# Toward a twistor action for chiral higher-spin gravity

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A covariant twistor action for chiral higher-spin theory in anti–de Sitter and flat space is constructed in term of a holomorphic Chern-Simons theory on twistor space. The action reproduces all known cubic vertices of chiral higher-spin theory in flat space. The spacetime action of the holomorphic Chern-Simons theory in flat space is also obtained.

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### I. INTRODUCTION

The idea of constructing viable interacting higher-spin theories that can avoid no-go theorems/results in flat space [\[1](#page-6-0)–[3](#page-6-1)] and anti–de Sitter (AdS) [[4](#page-6-2)–[6](#page-6-3)] has been going on for the past few decades. However, only a few higher-spin models can overcome the barriers posed by the no-go theorems [\[7](#page-6-4)]. They are the 3D topological higher-spin theories [[9](#page-6-5)–[16\]](#page-6-6), conformal higher-spin gravity [[17](#page-6-7)–[21](#page-6-8)], chiral higher-spin gravity (chiral HSGRA) [\[22](#page-6-9)–[28\]](#page-7-0) and its contractions [\[29](#page-7-1)–[31\]](#page-7-2), as well as higher-spin theories induced by the IKKT-matrix model [\[32](#page-7-3)–[35\]](#page-7-4). In order to retain locality, which is one of the crucial features of field theories, one often has to give up unitarity. As a consequence, local higher-spin theories with propagating degrees of freedom tend to be "chiral" in nature. It is, therefore, not surprising that twistor theory [\[36\]](#page-7-5) is one of ideal frameworks for constructing local higher-spin theories. See, e.g., [\[37\]](#page-7-6) for an expedition in this direction. We note, however, that the chiral higher-spin theories are consistent truncations of some hypothetical complete theories, which are unitary but are usually nonlocal.

There are two main reasons why chiral HSGRAs can avoid no-go theorems/results. Firstly, chiral theories are higher-spin extensions or at least closed cousins of selfdual theories [[38](#page-7-7)–[40\]](#page-7-8). Note that higher-spin extensions of self-dual Yang-Mills (SDYM), and self-dual gravity (SDGR) theories have been obtained recently in [\[29,](#page-7-1)[30](#page-7-9)]. Secondly, constraints from higher-spin symmetry force all possible interactions to cancel each other out in the physical amplitudes in flat space. For instance, both conformal higher-spin gravity and chiral HSGRA have been shown to

have vanishing tree-level amplitudes [\[41](#page-7-10)–[47](#page-7-11)]. Furthermore, the chiral HSGRA was shown to be UV-finite at one-loop [\[45](#page-7-12)–[47\]](#page-7-11).

In this work, we propose a covariant twistor action for the chiral HSGRA in AdS in response to the quest of covariantizing the light cone action of chiral HSGRA. Note that the proposed twistor action for chiral HSGRA is not fully determined as we do not know the explicit expression of what we call  $S_c$ —the correction accounts for the higherspin diffeomorphism of twistor coordinates. However, the twistor Chern-Simons action allows us to compute all three-point scattering amplitudes of chiral HSGRA. We also present the spacetime action of the holomorphic Chern-Simons theory in flat space. As a consistency check, we find that its cubic vertices, after projecting to the light cone gauge, match with the ones in [\[48](#page-7-13)[,49\]](#page-7-14). This partly solves a long-standing problem between the incompatibility of cubic interactions between higher-spin fields in Fronsdal's approach [[50](#page-7-15)–[54](#page-7-16)] and the light cone formalism [\[48](#page-7-13)[,49\]](#page-7-14). The twistor origin of chiral HSGRA indicates that it must be integrable and one-loop exact.

#### II. THE TWISTOR THEORY

#### A. Twistor geometry/correspondence

<span id="page-0-0"></span>Let  $\mathcal{P}T$  be the twistor space associated with a conformally flat Euclidean spacetime  $M$  with cosmological constant Λ,

$$
ds^{2} = \frac{dx_{\mu}dx^{\mu}}{(1 + \Lambda x^{2})^{2}} = \Omega^{2}dx_{\mu}dx^{\mu}, \qquad \mu = 1, 2, 3, 4. \qquad (1)
$$

 $\mathcal{P}T$  is defined as an open subset of  $\mathbb{P}^3$  [[55](#page-7-17)],

$$
\mathcal{P}\mathcal{T} = \{ \mathcal{Z}^A := (\lambda^\alpha, \mu^{\dot{\alpha}}) | I_{AB} \mathcal{Z}^A \hat{\mathcal{Z}}^B \neq 0 \}.
$$
 (2)

Here,  $\mathcal{Z}^A$  are homogeneous coordinates of  $\mathbb{P}^3$ . On  $\mathcal{PT}$ , there is a quaternionic conjugation that maps  $\mathcal{Z}^A$  to its dual twistor  $\hat{\mathcal{Z}}^{\overline{A}} = (\hat{\lambda}^{\alpha}, \hat{\mu}^{\dot{\alpha}})$ , where

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$$
\hat{\lambda}^{\alpha} = (-\overline{\lambda^1}, \overline{\lambda^0}), \qquad \hat{\mu}^{\dot{\alpha}} = (-\overline{\mu^1}, \overline{\mu^0}). \tag{3}
$$

Note that  $\overline{\lambda^1}$  is simply the complex conjugation of  $\lambda^1$ . Furthermore,  $I_{AB}$  is known as the infinity twistor that specifies the conformal factor  $\Omega$  in [\(1\)](#page-0-0). The infinity twistor is a skew bitwistor satisfying [\[60\]](#page-7-18)

$$
\frac{1}{2}I^{AB}\epsilon_{ABCD} = I_{CD}, \qquad I_{AC}I^{BC} = \Lambda \delta_A{}^B. \tag{4}
$$

The infinity twistor has the following representatives:

$$
I^{AB} = \begin{pmatrix} \Lambda \epsilon^{\alpha \beta} & 0 \\ 0 & \epsilon^{\dot{\alpha}\dot{\beta}} \end{pmatrix}, \qquad I_{AB} = \begin{pmatrix} \epsilon_{\alpha \beta} & 0 \\ 0 & \Lambda \epsilon_{\dot{\alpha}\dot{\beta}} \end{pmatrix}, \qquad (5)
$$

which induces a Poisson structure [\[61,](#page-7-19)[62\]](#page-7-20),

$$
\tilde{\Pi} = I^{AB} \overleftrightarrow{\partial_A} \wedge \overrightarrow{\partial_B} = \Lambda \frac{\overleftarrow{\partial}}{\partial \lambda^{\alpha}} \wedge \frac{\overrightarrow{\partial}}{\partial \lambda_{\alpha}} + \frac{\overleftarrow{\partial}}{\partial \mu^{\dot{\alpha}}} \wedge \frac{\overrightarrow{\partial}}{\partial \mu_{\dot{\alpha}}}, \quad (6)
$$

on  $\mathcal{P}\mathcal{T}$ . The above Poisson structure also induces the following star product on  $\mathcal{PT}$  [[20](#page-6-10),[21](#page-6-8)]:

$$
f \star g := f e^{\ell_p \tilde{\Pi}} \wedge g = \sum_{k=0}^{\infty} \frac{\ell_p^k}{k!} f \tilde{\Pi}^k g,\tag{7}
$$

where  $\ell_p$  is some natural length scale that plays the role of a deformation parameter. At  $k = 1$ , we recover the standard Poisson bracket,

$$
\{f,g\}_{\mathcal{P} \mathcal{T}} = f\tilde{\Pi} g. \tag{8}
$$

Note that we will sometime suppress the ∧-products to shorten our expressions.

### B. The twistor action

<span id="page-1-3"></span>In constructing the twistor action for chiral HSGRA, it is useful to define the following Euler operator [[20](#page-6-10)[,21\]](#page-6-8):

$$
\hat{\Sigma} = \mathcal{Z}^A \frac{\partial}{\partial \mathcal{Z}^A},\tag{9}
$$

to measure the weight in  $Z$  of any twistor expression. For instance,  $\hat{\Sigma}D^3 \mathcal{Z} = 4$ , where

$$
D^3 \mathcal{Z} = \epsilon_{ABCD} \mathcal{Z}^A d\mathcal{Z}^B \wedge d\mathcal{Z}^C \wedge d\mathcal{Z}^D \tag{10}
$$

<span id="page-1-0"></span>is the canonical measure on  $\mathcal{P}\mathcal{T}$ . Then, our proposed twistor action for chiral HSGRA in anti–de Sitter space is

$$
\mathbf{S}[\mathbf{A}] = S_{hCS} + S_c = \int D^3 \mathcal{Z} L[\mathbf{A}] + S_c. \quad (11)
$$

Here,

$$
L[\mathbb{A}] = \text{Tr}\left[\mathbb{A} \star \mathbb{A} + \frac{2}{3} \mathbb{A} \star \mathbb{A} \star \mathbb{A}\right]
$$

$$
= \text{Tr}\left[\sum_{h \in \mathbb{Z}} \mathbb{A}_{-h} \star \bar{\partial} \mathbb{A}_h + \frac{2}{3} \sum_{h_i \in \mathbb{Z}} \mathbb{A}_{h_1} \star \mathbb{A}_{h_2} \star \mathbb{A}_{h_3}\right] (12)
$$

<span id="page-1-2"></span>is the Lagrangian that obeys the constraint,

$$
\hat{\Sigma}L[\mathbb{A}] = -4,\tag{13}
$$

so that we have a well-defined integral on  $\mathcal{PT}$  [[63](#page-7-21)]. Note that on  $\mathcal{P}\mathcal{T}$ ,

$$
\bar{\partial} = d\hat{Z}^A \frac{\partial}{\partial \hat{Z}^A}
$$
, and  $A = \sum_{h \in \mathbb{Z}} A_h$ , (14)

where  $A_h \in \Omega^{0,1}(\mathcal{PT}, \text{End}(E) \otimes \mathcal{O}(2h-2))$  is a twistor field corresponding to a spacetime matrix-valued higherspin fields of helicity  $h$ , and  $E$  is some rank- $N$  vector bundle that is locally trivial on the restriction to any twistor line  $X \subset \mathcal{PT}$ .

Lastly, the term  $S_c$  in [\(11\)](#page-1-0) is the correction to the holomorphic Chern-Simons action  $S_{hCS}$  that accounts for higher-spin diffeomorphism of the coordinates  $\mathcal{Z}^A$  [\[66\]](#page-7-22),

$$
\delta \mathcal{Z}^A = \sum_{h \in \mathbb{Z}} \{ \mathcal{Z}^A, \xi_h \}, \qquad \xi_h \in \Gamma(\mathcal{PT}, \mathcal{O}(2h-2)), \quad (15)
$$

which results in a nongauge-invariant measure [\[67\]](#page-7-23).

### III. SCATTERING AMPLITUDES

All three-point tree-level amplitudes of chiral HSGRA in AdS can be computed as follows. By doing integration by parts, we observe that  $I^{AB}\partial_A\partial_Bf = 0$ . Hence, we can reduce the number of the star-products by one in each term of the action  $S_{hCS}$ . Therefore, [\(11\)](#page-1-0) can be cast into the following form:

$$
\mathbf{S} = \int \mathrm{Tr} \bigg[ \sum_{h} A_{-h} \bar{\partial} A_{h} + \frac{2}{3} \sum_{h_i} A_{h_1} A_{h_2} \star A_{h_3} \bigg] + S_c', \quad (16)
$$

<span id="page-1-1"></span>where  $S'_c$  accounts for  $S_c$  in [\(11\)](#page-1-0) and the remnants of  $\partial/\partial \mathcal{Z}^A$ when acting on the holomorphic measure  $D^3 \mathcal{Z}$ . The twistor representative of the (0,1)-form connection  $A_h$  on  $\mathcal{P}\mathcal{T}$  is chosen to be [\[20](#page-6-10)[,21\]](#page-6-8)

$$
\mathbb{A}_{h_i} = \int_{\mathbb{C}} \frac{dt_i}{t_i^{2h_i - 1}} \bar{\delta}^2(t_i \lambda - \lambda_i) e^{t_i[\mu \tilde{\lambda}_i]},\tag{17}
$$

in terms of the on shell four-momentum  $k_i^{\alpha \dot{\alpha}} = \lambda_i^{\alpha} \tilde{\lambda}_i^{\dot{\alpha}}$ , which<br>is a null vector on the tangent space of AdS. Here is a null vector on the tangent space of  $AdS<sub>4</sub>$ . Here,

$$
\bar{\delta}(az - b) = \frac{1}{2\pi i} d\bar{z} \frac{\partial}{\partial \bar{z}} \left( \frac{1}{az - b} \right)
$$
(18)

is a (0,1)-form holomorphic delta function [[68](#page-7-24)]. Parametrizing  $\lambda_{\alpha} = (1, z)$  and  $\lambda'_{\alpha} = (b, a)$ , we define

$$
\bar{\delta}(\langle \lambda \lambda' \rangle) = \frac{1}{2\pi i} d\overline{\lambda^{\dot{\alpha}}} \frac{\partial}{\partial \overline{\lambda^{\dot{\alpha}}}} \frac{1}{\langle \lambda \lambda' \rangle}.
$$
 (19)

Finally, the projective version of the holomorphic delta function is defined by [\[20\]](#page-6-10)

$$
\bar{\delta}_m(\lambda, \lambda') = \left[ \frac{\langle \xi \lambda \rangle}{\langle \xi \lambda' \rangle} \right]^m \bar{\delta}(\langle \lambda \lambda' \rangle) = \int_{\mathbb{C}} \frac{dt}{t^m} \bar{\delta}^2(t\lambda - \lambda'), \quad (20)
$$

which explains the origin of the twistor representative  $(17)$ . Here, the degree of  $t_i$  (or the helicity  $h_i$ ) essentially defines the weight in  $\lambda$  of  $A_{h_i}$ . Furthermore, the crucial difference between our setup and the setup in [\[20](#page-6-10)[,21\]](#page-6-8) is that there is a scalar field corresponding to  $h = 0$ , which is essential for quantum consistency of chiral HSGRA.

It was shown in [[69](#page-7-25)] that the plane wave solutions for higher-spin fields have the same structures with the ones in flat space. This explains why we can use the momentum eigenstates [\(17\)](#page-1-1) in AdS.

<span id="page-2-0"></span>A simple computation shows that

$$
\mathbb{A}_{h_2} \star \mathbb{A}_{h_3} \sim \frac{\ell_p^k}{k!} t_2^{d_{k,h_2}} t_3^{d_{k,h_3}} \left( [23] + \Lambda \left\langle \frac{\partial}{\partial \lambda_2} \frac{\partial}{\partial \lambda_3} \right\rangle \right)^k, \quad (21)
$$

where  $d_{k,h_i} = k + 1 - 2h_i$ . Following the steps in [\[21\]](#page-6-8), we rewrite  $D^3 Z$  as  $\frac{d^4 Z}{\text{VolC}^*}$ , which allows us to integrate out  $\mu$  and λ variables to obtain four-dimensional momentum delta function  $\delta^4(P)$ . Note that

$$
\left\langle \frac{\partial}{\partial \lambda_2} \frac{\partial}{\partial \lambda_3} \right\rangle \delta^4(P) = -[23] \Box_P \delta^4(P),
$$
  
where  $\Box_P = \frac{1}{2} \frac{\partial}{\partial P^{\alpha \dot{\alpha}}} \frac{\partial}{\partial P_{\alpha \dot{\alpha}}}.$  (22)

Furthermore, we use the conventions where  $\langle ab \rangle = a^{\alpha}b_{\alpha}$ and  $[ab] = a^{\dot{\alpha}} b_{\dot{\alpha}}$ . From [\(13\),](#page-1-2) we can read off the constraint<br>between the number of derivatives in (21) and helicities of between the number of derivatives in [\(21\)](#page-2-0) and helicities of the external states,

$$
k = h_1 + h_2 + h_3 - 1.
$$
 (23)

Integrating over  $\mu$ , we obtain

$$
\mathcal{M}_{3}^{\Lambda}(h_{1}, h_{2}, h_{3}) = \frac{\ell_{p}^{k}}{k!} \int d^{2} \lambda dt_{1} dt_{2} dt_{3} t_{1}^{1-2h_{1}} t_{2}^{d_{k,h_{2}}} t_{3}^{d_{k,h_{3}}} \times \left( [23] + \Lambda \left\langle \frac{\partial}{\partial \lambda_{2}} \frac{\partial}{\partial \lambda_{3}} \right\rangle \right)^{k} \times \bar{\delta}^{2} (t_{1} \tilde{\lambda}_{1} + t_{2} \tilde{\lambda}_{2} + t_{3} \tilde{\lambda}_{3}) \times \bar{\delta}^{2} (t_{1} \lambda - \lambda_{1}) \bar{\delta}^{2} (t_{2} \lambda - \lambda_{2}) \bar{\delta}^{2} (t_{3} \lambda - \lambda_{3}). \tag{24}
$$

The integrals over  $\lambda$  and  $t_i$  variables can be performed trivially as in [\[20](#page-6-10)[,21\]](#page-6-8). For instance, we can use  $VolC^*$  to fix  $t_1 = 1$ . Then, the integration over  $\lambda$  gives us delta functions on the support at

$$
t_2 = \frac{\langle 23 \rangle}{\langle 31 \rangle}, \qquad t_3 = \frac{\langle 32 \rangle}{\langle 12 \rangle}.
$$
 (25)

<span id="page-2-1"></span>From here, it is a simple computation to integrate over  $t_2$ and  $t_3$ . After a few more steps of manipulating spinors using momentum conservation, we arrive at

$$
\mathcal{M}_{3}^{\Lambda}(h_{1}, h_{2}, h_{3}) = \frac{[\ell_{p}(1 - \Lambda \Box_{p})]^{h_{1} + h_{2} + h_{3} - 1}}{\Gamma[h_{1} + h_{2} + h_{3}]} \delta^{4}(P) \times [12]^{h_{1} + h_{2} - h_{3}} [23]^{h_{2} + h_{3} - h_{1}} [31]^{h_{3} + h_{1} - h_{2}}.
$$
\n(26)

Note that we do not need to fix the kinetic part of the above three-point amplitudes by symmetry as in [\[69,](#page-7-25)[70](#page-7-26)]. Furthermore, in the flat limit where  $\Lambda \to 0$ , we obtain the standard  $MHV<sub>3</sub>$  amplitudes, which enables us to read off the cubic coupling constants,

$$
\mathcal{C}_{h_1, h_2, h_3} = \frac{\mathcal{C}_p^{h_1 + h_2 + h_3 - 1}}{\Gamma[h_1 + h_2 + h_3]}, \qquad h_1 + h_2 + h_3 > 0. \tag{27}
$$

A nice feature about our twistor construction is that the coupling constant  $\mathcal{C}_{h_1,h_2,h_3}$  is built in, and there is no need to derive it dynamically as in [\[22](#page-6-9)–[24](#page-6-11)].

