Single phonon detection for dark matter via quantum evaporation and sensing of ³He

S. A. Lyon[®],^{*} Kyle Castoria[®],[†] and Ethan Kleinbaum[‡] Department of Electrical and Computer Engineering, Princeton University, Princeton, New Jersey 08544, USA

Zhihao Qin[®], Arun Persaud[®], and Thomas Schenkel[®]

Acceleration Technology and Applied Physics, Lawrence Berkeley National Laboratory, 1 Cyclotron Road, California 94720, USA

Kathryn M. Zurek

Walter Burke Institute for Theoretical Physics, California Institute of Technology, Pasadena, California 91125, USA

(Received 11 January 2022; revised 6 September 2022; accepted 20 September 2023; published 8 January 2024)

Dark matter is five times more abundant than ordinary visible matter in our Universe. While laboratory searches hunting for dark matter have traditionally focused on the electroweak scale, theories of low mass hidden sectors motivate new detection techniques. Extending these searches to lower mass ranges, well below $1 \text{ GeV}/c^2$, poses new challenges as rare interactions with standard model matter transfer progressively less energy to electrons and nuclei in detectors. Here, we propose an approach based on phonon-assisted quantum evaporation combined with quantum sensors for detection of desorption events via tracking of spin coherence. The intent of our proposed dark matter sensors is to extend the parameter space to energy transfers in rare interactions to as low as a few meV for detection of dark matter particles in the keV/c² mass range.

DOI: 10.1103/PhysRevD.109.023010

I. INTRODUCTION

Dark matter (DM) direct detection experiments have focused on detecting weakly interacting massive particles via nuclear recoils (see, e.g., Ref. [1] for a review), where DM with mass in the 100 GeV range deposits energy by elastic scattering. However, in theories with low-mass hidden sectors (called a hidden valley), thermal DM can be much lighter, even down to a keV in mass where it carries meV of kinetic energy $(\frac{1}{2}m_X v_X^2, \text{ with } v_X \simeq 10^{-3}c)$. As the mass of the DM drops below approximately 10 GeV, the detection of rare scattering events with target nuclei falls below detection thresholds, and target nuclei absorb a very small fraction of the DM kinetic energy; see Ref. [2] for a review. At lower energies, electron recoils with energy transfer thresholds in the 1 eV range can be detected with sensitive charge coupled devices (CCD) counting electronhole pairs in semiconductors (e.g., [3]) or athermal phonon detectors (e.g., [4]). However, dark matter events have not yet been observed in these energy ranges, and it is desirable to probe thermal DM as light as 1 keV. Thus developing systems that can detect rare events with even lower deposited energy is an important goal.

In solids and liquids the lower energy excitations are generally phonons [5] (and rotons in superfluid helium [6,7]). Ionic crystals (polar materials) are especially interesting as detectors, since they enable new pathways for interaction with DM [5,8–10]. One challenge to sensing these phonons is that they are itinerant. Initially generated optical phonons rapidly decay to acoustic phonons, which disperse the deposited energy throughout the detection medium. The development of very sensitive and optimized detectors for quasiparticles and phonons using transition edge sensors and superconducting nanowire detectors is underway [11].

Here we propose an alternative, novel detection concept for single low-energy phonons based on the quantum sensing of the spin of ³He atoms which have been

^{*}Corresponding author: lyon@princeton.edu

Also at EeroQ Quantum Hardware, Chicago, Illinois 60651, USA.

[†]Present address: EeroQ Quantum Hardware, Chicago, Illinois 60651, USA.

[‡]Present address: Physical Electronics, Chanhassen, Minnesota 55317, USA.

Published by the American Physical Society under the terms of the Creative Commons Attribution 4.0 International license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP³.

evaporated from the surface of a He van der Waals film coating an ionic crystal. This is related to earlier proposals based upon He quantum evaporation [7,12], though here we consider ³He which is bound to a ⁴He surface with an energy of ~5 K [13], somewhat less than the ~7 K binding of a ⁴He atom. More importantly, the nuclear spin of ³He allows its quantum sensing at the level of single atoms. Calculations of DM scattering in an appropriate target crystal [8,9] and the analysis of the detector concept discussed here suggest a DM scattering count rate of about 40/hr for a 1 kg target.

A diagram of this concept is shown in Fig. 1. There are four major steps in the dark matter detector proposed here and shown in the figure: (1) production of phonons through the interaction with dark matter leading to the quantum evaporation of ³He atoms from Andreev bound states [14]; (2) trapping the 3 He on the detector surface using electrons bound to a film of isotopically enriched liquid ⁴He; (3) collecting and transporting the electrons and trapped ³He atoms to a detector structure; and (4) quantum sensing of the ³He atoms through their nuclear spin. An important feature of this detection concept is the separation of the dark matter absorber (i.e., target, such as a polar material) and the ³He detector, which opens the possibility to readily select and test a variety of absorber materials for specific dark matter searches. Further, this approach is compatible with large magnetic fields, a feature that enables testing of specific modes of proposed dark matter interactions. In addition, the disklike form factor of the absorber-sensor package that we envision as shown schematically in Fig. 1 might enable future adaptation of our concept using similar device integration concepts as developed, e.g., in SuperCDMS. In the remainder of this paper, we describe

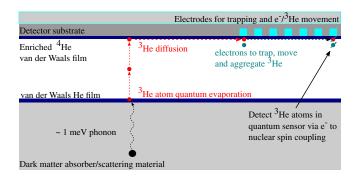


FIG. 1. Schematic of the DM detector concept. An interaction with DM in an ionic crystal generates ~ 1 meV phonons, which impinge on a surface covered with a van der Waals helium film. The phonon quantum evaporates a ³He atom from the surface of the film, which is then collected on the van der Waals film covering the detector structures. The ³He atoms diffuse until captured by an electron bound to the helium surface in a CCD-like structure. Periodically the collected ³He atoms are moved with the CCD to a readout device which operates via nuclear spin induced decoherence of an electron in a spin based quantum sensor.

each of the steps of our detector concept in detail. It spans a range of fields, including dark matter astrophysics, solid state physics for phonon propagation, quantum fluids for ³He evaporation, device physics for ³He trapping and transport, and quantum information and sensing for ³He detection. Here we show how it can be a viable complement to existing efforts for light DM detection with transition edge sensors, superconducting nanowire detectors, and CCDs.

II. HELIUM EVAPORATION VIA DM-PRODUCED PHONONS

Bulk superfluid He has been proposed for DM detection through the production of phonons and rotons [7,12]. By contrast, here we propose to use the helium as a means to detect the phonons produced in a solid target, and not as the target itself. This approach was also discussed in Ref. [12], though here we are specifically suggesting polar targets, such as NaI. Except for evidence that rotons do not efficiently evaporate ³He from bulk He [15], the remainder of this approach to detecting low-energy DM interactions could be utilized for a bulk He based detector. However, there are important advantages and complementary opportunities to interacting and generating phonons in crystals (notably reach to a broader range of dark matter theories and masses [8,9]), when evaporating ³He from these. We will focus on the case that a DM particle produces a single high-energy ($\gtrsim 10 \text{ meV}$) phonon by an interaction with an ion in a polar material target. The anticipated DM interaction rate is about 2/min in a 1 kg NaI crystal (detailed theoretical calculations can be found in Refs. [8,9]), with an expected background about 50 times lower. Practical details of detector crystal criteria, interaction rates, backgrounds, ³He detection approaches, and possible alternative adsorbates are discussed below.

Below about 80 mK a He surface is covered with ³He, both for bulk He and a van der Waals film. The athermal acoustic phonons resulting from the decay of the highenergy phonon, when interacting with the surface of the polar crystal coated with a thin helium film can lead to quantum evaporation. Heat pulse experiments with natural abundance He films on crystalline substrates have shown that about 5% of the detected atoms are directly evaporated by phonons from the heat pulse-the "phonoatomic" effect depicted in Fig. 1-while the remainder are evaporated by the overall temperature rise of the crystal [16,17]. However, these experiments have mostly used polished, rather than vacuum-cleaved surfaces. It is known that even wellpolished surfaces covered with helium lead to enhanced phonon thermalization [18] and inefficient transport of phonons across the interface into a film [19]. The efficiency of quantum evaporation from a van der Waals film of liquid helium on a freshly cleaved surface which has been protected from oxygen and humidity is not known. Boosting the evaporation efficiency may also be possible by depositing a thin film of Cs on a crystal and coating that with a monolayer of ³He, as suggested in Ref. [17], since ³He is bound to Cs by only about 2.4 K.

III. ³HELIUM TRAPPING

As shown in Fig. 1, the evaporated ³He atoms will be collected on an adjacent helium-covered surface. The helium in this collector film will be isotopically enriched to remove its ³He. Enrichment of ⁴He to less than 5 parts in 10^{13} (< 0.5 ppt) ³He has been demonstrated [20]. The enriched ⁴He film on this collector structure must be fully isolated from the ³He/⁴He mixture coating the DM target crystal. There are two well-established approaches to breaking a van der Waals film: a film burner as was employed in the HERON experiment [21] and a band of cold-evaporated Cs, since superfluid ⁴He does not wet Cs [22,23]. Here we expect that the Cs film will be preferable, since the film-burner could preferentially evaporate ³He atoms, which would appear as false events. After being captured onto this enriched ⁴He film, the ³He atoms diffuse across the film surface [24]. Our concept uses electrons held a few nanometers above the surface of the helium film by applied electric fields to localize ³He atoms in dimples under the electrons [25] and enable their transport to spin readout sensors for detection. It is essential that the ³He atoms be localized in the dimples for spin based ³He sensing, since if the ³He atoms are allowed to diffuse freely, motional narrowing causes them to have little effect on an electron's spin in a quantum sensor [26].

Electrons bound to the surface of superfluid helium have been studied for many years, [27] for both physics and devices. This physical system is the first in which a twodimensional (2D) Wigner crystal was observed, [28] and has demonstrated the highest mobility of any 2D electron system [29]. Devices used for transport experiments have often made use of "channel" technology [30,31]. Typically, an underlying metal layer is first deposited on a substrate and patterned to make gate electrodes, and this layer of electrodes is then covered with an insulator and a second metallic layer. This upper metal layer is patterned lithographically, and areas are removed to form the channels where the electrons will reside (see inset of Fig. 3). These channel devices are put into a vacuum tight cell, and cooled to below the λ point of helium. Helium is introduced into the cell, enough to form a small quantity of bulk helium but not enough to submerge the channel device. However, the helium covers the device through capillary action and fills the channels with superfluid helium. Electrons are then emitted onto the helium surface with the gate electrodes below the channels biased positive with respect to the top metal layer, causing electrons to accumulate on the liquid helium in the channels. Structures of this variety have been used to demonstrate a range of physics and devices, including CCDs with essentially perfect charge-transfer efficiency [32], reentrant melting of a quasi-1D Wigner crystal [33], and the isolation of individual [34] as well as pairs of electrons [35].

