

Magnetic-field-driven topological phase transition in the holographic Weyl semimetal

R. C. L. Bruni^{1,*}, Luiz F. Ferreira^{2,†} and Diego M. Rodrigues^{3,‡}

¹*The University of Queensland, School of Mathematics and Physics, Queensland 4072, Australia*

²*Instituto de Física y Astronomía, Universidad de Valparaíso, A. Gran Bretaña 1111, Valparaíso, Chile*

³*Instituto de Física Teórica, Universidade Estadual Paulista, R. Dr. Bento T. Ferraz 271, Bl. II, São Paulo 01140-070, São Paulo, Brazil*



(Received 29 March 2023; accepted 21 August 2023; published 14 September 2023; corrected 6 May 2025)

We study the magnetic field effects on the quantum critical point (QCP) in the holographic Weyl semimetal model. We show that it increases quadratically with the magnetic field for weak field and linearly with the magnetic field for strong field. Our findings are compatible with previous results in the literature from other approaches.

DOI: [10.1103/PhysRevD.108.066010](https://doi.org/10.1103/PhysRevD.108.066010)

I. INTRODUCTION

The investigation on the *topological states of matter*, *i.e.*, materials that do not fit in the usual Ginzburg-Landau criteria being classified by topology rather than symmetry (or breaking of it) has become increasingly stronger in condensed matter physics. In such materials, the momentum-space topology of the Fermionic ground states, most commonly manifested in their electronic band structure, plays a fundamental role [1,2]. The physical realization of these states includes topological insulators (TI) [3–5], topological superconductors (TSC) [6–8], and Weyl semimetals (WSM) [9–14].

The low energy effective field theory description of WSM's is a linear gapless effective Hamiltonian resembling a relativistic *Weyl fermion* from particle physics literature [15,16], whose geometric *locus* in momentum space defines a cone centered at a given value of momenta and energy, the *Weyl cone* [17–30]. These cones must always occur in an even number throughout the \mathbf{k} space, as a result of the *Fermion doubling theorem* [31,32]. This last feature, together with their associated topological charge [33,34], endows the Weyl cones with a topological stability against perturbations that preserves the original symmetries of the system and leads to a series of interesting

physical consequences, *e.g.*, Fermi-arcs [35–40] and the chiral magnetic effect [41–43].

The development of the AdS/CFT correspondence or, more generally known as the gauge/gravity duality, over the past years has become an important tool for studying strongly coupled systems. Originally, the AdS/CFT correspondence states a duality between a gravity theory in a $(d + 1)$ -dimensional anti-de Sitter (AdS) space and $\mathcal{N} = 4$ super-Yang-Mills theory, living on d -dimensional boundary of AdS space. Namely, the physics of the strongly coupled theory on d -dimensional boundary of AdS space is encoded in the weakly coupled classical gravity living on a $(d + 1)$ -dimensional AdS space. Nowadays, the extension of this duality has allowed to model a variety of condensed matter systems [44–46]. In particular, a holographic Weyl-semimetal model and its quantum phase transition to a trivial semimetal has been constructed in [47,48] by including deformations in the gravitational theory which mimic the essential features of a Weyl semimetal effective field theory.

At zero temperature, the holographic Weyl semimetal [48] exhibits a topological quantum phase transition from Weyl semimetal to a topological trivial phase, where the order parameter is the anomalous Hall conductivity. The quantum phase transition occurs at a critical value of the effective coupling $(M/b)_{\text{crit}}$, where M and b are parameters included, via proper boundary conditions, in the holographic model and which are dual to the mass deformation and the axial gauge field in the Weyl semimetal effective field theory. Thus, due to the inclusion of essential features, this approach provides a great tool to investigate various properties of Weyl semimetal physics, such as surface state [49], odd viscosity [50], conductivity [51], axial Hall conductivity [52], topological invariants [53], momentum relaxation [54,55], nodal line semimetal [56–58], and

*bruni.r.c.l@gmail.com

†lffaulhaber@gmail.com

‡diegomhrod@gmail.com

Published by the American Physical Society under the terms of the [Creative Commons Attribution 4.0 International](https://creativecommons.org/licenses/by/4.0/) license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP³.

Weyl- Z_2 semimetal [59]. Besides, see [60] for a recent review. Further applications of holography in Weyl semimetal physics can be found, for instance, in [61–70].

Motivated by recent experimental progress on the role played by the magnetic field as a driver of topological phase transitions for instance in *Fe*-doped heterostructures [71] and in a trivial semimetal [72], we extend the previous framework of the holographic Weyl semimetal model [47] and its quantum phase transition to a trivial semimetal [48] by including a finite magnetic field in order to study the behavior of the quantum critical point as a function of the magnetic field. The paper is organized as follows: in Sec. II, we revisit the Landau level picture in the effective field theory description of Weyl semimetals. In Sec. III, we present the holographic model for the Weyl semimetal in a magnetic field and our numerical results. Finally, in Sec. IV, we present our conclusions and discussions.

II. REVISITING THE EFFECTIVE DESCRIPTION OF WEYL SEMIMETALS IN A MAGNETIC FIELD: LANDAU LEVELS

In this section, we revisit the low-energy effective description of a Weyl semimetal in the presence of a uniform magnetic field and compute the energy spectrum associated with the Landau levels.

Weyl semimetals are exotic materials whose band structure shows points around which the low-energy excitations are a pair of left- and right-handed Weyl fermions, which are chiral massless relativistic particles, lying exactly at the Fermi level (for a short review, see [73]). Their low-energy effective field theory description can be captured by a Lorentz-breaking Dirac Lagrangian [74] coupled to an external electromagnetic potential A_μ given by

$$\mathcal{L} = \bar{\Psi}(\gamma^\mu(i\partial_\mu - eA_\mu) - \gamma^5\boldsymbol{\gamma}\cdot\mathbf{b} - M)\Psi, \quad (1)$$

where $\gamma^\mu = (\gamma^0, \boldsymbol{\gamma})$ are the gamma matrices which satisfy the Clifford algebra $\{\gamma^\mu, \gamma^\nu\} = 2\eta^{\mu\nu}\mathbb{1}_4$, where $\eta_{\mu\nu}$ is the Minkowski metric whose signature we take here to be $(+, -, -, -)$. γ^5 is the chiral matrix defined as $\gamma^5 = i\gamma^0\gamma^1\gamma^2\gamma^3$, and the electromagnetic potential is generally given by $A_\mu = (\phi, -\mathbf{A})$, where ϕ is the electric potential and \mathbf{A} is the vector potential. The vector \mathbf{b} is an axial vector responsible for breaking Lorentz symmetry, and which we will take it in the \hat{z} direction, i.e., $\mathbf{b} = b\hat{z}$, and will be the separation in momentum between the Weyl nodes. Finally, M is the mass deformation.

The Dirac equation coming from the variation of the above Lagrangian with respect to $\bar{\Psi}$ reads

$$(\gamma^\mu(i\partial_\mu - eA_\mu) - \gamma^5\boldsymbol{\gamma}\cdot\mathbf{b} - M)\Psi = 0. \quad (2)$$

Since in this section we are only interested in studying the system in a constant and uniform magnetic field in the \hat{z} direction, we take the electromagnetic potential to be $A_\mu = (0, -\mathbf{A})$, where the vector \mathbf{A} will be chosen in the symmetric gauge, which is given by $\mathbf{A} = \left(-\frac{By}{2}, \frac{Bx}{2}, 0\right)$, where B is the magnetic field. One can clearly see that this choice for the vector potential yields $\nabla \times \mathbf{A} = B\hat{z}$. Thus, the Dirac equation (2) takes the form,

$$(\gamma^\mu(i\partial_\mu) + e\boldsymbol{\gamma}\cdot\mathbf{A} - \gamma^5\boldsymbol{\gamma}\cdot\mathbf{b} - M)\Psi = 0. \quad (3)$$

Now, going to momentum space [using the correspondence principle $i\partial_\mu = p_\mu = (E, -\mathbf{p})$], we have

$$(\gamma^0 E + \boldsymbol{\gamma}\cdot(-\mathbf{p} + e\mathbf{A}) - \gamma^5\boldsymbol{\gamma}\cdot\mathbf{b} - M)\Psi = 0, \quad (4)$$

which can be written as

$$(\boldsymbol{\gamma}\cdot(\mathbf{p} - e\mathbf{A}) + \gamma^5\boldsymbol{\gamma}\cdot\mathbf{b} + M)\Psi = \gamma^0 E\Psi. \quad (5)$$

Multiplying both sides from the left by γ^0 and, using that $(\gamma^0)^2 = 1$, we can turn the Dirac equation (2) into a eigenvalue problem,

$$\hat{H}\Psi = E\Psi, \quad (6)$$

where the Hamiltonian operator \hat{H} is given by

$$\hat{H} = (\gamma^0\boldsymbol{\gamma}\cdot(\mathbf{p} - e\mathbf{A}) + b\gamma^0\gamma^5\gamma^3 + \gamma^0 M), \quad (7)$$

where we used that $\mathbf{b} = b\hat{z}$. For our particular case, it will be convenient to work in the Weyl representation for the gamma matrices, which is given by

$$\gamma^0 = \begin{pmatrix} 0 & \mathbb{1}_2 \\ \mathbb{1}_2 & 0 \end{pmatrix}, \quad \gamma^i = \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix}, \quad \gamma^5 = \begin{pmatrix} -\mathbb{1}_2 & 0 \\ 0 & \mathbb{1}_2 \end{pmatrix}, \quad (8)$$

where $\mathbb{1}_2$ is the 2×2 identity matrix, and σ_i are the Pauli matrices, given by

$$\sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma_y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}. \quad (9)$$

Using the gamma matrices defined above, the Hamiltonian operator takes the form,

$$\hat{H} = \begin{pmatrix} \boldsymbol{\sigma}\cdot(\mathbf{p} - e\mathbf{A}) + b\sigma_z & M\mathbb{1}_2 \\ M\mathbb{1}_2 & -\boldsymbol{\sigma}\cdot(\mathbf{p} - e\mathbf{A}) + b\sigma_z \end{pmatrix}. \quad (10)$$

Notice that the system have a well-defined *helicity* given by the eigenvalues of the operator $\hat{h} = \hat{\sigma} \cdot (\hat{\mathbf{p}} - e\hat{\mathbf{A}})^1$ —with $\hat{\mathbf{p}}$ being the momenta shifted by $\pm \mathbf{b} = b\hat{z}$, which coincides with the chirality in the massless limit, even though it is well known that in the absence of external electromagnetic field, it is also possible to define an effective chirality operator if $b > M$ [28].

