

**One vanishing minor in the neutrino mass matrix using trimaximal mixing**Iffat Ara Mazumder<sup>\*</sup> and Rupak Dutta<sup>†</sup>*National Institute of Technology Silchar, Silchar 788010, India*

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We investigate the implications of one vanishing minor in the neutrino mass matrix using trimaximal mixing matrix. In this context, we analyze all six patterns of one vanishing minor zero in the neutrino mass matrix and present correlations of the neutrino oscillation parameters. All the six patterns are found to be phenomenologically viable with the present neutrino oscillation data. We also predict the values of effective Majorana mass, the effective electron antineutrino mass, and the total neutrino mass for all the patterns. The value obtained for the effective neutrino mass is within the reach future neutrinoless double  $\beta$  decay experiments. We also propose a flavor model where such patterns can be generated within the seesaw model.

DOI: [10.1103/PhysRevD.107.115023](https://doi.org/10.1103/PhysRevD.107.115023)**I. INTRODUCTION**

Evidence of neutrino oscillations observed in multitude of experiments confirms that neutrinos mix with each other and have nonzero mass [1]. Neutrino oscillation phenomena can be parametrized in terms of six independent parameters, namely three mixing angles ( $\theta_{13}, \theta_{12}, \theta_{23}$ ), one Dirac  $CP$  violating phase ( $\delta_{CP}$ ), and two mass squared differences ( $\Delta m_{21}^2, \Delta m_{31}^2$ ). Although we have very precise values of the mixing angles and the absolute value of the mass squared differences, there are still some unknowns such as the octant of  $\theta_{23}$ ,  $\delta_{CP}$  and the sign of  $\Delta m_{31}^2$ . There are two possible mass ordering of the neutrino mass spectrum: normal mass ordering (NO)  $m_1 < m_2 \ll m_3$  and inverted mass ordering (IO)  $m_3 \ll m_1 < m_2$  depending on the sign of  $\Delta m_{31}^2$ .

Neutrino oscillation experiments are sensitive only to the mass squared differences. They cannot provide any information regarding the absolute mass scale of neutrinos which is one of the most sought after questions in particle physics today. Knowledge of the absolute mass scale of neutrinos is of great importance not only in particle physics but also in understanding the large scale structure of our universe. Neutrinos possess very tiny mass and unlike all other fermions in the Standard Model (SM), they do not seem to get their mass through Higgs mechanism. Hence, it may, in principle, help shape our understanding of the origin of particle mass which is still one of the most fundamental questions of particle physics. Unlike all other fermions in

the SM, we observe only left handed neutrinos and right handed antineutrinos. We have not found any right handed neutrino and left handed antineutrino so far in experiments. This brings us to the next relevant question whether neutrino is a Dirac particle or a Majorana particle. Neutrino interactions could violate  $CP$  as well which will be crucial in explaining the matter antimatter asymmetry in the universe. Moreover, there could be additional sterile neutrinos.

There are several experimental efforts to find the absolute mass of the neutrino. The  $\beta$  decay experiment performed at KATRIN can, in principle, measure the effective electron antineutrino mass by studying the endpoint region of the  $\beta$  decay spectrum. This is completely model independent determination, i.e., it depends neither on any cosmological models nor on the nature of neutrinos. At present, the improved upper bound on the effective electron antineutrino mass is reported to be  $m_\nu < 0.8$  eV at 90% confidence level. The KATRIN experiment will continue to take data over the next several years and it is expected that the mass sensitivity will reach up to 0.2 eV. Future experiments like Project 8 [2], designed to measure the absolute mass scale of the neutrino, hopes to reach a goal of 40 meV/ $c^2$  neutrino mass sensitivity. Indirectly, one can have information on neutrino mass from cosmological observations. These cosmological observations are sensitive to the total neutrino mass and to the number of neutrino species. There are several results related to the total neutrino mass coming from various cosmological observations. Most of these indirect methods put a limit on the total neutrino mass to be less than 0.2 eV. These results are, however, model dependent. They rely heavily on several cosmological assumptions. Current upper bound on the total neutrino mass is reported by the Planck satellite to be  $\sum m_i < 0.12$  eV at 95% confidence level combining BAO data with CMB data [3]. If KATRIN's mass sensitivity reach up to 0.2 eV in future, it can put severe constraint on several

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cosmological models. Rare double  $\beta$  decay process with two anti neutrinos in the final state is allowed in the SM. In general double beta decay processes are powerful probes of beyond the SM physics. More specifically, if one observes neutrinoless double  $\beta$  decay in experiments, it would confirm that neutrinos are Majorana in nature. One can determine the effective Majorana mass  $M_{ee}$  by studying neutrinoless double beta decay. There exists several limits on the value of  $M_{ee}$  using different isotopes. At present, the best limits are reported to be  $m_{ee} < (0.079-0.18)$  eV,  $m_{ee} < (0.075-0.35)$  eV and  $m_{ee} < (0.061-0.165)$  eV [4–7], respectively.

There are several theoretical efforts in explaining the origin of neutrino mass. The most natural way to understand neutrino mass is through seesaw mechanism. The neutrino mass matrix in the framework of type-I seesaw mechanism is given by  $M_\nu = -M_D M_R^{-1} M_D^T$ , where  $M_D$  is the Dirac neutrino mass matrix and  $M_R$  is the Majorana mass matrix of the right-handed neutrinos. Phenomenology of the Majorana neutrino mass matrix has been studied extensively assuming zero textures of the neutrino mass matrix which may be realized from the zeros in  $M_D$  or  $M_R$ . In literature, there have been phenomenological studies with texture one-zero [8–10], two-zeros [11–24] and more within the context of Pontecorvo Maki Nakagawa Sakata (PMNS), tribimaximal (TB), and trimaximal (TM) mixing matrix. Similarly, in Refs. [25–30], the authors have studied the phenomenological implication of vanishing minors in the neutrino mass matrix. Moreover, in Refs. [31–36] and Refs. [37–41], the authors have explored the implication of cofactor zero and hybrid texture of the neutrino mass matrix. In case of zero textures, it is found that three or more zeros in the neutrino mass matrix cannot accommodate the current neutrino oscillation data. In Ref. [11], the authors have found that out of fifteen possible two texture zeros cases only seven cases with PMNS mixing are allowed experimentally. Also out of fifteen possible two cofactors zero patterns only seven patterns are acceptable [25]. In case of TM mixing along with magic symmetry [22], the authors have found that only two cases are valid for two texture zero. TM mixing with one texture zero was studied in Ref. [10] and found that all six patterns are compatible with current neutrino oscillation data. For TB mixing along with the condition of texture zeros or vanishing minor [28] only five patterns are allowed. In this work we study the implication of one vanishing minor in the neutrino mass matrix using trimaximal mixing.

Our paper is organized as follows. In Sec. II, we briefly discuss the neutrino mass matrix using trimaximal mixing matrix. We find all the mixing parameters such as  $\theta_{13}$ ,  $\theta_{23}$ ,  $\theta_{12}$  and the  $CP$  violating parameter  $\delta_{CP}$  in terms of the unknown parameters  $\theta$  and  $\phi$  of the trimaximal mixing matrix. In Sec. III, we describe the formalism of one vanishing minor in the neutrino mass matrix and identify all the possible patterns of one vanishing minor. We provide all the detail numerical analysis and discussion of each

pattern in Secs. IV and V. The fine-tuning of the neutrino mass matrix is presented in Sec. VI. In Sec. VII, we present the symmetry realization and conclude in Sec. VIII.

## II. NEUTRINO MASS MATRIX

The most widely studied lepton flavor mixing is TB mixing pattern [42–45] introduced by Harrison, Perkins, and Scott. TB mixing pattern provides remarkable agreement with the atmospheric and solar neutrino oscillation data. The TB mixing pattern is given by

$$U_{TB} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \sqrt{\frac{1}{3}} & 0 \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & \sqrt{\frac{1}{2}} \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & -\sqrt{\frac{1}{2}} \end{pmatrix}. \quad (1)$$

The TB mixing matrix possesses two types of symmetries:  $\mu - \tau$  symmetry and magic symmetry. Although TB mixing matrix correctly predicted the value of atmospheric mixing angle  $\theta_{23}$  and the solar mixing angle  $\theta_{12}$ , it, however, failed to explain a nonzero value of the reactor mixing angle  $\theta_{13}$  that was experimentally confirmed by T2K [46], MINOS [47], Double Chooz [48], Daya Bay [49] and RENO [50] experiments. The possibility of an exact  $\mu - \tau$  symmetry in the mass matrix was completely ruled out by a relatively large value of  $\theta_{13}$ . Modifications in the TB mixing pattern [51–53] was made to accommodate the present data. The TM mixing matrix was constructed by multiplying the TB mixing matrix by an unitary matrix and can be written as

$$U_{TM_1} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} \cos \theta & \frac{1}{\sqrt{3}} \sin \theta \\ -\frac{1}{\sqrt{6}} & \frac{\cos \theta}{\sqrt{3}} - \frac{e^{i\phi} \sin \theta}{\sqrt{2}} & \frac{\sin \theta}{\sqrt{3}} + \frac{e^{i\phi} \cos \theta}{\sqrt{2}} \\ -\frac{1}{\sqrt{6}} & \frac{\cos \theta}{\sqrt{3}} + \frac{e^{i\phi} \sin \theta}{\sqrt{2}} & \frac{\sin \theta}{\sqrt{3}} - \frac{e^{i\phi} \cos \theta}{\sqrt{2}} \end{pmatrix}. \quad (2)$$

and

$$U_{TM_2} = \begin{pmatrix} \sqrt{\frac{2}{3}} \cos \theta & \frac{1}{\sqrt{3}} & \sqrt{\frac{2}{3}} \sin \theta \\ -\frac{\cos \theta}{\sqrt{6}} + \frac{e^{-i\phi} \sin \theta}{\sqrt{2}} & \frac{1}{\sqrt{3}} & -\frac{\sin \theta}{\sqrt{6}} - \frac{e^{-i\phi} \cos \theta}{\sqrt{2}} \\ -\frac{\cos \theta}{\sqrt{6}} - \frac{e^{-i\phi} \sin \theta}{\sqrt{2}} & \frac{1}{\sqrt{3}} & -\frac{\sin \theta}{\sqrt{6}} + \frac{e^{-i\phi} \cos \theta}{\sqrt{2}} \end{pmatrix}. \quad (3)$$

where  $\theta$  and  $\phi$  are two free parameters. The neutrino mass matrix corresponding to TM mixing matrix can be written as

$$M_{\rho\sigma} = (VM_{\text{diag}}V^T)_{\rho\sigma} \quad \text{with} \quad \rho, \sigma = e, \mu, \tau, \quad (4)$$

where  $M_{\text{diag}} = \text{diag}(m_1, m_2, m_3)$  is the diagonal matrix containing three mass state,  $V = U_{TM}P$  and  $P$  is the phase matrix written as

$$P = \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha} & 0 \\ 0 & 0 & e^{i\beta} \end{pmatrix}. \quad (5)$$

Here  $\alpha$  and  $\beta$  are the two  $CP$  violating Majorana phases.

### A. $TM_1$ mixing matrix

With  $TM_1$  mixing matrix, the elements of the neutrino mass matrix can be written as

$$\begin{aligned} M_{ee} &= \frac{2}{3}m_1 + \frac{1}{3}\cos^2\theta m_2 e^{2i\alpha} + \frac{1}{3}\sin^2\theta m_3 e^{2i\beta}, \\ M_{e\mu} &= \left(-\frac{1}{3}\right)m_1 + \left(\frac{1}{3}\cos^2\theta - \frac{1}{\sqrt{6}}\sin\theta\cos\theta e^{i\phi}\right)m_2 e^{2i\alpha} + \left(\frac{1}{3}\sin^2\theta + \frac{1}{\sqrt{6}}\sin\theta\cos\theta e^{i\phi}\right)m_3 e^{2i\beta}, \\ M_{e\tau} &= \left(-\frac{1}{3}\right)m_1 + \left(\frac{1}{3}\cos^2\theta + \frac{1}{\sqrt{6}}\sin\theta\cos\theta e^{i\phi}\right)m_2 e^{2i\alpha} + \left(\frac{1}{3}\sin^2\theta - \frac{1}{\sqrt{6}}\sin\theta\cos\theta e^{i\phi}\right)m_3 e^{2i\beta}, \\ M_{\mu\mu} &= \frac{1}{6}m_1 + \left(\frac{1}{\sqrt{3}}\cos\theta - \frac{1}{\sqrt{2}}\sin\theta e^{i\phi}\right)^2 m_2 e^{2i\alpha} + \left(\frac{1}{\sqrt{3}}\sin\theta + \frac{1}{\sqrt{2}}\cos\theta e^{i\phi}\right)^2 m_3 e^{2i\beta}, \\ M_{\mu\tau} &= \frac{1}{6}m_1 + \left(\frac{1}{3}\cos^2\theta - \frac{1}{2}\sin^2\theta e^{2i\phi}\right)m_2 e^{2i\alpha} + \left(\frac{1}{3}\sin^2\theta - \frac{1}{2}\cos^2\theta e^{2i\phi}\right)m_3 e^{2i\beta}, \\ M_{\tau\tau} &= \frac{1}{6}m_1 + \left(\frac{1}{\sqrt{3}}\cos\theta + \frac{1}{\sqrt{2}}\sin\theta e^{i\phi}\right)^2 m_2 e^{2i\alpha} + \left(\frac{1}{\sqrt{3}}\sin\theta - \frac{1}{\sqrt{2}}\cos\theta e^{i\phi}\right)^2 m_3 e^{2i\beta}. \end{aligned} \quad (6)$$

The three neutrino mixing angles  $\theta_{12}$ ,  $\theta_{23}$ , and  $\theta_{13}$  can be expressed in terms of  $\theta$  and  $\phi$ , the free parameters of the  $TM_1$  matrix, as

$$\begin{aligned} s_{12}^2 &= \frac{|(U_{12})_{TM_1}|^2}{1 - |(U_{13})_{TM_1}|^2} = 1 - \frac{2}{3 - \sin^2\theta}, \\ s_{23}^2 &= \frac{|(U_{23})_{TM_1}|^2}{1 - |(U_{13})_{TM_1}|^2} = \frac{1}{2} \left(1 + \frac{\sqrt{6}\sin 2\theta \cos\phi}{3 - \sin^2\theta}\right), \\ s_{13}^2 &= |(U_{13})_{TM_1}|^2 = \frac{1}{3}\sin^2\theta, \end{aligned} \quad (7)$$

where  $s_{ij} = \sin\theta_{ij}$  and  $c_{ij} = \cos\theta_{ij}$  for  $i, j = 1, 2, 3$ . Using the standard parametrization of the PMNS matrix, the Jarlskog invariant, a measure of  $CP$  violation, is defined as [54]

$$J = s_{12}s_{13}s_{23}c_{12}c_{13}^2c_{23}\sin\delta. \quad (8)$$

Again, using the elements from  $TM_1$  mixing matrix, the Jarlskog invariant can be expressed as

$$J = \frac{1}{6\sqrt{6}}\sin 2\theta \sin\phi. \quad (9)$$

Combining Eqs. (8) and (9), we can write  $\delta$  in terms of  $\theta$  and  $\phi$  as

$$\csc^2\delta = \csc^2\phi - \frac{6\sin^2 2\theta \cot^2\phi}{(3 - \sin^2\theta)^2}. \quad (10)$$

The nature of neutrino can be determined from the effective Majorana mass term. It also measures the rate of neutrinoless double beta decay. The effective Majorana mass  $|M_{ee}|$  for the  $TM_1$  mixing matrix can be written as

$$|M_{ee}| = \left| \frac{1}{3}(2m_1 + m_2\cos^2\theta e^{2i\alpha} + m_3\sin^2\theta e^{2i\beta}) \right|. \quad (11)$$

Similarly, the effective electron antineutrino mass can be expressed as

$$M_\nu^2 = \sum_{i=1}^3 U_{ie}^2 = \frac{1}{3}(2m_1^2 + m_2^2\cos^2\theta + m_3^2\sin^2\theta). \quad (12)$$

### B. $TM_2$ mixing matrix

Using  $TM_2$  mixing matrix, we can write the elements of the neutrino mass matrix as

$$\begin{aligned}
M_{ee} &= \left(\frac{2}{3}\cos^2\theta\right)m_1 + \frac{1}{3}m_2e^{2i\alpha} + \left(\frac{2}{3}\sin^2\theta\right)m_3e^{2i\beta}, \\
M_{e\mu} &= \left(-\frac{1}{3}\cos^2\theta + \frac{1}{\sqrt{3}}\sin\theta\cos\theta e^{-i\phi}\right)m_1 + \frac{1}{3}m_2e^{2i\alpha} + \left(-\frac{1}{3}\sin^2\theta - \frac{1}{\sqrt{3}}\sin\theta\cos\theta e^{-i\phi}\right)m_3e^{2i\beta}, \\
M_{e\tau} &= \left(-\frac{1}{3}\cos^2\theta - \frac{1}{\sqrt{3}}\sin\theta\cos\theta e^{-i\phi}\right)m_1 + \frac{1}{3}m_2e^{2i\alpha} + \left(-\frac{1}{3}\sin^2\theta + \frac{1}{\sqrt{3}}\sin\theta\cos\theta e^{-i\phi}\right)m_3e^{2i\beta}, \\
M_{\mu\mu} &= \left(-\frac{1}{\sqrt{6}}\cos\theta + \frac{1}{\sqrt{2}}\sin\theta e^{-i\phi}\right)^2 m_1 + \frac{1}{3}m_2e^{2i\alpha} + \left(\frac{1}{\sqrt{6}}\sin\theta + \frac{1}{\sqrt{2}}\cos\theta e^{-i\phi}\right)^2 m_3e^{2i\beta}, \\
M_{\mu\tau} &= \left(\frac{1}{6}\cos^2\theta - \frac{1}{2}\sin^2\theta e^{-2i\phi}\right)m_1 + \frac{1}{3}m_2e^{2i\alpha} + \left(\frac{1}{6}\sin^2\theta - \frac{1}{2}\cos^2\theta e^{-2i\phi}\right)m_3e^{2i\beta}, \\
M_{\tau\tau} &= \left(\frac{1}{\sqrt{6}}\cos\theta + \frac{1}{\sqrt{2}}\sin\theta e^{-i\phi}\right)^2 m_1 + \frac{1}{3}m_2e^{2i\alpha} + \left(-\frac{1}{\sqrt{6}}\sin\theta + \frac{1}{\sqrt{2}}\cos\theta e^{-i\phi}\right)^2 m_3e^{2i\beta}. \tag{13}
\end{aligned}$$

The three neutrino mixing angles  $\theta_{12}$ ,  $\theta_{23}$ , and  $\theta_{13}$  can be expressed as

$$\begin{aligned}
s_{12}^2 &= \frac{1}{3 - 2\sin^2\theta}, \\
s_{23}^2 &= \frac{1}{2} \left(1 + \frac{\sqrt{3}\sin 2\theta \cos \phi}{3 - 2\sin^2\theta}\right), \\
s_{13}^2 &= \frac{2}{3}\sin^2\theta. \tag{14}
\end{aligned}$$

Again, using the elements from the  $\text{TM}_2$  mixing matrix, the Jarlskog invariant can be expressed as

$$J = \frac{1}{6\sqrt{3}}\sin 2\theta \sin \phi. \tag{15}$$

We can express the Dirac  $CP$  violating parameter  $\delta_{CP}$  in terms of  $\theta$  and  $\phi$  as

$$\csc^2 \delta = \csc^2 \phi - \frac{3\sin^2 2\theta \cot^2 \phi}{(3 - 2\sin^2\theta)^2}. \tag{16}$$

The effective Majorana mass  $|M_{ee}|$  for the  $\text{TM}_2$  mixing matrix can be written as

$$|M_{ee}| = \left|\frac{1}{3}(2m_1\cos^2\theta + m_2e^{2i\alpha} + 2m_3\sin^2\theta e^{2i\beta})\right|. \tag{17}$$

The effective electron antineutrino mass can be expressed as

$$M_\nu^2 = \frac{1}{3}(2m_1^2 + m_2^2\cos^2\theta + 2m_3^2\sin^2\theta). \tag{18}$$

### III. ONE VANISHING MINOR IN NEUTRINO MASS MATRIX

There are six independent minors corresponding to six independent elements in the neutrino mass matrix. We

denote the minor corresponding to  $ij$ th element of  $M_{ij}$  as  $C_{ij}$ . The six possible patterns of one minor zero in the neutrino mass matrix are listed in Table I. The condition for one vanishing minor can be written as

$$M_{ab}M_{cd} - M_{uv}M_{wx} = 0. \tag{19}$$

More specifically, we can write Eq. (19) in terms of a complex equation as

$$m_1m_2X_3e^{2i\alpha} + m_2m_3X_1e^{2i(\alpha+\beta)} + m_3m_1X_2e^{2i\beta} = 0, \tag{20}$$

where

$$X_k = (U_{ai}U_{bi}U_{cj}U_{dj} - U_{ui}U_{vi}U_{wj}U_{xj}) + (i \leftrightarrow j), \tag{21}$$

with  $(i, j, k)$  as the cyclic permutation of  $(1, 2, 3)$ . Using Eq. (20), one can write the two mass ratios as

$$\begin{aligned}
\frac{m_1}{m_2} &= \frac{\Re(X_3e^{2i\alpha})\Im(X_1e^{2i(\alpha+\beta)}) - \Re(X_1e^{2i(\alpha+\beta)})\Im(X_3e^{2i\alpha})}{\Re(X_2e^{2i\beta})\Im(X_3e^{2i\alpha}) - \Re(X_3e^{2i\alpha})\Im(X_2e^{2i\beta})}, \\
\frac{m_3}{m_2} &= \frac{\Re(X_3e^{2i\alpha})\Im(X_1e^{2i(\alpha+\beta)}) - \Re(X_1e^{2i(\alpha+\beta)})\Im(X_3e^{2i\alpha})}{\Re(X_1e^{2i(\alpha+\beta)})\Im(X_2e^{2i\beta}) - \Re(X_2e^{2i\beta})\Im(X_1e^{2i(\alpha+\beta)})}. \tag{22}
\end{aligned}$$

TABLE I. One minor zero patterns.

Pattern	Constraining equation
I	$C_{33} = 0$
II	$C_{22} = 0$
III	$C_{31} = 0$
IV	$C_{21} = 0$
V	$C_{32} = 0$
VI	$C_{11} = 0$

The value of  $m_1$ ,  $m_2$ , and  $m_3$  can be calculated using Eq. (22) and mass square difference  $\Delta m_{21}^2$ . That is

$$\begin{aligned} m_1 &= \sqrt{\Delta m_{21}^2} \sqrt{\frac{\left(\frac{m_1}{m_2}\right)^2}{\left|1 - \left(\frac{m_1}{m_2}\right)^2\right|}}, \\ m_2 &= \sqrt{\Delta m_{21}^2} \sqrt{\frac{1}{\left|1 - \left(\frac{m_1}{m_2}\right)^2\right|}}, \\ m_3 &= \sqrt{\Delta m_{21}^2} \sqrt{\frac{\left(\frac{m_3}{m_2}\right)^2}{\left|1 - \left(\frac{m_1}{m_2}\right)^2\right|}}. \end{aligned} \quad (23)$$

Similarly, the ratio of squared mass difference is defined as

$$r \equiv \left| \frac{\Delta m_{21}^2}{\Delta m_{32}^2} \right| = \left| \frac{1 - \left(\frac{m_1}{m_2}\right)^2}{\left(\frac{m_3}{m_2}\right)^2 - 1} \right|, \quad (24)$$

where  $\Delta m_{21}^2$  and  $\Delta m_{32}^2$  represent solar and atmospheric mass squared difference, respectively. Value of  $r = (2.950 \pm 0.08) \times 10^{-2}$  is determined by using the measured values of  $\Delta m_{21}^2$  and  $\Delta m_{32}^2$  reported in Ref. [55].

#### IV. RESULTS AND DISCUSSION

For our numerical analysis, we use the measured values of the oscillation parameters reported in Ref. [55]. For completeness, we report them in Table II. We wish to find the value of the unknown parameters  $\theta$  and  $\phi$ . It is evident from Eqs. (7) and (14) that the neutrino oscillation parameters  $\theta_{12}$  and  $\theta_{13}$  depend only on  $\theta$ . To find the best fit value of  $\theta$ , we perform a naive  $\chi^2$  analysis. The relevant  $\chi^2$  is defined as

$$\chi^2(\theta) = \sum_{i=1}^2 \frac{(\theta_i^{\text{cal}} - \theta_i^{\text{exp}})^2}{(\sigma_i^{\text{exp}})^2}, \quad (25)$$

where  $\theta_i = (\theta_{12}, \theta_{13})$ . Here  $\theta_i^{\text{cal}}$  represents the theoretical value of  $\theta_i$  and  $\theta_i^{\text{exp}}$  represents measured central value of  $\theta_i$ . The corresponding uncertainties in the measured value of  $\theta_i$  is represented by  $\sigma_i^{\text{exp}}$ .

For the  $\text{TM}_1$  mixing matrix, the best fit value of  $\theta$  is obtained to be  $14.96^\circ$ . The corresponding best fit values of  $\theta_{12}$  and  $\theta_{13}$  are  $34.33^\circ$  and  $8.57^\circ$ , respectively. The  $3\sigma$  allowed range of  $\theta$  is found to be  $(14.26^\circ - 15.64^\circ)$ . Using the allowed range of  $\theta$ , we obtain the allowed ranges of  $\theta_{12}$  and  $\theta_{23}$  to be  $(34.25^\circ - 34.42^\circ)$  and  $(32.11^\circ - 57.88^\circ)$ , respectively. We show in Fig. 1(a) the correlation of  $\theta_{13}$  and  $\theta_{12}$  for the  $\text{TM}_1$  mixing matrix. To see the variation of  $\theta_{23}$  with  $\phi$ , we use the allowed range of  $\theta$  and vary  $\phi$  within its full range from  $0^\circ$  to  $360^\circ$ . We show in Fig. 1(b) the variation of  $\theta_{23}$  as a function of the unknown parameter  $\phi$ . We also obtain the best fit value of  $\phi$  by using the measured best fit value of  $\theta_{23}$ . The best fit value is shown with a \* mark in Fig. 1(b). The best fit values of  $\phi$  corresponding to the best fit value of  $\theta_{23} = 49.2^\circ$  are  $69.43^\circ$  and  $290.57^\circ$ , respectively. We show the variation of  $J$  and  $\delta$  as a function of  $\phi$  in Figs. 1(c) and 1(d), respectively. It is observed that the Jarlskog rephasing invariant  $J$  and the Dirac  $CP$  violating phase  $\delta$  are restricted to two regions. The corresponding best fit values of  $J$  and  $\delta$  are  $[-3.184 \times 10^{-2}, 3.185 \times 10^{-2}]$  and  $[71.11^\circ, 288.98^\circ]$ , respectively. We also obtain the  $3\sigma$  allowed ranges of  $J$  and  $\delta$  to be  $[0, \pm 3.53 \times 10^{-2}]$  and  $[(55.05, 124.95)^\circ, (235.05, 304.95)^\circ]$ , respectively.

For  $\text{TM}_2$  mixing matrix, the best fit value of  $\theta$  is obtained to be  $10.50^\circ$ . The corresponding best fit values of  $\theta_{12}$  and  $\theta_{13}$  are  $35.72^\circ$  and  $8.56^\circ$ , respectively. The  $3\sigma$  allowed range of  $\theta$  is found to be  $(10.03^\circ - 10.99^\circ)$ . Using the allowed range of  $\theta$ , we obtain the allowed ranges of  $\theta_{12}$  and  $\theta_{23}$  to be  $(35.68^\circ - 35.76^\circ)$  and  $(39.50^\circ - 51.40^\circ)$ , respectively. It should be noted that although the allowed range of  $\theta_{12}$  obtained with  $\text{TM}_2$  mixing matrix is consistent with the  $3\sigma$  experimental range, the best fit value obtained for  $\theta_{12}$ , however, deviates from the experimental best fit value at more than  $2\sigma$  significance. This is quite a generic feature of  $\text{TM}_2$  mixing matrix because, by default, value of  $\theta_{12}$  will be greater than or equal to the value obtained in case of TB mixing matrix. We show in Fig. 2(a) the correlation of  $\theta_{13}$  and  $\theta_{12}$  for the  $\text{TM}_2$  mixing matrix.

TABLE II. Neutrino oscillation parameters from NuFIT [55].