Using the map between the spinor-helicity formalism and the light cone formalism, see, e.g., [\[71,](#page-7-27)[72\]](#page-7-28) where

$$
i] = 2^{1/4} \begin{pmatrix} \bar{p}_i \beta_i^{-1/2} \\ -\beta_i^{1/2} \end{pmatrix}, \qquad i \rangle = 2^{1/4} \begin{pmatrix} p_i \beta_i^{-1/2} \\ -\beta_i^{1/2} \end{pmatrix}, \qquad (28)
$$

we can express the square and angle brackets as

$$
[ij] = \sqrt{\frac{2}{\beta_i \beta_j}} \bar{\mathbb{P}}_{ij}, \qquad \langle ij \rangle = \sqrt{\frac{2}{\beta_i \beta_j}} \mathbb{P}_{ij}.
$$
 (29)

Note that  $\bar{P}_{ij} = \bar{p}_i \beta_j - \bar{p}_j \beta_i$  with  $p_i = (\beta_i, p_i, p_i, \bar{p}_i)$  is the momentum of the external field that has helicity *h*. Then momentum of the external field that has helicity  $h_i$ . Then, we can show that the above three-point amplitudes in the flat limit reduce to

$$
\mathcal{M}_3^{\Lambda \to 0} = \mathcal{C}_{h_1, h_2, h_3} \frac{\bar{\mathbb{P}}_{23}^{h_1 + h_2 + h_3}}{\beta_1^{h_1} \beta_2^{h_2} \beta_3^{h_3}},
$$
(30)

which are the correct cubic vertices obtained previously in [\[22](#page-6-9)[,23,](#page-6-12)[48](#page-7-13)]. As a remark, it would be interesting to establish the map between [\(26\)](#page-2-1) and the cubic vertices in (A)dS found by Metsaev in [[25](#page-6-13)].

## IV. SPACETIME ACTION IN FLAT SPACE

In this section, we obtain the spacetime action of the holomorphic Chern-Simons action in flat space where  $M :=$  $\lim_{\Delta \to 0} \mathcal{M}$  from the action [\(11\)](#page-1-0). To simplify the problem, we assume that all deformations are sufficiently small so that they will not affect the complex structures on twistor space to avoid the complication that arises from Kodaira's theory; see details in [[73](#page-7-29)].

First of all, notice that the twistor space in this case reduces to the usual flat/undeformed twistor space where

$$
\mathbb{PT} = \{ Z^A = (\lambda^{\alpha}, \mu^{\dot{\alpha}} := F^{\dot{\alpha}}(\lambda, x)) | \lambda_{\alpha} \neq 0 \}. \tag{31}
$$

Here,  $\lambda_{\alpha}$  are coordinates on the Riemann sphere  $X \cong \mathbb{P}^1$ base of the fibration  $\pi: \mathbb{PT} \to \mathbb{P}^1$ , and  $\mu^{\dot{\alpha}}$  up the fibers of the normal bundle

$$
N_X := T(\mathbb{PT})|_{\mathbb{P}^1}/T(\mathbb{P}^1) \simeq \mathcal{O}(1) \oplus \mathcal{O}(1). \tag{32}
$$

<span id="page-3-0"></span>The correspondence between  $PT$  and M is given by the incidence relations,

$$
\mu^{\dot{\alpha}} = F^{\dot{\alpha}}(x,\lambda) = x^{\alpha \dot{\alpha}} \lambda_{\alpha},\tag{33}
$$

where x are complexified spacetime coordinates  $[74,75]$  $[74,75]$  $[74,75]$ . The inverse of the above reads

$$
x^{\alpha \dot{\alpha}} = \frac{\lambda^{\alpha} \hat{\mu}^{\dot{\alpha}} - \hat{\lambda}^{\alpha} \mu^{\dot{\alpha}}}{\langle \lambda \hat{\lambda} \rangle}.
$$
 (34)

Thus, each point  $x \in M$  corresponds to a holomorphic, linearly embedded Riemann sphere  $X \cong \mathbb{P}^1 \subset \mathbb{PT}$ , and any point  $Z \in \mathbb{PT}$  corresponds to a self-dual null  $\alpha$ -plane in M.

<span id="page-3-2"></span>It is convenient to define the following basis [\[76\]](#page-7-32) on the corresponding space PS, which is a projectivization of undotted spinor bundle,

$$
\bar{\partial}_0 = \langle \lambda \hat{\lambda} \rangle \lambda_\alpha \frac{\partial}{\partial \hat{\lambda}_\alpha}, \qquad \bar{\partial}_{\dot{\alpha}} = -\lambda^\alpha \partial_{\alpha \dot{\alpha}}, \qquad (35a)
$$

$$
\bar{e}^0 = \frac{\langle \hat{\lambda} d\hat{\lambda} \rangle}{\langle \hat{\lambda} \rangle^2}, \qquad \bar{e}^{\dot{\alpha}} = -\frac{\hat{\lambda}_a dx^{\alpha \dot{\alpha}}}{\langle \hat{\lambda} \hat{\lambda} \rangle}, \tag{35b}
$$

where  $\bar{\partial}_0$  and  $\bar{\partial}_{\dot{\alpha}}$  are (0,1)-vector fields, and  $\bar{e}^0$ ,  $\bar{e}^{\dot{\alpha}}$  are their dual (0,1)-forms, respectively. Note that the above basis can be defined according to the fact that  $T(\mathbb{P}^1) \cong \mathcal{O}(2)$  and  $T^*(\mathbb{P}^1) \cong \mathcal{O}(-2)$  with  $\lambda$  being our reference of weight. Since,

$$
\bar{\partial} \coloneqq \bar{e}^0 \bar{\partial}_0 + \bar{e}^{\dot{\alpha}} \bar{\partial}_{\dot{\alpha}} \quad \text{where } \bar{\partial}^2 = 0,
$$
 (36)

we will take  $\bar{\partial}$  to be our definition of integrable complex structure on  $\mathbb{PS} \cong \mathbb{P}^1 \times \mathbb{M}$ . Using [\(33\)](#page-3-0), we can check that

$$
\bar{\partial} = d\hat{\lambda}^{\alpha} \frac{\partial}{\partial \hat{\lambda}^{\alpha}} + d\hat{\mu}^{\dot{\alpha}} \frac{\partial}{\partial \hat{\mu}^{\dot{\alpha}}} = d\hat{Z}^{A} \frac{\partial}{\partial \hat{Z}^{A}},\tag{37}
$$

which is the usual definition of the Dolbeault operator on PT. Note that we will use the Dolbeault operator  $\overline{\partial}$  to define twistor "background".

The analog of the Euler operator [\(9\)](#page-1-3) on PS is

$$
\hat{\Sigma}_{\lambda} = \lambda^{\alpha} \frac{\partial}{\partial \lambda^{\alpha}},\tag{38}
$$

<span id="page-3-3"></span>where the constraint  $(13)$  becomes

$$
\hat{\Sigma}_{\lambda}L[\mathbb{A}] = -4. \tag{39}
$$

For later convenience, we also define the following (1,0) vector fields and their dual (1,0)-forms on PS:

$$
\partial_0 := \frac{\hat{\lambda}_{\alpha}}{\langle \lambda \hat{\lambda} \rangle} \frac{\partial}{\partial \lambda_{\alpha}}, \qquad \partial_{\dot{\alpha}} := -\frac{\hat{\lambda}^{\alpha}}{\langle \lambda \hat{\lambda} \rangle} \partial_{\alpha \dot{\alpha}}, \qquad (40a)
$$

$$
e^{0} := \langle \lambda d\lambda \rangle, \qquad e^{\dot{\alpha}} := \lambda_{\alpha} dx^{\alpha \dot{\alpha}}.
$$
 (40b)

Here,  $e^0$  is the holomorphic top-form of the fiber  $\mathbb{P}^1$ . The following relations are useful:

$$
[\bar{\partial}_0, \partial_{\dot{\alpha}}] = \bar{\partial}_{\dot{\alpha}}, \qquad [\bar{\partial}_{\dot{\alpha}}, \partial_0] = \partial_{\dot{\alpha}}.
$$
 (41)

To obtain spacetime action of the chiral HSGRA, it is more convenient to work on PS that has the following Poisson structure:

$$
f\Pi g = \epsilon^{\dot{\alpha}\dot{\beta}}\partial_{\dot{\alpha}}f \wedge \partial_{\dot{\beta}}g. \tag{42}
$$

<span id="page-3-1"></span>When  $\partial_{\alpha}$ -vector field acting on any  $(p, q)$ -form, we must promote the above Poisson structure to

$$
\omega \Pi \eta := {\omega, \eta}_h = \epsilon^{\dot{\alpha}\dot{\beta}} \mathcal{L}_{\partial_{\dot{\alpha}}} \omega \wedge \mathcal{L}_{\partial_{\dot{\beta}}} \eta. \tag{43}
$$

<span id="page-3-4"></span>The holomorphic Poisson structure [\(43\)](#page-3-1) then induces

$$
\omega * \eta := \omega e^{\ell_p \Pi} \wedge \eta = \sum_{k=0}^{\infty} \frac{\ell_p^k}{k!} \omega \Pi^k \eta. \tag{44}
$$

For simplicity, we will set  $\ell_p = 1$  from now and require any  $\omega \in \Omega^{p,q}(\mathcal{PT}, \mathcal{O}(n))$  to satisfy

$$
\hat{\Sigma}_{\lambda} \lrcorner \omega = 0, \qquad \mathcal{L}_{\hat{\Sigma}_{\lambda}} \omega = n\omega. \tag{45}
$$

Here, the notation  $\hat{\Sigma}_{\lambda} = i_{\hat{\Sigma}}$  is the interior product with respect to the  $\hat{\Sigma}_{\lambda}$  vector field. Furthermore, the Lie derivative  $\mathcal{L}_{\partial_{\alpha}}$  acting on  $\mathcal{O}(n)$ -valued  $(p, q)$  forms can be defined via the Cartan's magic formula as

$$
\mathcal{L}_{\partial_{\dot{\alpha}}}\omega = \partial_{\dot{\alpha}}\,\Box\,\eth\omega + \eth(\partial_{\dot{\alpha}}\,\Box\,\omega),\tag{46}
$$

where we denote the usual exterior derivative on PS by  $d_{\text{PS}} \equiv \delta$ . Note that  $\delta$  is defined via a unique Chern connection on the bundle  $\mathcal{O}(n) \to \mathbb{P}^1$  where [\[77\]](#page-7-33)

$$
\delta \coloneqq \partial + \bar{\partial} = d_{\mathbb{S}} + n \frac{\langle \hat{\lambda} d\lambda \rangle}{\langle \lambda \hat{\lambda} \rangle} \wedge . \tag{47}
$$

Here,

$$
d_{\mathbb{S}} \coloneqq e^0 \partial_0 + \bar{e}^0 \bar{\partial}_0 + dx^{\alpha \dot{\alpha}} \frac{\partial}{\partial x^{\alpha \dot{\alpha}}} \tag{48}
$$

is the exterior derivative on the unprojective spinor bundle S. It is easy to check that

$$
\delta e^0 = \delta \bar{e}^0 = 0, \qquad \delta e^{\dot{\alpha}} = e^0 \wedge \bar{e}^{\dot{\alpha}}, \qquad \delta \bar{e}^{\dot{\alpha}} = e^{\dot{\alpha}} \wedge \bar{e}^0.
$$
\n(49)

This is the "frame-dragging" effect caused by the Lie derivative when it acts on vielbeins. In addition, since  $\delta = \partial + \overline{\partial}$ , we get

$$
\partial \bar{e}^0 = 0, \qquad \bar{e}^{\dot{\alpha}} = e^{\dot{\alpha}} \wedge \bar{e}^0. \tag{50}
$$

Using the basis [\(35\)](#page-3-2), we decompose each twistor field as

$$
\mathbf{A} = \mathbf{A}_0 \bar{e}^0 + \mathbf{A}_{\dot{\alpha}} \bar{e}^{\dot{\alpha}}.
$$
 (51)

Furthermore, we have the following rules:

$$
\partial_0 \Box e^0 = 1, \qquad \bar{\partial}_0 \Box \bar{e}^0 = 1,
$$
\n(52a)

$$
\partial_{\dot{\alpha}} \, \lrcorner \, e^{\dot{\beta}} = \delta_{\dot{\alpha}}{}^{\dot{\beta}}, \qquad \bar{\partial}_{\dot{\alpha}} \, \lrcorner \, \bar{e}^{\dot{\beta}} = \delta_{\dot{\alpha}}{}^{\dot{\beta}}. \tag{52b}
$$

Note that to reduce the number of  $*$  product by one as in the case of holomorphic Chern-Simons action on  $\mathcal{PT}$ , we need to assume that [[78](#page-7-34)]

$$
\mathcal{L}_{\partial_{\dot{\alpha}}}\mathbf{A}^{\dot{\alpha}} = \partial_{\dot{\alpha}}\mathbf{A}^{\dot{\alpha}} = 0. \tag{53}
$$

This will be our gauge condition. In this particular gauge, we can check that

$$
\mathcal{L}_{\partial_{\dot{a}}} \mathcal{L}_{\partial^{\dot{a}}} \mathsf{A} = 0. \tag{54}
$$

<span id="page-4-0"></span>As a consequence, the origin Chern-Simons twistor action can be written as

$$
S[\mathbb{A}] = \int_{\mathbb{P}\mathbb{S}} \text{Tr}\bigg[\sum_{h \in \mathbb{Z}} \mathbb{A}_{-h} \bar{\partial} \mathbb{A}_h + \frac{2}{3} \sum_{h_i \in \mathbb{Z}} \mathbb{A}_{h_1} \mathbb{A}_{h_2} * \mathbb{A}_{h_3}\bigg]. \quad (55)
$$

This is, indeed, a crucial fact since we want to have only one derivative, i.e.,  $\overline{\partial}$ , in the kinetic term. Furthermore, we must have at least one positive-helicity field in [\(55\)](#page-4-0) so that the constraint [\(39\)](#page-3-3) can be implemented.

A simple computation results in

$$
\mathcal{L}_{\partial_{\dot{\alpha}}}\mathbb{A} = (\partial_{\dot{\alpha}}\mathbb{A}_0 + \mathbb{A}_{\dot{\alpha}})\bar{e}^0 + \partial_{\dot{\alpha}}\mathbb{A}_{\dot{\beta}}\bar{e}^{\dot{\beta}},\tag{56}
$$

<span id="page-4-2"></span>where  $\partial_{\dot{\alpha}}$  ⊿ A = 0. Then, it can be shown, albeit with some tedium, that

$$
\mathbf{A} \ast \mathbf{A} = \frac{1}{k!} \left[ \partial_{\dot{\alpha}(k)} \mathbf{A}_{\dot{\beta}}, \partial^{\dot{\alpha}(k)} \mathbf{A}_{\dot{\gamma}} \right] \bar{e}^{\dot{\beta}} \bar{e}^{\dot{\gamma}}
$$
  
+ 
$$
\frac{1}{k!} \left[ \partial_{\dot{\alpha}(k)} \mathbf{A}_{0}, \partial^{\dot{\alpha}(k)} \mathbf{A}_{\dot{\beta}} \right] \bar{e}^0 \bar{e}^{\dot{\beta}}
$$
  
+ 
$$
\bar{e}^0 \bar{e}^{\dot{\alpha}} \frac{\partial_{\dot{\gamma}(k-1)} \mathbf{A}_{\dot{\gamma}} \partial^{\dot{\gamma}(k)} \mathbf{A}_{\dot{\alpha}} - \partial_{\dot{\gamma}(k)} \mathbf{A}_{\dot{\alpha}} \partial^{\dot{\gamma}(k-1)} \mathbf{A}^{\dot{\gamma}}}{(k-1)!}, (57)
$$

where we used the convention  $\partial_{\dot{\alpha}(k)} \equiv \partial_{\dot{\alpha}_1} \dots \partial_{\dot{\alpha}_k}$  to shorten our expressions.

Unlike the case of SD HS-YM [\[31\]](#page-7-2), where we can gauge fix  $A_0 \in \Omega^{0,1}(\mathbb{P}^1, \mathcal{O}(n)) = 0$  for  $n \ge -1$  in Woodhouse gauge [[79](#page-7-35)]. The situation here is significantly different since the deformation of twistor geometry is related to derivatives along the horizontal direction with respect to  $T<sub>x</sub>$ M. In particular, there is an in-homogeneous contribution to the  $\bar{e}^{\dot{0}} \wedge \bar{e}^{\dot{\alpha}}$  component of the equation,

$$
\bar{\partial}\mathbb{A} + \mathbb{A} * \mathbb{A} = 0. \tag{58}
$$

<span id="page-4-3"></span><span id="page-4-1"></span>From [\(58\)](#page-4-1), we deduce that [\(57\)](#page-4-2)

$$
0 = \bar{\partial}_0 \mathbb{A}_{\dot{\alpha}} - \bar{\partial}_{\dot{\alpha}} \mathbb{A}_0 + \frac{1}{2} \bar{\partial}_0 \lrcorner \bar{\partial}_{\dot{\alpha}} \lrcorner (57), \tag{59a}
$$

$$
0 = \bar{\partial}_{\dot{\alpha}} \mathbb{A}_{\dot{\beta}} + \frac{1}{k!} \partial_{\dot{\gamma}(k)} \mathbb{A}_{\dot{\alpha}} \partial^{\dot{\gamma}(k)} \mathbb{A}_{\dot{\beta}}.
$$
 (59b)

<span id="page-4-4"></span>Using the fact that  $\bar{\partial}_0 \partial_{\dot{\alpha}} = -\bar{\partial}_{\dot{\alpha}}$ , we obtain the following bution for  $(59a)$ . solution for [\(59a\):](#page-4-3)

$$
\partial_{\dot{\alpha}} \mathcal{A}_0 = -\mathcal{A}_{\dot{\alpha}}.\tag{60}
$$

From [\(60\)](#page-4-4), we can further show that

$$
\bar{\partial}_0 \mathbb{A}_0 = 0, \qquad \bar{\partial}_0 \mathbb{A}_{\dot{\alpha}} = 0. \tag{61}
$$

Namely,  $A_{\dot{\alpha}}$ ,  $A_0$  must be holomorphic in  $\lambda$  when they have positive weight. We can now consider

$$
\mathbb{A}_{h,\dot{\alpha}}\bar{e}^{\dot{\alpha}} = \lambda^{\alpha(2h-1)}A_{\alpha(2h-1),\dot{\alpha}}\bar{e}^{\dot{\alpha}}.\tag{62}
$$

<span id="page-4-5"></span>This can be used to solve for the zero component of A as

$$
\mathbb{A}_{h,0}^+ \bar{e}^0 = -\frac{\bar{\partial}^{\dot{\alpha}}}{\Box} \mathbb{A}_{\dot{\alpha}} \bar{e}^0, \qquad h > 0,\tag{63a}
$$

$$
\mathbb{A}_{h,0}^- \bar{e}^0 = \frac{\hat{\lambda}_{\alpha(2|h|)}}{\langle \lambda \hat{\lambda} \rangle^{2|h|}} B^{\alpha(2|h|)} \bar{e}^0, \qquad h \le 0, \qquad (63b)
$$

where  $\lambda^{\alpha(s)} = \lambda^{(\alpha_1} \dots \lambda^{\alpha_s)}, \hat{\lambda}_{\alpha(s)} = \hat{\lambda}_{(\alpha_1} \dots \hat{\lambda}_{\alpha_s)}.$  Note that while the nonlocal  $\square^{-1}$  may look "dangerous" at the moment, it will disappear after we integrating out all fiber coordinates (see below).

