Implications for the ΔA_{FB} anomaly in $\bar{B}^0 \to D^{*+} \mathscr{C}^- \bar{\nu}$ using a new Monte Carlo event generator

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Recent experimental results in B physics from Belle, BABAR, and LHCb suggest new physics (NP) in the weak $b \to c$ charged-current processes. Here we focus specifically on the decay modes $\bar{B}^0 \to D^{*+} \ell^- \bar{\nu}$ with $\ell = e$ and μ . The world averages of the ratios R_D and R_D^* currently differ from the Standard Model (SM) predictions by 3.4σ while recently a new anomaly has been observed in the forward-backward asymmetry measurement, $A_{\rm FB}$, in $\bar{B}^0 \to D^{*+} \mu^- \bar{\nu}$ decay. It is found that $\Delta A_{\rm FB} = A_{\rm FB}(B \to D^* \mu \nu) - D^* \mu \nu$ $A_{\rm FB}(B \to D^* e\nu)$ is around 4.1 σ away from the SM prediction in an analysis of 2019 Belle data. In this work we explore possible solutions to the $\Delta A_{\rm FB}$ anomaly and point out correlated NP signals in other angular observables. These correlations between angular observables must be present in the case of beyond the Standard Model physics. We stress the importance of Δ type observables that are obtained by taking the difference of the observable for the muon and the electron mode. These quantities cancel form-factor uncertainties in the SM and allow for clean tests of NP. These intriguing results also suggest an urgent need for improved simulation and analysis techniques in $\bar{B}^0 \to D^{*+} \ell^- \bar{\nu}$ decays. Here we also describe a new Monte Carlo event generator tool based on EVTGEN that we developed to allow simulation of the NP signatures in $\bar{B}^0 \to D^{*+} \ell^- \nu$, which arise due to the interference between the SM and NP amplitudes. We then discuss prospects for improved observables sensitive to NP couplings with 1, 5, 50, and 250 ab^{-1} of Belle II data, which seem to be ideally suited for this class of measurements.

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I. INTRODUCTION

A powerful way to study physics beyond the Standard Model (SM) is via virtual effects of new particles, not present in the SM, in low energy experiments. These virtual effects can in many cases probe mass scales beyond the reach of present or proposed colliders, where the new particles are expected to appear. There is also the possibility that beyond the Standard Model physics comes in the form of weakly coupled light new states. These new states are

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Published by the American Physical Society under the terms of the Creative Commons Attribution 4.0 International license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP³. more likely to be detected at low energy, high precision experiments. In this work we will focus on charged current semileptonic *B* decays, $\bar{B}^0 \to D^{*+} \ell^- \bar{\nu}$ with $\ell = e$ and μ . These decays originate from the underlying quark-level transitions $b \to c\ell^-\bar{\nu}_\ell$, where $\ell = e, \mu$, or τ . At the hadron level they manifest as decays such as $\bar{B} \to D^{(*)}\ell^-\bar{\nu}_\ell$.

The charged-current decays $B \to D^{(*)}\tau\nu_{\tau}$ have been measured by the *BABAR*, Belle, and LHCb experiments. Discrepancies with SM predictions of $R_{D^{(*)}}^{\tau\ell} \equiv \mathcal{B}(\bar{B} \to D^{(*)}\tau^-\bar{\nu}_{\tau})/\mathcal{B}(\bar{B} \to D^{(*)}\ell^-\bar{\nu}_{\ell})$ ($\ell' = e, \mu$) [1–10] have been observed thus far. The SM predictions and the corresponding world-averaged experimental results from the heavy flavor averaging group (HFLAV) [11] are shown in Table I. The deviation from the SM in $R_D^{\tau\ell}$ and $R_{D^*}^{\tau\ell}$ (combined) has a significance of 3.4 σ [11]. These measurements suggest the presence of new physics (NP) that is lepton-flavor universality violating (LFUV) in $b \to c\tau\nu_{\tau}$ decays.

We will focus on the decay $\bar{B}^0 \to D^{*+} \ell^- \bar{\nu}$ as a laboratory to explore NP effects in $b \to c \ell^- \bar{\nu}_{\ell}$ transitions. At leading order, the $\bar{B}^0 \to D^{*+} \ell^- \bar{\nu}$ transitions proceed via the SM. However, new interactions can affect these decays.

TABLE I. Measured values of observables that suggest NP in $b \rightarrow c\tau\nu_{\tau}$. Measurements presented in this table refer to world averages (WA). Note that in [12], the most recent lattice data from [13] on $B \rightarrow D^* \ell \nu$ form factors were used to obtain the SM prediction for $R_{D^*}^{\tau/\ell}$, 0.2586 \pm 0.0030.

Observable	SM prediction	Measurement (WA)
$R_{D^*}^{ au/\ell}$	0.258 ± 0.005 [11]	0.295 ± 0.011 ± 0.008 [11]
$R_D^{ au/\ell}$	$0.299 \pm 0.003 \ \text{[11]}$	$0.340 \pm 0.027 \pm 0.013$ [11]
$R_{D^*}^{\mu/e}$	~1.0	$1.04 \pm 0.05 \pm 0.01$ [14]

In experiment, the underlying transition is $b \rightarrow c\ell X$ where the invisible state X can be a left-handed (LH) neutrino (part of the SM LH doublet of leptons) or a light right-handed (RH) singlet neutrino. Here we will focus on NP scenarios that produce only LH neutrinos in the final state.

Although theoretical work on NP has concentrated on the semileptonic τ modes, where experimental statistics are limited, attention is now also being paid to the semileptonic muon and electron modes where large data samples will be available. For example, scaling the Belle results in [15] to Belle II at 50 ab⁻¹ we expect a yield of 8×10^6 events in each of the muon and electron modes. Similarly, scaling the *BABAR* results in [16] on $B \rightarrow D^* \ell \nu$ with a fully reconstructed hadronic tag, we expect 3×10^5 events with no background.

An additional advantage is that the missing neutrino momentum can be calculated from kinematic constraints of e^+e^- production at the $\Upsilon(4S)$ and the angular distributions can be fully reconstructed. Unlike the τ , which is detected through its decay products, the muon and electron are directly detected in experiment. In contrast, for semileptonic B decays to the τ lepton, the final state contains one or more additional neutrinos from the τ decay, which complicates the situation. Examining NP in the muon mode is further motivated by the anomalous $(g-2)_{\mu}$ measurements [17] as well as by the neutral-current LFUV B anomalies in the $b \rightarrow s\mu^+\mu^-$ decays (see, for example, Ref. [18]). At first glance, when studying the B anomalies within the framework of an effective field theory (EFT), these anomalies may appear unrelated. However, within an SMEFT framework NP in the $b \rightarrow s\mu^+\mu^-$ transition could imply NP in the $b \to c \mu^- \bar{\nu}_{\mu}$ decay [19]. In this article, therefore, we will focus on the muon and electron modes, assuming that the electron decay mode is well described by the SM, but NP contributions are allowed in the muon mode.

Although hints for NP have appeared in the ratio of rates such as $R_{D^{(*)}}$, establishing NP and diagnosing the type of NP will require examination of deviations from the SM in other observables as well. Several observables can be constructed from a complete differential distribution of events using helicity angles. Figure 1 shows a schematic definition of the three helicity angles in $\bar{B} \rightarrow D^*(\rightarrow D\pi) \ell^- \bar{\nu}$.



FIG. 1. Schematic diagram defining various angles in $\bar{B} \rightarrow D^*(\rightarrow D\pi)\ell^-\bar{\nu}$ decay [20]. We have aligned the coordinate axes so that the decaying \bar{B} meson is at rest at the origin and in this frame the momentum of the D^* meson is oriented along the *z* axis. Subsequent decays are shown in the rest frames of the corresponding object that is decaying— $D^* \rightarrow D\pi$ is in the rest frame of the D^* and a virtual particle decays into $\ell^-\bar{\nu}$. The polar angles, θ^* and θ_{ℓ} , are, respectively, defined in these subsequent rest frames, while the azimuthal angle, χ , is defined in the rest frame of the \bar{B} meson.

Angular observables are even more interesting as these may provide one or more unambiguous signals for NP. One such sensitive angular observable is the forward-backward asymmetry of the charged lepton, $A_{\rm FB}$, which can be reconstructed as the difference between the number of leptons with the lepton's helicity angle, θ_{ℓ} (see Fig. 1), greater and less than $\pi/2$. Another observable is S_3 , which can be reconstructed as an asymmetric integral over the angle χ , which measures the difference between the decay planes of the D^* and the lepton-neutrino system (see Fig. 1). There are additional interesting and correlated angular observables, such as S_5 and S_7 , which require asymmetric integrals over multiple helicity angles. In Ref. [20], it was shown that NP in the μ modes can also be detected in the CP-violating triple-product terms, like S_7 , in the angular distribution [21,22]. