and we now examine this possibility.⁶ Since the spin of the possible DM particles is not known, we consider three cases; a spin-0 scalar ϕ , a spin-1/2 fermion χ , or a spin-1 vector X , respectively. We work within the framework of LEFT [[16](#page-11-2),[17](#page-11-3)], so we only impose the SM broken-phase symmetry $SU(3)_c \times U(1)_{em}$ on the effective operators. In this way, the LEFT framework covers scenarios that contain both light DM as well as new weak scale mediators that have been integrated out. The complete LEFT operator basis with a pair of light DM fields at leading order was recently provided by us in Ref. [[14](#page-11-0)]. In the following, we will adopt the notation in that paper and list the relevant FCNC local quark-DM operators mediating $B^+ \rightarrow K^+ +$ *invisible* decays for all cases.

For this discussion, it is not convenient to use the ratio $R_{\nu\nu}^{K}$ of Eq. [\(1\).](#page-0-4) Instead, it is better to look at the difference between the measurement and the SM, and refer to it as the "NP window". Using the new Belle II measurement this is

$$
\mathcal{B}(B^+ \to K^+ + invisible)_{\text{NP}}
$$

\n
$$
\equiv \mathcal{B}(B^+ \to K^+ \nu \bar{\nu})_{\text{exp}} - \mathcal{B}(B^+ \to K^+ \nu \bar{\nu})_{\text{SM}}
$$

\n
$$
= (1.9 \pm 0.7) \times 10^{-5}, \tag{17}
$$

and it becomes $(1.0 \pm 0.4) \times 10^{-5}$ using instead the new average experimental value.

In keeping with our previous discussion, we explore the possibility of attributing the new Belle II result to light dark particles while satisfying other constraints. We thus illustrate the implications of requiring the new contribution to the rate $B^+ \rightarrow K^+ + invisible$ to fall within the 1σ range of Eq. [\(17\).](#page-5-0) At the same time, the effective operators studied below will also contribute to the decays $B^0 \rightarrow K^0 +$ *invisible* and $B^{+(0)} \rightarrow K^{*+(0)} +$ *invisible*. There exist experimental upper bounds on these modes (at 90% C.L.) that we will use to constrain the parameter space. A final consideration relevant to this scenario is the observation by Belle II that the excess of events is predominantly concentrated in the region with $3 \le q^2 \le 7$ GeV² [\[1\]](#page-10-0).

It has been pointed out recently that the excess over the SM from the Belle II measurement is model dependent [\[43](#page-11-24),[44\]](#page-11-25). A recasting of the Belle II analysis for the different types of particles, different masses, and different operators that we discuss below is beyond the scope of the present study. Each choice of particle, mass or operator leads to a different kinematic (q^2) distribution and this can be exploited by future experimental analysis to differentiate between benchmarks. In this study we rely on the branching ratio under the assumption that the number reported by Belle II is model independent. Clearly this is not exactly true, but a reliable model independent branching ratio can only be extracted by the experimental collaboration (if at all). Our results should be interpreted as a comparative guide to different scenarios to select preferred benchmarks that can be studied further if the excess persists.

⁵Which appeared as we were finalizing our manuscript. ⁶Similar considerations for the charm decay can be found in [[41](#page-11-26),[42](#page-11-27)].

FIG. 4. The pink [purple] region is the parameter space that could explain the recent Belle II excess [new average] with scalar DM for scalar (vector) quark current operators $\mathcal{O}_{q\phi}^{S, sb}(\mathcal{O}_{q\phi}^{V, sb})$. The gray reg plots by color lines.

A. Scalar DM case

We begin with the spin-0 scalar DM case. The leadingorder operators mediating $B^+ \to K^+ \phi \phi$ consist of quark scalar and vector currents with s, b flavors coupled to two scalar fields,

$$
\mathcal{O}_{q\phi}^{S,sb} = (\bar{s}b)(\phi^{\dagger}\phi), \quad \mathcal{O}_{q\phi}^{V,sb} = (\bar{s}\gamma^{\mu}b)(\phi^{\dagger}i\overset{\leftrightarrow}{\partial}_{\mu}\phi), (\times). \quad (18)
$$

The symbol " (x) " indicates an operator that vanishes for real scalar fields, and $\phi^{\dagger} \overleftrightarrow{\partial}_{\mu} \phi = \phi^{\dagger} (\partial_{\mu} \phi) - (\partial_{\mu} \phi^{\dagger}) \phi$. These operators occur at dimension six in the SMEFT framework [[45](#page-11-28)]. In addition to $B^+ \to K^+ \phi \phi$, the scalar current operator also induces the decay $B^0 \to K^0 \phi \phi$. The vector operator can also mediate the modes $B^0 \to K^{*0} \phi \phi$ and $B^+ \to K^{*+} \phi \phi$. The differential decay widths for these four modes have been given in [\[14\]](#page-11-0). For the hadronic form factors involved, we use the recent lattice calculation in [\[46\]](#page-11-29) for $B \to K$, and those from LCSR [\[21\]](#page-11-7) for $B \to K^*$. To quantify the allowed parameter space, we define an effective heavy scale associated with each operator as $C_{q\phi}^{S, sb} \equiv \Lambda_{\text{eff}}^{-1}$ and $C_{q\phi}^{V, sb} \equiv \Lambda_{\text{eff}}^{-2}$.

Figure [4](#page-6-0) shows our results for the operators in Eq. [\(18\)](#page-6-1). The green (solid and dashed) lines and the blue dashed line mark the lower limit on the scale Λ_{eff} associated with the scalar (vector) operator in the left (right) panel resulting from the upper bound on $B^{+(0)} \to K^{*+(0)} \phi \phi$ and $B^0 \rightarrow K^0 \phi \phi$, respectively. The pink region covers the parameter space in the m_{ϕ} - Λ _{eff} plane that accommodates the Belle II 1σ range. The purple region indicates the allowed parameter space once the new average number is adopted, the thin dark purple strip is the overlapping region. The gray area is excluded by the experimental upper limits and is mostly due to the neutral modes $B^0 \to K^{(*)0} \nu \bar{\nu}$ for these two operators. The scalar operator provides a larger parameter region consistent with the new Belle II measurement than the vector operator. The vector operator, however, results in a q^2 distribution more in tune with the reported excess, as we illustrate in Fig. [7](#page-8-0).

B. Fermion DM case

We now turn to fermion DM particles.⁷ Denoting the DM field by χ , there are six leading, dimension-six operators that are responsible for the $B^+ \to K^+ \chi \chi$ transition [\[14\]](#page-11-0),

$$
\mathcal{O}_{q\chi 1}^{S, sb} = (\bar{s}b)(\bar{\chi}\chi), \n\mathcal{O}_{q\chi 2}^{S, sb} = (\bar{s}b)(\bar{\chi}i\gamma_5\chi),
$$
\n(19a)

$$
\mathcal{O}_{q\chi^{1}}^{V,sb} = (\bar{s}\gamma^{\mu}b)(\bar{\chi}\gamma_{\mu}\chi), (\times)
$$

\n
$$
\mathcal{O}_{q\chi^{2}}^{V,sb} = (\bar{s}\gamma^{\mu}b)(\bar{\chi}\gamma_{\mu}\gamma_{5}\chi),
$$
\n(19b)

$$
\mathcal{O}_{q\chi 1}^{T,sb} = (\bar{s}\sigma^{\mu\nu}b)(\bar{\chi}\sigma_{\mu\nu}\chi), (\times)
$$

$$
\mathcal{O}_{q\chi 2}^{T,sb} = (\bar{s}\sigma^{\mu\nu}b)(\bar{\chi}\sigma_{\mu\nu}\gamma_5\chi), (\times)
$$
 (19c)

where the " (x) " indicates the accompanying operator vanishes for Majorana fermions. The implications for $B \to K$ transitions from such operators have been partially considered before [[49](#page-11-30)]. In Ref. [\[49\]](#page-11-30), only the charged processes $B^+ \to K^{(*)+} \chi \chi$ are considered and this was done before the new Belle II measurement. Here we use the new experimental result and include all possible modes including the neutral ones. We also use the new lattice form factors [\[46\]](#page-11-29).

The differential decay widths for this case have been given previously in [\[45\]](#page-11-28), and we reproduce them in terms of our convention in the Appendix for reference. To study the parameter space, we rewrite the Wilson coefficient C_i^j of the operator \mathcal{O}_i^j as an effective scale $C_i^j \equiv \Lambda_{\text{eff}}^{-2}$. For each operator in Eq. [\(19\),](#page-6-2) Fig. [5](#page-7-0) shows the parameter space in the m_{χ} - Λ _{eff} plane that reproduces the new Belle II 1 σ (new average) result as the pink (purple) shaded region. The gray region is excluded by other B meson decay modes. It can be seen that the two operators with tensor quark

 $7A$ general study of the three-body decay with sterile neutrinos in the ν SMEFT framework [\[47\]](#page-11-31) can be found in [[8,](#page-10-9)[48\]](#page-11-32).

