

Anisotropy of magnetized quark matter

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Strong transient magnetic fields are generated in noncentral relativistic heavy-ion collisions. These fields induce anisotropy within the strongly interacting medium that, in principle, can affect the thermodynamic properties of the medium. We use the Polyakov loop extended Nambu Jona-Lasinio model to study the quark matter subjected to an external magnetic field at vanishing baryon chemical potential (μ_B). We have estimated the degree of anisotropy in the speed of sound and isothermal compressibility within the magnetized quark matter as a function of temperature (T) and magnetic field (eB). This study helps us to understand the extent of directionality generated in the initial stages of noncentral collisions while giving us useful information about the system.

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I. INTRODUCTION

One of the primary goals of relativistic heavy-ion collision experiments is to study the deconfined state of strongly interacting quarks and gluons in local thermal equilibrium, known as the quark-gluon plasma (QGP). In the noncentral heavy-ion collision, the charged spectators move past the fireball at a relativistic speed. According to the Biot-Savart law, these moving charge particles create a large transient electromagnetic field of the order of 10^{18} G at the Large Hadron Collider (LHC) [1,2]. Direct experimental evidence for the strength of the magnetic field (eB) is yet to be discovered. However, recent measurements of the directed flow of D^0 and \bar{D}^0 at the Relativistic Heavy Ion Collider (RHIC) [3] and Large Hadron Collider [4], indicate the creation of a strong magnetic field during the collision.

The first principle nonperturbative theory of strong interaction, lattice quantum chromodynamics (LQCD), found many interesting phenomena in QGP in the presence of a magnetic field such as the chiral magnetic effect [5], and magnetic and inverse magnetic catalysis [6]. However, experimental verification of these phenomena are yet to be observed. Under an external magnetic field, the energy levels of the charged particles get quantized following the Landau quantization, which creates momentum anisotropy

affecting various thermodynamical [7–9] and dissipative quantities [10–12]. For instance, the thermodynamical pressure becomes anisotropic with a longitudinal component along the magnetic field and a transverse component in the transverse plane of the field. This leads to anisotropy in many other thermodynamical quantities, such as speed of sound (c_s) and isothermal compressibility (κ_T). Very recently, the anisotropy in the speed of sound in a magnetized hybrid neutron star was explored in Ref. [13]. The authors considered a magnetized hybrid neutron star with three density phases, employed three different models to study the speed of sound, and explored the anisotropy created due to a magnetic field at zero temperature. The MIT bag model was considered for high-density regimes where quark matter could possibly exist. Reference [14] also explored the anisotropy in the equation of state for magnetized baryon matter at finite temperature and chemical potential using a two-flavor Nambu Jona-Lasinio (NJL) model. The study aimed to locate the critical end point with the help of isothermal compressibility and quark number susceptibility. Such studies in low-temperature and high- μ_B regimes motivated us to understand and quantify the anisotropy generated in peripheral heavy-ion collision at the LHC energies. In the strong field limit, Landau quantization leads to a dimensional reduction in the phase space ($3 \rightarrow 1$ dimension) at the lowest Landau level. Moreover, the magnetic field influences the QCD phase diagram. Different effective QCD models were used to study the phase diagram in the $B - T$ plane, such as the linear sigma model [15], NJL model, and its extended version Polyakov loop extended NJL (PNJL) model [16,17]. Initial prediction of the PNJL model showed that the transition temperature (T_c) and its strength increase with

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eB , leading to a first-order phase transition. However, the LQCD calculation [6] found an opposite trend, which means T_c decreases with eB . The same was also reproduced in model calculations such as NJL and PNJL models [18,19] with a magnetic field-dependent coupling constant. The magnetic field-dependent coupling constant was introduced in these models to allow them to explain the phenomena of the inverse magnetic catalysis, as observed in LQCD.

Many indirect probes and observables are suggested to study the microscopic and bulk features of QGP under the effect of the magnetic field. Theoretically, one can study the change in the thermodynamic observables to understand the changes in the deconfined medium in the presence of an external magnetic field. The behavior of certain thermodynamic observables such as κ_T and c_s^2 provide useful information about the nature of the phase transition of the system. κ_T represents the rate of change in the volume of the system concerning pressure at a constant temperature. Precisely, κ_T measures the extent to which the density of quarks and gluons changes in response to changes in external pressure, which is an essential factor in determining the equation of state of the medium [20–22]. Moreover, it can tell us about the degree of deviation of a system from a perfect fluid. κ_T is expected to show a sudden jump near the critical end-point (CEP), where the smooth cross-over in the QCD phase diagram meets the first-order phase transition. Thus, it is an interesting observable to explore the QCD phase diagram [23]. In the literature, κ_T has been studied as a function of temperature and charged particle multiplicity [24–29]. In Ref. [25], the high-temperature QCD matter has been found to be the closest to a perfect fluid.

On the other hand, the speed of sound reflects the propagation of small perturbations produced in the system in its local rest frame. Its dependence on the environment, i.e., temperature, density, and baryon chemical potential, means that it is an ideal probe to explore the evolution of the fireball [30]. The studies in Refs. [29,31–33] reveal that exploring c_s^2 as a function of charged particle multiplicity can be utilized to study the dynamics of heavy-ion collisions. Moreover, it shows minima near the phase transition. This also makes c_s^2 a suitable observable to study the QCD phase space. In literature, many phenomenological models are employed to compute the speed of sound such as the quasiparticle model [34], dynamic quasiparticle model [35], color string percolation model [36,37], hadron resonance gas model [38,39], NJL model [40,41], and PNJL model [42–44]. It has also been estimated using the LQCD calculations [45–48].

The lattice QCD model is a first principle theory of strongly interacting matter. Despite its many successes, it fails to explain the high-baryon-rich environment due to its fermion sign problem [23]. Hence, exploring the CEP within the LQCD formalism is not feasible. Thus, the NJL

and PNJL models are good alternatives in this regard. The PNJL model is an extended version of the NJL model, initially developed for nuclear matter. Because of its respect for the global symmetries of QCD, particularly the chiral symmetry, this model has been widely utilized to explore some of the nonperturbative features of the QCD vacuum. The NJL model is not renormalizable due to the pointlike interaction between the quarks [49,50]. A suitable regularization method must be established to deal with the divergent integrals. The three-momentum cutoff scheme is widely accepted among many others in the literature [51]. The model's parameters must be fixed to replicate well-known phenomenological quantities, such as the pion-decay constant and quark condensate density [51]. The NJL model is based on the idea that the chiral symmetry of quantum chromodynamics (QCD) is spontaneously broken at low temperatures and densities, which leads to the appearance of a nonzero quark condensate. The NJL model incorporates this symmetry breaking by introducing a four-fermion interaction between quarks. The interaction is attractive in the scalar channel and repulsive in the pseudoscalar channel, which leads to the formation of a chiral condensate and the generation of quark masses [52].

Albeit being a successful model, one drawback of the NJL model is that it does not incorporate the gluon dynamics, and as a result, quark confinement is absent in this model. As suggested in [53], the Polyakov loop characterizes the effects of confinement, which prevents quarks from existing as free particles. The trace of the Polyakov loop (Φ) is the order parameter for the confinement transition in a pure gauge theory [54–57]. It can be understood that Φ is related to $e^{-F/T}$, where F is the free energy of the static quark. The free energy of a confined single quark is infinite, which leads to $\Phi = 0$, and in the deconfined phase, free energy is finite, which makes $\Phi \neq 0$ [58]. In the PNJL model, the Polyakov loop is included as a background temporal gluon field that interacts with the quarks. The PNJL model has been successful in reproducing lattice results qualitatively [30,43,59–61]. The PNJL model has also been used to study a wide range of phenomena in QCD matter, including thermodynamic and transport properties [35,62–66], probing the critical point in the $(T - \mu_B)$ plane [67–70] and exploring the QCD medium under the effect of an external magnetic field [9,18,71–73].

In this work, we investigate the behavior of the speed of sound and isothermal compressibility and explore the degree of anisotropy generated in baryon-free quark matter in the presence of an external magnetic field. This paper is organized as follows. In Secs. II A and II B, we have discussed the PNJL model formulation with and without an external magnetic field, respectively. Section III comprises the results of this work. Finally, Sec. IV summarizes this work.

II. FORMULATION

A. Quark matter in zero magnetic field

The PNJL model was introduced as an improvement to the NJL model [53,59,74,75]. The NJL model effectively takes care of the chiral symmetry breaking. However, due to the lack of gluonic interaction in the model, it fails to explain the deconfinement dynamics of quarks. The PNJL model takes care of this problem by adding a Polyakov loop effective potential to the NJL Lagrangian. This model simplifies the interaction between quarks and gluons by considering chiral point couplings between quarks and a temporal background gauge field. Extensive studies have been done using the PNJL model with 2 and 2 + 1 flavors [17,43,59,67,68,76–78]. The PNJL Lagrangian for 2 + 1 flavors is given by [68,78,79]

$$\begin{aligned} \mathcal{L}_{PNJL} = & \bar{q}(i\gamma^\mu D_\mu - \hat{m})q + G \sum_{a=0}^8 [(\bar{q}\lambda^a q)^2 + (\bar{q}i\gamma_5\lambda^a q)^2] \\ & - K \{ \det[\bar{q}(1 + \gamma_5)q] + \det[\bar{q}(1 - \gamma_5)q] \} \\ & - \mathcal{U}(\Phi, \bar{\Phi}, T) \end{aligned} \quad (1)$$

where q is the three-flavor quark field, $q = (u, d, s)$, λ^a are the Gell-Mann matrices in flavor space, and \hat{m} is the current quark mass, $\hat{m} = \text{diag}(m_u, m_d, m_s)$. The covariant derivative is defined as

$$D^\mu = \partial^\mu - iA^\mu,$$

where $A^\mu = \delta_0^\mu A^0$ and in the Euclidian notation, we can write $A^0 = -iA_4$. The four-point interaction and six-point

interaction of the quark field are incorporated with an effective coupling strength, G and K , respectively. $\mathcal{U}(\Phi, \bar{\Phi}, T)$ is the effective Polyakov loop potential. The parameters in the Lagrangian are the cutoff parameter λ , the coupling constants G and K , and in the current quark masses m_u and m_s , we use the value set in Ref. [80] as $\lambda = 602.3$ MeV, $G\lambda^2 = 1.835$, $K\lambda^5 = 12.36$, $m_u = m_d = 5.5$ MeV, and $m_s = 140.7$ MeV. These parameters are obtained by fixing the values of $m_\pi = 135$ MeV, $m_K = 497.7$ MeV, $m_\eta = 957.8$ MeV and $f_\pi = 92.4$ MeV.

This model takes the analogy from the BCS theory of superconductor where pairing between electrons in a spin-singlet state leads to a condensation which introduces a gap in the energy spectra. Similarly, the condensation $\langle \bar{q}_R q_L \rangle$ (or $\langle \bar{q}_L q_R \rangle$) arises due to the pairing between the quark-antiquark of the same chirality. Thus, $\langle \bar{q}q \rangle = \langle \bar{q}_R q_L + \bar{q}_L q_R \rangle \sim \sigma$ can be taken as the order parameter and in the mean field approximation gives rise to a dynamical mass $M \sim \langle \bar{q}q \rangle$ [81]. Thus, the gap equation for the PNJL model can be expressed as [78]

$$M_i = m_i - 4G\sigma_i + 2K\sigma_j\sigma_k \quad (2)$$

where $i, j, k = u, d, s$ in cyclic order, M_i is the constituent quark mass, m_i is the bare quark mass, and $\sigma_{i,j,k}$ is the quark condensate for different flavors. The third term exists as a result of introducing six-fermion interactions and in turn, generating a flavor mixing in the quark mass.

The thermodynamic potential is given as [78,82]

$$\begin{aligned} \Omega = & \mathcal{U}(\Phi, \bar{\Phi}, T) + 2G \sum_{f=u,d,s} \sigma_f^2 - 4K\sigma_u\sigma_d\sigma_s - 2N_c \sum_{f=u,d,s} \int_\Lambda \frac{d^3p}{(2\pi)^3} \sqrt{p^2 + M_f^2} \\ & - 2T \sum_{f=u,d,s} \int \frac{d^3p}{(2\pi)^3} \ln[1 + 3\Phi e^{-\beta(E-\mu)} + 3\bar{\Phi} e^{-2\beta(E-\mu)} + e^{-3\beta(E-\mu)}] \\ & - 2T \sum_{f=u,d,s} \int \frac{d^3p}{(2\pi)^3} \ln[1 + 3\bar{\Phi} e^{-\beta(E+\mu)} + 3\Phi e^{-2\beta(E+\mu)} + e^{-3\beta(E+\mu)}], \end{aligned} \quad (3)$$

where $\beta = 1/T$, $E = \sqrt{p^2 + M^2}$ is the energy of the quark and $\mathcal{U}(\Phi, \bar{\Phi}, T)$ is the effective Polyakov loop potential for the Φ and $\bar{\Phi}$ fields. Different versions of this potential exist in the literature [67,83,84]. For our study, we choose the following form of the potential, which mathematically limits the value of Φ to be unity as $T \rightarrow \infty$ [67],

$$\begin{aligned} \frac{\mathcal{U}(\Phi, \bar{\Phi}, T)}{T^4} = & -\frac{a(T)}{2} \Phi \bar{\Phi} + b(T) \ln[1 - 6\Phi \bar{\Phi} \\ & + 4(\Phi^3 + \bar{\Phi}^3) - 3(\Phi \bar{\Phi})^2], \end{aligned} \quad (4)$$

where $a(T)$ and $b(T)$ reads

$$\begin{aligned} a(T) = & a_0 + a_1 \left(\frac{T_0}{T}\right) + a_2 \left(\frac{T_0}{T}\right)^2, \\ b(T) = & b_3 \left(\frac{T_0}{T}\right)^3, \end{aligned} \quad (5)$$

where the parameters a_0 , a_1 , a_2 , and b_3 are fixed by performing a simultaneous fit of the thermodynamic observables obtained from the lattice QCD calculation and given as [67]

$$\begin{aligned} a_0 &= 3.51, & a_1 &= -2.47 \\ a_2 &= 15.2, & b_3 &= -1.75. \end{aligned} \quad (6)$$

The parameter T_0 is the critical temperature for the deconfinement phase transition in a pure-gauge system. In a pure-gauge approach, T_0 is fixed to be 270 MeV. One can find different values of T_0 in the literature. In Ref. [85], a N_f dependent form of T_0 is taken. In Ref. [86], the authors studied the effect of different T_0 values on the thermodynamic observables. However, following Ref. [59], for a quantitative comparison of results with LQCD data, we choose $T_0 = 190$ MeV.

In order to estimate Φ , $\bar{\Phi}$, and σ_f , one needs to minimize Ω with respect to the above-mentioned quantities [78,87],

$$\frac{\partial \Omega}{\partial \sigma_f} = 0, \quad \frac{\partial \Omega}{\partial \Phi} = 0, \quad \frac{\partial \Omega}{\partial \bar{\Phi}} = 0.$$

Solving $\frac{\partial \Omega}{\partial \sigma_f} = 0$, we get the condensates as [87]

$$\sigma_f = -2N_c \int_{\Lambda} \frac{d^3 p}{(2\pi)^3} \frac{M_f}{E_f} (1 - f_{\Phi}^+ - f_{\bar{\Phi}}^-), \quad (7)$$

where f_{Φ}^+ and $f_{\bar{\Phi}}^-$ are the quark and antiquark distribution functions, respectively [76,82],

$$\begin{aligned} f_{\Phi}^+ &= \frac{(\Phi + 2\bar{\Phi}e^{-\beta(E_f - \mu_f)})e^{-\beta(E_f - \mu_f)} + e^{-\beta(E_f - \mu_f)}}{1 + 3(\Phi + \bar{\Phi}e^{-\beta(E_f - \mu_f)})e^{-\beta(E_f - \mu_f)} + e^{-3\beta(E_f - \mu_f)}}, \\ f_{\bar{\Phi}}^- &= \frac{(\bar{\Phi} + 2\Phi e^{-\beta(E_f + \mu_f)})e^{-\beta(E_f + \mu_f)} + e^{-\beta(E_f + \mu_f)}}{1 + 3(\bar{\Phi} + \Phi e^{-\beta(E_f + \mu_f)})e^{-\beta(E_f + \mu_f)} + e^{-3\beta(E_f + \mu_f)}}. \end{aligned}$$

Finally, pressure can be estimated as $P = -\Omega$. We take care of the vacuum pressure, taken numerically at

$T = 0.001$ and $\mu = 0$, by subtracting its contribution from the total pressure given as [88]

$$P = -(\Omega(T, \mu) - \Omega(0.001, 0)). \quad (8)$$

Now, the entropy density (s) and the energy density (ϵ) are estimated as

$$\begin{aligned} s &= -\frac{\partial \Omega}{\partial T} = \frac{\partial P}{\partial T}, \\ \epsilon &= -P + Ts. \end{aligned}$$

Furthermore, the speed of sound c_s and the isothermal compressibility κ_T can be obtain as

$$c_s^2 = \frac{\partial P}{\partial \epsilon} = \frac{\frac{\partial P}{\partial T}}{\frac{\partial \epsilon}{\partial T}}, \quad (9)$$

$$\kappa_T = -\frac{1}{V} \frac{\partial V}{\partial P}, \quad (10)$$

where volume, $V = N/n$ with the total number of partons N and the number density n . Plugging this into the above equation, we get

$$\kappa_T = \frac{1}{n} \frac{\partial n}{\partial P} = \frac{1}{n} \frac{\frac{\partial n}{\partial \mu}}{\frac{\partial P}{\partial \mu}}. \quad (11)$$

B. Quark matter in a finite magnetic field

Under the effect of an external magnetic field, the PNJL Lagrangian changes as [18,89]

$$\begin{aligned} \mathcal{L}_{PNJL}^B &= \bar{q}(i\gamma^\mu D_\mu - \hat{m})q + G_B(eB) \sum_{a=0}^8 [(\bar{q}\lambda^a q)^2 + (\bar{q}i\gamma_5 \lambda^a q)^2] - K \{ \det[\bar{q}(1 + \gamma_5)q] + \det[\bar{q}(1 - \gamma_5)q] \} \\ &\quad - \mathcal{U}(\Phi, \bar{\Phi}, T) - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}. \end{aligned} \quad (12)$$

The coupling of the quarks and the temporal effective gluon field with the external magnetic field is incorporated by adding the $\frac{1}{4} F_{\mu\nu} F^{\mu\nu}$ term, where $F_{\mu\nu} = \partial_\mu A_\nu^{\text{EM}} - \partial_\nu A_\mu^{\text{EM}}$, and altering the covariant derivative as $D^\mu = \partial^\mu - iq_f A_{\text{EM}}^\mu - iA^\mu$, where e_f is the electric charge of the quark of flavor f . The coupling constant G is taken as $G_B(eB)$ a function of eB to incorporate the effect of an external magnetic field. Various forms of $G_B(eB)$ exist in the literature [18,71,90–94]. For our study, we use the form given in [18]

$$G_B(eB) = G \left(\frac{1 + a\zeta^2 + b\zeta^3}{1 + c\zeta^2 + d\zeta^4} \right). \quad (13)$$

This form is obtained by reproducing the chiral transition temperature, $T_C^\chi(eB)$, obtained in LQCD [6]. This form of coupling constant is able to yield qualitatively precise results for quark condensate values at a finite magnetic field, where $a = 0.0108805$, $b = -1.0133 \times 10^{-4}$, $c = 0.02228$,

$d = 1.84558 \times 10^{-4}$, and $\zeta = eB/\Lambda_{\text{QCD}}^2$ with $\Lambda_{\text{QCD}} = 300$ MeV. Finally, the thermodynamic potential due to an external magnetic field changes as [18,89]

$$\Omega(T, \mu, eB) = \mathcal{U}(\Phi, \bar{\Phi}, T) + 2G_B(eB) \sum_{f=u,d,s} \sigma_f^2 - 4K\sigma_u\sigma_d\sigma_s + \left(\sum_{f=u,d,s} \Omega_f^{\text{vac}} + \Omega_f^{\text{med}} + \Omega_f^{\text{mag}} \right) \quad (14)$$

where contributions from the vacuum, Ω_f^{vac} , medium, Ω_f^{med} , and magnetic field, Ω_f^{mag} are given as [18,95,96]

$$\begin{aligned} \Omega_f^{\text{vac}} &= -6 \int_{\Lambda} \frac{d^3 p}{(2\pi)^3} \sqrt{p^2 + M_f^2}, \\ \Omega_f^{\text{med}} &= -T \frac{|e_f B|}{4\pi^2} \sum_k a(k) \int_{-\infty}^{\infty} dp_z (Z_{\Phi}^+ + Z_{\Phi}^-), \\ \Omega_f^{\text{mag}} &= -\frac{3|e_f B|^2}{2\pi^2} \left[\zeta'(-1, x_f) - \frac{1}{2}(x_f^2 - x_f) \ln(x_f) + \frac{x_f^2}{4} \right]. \end{aligned} \quad (15)$$

Here, the contribution from the vacuum (Ω_f^{vac}) is regularized with a three-momentum cutoff scheme. In this scheme, the particles above a sharp momentum cutoff ($\Lambda = 602.3$ MeV) are excluded. Thus, allowing the vacuum term to restrict itself to a finite value. For Ω_f^{mag} , the authors in Ref. [95] have taken care of the divergent part by using the standard dimensional regularization formula. Moreover, the contribution coming from medium (Ω_f^{med}) is a converging function, and thus no regularization scheme has been applied for this part. However, in the literature, there exist a few studies where the authors have considered the Λ cutoff for the Ω_f^{med} part as well [97,98]. It has been observed in these studies that considering the medium contributions without any three-momentum cutoff yields better results for quantities such as pressure, energy density, specific heat, and speed of sound as well as deconfinement temperature. The details of the three-momentum cutoff and other possible regularization schemes have been explained in Refs. [51,95,99].

In Eq. (15), Z_{Φ}^+ and Z_{Φ}^- are the partition function density given as [76,82]

$$\begin{aligned} Z_{\Phi}^+ &= \ln[1 + 3(\Phi + \bar{\Phi} e^{-\beta(E_f - \mu_f)}) e^{-\beta(E_f - \mu_f)} + e^{-3\beta(E_f - \mu_f)}] \\ Z_{\Phi}^- &= \ln[1 + 3(\bar{\Phi} + \Phi e^{-\beta(E_f + \mu_f)}) e^{-\beta(E_f + \mu_f)} + e^{-3\beta(E_f + \mu_f)}]. \end{aligned}$$

$x_f = M_f^2/(2|e_f B|)$, $\zeta(z, x)$ is the Riemann-Hurwitz zeta function, and $\zeta'(-1, x_f) = d\zeta(z, x_f)/dz|_{z=-1}$ can be expressed as [100,101]

$$\zeta'(-1, x_f) = \frac{1}{12} - \frac{x_f^2}{4} + \left(\frac{1}{12} - \frac{x_f}{2} + \frac{x_f^2}{2} \right) \log(x_f) + \mathcal{O}(x_f^{-2}).$$

In the presence of a magnetic field, the dispersion relation of quarks will be modified due to the Landau quantization as

$$E_f = \sqrt{p^2 + M_f^2 + 2k|e_f|B}. \quad (16)$$

Minimizing the thermodynamic potential with respect to Φ , $\bar{\Phi}$, and σ_f , we get a set of coupled equations to be solved simultaneously [78,87]

$$\frac{\partial \Omega}{\partial \sigma_f} = 0, \quad \frac{\partial \Omega}{\partial \Phi} = 0, \quad \frac{\partial \Omega}{\partial \bar{\Phi}} = 0.$$

Further, the condensates split under the effect of an external magnetic field as

$$\sigma_f = \sigma_f^{\text{vac}} + \sigma_f^{\text{med}} + \sigma_f^{\text{mag}}$$

where σ_f^{vac} , σ_f^{med} , and σ_f^{mag} are the contributions from vacuum, medium, and the magnetic field, respectively, which are given as [9,73,82,91,95]

$$\begin{aligned} \sigma_f^{\text{vac}} &= -6 \int_{\Lambda} \frac{d^3 p}{(2\pi)^3} \frac{M_f}{E_f}, \\ \sigma_f^{\text{med}} &= \frac{3(|e_f B|)^2}{4\pi^2} \sum_k a(k) \int_{-\infty}^{\infty} \frac{dp_z}{E_f} (f_{\Phi}^+ + f_{\Phi}^-), \\ \sigma_f^{\text{mag}} &= -\frac{M|e_f B N_c}{2\pi^2} \left[\ln \Gamma(x_f) - \frac{1}{2} \ln(2\pi) \right. \\ &\quad \left. + x_f - \frac{1}{2}(2x_f - 1) \ln(x_f) \right]. \end{aligned} \quad (17)$$

Any thermodynamic quantity can be derived from the thermodynamic potential Ω . In the presence of a magnetic field, pressure becomes anisotropic due to the Landau quantization. Pressure along the direction of the magnetic field is longitudinal pressure (P^{\parallel}) and in the direction perpendicular to the magnetic field it is transverse pressure (P^{\perp}), and can be expressed as [14,102]

$$\begin{aligned} P^{\parallel} &= -\Omega - \frac{(eB)^2}{2}, \\ P^{\perp} &= -\Omega - eB\mathcal{M} + \frac{(eB)^2}{2}, \end{aligned} \quad (18)$$

where \mathcal{M} is the magnetization,

$$\mathcal{M} = -\left(\frac{\partial \Omega}{\partial eB} \right)_{\mu} \quad (19)$$