Parameter	Normal ordering (best fit)		Inverted ordering ( $\Delta\chi^2 = 7.1$ )	
	bfp $\pm 1\sigma$	$3\sigma$ ranges	bfp $\pm 1\sigma$	$3\sigma$ ranges
$\theta_{12}^\circ$	$33.44^{+0.77}_{-0.74}$	$31.27 \rightarrow 35.86$	$33.45^{+0.77}_{-0.74}$	$31.27 \rightarrow 35.87$
$\theta_{23}^\circ$	$49.2^{+1.0}_{-1.3}$	$39.5 \rightarrow 52.0$	$49.5^{+1.0}_{-1.2}$	$39.8 \rightarrow 52.1$
$\theta_{13}^\circ$	$8.57^{+0.13}_{-0.12}$	$8.20 \rightarrow 8.97$	$8.60^{+0.12}_{-0.12}$	$8.24 \rightarrow 8.98$
$\delta^\circ$	$194^{+52}_{-25}$	$105 \rightarrow 405$	$287^{+27}_{-32}$	$192 \rightarrow 361$
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$
$\frac{\Delta m_{3l}^2}{10^{-3} \text{ eV}^2}$	$+2.515^{+0.028}_{-0.028}$	$+2.431 \rightarrow +2.599$	$-2.498^{+0.028}_{-0.029}$	$-2.584 \rightarrow -2.413$



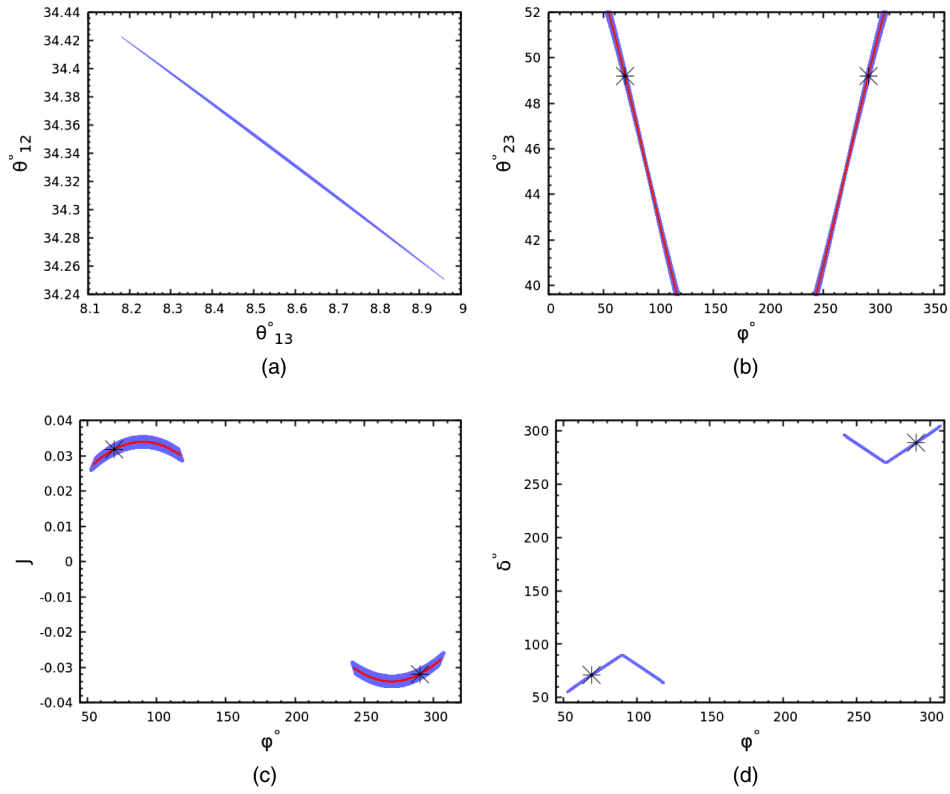


FIG. 1. (a) Correlation of  $\theta_{13}$  and  $\theta_{12}$ , (b) variation of  $\theta_{23}$  as a function of  $\phi$ , (c) variation of  $J$  as a function of  $\phi$ , and (d) variation of  $\delta$  as a function of  $\phi$  for  $TM_1$  mixing matrix.

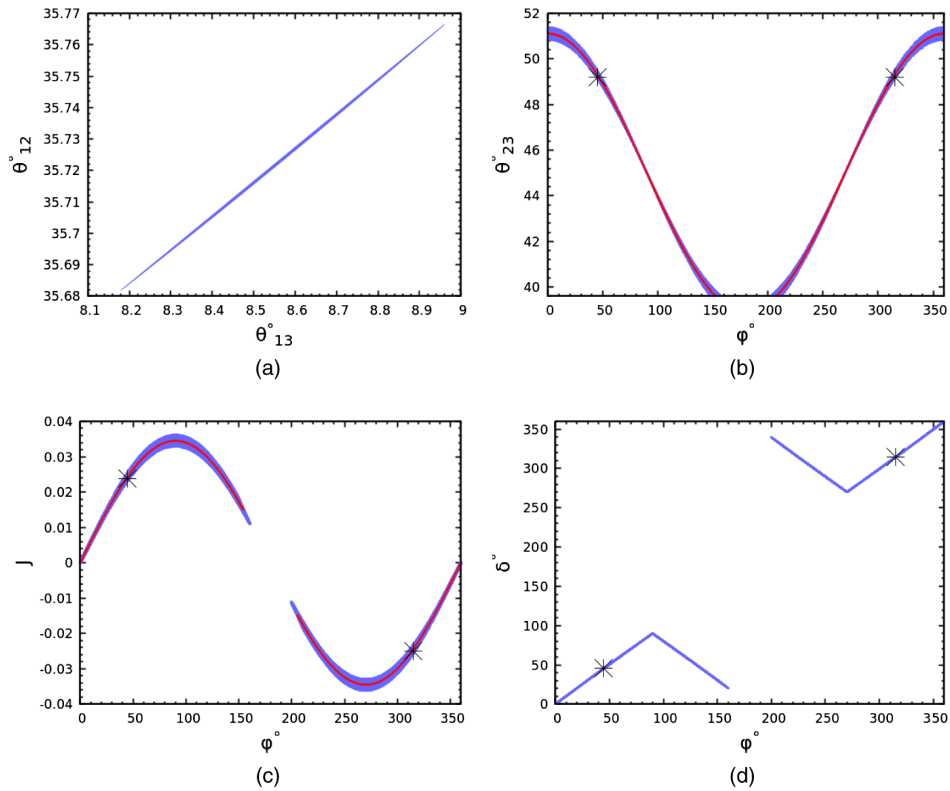


FIG. 2. (a) Correlation of  $\theta_{13}$  and  $\theta_{12}$ , (b) variation of  $\theta_{23}$  as a function of  $\phi$ , (c) variation of  $J$  as a function of  $\phi$ , and (d) variation of  $\delta$  as a function of  $\phi$  for  $TM_2$  mixing matrix.

To see the variation of  $\theta_{23}$  with  $\phi$ , we use the allowed range of  $\theta$  and vary  $\phi$  within its full range from  $0^\circ$  to  $360^\circ$ . We show in Fig. 2(b) the variation of  $\theta_{23}$  as a function of the unknown parameter  $\phi$ . The best fit value is shown with a \* mark in Fig. 2(b). We obtain the best fit value of  $\phi$  by using the measured best fit value of  $\theta_{23}$ . The best fit values of  $\phi$  corresponding to the best fit value of  $\theta_{23} = 49.2^\circ$  are  $44.86^\circ$  and  $315.19^\circ$ , respectively. We get two best fit values of  $\phi$  because  $\theta_{23}$  is invariant under the transformation  $\phi \rightarrow (2\pi - \phi)$  which is evident from Eqs. (7) and (14). We also show the variation of  $J$  and  $\delta$  as a function of  $\phi$  in Figs. 2(c) and 2(d), respectively. It is observed that the Jarlskog rephasing invariant  $J$  and the Dirac  $CP$  violating phase  $\delta$  are restricted to two regions. The corresponding best fit values of  $J$  and  $\delta$  are  $[2.37 \times 10^{-2}, -2.50 \times 10^{-2}]$  and  $[45.48^\circ, 314.44^\circ]$ , respectively. We also obtain the  $3\sigma$  allowed ranges of  $J$  and  $\delta$  to be  $[0, \pm 3.60 \times 10^{-2}]$  and  $[(0, 90)^\circ, (270, 360)^\circ]$ , respectively.

In case of inverted mass ordering the  $3\sigma$  allowed range of  $\theta$  are found to be  $(14.37^\circ - 15.64^\circ)$  and  $(10.10^\circ - 10.99^\circ)$  for both  $TM_1$  and  $TM_2$  mixing matrix, respectively. Using the  $3\sigma$  allowed range of  $\theta$ , we obtain the  $3\sigma$  allowed ranges of  $\theta_{12}$  to be  $(34.25^\circ - 34.41^\circ)$  and  $(35.68^\circ - 35.76^\circ)$ , respectively for both  $TM_1$  and  $TM_2$  mixing matrix. The  $3\sigma$  allowed ranges of  $\theta_{23}$  to be  $(32.12^\circ - 57.87^\circ)$  and  $(38.60^\circ - 51.39^\circ)$ , respectively for both  $TM_1$  and  $TM_2$  mixing matrix. It is to be noted that the mixing angles are almost similar for both the normal and inverted mass ordering. So for our later discussion we will use the values of mixing angles for the normal mass ordering reported in Ref. [55].

## V. PHENOMENOLOGY OF ONE VANISHING MINOR

We wish to investigate the phenomenological implication of one vanishing minor in the neutrino mass matrix on the total neutrino mass, the effective Majorana mass term and the electron antineutrino mass. It is evident from Eq. (23) that neutrino mass  $m_i$  depends on  $\theta$ ,  $\phi$ ,  $\alpha$ ,  $\beta$ , and the mass squared difference  $\Delta m_{21}^2$ . We use the best fit value and the

$3\sigma$  allowed range of  $\theta$  and  $\phi$  of Sec. IV that are determined by the measured values of the mixing angles  $\theta_{13}$ ,  $\theta_{12}$ , and  $\theta_{23}$ . The two unknown Majorana phases  $\alpha$  and  $\beta$  are varied within their full range from  $0^\circ$  to  $360^\circ$ . Moreover, we use the  $3\sigma$  allowed ranges of  $\Delta m_{21}^2$  and  $r$  to constrain the values of the neutrino masses. Now we proceed to analyze all the six patterns of one vanishing minor one by one.

### A. Pattern I: $C_{33} = 0$

Let us first consider minor zero for the (3,3) element of the neutrino mass matrix. The equation corresponding to this pattern can be expressed in terms of the elements of the neutrino mass matrix as

$$(M_\nu)_{ee}(M_\nu)_{\mu\mu} - (M_\nu)_{e\mu}(M_\nu)_{e\mu} = 0. \quad (26)$$

Using Eq. (22), the two mass ratios for  $TM_1$  can be expressed as

$$\frac{m_1}{m_2} = \frac{\mathcal{A}_1 \sin 2\beta + \mathcal{A}_2 \cos 2\beta}{(\mathcal{A}_3 + \mathcal{A}_4) \sin 2(\alpha - \beta) + (\mathcal{A}_5 - \mathcal{A}_6) \cos 2(\alpha - \beta)},$$

$$\frac{m_3}{m_2} = \frac{\mathcal{A}_1 \sin 2\beta + \mathcal{A}_2 \cos 2\beta}{\mathcal{A}_7 \sin 2\alpha + \mathcal{A}_8 \cos 2\alpha}, \quad (27)$$

Similarly, for  $TM_2$  mixing matrix, the mass ratios can be expressed as

$$\frac{m_1}{m_2} = \frac{(\tilde{\mathcal{A}}_1 + \tilde{\mathcal{A}}_2) \sin 2(\beta - \phi) + (\tilde{\mathcal{A}}_3 + \tilde{\mathcal{A}}_4) \cos 2(\beta - \phi)}{\tilde{\mathcal{A}}_5 \sin 2(\alpha - \beta) - \tilde{\mathcal{A}}_6 \cos 2(\alpha - \beta)},$$

$$\frac{m_3}{m_2} = \frac{(\tilde{\mathcal{A}}_1 + \tilde{\mathcal{A}}_2) \sin 2(\beta - \phi) + (\tilde{\mathcal{A}}_3 + \tilde{\mathcal{A}}_4) \cos 2(\beta - \phi)}{\tilde{\mathcal{A}}_7 \sin 2(\phi - \alpha) - \tilde{\mathcal{A}}_8 \cos 2(\phi - \alpha)}. \quad (28)$$

All the relevant expressions for  $\mathcal{A}_i$  and  $\tilde{\mathcal{A}}_i$  are reported in Eqs. (A1) and (A6) of appendix, respectively. We show the variation of neutrino masses  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$  in Figs. 3(a) and 4(a) for  $TM_1$  and  $TM_2$  mixing matrix,

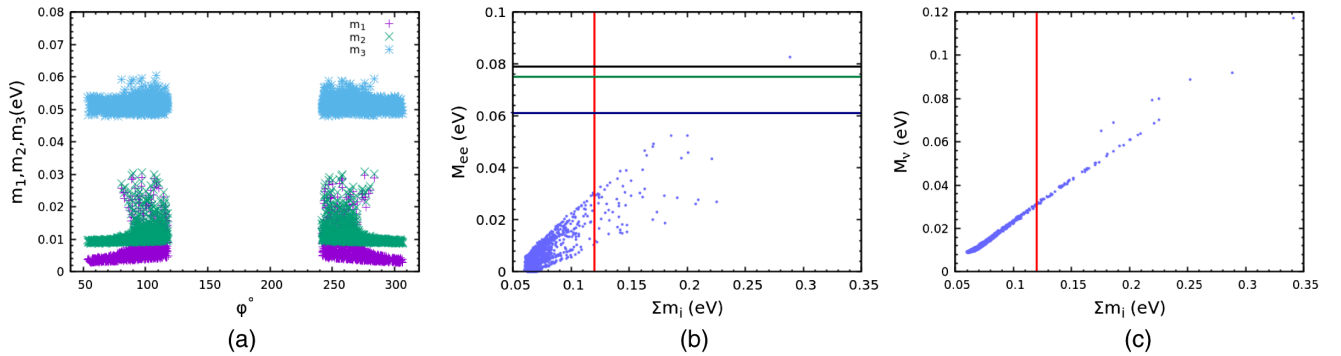


FIG. 3. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{33}$  pattern using  $TM_1$  mixing matrix. The vertical red line is the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines are the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

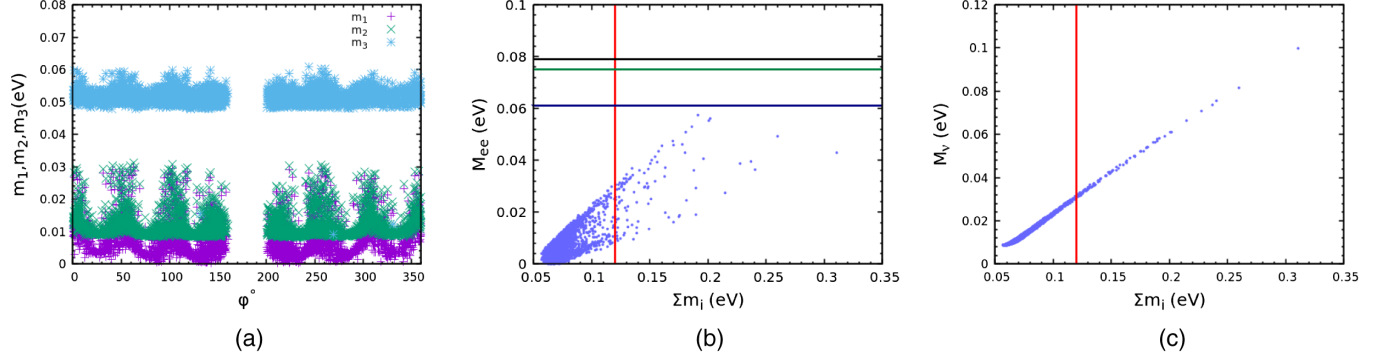


FIG. 4. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{33}$  pattern using  $TM_2$  mixing matrix. The vertical red line represents the upper bound of  $\sum m_i$  reported in Ref. [3]. The black, green, and blue lines are the experimental upper bounds of  $M_{ee}$  obtained from Refs. [4–7].

respectively. It shows normal mass ordering for  $TM_1$  mixing matrix while for  $TM_2$  mixing matrix, it shows both normal and inverted mass ordering. The correlation of  $M_{ee}$  and  $\sum m_i$  for  $TM_1$  and  $TM_2$  mixing matrix are shown in Figs. 3(b) and 4(b), respectively. The vertical red line shows the upper bound of the total neutrino mass reported in Ref. [3]. The black, green and blue lines are the experimental upper bounds of the effective Majorana mass as reported in Refs. [4–7]. In Figs. 3(c) and 4(c), we have shown the correlation of  $M_\nu$  with  $\sum m_i$  for  $TM_1$  and  $TM_2$  mixing matrix, respectively. It is observed that the total neutrino mass  $\sum m_i$  puts severe constraints on effective Majorana mass and  $M_\nu$ .

The range of the absolute neutrino mass scale, the effective Majorana neutrino mass and the effective electron antineutrino mass obtained for both the mixing matrix are listed in Table III. The calculated upper bound of  $m_{ee}$  is obtained to be of  $\mathcal{O}(10^{-2})$  which is within the reach of neutrinoless double beta decay experiment. The calculated upper bound of  $m_\nu < 0.06$  eV may not be within the reach of KATRIN experiment. It may, however, be within the reach of next generation experiment such as Project 8. It should, however, be mentioned that once the total neutrino mass constraint is imposed, the calculated upper bound of  $m_{ee}$  and  $m_\nu$  is found to be less than 0.04 eV.

### B. Pattern II: $C_{22} = 0$

The vanishing minor condition for this pattern corresponding to element (2,2) is given by

$$(M_\nu)_{ee}(M_\nu)_{\tau\tau} - (M_\nu)_{e\tau}(M_\nu)_{e\tau} = 0. \quad (29)$$

The two mass ratios for  $TM_1$  and  $TM_2$  mixing matrix can be expressed as

$$\frac{m_1}{m_2} = \frac{\mathcal{B}_1 \sin 2\beta + \mathcal{B}_2 \cos 2\beta}{(\mathcal{B}_3 + \mathcal{B}_4) \sin 2(\alpha - \beta) - (\mathcal{B}_5 - \mathcal{B}_6) \cos 2(\alpha - \beta)},$$

$$\frac{m_3}{m_2} = \frac{\mathcal{B}_1 \sin 2\beta + \mathcal{B}_2 \cos 2\beta}{\mathcal{B}_7 \sin 2\alpha + \mathcal{B}_8 \cos 2\alpha}, \quad (30)$$

and

$$\frac{m_1}{m_2} = \frac{(\tilde{\mathcal{B}}_1 + \tilde{\mathcal{B}}_2) \sin 2(\beta - \phi) + (\tilde{\mathcal{B}}_3 + \tilde{\mathcal{B}}_4) \cos 2(\beta - \phi)}{\tilde{\mathcal{B}}_5 \sin 2(\alpha - \beta) - \tilde{\mathcal{B}}_6 \cos 2(\alpha - \beta)},$$

$$\frac{m_3}{m_2} = \frac{(\tilde{\mathcal{B}}_1 + \tilde{\mathcal{B}}_2) \sin 2(\beta - \phi) + (\tilde{\mathcal{B}}_3 + \tilde{\mathcal{B}}_4) \cos 2(\beta - \phi)}{\tilde{\mathcal{B}}_7 \sin 2(\phi - \alpha) - \tilde{\mathcal{B}}_8 \cos 2(\phi - \alpha)}. \quad (31)$$

All the relevant expressions for  $\mathcal{B}_i$  and  $\tilde{\mathcal{B}}_i$  are reported in Eqs. (A2) and (A7) of appendix. We show the variation of neutrino masses  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$  in Figs. 5(a) and 6(a) for  $TM_1$  and  $TM_2$  mixing matrix, respectively. It shows normal mass ordering for  $TM_1$  mixing matrix while for  $TM_2$  mixing matrix, it shows both normal and inverted mass ordering. The correlation of  $M_{ee}$  and  $\sum m_i$  for  $TM_1$  and  $TM_2$  mixing matrix are shown in Figs. 5(b) and 6(b), respectively. In Figs. 5(c) and 6(c), we have shown the correlation of  $M_\nu$  with  $\sum m_i$  for  $TM_1$  and  $TM_2$  mixing matrix, respectively. The phenomenology of this pattern is quite similar to  $C_{33}$ .

The allowed range of the absolute neutrino mass scale, the effective Majorana mass and the effective electron

TABLE III. Allowed range of  $\sum m_i$ ,  $m_{ee}$ , and  $m_\nu$  for  $C_{33}$  pattern.

Mixing matrix	Mass ordering	$\sum m_i$ (eV)	$m_{ee}$ (eV)	$m_\nu$ (eV)
$TM_1$	NO	[0.059, 0.288]	$[1.955 \times 10^{-5}, 0.048]$	$[8.889 \times 10^{-3}, 0.053]$
$TM_2$	NO	[0.056, 0.310]	$[1.462 \times 10^{-5}, 0.057]$	$[8.575 \times 10^{-3}, 0.099]$
	IO	[0.094, 0.425]	[0.014, 0.116]	[0.044, 0.144]



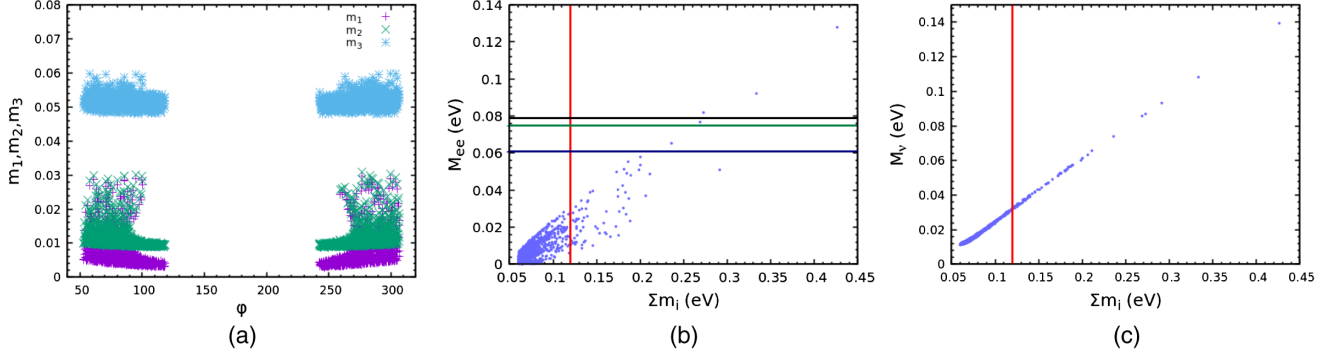


FIG. 5. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{22}$  pattern using  $TM_1$  mixing matrix. The vertical red line shows the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines are the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

antineutrino mass for this pattern are listed in Table IV. It is evident from Figs. 5 and 6 that the upper bound of  $\sum m_i < 1.2$  eV put severe constraint on the value of the effective Majorana mass term  $m_{ee}$  and the value of the effective electron antineutrino mass  $m_\nu$ . The estimated upper bound of  $m_{ee}$  and  $m_\nu$  is found to be less than 0.04 eV.

### C. Pattern III: $C_{31} = 0$

This pattern corresponds to the matrix element (3,1) of the neutrino mass matrix. The vanishing minor condition is given by

$$(M_\nu)_{e\mu}(M_\nu)_{\mu\tau} - (M_\nu)_{e\tau}(M_\nu)_{\mu\mu} = 0. \quad (32)$$

Using the elements from the neutrino mass matrix, one can write the neutrino mass ratios for this pattern. With  $TM_1$  mixing matrix, we have

$$\frac{m_1}{m_2} = \frac{C_1 \sin 2\beta + C_2 \cos 2\beta}{C_3 \sin 2(\alpha - \beta) - C_4 \cos 2(\alpha - \beta)},$$

$$\frac{m_3}{m_2} = \frac{C_1 \sin 2\beta + C_2 \cos 2\beta}{C_5 \sin 2\alpha + C_6 \cos 2\alpha}, \quad (33)$$

and for  $TM_2$  mixing matrix, we have

$$\frac{m_1}{m_2} = \frac{\tilde{C}_1 \sin 2(\beta - \phi) + \tilde{C}_2 \cos 2(\beta - \phi)}{\tilde{C}_3 \sin 2(\alpha - \beta) - \tilde{C}_4 \cos 2(\alpha - \beta)},$$

$$\frac{m_3}{m_2} = \frac{\tilde{C}_1 \sin 2(\beta - \phi) + \tilde{C}_2 \cos 2(\beta - \phi)}{\tilde{C}_5 \sin 2(\phi - \alpha) + \tilde{C}_6 \cos 2(\phi - \alpha)}, \quad (34)$$

where all the relevant  $C_i$  and  $\tilde{C}_i$  are reported in Eqs. (A3) and (A8) of appendix. We show in Figs. 7(a) and 8(a) the correlation of neutrino masses  $m_1$ ,  $m_2$  and  $m_3$  with the unknown parameter  $\phi$  for  $TM_1$  and  $TM_2$  mixing matrix. It shows both normal and inverted mass ordering for  $TM_1$  and  $TM_2$  mixing matrix. Similarly, the correlation of  $M_{ee}$  against  $\sum m_i$  for  $TM_1$  and  $TM_2$  mixing patterns are shown in Figs. 7(b) and 8(b), respectively. Moreover, in the Figs. 7(c) and 8(c), we have shown the correlation of  $M_\nu$  with  $\sum m_i$  for  $TM_1$  and  $TM_2$ , respectively.

The allowed ranges of the absolute neutrino mass, the effective Majorana mass and the effective electron antineutrino mass for both the maxing matrix are listed in the Table V.

### D. Pattern IV: $C_{21} = 0$

The vanishing minor condition for this pattern corresponding to element (2,1) of the neutrino mass matrix is given by

$$(M_\nu)_{\mu e}(M_\nu)_{\tau\tau} - (M_\nu)_{\tau e}(M_\nu)_{\mu\tau} = 0. \quad (35)$$

The two neutrino mass ratios for this pattern for  $TM_1$  and  $TM_2$  mixing matrix are given by

$$\frac{m_1}{m_2} = \frac{\mathcal{D}_1 \sin 2\beta + \mathcal{D}_2 \cos 2\beta}{\mathcal{D}_3 \sin 2(\alpha - \beta) + \mathcal{D}_4 \cos 2(\alpha - \beta)},$$

$$\frac{m_3}{m_2} = \frac{\mathcal{D}_1 \sin 2\beta + \mathcal{D}_2 \cos 2\beta}{\mathcal{D}_5 \sin 2\alpha + \mathcal{D}_6 \cos 2\alpha}, \quad (36)$$

and

TABLE IV. Allowed range of  $\sum m_i$ ,  $m_{ee}$ , and  $m_\nu$  for  $C_{22}$  pattern.

Mixing matrix	Mass ordering	$\sum m_i$ (eV)	$m_{ee}$ (eV)	$m_\nu$ (eV)
$TM_1$	<b>NO</b>	[0.059, 0.425]	$[2.097 \times 10^{-5}, 0.127]$	[0.011, 0.139]
$TM_2$	<b>NO</b>	[0.056, 0.350]	$[4.185 \times 10^{-6}, 0.081]$	$[8.415 \times 10^{-3}, 0.113]$
	<b>IO</b>	[0.098, 0.427]	[0.015, 0.082]	[0.045, 0.145]

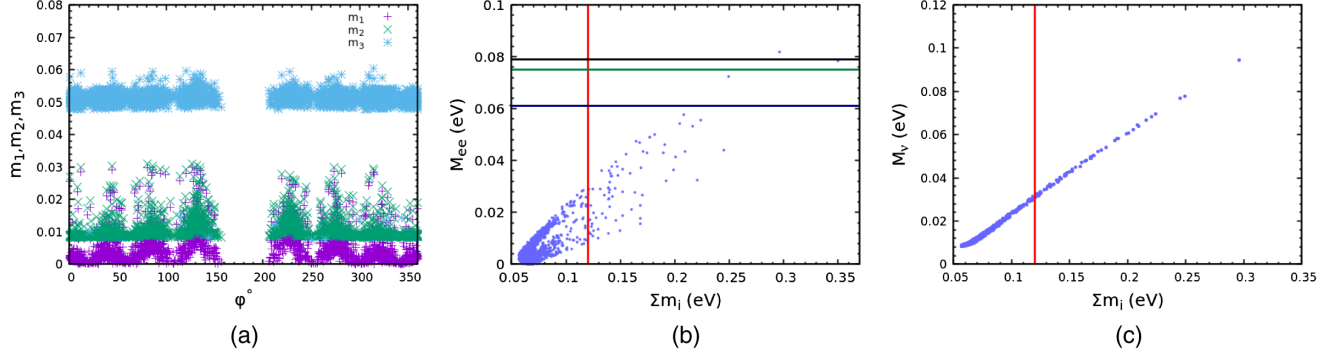


FIG. 6. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{22}$  pattern using  $TM_2$  mixing matrix. The vertical red line shows the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines show the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

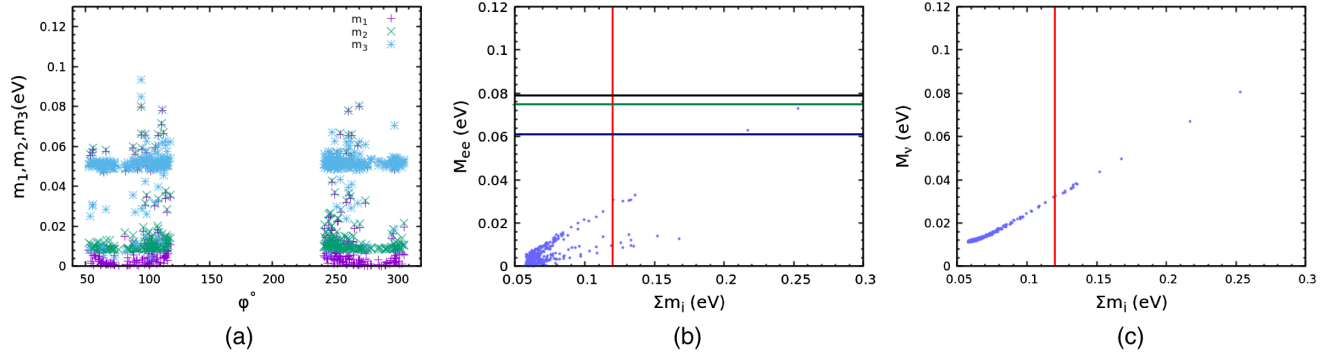


FIG. 7. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{31}$  pattern using  $TM_1$  mixing matrix. The vertical red line is the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines are the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

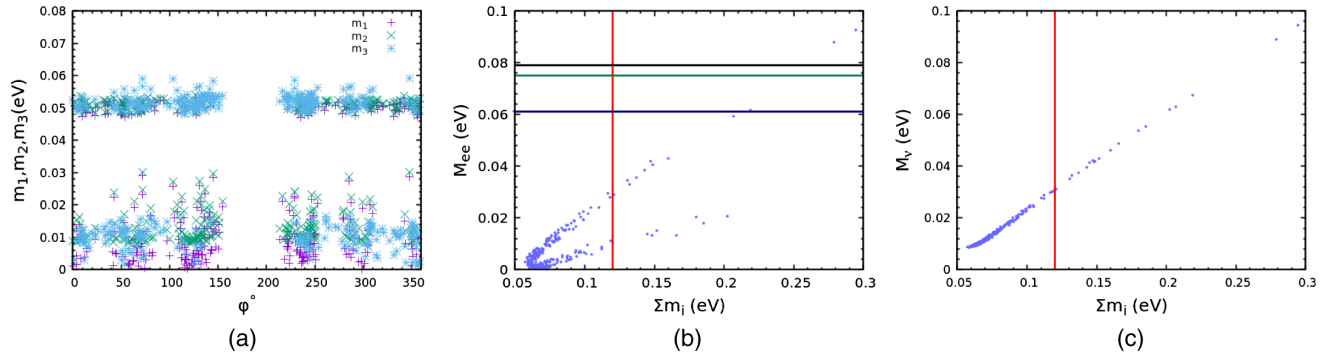


FIG. 8. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{31}$  pattern using  $TM_2$  mixing matrix. The vertical red line represents the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines represent the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

TABLE V. Allowed range of  $\sum m_i$ ,  $m_{ee}$ , and  $m_\nu$  for  $C_{31}$  pattern.