FIG. 5. The pink [purple] region shows the parameter space that could explain the recent Belle II excess [new average] with fermion DM for the operators in Eq. [\(19\).](#page-6-2) The gray region is excluded by other B meson decay modes that are indicated in the plots by colored lines.

currents have almost no viable region to explain the excess, unlike the operators with scalar or vector currents. Most of the exclusion for these four operators arises from the upper limit on the neutral mode $B^0 \to K^0 \chi \chi$. If we insist on an excess of events concentrated in the $3 \le q^2 \le 7$ GeV² bins, only the vector current operators $\mathcal{O}_{q\chi_1,2}^{V,sb}$ with certain DM masses remain viable as we illustrate in Fig. [7](#page-8-0).

C. Vector DM case

Finally, we consider the vector DM. There are two parametrizations that can be used in this case as discussed in [[14](#page-11-0)]. Here we adopt the one with a four-vector field X_u for simplicity. The operators have been classified by us in [\[14\]](#page-11-0), and the ones relevant for $B^+ \to K^+XX$ transitions are

$$
\mathcal{O}_{qX}^{S,sb} = (\bar{s}b)(X_{\mu}^{\dagger}X^{\mu}),\tag{20a}
$$

$$
\mathcal{O}_{qX1}^{T,sb} = \frac{i}{2} (\bar{s}\sigma^{\mu\nu}b)(X_{\mu}^{\dagger}X_{\nu} - X_{\nu}^{\dagger}X_{\mu}), (\times)
$$
 (20b)

$$
\mathcal{O}_{qX2}^{T,sb} = \frac{1}{2} (\bar{s}\sigma^{\mu\nu}\gamma_5 b)(X_\mu^\dagger X_\nu - X_\nu^\dagger X_\mu), (\times) \qquad (20c)
$$

$$
\mathcal{O}_{qX2}^{V,sb} = (\bar{s}\gamma_\mu b)\partial_\nu (X^{\mu\dagger}X^\nu + X^{\nu\dagger}X^\mu),\tag{20d}
$$

$$
\mathcal{O}_{qX3}^{V,sb} = (\bar{s}\gamma_{\mu}b)(X_{\rho}^{\dagger}\overset{\leftrightarrow}{\partial}_{\nu}X_{\sigma})\epsilon^{\mu\nu\rho\sigma},\tag{20e}
$$

$$
\mathcal{O}_{qX4}^{V,sb} = (\bar{s}\gamma^{\mu}b)(X_{\nu}^{\dagger}i\overset{\leftrightarrow}{\partial}_{\mu}X^{\nu}), (\times)
$$
 (20f)

$$
\mathcal{O}_{qX5}^{V,sb} = (\bar{s}\gamma_\mu b)i\partial_\nu (X^{\mu\dagger}X^\nu - X^{\nu\dagger}X^\mu), (\times) \qquad (20g)
$$

$$
\mathcal{O}_{qX6}^{V,sb} = (\bar{s}\gamma_{\mu}b)i\partial_{\nu}(X_{\rho}^{\dagger}X_{\sigma})\epsilon^{\mu\nu\rho\sigma} . (\times)
$$
 (20h)

The symbol " (x) " indicates that the corresponding operator vanishes for real vector fields. To address the well-known singularity problem that affects vector fields in the limit of vanishing mass, for our numerical analysis, we scale the Wilson coefficients of these operators in the following manner:

$$
C_{qX}^{S} \equiv \frac{m^2}{\Lambda_{\text{eff}}^3}, \qquad C_{qX1,2}^{T} \equiv \frac{m^2}{\Lambda_{\text{eff}}^3},
$$

$$
C_{qX2,4,5}^{V} \equiv \frac{m^2}{\Lambda_{\text{eff}}^4}, \qquad C_{qX3,6}^{V} \equiv \frac{m}{\Lambda_{\text{eff}}^3}.
$$
 (21)

In Fig. [6](#page-8-1), we show the parameter space resulting in a branching ratio in agreement with Eq. [\(17\)](#page-5-0). It can be seen, that except for the two operators with tensor quark currents $\mathcal{O}_{qX1,2}^{T,sb}$, the remaining operators contain a large acceptable parameter region.

D. The q^2 distribution

The excess of events observed by Belle II appears to occur mainly for q^2 values between 3–7 GeV² [\[1](#page-10-0)].⁸ It is

⁸We thank Eldar Ganiev for confirming this observation.

FIG. 6. The pink (purple) region shows the parameter space that could explain the recent Belle II excess (new average) with vector DM for the operators in Eq. (20) . The gray region is excluded by other B meson decay modes that are indicated in the plots by colored lines.

FIG. 7. The q^2 distribution of normalized differential decay widths from all three cases: scalar DM (left panel), fermion DM (middle panel), vector DM (right panel) for selected masses.