To include the scalar field in the spectrum of chiral HSGRA, we can consider the following twistor field:

$$
\mathbb{A}_{h=0^+} := \vartheta = \frac{\hat{\lambda}_{\alpha}}{\langle \lambda \hat{\lambda} \rangle} \vartheta^{\alpha}{}_{\dot{\alpha}} \bar{e}^{\dot{\alpha}}.
$$
 (64)

<span id="page-5-0"></span>Here,  $\theta^{\alpha}{}_{\dot{\alpha}}$  is the auxiliary field associated with the scalar field, which can be integrated out by its own equation of motion as observed in [[80](#page-7-36)]. It is not hard to show that

$$
\theta^{\alpha}{}_{\dot{\alpha}} = (\partial^{\alpha}{}_{\dot{\alpha}} + A^{\alpha}{}_{\dot{\alpha}}*)A_0 + \frac{1}{k!} \partial_{\dot{\gamma}_1 \dots \dot{\gamma}_{k+1}} A_{\dot{\alpha}} \partial^{(\dot{\gamma}_1} \dots \partial^{\dot{\gamma}_k} A^{\dot{\gamma}_{k+1})},
$$
\n(65)

<span id="page-5-1"></span>where  $A_{\alpha\dot{\alpha}} \in {\bigoplus_{s} \Gamma(\mathbb{PT}, \text{End}(E) \otimes \mathcal{O}(2s-2)) | s \geq 1 \}.$ The twistor action [\(55\)](#page-4-0) reads

$$
\mathbf{S} = \int_{\mathbb{P}\mathbb{S}} \mathbf{U} \text{Tr} \left[ \mathbb{A}_0 (\bar{\partial}_{\dot{\alpha}} + \mathbb{A}_{\dot{\alpha}}*) \mathbb{A}^{\dot{\alpha}} + \frac{\lambda^{\gamma} \hat{\lambda}^{\beta}}{\langle \lambda \hat{\lambda} \rangle} \partial_{\alpha \dot{\alpha}} \partial_{\beta}^{\dot{\alpha}} \right] + S_c, \quad (66)
$$

where the measure  $\sigma$  is [[76](#page-7-32)]

$$
\mathfrak{V} = D^3 Z \bar{e}^0 [\bar{e}^{\dot{\alpha}} \bar{e}_{\dot{\alpha}}] = d^4 x \frac{\langle \lambda d\lambda \rangle \wedge \langle \hat{\lambda} d\hat{\lambda} \rangle}{\langle \lambda \hat{\lambda} \rangle^2} = d^4 x \mathbf{K}, \quad (67)
$$

and K is the top form on  $\mathbb{P}^1$ . The spacetime action for chiral HSGRA on a flat background can be obtained by integrating out fibre coordinates using [\[56,](#page-7-37)[79](#page-7-35)[,80\]](#page-7-36),

$$
\int_{\mathbb{P}^1} \mathcal{K} \frac{\hat{\lambda}_{\alpha(m)} \lambda^{\beta(m)}}{\langle \lambda \hat{\lambda} \rangle^m} = -\frac{2\pi i}{(m+1)} \epsilon_{\alpha}^{\beta} \dots \epsilon_{\alpha}^{\beta},\tag{68}
$$

where we adopted the same convention in [\[31\]](#page-7-2).

<span id="page-5-2"></span>The resulting spacetime action [after substituting [\(65\)](#page-5-0) to [\(66\)](#page-5-1) and do some suitable rescaling] is the following action:

$$
S = \langle \mathbf{B}_{\alpha\alpha} | D^{\alpha}_{\dot{\alpha}} \mathbf{A}^{\dot{\alpha}\dot{\alpha}} \rangle - \frac{1}{2} \langle D_{\alpha\dot{\alpha}} \tilde{\Phi} | D^{\alpha\dot{\alpha}} \tilde{\Phi} \rangle
$$
  
+ \langle \tilde{\Phi} | [[\mathbf{A}\_{\dot{\alpha}}, \mathbf{A}^{\dot{\alpha}}]] \rangle - \frac{1}{2} \langle [[\mathbf{A}\_{\dot{\gamma}}, \partial^{\dot{\gamma}} \mathbf{A}\_{\alpha\dot{\alpha}}]] | [[\mathbf{A}\_{\dot{\beta}}, \partial^{\dot{\beta}} \mathbf{A}^{\alpha\dot{\alpha}}]] \rangle, (69)

where we define

$$
\langle X|Y\rangle := \int dx^4 X_{\alpha(n)} Y^{\alpha(n)}.
$$
 (70)

<span id="page-5-3"></span>To understand the above angled bracket notation, it is convenient to introduce

$$
\mathbf{A}^{\alpha \dot{\alpha}} := \sum_{s \ge 1} A^{\alpha(2s-1), \dot{\alpha}} y_{\alpha(2s-2)}, \qquad y_{\alpha(n)} = y_{\alpha} \dots y_{\alpha} \qquad (71a)
$$

$$
\mathbf{A}^{\dot{\alpha}} := \sum_{s \ge 1} A^{\alpha(2s-1), \dot{\alpha}} y_{\alpha(2s-1)},
$$
\n(71b)

as the generating function for positive-helicity fields, and

$$
\mathbf{B}_{\alpha\alpha} := \sum_{s \ge 1} B_{\alpha(2s)} \tilde{\mathbf{y}}^{\alpha(2s-2)}, \qquad \tilde{\mathbf{y}}^{\alpha(n)} = \tilde{\mathbf{y}}^{\alpha} \dots \tilde{\mathbf{y}}^{\alpha}, \qquad (72)
$$

<span id="page-5-4"></span>as generating functions for nonpositive helicity fields. Lastly, the field  $\tilde{\Phi}$  contains both positive and negative helicity fields since its originated from  $A_0$ ,

$$
\tilde{\Phi} := \sum_{s \ge 0} \frac{\partial^{\beta \dot{\alpha}}}{\Box} A^{\alpha(2s-1)}{}_{\dot{\alpha}} y_{\beta \alpha(2s-1)} + B_{\alpha(2s)} \tilde{y}^{\alpha(2s)}, \quad (73)
$$

where we note that the coefficients come with  $y^{\alpha}$  are positive helicity fields and the coefficients come with  $\tilde{y}^{\alpha}$  are negative helicity fields, and  $B_0 = \phi$  the scalar field.