Mixing matrix	Mass ordering	$\sum m_i$ (eV)	$m_{ee}$ (eV)	$m_\nu$ (eV)
TM <sub>1</sub>	NO	[0.057, 0.253]	$[1.003 \times 10^{-4}, 0.073]$	[0.011, 0.080]
	IO	[0.096, 0.329]	[0.014, 0.087]	[0.044, 0.112]
TM <sub>2</sub>	NO	[0.057, 0.294]	$[3.658 \times 10^{-5}, 0.092]$	$[8.624 \times 10^{-3}, 0.094]$
	IO	[0.092, 0.548]	[0.014, 0.126]	[0.045, 0.185]

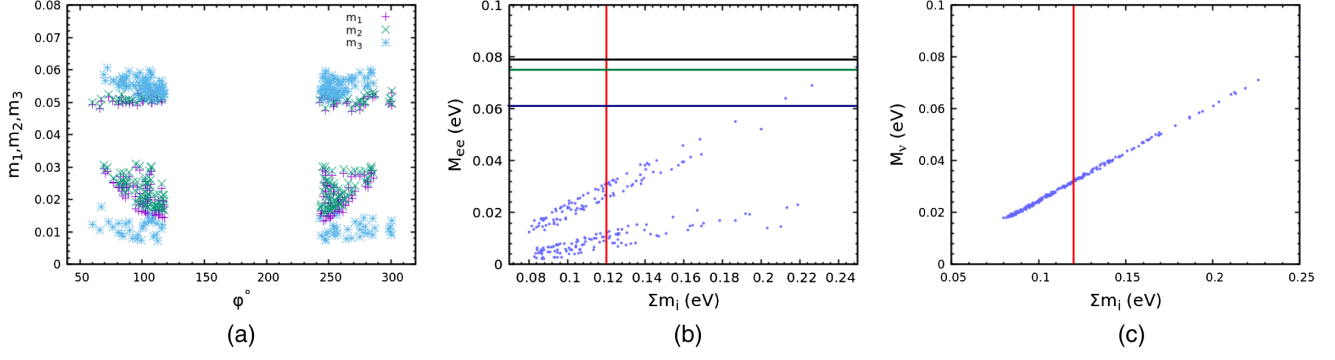


FIG. 9. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{21}$  pattern using  $TM_1$  mixing matrix. The vertical red line is the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines are the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

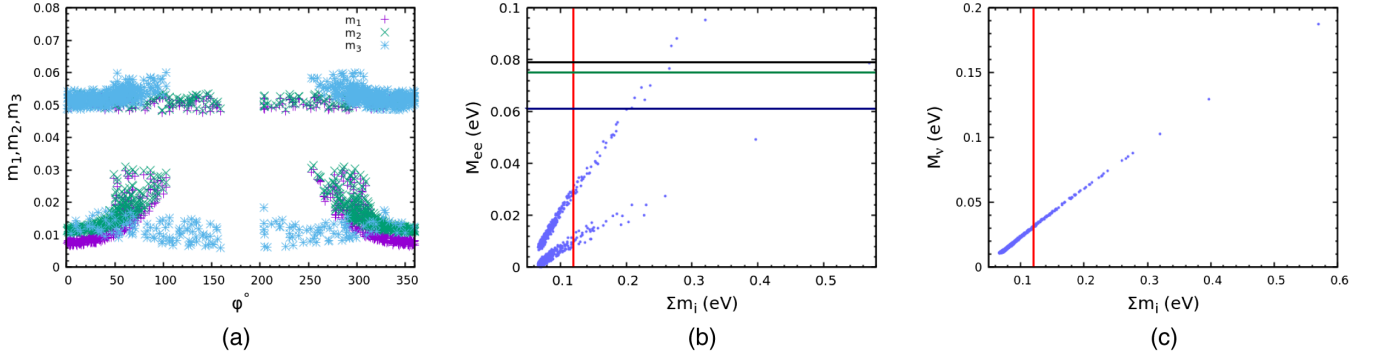


FIG. 10. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{21}$  pattern using  $TM_2$  mixing matrix. The vertical red line shows the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines show the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

$$\begin{aligned} \frac{m_1}{m_2} &= \frac{\tilde{\mathcal{D}}_1 \sin 2(\beta - \phi) + \tilde{\mathcal{D}}_2 \cos 2(\beta - \phi)}{\tilde{\mathcal{D}}_3 \sin 2(\alpha - \beta) + \tilde{\mathcal{D}}_4 \cos 2(\alpha - \beta)}, \\ \frac{m_3}{m_2} &= \frac{\tilde{\mathcal{D}}_1 \sin 2(\beta - \phi) + \tilde{\mathcal{D}}_2 \cos 2(\beta - \phi)}{\tilde{\mathcal{D}}_5 \sin 2(\phi - \alpha) - \tilde{\mathcal{D}}_6 \cos 2(\phi - \alpha)}, \end{aligned} \quad (37)$$

where all the relevant  $\mathcal{D}_i$  and  $\tilde{\mathcal{D}}_i$  are reported in Eqs. (A4) and (A9) of appendix. In Figs. 9(a) and 10(a) we have shown the correlation of neutrino masses  $m_1$ ,  $m_2$ , and  $m_3$  with the unknown parameter  $\phi$  for  $TM_1$  and  $TM_2$  mixing matrix. It shows both normal and inverted mass ordering for  $TM_1$  and  $TM_2$  mixing matrix. The correlation of  $M_{ee}$  against  $\sum m_i$  for  $TM_1$  and  $TM_2$  mixing patterns are shown in Figs. 9(b) and 10(b), respectively. In the Figs. 9(c)

and 10(c), we have shown the correlation of  $M_\nu$  with  $\sum m_i$  for  $TM_1$  and  $TM_2$ , respectively.

We also report the allowed ranges of the absolute neutrino mass, the effective Majorana mass and the effective electron antineutrino mass for both the mixing matrix in Table VI. The phenomenology of this pattern is very similar to that of  $C_{31}$ .

### E. Pattern V: $C_{32} = 0$

The condition of vanishing minor for this pattern is given by

$$(M_\nu)_{ee}(M_\nu)_{\mu\tau} - (M_\nu)_{\mu e}(M_\nu)_{e\tau} = 0. \quad (38)$$

TABLE VI. Allowed range of  $\sum m_i$ ,  $m_{ee}$ , and  $m_\nu$  for  $C_{21}$  pattern.

Mixing matrix	Mass ordering	$\sum m_i$ (eV)	$m_{ee}$ (eV)	$m_\nu$ (eV)
$TM_1$	NO	[0.080, 0.249]	$[2.213 \times 10^{-3}, 0.076]$	[0.017, 0.079]
	IO	[0.101, 0.720]	[0.014, 0.125]	[0.044, 0.241]
$TM_2$	NO	[0.066, 0.569]	$[1.655 \times 10^{-4}, 0.095]$	[0.010, 0.187]
	IO	[0.097, 0.458]	[0.015, 0.137]	[0.045, 0.154]

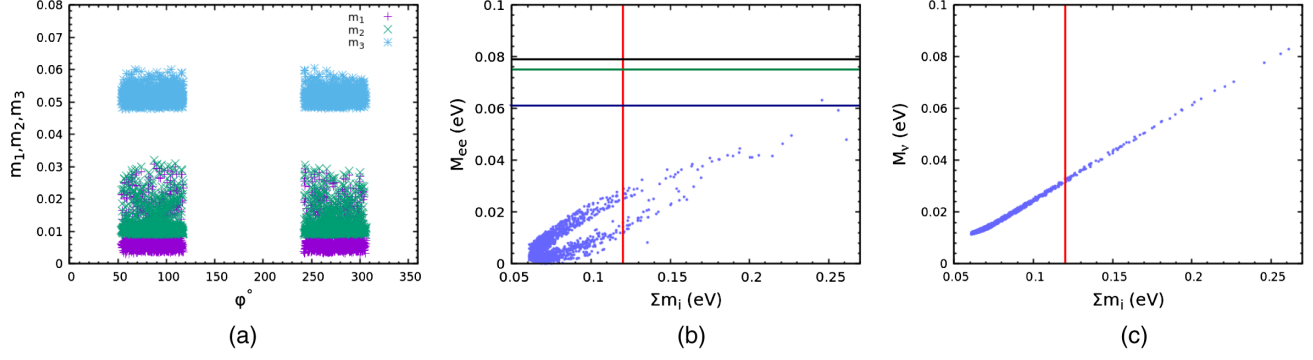


FIG. 11. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{32}$  pattern using  $TM_1$  mixing matrix. The vertical red line represents the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines represent the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

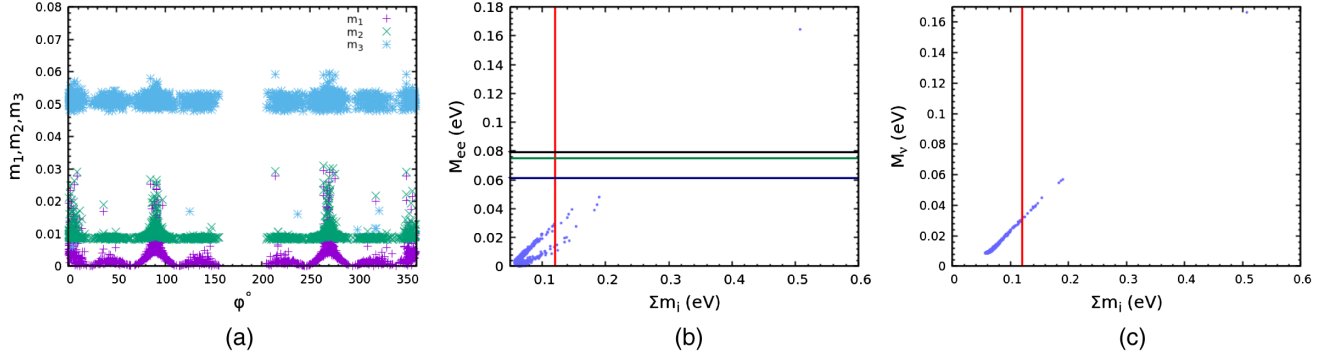


FIG. 12. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{32}$  pattern using  $TM_2$  mixing matrix. The vertical red line is the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines are the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

For this pattern, the two neutrino mass ratios for  $TM_1$  and  $TM_2$  mixing matrix are given by

$$\begin{aligned} \frac{m_1}{m_2} &= \frac{\mathcal{E}_1 \sin 2\beta + \mathcal{E}_2 \cos 2\beta}{\mathcal{E}_3 \sin 2(\alpha - \beta) + \mathcal{E}_4 \cos 2(\alpha - \beta)}, \\ \frac{m_3}{m_2} &= \frac{\mathcal{E}_1 \sin 2\beta + \mathcal{E}_2 \cos 2\beta}{\mathcal{E}_5 \sin 2\alpha + \mathcal{E}_6 \cos 2\alpha}, \end{aligned} \quad (39)$$

and

$$\begin{aligned} \frac{m_1}{m_2} &= \frac{\tilde{\mathcal{E}}_1 \sin 2(\beta - \phi) + \tilde{\mathcal{E}}_2 \cos 2(\beta - \phi)}{\tilde{\mathcal{E}}_3 \sin 2(\alpha - \beta) + \tilde{\mathcal{E}}_4 \cos 2(\alpha - \beta)}, \\ \frac{m_3}{m_2} &= \frac{\tilde{\mathcal{E}}_1 \sin 2(\beta - \phi) + \tilde{\mathcal{E}}_2 \cos 2(\beta - \phi)}{\tilde{\mathcal{E}}_5 \sin 2(\phi - \alpha) - \tilde{\mathcal{E}}_6 \cos 2(\phi - \alpha)}, \end{aligned} \quad (40)$$

respectively. The relevant expressions for  $\mathcal{E}_i$  and  $\tilde{\mathcal{E}}_i$  are reported in Eqs. (A5) and (A10) of appendix. The correlation of neutrino masses  $m_1$ ,  $m_2$ , and  $m_3$  with the unknown parameter  $\phi$  are shown in Figs. 11(a) and 12(a), respectively for  $TM_1$  and  $TM_2$  mixing matrix. It is observed that, it shows normal mass ordering for  $TM_1$  mixing matrix, whereas, for  $TM_2$  mixing matrix, it shows both normal and inverted mass ordering. The correlation of  $M_{ee}$  and  $\sum m_i$  are shown in Figs. 11(b) and 12(b), respectively using  $TM_1$  and  $TM_2$  mixing matrix. In Figs. 11(c) and 12(c), we have shown the correlation of  $M_\nu$  with  $\sum m_i$  using  $TM_1$  and  $TM_2$  mixing matrix, respectively.

The allowed ranges of all the relevant parameters such as the absolute neutrino mass, the effective Majorana mass and the effective electron antineutrino mass under normal and inverted ordering for both the mixing matrix are reported in Table VII.

TABLE VII. Allowed range of  $\sum m_i$ ,  $m_{ee}$ , and  $m_\nu$  for  $C_{32}$  pattern.

Mixing matrix	Mass ordering	$\sum m_i$ (eV)	$m_{ee}$ (eV)	$m_\nu$ (eV)
$TM_1$	NO	[0.061, 0.261]	$[1.309 \times 10^{-4}, 0.063]$	[0.011, 0.082]
$TM_2$	NO	[0.056, 0.507]	$[6.054 \times 10^{-5}, 0.164]$	[0.084, 0.166]
	IO	[0.101, 0.346]	[0.017, 0.078]	[0.045, 0.118]

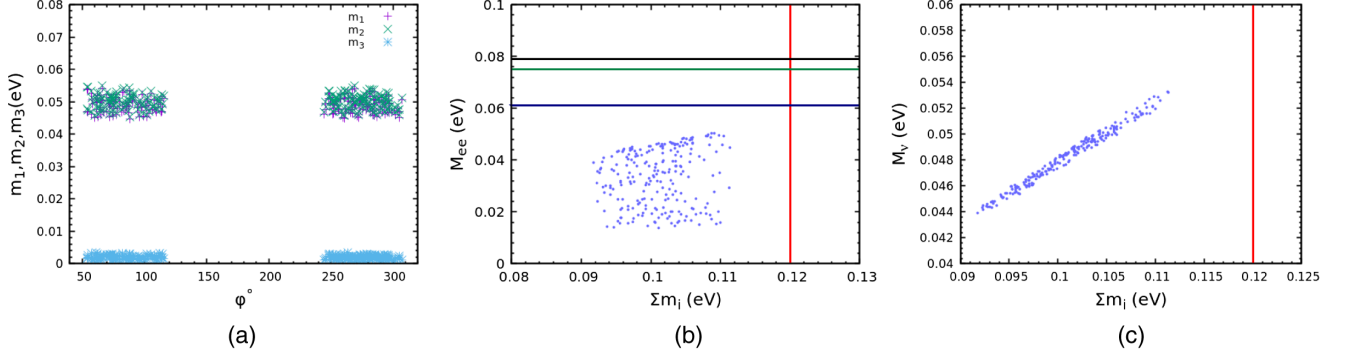


FIG. 13. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{11}$  pattern using  $TM_1$  mixing matrix. The vertical red line represents the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines represent the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].

### F. Pattern VI: $C_{11} = 0$

The vanishing minor condition for this pattern is given by

$$(M_\nu)_{\mu\mu}(M_\nu)_{\tau\tau} - (M_\nu)_{\tau\mu}(M_\nu)_{\mu\tau} = 0. \quad (41)$$

The two neutrino mass ratios can be obtained using the elements from the neutrino mass matrix. For  $TM_1$  mixing matrix, we have

$$\begin{aligned} \frac{m_1}{m_2} &= \frac{2 \sin 2\beta}{\cos^2 \theta \sin 2(\alpha - \beta)}, \\ \frac{m_3}{m_2} &= -\frac{\tan^2 \theta \sin 2\beta}{\sin 2\alpha}, \end{aligned} \quad (42)$$

and for  $TM_2$  mixing matrix, we have

$$\begin{aligned} \frac{m_1}{m_2} &= \frac{2 \cos^2 \theta \sin 2\beta}{\sin 2(\alpha - \beta)}, \\ \frac{m_3}{m_2} &= -\frac{2 \sin^2 \theta \sin 2\beta}{\sin 2\alpha}. \end{aligned} \quad (43)$$

Using Eq. (42), we obtain the mass relation for  $TM_1$  mixing matrix as

$$m_1 \sin 2(\alpha - \beta) - 2m_2 \sin 2\beta + m_3 \sin 2\alpha = 0 \quad (44)$$

and using Eq. (43), we obtain the mass relation for  $TM_2$  mixing matrix as

$$m_1 \sin 2(\alpha - \beta) - 2m_2 \sin 2\beta - m_3 \sin 2\alpha = 0. \quad (45)$$

This pattern gives a clear inverted mass ordering for both  $TM_1$  and  $TM_2$  mixing matrix. The correlation of the neutrino masses  $m_1$ ,  $m_2$  and  $m_3$  for both the mixing patterns with the unknown parameter  $\phi$  is shown in Figs. 13(a) and 14(a), respectively. The correlation of  $M_{ee}$  with  $\sum m_i$  for  $TM_1$  and  $TM_2$  are shown in Figs. 13(b) and 14(b), respectively. In Figs. 13(c) and 14(c), we have shown the correlation of  $M_\nu$  with  $\sum m_i$  for  $TM_1$  and  $TM_2$  mixing matrix, respectively.

The allowed ranges of the absolute neutrino mass, the effective Majorana mass and the effective electron anti-neutrino mass obtained for both the mixing matrix are listed in the Table VIII.

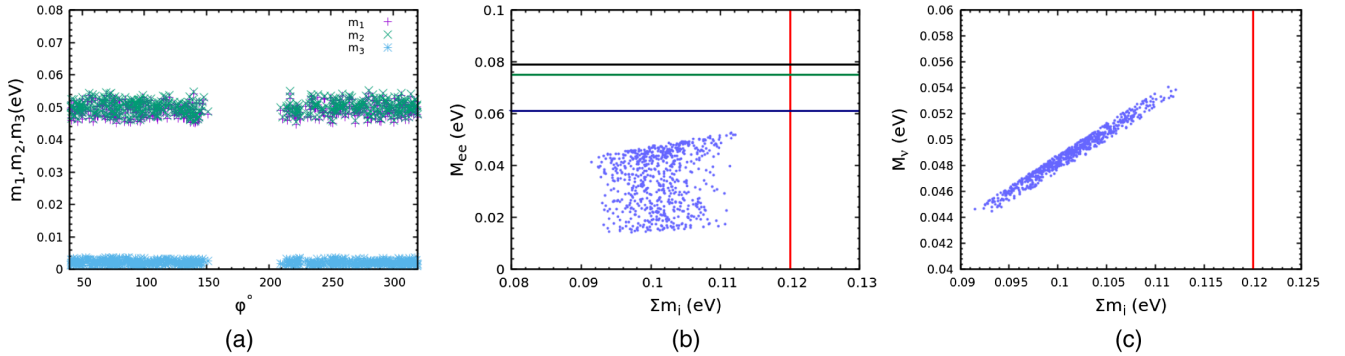


FIG. 14. (a) Variation of  $m_1$ ,  $m_2$ , and  $m_3$  as a function of  $\phi$ , (b) correlation of  $\sum m_i$  and  $m_{ee}$ , and (c) correlation of  $\sum m_i$  and  $m_\nu$  for  $C_{11}$  pattern using  $TM_2$  mixing matrix. The vertical red line is the upper bound of the total neutrino mass reported in Ref. [3]. The black, green, and blue lines are the experimental upper bounds of the effective Majorana mass reported in Refs. [4–7].



TABLE VIII. Allowed range of  $\sum m_i$ ,  $m_{ee}$ , and  $m_\nu$  for  $C_{11}$  pattern.

Mixing matrix	Mass ordering	$\sum m_i$ (eV)	$m_{ee}$ (eV)	$m_\nu$ (eV)
TM <sub>1</sub>	<b>IO</b>	[0.092, 0.111]	[0.013, 0.050]	[0.043, 0.053]
TM <sub>2</sub>	<b>IO</b>	[0.091, 0.112]	[0.014, 0.052]	[0.044, 0.054]

## VI. DEGREE OF FINE TUNING IN THE NEUTRINO MASS MATRIX

In this section, we wish to determine whether the entries of the neutrino mass matrix are fine tuned or not. In order to determine the degree of fine tuning of the mass matrix elements, we define a parameter  $d_{\text{FT}}$  [19,56] which is obtained as the sum of the absolute values of the ratios between each parameter and its error. We follow Refs. [19,56] and define the fine tuning parameter as

$$d_{\text{FT}} = \sum \left| \frac{\text{par}_i}{\text{err}_i} \right|, \quad (46)$$

where  $\text{par}_i = (\theta_{12}, \theta_{13}, \theta_{23}, \Delta m_{21}^2, \Delta m_{31}^2)$  is the best fit values of the parameters. The error  $\text{err}_i$  for each parameter is obtained from the shift of best fit value that changes the  $\chi^2_{\text{min}}$  value by one unit keeping all other parameters fixed at their best fit values. To determine the best fit values of all the parameters, we perform a  $\chi^2$  analysis of all the classes of one minor zero and find the  $\chi^2_{\text{min}}$ . We define the  $\chi^2$  as follows:

$$\chi^2 = \sum_{i=1}^3 \frac{(\theta_i^{\text{cal}} - \theta_i^{\text{exp}})^2}{(\sigma_i^{\text{exp}})^2} + \sum_{j=21,31} \frac{(\Delta m_j^{\text{cal}} - \Delta m_j^{\text{exp}})^2}{(\sigma_j^{\text{exp}})^2}, \quad (47)$$

where  $\theta_i = (\theta_{12}, \theta_{13}, \theta_{23})$  and  $\Delta m_j = (\Delta m_{21}^2, \Delta m_{31}^2)$ . Here  $\theta_i^{\text{cal}}$  and  $\Delta m_j^{\text{cal}}$  represent the theoretical value of  $\theta_i$  and  $\Delta m_j$ , respectively, whereas  $\theta_i^{\text{exp}}$  and  $\Delta m_j^{\text{exp}}$  represent measured central value of  $\theta_i$  and  $\Delta m_j$ , respectively. It should be noted that  $\theta_i^{\text{cal}}$  and  $\Delta m_j^{\text{cal}}$  depend on four unknown model parameters, namely  $\theta$ ,  $\phi$ ,  $\alpha$ , and  $\beta$ .

Similarly, the uncertainties in the measured value of  $\theta_i$  and  $\Delta m_j$  are represented by  $\sigma_i^{\text{exp}}$  and  $\sigma_j^{\text{exp}}$ , respectively. The central values and the corresponding uncertainties in each parameter are reported in Table II.

We first compute  $d_{\text{Data}}$  which is defined as the sum of the absolute values of the ratios between the measured values of each parameter and its error from Table II. We obtain the value of  $d_{\text{Data}}$  to be around 200 for both normal and inverted ordering case. The degree of fine tuning can be roughly estimated from the value of  $d_{\text{FT}}$  because if the  $d_{\text{FT}}$  value is large then a minimal variation of the corresponding parameters give large difference on the value of  $\chi^2$ . Hence a large value of  $d_{\text{FT}}$  corresponds to a strong fine tuning of the mass matrix elements and vice versa. The  $\chi^2_{\text{min}}$  value and the corresponding best fit values of the unknown parameters of the neutrino mass matrix  $\theta$ ,  $\phi$ ,  $\alpha$ ,  $\beta$  and the value of  $d_{\text{FT}}$  parameter for each patterns are listed in the Tables IX and X for the TM<sub>1</sub> and TM<sub>2</sub> mixing matrix respectively. We also report the best fit values of several observables such as  $\theta_{12}$ ,  $\theta_{13}$ ,  $\theta_{23}$ ,  $\Delta m_{21}^2$ , and  $\Delta m_{31}^2$  for each pattern. For the patterns  $C_{33}$ ,  $C_{22}$ ,  $C_{31}$ ,  $C_{32}$ , and  $C_{21}$ , the results are for NO case and for the pattern  $C_{11}$ , the results are for the IO case. As the pattern  $C_{11}$  follows the IO, the  $\chi^2_{\text{min}}$  value obtained for this pattern is large for both TM<sub>1</sub> and TM<sub>2</sub> mixing matrix. The best fit values of the mixing angles  $\theta_{23}$ ,  $\theta_{12}$ ,  $\theta_{13}$  and the mass squared differences  $\Delta m_{21}^2$ ,  $\Delta m_{31}^2$  obtained for each pattern are compatible with the experimentally measured values reported in Table II.

In case of TM<sub>1</sub> mixing matrix, pattern  $C_{23}$  shows very good agreement with the data with a very small  $d_{\text{FT}}$  value. Although, the pattern  $C_{31}$  also have same  $\chi^2$  as pattern  $C_{23}$ , it, however, has a much larger  $d_{\text{FT}}$  value compared to pattern  $C_{23}$ . It can be concluded that for the pattern  $C_{31}$ , there is a strong fine tuning among the elements of the mass

TABLE IX. The values of  $\chi^2_{\text{min}}$ ,  $d_{\text{FT}}$ , the best fit values of  $\theta_{13}^\circ$ ,  $\theta_{12}^\circ$ ,  $\theta_{23}^\circ$ ,  $\theta^\circ$ ,  $\phi^\circ$ ,  $\alpha^\circ$ ,  $\beta^\circ$ ,  $\Delta m_{21}^2 (10^{-5} \text{ eV}^2)$ , and  $\Delta m_{31}^2 (10^{-3} \text{ eV}^2)$  for TM<sub>1</sub> mixing matrix.

Type	$\chi^2_{\text{min}}$	$d_{\text{FT}}$	$\theta_{13}^\circ$	$\theta_{12}^\circ$	$\theta_{23}^\circ$	$\theta^\circ$	$\phi^\circ$	$\alpha^\circ$	$\beta^\circ$	$\Delta m_{21}^2 (10^{-5} \text{ eV}^2)$	$\Delta m_{31}^2 (10^{-3} \text{ eV}^2)$
$C_{33}$	2.66	$8.38 \times 10^3$	8.48	34.35	48.59	14.80	287.67	290.89	320.72	7.51	2.42
$C_{22}$	2.80	$2.35 \times 10^5$	8.46	34.36	48.77	14.77	251.38	313.24	74.37	7.48	2.45
$C_{31}$	1.69	$8.43 \times 10^3$	8.57	34.33	48.83	14.96	108.69	7.46	338.00	7.47	2.44
$C_{32}$	1.69	83.72	8.60	34.33	49.56	15.02	76.68	84.95	357.14	7.47	2.44
$C_{21}$	3.24	$1.36 \times 10^5$	8.67	34.31	50.00	15.07	63.50	269.82	216.81	7.47	2.41
$C_{11}$	4.05	$4.88 \times 10^2$	8.41	34.37	49.33	14.68	291.60	188.42	83.27	7.40	-2.46

TABLE X. The values of  $\chi^2_{\min}$ ,  $d_{\text{FT}}$ , the best fit values of  $\theta_{13}^\circ$ ,  $\theta_{12}^\circ$ ,  $\theta_{23}^\circ$ ,  $\theta^\circ$ ,  $\phi^\circ$ ,  $\alpha^\circ$ ,  $\beta^\circ$ ,  $\Delta m_{21}^2 (10^{-5} \text{ eV}^2)$ , and  $\Delta m_{3l}^2 (10^{-3} \text{ eV}^2)$  for  $\text{TM}_2$  mixing matrix.

Type	$\chi^2_{\min}$	$d_{\text{FT}}$	$\theta_{13}^\circ$	$\theta_{12}^\circ$	$\theta_{23}^\circ$	$\theta^\circ$	$\phi^\circ$	$\alpha^\circ$	$\beta^\circ$	$\Delta m_{21}^2 (10^{-5} \text{ eV}^2)$	$\Delta m_{3l}^2 (10^{-3} \text{ eV}^2)$
$C_{33}$	8.87	$5.86 \times 10^2$	8.57	35.72	48.95	10.52	310.43	178.66	169.70	7.42	2.43
$C_{22}$	9.08	$4.31 \times 10^2$	8.59	35.72	48.85	10.54	230.78	63.70	184.68	7.37	4.43
$C_{31}$	9.23	$4.78 \times 10^4$	8.55	35.72	48.50	10.49	234.75	185.56	217.34	7.37	2.44
$C_{32}$	9.29	$1.09 \times 10^2$	8.56	35.72	48.94	10.51	310.43	355.65	82.48	7.50	2.44
$C_{21}$	9.12	$2.07 \times 10^3$	8.54	35.70	49.80	10.47	142.15	279.77	59.23	7.40	2.43
$C_{11}$	9.82	$2.29 \times 10^2$	8.65	35.73	50.55	10.62	25.81	341.08	103.82	7.42	-2.49

matrix. Similarly,  $C_{22}$ ,  $C_{21}$ , and  $C_{33}$  have larger  $d_{\text{FT}}$  value compared to  $C_{11}$  pattern, although they have less  $\chi^2$  value than  $C_{11}$ . For  $C_{22}$ ,  $C_{21}$ , and  $C_{33}$  also the degree of fine tuning among the mass matrix elements is very strong. Moreover, it is very clear from Table IX that all these patterns prefer the atmospheric mixing angle  $\theta_{23}$  to be greater than  $\pi/4$ . Based on the  $d_{\text{FT}}$  values, it is clear that it requires less fine tuning of the mass matrix elements for patterns  $C_{23}$  and  $C_{11}$ .