FIG. 8. The q^2 distribution of normalized differential decay widths after taking into account the experimental efficiency. The left panel for scalar DM and right panel for fermion DM.

thus interesting to compare the q^2 distributions that follow from the different DM cases. Because the new particles would add incoherently to the SM rate, we combine the two to obtain a normalized q^2 distribution as

$$
\frac{d\tilde{\Gamma}}{dq^2} \equiv \frac{\frac{d\Gamma_{\text{SM}}}{dq^2} + \frac{d\Gamma_{\text{NP}}}{dq^2}}{\Gamma_{\text{SM}} + \Gamma_{\text{NP}}} = \frac{\frac{d\tilde{\Gamma}_{\text{SM}}}{dq^2} + (R_{\nu\nu}^K - 1)\frac{d\tilde{\Gamma}_{\text{NP}}}{dq^2}}{R_{\nu\nu}^K},\qquad(22)
$$

where $d\tilde{\Gamma}_{SM,NP}/dq^2 \equiv (d\Gamma_{SM,NP}/dq^2)/\Gamma_{SM,NP}$.
In Fig. 7, we show representative cases from

In Fig. [7](#page-8-0), we show representative cases from the insertion of all the DM operators discussed above for selected DM mass value $m_{\phi} = 300, 700 \text{ MeV}$ (left panel), $m_{\chi} =$ 700 MeV (center panel) and $m_V = 700$ MeV (right panel). For comparison, the SM distribution is also shown (black line). The figures indicate that the cases of $\mathcal{O}_{q\phi}^{V,sb}$ with $m_{\phi} \sim$ 300 MeV and $\mathcal{O}_{q\chi_1,2}^{V,sb}$ with $m_\chi \sim 700$ MeV would more closely match the preliminary q^2 distribution of the excess. This can be seen more clearly when the experimental detection efficiency is taken into account. Figure [8](#page-8-2) shows how the inclusion of the Belle II signal-selection efficiency for inclusive tagging analysis (ITA) affects the distribution for scalar and fermion DM case. Since the selection efficiency is given in each q^2 -bin, we show the new distributions in histogram according to $\left(\frac{d\tilde{\Gamma}}{dq^2}\right)_{i\in I}$ $(2 \text{ GeV}^2 \sum_i (d\tilde{\Gamma}/dq^2)_{i} \epsilon_i)$. Here ϵ_i is the efficiency taken from Fig. 6 in [\[1\]](#page-10-0), 2 GeV² is the bin width, and $\left(\frac{d\tilde{\Gamma}}{dq^2}\right)$ is evaluated at the central q^2 value of each bin. The study of the shape of this distribution will help narrow down possible explanations if the excess is confirmed.

IV. SUMMARY AND CONCLUSIONS

The recent measurement of the branching ratio $\mathcal{B}(B^+ \to K^+ \nu \bar{\nu})$ by the Belle II Collaboration is larger than the SM prediction at the 2.7σ level and has attracted some attention. This new Belle II result is consistent within errors with the average of measurements, and the discrepancy with the SM is not sufficiently large to require a new physics explanation. In anticipation of future, more precise, measurements, the new number invites speculation on the possibility of accommodating a rate that exceeds the SM. In particular it is important to understand whether models that can enhance $R_K^{\nu\nu}$ can remain consistent with a lower $R_{K^*}^{\nu\nu}$, closer to the SM.

Since the neutrinos in the final state are not identified, it is possible to enhance $R_K^{\nu\nu}$ with models that contain other invisible particles, such as light dark matter. It is also possible to do so with additional neutrinos or by introducing lepton flavor violating neutrino couplings as in models previously discussed in the literature.

We began with a model-independent description of the interactions in $b \rightarrow s \nu \bar{\nu}$. We considered four types of dimension six operators, two of which can interfere with the SM (when diagonal in neutrino flavor) and two that can not. We then scanned the parameter spaces to see if it was possible to cover the new Belle II 1σ range for $\mathcal{B}/B^+ \rightarrow$ $K^+\nu\bar{\nu}$) (or the new average) as well as existing bounds on $\mathcal{B}(B \to K^* \nu \bar{\nu})$. Our results, shown in Figs. [1](#page-3-0) and [2](#page-4-0) indicate that whereas it is relatively easy to reproduce the new 1σ range of $R_K^{\nu\nu}$, it is much harder to simultaneously obtain a low value of $R_{K^*}^{\nu\nu}$ < 1.9. This conclusion is stronger for operators which have no interference with the SM. The relative size of $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ will play an important role in untangling any hint for NP in these modes.

We then restricted the parameter scans to models with one leptoquark, and found that only $\tilde{S}_{1/2}$ or $V_{1/2}$ can produce $R_K^{\nu\nu} \neq R_{K^*}^{\nu\nu}$ and satisfy the constraints from both $R_{K^*}^{\nu\nu}$ and $b \to s\ell^+\ell^-$ modes. However, they cannot explain the $b \to s\ell^+\ell^-$ anomalies. These two cases predict significant enhancements for the modes $B^+ \to K^+\mu^-\tau^+,$ $B_s \to \mu^- \tau^+$, $B^+ \to K^+ \tau^- \tau^+$, and $B_s \to \tau^- \tau^+$. In contrast, we found that models with only C_L^{ij} coefficients, such as those with S_0 , S_1 or V_1 leptoquarks, cannot satisfy both conditions on $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ because they give $R_K^{\nu\nu} = R_{K^*}^{\nu\nu}$. Furthermore, if S_0 is used to explain the current values of $R_{D^{(*)}}$, it produces values of $R_K^{\nu\nu}$ and $R_{K^*}^{\nu\nu}$ much larger than the existing constraints. We also found that solutions with light right-handed neutrinos that couple to SM fields via a Z' are excluded when B_s mixing constraints are imposed.

In scenarios with light dark matter resulting in $B^+ \rightarrow$ $K^+ + \cancel{E}$, we present constraints on the effective scale of all lowest-dimensional operators coupling b , s quark bilinears to pairs of scalar, fermion or vector dark matter particles. We find that existing experimental upper bounds on $B^0 \to K^0 + \mathcal{E}, B^0 \to K^{0*} + \mathcal{E}, \text{ and } B^+ \to K^{++} + \mathcal{E} \text{ rule}$ out much of the parameter space where these operators could enhance $B^+ \to K^+ + \not{E}$ to the level of the new Belle II result.

We illustrated the remaining parameter space that can survive the constraints, finding that the effective scales are typically constrained to be in the multi-TeV range for scalar and fermion dark matter. For vector dark matter, the effective scales are in the several hundred GeV range when the masses are less than 100 MeV and become weaker as the mass increases due to the reduced phase space.

We confronted the different operators with the preliminary q^2 distribution reported by Belle II. We found that three operators with specific values for dark matter mass would best accommodate the spectrum, $\mathcal{O}_{q\phi}^{V, sb}$ with $m_{\phi} \sim 300$ MeV and $\mathcal{O}_{q\chi_1,2}^{V,sb}$ with $m_{\chi} \sim 700$ MeV. Clearly, the experimental result is preliminary, but our exercise illustrates how the shape of the spectrum can be used to narrow the possible explanations for an excess in the rate.

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APPENDIX: THE DIFFERENTIAL WIDTH FOR $B \to K^{(*)}\chi\chi$

For the light invisible fermionic DM case, based on the parametrization for the hadronic matrix elements given in [[14](#page-11-0),[21](#page-11-7),[50](#page-11-33),[51](#page-11-34)], the differential decay widths for $B \rightarrow$ $K^+(K^0)\chi\chi$ and $B \to K^{*+}(K^{*0})\chi\chi$ transitions induced from the interactions in Eq. [\(19\)](#page-6-2) are calculated to take the following general form:

$$
\frac{d\Gamma_{B\to P\chi\chi}}{dq^2} = \frac{\lambda^{\frac{1}{2}}(m_B^2, m_P^2, s) \kappa^{\frac{1}{2}}(m^2, s)}{384\pi^3 m_B^3} \left\{ \frac{3(m_B^2 - m_P^2)^2}{(m_b - m_s)^2} f_0^2 \left[(s - 4m^2) |C_{q\chi^1}^{S, sb}|^2 + s |C_{q\chi^2}^{S, sb}|^2 \right] + \frac{2(s + 2m^2)\lambda(m_B^2, m_P^2, s)}{s} f_+^2 |C_{q\chi^1}^{V, sb}|^2 \right. \\ \left. + \frac{2}{s} \left[6m^2(m_B^2 - m_P^2)^2 f_0^2 + (s - 4m^2)\lambda(m_B^2, m_P^2, s) f_+^2 \right] |C_{q\chi^2}^{V, sb}|^2 \right. \\ \left. + \frac{4\lambda(m_B^2, m_P^2, s)}{(m_B + m_P)^2} f_T^2 \left[(s + 8m^2) |C_{q\chi^1}^{T, sb}|^2 + (s - 4m^2) |C_{q\chi^2}^{T, sb}|^2 \right] \\ - \frac{12m(m_B^2 - m_P^2)^2}{m_b - m_s} f_0^2 \mathfrak{F} \left[C_{q\chi^2}^{S, sb} C_{q\chi^2}^{V, sb}| + \frac{24m\lambda(m_B^2, m_P^2, s)}{m_B + m_P} f_+ f_T \mathfrak{R} \left[C_{q\chi^1}^{V, sb} C_{q\chi^1}^{T, sb}|^2 \right] \right\}, \tag{A1}
$$

$$
\frac{d\Gamma_{B\to V_{XX}}}{dq^2} = \frac{\lambda_2^2 (m_B^2, m_P^2, s) \kappa_2^1 (m^2, s)}{96\pi^3 m_B^3 (m_B + m_V)^2} V_0^2 \Big[(s + 2m^2) |C_{qX1}^{V,sb}|^2 + (s - 4m^2) |C_{qX2}^{V,sb}|^2 \Big] \n+ \frac{\lambda_2^1 (m_B^2, m_V^2, s) \kappa_2^1 (m^2, s)}{48\pi^3 m_B^3 s} \Big\{ (s + 8m^2) \lambda (m_B^2, m_V^2, s) T_1^2 + (s - 4m^4) \Big[(m_B^2 - m_V^2)^2 T_2^2 + \frac{8m_B^2 m_V^2 s}{(m_B + m_V)^2} T_{23}^2 \Big] \Big\} |C_{qX1}^{T,sb}|^2 \n+ \frac{\lambda_2^1 (m_B^2, m_V^2, s) \kappa_2^1 (m^2, s)}{48\pi^3 m_B^3 s} \Big\{ (s - 4m^2) \lambda (m_B^2, m_V^2, s) T_1^2 + (s + 8m^2) \Big[(m_B^2 - m_V^2)^2 T_2^2 + \frac{8m_B^2 m_V^2 s}{(m_B + m_V)^2} T_{23}^2 \Big] \Big\} |C_{qX2}^{T,sb}|^2 \n+ \frac{m \lambda_2^2 (m_B^2, m_V^2, s) \kappa_2^1 (m^2, s)}{8\pi^3 m_B^3 (m_B + m_V)} V_0 T_1 \Re \Big[C_{qX1}^{V,sb} C_{qX1}^{T,sb*} \Big], \tag{A2}
$$

where $P = K^+, K^0$, and $V = K^{*,}, K^{*0}, C_i^j$ is the corresponding Wilson coefficient of the operator \mathcal{O}_i^j . The Källen function $\lambda(x, y, z) \equiv x^2 + y^2 + z^2 - 2(xy + yz + zx)$ and $\kappa(m^2, s) = \sqrt{1 - 4m^2/s}$, with m being the DM mass while s the invariant mass of DM pair. m_B , m_V are the masses of mesons B^+ , P, V and m_b , m_s are masses of b, s quarks, respectively. $f_0, f_+, V_0, f_T, T_1, T_2, T_{23}$ are hadronic form factors associated with different quark currents, with definitions that can be found in [\[14,](#page-11-0)[21\]](#page-11-7).