Now, in order to proceed with the computation, let us take a look on the term $\hat{\sigma} \cdot (\hat{\mathbf{p}} - e\hat{\mathbf{A}}) + b\hat{\sigma}_z$. In components and, using the symmetric gauge for the vector potential $\hat{\mathbf{A}}$, we have

$$\hat{\sigma}_x \hat{\Pi}_x + \hat{\sigma}_y \hat{\Pi}_y + \hat{\sigma}_z (\hat{p}_z + b\mathbb{1}), \quad (11)$$

where we introduced the operators $\hat{\Pi}_x$ and $\hat{\Pi}_y$, defined by

$$\hat{\Pi}_x := \hat{p}_x + \frac{eB\hat{y}}{2}; \quad \hat{\Pi}_y := \hat{p}_y - \frac{eB\hat{x}}{2}, \quad (12)$$

which satisfy the canonical commutation relation $[\hat{\Pi}_x, \hat{\Pi}_y] = ieB$. Now, in terms of the “ Π ’s” operators above defined, we introduce the annihilation and creation operators, \hat{a} and \hat{a}^\dagger [75–77],

$$\hat{a} := \sqrt{\frac{1}{2eB}} (\hat{\Pi}_x + i\hat{\Pi}_y); \quad \hat{a}^\dagger := \sqrt{\frac{1}{2eB}} (\hat{\Pi}_x - i\hat{\Pi}_y), \quad (13)$$

which satisfy the canonical commutation relation $[\hat{a}, \hat{a}^\dagger] = 1$. Thus, Eq. (11) can be written as

$$\sqrt{2eB} (\hat{S}_- \hat{a} + \hat{S}_+ \hat{a}^\dagger) + 2\hat{S}_z (b + \hat{p}_z), \quad (14)$$

where $\hat{S}_\pm = \frac{1}{2}\hat{\sigma}_\pm := \frac{1}{2}(\hat{\sigma}_x \pm i\hat{\sigma}_y)$ and $\hat{S}_z := \frac{1}{2}\hat{\sigma}_z$ are the spin operators. Therefore, in terms of these definitions, the Hamiltonian operator, given by Eq. (10), becomes

$$\hat{H} = \begin{pmatrix} \sqrt{2eB} (\hat{S}_- \hat{a} + \hat{S}_+ \hat{a}^\dagger) + 2\hat{S}_z \hat{p}_z + 2b\hat{S}_z & M\hat{\mathbb{1}}_2 \\ M\hat{\mathbb{1}}_2 & -(\sqrt{2eB} (\hat{S}_- \hat{a} + \hat{S}_+ \hat{a}^\dagger) + 2\hat{S}_z \hat{p}_z) + 2b\hat{S}_z \end{pmatrix}. \quad (15)$$

Now, we have to define the ground state, which is going to correspond to the lowest Landau level (LLL). It is characterized by the tensor product,²

$$|0, k_z\rangle \otimes \left| \frac{1}{2}, +\frac{1}{2} \right\rangle := \left| 0, k_z, +\frac{1}{2} \right\rangle,$$

where $|\frac{1}{2}, +\frac{1}{2}\rangle$ is the eigenstate of the spin-1/2 operator S_z with eigenvalue $+\frac{1}{2}$.³ Note that the term $(\sqrt{2eB}(\hat{S}_- \hat{a} + \hat{S}_+ \hat{a}^\dagger) + 2\hat{S}_z \hat{p}_z + 2b\hat{S}_z)$ in (15) is annihilated by the

ground state $|0, k_z, +\frac{1}{2}\rangle$ since $\hat{S}_- \hat{a} |0, k_z, +\frac{1}{2}\rangle = 0$ and $\hat{S}_+ \hat{a}^\dagger |0, k_z, +\frac{1}{2}\rangle = 0$. Thus, the matrix elements of (15), $\langle +\frac{1}{2}, k_z, 0 | \hat{H} | 0, k_z, +\frac{1}{2} \rangle$, are given by

$$H = \begin{pmatrix} k_z + b & M \\ M & -(k_z - b) \end{pmatrix}, \quad (16)$$

where we used that $\hat{p}_z |0, k_z, +\frac{1}{2}\rangle = k_z |0, k_z, +\frac{1}{2}\rangle$ and $\hat{S}_z |0, k_z, +\frac{1}{2}\rangle = +\frac{1}{2} |0, k_z, +\frac{1}{2}\rangle$. Thus, diagonalizing this matrix, we find that the ground state is linearly dispersing (see Fig. 1) and given by

$$E = b \pm \sqrt{k_z^2 + M^2}; \quad n = 0. \quad (17)$$

¹Usually the helicity operator is defined to be $\hat{h} = \hat{\sigma} \cdot \hat{\mathbf{p}}$, which is generalized to this definition according to the usual minimal coupling prescription.

²Since the spin does not depend on the spatial degrees of freedom, the operator \hat{S} commutes with all the spatial operators. In particular, it commutes with the momentum operator \hat{p} , i.e., $[S_i, p_j] = 0$. Therefore, it acts only on the spin part $|\frac{1}{2}, +\frac{1}{2}\rangle$ and leaves the rest unchanged.

³In this case, the spin up with $m_s = \frac{1}{2}$ is the ground state because the spinning motion gives rise to a magnetic dipole moment given by $\vec{\mu}_S = \frac{e}{2mc} \vec{S}$ which, when interacting with the external magnetic field \vec{B} , in turn generates a coupling term $-\vec{\mu}_S \cdot \vec{B}$, which is minimized when the spin is aligned with the magnetic field.

Note that the energy dispersion for the LLL (17) is independent of the magnetic field B and purely chiral, i.e., right (left) movers for the negative (positive) chirality node [78].

For the excited Landau levels, $n \geq 1$, the most general eigenstate is characterized by $|n, k_z, \pm\frac{1}{2}\rangle$. Therefore, in order to construct the matrix elements, it will be convenient to square the Hamiltonian operator (15) and writing them as $\langle \pm\frac{1}{2}, k_z, n | \hat{H}^2 | n, k_z, \pm\frac{1}{2} \rangle$. By doing this computation, we find that the Hamiltonian operator (15) squared is given by

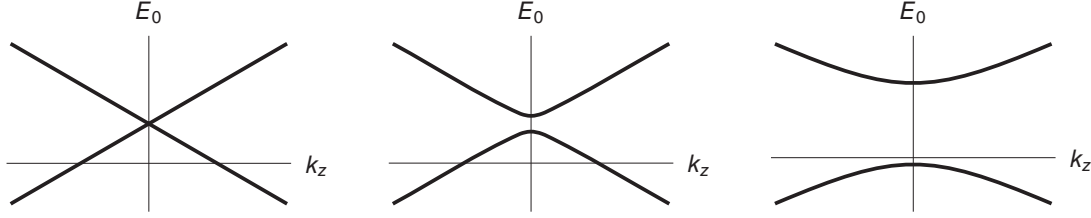


FIG. 1. Energy dispersion for the lowest Landau level (LLL), E_0 , in three different scenarios: (i) $M = 0$; (ii) $\frac{M}{b} < 1$ and (iii) $\frac{M}{b} > 1$. Note that between $\frac{M}{b} < 1$ and $\frac{M}{b} > 1$, a quantum phase transition takes place and, for $\frac{M}{b} > 1$, the system is gapped and become a trivial semimetal, i.e., an insulator.

$$\hat{H}^2 = \begin{pmatrix} 2eB\left(\hat{a}^\dagger\hat{a} + \frac{1}{2} - \hat{S}_z\right) + (\hat{p}_z + b)^2 + M^2 & 4Mb\hat{S}_z \\ 4Mb\hat{S}_z & 2eB\left(\hat{a}^\dagger\hat{a} + \frac{1}{2} - \hat{S}_z\right) + (\hat{p}_z - b)^2 + M^2 \end{pmatrix}. \quad (18)$$

Now, acting with this operator on the spin-up sector, with matrix elements given by $\langle +\frac{1}{2}, k_z, n | \hat{H}^2 | n, k_z, +\frac{1}{2} \rangle$, we find

$$H_{\uparrow}^2 = \begin{pmatrix} 2eBn + (k_z + b)^2 + M^2 & 2Mb \\ 2Mb & 2eBn + (k_z - b)^2 + M^2 \end{pmatrix}. \quad (19)$$

Analogously, for the spin-down sector, with matrix elements given by $\langle -\frac{1}{2}, k_z, n | \hat{H}^2 | n, k_z, -\frac{1}{2} \rangle$, we obtain

$$H_{\downarrow}^2 = \begin{pmatrix} 2eB(n+1) + (k_z + b)^2 + M^2 & -2Mb \\ -2Mb & 2eB(n+1) + (k_z - b)^2 + M^2 \end{pmatrix}. \quad (20)$$

Therefore, the excited Landau levels ($n \geq 1$) for both spin up and down sectors, respectively, are given by

$$E_n^{\uparrow} = \pm \sqrt{E_0^2 + \beta n}; \quad n \geq 1, \quad (21)$$

$$E_n^{\downarrow} = \pm \sqrt{(E_n^{\uparrow})^2 + \beta}, \quad (22)$$

where $E_0 = b \pm \sqrt{k_z^2 + M^2}$ is the LLL, and $\beta = 2eB$. From the above equations, one can see that there is an energy gap ΔE_n^2 between the up and down spins, given by

$$\Delta E_n^2 := (E_n^{\downarrow})^2 - (E_n^{\uparrow})^2 = \beta = 2eB. \quad (23)$$

Finally, in Fig. 2, we display the energy dispersion corresponding to the first excited Landau level for both spins up and down.