The normalized pressure is taken in a way that it vanishes at vacuum [9,91,93,95]

$$P_N^{\parallel} = P^{\parallel}(T, \mu, eB) - P^{\parallel}(0.001, 0, eB)$$

$$P_N^{\perp} = P^{\perp}(T, \mu, eB) - P^{\perp}(0.001, 0, eB).$$

In the weak magnetic field limit, for simplicity, we can take $P_N^{\perp} \simeq P_N^{\parallel} = P_N$. Using the Euler thermodynamic relation, energy density can be written as

$$\epsilon = -P_N + Ts + MeB. \quad (20)$$

The speed of sound also becomes anisotropic in the presence of a magnetic field due to changes in pressure in the longitudinal and transverse directions. The anisotropic squared speed of sound reads as [103,104]

$$c_s^{2\parallel(\perp)} = \left(\frac{dP_N^{\parallel(\perp)}}{d\epsilon} \right)_{s/n}, \quad (21)$$

$$c_s^{2\parallel(\perp)} = \frac{\frac{\partial P_N^{\parallel(\perp)}}{\partial T} + \frac{\partial P_N^{\parallel(\perp)}}{\partial eB} \frac{deB}{dT}}{\frac{\partial \epsilon}{\partial T} + \frac{\partial \epsilon}{\partial eB} \frac{deB}{dT}}, \quad (22)$$

where

$$\frac{deB}{dT} = \frac{s \frac{\partial n}{\partial T} - n \frac{\partial s}{\partial T}}{n \frac{\partial s}{\partial eB} - s \frac{\partial n}{\partial eB}}.$$

Isothermal compressibility can be expressed as in Eq. (11) for the vanishing magnetic field. For a system with a finite magnetic field, the pressure term in the denominator splits isothermal compressibility into longitudinal and transverse components, which can be expressed as [14]

$$\kappa_T^{\parallel} = \frac{1}{n} \frac{\frac{\partial n}{\partial \mu}}{\frac{\partial P^{\parallel}}{\partial \mu}} = \frac{1}{n^2} \frac{\partial n}{\partial \mu}.$$

$$\kappa_T^{\perp} = \frac{1}{n} \frac{\frac{\partial n}{\partial \mu}}{\frac{\partial P^{\perp}}{\partial \mu}} = \frac{1}{n} \frac{\partial n}{\partial \mu} \left(\frac{1}{n - \frac{\partial \mathcal{M}}{\partial \mu} B} \right). \quad (23)$$

III. RESULTS AND DISCUSSION

In the presence of a magnetic field, c_s^2 becomes anisotropic. It signifies that any disturbance in the medium will not travel at the same speed in all directions in the presence of an external magnetic field. In Fig. 1, longitudinal (\parallel) (left) and transverse (\perp) (right) components of c_s^2 are plotted against temperature for different values of magnetic fields; $eB = 0$ (black dotted line), $eB = 0.2 \text{ GeV}^2$ (red dashed line), and $eB = 0.4 \text{ GeV}^2$ (solid blue line). Lattice estimation at the zero magnetic fields is also presented here from the Wuppertal-Budapest (WB) group [105]. We also compare with the latest CMS estimation of speed of sound in a central Pb-Pb collision at $\sqrt{s_{NN}} = 5.02 \text{ TeV}$ [106], where a relation between $\partial P / \partial \epsilon$ and experimental observables such as N_{ch} and $\langle p_T \rangle$ has been established to obtain a c_s^2 estimation from the CMS data. The study is done for a head-on collision where no magnetic field is expected and as a result, our estimations for the $eB = 0.0 \text{ GeV}^2$ case match with the results taken from the CMS data. Moreover, it will be interesting to perform similar studies for peripheral heavy-ion collisions where strong transient magnetic field is expected. Estimating the speed of sound for such collisions will shed light on the effect of the magnetic field on the speed of sound. At zero magnetic field, the PNJL result agrees with lattice estimation at the high temperature. The dip around the transition temperature (T_c) in the PNJL result is the consequence of the deconfinement phase transition that is absent in NJL estimations [40]. With the increasing magnetic field, the dip shifts towards lower temperatures. This signifies that with the increasing magnetic field, the deconfinement temperature decreases, which agrees with the lattice results obtained in Ref. [107].

In a massless noninteracting (ideal) gas, the speed of sound is estimated to be $1/3$ [108]. In Fig. 1, we observe that as the temperature increases, chiral symmetry is restored, and c_s^2 approaches the ideal gas limit. We observe an interesting trend of c_s^2 near the transition temperature.

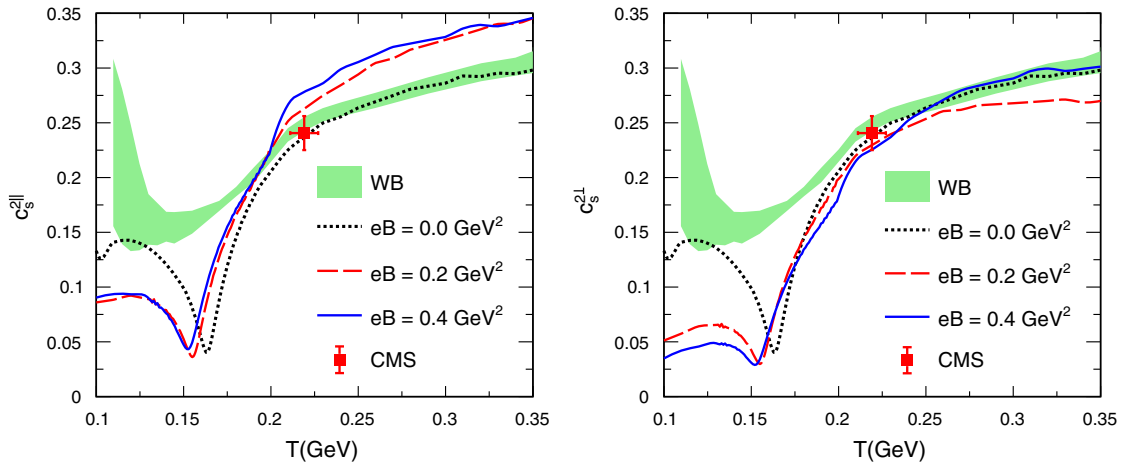


FIG. 1. Squared speed of sound's longitudinal ($c_s^{2\parallel}$) (left) and transverse ($c_s^{2\perp}$) (right) components against temperature (T).

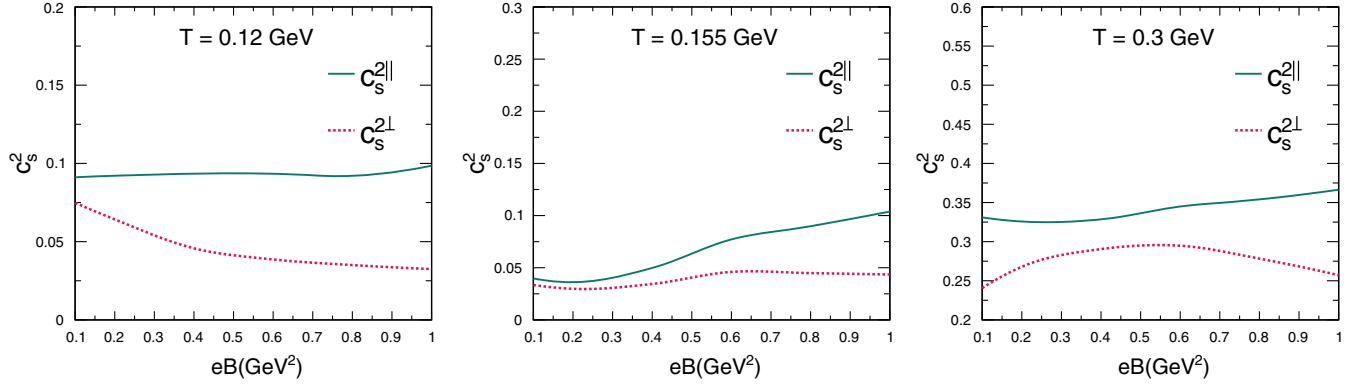


FIG. 2. Squared speed of sound ($c_s^{2||}$ and $c_s^{2\perp}$) against the magnetic field (eB) for different temperatures: $T = 0.12$ GeV (left), $T = 0.155$ GeV (middle), and $T = 0.3$ GeV (right).

For $T < T_c$ (or below the dip in the speed of sound), c_s^2 decreases with increasing magnetic field, i.e., the speed of sound shows an order as $(c_s^2)^{B=0} > (c_s^2)^{B=0.2} > (c_s^2)^{B=0.4}$. However, slightly above T_c , the speed of sound is higher at a larger magnetic field, i.e., the speed of sound shows an order as $(c_s^2)^{B=0.4} > (c_s^2)^{B=0.2} > (c_s^2)^{B=0}$. This behavior is more prominent in the parallel component of c_s^2 . This behavior is a consequence of magnetic catalysis (MC) and inverse magnetic catalysis (IMC) phenomena. Because of magnetic catalysis, at low temperatures, the quark condensate is enhanced in the presence of a magnetic field, and as a result, the dynamical mass of the quarks is directly proportional to the magnetic field strength. However, due to inverse magnetic catalysis, beyond T_c , the quark condensate is reduced with an increasing magnetic field and the mass is inversely proportional to the external magnetic field. The same behavior can be seen for the speed of sound near T_c , thus hinting at the fact that speed of sound is sensitive to mass ordering. These phenomena have been widely explored in the literature [18,91,109].