- [1] I. Adachi et al. (Belle-II Collaboration), [arXiv:2311.14647.](https://arXiv.org/abs/2311.14647)
- [2] F. Dattola (Belle-II Collaboration), in Proceedings of the 55th Rencontres de Moriond on Electroweak Interactions and Unified Theories (2021), [arXiv:2105.05754](https://arXiv.org/abs/2105.05754).
- [3] D. Bečirević, G. Piazza, and O. Sumensari, [Eur. Phys. J. C](https://doi.org/10.1140/epjc/s10052-023-11388-z) 83[, 252 \(2023\).](https://doi.org/10.1140/epjc/s10052-023-11388-z)
- [4] P. Asadi, A. Bhattacharya, K. Fraser, S. Homiller, and A. Parikh, [J. High Energy Phys. 10 \(2023\) 069.](https://doi.org/10.1007/JHEP10(2023)069)
- [5] R. Bause, H. Gisbert, and G. Hiller, [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.109.015006) 109, [015006 \(2024\).](https://doi.org/10.1103/PhysRevD.109.015006)
- [6] L. Allwicher, D. Becirevic, G. Piazza, S. Rosauro-Alcaraz, and O. Sumensari, Phys. Lett. B 848[, 138411 \(2024\)](https://doi.org/10.1016/j.physletb.2023.138411).
- [7] P. Athron, R. Martinez, and C. Sierra, [J. High Energy Phys.](https://doi.org/10.1007/JHEP02(2024)121) [02 \(2024\) 121.](https://doi.org/10.1007/JHEP02(2024)121)
- [8] T. Felkl, A. Giri, R. Mohanta, and M. A. Schmidt, [Eur. Phys.](https://doi.org/10.1140/epjc/s10052-023-12326-9) J. C 83[, 1135 \(2023\)](https://doi.org/10.1140/epjc/s10052-023-12326-9).
- [9] M. Abdughani and Y. Reyimuaji, [arXiv:2309.03706](https://arXiv.org/abs/2309.03706).
- [10] H. K. Dreiner, J. Y. Günther, and Z. S. Wang, [arXiv:2309](https://arXiv.org/abs/2309.03727) [.03727.](https://arXiv.org/abs/2309.03727)
- [11] A. J. Buras, [Eur. Phys. J. C](https://doi.org/10.1140/epjc/s10052-023-11222-6) 83, 66 (2023).
- [12] W. G. Parrott, C. Bouchard, and C. T. H. Davies (HPQCD) Collaboration), Phys. Rev. D 107[, 014511 \(2023\);](https://doi.org/10.1103/PhysRevD.107.014511) [107](https://doi.org/10.1103/PhysRevD.107.119903), [119903\(E\) \(2023\).](https://doi.org/10.1103/PhysRevD.107.119903)
- [13] X. G. He and G. Valencia, Phys. Lett. B 821[, 136607 \(2021\).](https://doi.org/10.1016/j.physletb.2021.136607)
- [14] X.-G. He, X.-D. Ma, and G. Valencia, [J. High Energy Phys.](https://doi.org/10.1007/JHEP03(2023)037) [03 \(2023\) 037.](https://doi.org/10.1007/JHEP03(2023)037)
- [15] J. Grygier et al. (Belle Collaboration), [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.96.091101) 96, [091101 \(2017\);](https://doi.org/10.1103/PhysRevD.96.091101) 97[, 099902\(A\) \(2018\).](https://doi.org/10.1103/PhysRevD.97.099902)
- [16] E. E. Jenkins, A. V. Manohar, and P. Stoffer, [J. High Energy](https://doi.org/10.1007/JHEP03(2018)016) [Phys. 03 \(2018\) 016.](https://doi.org/10.1007/JHEP03(2018)016)
- [17] Y. Liao, X.-D. Ma, and Q.-Y. Wang, [J. High Energy Phys.](https://doi.org/10.1007/JHEP08(2020)162) [08 \(2020\) 162.](https://doi.org/10.1007/JHEP08(2020)162)
- [18] G. Buchalla and A. J. Buras, Nucl. Phys. B548[, 309 \(1999\).](https://doi.org/10.1016/S0550-3213(99)00149-2)
- [19] J. Brod, M. Gorbahn, and E. Stamou, [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.83.034030) 83, [034030 \(2011\).](https://doi.org/10.1103/PhysRevD.83.034030)
- [20] T. E. Browder, N. G. Deshpande, R. Mandal, and R. Sinha, Phys. Rev. D 104[, 053007 \(2021\)](https://doi.org/10.1103/PhysRevD.104.053007).
- [21] N. Gubernari, A. Kokulu, and D. van Dyk, [J. High Energy](https://doi.org/10.1007/JHEP01(2019)150) [Phys. 01 \(2019\) 150.](https://doi.org/10.1007/JHEP01(2019)150)
- [22] D. M. Straub, [arXiv:1810.08132](https://arXiv.org/abs/1810.08132).
- [23] A. J. Davies and X.-G. He, [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.43.225) 43, 225 (1991).
- [24] S. Davidson, D. C. Bailey, and B. A. Campbell, [Z. Phys. C](https://doi.org/10.1007/BF01552629) 61[, 613 \(1994\).](https://doi.org/10.1007/BF01552629)
- [25] I. Doršner, S. Fajfer, A. Greljo, J. F. Kamenik, and N. Košnik, [Phys. Rep.](https://doi.org/10.1016/j.physrep.2016.06.001) 641, 1 (2016).
- [26] J. M. Lizana, J. Matias, and B. A. Stefanek, [J. High Energy](https://doi.org/10.1007/JHEP09(2023)114) [Phys. 09 \(2023\) 114.](https://doi.org/10.1007/JHEP09(2023)114)
- [27] Y. Amhis, S. Banerjee, E. Ben-Haim, E. Bertholet, F. U. Bernlochner, M. Bona, A. Bozek, C. Bozzi, J. Brodzicka, V. Chobanova et al. (Heavy Flavor Averaging Group Collaboration), Phys. Rev. D 107[, 052008 \(2023\).](https://doi.org/10.1103/PhysRevD.107.052008)
- [28] R. Bause, H. Gisbert, M. Golz, and G. Hiller, [J. High Energy](https://doi.org/10.1007/JHEP12(2021)061) [Phys. 12 \(2021\) 061.](https://doi.org/10.1007/JHEP12(2021)061)
- [29] N. Rajeev and R. Dutta, Phys. Rev. D 105[, 115028 \(2022\).](https://doi.org/10.1103/PhysRevD.105.115028)
- [30] R. Aaij et al. (LHCb Collaboration), [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.131.051803) 131, [051803 \(2023\).](https://doi.org/10.1103/PhysRevLett.131.051803)
- [31] B. Capdevila, A. Crivellin, and J. Matias, [Eur. Phys. J. Spec.](https://doi.org/10.1140/epjs/s11734-023-01012-2) Top. 1[, 20 \(2023\)](https://doi.org/10.1140/epjs/s11734-023-01012-2).
- [32] U. Laa, D. Cook, and G. Valencia, [J. Comput. Graph. Stat.](https://doi.org/10.1080/10618600.2020.1777140) 29[, 681 \(2020\).](https://doi.org/10.1080/10618600.2020.1777140)
- [33] U. Laa and G. Valencia, [Eur. Phys. J. Plus](https://doi.org/10.1140/epjp/s13360-021-02310-1) 137, 145 (2022).
- [34] U. Laa, A. Aumann, D. Cook, and G. Valencia, [J. Comput.](https://doi.org/10.1080/10618600.2023.2206459) Graph. Stat. 32[, 1229 \(2023\)](https://doi.org/10.1080/10618600.2023.2206459).
- [35] See Supplemental Material at [http://link.aps.org/](http://link.aps.org/supplemental/10.1103/PhysRevD.109.075019) [supplemental/10.1103/PhysRevD.109.075019](http://link.aps.org/supplemental/10.1103/PhysRevD.109.075019) for a video demonstrating how to find this slice and projection.
- [36] R. L. Workman et al. (Particle Data Group), [Prog. Theor.](https://doi.org/10.1093/ptep/ptac097) Exp. Phys. 2022[, 083C01 \(2022\)](https://doi.org/10.1093/ptep/ptac097).
- [37] X.-G. He and G. Valencia, Phys. Rev. D 66[, 013004 \(2002\)](https://doi.org/10.1103/PhysRevD.66.013004); 66[, 079901\(E\) \(2002\).](https://doi.org/10.1103/PhysRevD.66.079901)
- [38] X.-G. He and G. Valencia, [Phys. Lett. B](https://doi.org/10.1016/j.physletb.2018.01.073) 779, 52 (2018).
- [39] X.-G. He and G. Valencia, [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.70.053003) 70, 053003 [\(2004\).](https://doi.org/10.1103/PhysRevD.70.053003)
- [40] X.-G. He and G. Valencia, Phys. Rev. D **74**[, 013011 \(2006\).](https://doi.org/10.1103/PhysRevD.74.013011)
- [41] S. Fajfer and A. Novosel, [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.104.015014) 104, 015014 [\(2021\).](https://doi.org/10.1103/PhysRevD.104.015014)
- [42] G. Li and J. Tandean, [J. High Energy Phys. 11 \(2023\) 205.](https://doi.org/10.1007/JHEP11(2023)205)
- [43] K. Fridell, M. Ghosh, T. Okui, and K. Tobioka, [arXiv:](https://arXiv.org/abs/2312.12507) [2312.12507.](https://arXiv.org/abs/2312.12507)
- [44] L. Gärtner, N. Hartmann, L. Heinrich, M. Horstmann, T. Kuhr, M. Reboud, S. Stefkova, and D. van Dyk, [arXiv:2402](https://arXiv.org/abs/2402.08417) [.08417.](https://arXiv.org/abs/2402.08417)
- [45] J. F. Kamenik and C. Smith, [J. High Energy Phys. 03](https://doi.org/10.1007/JHEP03(2012)090) [\(2012\) 090.](https://doi.org/10.1007/JHEP03(2012)090)
- [46] W. G. Parrott, C. Bouchard, and C. T. H. Davies ((HPQCD) collaboration) and HPQCD Collaboration), [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.107.014510) 107[, 014510 \(2023\).](https://doi.org/10.1103/PhysRevD.107.014510)
- [47] Y. Liao and X.-D. Ma, Phys. Rev. D 96[, 015012 \(2017\)](https://doi.org/10.1103/PhysRevD.96.015012).
- [48] T. Felkl, S. L. Li, and M. A. Schmidt, [J. High Energy Phys.](https://doi.org/10.1007/JHEP12(2021)118) [12 \(2021\) 118.](https://doi.org/10.1007/JHEP12(2021)118)
- [49] G. Li, T. Wang, Y. Jiang, J.-B. Zhang, and G.-L. Wang, Phys. Rev. D 102[, 095019 \(2020\)](https://doi.org/10.1103/PhysRevD.102.095019).
- [50] P. Ball and R. Zwicky, Phys. Rev. D 71[, 014015 \(2005\)](https://doi.org/10.1103/PhysRevD.71.014015).
- [51] A. Bharucha, D. M. Straub, and R. Zwicky, [J. High Energy](https://doi.org/10.1007/JHEP08(2016)098) [Phys. 08 \(2016\) 098.](https://doi.org/10.1007/JHEP08(2016)098)