Figures 1 and 2 should be compared with the energy dispersion when the magnetic field is absent. In this case, the energy dispersion is given by

$$E|_{B=0} = \pm \left(b \pm \sqrt{k_z^2 + M^2} \right). \quad (24)$$

This energy dispersion is displayed in Fig. 3. One can see from Fig. 1 that, even though the presence of an external magnetic field “gaps out” the band touching points present in the $B = 0$ case (Fig. 3), the LLL still supports a Weyl semimetal phase, once it still hosts a linearly dispersion branch with a well-defined chirality close to Fermi level (set to zero for convenience), which is in fact the most general condition for the establishment of the phase, since it automatically yields a nonzero flux of the corresponding Berry curvature [13,79]. The maintenance of Weyl points under weak, intermediate, and strong magnetic fields is a far general phenomenon with roots in the topological nature of Weyl points, *i.e.*, the fact that they are robust against perturbations that preserve the original symmetries of the system. To be more precise, in the absence of an external magnetic field perpendicular to the separation of Weyl points, the momenta along the latter is still a good quantum number. Therefore, the many-body wave function of the electron occupying the state of a Weyl node has a definitive momenta eigenvalue [80]. The situation changes dramatically if one includes a magnetic field perpendicular to the Weyl nodes. Now, the translation symmetry along the Weyl points is broken, the corresponding momenta is not a good

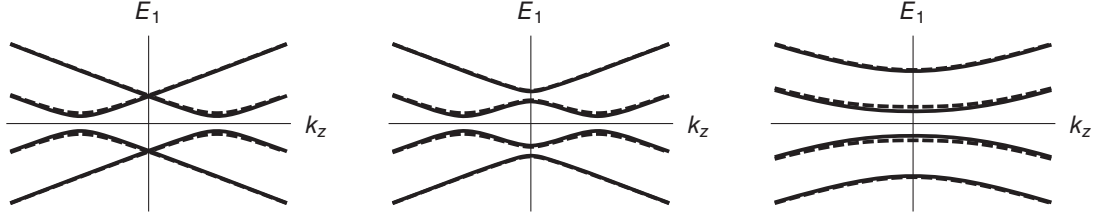


FIG. 2. Energy dispersion for the first excited Landau level, E_1 , for both spin up (black curve) and down (dashed black curve) in three different scenarios: (i) $M = 0$, (ii) $\frac{M}{b} < 1$, and (iii) $\frac{M}{b} > 1$. Note that for the excited Landau levels ($n \geq 1$) the Weyl semimetal state is destroyed. Also, there is an energy gap ΔE_n^2 between the up and down spins, which is of order $\Delta E_n^2 = \beta = 2eB$.

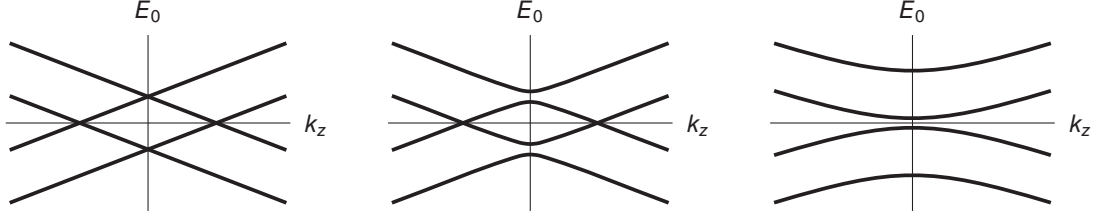


FIG. 3. Energy dispersion for zero magnetic field ($B = 0$) in three different scenarios: (left panel) $M = 0$; (middle panel) $\frac{M}{b} < 1$, and (right panel) $\frac{M}{b} > 1$. Note that there is a quantum phase transition between $\frac{M}{b} < 1$ and $\frac{M}{b} > 1$. For $\frac{M}{b} > 1$, the system is gapped and become a trivial semimetal, i.e., an insulator.

quantum number anymore, and the many body wave function is broadened into a finite peak around each node, allowing for the corresponding collective counterpropagating modes to hybridize and establish an electronic gap.⁴ If the Weyl nodes separation is smaller than the inverse magnetic length, $(k_{\text{mag}})^{-1} = (eB)^{-1/2}$, the aforementioned electronic gap is not negligible, which effectively “gaps out” the system leading to the magnetic tunneling effect [81].

III. HOLOGRAPHIC SEMIMETAL MODEL IN A UNIFORM MAGNETIC FIELD

A. Holographic setup

The holographic model for the Weyl semimetal we consider here is an extension based on [47,48] (for a review, see, e.g., [60]), in which we include a background magnetic field.⁵ The bulk action is given by

$$S = \int d^5x \sqrt{-g} \left[\frac{1}{2\kappa^2} \left(R + \frac{12}{L^2} \right) - \frac{1}{4} (F^2 + \mathcal{F}^2) + \frac{\alpha}{3} \epsilon^{abcde} A_a (\mathcal{F}_{bc} \mathcal{F}_{de} + 3F_{bc} F_{de}) - (D\Phi)^2 - V(\Phi) \right], \quad (25)$$

⁴One can borrow a classical analogue: when one introduces the magnetic field, each electron now describes a cyclotron motion in the plane perpendicular to the applied magnetic field. This motion delocalizes the electrons and their corresponding wave functions, allowing for their superposition and a consequent energy difference, i.e., an electronic gap.

⁵Since this holographic model is purely bosonic, i.e., does not contain fermions (in particular, Weyl fermions), it is not clear how this model captures the physics of a strongly coupled Weyl semimetal. However, this model does describe very well the topological quantum phase transition between a nontrivial semimetal phase with nonvanishing anomalous Hall conductivity to a trivial insulating phase with vanishing anomalous Hall conductivity [47,48,60]. In order to make a connection which is closer to the physics of a strongly interacting Weyl semimetal in terms of the spectral functions and energy dispersion relations, one should consider a different holographic approach, like the one adopted in [61].

where, κ^2 is the Newton’s constant, L is the AdS radius, $F = dV$ and $\mathcal{F} = dA$ are the vector and axial gauge field strengths, α is the Chern-Simons coupling constant, $D_a \Phi = (\nabla_a - iqA_a) \Phi$ is the covariant derivative, and Φ is a complex scalar field charged under the $U(1)$ axial gauge symmetry. The scalar potential $V(\Phi)$ is given by

$$V(|\Phi|) = m^2 |\Phi|^2 + \frac{\lambda}{2} |\Phi|^4, \quad (26)$$

where, according to the holographic dictionary, $(mL)^2 = \Delta(\Delta - 4)$, and where Δ is the scaling dimension of the operator dual to Φ and is chosen such that it has scaling dimension $\Delta = 3$, which gives $m^2 = -3$.⁶

⁶This matches the dimension of the boundary field theory operator $\langle \bar{\Psi} \Psi \rangle$. Therefore, the dimension of the resulting deformation of the dual field theory agrees with the dimension of the fermionic mass deformation, $M \bar{\Psi} \Psi$, appearing in (1).

Throughout this work, we will work in units in which $2\kappa^2 = L = 1$. Furthermore, we will set $\alpha = q = \lambda = 1$,⁷ so that the quantum critical point at zero magnetic field will be at $\mathcal{M}_{\text{crit}} := \frac{M}{b}|_{\text{crit}} \approx 0.954$, where M and b are the mass and time-reversal breaking parameters appearing in (1) and, in the holographic model described by (25). They are implemented via the following boundary conditions on the scalar and axial vector fields Φ and A_z :

$$\lim_{r \rightarrow \infty} r\Phi(r) = M; \quad \lim_{r \rightarrow \infty} A_z(r) = b. \quad (27)$$

The field equations coming from the action (25) are given by

$$T_{ab} = \frac{1}{2} \left(\mathcal{F}_{ac} \mathcal{F}_b^c - \frac{1}{4} g_{ab} \mathcal{F}^2 \right) + \frac{1}{2} \left(F_{ac} F_b^c - \frac{1}{4} g_{ab} F^2 \right) + \frac{1}{2} \left((D_a \Phi)^* D_b \Phi + (D_b \Phi)^* D_a \Phi \right) - \frac{1}{2} g_{ab} \left((D_c \Phi)^* D^c \Phi + V(\Phi) \right). \quad (32)$$