Trapping ³He in dimples under electrons bound to superfluid ⁴He has not been discussed previously, but will arise from ³He reducing the surface tension at mK temperatures [14]. The addition of a ³He atom will deepen the dimple, lowering the electron in the applied electric field, increasing its potential energy and trapping the ³He. The depth and shape of the dimple in the He surface will be determined by the equilibrium between electrostatic forces pulling the electron against the helium and capillary forces resisting the deformation.

To understand the temperature needed for stable trapping of the ³He we have performed numerical calculations of the helium dimple as shown in Fig. 2. These devices will use the channel technology described above. Figure 2 shows calculated electrostatic potentials for a 200 nm wide channel that is 110 nm deep. The lower metal electrode is biased to +20 V, and the upper metal is at ground (0 V), with the potential contours at 1 V steps. The change in the dimple with the addition of a ³He atom is too small to be seen in the figure, but the vertical electric field is calculated to be about 0.8×10^6 V/cm at the electron, so a very small change in dimple depth can produce a significant change in electrostatic energy. For the parameters of Fig. 2, the calculated energy change per ³He atom is about 27 K. A variety of channel geometries and

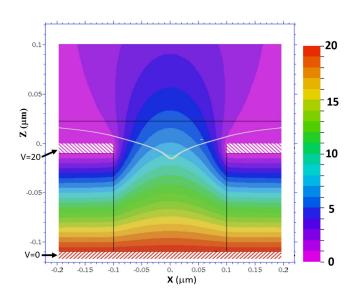


FIG. 2. Finite element calculation of the potential in a 0.2 μ m wide He-filled channel with metal gates biased as shown. The channel is assumed to extend in the Y direction (into the page) and electrons are placed with a periodicity of 0.2 μ m. The contours are at 1 V steps. The black horizontal line at a height of ~0.02 μ m shows the helium surface without the electron or electric fields. The white curve superimposed on the potential image is the calculated helium surface for an electron held in the channel with the applied voltages.

applied voltages have been modeled: higher voltages are required for narrow channels where capillary forces are stronger, while the helium surface becomes unstable if the channel becomes too wide. The calculations suggest that stable trapping of ³He is possible over at least a factor of 4 range in channel widths.

It is expected that this detector will be operated at \sim 35 mK, or colder, since the background pressure of ³He must be kept very low. At this temperature, if the trapping energy is 2 K, then the calculated density of free ³He atoms is \sim 10⁻¹² cm⁻² for every trapped ³He atom. Thus, trapping energies in the range of 1–2 K will be sufficient for localizing the ³He atoms.

Another consideration is the cross section for an electron to capture a ³He atom. Again, we have calculated the cross section numerically, here by introducing a change in the surface tension some distance from the electron to determine how the energy changes. Results for these same parameters (0.2 μ m channel with a 20 V bias) are shown in Fig. 3. Taking the criterion for capture that the energy falls below kT, then we have a capture distance of ~6 nm. In the orthogonal (Y) direction, calculations show that the capture distance is smaller, of order 1 nm. This can be readily understood, since the surface curvature is large in the X direction, where van der Waals forces require the helium surface to bend tightly around the edges of the channel, as seen in Fig. 2. In the Y direction the channel is long,

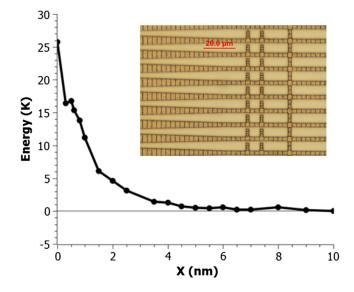


FIG. 3. Calculation of the binding energy of a ³He atom to an electron in a channel like that shown in Fig. 3, but with the ³He displaced by a distance (X) in the X direction (across the channel). Lines connecting dots are guides to the eye. The inset shows a CCD used for moving electrons along helium filled channels (the micrograph shows the metal layers). The main channels run horizontally and are similar, though wider and deeper, to the CCDs needed for the detectors. The underlying gate electrodes run vertically as seen at the bottoms of the channels (after Ref. [32]).

the dimple is more gradual, and the change in surface tension is only felt very close to the electron.

The ³He forms a Fermi gas on the ⁴He surface at low densities, and its motion is diffusive. Measurements of the spin diffusion at low coverage (~0.1 monolayers) in high surface area substrates finds a diffusivity of about $0.015 \text{ cm}^2 \text{ s}^{-1}$ at 40 mK [36]. These measurements had five monolayers of ⁴He below the ³He, which is sufficient to be a superfluid and avoid localizing the ³He atoms. For simplicity we can approximate the capture perimeter as an ellipse with a minor axis of 1 nm and major axis of 6 nm and determine an effective isotropic capture cross section [37]. Assuming electrons are spaced 0.2 µm in Y, and the channels are spaced 0.2 µm in X (so electrons are 0.4 µm apart), we use this electron (trap) density, the capture cross section, and thermal velocity to calculate a capture time of about 100 ns and a diffusion length of about 1 µm. Thus, a reasonable density of electrons can rapidly capture the ³He, and the location of where the ³He arrived can be determined with a few micron accuracy. Micronlevel resolution is unlikely to be necessary, and a lower density of channels and electrons should be adequate. For example, if the electrons are spaced 100 µm apart in both X and Y, the capture time for a 3 He atom becomes about 0.1 s. and the spatial resolution is about 0.4 mm. The frequency of readout operation cycles with ³He collection and quantum sensing will be adjusted to match event and background rates. Readout times of a few ms are slow, on the typical scale of CCDs and electronics, and nearly all the power will be dissipated in driver circuitry at higher temperature. Thus the heat load is expected to be small enough to allow operation at 35 mK, or colder.

IV. ³HELIUM TRANSPORT

An important feature of this detector concept is the ability to collect ³He atoms over a large area and bring them to one or a few optimized quantum sensors. As discussed earlier, phonons are created in the bulk of a detector crystal, and they rapidly disperse the energy throughout the volume of the material, making their direct detection challenging. Our concept uses CCDs for electrons bound to helium to transport ³He, similar to that shown in the inset of Fig. 3 [32]. In this device the channels run horizontally, while the gate electrodes can be seen running vertically under the channels. As the gate voltages are controlled to move the electrons, they will drag the ³He atoms along with them in moving dimples. The CCD device in the inset of Fig. 3 was made at a standard silicon processing facility, which can fabricate similar electrode structures over large areas. The assumed 1 kg NaI crystal forms a disk about 2 cm thick and 13 cm in diameter, and thus the collector must be similar in size. Silicon devices with areas in that range are practical. Either a single large device could be fabricated, or a number of smaller ones can be tiled together, as is done for large-area optical CCDs.

V. ³He DETECTION

After ³He atoms have been evaporated, captured, and collected with the CCD, it is necessary to detect single atoms. An electron's spin without a ³He trapped in the dimple below it is expected to have long phase coherence, since the spin-orbit interaction for an electron in the vacuum is particularly small [26]. However, if a ³He atom is trapped by an electron, then the nuclear spin will rapidly decohere the spin of an electron initially prepared in a superposition of up and down spin. This decoherence will happen in less than 1 ms, while the spin coherence of electrons bound to helium is thought to be at least seconds [26]. This is a quantum nondemolition process and can be repeated as long as the ³He remains trapped, allowing multiple interrogations to ensure reliable detection. Several approaches to detecting the spin of single electrons bound to helium are under active investigation, driven by quantum computing applications, and are discussed in the next section. Once the measurement is complete, any detected ³He atoms will be clocked with the CCD to a region with a large number of electrons tightly bound in circular (~200 nm diameter) "quantum dots." The ³He atoms will be trapped and gettered by these electrons. Residual ³He atoms present when the detector is initialized will similarly be collected and moved to the getter region with the CCD. A 100 nm thick enriched He layer over the collector area (for the 13 cm diameter target) can be expected to have $\sim 10^7$ ³He atoms, but it is quite straightforward to fabricate 10⁸ or more quantum dots in an area of $\sim 10 \text{ mm}^2$, and each dot can trap multiple ³He atoms.

VI. PRACTICAL CONSIDERATIONS

As noted earlier, detected count rates in the range of tens/hr are anticipated for a detector of of this variety utilizing a 1 kg NaI crystal. The density of low-energy DM particles is much larger than for massive ones. Thus we require low backgrounds, but not at the extreme level as in weakly interacting massive particle searches, where minimum detectable signals can be tens of counts/year in a 1 ton detector [38]. However, it will be assumed that the detector will be operated in a well-shielded underground low-background facility. The estimate of 2 DM scattering events/minute, with the probability of detecting an event being about 35%, compares favorably with the expected background of about one event per hour.

It should also be emphasized that this detector is not operating as a bolometer. Rather, nonequilibrium phonons with energies about 200 times the operating temperature are being sensed. Thus the detector is insensitive to small temperature fluctuations and background processes which can lead to minute heating; low levels of electrical noise on gate lines, for example.

An initial phonon with an energy of a few 10s of meV created in a dark matter scattering event in a high-quality crystal typically decays through a sequence of inelastic processes to acoustic phonons with frequencies of order 1 THz, where thermalization is slowed by the decreasing phonon density of states [39]. For concreteness we will consider light DM detection by a 1 kg NaI crystal with ³He quantum sensing of the resulting phonons. Other crystals may prove to be superior, but from a cursory look NaI satisfies several criteria: (1) it has low energy cut-off $(\sim 20 \text{ meV})$ for phonons generated by DM [8]; (2) it can be purified to have a low radioactive background; (3) neither Na nor I have multiple naturally occurring isotopes, thus eliminating isotopic scattering of the acoustic phonons; and (4) it can be cleaved, which will reduce the phonon thermalization at surfaces and may increase the yield of evaporated ³He atoms. From calculations of the cross section for DM interaction within a dark photon interaction model and a freeze-in model of the DM flux [9], one finds that the rate of DM events is about 2/minute at a DM mass of about 20 keV in 1 kg of NaI with a minimum energy cut-off of 20 meV. A 20 meV phonon in the NaI will decay to about 20 acoustic phonons with enough energy to quantum evaporate the 3 He. If we assume that the efficiency for an acoustic phonon to desorb a helium atom is $\sim 5\%$ [16,17], and the probability of that atom being a ³He is about 1/3 [15], there is thus about a 1/60 chance of a single acoustic phonon being detected through ³He evaporation. With each DM event producing ~20 acoustic phonons, we estimate about one ³He atom will be produced every 1.5 minutes. Improved preparation of the NaI surfaces or better ionic crystals may increase the ³He evaporation rate.

Backgrounds for this detector are expected to be similar to those seen by other DM experiments. Background sources for this class of quantum evaporation detectors have been identified and modeled as part of the HeRALD experiment, [7] including the layers of shielding required. Background excitation of the helium is suppressed by its large bandgap for electronic excitations. NaI has a smaller gap, of about 5.8 eV, but most of that analysis carries over to this case. The gap still protects against low-energy processes. If an event does excite an electron across the gap, a large number of phonons will be produced when the electron-hole pair recombine or are trapped, and again these events can be identified. There is evidence in some other ultra-sensitive DM detection devices that stresses built up in materials can slowly relax by emitting phonons. In the detectors discussed here no thin films are deposited on the target crystals. Such films can exhibit thermal expansion mismatch stresses. Low-stress mounting will still be important.

Our estimates of specific background processes are guided by analyses for other proposed low-mass DM

detectors [7,11]. The DAMA/LIBRA and KIMS experiments [40,41] have established that an important background source in NaI is residual ⁴⁰K. Large NaI crystals with no more than 20 ppb of potassium impurities [40] imply a decay rate of about 1.2/kg/hr. The decay of cosmogenic tritium will also contribute to the background. The CDMSlite experiment [42] has found a tritium production rate of \sim 75 atoms/kg/day in a Ge detector. Calculations of the tritium production for NaI find a rate of about 83 atoms/kg/day at sea level [43]. Assuming 60 days at sea level for detector crystal preparation before installation underground, the cosmogenic tritium will contribute about one decay every 30 hours. Each decay of ⁴⁰K and tritium will generate many phonons and thus many ³He atoms. If the detector is read out more frequently than these background events, the large signals can be used as a veto. These high energy events will also produce scintillation photons which can provide another avenue for vetoing them. Taken together it is anticipated that there will be about 50 detectable DM events between ⁴⁰K and tritium decays under the assumptions discussed above. Since these decays can be vetoed based on their deposited energy, they will contribute to detector dead time, but will not otherwise interfere with the DM signal.

The collector device (³He trapping, transport, and readout) can also introduce backgrounds from radioactive decay. It appears best to avoid Al metallization in the collector chip, since cosmogenic ²⁶Al could add a significant background. A copper process will avoid this issue. If the collector is made as a standard silicon device, it will also introduce signals from the decay of ³²Si, as has been seen in other DM experiments. The DAMIC experiment [44] has quantified the radioactivity of ³²Si, and finds it contributes about 80 decays/kg/day. A typical Si wafer with the 13 cm diameter discussed above weighs about 30 gm, and thus the ³²Si can be expected to cause about 3 events/day, but it is unlikely that much of the liberated energy can reach the target crystal. Again, these are high energy events which can be vetoed. The collector substrate is considerably lighter than the target crystal, and thus the tritium background from it is not expected to be a major contributor, with an event rate comparable to ³²Si, and a similar difficulty in the energy reaching the target crystal.

Compton scattering of MeV-scale photons will deposit high energies in the target crystal, which can be vetoed as described above, but Robinson has pointed out that coherent photon scattering can deposit much smaller energies, of the same order as DM events [45]. Data from the IGEX experiment shows that background events with energies above about 30 keV can be accounted for by radioactivity of experimental components, once sufficient shielding is in place in a low-background underground facility [46]. Using this data Robinson calculated an integrated coherent photon scattering rate of ~0.34 recoils/kg/day for recoil energies below 1 eV in Ge, assuming such a passive radiation shield and neglecting both coherence between atoms and phonon quantization in the crystal. Again, higher energy events can be vetoed. The iodine in the NaI crystal will dominate the coherent photon scattering, having a 3.8x larger atomic cross section than Ge through the relevant energy range [47]. Under similar assumptions we estimate that coherent photon scattering will produce ~0.6 recoils/kg/day. Being of similar energy as the DM events, it is not possible to veto these recoils, but their rate is over 3 orders of magnitude lower than the calculated DM rate (2/kg/min. in NaI). Coherent neutrino scattering will similarly generate recoils which cannot be vetoed based on deposited energy. However, for recoil energies below $\sim 1 \text{ eV}$, the coherent photon scattering rates as estimated by Robinson exceed the expected coherent neutrino rates [45,48]. Thus, while the photons and neutrinos will add a small offset to the DM signal, this background is expected to be smaller by several orders of magnitude.

To summarize the background considerations, they fall into two categories: (1) high-energy processes, like radioactive decay, which will generate large numbers of phonons and ³He and thus can be vetoed, but will still introduce a dead-time; and (2) low-energy processes which are essentially indistinguishable from DM events and thus will lead to a background. For a 1 kg NaI detector crystal the total rate of high-energy background processes (type 1) is estimated to be ≤ 1.5 events/hr (~1.2/hr from 40 K, ~ 0.03 /hr from cosmogenic tritium in the target, and < 0.12/hr each from ³²Si and cosmogenic tritium in the collector device). The total rate of low-energy background processes (type 2) is estimated to be $\sim 0.025/hr$, dominated by radioactivity in the experimental setup. Again, these estimated background rates are considerably lower than the expected DM event rate of $\sim 120/hr$, with a detection efficiency of about 35%.

Multiple approaches are being taken by different groups for measuring spins on helium. Detection of single nuclear spins in other systems has been accomplished with quantum sensors in recent years. For example, nitrogen-vacancy centers have been used to sense the presence of nearby ²⁹Si atoms [49]. However, it is not clear whether direct nuclear spin detection can be adapted to the situation of a ³He atom on ⁴He, since the direct sensing of nuclear spins has relied on extremely close and stable positioning of the nucleus and the sensor. Converting to an electron spin, with its much larger magnetic moment, appears easier as discussed earlier. The nuclear spin can be entangled with the electron's spin, which constitutes a decoherence process from the point of view of the electron. Detection of single electron spins has been demonstrated in a range of quantum sensor and qubit platforms, from quantum dots to color centers [50]. It has been shown that the electron motion can be coupled to a superconducting micro-resonator with a coupling constant of ~5 MHz [51]. However, these first experiments were limited by decoherence of the motional states, apparently due to vibrations exciting fluctuations in the helium surface. Recently, strong coupling of the electron motion to a superconducting micro-resonator while bound to solid neon has been demonstrated [52]. Isolating the helium from the vibrations is being investigated in several labs and a high degree of vibration isolation will be central to the integration of our detector concept. With the motion strongly coupled to the resonator, an inhomogeneous magnetic field can provide the spin interaction, as has been demonstrated for electrons in silicon quantum dots [53,54]. In an alternative configuration, one could utilize a pair of electrons initialized to a spin singlet in a nanofabricated quantum dot, separating the two electrons, trapping the ³He under one to shift its phase, and then bringing the electrons back together to determine whether they are still a singlet. Decoherence from (single) ³He atoms will drive them from the singlet to the triplet with m = 0. A third approach would be to use a color center, like a NV⁻ or SiV⁰ in diamond to sense the electron spin (much less demanding than sensing a nuclear spin) [55]. Direct ESR techniques may also be possible, where sensitivity to a single electron's spin has recently been demonstrated [56]. The signal could be enhanced by using one ³He atom to sequentially decohere multiple electrons, since the atom is preserved in the process (its spin need not be preserved).

Here we have concentrated on using the ³He nuclear spin for quantum sensing, but there may be other ways to utilize the unique signatures of ³He. For example, the CCDs could be arranged to transport all of the ³He atoms to one place, where they are ejected from the surface with a heat pulse. With the atoms all emerging in one place, an ionization process like that described by Maris *et al.* [12], but with isotope-selective (perhaps optical) excitation, could be employed. Alternatively, ultra-sensitive mass spectrometry or other sensing technique might be enabled with the localized He source.

Here we have concentrated on the evaporation of ³He from the surface of liquid ⁴He, since it has the lowest surface binding energy (~5 K). However, many other atomic and molecular species (as well as electrons) can be bound to a liquid He surface, and their evaporation may prove useful as phonon detectors. An isolated electron binds with an energy of $\sim 0.6 \text{ meV}$ [57], but a high electron density is necessary if the ejection of an electron is to have a high probability. However, large holding fields are then required to keep the electrons on the surface, and electron emission is limited by electron-electron interactions [58]. Alkali metals are predicted to bind to helium with energies of 10-20 K [59], and experimentally found to bind to the surface of He nanodroplets [60]. Being uncharged they do not require holding fields, but at high densities they form dimers and clusters. It has also been reported that other species, such as HD, can be desorbed from alkali halides with a single phonon [61]. Such species may be useful as detectors for particular energy ranges of proposed dark matter candidates and interactions. Being much more polarizable than He, it may also be possible to tune their desorption energy with an applied electric field.

VII. SUMMARY

In summary we have presented a new concept for detecting low energy (~meV) excitations, in particular those which might be generated in target materials through the interaction with low-mass dark matter. The approach begins with a DM interaction producing phonons in an ionic crystal, which cause the quantum evaporation of ³He from the surface. The ³He is then caught on an adjacent surface, where there is an isotopically enriched van der Waals ⁴He film covering a layer of metallic electrodes and etched microchannels holding electrons on the film. We calculate that the electrons on the helium can trap ³He atoms, and they will drag ³He atoms as they are clocked across the helium surface in a CCD, allowing ³He atoms to be collected for detection by quantum sensors. We suggest that the spin of ³He atoms can be coupled to electron spins for sensitive detection—to the level of a single ³He atom. Thus the difficult balance of efficient detection of rare lowenergy events occurring throughout a large volume is solved in our approach through the trapping, collection, and quantum sensing of the ³He atoms. Calculations of dark photon mediated interactions and estimates of the various background processes show that with a kg-sized ionic crystal a detected DM event rate of about 40/hr can be achievable, while high-energy radioactive decays and Compton events will be about 50 times less frequent. These high-energy events can be distinguished by the detector, and thus vetoed. Coherent photon and neutrino scattering will produce low-energy events, similar to DM, but their estimated rates are 2 orders of magnitude lower than the DM. Assemblies of dark matter sensors of this design could operate for long periods with periodic readout of accumulated ³He atoms.