In Fig. 2, c_s^2 is plotted against eB for three different values of temperatures, such that $T = 0.12$ GeV (left) represents the confined or hadronic medium, $T = 0.155$ GeV (middle) is around T_c , and $T = 0.3$ GeV (right) represents the deconfined or partonic medium. \parallel and \perp components are represented by a solid green and red dotted line, respectively. c_s^2 shows a nonmonotonic behavior with respect to eB . The \parallel component is almost independent of the magnetic field because, along the magnetic field, direct contribution from Lorentz force vanishes. In the confinement zone, the effect of the magnetic field is large in the \perp component. We observe a decrease in the transverse component of the speed of sound with an increase in magnetic field strength. Around the transition temperature, the effect of the magnetic field on both of the components is not significant. However, there is a steady rise for the longitudinal component at a strong magnetic field. On the right, for the partonic phase, we have both weak and strong field regions. In the weak field region, we observe an

increasing and then decreasing behavior that could be attributed to the interplay of the decreasing coupling constant with an increase in eB and dependence of dynamical mass on the magnetic field strength.

To better understand the anisotropy in the components of c_s^2 due to magnetic field, we have plotted $\Delta c_s^2 / c_s^{2||}$ with T for three different values of magnetic field ($eB = 0.1, 0.2, 0.4$ GeV²) in Fig. 3 (left). On the right panel, we have plotted the same against eB for different temperatures ($T = 0.12, 0.155, 0.3$ GeV). Here, $\Delta c_s^2 = c_s^{2||} - c_s^{2\perp}$ represents normalized anisotropy in the components of c_s^2 . In other words, Fig. 3 represents the deviation of the \perp component from the \parallel component in the presence of magnetic fields. In the left panel, anisotropy is larger at a low temperature and decreases as temperature increases. For $eB = 0.2, 0.4$ GeV², anisotropy is around 20% to 60%, respectively, at $T = 0.1$ GeV. The degree of anisotropy decreases with temperature and obtains minima near the transition temperature. The position of the minima changes with magnetic field strength. With an increase in temperature, the deviation increases as well. Moreover, we observe two kink-like structures, one near 150–160 MeV and a second one near 200–220 MeV. The first one is due to the effect of the deconfinement temperature, and the second one emerges due to the change in the system near the chiral transition temperature. In Ref. [110], the authors have explicitly shown the effect of the deconfinement transition temperature (maxima of the Polyakov loop susceptibility) and chiral transition temperature (maxima of the quark condensate susceptibility) in the speed of sound and specific heat.

The figure on the right shows that anisotropy increases as the magnetic field increases. Anisotropy is smaller at low eB for all three temperatures and increases gradually with increasing eB . Therefore, from both figures, we can conclude that anisotropy is higher in the strong field zone, i.e., low T and high eB . With the decreasing magnetic field, anisotropy diminishes, and \parallel and \perp components merge to

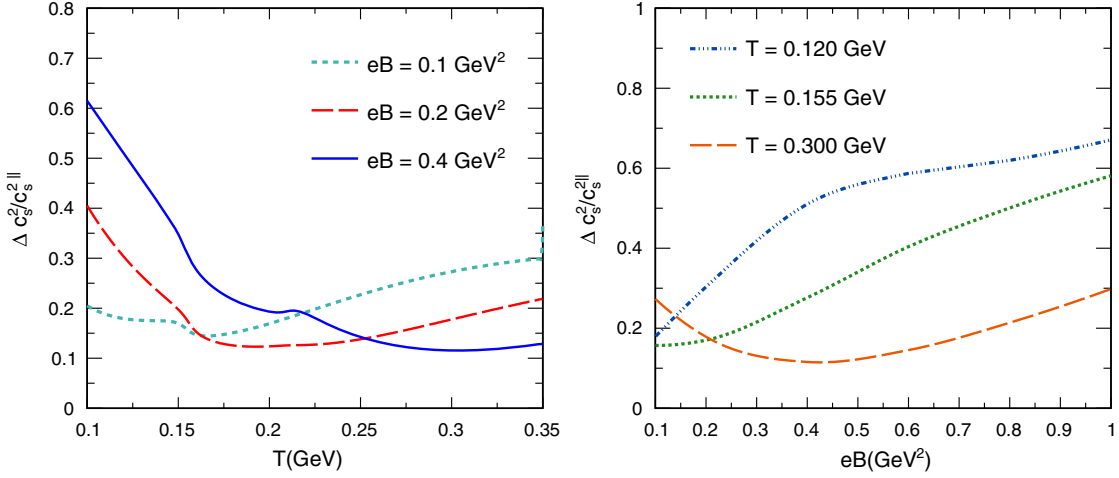


FIG. 3. Anisotropy in squared speed of sound ($\Delta c_s^2/c_s^2||$) against T (left) and eB (right). Here, $\Delta c_s^2 = c_s^{2||} - c_s^{2\perp}$.

the isotropic component of c_s^2 as soon as the magnetic field vanishes.

In Fig. 4, κ_T of QCD matter is plotted against temperature for QGP using the PNJL model. We compare our results with experimental results and different models. Experimentally, the isothermal compressibility is estimated using charged particle multiplicity fluctuations in Pb-Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV [111]. However, this estimation requires the information of temperature and volume of the system at chemical freeze-out which is taken from the statistical hadronization model. Results from hadronic models Hadron resonance gas (HRG) (green dotted line), Excluded volume hadron resonance gas (EV-HRG) (solid black line) [26], and partonic models CSPM (solid violet circle) [25], and the PNJL (blue dash

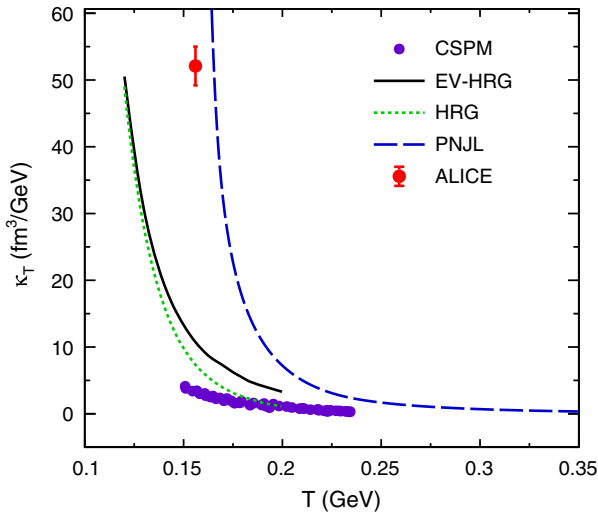


FIG. 4. Isothermal compressibility (κ_T) against temperature (T) at zero magnetic field. A comparison of this work (PNJL) with different models (HRG, EV-HRG [26], and CSPM [25]) and experimental results from ALICE [111].

line) in the absence of magnetic field are presented for comparison. κ_T indicates stiffness in the equation of state (EOS). The smaller the value of κ_T , the stiffer the EOS is. Here, κ_T decreases as T increases, which means at very high T , the QCD matter is highly incompressible. This makes the QGP matter almost a perfect fluid. In the absence of magnetic fields, κ_T is isotropic. The EOS is the same in all directions. As we introduce the magnetic field, κ_T becomes anisotropic, and we get two independent components. Along the magnetic field, EOS is defined by the $\kappa_T^||$ component, and in the plane perpendicular to the magnetic field, it is κ_T^\perp .

In Fig. 5, we have plotted $\kappa_T^||$ (left) and κ_T^\perp (right) as a function of T for $eB = 0, 0.2, 0.4 \text{ GeV}^2$. κ_T is a proxy to the phase transition. For the baryon-free ($\mu_B = 0$) QCD matter, our results show a smooth transition from the deconfinement to confinement temperature zone. Even in the presence of a magnetic field, both the components show a smooth transition or crossover phase transition. A similar behavior as Fig. 3 of a kinklike structure can be seen here as well, where the kinks arise as a result of the effect of the deconfinement temperature and chiral transition temperature in the medium. However, one recent study [14] in the NJL model discovered a first-order phase transition in the $\kappa_T^||$ component in the presence of a magnetic field at very low temperature and finite baryon chemical potential, which may not be realized in RHIC or LHC energies. Here, we see that the magnetic field reduces the compressibility, and the effect is nearly independent of T .

In Fig. 6, we have plotted $\kappa_T^||$ (green solid line) and κ_T^\perp (red dotted line) as a function of eB for $T = 0.12 \text{ GeV}$ (left), $T = 0.155 \text{ GeV}$ (middle), and $T = 0.3 \text{ GeV}$ (right). At any fixed temperature, both the components decrease with increasing magnetic field. In the middle figure, where T is around T_c , a peaklike structure is seen in $\kappa_T^||$. The peak emerges as a result of deconfinement in the system.

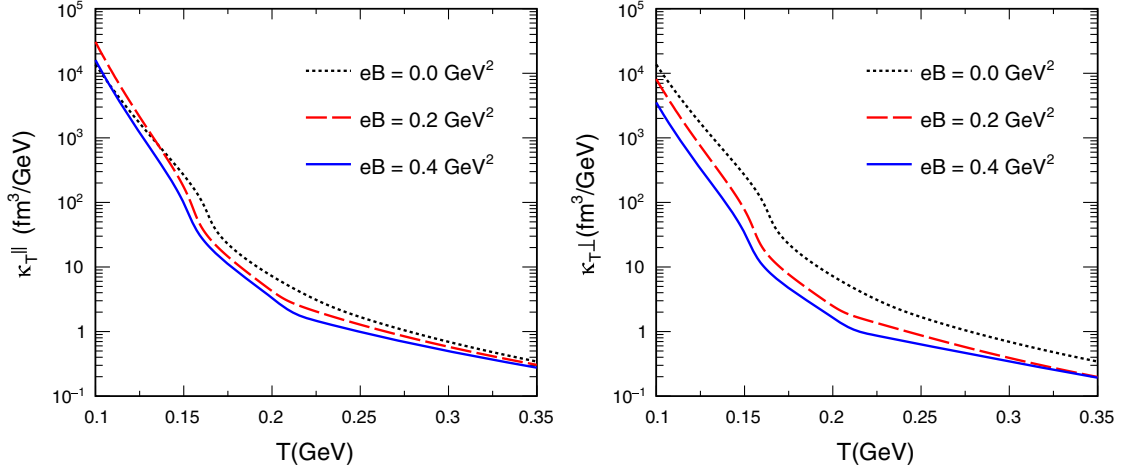


FIG. 5. Isothermal compressibility: longitudinal component (κ_T^{\parallel}) (left) and transverse component (κ_T^{\perp}) (right) against temperature (T) for different value of magnetic field (eB).

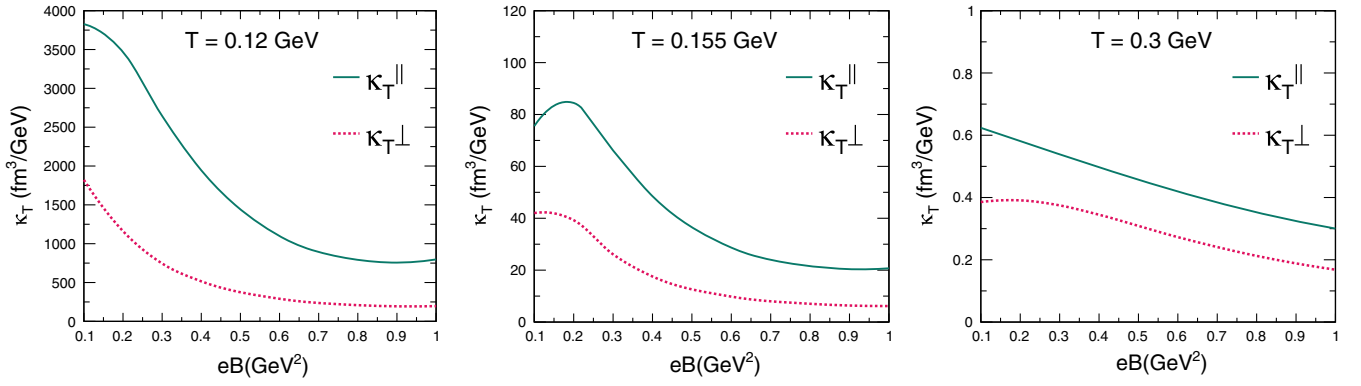


FIG. 6. κ_T^{\parallel} and κ_T^{\perp} against eB at $T = 0.12 \text{ GeV}$ (left), $T = 0.155 \text{ GeV}$ (middle), and $T = 0.3 \text{ GeV}$ (right).

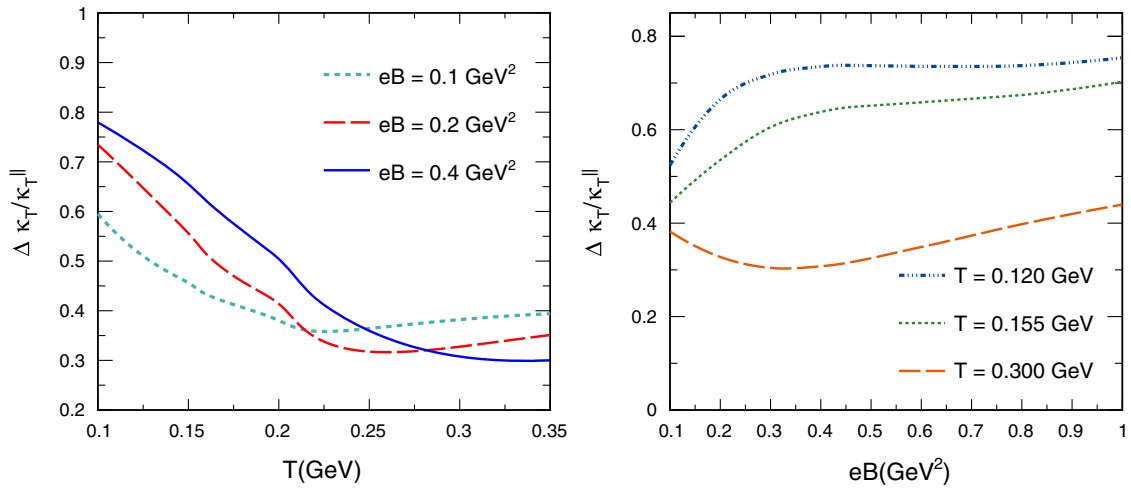


FIG. 7. Anisotropy in isothermal compressibility ($\Delta \kappa_T / \kappa_T^{\parallel}$) as a function of temperature (T) (left) and magnetic field (eB) (right), $\Delta \kappa_T = \kappa_T^{\parallel} - \kappa_T^{\perp}$.

Similarly, a kinklike structure appears when it is plotted against temperature (Fig. 5), and the kink moves towards a lower temperature for a higher magnetic field. Here, for $eB \simeq 0.2 \text{ GeV}^2$, the deconfinement temperature is around $T \simeq 0.155 \text{ GeV}$, which gives rise to the observed peak. For a lower temperature, the peak will shift towards a higher eB value. In the low-temperature regime, saturation in κ_T can be observed for both parallel and perpendicular components. The value of the \perp component for κ_T is smaller compared to the \parallel component. This signifies that along the magnetic field, the EOS is stiffer.

The anisotropy between the two components is very prominent. In order to quantify the anisotropy in κ_T we have plotted $\Delta\kappa_T/\kappa_T^\parallel$ as a function of T (left) and eB (right) in Fig. 7, where $\Delta\kappa_T = \kappa_T^\parallel - \kappa_T^\perp$. Similar to c_s^2 in Fig. 3, here also we see that anisotropy decreases with increasing T at a fixed eB , reaching minima. For temperatures higher than T_c , there is an increase in anisotropic behavior with temperature for each magnetic field value. On the right, we have studied the degree of anisotropy as a function of eB . Anisotropy increases with eB at a fixed T . Anisotropy is maximum in the strong field zone. Compared to c_s^2 , anisotropy is higher in κ_T . In the low T zone, anisotropy in κ_T varies from 60% to 80%, and in high T , it is 30% to 40% in the mentioned magnetic field regime. In the case of c_s^2 , the effect is comparatively 10% to 20% less. This signifies that the magnetic field effect is more prominent in the anisotropy of κ_T than that of c_s^2 .

IV. SUMMARY AND OUTLOOK

In summary, we have estimated the squared speed of sound (c_s^2) and isothermal compressibility (κ_T) of baryon-free QCD matter in the presence of a magnetic field using the PNJL model. With the increasing temperature, c_s^2 approaches the conformal limit. It shows a dip around the T_c , which shifts to a lower temperature with the increasing magnetic field. The effect of MC and IMC is also found in the components of c_s^2 near the transition temperature. The isothermal compressibility is estimated within the PNJL model for the first time. Our results match with other results in the literature qualitatively, confirming that the QCD matter is the most perfect fluid found in nature. With the increasing temperature, κ_T decreases. Smooth transition of longitudinal and transverse components from the confinement to deconfinement phase indicates crossover transition in the presence of magnetic field at zero chemical potential. Furthermore, we noticed that

anisotropy is higher in κ_T compared to c_s^2 , and in the strong field limit, the lowest Landau level is a good approximation. We found that the degree of anisotropy increases with the magnetic field and shows minima near T_c when plotted against temperature for both thermodynamical quantities.