The finite-temperature ansatz we consider for the background metric, scalar field, vector, and axial vector fields are given by

$$ds^2 = -f(r)dt^2 + \frac{dr^2}{f(r)} + u(r)(dx^2 + dy^2) + h(r)dz^2, \\ \Phi = \Phi(r), \\ V = V_t(r)dt + \frac{B}{2}(-ydx + xdy); \quad A = A_z(r)dz, \quad (33)$$

where B is a constant magnetic field along the z direction. Using this ansatz, the equations of motion for the background functions u, f, h, A_z, V_t, Φ are given by

$$\frac{u''}{u} + \frac{f''}{2f} + \left(\frac{f'}{f} - \frac{u'}{2u} \right) \frac{u'}{2u} - \frac{1}{4} \left(\frac{V_t'^2}{f} + \frac{A_z'^2}{h} \right) + \frac{1}{2} \Phi'^2 \\ + (h(\lambda\Phi^2 - 6) - 2q^2 A_z^2) \frac{\Phi^2}{4hf} - \frac{6}{f} + \frac{B^2}{4u^2 f} = 0, \quad (34)$$

$$\frac{f''}{f} - \frac{u''}{u} + \left(\frac{f'}{f} - \frac{u'}{u} \right) \frac{h'}{2h} - \frac{V_t'^2}{f} - \frac{B^2}{fu^2} = 0, \quad (35)$$

$$\Phi'' + \left(\frac{f'}{f} + \frac{h'}{2h} + \frac{u'}{u} \right) \Phi' + \left(3 - q^2 \frac{A_z^2}{h} - \lambda\Phi^2 \right) \frac{\Phi}{f} = 0, \quad (36)$$

⁷In [52], it was shown that for fixed m^2 , which is the case in the present paper, that the quantum phase transition persists for a wide range of the quartic coupling λ , and that there is an upper bound above which the transition to the trivial phase can not be reached anymore. Since we are interested in studying the effects of the magnetic field on the quantum critical point, we have fixed the quartic coupling to a convenient value.

$$R_{ab} - \frac{1}{2} g_{ab} (R + 12) - T_{ab} = 0, \quad (28)$$

$$\nabla_b (F^{ba}) + 2\alpha \epsilon^{abcde} \mathcal{F}_{bc} F_{de} = 0, \quad (29)$$

$$\nabla_b (\mathcal{F}^{ba}) + \alpha \epsilon^{abcde} (\mathcal{F}_{bc} \mathcal{F}_{de} + F_{bc} F_{de}) \\ - iq(\Phi^* D^a \Phi - (D^a \Phi)^* \Phi) = 0, \quad (30)$$

$$D_a D^a \Phi - \partial_{\Phi^*} V(\Phi) = 0, \quad (31)$$

where T_{ab} is the total energy-momentum tensor, which is given by

$$A_z'' + \left(\frac{f'}{f} - \frac{h'}{2h} + \frac{u'}{u} \right) A_z' - \frac{2q^2 \Phi^2}{f} A_z + \frac{8\alpha B \sqrt{h}}{fu} V_t' = 0, \quad (37)$$

$$\left(\sqrt{h} u V_t' + 8\alpha B A_z \right)' = 0, \quad (38)$$

and a first-order constraint given by

$$\left(\frac{h'}{h} + \frac{u'}{2u} \right) \frac{u'}{2u} + \left(\frac{h'}{2h} + \frac{u'}{u} \right) \frac{f'}{2f} + \frac{1}{4} \left(\frac{V_t'^2}{f} - \frac{A_z'^2}{h} \right) - \frac{1}{2} \Phi'^2 \\ + (h(\lambda\Phi^2 - 6) + 2q^2 A_z^2) \frac{\Phi^2}{4hf} - \frac{6}{f} + \frac{B^2}{4u^2 f} = 0. \quad (39)$$

The prime $'$ denotes derivative with respect to the radial coordinate r . This system of differential equations can only be solved numerically with proper boundary conditions (see Appendix A for more details on the asymptotic expansions and the numerical integration method). Note that an analytic solution for V_t in (38) can be formally obtained as

$$V_t(r) = \mu \left(1 - \frac{\int_{\infty}^r \frac{1}{u(x)\sqrt{h(x)}} dx}{\int_{\infty}^{r_h} \frac{1}{u(x)\sqrt{h(x)}} dx} \right) \\ + 8\alpha B \left(\frac{\int_{\infty}^{r_h} \frac{A_z(x)}{\sqrt{h(x)u(x)}} dx}{\int_{\infty}^{r_h} \frac{1}{\sqrt{h(x)u(x)}} dx} \int_{\infty}^r \frac{1}{\sqrt{h(x)u(x)}} dx \right. \\ \left. - \int_{\infty}^r \frac{A_z(x)}{\sqrt{h(x)u(x)}} dx \right), \quad (40)$$

where we used the boundary conditions $V_t(\infty) = \mu$ and $V_t(r_h) = 0$ and where μ is the chemical potential and r_h is the horizon position, such that $f(r_h) = 0$. Since the Weyl semimetal phase is a zero density state, in this work, we will take $\mu = 0$. Then, the formal solution for V_t becomes

$$V_t(r) = 8\alpha B \left(\frac{\int_{\infty}^{r_h} \frac{A_z(x)}{\sqrt{h(x)u(x)}} dx}{\int_{\infty}^{r_h} \frac{1}{\sqrt{h(x)u(x)}} dx} \int_{\infty}^r \frac{1}{\sqrt{h(x)u(x)}} dx - \int_{\infty}^r \frac{A_z(x)}{\sqrt{h(x)u(x)}} dx \right).$$

B. Numerical results

In this subsection, we present our numerical results for the bulk profile of the background fields and for the quantum critical point (QCP) as a function of the magnetic field. The results were obtained by numerically solving the system of equations (34)–(38) at small but finite

temperature. The reason behind this is that the boundary conditions for the full zero-temperature solution with magnetic field are very tricky, and we will explore this analysis further in an upcoming publication.

Due to the scaling symmetries of the background, the only relevant parameters are the dimensionless ratios which we are going to define, for convenience, as follows:

$$\begin{aligned} \mathcal{T} &:= \frac{T}{b} = 0.05; & \text{Temperature (fixed),} \\ \mathcal{M} &:= \frac{M}{b}; & \text{Effective coupling,} \\ \mathcal{B} &:= \frac{B}{T^2}; & \text{Magnetic field,} \\ \mathcal{M}_{\text{crit}} &:= \left(\frac{M}{b} \right)_{\text{crit}}; & \text{Quantum critical point (QCP),} \end{aligned}$$

with \mathcal{M} and \mathcal{B} being the free parameters of the model.

1. Background profiles

In Figs. 4–6, we display the holographic RG flow of the axial vector and scalar fields at all topological phases,

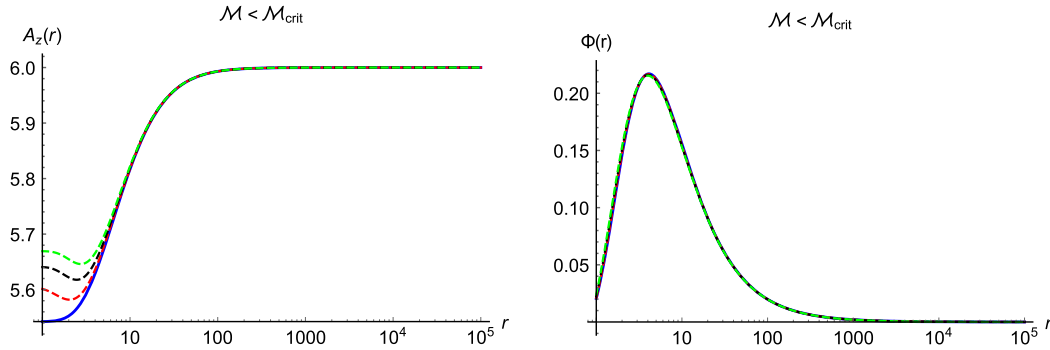


FIG. 4. Bulk profile of the axial vector and scalar field in the topologically nontrivial (semimetal) phase at $T = 0.05$ for different values of magnetic field: $\mathcal{B} = 0$ (blue curve), $\mathcal{B} = \pi^2$ (dashed red), $\mathcal{B} = 2\pi^2$ (dashed black), $\mathcal{B} = 3\pi^2$ (dashed green). Here, we took $\mathcal{M} = 1/3$.

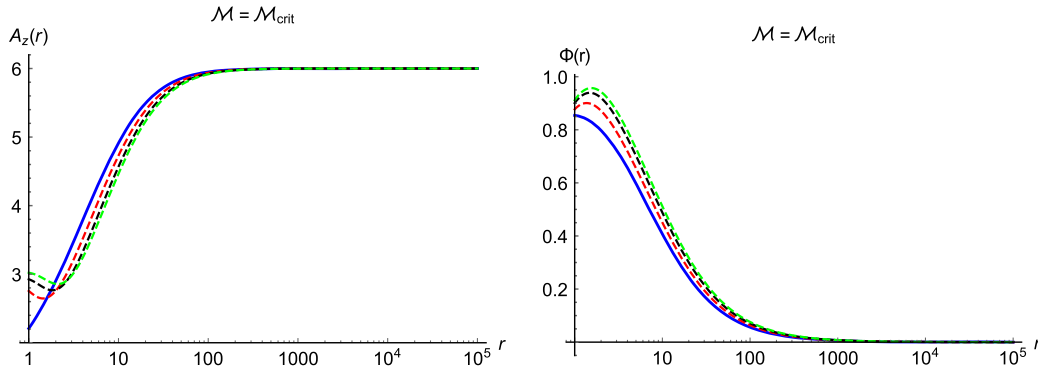


FIG. 5. Bulk profile of the axial vector and scalar field in the quantum critical phase at $T = 0.05$ and for different values of magnetic field: $\mathcal{B} = 0$ (blue curve), $\mathcal{B} = \pi^2$ (dashed red), $\mathcal{B} = 2\pi^2$ (dashed black), and $\mathcal{B} = 3\pi^2$ (dashed green). Here, we took $\mathcal{M} = \mathcal{M}_{\text{crit}} \approx (0.954, 1.09, 1.21, 1.30)$ for $\mathcal{B} = (0, \pi^2, 2\pi^2, 3\pi^2)$, respectively.