All major aspects of this detector concept are based on established experimental results, or in the cases of single spin measurement and ³He trapping (also suggested for electron bubbles [62]), they are being actively pursued in the context of quantum computer development with electrons on liquid helium [26]. Experimental verification of spin measurement and ³He trapping will enable first-generation detectors and open the door to this path of quantum sensing of phonons for DM detection.

ACKNOWLEDGMENTS

This work was supported by Quantum Information Science Enabled Discovery (QuantISED) for High Energy Physics and by the Office of Science of the U.S. Department of Energy under Contract No. DE-AC02-05CH11231.

- [1] P. Cushman *et al.*, arXiv:1310.8327.
- [2] M. Battaglieri et al., arXiv:1707.04591.
- [3] L. Barak *et al.* (SENSEI Collaboration), Phys. Rev. Lett. 125, 171802 (2020).
- [4] I. Alkhatib *et al.* (SuperCDMS Collaboration), Phys. Rev. Lett. **127**, 061801 (2021).
- [5] S. Knapen, T. Lin, M. Pyle, and K. M. Zurek, Phys. Lett. B 785, 386 (2018).
- [6] K. Schutz and K. M. Zurek, Phys. Rev. Lett. 117, 121302 (2016).
- [7] S. A. Hertel, A. Biekert, J. Lin, V. Velan, and D. N. McKinsey, Phys. Rev. D 100, 092007 (2019).
- [8] T. Trickle, Z. Zhang, K. M. Zurek, K. Inzani, and S. M. Griffin, J. High Energy Phys. 3 (2020) 036.
- [9] S. M. Griffin, K. Inzani, T. Trickle, Z. Zhang, and K. M. Zurek, Phys. Rev. D 101, 055004 (2020).
- [10] A. Mitridate, T. Trickle, Z. Zhang, and K. M. Zurek, Phys. Rev. D 102, 095005 (2020).
- [11] Y. Hochberg, M. Pyle, Y. Zhao, and K. M. Zurek, J. High Energy Phys. 8 (2016) 057.
- [12] H. J. Maris, G. M. Seidel, and D. Stein, Phys. Rev. Lett. 119, 181303 (2017).
- [13] D. O. Edwards and W. F. Saam, Prog. Low Temp. Phys. 7A, 284 (1978).
- [14] A. F. Andreev, Zh. Eksp. Teor. Fiz. 50, 1415 (1966) [For English translation see Sov. Phys. JETP 50 (1966)].
- [15] J. P. Warren and C. D. Williams, Physica (Amsterdam) 284-288B, 160 (2000).
- [16] D. L. Goodstein, R. Maboudian, F. Scaramuzzi, M. Sinvani, and G. Vidali, Phys. Rev. Lett. 54, 2034 (1985).
- [17] T. More, J. S. Adams, S. R. Bandler, S. M. Brouër, R. E. Lanou, H. J. Maris, and G. M. Seidel, Phys. Rev. B 54, 534 (1996).
- [18] F. Türk, G. Ullrich, and H. Kinder, Ann. Phys. (N.Y.) 507, 165 (1995).
- [19] C. H. Anderson and E. S. Sabisky, Phys. Rev. Lett. 24, 1049 (1970).
- [20] P. Hendry and P. McClintock, Cryogenics 27, 131 (1987).
- [21] R. Torii, S. R. Bandler, T. More, F. S. Porter, R. E. Lanou, H. J. Maris, and G. M. Seidel, Rev. Sci. Instrum. 63, 230 (1992).
- [22] P.J. Nacher and J. Dupont-Roc, Phys. Rev. Lett. 67, 2966 (1991).
- [23] P. Taborek and J. E. Rutledge, Phys. Rev. Lett. 68, 2184 (1992).
- [24] P. A. Sheldon and R. B. Hallock, Phys. Rev. Lett. 77, 2973 (1996).
- [25] R. Williams and R. Crandall, Phys. Lett. 36A, 35 (1971).
- [26] S. A. Lyon, Phys. Rev. A 74, 052338 (2006).
- [27] Two-Dimensional Electron Systems on Helium and other Cryogenic Substrates, edited by E. Y. Andrei (Springer, Dordrecht, 1997), Vol. 19.
- [28] C. C. Grimes and G. Adams, Phys. Rev. Lett. 42, 795 (1979).
- [29] K. Shirahama, S. Ito, H. Suto, and K. Kono, J. Low Temp. Phys. 101, 439 (1995).
- [30] D. Marty, J. Phys. C 19, 6097 (1986).
- [31] R. van Haren, G. Acres, P. Fozooni, A. Kristensen, M. Lea, P. Richardson, A. Valkering, and R. van der Heijden, Physica (Amsterdam) 249-251B, 656 (1998).

- [32] F. R. Bradbury, M. Takita, T. M. Gurrieri, K. J. Wilkel, K. Eng, M. S. Carroll, and S. A. Lyon, Phys. Rev. Lett. 107, 266803 (2011).
- [33] D. G. Rees, H. Ikegami, and K. Kono, J. Phys. Soc. Jpn. 82, 124602 (2013).
- [34] G. Papageorgiou, P. Glasson, K. Harrabi, V. Antonov, E. Collin, P. Fozooni, P.G. Frayne, M.J. Lea, D.G. Rees, and Y. Mukharsky, Appl. Phys. Lett. 86, 153106 (2005).
- [35] M. Takita and S. A. Lyon, J. Phys. Conf. Ser. 568, 052034 (2014).
- [36] P. A. Sheldon and R. B. Hallock, Phys. Rev. Lett. 85, 1468 (2000).
- [37] C. R. Crowell, Appl. Phys. 9, 79 (1976).
- [38] E. Aprile et al., Phys. Rev. D 102, 072004 (2020).
- [39] J. P. Wolfe, Imaging Phonons: Acoustic Wave Propagation in Solids (Cambridge University Press, Cambridge, England, 1998), 10.1017/CBO9780511665424.
- [40] R. Bernabei, P. Belli, A. Bussolotti, F. Cappella, R. Cerulli, C. Dai, A. d'Angelo, H. He, A. Incicchitti, H. Kuang, J. Ma, A. Mattei, F. Montecchia, F. Nozzoli, D. Prosperi, X. Sheng, and Z. Ye, Nucl. Instrum. Methods Phys. Res., Sect. A 592, 297 (2008).
- [41] G. Adhikari, P. Adhikari, C. Ha, E. Jeona, N. Kim, Y. Kim, S. Kong, H. Lee, S. Oh, J. Park, and K. Park, Eur. Phys. J. C 77, 437 (2017).
- [42] R. Agnese et al., Astropart. Phys. 104, 1 (2019).
- [43] J. Amare, J. Castel, S. Cebrian, I. Coarasa, C. Cuesta, T. Dafni, J. Galan, E. Garcıa, J. Garza, F. Iguaz, I. Irastorza, G. Luzon, M. Martinez, H. Mirallas, M. Olivan, Y. Ortigoza, A. Ortiz de Solorzano, J. Puimedon, E. Ruiz-Choliz, M. Sarsa, J. Villar, and P. Villar, Astropart. Phys. 97, 96 (2018).
- [44] A. Aguilar-Arevalo et al., J. Instrum. 10, P08014 (2015).
- [45] A. E. Robinson, Phys. Rev. D 95, 021301 (2017).
- [46] S. Cebrián, J. Amaré, B. Beltrán, J. Carmona, E. García, I. Irastorza, G. Luzón, M. Martínez, A. Morales, J. Morales, A. O. de Solórzano, C. Pobes, J. Puimedón, J. Ruz, M. Sarsa, L. Torres, and J. Villar, Nucl. Phys. B, Proc. Suppl. 138, 147 (2005). Proceedings of the Eighth International Workshop on Topics in Astroparticle and Undeground Physics.
- [47] B. K. Chatterjee and S. C. Roy, J. Phys. Chem. Ref. Data 27, 1011 (1998).
- [48] J. Billard, E. Figueroa-Feliciano, and L. Strigari, Phys. Rev. D 89, 023524 (2014).
- [49] C. Müller, X. Kong1, J.-M. Cai, K. Melentijević, A. Stacey, M. Markham, D. Twitchen, J. Isoya, S. Pezzagna, J. Meijer, J. Du, M. Plenio, B. Naydenov, L. McGuinness, and F. Jelezko, Nat. Commun. 5, 4703 (2014).
- [50] C. L. Degen, F. Reinhard, and P. Cappellaro, Rev. Mod. Phys. 89, 035002 (2017).
- [51] G. Koolstra, G. Yang, and D. I. Schuster, Nat. Commun. 10, 5323 (2019).
- [52] X. Zhou, G. Koolstra, X. Zhang, G. Yang, X. Han, B. Dizdar, D. Ralu, W. Guo, K. W. Murch, D. I. Schuster, and D. Jin, Nature (London) 605, 46 (2022).
- [53] X. Mi, M. Benito, S. Putz, D. M. Zajac, J. M. Taylor, G. Burkard, and J. R. Petta, Nature (London) 555, 599 (2018).

- [54] N. Samkharadze, G. Zheng, N. Kalhor, D. Brousse, A. Sammak, U. C. Mendes, A. Blais, G. Scappucci, and L. M. K. Vandersypen, Science 359, 1123 (2018).
- [55] J. M. Taylor, P. Cappellaro, L. Childress, L. Jiang, D. Budker, P. R. Hemmer, A. Yacoby, R. Walsworth, and M. D. Lukin, Nat. Phys. 4, 810 (2008).
- [56] Z. Wang, L. Balembois, M. Rančić, E. Billaud, M. L. Dantec, A. Ferrier, P. Goldner, S. Bertaina, T. Chanelière, D. Estève, D. Vion, P. Bertet, and E. Flurin, Nature (London) 619, 276 (2023).
- [57] M. W. Cole, Rev. Mod. Phys. 46, 451 (1974).
- [58] K. Kono, K. Kajita, S.-i. Kobayashi, and W. Sasaki, J. Low Temp. Phys. 46, 195 (1982).
- [59] F. Ancilotto, E. Cheng, M. W. Cole, and F. Toigo, Z. Phys. B 98, 323 (1995).
- [60] F. Stienkemeier, J. Higgins, C. Callegari, S. I. Kanorsky, W. E. Ernst, and G. Scoles, Z. Phys. D 38, 253 (1996).
- [61] P. M. Ferm, S. R. Kurtz, K. A. Pearlstine, and G. M. McClelland, Phys. Rev. Lett. 58, 2602 (1987).
- [62] A. J. Dahm, Phys. Rev. 180, 259 (1969).