The anisotropy in the EOS, due to the magnetic field, is not directly measurable in experiments. However, the observable, directed flow is a promising probe to the initial magnetic field [3,4]. Recent studies [112,113] show that the directed flow is sensitive to the EOS. The EOS is affected significantly by the external magnetic field. In principle, one should be able to extract information about the effect of the magnetic field on the hot QCD matter by studying the splitting of the directed flow of particle and antiparticle pair. Experimentally, at LHC and RHIC, the splitting of the directed flow of D^0 and \bar{D}^0 has been observed [3,4].

As a preliminary study, our work concentrates only on baryon-free QCD matter at a constant magnetic field. This work can be extended to the baryonic matter with a time-dependent magnetic field, which applies to matter created at the RHIC and energies achievable below the RHIC. Moreover, in this study, the upper limit of the magnetic field is around 1 GeV owing to the fact that $G(eB)$ is obtained by fitting LQCD data that is available up to 1 GeV² only. However, the magnetic field generated at the RHIC and LHC is around 0.2–0.4 GeV². Thus making 1 GeV² a suitable upper limit. A further theoretical improvement would also enable us to explore the high eB region. In addition, as mentioned in [53], the dependence of G on μ and Φ has been neglected for simplicity. One can find different forms of G in use, but for practical purposes, G should be a function of all the parameters, i.e., T , μ , eB , and Φ , to successfully explain the hot QCD matter. However, considering that the eB dependent coupling constant contemplated in this work can reproduce LQCD results qualitatively well, one can proceed with this approximation for further phenomenological studies to explore the QCD medium.

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- [1] W.-T. Deng and X.-G. Huang, *Phys. Rev. C* **85**, 044907 (2012).
- [2] V. Skokov, A. Y. Illarionov, and V. Toneev, *Int. J. Mod. Phys. A* **24**, 5925 (2009).
- [3] J. Adam *et al.* (STAR Collaboration), *Phys. Rev. Lett.* **123**, 162301 (2019).
- [4] S. Acharya *et al.* (ALICE Collaboration), *Phys. Rev. Lett.* **125**, 022301 (2020).
- [5] K. Fukushima, D. E. Kharzeev, and H. J. Warringa, *Phys. Rev. D* **78**, 074033 (2008).
- [6] G. S. Bali, F. Bruckmann, G. Endrodi, Z. Fodor, S. D. Katz, S. Krieg, A. Schafer, and K. K. Szabo, *J. High Energy Phys.* **02** (2011) 044.
- [7] A. N. Tawfik, A. M. Diab, N. Ezzelarab, and A. G. Shalaby, *Adv. High Energy Phys.* **2016**, 1381479 (2016).
- [8] G. S. Bali, F. Bruckmann, G. Endrödi, S. D. Katz, and A. Schäfer, *J. High Energy Phys.* **08** (2014) 177.
- [9] M. Ferreira, P. Costa, and C. Providência, *Phys. Rev. D* **89**, 036006 (2014).
- [10] A. Bandyopadhyay, S. Ghosh, R. L. S. Farias, J. Dey, and G. a. Krein, *Phys. Rev. D* **102**, 114015 (2020).
- [11] A. Dash, S. Samanta, J. Dey, U. Gangopadhyaya, S. Ghosh, and V. Roy, *Phys. Rev. D* **102**, 016016 (2020).
- [12] J. Dey, S. Samanta, S. Ghosh, and S. Satapathy, *Phys. Rev. C* **106**, 044914 (2022).
- [13] E. J. Ferrer and A. Hackebill, *Nucl. Phys.* **A1031**, 122608 (2023).
- [14] L. Yang and X.-J. Wen, *Phys. Rev. D* **104**, 114010 (2021).
- [15] A. J. Mizher, M. N. Chernodub, and E. S. Fraga, *Phys. Rev. D* **82**, 105016 (2010).
- [16] R. Gatto and M. Ruggieri, *Phys. Rev. D* **82**, 054027 (2010).
- [17] R. Gatto and M. Ruggieri, *Phys. Rev. D* **83**, 034016 (2011).
- [18] M. Ferreira, P. Costa, O. Lourenço, T. Frederico, and C. Providência, *Phys. Rev. D* **89**, 116011 (2014).
- [19] W. R. Tavares, R. L. S. Farias, S. S. Avancini, V. S. Timóteo, M. B. Pinto, and G. a. Krein, *Eur. Phys. J. A* **57**, 278 (2021).
- [20] B. Stokić, B. Friman, and K. Redlich, *Phys. Lett. B* **673**, 192 (2009).
- [21] A. Mekjian, S. Lee, and L. Zamick, *Phys. Lett. B* **621**, 239 (2005).
- [22] M. Mukherjee, S. Basu, A. Chatterjee, S. Chatterjee, S. P. Adhya, S. Thakur, and T. K. Nayak, *Phys. Lett. B* **784**, 1 (2018).
- [23] F. Karsch, Lattice QCD at high temperature and density, in *Lectures on Quark Matter*, edited by W. Plessas and L. Mathelitsch (Springer, Berlin, Heidelberg, 2002), pp. 209–249.
- [24] S. Jain, R. Gupta, and S. Jena, *Universe* **9**, 170 (2023).
- [25] D. Sahu, S. Tripathy, R. Sahoo, and S. K. Tiwari, *Eur. Phys. J. A* **58**, 78 (2022).
- [26] A. Khuntia, S. K. Tiwari, P. Sharma, R. Sahoo, and T. K. Nayak, *Phys. Rev. C* **100**, 014910 (2019).
- [27] S. Pal, G. Kadam, and A. Bhattacharyya, *Nucl. Phys.* **A1023**, 122464 (2022).
- [28] M. Mukherjee, S. Basu, A. Chatterjee, S. Chatterjee, S. P. Adhya, S. Thakur, and T. K. Nayak, *Phys. Lett. B* **784**, 1 (2018).
- [29] D. Sahu, S. Tripathy, R. Sahoo, and A. R. Dash, *Eur. Phys. J. A* **56**, 187 (2020).
- [30] M. Albright and J. I. Kapusta, *Phys. Rev. C* **93**, 014903 (2016).
- [31] R. Campanini and G. Ferri, *Phys. Lett. B* **703**, 237 (2011).
- [32] F. G. Gardim, G. Giacalone, M. Luzum, and J.-Y. Ollitrault, *Nat. Phys.* **16**, 615 (2020).
- [33] D. Biswas, K. Deka, A. Jaiswal, and S. Roy, *Phys. Rev. C* **102**, 014912 (2020).
- [34] V. Mykhaylova and C. Sasaki, *Phys. Rev. D* **103**, 014007 (2021).
- [35] O. Soloveva, J. Aichelin, and E. Bratkovskaya, *Phys. Rev. D* **105**, 054011 (2022).
- [36] D. Sahu and R. Sahoo, *J. Phys. G* **48**, 125104 (2021).
- [37] P. Sahoo, S. K. Tiwari, S. De, R. Sahoo, R. P. Scharenberg, and B. K. Srivastava, *Mod. Phys. Lett. A* **34**, 1950034 (2019).
- [38] R. Venugopalan and M. Prakash, *Nucl. Phys.* **A546**, 718 (1992).
- [39] M. Bluhm, P. Alba, W. Alberico, A. Beraudo, and C. Ratti, *Nucl. Phys.* **A929**, 157 (2014).
- [40] P. Deb, G. P. Kadam, and H. Mishra, *Phys. Rev. D* **94**, 094002 (2016).
- [41] R. Marty, E. Bratkovskaya, W. Cassing, J. Aichelin, and H. Berrehrah, *Phys. Rev. C* **88**, 045204 (2013).
- [42] M. Motta, R. Stiele, W. M. Alberico, and A. Beraudo, *Eur. Phys. J. C* **80**, 770 (2020).
- [43] S. K. Ghosh, T. K. Mukherjee, M. G. Mustafa, and R. Ray, *Phys. Rev. D* **73**, 114007 (2006).
- [44] W.-b. He, G.-y. Shao, X.-y. Gao, X.-r. Yang, and C.-l. Xie, *Phys. Rev. D* **105**, 094024 (2022).
- [45] Y. Aoki, G. Endrodi, Z. Fodor, S. D. Katz, and K. K. Szabo, *Nature (London)* **443**, 675 (2006).
- [46] A. Bazavov *et al.* (HotQCD Collaboration), *Phys. Rev. D* **90**, 094503 (2014).
- [47] S. Borsanyi, Z. Fodor, C. Hoelbling, S. D. Katz, S. Krieg, and K. K. Szabo, *Phys. Lett. B* **730**, 99 (2014).
- [48] S. Borsanyi, Z. Fodor, J. N. Guenther, R. Kara, S. D. Katz, P. Parotto, A. Pasztor, C. Ratti, and K. K. Szabo, *Phys. Rev. Lett.* **125**, 052001 (2020).
- [49] Y. Nambu and G. Jona-Lasinio, *Phys. Rev.* **122**, 345 (1961).
- [50] Y. Nambu and G. Jona-Lasinio, *Phys. Rev.* **124**, 246 (1961).
- [51] S. P. Klevansky, *Rev. Mod. Phys.* **64**, 649 (1992).
- [52] P. N. Meisinger and M. C. Ogilvie, *Phys. Rev. D* **65**, 056013 (2002).
- [53] K. Fukushima, *Phys. Lett. B* **591**, 277 (2004).
- [54] A. M. Polyakov, *Phys. Lett.* **72B**, 477 (1978).
- [55] L. Susskind, *Phys. Rev. D* **20**, 2610 (1979).
- [56] B. Svetitsky and L. G. Yaffe, *Nucl. Phys.* **B210**, 423 (1982).
- [57] B. Svetitsky, *Phys. Rep.* **132**, 1 (1986).
- [58] K. Fukushima, *J. Phys. G* **39**, 013101 (2012).
- [59] C. Ratti, M. A. Thaler, and W. Weise, *Phys. Rev. D* **73**, 014019 (2006).
- [60] S. Mukherjee, M. G. Mustafa, and R. Ray, *Phys. Rev. D* **75**, 094015 (2007).
- [61] C. Ratti, S. Roessner, and W. Weise, *Phys. Lett. B* **649**, 57 (2007).

- [62] A. Bhattacharyya, P. Deb, S. K. Ghosh, and R. Ray, *Phys. Rev. D* **82**, 014021 (2010).
- [63] A. Bhattacharyya, S. K. Ghosh, S. Maity, S. Raha, R. Ray, K. Saha, and S. Upadhaya, *Phys. Rev. D* **95**, 054005 (2017).
- [64] A. Bhattacharyya, S. K. Ghosh, S. Maity, S. Raha, R. Ray, K. Saha, S. Samanta, and S. Upadhaya, *Phys. Rev. C* **99**, 045207 (2019).
- [65] A. Bhattacharyya, P. Deb, S. K. Ghosh, S. Maity, S. Raha, R. Ray, K. Saha, and S. Upadhaya, *Phys. Rev. D* **102**, 074006 (2020).
- [66] C. A. Islam, J. Dey, and S. Ghosh, *Phys. Rev. C* **103**, 034904 (2021).
- [67] S. Roessner, C. Ratti, and W. Weise, *Phys. Rev. D* **75**, 034007 (2007).
- [68] K. Fukushima, *Phys. Rev. D* **77**, 114028 (2008); **78**, 039902(E) (2008).
- [69] S. Carignano, D. Nickel, and M. Buballa, *Phys. Rev. D* **82**, 054009 (2010).
- [70] T. Hell, S. Roessner, M. Cristoforetti, and W. Weise, *Phys. Rev. D* **79**, 014022 (2009).
- [71] M. Ferreira, P. Costa, D. P. Menezes, C. Providência, and N. Scoccola, *Phys. Rev. D* **89**, 016002 (2014); **89**, 019902 (A) (2014).
- [72] N. Chaudhuri, S. Ghosh, P. Roy, and S. Sarkar, *Phys. Rev. D* **106**, 056020 (2022).
- [73] Y. Wang and X.-J. Wen, *J. Phys. G* **47**, 105201 (2020).
- [74] P. N. Meisinger and M. C. Ogilvie, *Phys. Lett. B* **379**, 163 (1996).
- [75] P. N. Meisinger and M. C. Ogilvie, *Nucl. Phys. B, Proc. Suppl.* **47**, 519 (1996).
- [76] H. Hansen, W. M. Alberico, A. Beraudo, A. Molinari, M. Nardi, and C. Ratti, *Phys. Rev. D* **75**, 065004 (2007).
- [77] K. Fukushima, M. Ruggieri, and R. Gatto, *Phys. Rev. D* **81**, 114031 (2010).
- [78] W.-j. Fu, Z. Zhang, and Y.-x. Liu, *Phys. Rev. D* **77**, 014006 (2008).
- [79] M. Ciminale, R. Gatto, N. D. Ippolito, G. Nardulli, and M. Ruggieri, *Phys. Rev. D* **77**, 054023 (2008).
- [80] P. Rehberg, S. P. Klevansky, and J. Hufner, *Phys. Rev. C* **53**, 410 (1996).
- [81] T. Hatsuda and T. Kunihiro, *Phys. Rep.* **247**, 221 (1994).
- [82] J. M. Torres-Rincon and J. Aichelin, *Phys. Rev. C* **96**, 045205 (2017).
- [83] C. Sasaki, B. Friman, and K. Redlich, *Phys. Rev. D* **75**, 074013 (2007).
- [84] S. K. Ghosh, T. K. Mukherjee, M. G. Mustafa, and R. Ray, *Phys. Rev. D* **77**, 094024 (2008).
- [85] B.-J. Schaefer, J. M. Pawłowski, and J. Wambach, *Phys. Rev. D* **76**, 074023 (2007).
- [86] J.-L. Zhang, C.-M. Li, and H.-S. Zong, *Chin. Phys. C* **42**, 123105 (2018).
- [87] P. Costa, M. C. Ruivo, C. A. de Sousa, H. Hansen, and W. M. Alberico, *Phys. Rev. D* **79**, 116003 (2009).
- [88] D. Fuseau, T. Steinert, and J. Aichelin, *Phys. Rev. C* **101**, 065203 (2020).
- [89] S. S. Avancini, D. P. Menezes, and C. Providencia, *Phys. Rev. C* **83**, 065805 (2011).
- [90] R. L. S. Farias, K. P. Gomes, G. I. Krein, and M. B. Pinto, *Phys. Rev. C* **90**, 025203 (2014).
- [91] R. L. S. Farias, V. S. Timoteo, S. S. Avancini, M. B. Pinto, and G. Krein, *Eur. Phys. J. A* **53**, 101 (2017).
- [92] L. Yang and X.-J. Wen, *Phys. Rev. D* **96**, 056023 (2017).
- [93] C.-F. Li, L. Yang, X.-J. Wen, and G.-X. Peng, *Phys. Rev. D* **93**, 054005 (2016).
- [94] J. a. Moreira, P. Costa, and T. E. Restrepo, *Phys. Rev. D* **102**, 014032 (2020).
- [95] D. P. Menezes, M. Benghi Pinto, S. S. Avancini, A. Perez Martinez, and C. Providencia, *Phys. Rev. C* **79**, 035807 (2009).
- [96] D. P. Menezes, M. Benghi Pinto, S. S. Avancini, and C. Providencia, *Phys. Rev. C* **80**, 065805 (2009).
- [97] P. Costa, H. Hansen, M. C. Ruivo, and C. A. de Sousa, *Phys. Rev. D* **81**, 016007 (2010).
- [98] K. Xue, X. Yu, and X. Wang, *Chin. Phys. C* **46**, 013103 (2022).
- [99] H. Kohyama, D. Kimura, and T. Inagaki, *Nucl. Phys.* **B896**, 682 (2015).
- [100] G. S. Pradhan, D. Sahu, S. Deb, and R. Sahoo, *J. Phys. G* **50**, 055104 (2023).
- [101] E. E, *Math. Comp.* 47 347, Digital Library of Mathematical Functions, National Institute of Standards and Technology <https://dlmf.nist.gov/> **132**, 1 (1986).
- [102] P.-C. Chu, B. Wang, Y.-Y. Jia, Y.-M. Dong, S.-M. Wang, X.-H. Li, L. Zhang, X.-M. Zhang, and H.-Y. Ma, *Phys. Rev. D* **94**, 123014 (2016).
- [103] S. Borsanyi, G. Endrodi, Z. Fodor, S. D. Katz, S. Krieg, C. Ratti, and K. K. Szabo, *J. High Energy Phys.* **08** (2012) 053.
- [104] O. F. Piattella, J. C. Fabris, and N. Bilić, *Classical Quantum Gravity* **31**, 055006 (2014).
- [105] S. Borsányi, Z. Fodor, C. Hoelbling, S. D. Katz, S. Krieg, and K. K. Szabó, *Phys. Lett. B* **730**, 99 (2014).
- [106] C. A. Bernardes, *arXiv:2312.11758*.
- [107] V. V. Braguta, M. N. Chernodub, A. Y. Kotov, A. V. Molochkov, and A. A. Nikolaev, *Phys. Rev. D* **100**, 114503 (2019).
- [108] M. Pardy, *Results Phys.* **3**, 70 (2013).
- [109] G. S. Bali, F. Bruckmann, G. Endrodi, Z. Fodor, S. D. Katz, and A. Schafer, *Phys. Rev. D* **86**, 071502 (2012).
- [110] M. Ferreira, P. Costa, and C. Providência, *Phys. Rev. D* **90**, 016012 (2014).
- [111] S. Acharya *et al.* (ALICE Collaboration), *Eur. Phys. J. C* **81**, 1012 (2021).
- [112] Y. Nara, H. Niemi, J. Steinheimer, and H. Stöcker, *Phys. Lett. B* **769**, 543 (2017).
- [113] P. Božek, *Phys. Rev. C* **106**, L061901 (2022).