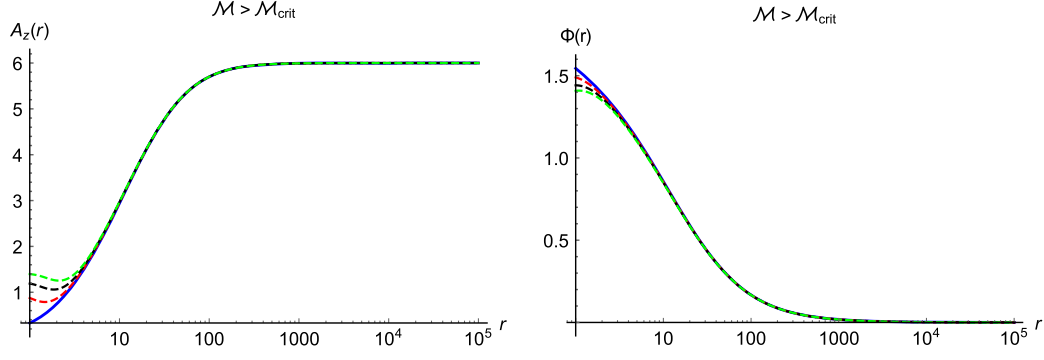


FIG. 6. Bulk profile of the axial vector and scalar field in the topologically trivial (insulator) phase at $\mathcal{T} = 0.05$ and for different values of magnetic field: $B = 0$ (blue curve), $B = \pi^2$ (dashed red), $B = 2\pi^2$ (dashed black), and $B = 3\pi^2$ (dashed green). Here, we took $\mathcal{M} = 3$.

namely the semimetal phase ($\mathcal{M} < \mathcal{M}_{\text{crit}}$), the critical phase ($\mathcal{M} = \mathcal{M}_{\text{crit}}$), and the topologically trivial phase ($\mathcal{M} > \mathcal{M}_{\text{crit}}$) for different values of B and small \mathcal{M} . Notice that the value of scalar field is almost zero at the horizon in the semimetal phase, it jumps at the critical point, and in the trivial phase, it remains increasing near the horizon. Meanwhile, the axial field has a continuous behavior in all phases for all values of the magnetic field B . However, at a nonzero magnetic field, one can observe that the axial vector field develops a dip near the horizon, which makes its behavior nonmonotonic for a nonzero field.

Finally, in Fig. 7, we display the bulk profile of the vector field in all topological phases. One can observe that it increases as we increase the magnetic field, as can be seen from the formal solution (41) that is proportional to B . Also, note that the vector field has its peak increased substantially in the critical region $\mathcal{M} = \mathcal{M}_{\text{crit}}$, and it remains stable in the trivial phase $\mathcal{M} > \mathcal{M}_{\text{crit}}$. Finally, it is worthy to mention that at the UV boundary, the vector field is always zero because we have set the chemical potential to zero ($\mu = 0$); i.e., there is no electric charge density.

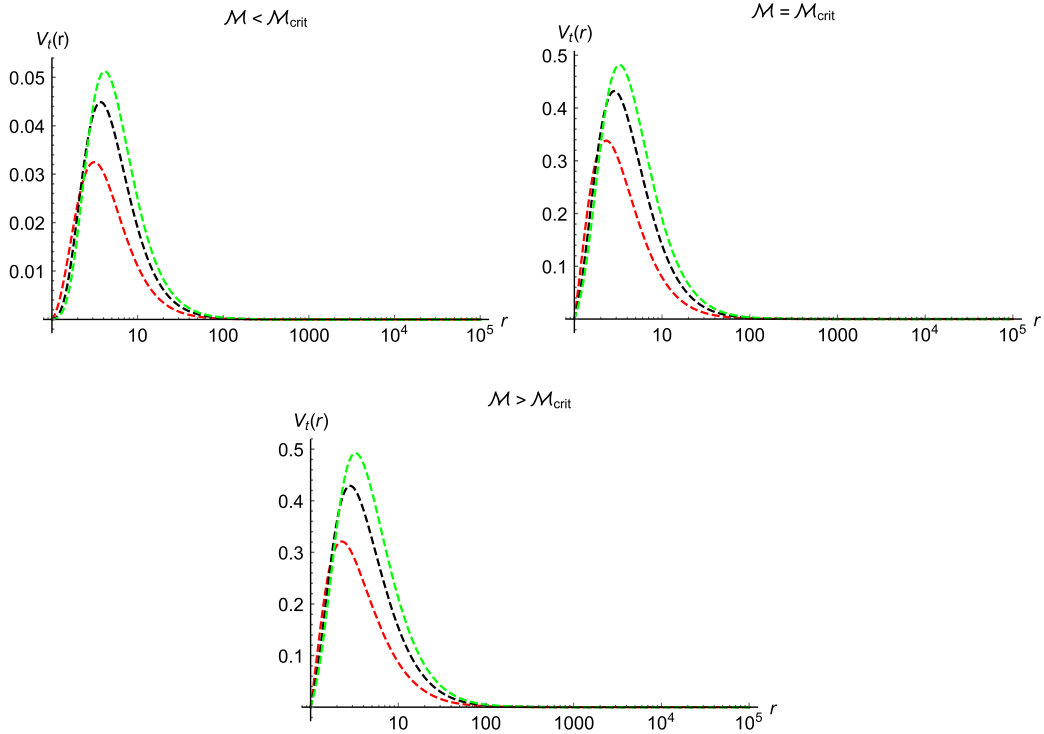


FIG. 7. Bulk profile of the vector potential in all topological phases at $\mathcal{T} = 0.05$ and for different values of magnetic field: $B = \pi^2$ (dashed red), $B = 2\pi^2$ (dashed black), and $B = 3\pi^2$ (dashed green). Here, we used the same values of \mathcal{M} in the previous figures.

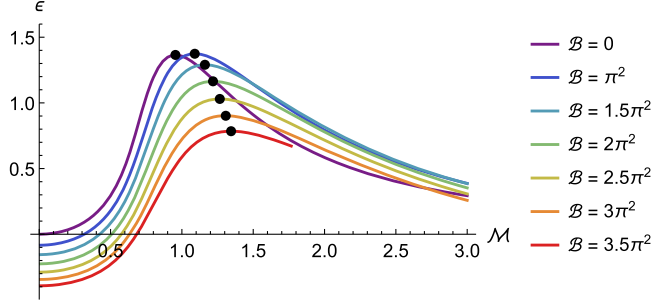


FIG. 8. Horizon spatial anisotropy parameter ϵ as a function of the effective coupling \mathcal{M} for different values of the magnetic field \mathcal{B} at $T = 0.05$. The black dots represent the QCP for each value magnetic field.

2. QCP as a function of \mathcal{B}

It has been known that the order parameter of the quantum phase transition is the anomalous Hall conductivity σ_{AHE} , which was obtained in holography in terms of the horizon data [47,48]. Thus, the quantum critical point (QCP) $\mathcal{M}_{\text{crit}}$ can be extracted from the behavior of σ_{AHE} as a function of the effective coupling \mathcal{M} at both a zero and finite temperature. However, in the presence of the magnetic field, we found it more convenient to use another probe of the QCP, which is the spatial anisotropy along the z direction in terms of the horizon data. It is defined as follows:

$$\epsilon := \frac{h(r_h)}{u(r_h)} - 1. \quad (41)$$

The anisotropy parameter ϵ as a function of the effective coupling \mathcal{M} has proven to be a good probe to detect the location of the QCP as it is peaked around it, suggesting it is a point or region (in the case of finite temperature) of maximum anisotropy and strong divergences at zero temperature, as shown in [82].

Note, however, that the (sharp) quantum phase transition is expected to happen at zero temperature, and at a finite temperature, it turns into a smooth crossover. Here, since we are working at finite but small temperature ($T \sim 10^{-2}$),

the quantum critical point can still be estimated with certain reliability compared to the zero temperature case.

From the numerical solution of the background, we extracted the anisotropy parameter ϵ as a function of the effective coupling \mathcal{M} for several values magnetic field \mathcal{B} , as displayed in Fig. 8. From those anisotropy curves, we extracted the location of the QCP for each magnetic field by computing the value of \mathcal{M} for which ϵ is maximum. These points are represented by the black dots in Fig. 8. From this, one can clearly see that the QCP increases as we increase the value of the magnetic field \mathcal{B} , suggesting that it is favorable for the system to stay in the WSM phase (nontrivial topological phase) in the presence of a finite magnetic field, which is in agreement with the experimental results on the effects of the magnetic field on other kinds of topological materials [71,72].

In order to comprehend in more detail how the QCP behaves as a function of the magnetic field, in the left panel of Fig. 9, we display the behavior of $\mathcal{M}_{\text{crit}}$ as a function of \mathcal{B} . Note that it presents different scalings depending on the regime of the magnetic field \mathcal{B} . For instance, at small \mathcal{B} , it has a quadratic-in- \mathcal{B} dependence, while for large \mathcal{B} , it has a linear-in- \mathcal{B} dependence, as shown in the fittings presented in the right panel of Fig. 9. Those behaviors seem to agree with the semiclassical and ultraquantum magnetotransport regimes previously described in the literature within the framework of Boltzmann transport theory (see [78] and references therein). It is interesting to see that this holographic model can capture those behaviors through the QCP. Ultimately, they are manifestations of quantum anomalies, notably the chiral anomaly, which is responsible for the quadratic-in- \mathcal{B} in the magnetoconductivity, as experimentally shown in [83]. However, in the intermediate range of \mathcal{B} , it should be expected a quantum oscillatory behavior periodic in $1/\mathcal{B}$ in the magnetotransport coefficients as described in [78], in which the present holographic model is not capable of capturing it. For a more recent discussion on the role played by the quantum oscillations within the nonlinear response regime in WSMs, see [84].