For the  $\text{TM}_2$  mixing matrix, the fine-tuned parameter  $d_{\text{FT}}$  is small for the patterns  $C_{33}$ ,  $C_{22}$ ,  $C_{32}$ , and  $C_{11}$ . Among all these patterns  $C_{32}$  has the lowest  $d_{\text{FT}}$  value. However, for patterns  $C_{31}$  and  $C_{21}$ ,  $d_{\text{FT}}$  value is quite large and hence the degree of fine tuning among the elements of the mass matrix is quite strong for these patterns. All the patterns prefer the best fit value of  $\theta_{23}$  to be larger than  $\pi/4$ .

## VII. SYMMETRY REALIZATION

The symmetry of one vanishing minor can be realized through type-I seesaw mechanism [57,58] along with Abelian symmetry. One vanishing minor in the neutrino mass matrix can easily be obtained if one element in the Majorana matrix  $M_R$  is zero along with diagonal Dirac mass matrix  $M_D$ . In order to fulfill this condition, we need three right-handed charged lepton  $l_{Rp}$  ( $p = 1, 2, 3$ ), three right-handed neutrinos  $\nu_{Rp}$  and three left-handed lepton doublets  $D_{Lp}$ . We present the symmetry realization of pattern V. The symmetry of this pattern can be realized through the Abelian symmetry group  $(Z_{12} \times Z_2)$  that is discussed in Refs. [27,59].

The leptonic fields under  $Z_{12}$  transform as

$$\begin{aligned} \bar{l}_{R1} &\rightarrow \omega \bar{l}_{R1}, & \bar{\nu}_{R1} &\rightarrow \omega \bar{\nu}_{R1}, & D_{L1} &\rightarrow \omega \bar{D}_{L1}, \\ \bar{l}_{R2} &\rightarrow \omega^2 \bar{l}_{R2}, & \bar{\nu}_{R2} &\rightarrow \omega^2 \bar{\nu}_{R2}, & D_{L2} &\rightarrow \omega^3 \bar{D}_{L2}, \\ \bar{l}_{R3} &\rightarrow \omega^5 \bar{l}_{R3}, & \bar{\nu}_{R3} &\rightarrow \omega^5 \bar{\nu}_{R3}, & D_{L3} &\rightarrow \omega^8 \bar{D}_{L3}, \end{aligned} \quad (48)$$

where  $\omega = \exp(\frac{i\pi}{6})$ . The bilinears  $\bar{l}_{Rp} D_{Lq}$  and  $\bar{\nu}_{Rp} D_{Lq}$ , where  $p, q = 1, 2, 3$ , relevant for  $(M_I)_{pq}$  and  $(M_D)_{pq}$  transform as  $\bar{l}_{Rp} D_{Lq} \rightarrow \Omega l_{Rp} D_{Lq}$ , where

$$\Omega = \begin{pmatrix} \omega^2 & \omega^4 & \omega^9 \\ \omega^3 & \omega^5 & \omega^{10} \\ \omega^6 & \omega^8 & \omega \end{pmatrix} \quad (49)$$

and the bilinears  $\bar{\nu}_{Rp} C \bar{\nu}_{Rq}^T$  relevant for  $(M_R)_{pq}$  transform as  $\bar{\nu}_{Rp} C \bar{\nu}_{Rq}^T \rightarrow \Lambda \bar{\nu}_{Rp} C \bar{\nu}_{Rq}^T$ , where

$$\Lambda = \begin{pmatrix} \omega^2 & \omega^3 & \omega^6 \\ \omega^3 & \omega^4 & \omega^7 \\ \omega^6 & \omega^7 & \omega^{10} \end{pmatrix}. \quad (50)$$

For each nonzero element in  $M_R$ , we need a scalar singlet  $\chi_{pq}$  and for each nonzero element in  $(M_I)_{pq}$  or  $(M_D)_{pq}$ , we need Higgs scalar  $\phi_{pq}$  or  $\tilde{\phi}_{pq}$ , respectively. The scalar singlets get the vacuum expectation values (vevs) at the seesaw scale, while Higgs doublets get vevs at the electro-weak scale. Under  $Z_2$  transformation, the sign of  $\tilde{\phi}_{pq}$  and  $\nu_{Rp}$  changes, while other multiplets remain invariant. The diagonal charged lepton mass matrix can be obtained by introducing only three Higgs doublets namely  $\phi_{11}$ ,  $\phi_{22}$ , and  $\phi_{33}$ , similarly, the diagonal Dirac neutrino mass matrix can be obtained by introducing three Higgs doublets  $\tilde{\phi}_{11}$ ,  $\tilde{\phi}_{22}$ , and  $\tilde{\phi}_{33}$ . The nonzero elements of  $M_R$  can be obtained by introducing scalar fields  $\chi_{11}$ ,  $\chi_{12}$ ,  $\chi_{13}$ ,  $\chi_{22}$ , and  $\chi_{33}$  which under  $Z_{12}$  transformation gets multiplied by  $\omega^{10}$ ,  $\omega^9$ ,  $\omega^6$ ,  $\omega^8$ , and  $\omega^2$ , respectively. The Majorana mass matrix  $M_R$  can be written as

$$M_R = \begin{pmatrix} a & b & c \\ b & d & 0 \\ c & 0 & e \end{pmatrix}. \quad (51)$$

This provides minor zero corresponding to (3,2) element in the neutrino mass matrix. Other patterns can also be realized similarly for different  $M_R$ .

### VIII. CONCLUSION

We explore the implication of one minor zero in the neutrino mass matrix obtained using trimaximal mixing matrix. There are total six possible patterns and all the patterns are found to be phenomenologically compatible with the present neutrino oscillation data. The two unknown parameters  $\theta$  and  $\phi$  of the trimaximal mixing matrix are determined by using the experimental values of the mixing angles  $\theta_{12}$ ,  $\theta_{23}$ , and  $\theta_{13}$ . It is found that  $TM_1$  mixing matrix provides a better fit to the experimental results than  $TM_2$  mixing matrix. The Jarlskog invariant measure of  $CP$  violation is nonzero for all the pattern, so they are necessarily  $CP$  violating. Patterns I, II, and V show normal mass ordering for  $TM_1$  mixing matrix while these patterns show both normal and inverted mass ordering for  $TM_2$  mixing matrix. Patterns III and IV show both normal and inverted mass ordering for both  $TM_1$  and  $TM_2$  mixing matrix. Pattern VI predicts inverted mass ordering for both the mixing matrix. We predict the unknown parameters such as the absolute neutrino mass scale, the effective Majorana

mass and the effective electron antineutrino mass using both  $TM_1$  and  $TM_2$  mixing matrix. The effective Majorana mass obtained for each pattern is within the reach of neutrinoless double beta decay experiment. Similarly, the value obtained for the effective electron antineutrino mass may be within the reach of future Project 8 experiment. We also discuss the fine tuning of the elements of the mass matrix for all the patterns by introducing a new parameter  $d_{FT}$ . We observe that for the pattern  $C_{23}$ , the fine tuning among the elements of the mass matrix is small compared to other patterns. Moreover, we also discuss the symmetry realization of pattern V using Abelian symmetry group  $Z_{12} \times Z_2$  in the framework of type-I seesaw model which can be easily generalized to all the other patterns as well.

### APPENDIX

The coefficients in the mass ratios for the  $TM_1$  mixing matrix can be expressed in terms of the two unknown parameters  $\theta$  and  $\phi$  as

$$\begin{aligned}
 \mathcal{A}_1 &= \frac{1}{6} \left( \frac{1}{3} \sin^2 \theta + \frac{1}{2} \cos^2 \theta \cos 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
 \mathcal{A}_2 &= \frac{1}{6} \left( \frac{1}{2} \cos^2 \theta \sin 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right), \\
 \mathcal{A}_3 &= \left( \frac{1}{3} \cos^2 \theta + \frac{1}{2} \sin^2 \theta \cos 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right) \left( \frac{1}{3} \sin^2 \theta + \frac{1}{2} \cos^2 \theta \cos 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
 \mathcal{A}_4 &= \left( \frac{1}{2} \sin^2 \theta \sin 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right) \left( \frac{1}{2} \cos^2 \theta \sin 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
 \mathcal{A}_5 &= \left( \frac{1}{2} \sin^2 \theta \sin 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right) \left( \frac{1}{3} \sin^2 \theta + \frac{1}{2} \cos^2 \theta \cos 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
 \mathcal{A}_6 &= \left( \frac{1}{3} \cos^2 \theta + \frac{1}{2} \sin^2 \theta \cos 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right) \left( \frac{1}{2} \cos^2 \theta \sin 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right), \\
 \mathcal{A}_7 &= -\frac{1}{6} \left( \frac{1}{3} \cos^2 \theta + \frac{1}{2} \sin^2 \theta \cos 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
 \mathcal{A}_8 &= -\frac{1}{6} \left( \frac{1}{2} \sin^2 \theta \sin 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right). \tag{A1}
 \end{aligned}$$

$$\begin{aligned}
\mathcal{B}_1 &= \frac{1}{6} \left( \frac{1}{3} \sin^2 \theta + \frac{1}{2} \cos^2 \theta \cos 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{B}_2 &= \frac{1}{6} \left( \frac{1}{2} \cos^2 \theta \sin 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right), \\
\mathcal{B}_3 &= \left( \frac{1}{3} \cos^2 \theta + \frac{1}{2} \sin^2 \theta \cos 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right) \left( \frac{1}{3} \sin^2 \theta + \frac{1}{2} \cos^2 \theta \cos 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{B}_4 &= \left( \frac{1}{2} \sin^2 \theta \sin 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right) \left( \frac{1}{2} \cos^2 \theta \sin 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{B}_5 &= \left( \frac{1}{2} \cos^2 \theta \sin 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right) \left( \frac{1}{3} \cos^2 \theta + \frac{1}{2} \sin^2 \theta \cos 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{B}_6 &= \left( \frac{1}{3} \sin^2 \theta + \frac{1}{2} \cos^2 \theta \cos 2\phi + \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right) \left( \frac{1}{2} \sin^2 \theta \sin 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right), \\
\mathcal{B}_7 &= -\frac{1}{6} \left( \frac{1}{3} \cos^2 \theta + \frac{1}{2} \sin^2 \theta \cos 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{B}_8 &= -\frac{1}{6} \left( \frac{1}{2} \sin^2 \theta \sin 2\phi - \sqrt{\frac{2}{3}} \sin \theta \cos \theta \sin \phi \right). \tag{A2}
\end{aligned}$$

$$\begin{aligned}
\mathcal{C}_1 &= -\frac{1}{3} \left( \frac{1}{3} \sin \theta \cos \theta - \sqrt{\frac{1}{6}} \cos^2 \theta \cos \phi \right), \\
\mathcal{C}_2 &= -\frac{1}{3\sqrt{6}} (\cos^2 \theta \sin \phi), \\
\mathcal{C}_3 &= \left( \frac{1}{6} \sin \theta \cos^3 \theta \sin^2 \phi \right) - \left( \frac{1}{3} \cos^2 \theta + \sqrt{\frac{1}{6}} \sin \theta \cos \theta \sin \phi \right) \left( \frac{1}{3} \sin \theta \cos \theta - \sqrt{\frac{1}{6}} \cos^2 \theta \cos \phi \right), \\
\mathcal{C}_4 &= \sqrt{\frac{1}{6}} \cos^2 \theta \sin \phi \left( \frac{1}{3} \cos^2 \theta + \sqrt{\frac{1}{6}} \sin \theta \cos \theta \cos \phi \right) + \sqrt{\frac{1}{6}} \sin \theta \cos \theta \cos \phi \left( \frac{1}{3} \sin \theta \cos \theta - \sqrt{\frac{1}{6}} \cos^2 \theta \cos \phi \right), \\
\mathcal{C}_5 &= \frac{1}{3} \left( \frac{1}{3} \cos^2 \theta + \sqrt{\frac{1}{6}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{C}_6 &= \frac{1}{3\sqrt{6}} \sin \theta \cos \theta \sin \phi. \tag{A3}
\end{aligned}$$

$$\begin{aligned}
\mathcal{D}_1 &= -\frac{1}{3} \left( \frac{1}{3} \sin^2 \theta - \sqrt{\frac{1}{6}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{D}_2 &= -\frac{1}{3\sqrt{6}} (\cos \theta \sin \phi), \\
\mathcal{D}_3 &= \left( \frac{1}{6} \sin^2 \theta \cos^2 \theta \sin^2 \phi \right) + \left( \frac{1}{3} \cos^2 \theta - \sqrt{\frac{1}{6}} \sin \theta \cos \theta \cos \phi \right) \left( \frac{1}{3} \sin^2 \theta - \sqrt{\frac{1}{6}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{D}_4 &= \sqrt{\frac{1}{6}} \sin \theta \cos \theta \sin \phi \left( \frac{1}{3} \sin^2 \theta + \sqrt{\frac{1}{6}} \sin \theta \cos \theta \cos \phi \right) - \sqrt{\frac{1}{6}} \sin \theta \cos \theta \sin \phi \left( \frac{1}{3} \cos^2 \theta - \sqrt{\frac{1}{6}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{D}_5 &= \frac{1}{3} \left( \frac{1}{3} \cos^2 \theta - \sqrt{\frac{1}{6}} \sin \theta \cos \theta \cos \phi \right), \\
\mathcal{D}_6 &= \frac{1}{3\sqrt{6}} \sin \theta \cos \theta \sin \phi. \tag{A4}
\end{aligned}$$

$$\begin{aligned}
\mathcal{E}_1 &= -\frac{1}{6} \left( \frac{1}{2} \cos^2 \theta \sin 2\phi - \frac{1}{2} \sin^2 \theta \right), \\
\mathcal{E}_2 &= -\frac{1}{12} (\cos^2 \theta \sin 2\phi), \\
\mathcal{E}_3 &= \left( \frac{1}{4} \sin^2 \theta \cos^2 \theta \sin^2 2\phi \right) - \left( \frac{1}{3} \cos^2 \theta - \frac{1}{2} \sin^2 \theta \cos 2\phi \right) \left( \frac{1}{2} \cos^2 \theta \cos 2\phi - \frac{1}{3} \sin^2 \theta \right), \\
\mathcal{E}_4 &= \frac{1}{2} \cos^2 \theta \sin 2\phi \left( \frac{1}{3} \cos^2 \theta - \frac{1}{2} \sin^2 \theta \cos 2\phi \right) + \frac{1}{2} \sin^2 \theta \sin 2\phi \left( \frac{1}{2} \cos^2 \theta \cos 2\phi - \frac{1}{3} \sin^2 \theta \right), \\
\mathcal{E}_5 &= -\frac{1}{6} \left( \frac{1}{3} \cos^2 \theta - \frac{1}{2} \sin^2 \theta \cos 2\phi \right), \\
\mathcal{E}_6 &= \frac{1}{12} \sin^2 \theta \sin 2\phi.
\end{aligned} \tag{A5}$$

Similarly, the coefficients in the mass ratios for the  $\text{TM}_2$  mixing can be written as

$$\begin{aligned}
\tilde{\mathcal{A}}_1 &= \frac{10}{36} \sin^2 \theta \cos^2 \theta \cos 2\phi + \frac{1}{12} \sin^4 \theta - \frac{1}{6\sqrt{3}} \sin^2 \theta \sin 2\theta \cos \phi + \frac{1}{12} \cos^4 \theta (\cos^2 2\phi - \sin^2 2\phi), \\
\tilde{\mathcal{A}}_2 &= \frac{1}{6\sqrt{3}} \cos^2 \theta \sin 2\theta \cos 3\phi - \frac{1}{12} \sin^2 2\theta \cos 2\phi, \\
\tilde{\mathcal{A}}_3 &= \frac{10}{36} \sin^2 \theta \cos^2 \theta \sin 2\phi - \frac{1}{6\sqrt{3}} \sin^2 \theta \sin 2\theta \sin \phi + \frac{1}{6} \cos^4 \theta \cos 2\phi \sin 2\phi, \\
\tilde{\mathcal{A}}_4 &= \frac{1}{6\sqrt{3}} \cos^2 \theta \sin 2\theta \sin 3\phi - \frac{1}{12} \sin^2 2\theta \sin 2\phi, \\
\tilde{\mathcal{A}}_5 &= \frac{1}{18} \sin^2 \theta + \frac{1}{6} \cos^2 \theta \cos 2\phi - \frac{1}{6\sqrt{3}} \sin 2\theta \cos \phi, \\
\tilde{\mathcal{A}}_6 &= \frac{1}{6} \cos^2 \theta \sin 2\phi - \frac{1}{6\sqrt{3}} \sin 2\theta \sin \phi, \\
\tilde{\mathcal{A}}_7 &= \frac{1}{6} \sin^2 \theta + \frac{1}{18} \cos^2 \theta \cos 2\phi + \frac{1}{6\sqrt{3}} \sin 2\theta \cos \phi, \\
\tilde{\mathcal{A}}_8 &= \frac{1}{18} \cos^2 \theta \sin 2\phi + \frac{1}{6\sqrt{3}} \sin 2\theta \sin \phi.
\end{aligned} \tag{A6}$$

$$\begin{aligned}
\tilde{\mathcal{B}}_1 &= \frac{10}{36} \sin^2 \theta \cos^2 \theta \cos 2\phi + \frac{1}{12} \sin^4 \theta + \frac{1}{6\sqrt{3}} \sin^2 \theta \sin 2\theta \cos \phi + \frac{1}{12} \cos^4 \theta (\cos^2 2\phi - \sin^2 2\phi), \\
\tilde{\mathcal{B}}_2 &= -\frac{1}{6\sqrt{3}} \cos^2 \theta \sin 2\theta \cos 3\phi - \frac{1}{12} \sin^2 2\theta \cos 2\phi, \\
\tilde{\mathcal{B}}_3 &= \frac{10}{36} \sin^2 \theta \cos^2 \theta \sin 2\phi + \frac{1}{6\sqrt{3}} \sin^2 \theta \sin 2\theta \sin \phi + \frac{1}{6} \cos^4 \theta \cos 2\phi \sin 2\phi, \\
\tilde{\mathcal{B}}_4 &= -\frac{1}{6\sqrt{3}} \cos^2 \theta \sin 2\theta \sin 3\phi - \frac{1}{12} \sin^2 2\theta \sin 2\phi, \\
\tilde{\mathcal{B}}_5 &= \frac{1}{18} \sin^2 \theta + \frac{1}{6} \cos^2 \theta \cos 2\phi + \frac{1}{6\sqrt{3}} \sin 2\theta \sin \phi, \\
\tilde{\mathcal{B}}_6 &= \frac{1}{6} \cos^2 \theta \sin 2\phi + \frac{1}{6\sqrt{3}} \sin 2\theta \sin \phi, \\
\tilde{\mathcal{B}}_7 &= \frac{1}{6} \sin^2 \theta + \frac{1}{18} \cos^2 \theta \cos 2\phi - \frac{1}{6\sqrt{3}} \sin 2\theta \cos \phi, \\
\tilde{\mathcal{B}}_8 &= \frac{1}{18} \cos^2 \theta \sin 2\phi - \frac{1}{6\sqrt{3}} \sin 2\theta \sin \phi.
\end{aligned} \tag{A7}$$



$$\begin{aligned}
\tilde{C}_1 &= \frac{1}{9} \sin^2 \theta \cos^2 \theta \cos 2\phi - \frac{1}{6\sqrt{3}} \cos^2 \theta \sin 2\theta \cos 3\phi + \frac{1}{6\sqrt{3}} \sin^2 \theta \sin 2\theta \cos \phi - \frac{1}{12} \sin^2 2\theta \cos 2\phi, \\
\tilde{C}_2 &= \frac{1}{9} \sin^2 \theta \cos^2 \theta \sin 2\phi - \frac{1}{6\sqrt{3}} \cos^2 \theta \sin 2\theta \sin 3\phi + \frac{1}{6\sqrt{3}} \sin^2 \theta \sin 2\theta \sin \phi - \frac{1}{12} \sin^2 2\theta \sin 2\phi, \\
\tilde{C}_3 &= \frac{1}{6\sqrt{3}} \sin 2\theta \cos \phi - \frac{1}{9} \sin^2 \theta, \\
\tilde{C}_4 &= \frac{1}{6\sqrt{3}} \sin 2\theta \sin \phi, \\
\tilde{C}_5 &= \frac{1}{9} \cos^2 \theta \cos 2\phi + \frac{1}{6\sqrt{3}} \sin 2\theta \cos \phi, \\
\tilde{C}_6 &= \frac{1}{9} \cos^2 \theta \sin 2\phi + \frac{1}{6\sqrt{3}} \sin 2\theta \sin \phi.
\end{aligned} \tag{A8}$$

$$\begin{aligned}
\tilde{D}_1 &= \frac{1}{9} \sin^2 \theta \cos^2 \theta \cos 2\phi - \frac{1}{6\sqrt{3}} \sin^2 \theta \sin 2\theta \cos 3\phi + \frac{1}{6\sqrt{3}} \cos^2 \theta \sin 2\theta \cos \phi - \frac{1}{12} \sin^2 2\theta \cos 2\phi, \\
\tilde{D}_2 &= \frac{1}{9} \sin^2 \theta \cos^2 \theta \sin 2\phi - \frac{1}{6\sqrt{3}} \sin^2 \theta \sin 2\theta \sin \phi + \frac{1}{6\sqrt{3}} \cos^2 \theta \sin 2\theta \sin 3\phi - \frac{1}{12} \sin^2 2\theta \sin 2\phi, \\
\tilde{D}_3 &= \frac{1}{6\sqrt{3}} \sin 2\theta \cos \phi + \frac{1}{9} \sin^2 \theta, \\
\tilde{D}_4 &= \frac{1}{6\sqrt{3}} \sin 2\theta \sin \phi, \\
\tilde{D}_5 &= \frac{1}{9} \cos^2 \theta \cos 2\phi - \frac{1}{6\sqrt{3}} \sin 2\theta \cos \phi, \\
\tilde{D}_6 &= \frac{1}{6\sqrt{3}} \sin 2\theta \sin \phi - \frac{1}{9} \cos^2 \theta \sin 2\phi.
\end{aligned} \tag{A9}$$

$$\begin{aligned}
\tilde{\mathcal{E}}_1 &= \frac{10}{36} \sin^2 \theta \cos^2 \theta \cos 2\phi - \frac{1}{12} \cos^4 \theta \cos^2 2\phi - \frac{1}{12} \sin^4 \theta + \frac{1}{12} \cos^4 \theta \sin^2 2\phi, \\
\tilde{\mathcal{E}}_2 &= \frac{5}{18} \sin^2 \theta \cos^2 \theta \sin 2\phi - \frac{1}{6} \cos^4 \theta \sin 2\phi \cos 2\phi, \\
\tilde{\mathcal{E}}_3 &= \frac{1}{6} \cos^2 \theta \cos 2\phi - \frac{1}{18} \sin^2 \theta, \\
\tilde{\mathcal{E}}_4 &= \frac{1}{6} \cos^2 \theta \sin 2\phi, \\
\tilde{\mathcal{E}}_5 &= \frac{1}{6} \sin^2 \theta - \frac{1}{18} \cos^2 \theta \cos 2\phi, \\
\tilde{\mathcal{E}}_6 &= \frac{1}{6} \cos^2 \theta \sin 2\phi.
\end{aligned} \tag{A10}$$

- [1] Y. Fukuda *et al.* (Super-Kamiokande Collaboration), *Phys. Rev. Lett.* **81**, 1562 (1998).
- [2] A. A. Esfahani *et al.* (Project 8 Collaboration), *arXiv:2203.07349*.
- [3] M. Zhang, J. F. Zhang, and X. Zhang, *Commun. Theor. Phys.* **72**, 125402 (2020).
- [4] M. Agostini *et al.* (GERDA Collaboration), *Phys. Rev. Lett.* **111**, 122503 (2013).
- [5] K. Alfonso *et al.* (CUORE Collaboration), *Phys. Rev. Lett.* **115**, 102502 (2015).
- [6] J. B. Albert *et al.* (EXO-200 Collaboration), *Nature (London)* **510**, 229 (2014).
- [7] A. Gando *et al.* (KamLAND-Zen Collaboration), *Phys. Rev. Lett.* **110**, 062502 (2013).
- [8] E. I. Lashin and N. Chamoun, *Phys. Rev. D* **85**, 113011 (2012).
- [9] M. Singh, *Adv. High Energy Phys.* **2018**, 2863184 (2018).
- [10] R. R. Gautam, *Phys. Rev. D* **97**, 055022 (2018).
- [11] P. H. Frampton, S. L. Glashow, and D. Marfatia, *Phys. Lett. B* **536**, 79 (2002).
- [12] Z. z. Xing, *Phys. Lett. B* **530**, 159 (2002).
- [13] Z. z. Xing, *Phys. Lett. B* **539**, 85 (2002).
- [14] L. Lavoura, *Phys. Lett. B* **609**, 317 (2005).
- [15] S. Dev, S. Kumar, S. Verma, and S. Gupta, *Phys. Rev. D* **76**, 013002 (2007).
- [16] S. Kumar, *Phys. Rev. D* **84**, 077301 (2011).
- [17] H. Fritzsch, Z. z. Xing, and S. Zhou, *J. High Energy Phys.* **09** (2011) 083.
- [18] P. O. Ludl, S. Morisi, and E. Peinado, *Nucl. Phys.* **B857**, 411 (2012).
- [19] D. Meloni and G. Blankenburg, *Nucl. Phys.* **B867**, 749 (2013).
- [20] W. Grimus and P. O. Ludl, *J. Phys. G* **40**, 055003 (2013).
- [21] S. Dev, R. R. Gautam, L. Singh, and M. Gupta, *Phys. Rev. D* **90**, 013021 (2014).
- [22] R. R. Gautam and S. Kumar, *Phys. Rev. D* **94**, 036004 (2016); **100**, 039902(E) (2019).
- [23] K. S. Channey and S. Kumar, *J. Phys. G* **46**, 015001 (2019).
- [24] M. Singh, *Europhys. Lett.* **129**, 1 (2020).
- [25] E. I. Lashin and N. Chamoun, *Phys. Rev. D* **78**, 073002 (2008).
- [26] E. I. Lashin and N. Chamoun, *Phys. Rev. D* **80**, 093004 (2009).
- [27] S. Dev, S. Verma, S. Gupta, and R. R. Gautam, *Phys. Rev. D* **81**, 053010 (2010).
- [28] S. Dev, S. Gupta, and R. R. Gautam, *Mod. Phys. Lett. A* **26**, 501 (2011).
- [29] Z. Tavartkiladze, *Phys. Rev. D* **106**, 115002 (2022).
- [30] T. Araki, J. Heeck, and J. Kubo, *J. High Energy Phys.* **07** (2012) 083.
- [31] J. Liao, D. Marfatia, and K. Whisnant, *J. High Energy Phys.* **09** (2014) 013.
- [32] S. Dev, R. R. Gautam, and L. Singh, *Phys. Rev. D* **87**, 073011 (2013).
- [33] W. Wang, *Eur. Phys. J. C* **73**, 2551 (2013).
- [34] K. Whisnant, J. Liao, and D. Marfatia, *AIP Conf. Proc.* **1604**, 273 (2015).
- [35] S. Dev, L. Singh, and D. Raj, *Eur. Phys. J. C* **75**, 394 (2015).
- [36] W. Wang, S. Y. Guo, and Z. G. Wang, *Mod. Phys. Lett. A* **31**, 1650080 (2016).
- [37] S. Goswami, S. Khan, and A. Watanabe, *Phys. Lett. B* **693**, 249 (2010).
- [38] S. Dev, S. Verma, and S. Gupta, *Phys. Lett. B* **687**, 53 (2010).
- [39] S. Dev, S. Gupta, and R. R. Gautam, *Phys. Rev. D* **82**, 073015 (2010).
- [40] J. Y. Liu and S. Zhou, *Phys. Rev. D* **87**, 093010 (2013).
- [41] S. Dev, R. R. Gautam, and L. Singh, *Phys. Rev. D* **88**, 033008 (2013).
- [42] P. F. Harrison, D. H. Perkins, and W. G. Scott, *Phys. Lett. B* **530**, 167 (2002).
- [43] P. F. Harrison and W. G. Scott, *Phys. Lett. B* **535**, 163 (2002).
- [44] Z. z. Xing, *Phys. Lett. B* **533**, 85 (2002).
- [45] P. F. Harrison and W. G. Scott, *Phys. Lett. B* **557**, 76 (2003).
- [46] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. Lett.* **107**, 041801 (2011).
- [47] P. Adamson *et al.* (MINOS Collaboration), *Phys. Rev. Lett.* **107**, 181802 (2011).
- [48] Y. Abe *et al.* (Double Chooz Collaboration), *Phys. Rev. Lett.* **108**, 131801 (2012).
- [49] F. P. An *et al.* (Daya Bay Collaboration), *Phys. Rev. Lett.* **108**, 171803 (2012).
- [50] J. K. Ahn *et al.* (RENO Collaboration), *Phys. Rev. Lett.* **108**, 191802 (2012).
- [51] S. Kumar, *Phys. Rev. D* **82**, 013010 (2010); **85**, 079904(E) (2012).
- [52] X. G. He and A. Zee, *Phys. Rev. D* **84**, 053004 (2011).
- [53] W. Grimus and L. Lavoura, *J. High Energy Phys.* **09** (2008) 106.
- [54] C. Jarlskog, *Phys. Rev. Lett.* **55**, 1039 (1985).
- [55] I. Esteban, M. C. Gonzalez-Garcia, M. Maltoni, T. Schwetz, and A. Zhou, *J. High Energy Phys.* **09** (2020) 178.
- [56] G. Altarelli and G. Blankenburg, *J. High Energy Phys.* **03** (2011) 133.
- [57] P. Minkowski, *Phys. Lett.* **67B**, 421 (1977).
- [58] R. N. Mohapatra and G. Senjanovic, *Phys. Rev. Lett.* **44**, 912 (1980).
- [59] W. Grimus, A. S. Joshipura, L. Lavoura, and M. Tanimoto, *Eur. Phys. J. C* **36**, 227 (2004).