To proceed, we will treat the commuting auxiliary variables  $y_\alpha$ ,  $\tilde{y}^\alpha$  as creation and annihilation oscillators with the property that each  $\tilde{y}$  will consume one y and give us a Kronecker delta for contraction. Therefore, at free level, the second term in [\(69\)](#page-5-2) reduces to the usual kinetic term of free fields in spacetime. Furthermore,

$$
D_{\dot{\alpha}}^{\alpha} \bullet := \partial_{\dot{\alpha}}^{\alpha} \bullet + [\![\mathbf{A}_{\dot{\alpha}}^{\alpha}, \bullet]\!],\tag{74}
$$

where the double square bracket takes the following form by virtue of the  $*$  product [\(44\)](#page-3-4), e.g.,

$$
\begin{split} [\![A^{\alpha(2s-1)},\! \dot{\gamma}, A^{\alpha(2s'-1),\dot{\gamma}}]\!] \\ &:= \frac{1}{k!} [\partial_{\alpha \dot{\beta}_1} \dots \partial_{\alpha \dot{\beta}_k} A^{\alpha(2s-1)},\! \dot{\gamma}, \partial_{\alpha}^{\dot{\beta}_1} \dots \partial_{\alpha}^{\dot{\beta}_k} A^{\alpha(2s'-1),\dot{\gamma}}], \end{split} \tag{75}
$$

where we recall that we have set  $\ell_p = 1$  in this section for simplicity. Here, all undotted indices of the partial derivatives are understood to contract with the ones of physical fields in every possible way.

To this end, let us explain how the  $(+, +, +)$  cubic vertices come to be. Due to the contraction between derivatives originated from the  $*$  product and the one in  $\Phi_{h>0}$ , we can form  $\Box$  to cancel out the nonlocal  $\Box^{-1}$  in [\(63a\).](#page-4-5) To illustrate, let us look at the term  $\langle \tilde{\Phi} | [[\mathbf{A}_{\dot{\alpha}}, \mathbf{A}^{\dot{\alpha}}]] \rangle$ .<br>Using the representatives (71) and (73), we obtain Using the representatives [\(71\)](#page-5-3) and [\(73\)](#page-5-4), we obtain

<span id="page-5-5"></span>
$$
\frac{\partial^{i\dot{\gamma}}}{\Box}A^{\alpha(2s_1-1)}{}_{\dot{\gamma}}[\partial_{\alpha\dot{\beta}}...\partial_{\alpha\dot{\beta}}A^{\alpha(2s_2-1)}{}_{\dot{\circ}},\partial_{\bullet}\dot{\beta}...\partial_{\alpha}\dot{\beta}A^{\alpha(2s_3-1)\dot{\circ}}].
$$
 (76)

Upon integrating by part and applying the identity  $\frac{\partial^{\dot{\gamma}}\partial_{\dot{\gamma}}\partial_{\dot{\gamma}}}{\partial \dot{\gamma}} \sim \Box \dot{\epsilon}^{\dot{\gamma}}{}_{\dot{\beta}}$ , we can cancel the  $\Box^{-1}$  in [\(76\)](#page-5-5). This leaves us with the following  $(+, +, +)$  vertices:

$$
A^{\alpha(2s_1-1)} \cdot [\partial_{\alpha\dot{\gamma}} \dots \partial_{\alpha\dot{\gamma}} A^{\alpha(2s_2-1)} \dot{\alpha}, \partial_{\alpha} \dot{\gamma} \dots \partial_{\alpha} \dot{\gamma} \partial_{\alpha} \dot{\gamma} A^{\alpha(2s_3-1), \dot{\alpha}}].
$$
 (77)

Notice that there must be at least one extra pair of derivatives coming from the  $*$  product to generate the  $(+, +, +)$  vertices. In addition, the all-plus vertices have maximal number of derivatives allowed by kinematics. As such,  $(+, +, +)$  vertices represent nonminimal couplings.

### V. DISCUSSION

In this paper, we have constructed a covariant action for chiral HSGRA in AdS from a Chern-Simons action on twistor space. The twistor origin of chiral HSGRA indicates that it must be integrable and one-loop exact. It is intriguing to ask whether we can have a world sheet description for the chiral HSGRA to gain control over nonlocality issues the moment we step outside the self-dual sectors. It will be a crucial step in finding a higher-spin theory with unitary completion where we have total control of nonlocal interactions.

One of the fundamental questions the results of this paper can address is whether having a covariant form for interacting higher-spin theories is an advantage. As in other covariant formulations of higher-spin theories, the twistor construction provides us a clear view about higher-spin symmetry that governs the chiral HSGRA. The covariant action of the chiral HSGRA also enable us to see some vertices in terms of spacetime derivatives that can not be obtained using Fronsdal's approach. However, note that the computation of scattering amplitudes of the action [\(69\)](#page-5-2) is more involved compared to the calculations in the light cone gauge even at tree level.

Contrary to the old folklore, recent developments in constructing higher-spin interactions by means of the lightfront formalism or free differential algebra show that there exists a smooth deformation between the vertices in flat space and AdS [\[25,](#page-6-13)[28\]](#page-7-0). In this paper, we also confirm that such deformation exists at the level of the twistor action  $(11)$ .

Lastly, chiral HSGRA is known to admit  $U(N)$ ,  $O(N)$ and  $USp(N)$  gaugings [[22](#page-6-9),[23](#page-6-12)[,46\]](#page-7-38); see also [\[81\]](#page-7-39). From this perspective, it is plausible that chiral HSGRA will become the self-dual part of the HS-IKKT matrix model with truncated higher-spin spectrum [\[32](#page-7-3)–[35\]](#page-7-4) in the deep quantum regime where spacetime coordinates no longer commute. We postpone the study of finding the connection between chiral HSGRA and HS-IKKT for future work.

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