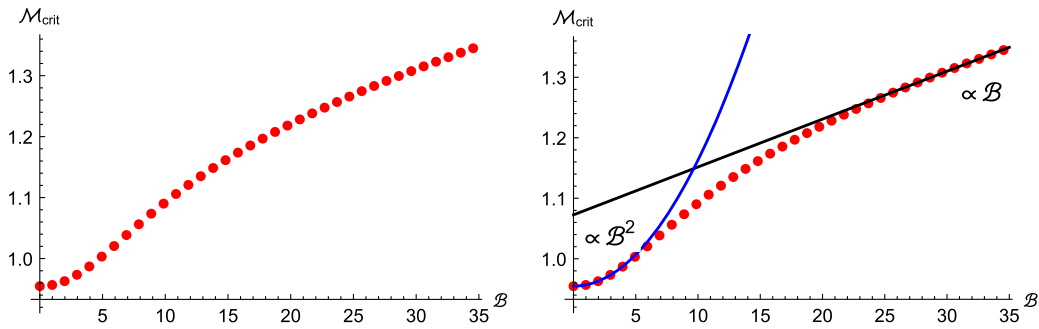


FIG. 9. Left panel: Quantum critical point (QCP), $\mathcal{M}_{\text{crit}}$, as a function of the magnetic field \mathcal{B} at fixed small temperature $T = 0.05$. Right panel: Quantum critical point (QCP), $\mathcal{M}_{\text{crit}}$, as a function of the magnetic field \mathcal{B} at fixed small temperature $T = 0.05$ together with a quadratic-in- \mathcal{B} and linear-in- \mathcal{B} fittings.

IV. CONCLUSION AND DISCUSSION

In this work, we have considered an extension of the holographic Weyl semimetal model and its quantum phase transition to a trivial semimetal by including a finite magnetic field at a small but finite-temperature in order to investigate the effects of the magnetic field on the quantum critical point. For this purpose, we have made use of the spatial anisotropy parameter in terms of the horizon data and, from it, we extracted the QCP, $\mathcal{M}_{\text{crit}}$, as a function of the magnetic field \mathcal{B} . We have found that $\mathcal{M}_{\text{crit}}(\mathcal{B})$ displays a quadratic-in- \mathcal{B} for a weak field and linear-in- \mathcal{B} for a strong field compatible with results found in the literature within the Boltzmann transport theory. Since we have considered a finite but small temperature holographic model, it would be interesting to try to construct a background at zero temperature in the presence of a magnetic field [85] and compare the results with the ones obtained here for $\mathcal{M}_{\text{crit}}(\mathcal{B})$. Furthermore, one could study the effects of the magnetic field on the QCP by analyzing other probes that have been proposed recently, such as the c function [67] and the entanglement entropy (EE) for the case of nodal line WSMs [86].

Despite the fact the electrical transport coefficients, such as the electrical conductivity, has been extensively studied in the holographic semimetal model [48,51,52,55,58,62], much less attention has been given to the holographic calculation of magnetotransport in WSMs. From the experimental observations, we know that the longitudinal magnetoconductivity in the Weyl semimetal is enhanced by the chiral anomaly, which is called as positive magnetoconductivity or, equivalently, negative magnetoresistivity. Experimentally, it was observed that the longitudinal conductivity in this topological material has a B^2 dependence at a small magnetic field as a consequence of the chiral magnetic effect (CME), which is the generation of a current along the direction of the magnetic field. Therefore, it would be interesting as a possible extension of this work to consider the effects of including vector and chiral chemical potentials in the current holographic WSM model, in order to study the magnetoconductivity and see if one can observe the presence of the CME.

Finally, another possible extension of this work is the study of the role played by the mixed-chiral gravitational

anomaly, which yields a nonconservation of the chiral charge and energy [78,87,88]. This anomaly leads to the enhancement of the longitudinal magnetothermal conductivity and magnetothermoelectric transport coefficients [89]. Thus, it would be interesting to consider this term in the holographic model, as done in [66], to study the chiral-vortical conductivity in holography, and investigate the magnetothermoelectric transport coefficients and compare with experimental predictions [90–92]. We will leave these further studies for the future.

ACKNOWLEDGMENTS

We would like to thank Yan Liu for useful discussions and motivation in the early stage of this project and for providing us interesting references on experimental results in topological materials. D. M. R. is supported by Grant No. 2021/01565-8, São Paulo Research Foundation (FAPESP). L. F. F. is supported by ANID Fondecyt Postdoctoral Grant Folio No. 3220304.

APPENDIX A: IR AND UV ASYMPTOTIC EXPANSIONS AND NUMERICS

Here, we provide the asymptotic expansions for the background fields near the horizon $r \rightarrow r_h$ (IR) and near the boundary $r \rightarrow \infty$ (UV) as well as we briefly explain how we numerically integrate the equations of motion (34)–(38).

Near the horizon, we assume the background fields can be expressed in a regular Taylor expansion around $r = r_h$, which can be written as

$$A_z = A_0 + \frac{-4\alpha B \sqrt{h_0} V_1 + A_0 q^2 u_0 \Phi_0^2}{2\pi T u_0} (r - r_h) + \dots, \quad (\text{A1})$$

$$V_t = V_1 (r - r_h) + \dots, \quad (\text{A2})$$

$$\Phi = \Phi_0 + \frac{\Phi_0 (A_0^2 q^2 + h_0 (-3 + \lambda \Phi_0^2))}{4\pi T h_0} (r - r_h) + \dots, \quad (\text{A3})$$

$$f = 4\pi T (r - r_h) + \dots, \quad (\text{A4})$$

$$u = u_0 + \frac{2B^2 + u_0^2 (-24 + V_1^2 + \Phi_0^2 (-6 + \lambda \Phi_0^2))}{12\pi T u_0} (r - r_h) + \dots, \quad (\text{A5})$$

$$h = h_0 + \frac{B^2 h_0 - u_0^2 (6A_0^2 q^2 \Phi_0^2 + h_0 [-24 + V_1^2 + \Phi_0^2 (-6 + \lambda \Phi_0^2)])}{12\pi T u_0^2} (r - r_h) + \dots, \quad (\text{A6})$$

where in the numerics, we set $r_h = 1$ and $T = \frac{1}{\pi}$, and the shooting parameters near the horizon are A_0 , V_1 , Φ_0 , u_0 , and h_0 .

On the other hand, near the boundary $r \rightarrow \infty$, we impose that the background is asymptotically AdS. Thus, the asymptotic expansion for the various fields can be written as

$$A_z = b + \frac{bM^2q^2(9b^2M^2q^2 - 192\alpha^2B^2 + M^4(9\lambda + 9q^2 + 5))}{12(M^4(q^2 + 1) - 64\alpha^2B^2)r^2} - \frac{bM^2q^2 \ln(r)}{r^2} + \dots, \quad (\text{A7})$$

$$V_t = \mu + \frac{2abBM^2q^2(9b^2q^2 + M^2(9\lambda + 6q^2 + 2))}{(192\alpha^2B^2 - 3M^4(q^2 + 1))r^2} - \frac{2abBM^2q^2 \ln(r)}{r^4} + \dots, \quad (\text{A8})$$

$$\Phi = \frac{M}{r} - \frac{M(3b^2q^2 + (3\lambda + 2)M^2) \ln(r)}{6r^3} + \dots, \quad (\text{A9})$$

$$f = r^2 - \frac{M^2}{3} + \frac{(-3B^2 + 3\lambda M^4 + 2M^4) \ln(r)}{18r^2} + \dots, \quad (\text{A10})$$

$$u = r^2 - \frac{M^2}{3} + \frac{(3B^2 + 6\lambda M^4 + 4M^4) \ln(r)}{36r^2} + \dots, \quad (\text{A11})$$

$$h = r^2 - \frac{M^2}{3} + \frac{9b^2M^2q^2 + (9\lambda + 14)M^4}{72r^2} + \frac{(9b^2M^2q^2 - 3B^2 + (3\lambda + 2)M^4) \ln(r)}{18r^2}. \quad (\text{A12})$$

Since we have considered in this work the case of zero chemical potential, we have set $\mu = 0$ in numerics. In addition, due to the underlying conformal symmetry of the background, the relevant model parameters are the dimensionless ratios previously defined in the main text ($T/b := \mathcal{T}$; $M/b := \mathcal{M}$; $B/T^2 := \mathcal{B}$).

The numerical integration of the background equations of motion (34)–(38) were obtained using a matching technique (for more details and practical examples, we refer the reader to [93,94]). In the following, we briefly summarize the numerical procedure,

- (i) One solves numerically the equations from the boundary to an intermediate point by imposing the asymptotic expansion at boundary;
- (ii) One solves numerically the equations from the horizon to an intermediate point by imposing the asymptotic expansion at horizon;
- (iii) One matches the numerical solutions from the first and second steps at the intermediate point using a `FindRoot` routine in *Mathematica*. This will give five equations ($\Psi_{r \rightarrow \infty}^i - \Psi_{r \rightarrow r_h}^i$, where $\Psi^i = A, V_t, \Phi, u, h$) for the five parameters, namely A_0, V_1, Φ_0, u_0 , and h_0 . For fixed \mathcal{T} , these parameters

will vary as we vary the effective coupling \mathcal{M} and the magnetic field \mathcal{B} .

One could also numerically solve the equations of motions by making use of the scaling symmetries presented in the background (for more details see, for instance, [48,57]).

APPENDIX B: ON THE FREE ENERGY AND PHASE TRANSITION

In this appendix, we compute the free energy as a function of \mathcal{M} at finite \mathcal{T} and \mathcal{B} in the holographic Weyl semimetal model using the holographic renormalization prescription. This can be done by evaluating the Euclidean bulk on shell action plus the Gibbons-Hawking boundary action together with the appropriate counterterms to cancel the UV ($r \rightarrow \infty$) divergences. For this purpose, we consider the renormalized gravitational action as given by

$$S_{\text{ren}} = S + S_{\text{GH}} + S_{\text{c.t.}}, \quad (\text{B1})$$

where $S_{\text{GH}} = \int_0^{\beta=1/T} d\tau \int d^3x \sqrt{-\gamma} 2K|_{r=\epsilon}$ is the Gibbons-Hawking boundary term, and ϵ is a UV cutoff. The counterterms are written in the following form:

$$S_{\text{c.t.}} = \int_0^{\beta=1/T} d\tau \int d^3x \sqrt{-\gamma} \left(-6 - |\Phi|^2 + \frac{1}{2} (\log r^2) \left[\frac{1}{4} F^2 + \frac{1}{4} \mathcal{F}^2 + |D_m \Phi|^2 + |\Phi|^4 \left(\frac{1}{3} + \frac{\lambda}{2} \right) \right] \right) \Big|_{r=\epsilon}, \quad (\text{B2})$$

where $\gamma_{\mu\nu}$ is the induced metric on the boundary $r = \infty$, and the trace of the extrinsic curvature $K = \gamma^{\mu\nu} \nabla_\mu n_\nu$ with n_ν being the outward unit vector normal to the boundary. Then, using the equations of motion, the on shell action becomes

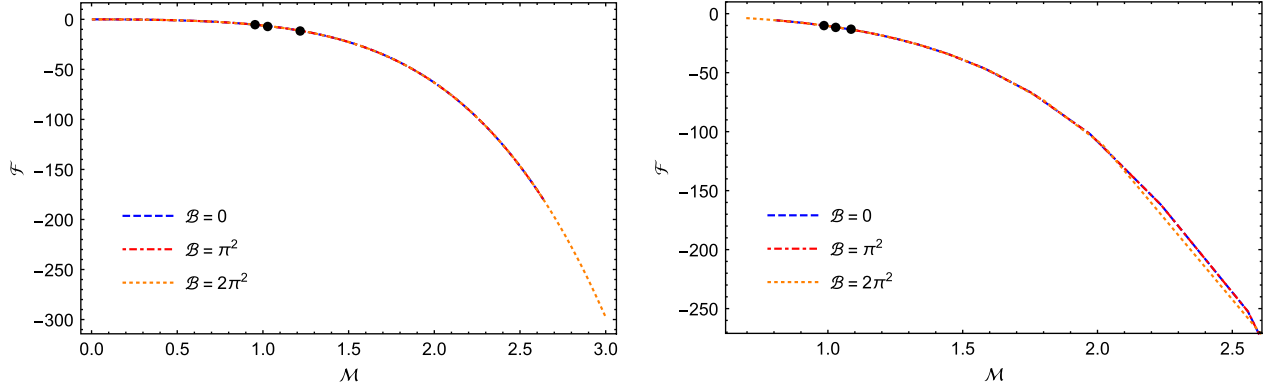


FIG. 10. Free energy density \mathcal{F} as a function of \mathcal{M} for $\mathcal{T} = 0.05$ (left panel) and $\mathcal{T} = 0.01$ (right panel) and different values of magnetic field: $B = 0$ (blue), $B = \pi^2$ (red), and $B = 2\pi^2$ (orange). The black dots represent the location of the QCP. Note that \mathcal{F} is continuous and smooth around and at the critical points $\mathcal{M}_{\text{crit}}$ at finite \mathcal{T} , which implies that the topological phase transition is a crossover at finite temperature.

$$\begin{aligned} \frac{S_{\text{on-shell}}}{V} = & -\beta \int_{\epsilon}^{r_h} \sqrt{hu} \left(8 + \frac{B^2}{3u^2} + 2\Phi^2 - \frac{\lambda}{3}\Phi^4 + \frac{fA'_z}{3h} - \frac{V'^2}{3} \right) dr + \beta \sqrt{hu} \left(f' + f \left(\frac{h'}{h} + \frac{2u'}{u} \right) \right) \Big|_{r=\epsilon} \\ & + \beta \left[\sqrt{fhu} \left[-6 - \Phi^2 + \log r \left(\frac{A_z^2 \Phi^2}{h} + \Phi'^2 + \left(\frac{1}{3} + \frac{\lambda}{2} \right) \Phi^4 - \frac{1}{2} V'^2 + \frac{fA_z'^2}{2h} + \frac{B^2}{2u^2} \right) \right] \right] \Big|_{r=\epsilon}, \quad (\text{B3}) \end{aligned}$$

where $V \equiv \int d^3x$ is the spatial volume. Thus, the free energy density is given by

$$F = -T \frac{S_{\text{on shell}}}{V}. \quad (\text{B4})$$

Now, by numerically evaluating the renormalized on shell action, given by (B3), and using the definition (B4), one can clearly see that $\mathcal{F} := \frac{F}{V}$ is negative and continuous for all \mathcal{M} finite, as shown in Fig. 10 for different values of magnetic field and two fixed low temperatures, $\mathcal{T} = 0.05$ and $\mathcal{T} = 0.01$. The results show that there is no *first-order* phase transition and, in this case, the system is always in the black hole phase for any finite \mathcal{T} , according with the criteria

in [95]. Furthermore, as also a established criterion in [95], all nonconfining geometries at zero- T are always in the black hole phase for any finite T , and this includes the holographic geometry in the present paper, since it is a nonconfining background.

Finally, it is worth to emphasize that free energy density does not capture the topological nature of the quantum phase transition between the nontrivial topological phase and the trivial one in the holographic Weyl semimetal. As shown in Fig. 10, the free energy density is continuous and smooth at the QCPs. This is different from the traditional concept of thermal phase transition, which is formulated based on the Landau-Ginsburg-Wilson paradigm, and is a consequence of symmetry breaking.

-
- [1] A. P. Schnyder, S. Ryu, A. Furusaki, and A. W. W. Ludwig, *Phys. Rev. B* **78**, 195125 (2008).
 - [2] C.-K. Chiu, J. C. Teo, A. P. Schnyder, and S. Ryu, *Rev. Mod. Phys.* **88**, 035005 (2016).
 - [3] C. L. Kane and E. J. Mele, *Phys. Rev. Lett.* **95**, 146802 (2005).
 - [4] L. Fu, C. L. Kane, and E. J. Mele, *Phys. Rev. Lett.* **98**, 106803 (2007).
 - [5] M. Z. Hasan and C. L. Kane, *Rev. Mod. Phys.* **82**, 3045 (2010).
 - [6] X.-L. Qi and S.-C. Zhang, *Rev. Mod. Phys.* **83**, 1057 (2011).
 - [7] C. Kallin and J. Berlinsky, *Rep. Prog. Phys.* **79**, 054502 (2016).
 - [8] M. Sato and Y. Ando, *Rep. Prog. Phys.* **80**, 076501 (2017).
 - [9] N. Armitage, E. Mele, and A. Vishwanath, *Rev. Mod. Phys.* **90**, 015001 (2018).
 - [10] P. Hosur and X. Qi, *C.R. Phys.* **14**, 857 (2013).
 - [11] S. Jia, S.-Y. Xu, and M. Z. Hasan, *Nat. Mater.* **15**, 5 (2016).
 - [12] T. O. Wehling, A. M. Black-Schaffer, and A. V. Balatsky, *Adv. Phys.* **63**, 1 (2014).
 - [13] H. Gao, J. W. F. Venderbos, Y. Kim, and A. M. Rappe, *Annu. Rev. Mater. Res.* **49**, 153 (2019).

- [14] O. Vafek and A. Vishwanath, *Annu. Rev. Condens. Matter Phys.* **5**, 83 (2014).
- [15] H. Weyl, *Z. Phys.* **56**, 330 (1929).
- [16] C. Herring, *Phys. Rev.* **52**, 365 (1937).
- [17] A. A. Burkov, M. D. Hook, and L. Balents, *Phys. Rev. B* **84**, 235126 (2011).
- [18] G. B. Halász and L. Balents, *Phys. Rev. B* **85**, 035103 (2012).
- [19] P. Delplace, J. Li, and D. Carpentier, *Europhys. Lett.* **97**, 67004 (2012).
- [20] D. Pesin and L. Balents, *Nat. Phys.* **6**, 376 (2010).
- [21] X. Wan, A. M. Turner, A. Vishwanath, and S. Y. Savrasov, *Phys. Rev. B* **83**, 205101 (2011).
- [22] K.-Y. Yang, Y.-M. Lu, and Y. Ran, *Phys. Rev. B* **84**, 075129 (2011).
- [23] Shuichi Murakami, *New J. Phys.* **9**, 356 (2007).
- [24] P. Hosur, S. A. Parameswaran, and A. Vishwanath, *Phys. Rev. Lett.* **108**, 046602 (2012).
- [25] A. A. Burkov and L. Balents, *Phys. Rev. Lett.* **107**, 127205 (2011).
- [26] A. A. Zyuzin, S. Wu, and A. A. Burkov, *Phys. Rev. B* **85**, 165110 (2012).
- [27] A. A. Zyuzin and A. A. Burkov, *Phys. Rev. B* **86**, 115133 (2012).
- [28] P. Goswami and S. Tewari, *Phys. Rev. B* **88**, 245107 (2013).
- [29] A. G. Grushin, *Phys. Rev. D* **86**, 045001 (2012).
- [30] M. M. Vazifeh and M. Franz, *Phys. Rev. Lett.* **111**, 027201 (2013).
- [31] H. B. Nielsen and M. Ninomiya, *Nucl. Phys.* **B185**, 20 (1981).
- [32] H. B. Nielsen and M. Ninomiya, *Nucl. Phys.* **B193**, 173 (1981).
- [33] G. E. Volovik, in *Quantum Analogues: From Phase Transitions to Black Holes and Cosmology*, Lecture Notes in Physics, edited by W. G. Unruh and R. Schützhold (Springer, Berlin, Heidelberg, 2007), pp. 31–73.
- [34] F. R. Klinkhamer and G. E. Volovik, *Int. J. Mod. Phys. A* **20**, 2795 (2005).
- [35] S.-Y. Xu *et al.*, *Science* **349**, 613 (2015).
- [36] L. Lu, Z. Wang, D. Ye, L. Ran, L. Fu, J. D. Joannopoulos, and M. Solja i, *Science* **349**, 622 (2015).
- [37] I. Belopolski, S.-Y. Xu, D. Sanchez, G. Chang, C. Guo, M. Neupane, H. Zheng, C.-C. Lee, S.-M. Huang, G. Bian, N. Alidoust, T.-R. Chang, B. Wang, X. Zhang, A. Bansil, H.-T. Jeng, H. Lin, S. Jia, and M. Z. Hasan, *Phys. Rev. Lett.* **116**, 066802 (2016).
- [38] C.-X. Liu, P. Ye, and X.-L. Qi, *Phys. Rev. B* **87**, 235306 (2013).
- [39] K. Fukushima, D. E. Kharzeev, and H. J. Warringa, *Phys. Rev. D* **78**, 074033 (2008).
- [40] D. T. Son and B. Z. Spivak, *Phys. Rev. B* **88**, 104412 (2013).
- [41] D. E. Kharzeev and H. J. Warringa, *Phys. Rev. D* **80**, 034028 (2009).
- [42] X. Huang, L. Zhao, Y. Long, P. Wang, D. Chen, Z. Yang, H. Liang, M. Xue, H. Weng, Z. Fang, X. Dai, and G. Chen, *Phys. Rev. X* **5**, 031023 (2015).
- [43] C.-L. Zhang *et al.*, *Nat. Commun.* **7**, 10735 (2016).
- [44] J. Zaanen, Y.-W. Sun, Y. Liu, and K. Schalm, *Holographic Duality in Condensed Matter Physics* (Cambridge University Press, Cambridge, England, 2015).
- [45] S. A. Hartnoll, A. Lucas, and S. Sachdev, *Holographic Quantum Matter*, illustrated edição ed. (The MIT Press, Cambridge, Massachusetts, London, England, 2018).
- [46] H. Nastase, *String Theory Methods for Condensed Matter Physics* (Cambridge University Press, Cambridge, England, 2017).
- [47] K. Landsteiner and Y. Liu, *Phys. Lett. B* **753**, 453 (2016).
- [48] K. Landsteiner, Y. Liu, and Y.-W. Sun, *Phys. Rev. Lett.* **116**, 081602 (2016).
- [49] M. Ammon, M. Heinrich, A. Jiménez-Alba, and S. Moeckel, *Phys. Rev. Lett.* **118**, 201601 (2017).
- [50] K. Landsteiner, Y. Liu, and Y.-W. Sun, *Phys. Rev. Lett.* **117**, 081604 (2016).
- [51] G. Grignani, A. Marini, F. Pena-Benitez, and S. Speziali, *J. High Energy Phys.* **03** (2017) 125.
- [52] C. Copetti, J. Fernández-Pendás, and K. Landsteiner, *J. High Energy Phys.* **02** (2017) 138.
- [53] Y. Liu and Y.-W. Sun, *J. High Energy Phys.* **10** (2018) 189.
- [54] J. Zhao, *Eur. Phys. J. C* **82**, 300 (2022).
- [55] J. Zhao, *Phys. Rev. D* **104**, 066003 (2021).
- [56] Y. Liu and Y.-W. Sun, *J. High Energy Phys.* **12** (2018) 072.
- [57] Y. Liu and X.-M. Wu, *J. High Energy Phys.* **05** (2021) 141.
- [58] R. Rodgers, E. Mauri, U. Gürsoy, and H. T. C. Stoof, *J. High Energy Phys.* **11** (2021) 191.
- [59] X. Ji, Y. Liu, Y.-W. Sun, and Y.-L. Zhang, *J. High Energy Phys.* **12** (2021) 066.
- [60] K. Landsteiner, Y. Liu, and Y.-W. Sun, *Sci. China Phys. Mech. Astron.* **63**, 250001 (2020).
- [61] U. Gürsoy, V. Jacobs, E. Plauschinn, H. Stoof, and S. Vandoren, *J. High Energy Phys.* **04** (2013) 127.
- [62] V. P. J. Jacobs, P. Betzios, U. Gürsoy, and H. T. C. Stoof, *Phys. Rev. B* **93**, 195104 (2016).
- [63] K. Hashimoto, S. Kinoshita, K. Murata, and T. Oka, *J. High Energy Phys.* **05** (2017) 127.
- [64] Y. Liu and J. Zhao, *J. High Energy Phys.* **12** (2018) 124.
- [65] G. Song, J. Rong, and S.-J. Sin, *J. High Energy Phys.* **10** (2019) 109.
- [66] X. Ji, Y. Liu, and X.-M. Wu, *Phys. Rev. D* **100**, 126013 (2019).
- [67] M. Baggioli and D. Giataganas, *Phys. Rev. D* **103**, 026009 (2021).
- [68] V. Juričić, I. Salazar Landea, and R. Soto-Garrido, *J. High Energy Phys.* **07** (2020) 052.
- [69] K. Bitaghsir Fadafan, A. O'Bannon, R. Rodgers, and M. Russell, *J. High Energy Phys.* **04** (2021) 162.
- [70] L.-L. Gao, Y. Liu, and H.-D. Lyu, *J. High Energy Phys.* **03** (2023) 034.
- [71] Y. Satake, J. Shiogai, G. P. Mazur, S. Kimura, S. Awaji, K. Fujiwara, T. Nojima, K. Nomura, S. Souma, T. Sato, T. Dietl, and A. Tsukazaki, *Phys. Rev. Mater.* **4**, 044202 (2020).
- [72] A. H. Mayo, H. Takahashi, M. S. Bahramy, A. Nomoto, H. Sakai, and S. Ishiwata, *Phys. Rev. X* **12**, 011033 (2022).
- [73] S. Rao, Weyl semi-metals: A short review (2016).
- [74] D. Colladay and V. A. Kostelecky, *Phys. Rev. D* **58**, 116002 (1998).
- [75] A. A. Zyuzin and A. A. Burkov, *Phys. Rev. B* (2011).
- [76] A. H. Castro Neto, F. Guinea, N. M. R. Peres, K. S. Novoselov, and A. K. Geim, *Rev. Mod. Phys.* (2009).
- [77] P. E. C. Ashby and J. P. Carbotte, *Phys. Rev. B* (2013).

- [78] K. Das, S. K. Singh, and A. Agarwal, *Phys. Rev. Res.* **2**, 033511 (2020).
- [79] K. Landsteiner, *Acta Phys. Pol. B* **47**, 2617 (2016).
- [80] H.-Z. Lu, S.-B. Zhang, and S.-Q. Shen, *Phys. Rev. B* **92**, 045203 (2015).
- [81] C.-L. Zhang *et al.*, *Nat. Phys.* **13**, 979 (2017).
- [82] M. Baggioli, B. Padhi, P. W. Phillips, and C. Setty, *J. High Energy Phys.* **07** (2018) 049.
- [83] Q. Li, D. E. Kharzeev, C. Zhang, Y. Huang, I. Pletikosić, A. V. Fedorov, R. D. Zhong, J. A. Schneeloch, G. D. Gu, and T. Valla, *Nat. Phys.* **12**, 550 (2016).
- [84] C. Zeng, S. Nandy, P. Liu, S. Tewari, and Y. Yao, *Phys. Rev. B* **107** L081107 (2023).
- [85] Y.-W. Sun and Q. Yang, *J. High Energy Phys.* **09** (2016) 122.
- [86] M. Baggioli, Y. Liu, and X.-M. Wu, *J. High Energy Phys.* **05** (2023) 221.
- [87] K. Das and A. Agarwal, *Phys. Rev. Res.* **2**, 013088 (2020).
- [88] S. Das, K. Das, and A. Agarwal, *Phys. Rev. B* **108**, 045405 (2023).
- [89] M. N. Chernodub, Y. Ferreiros, A. G. Grushin, K. Landsteiner, and M. A. H. Vozmediano, *Phys. Rep.* **977**, 1 (2022).
- [90] J. Gooth, A. C. Niemann, T. Meng, A. G. Grushin, K. Landsteiner, B. Gotsmann, F. Menges, M. Schmidt, C. Shekhar, V. Süß, R. Hühne, B. Rellinghaus, C. Felser, B. Yan, and K. Nielsch, *Nature (London)* **547**, 324 (2017).
- [91] C. Schindler *et al.*, *Phys. Rev. B* **101**, 125119 (2020).
- [92] D. Vu, W. Zhang, C. Şahin, M. E. Flatté, N. Trivedi, and J. P. Heremans, *Nat. Mater.* **20**, 1525 (2021).
- [93] T. Andrade, [arXiv:1712.00548](https://arxiv.org/abs/1712.00548).
- [94] M. Baggioli, *Applied Holography* (Springer International Publishing, New York, 2019).
- [95] U. Gursoy, E. Kiritsis, L. Mazzanti, and F. Nitti, *J. High Energy Phys.* **05** (2009) 033.

Correction: Support information in the last sentence in the Acknowledgements has been fixed.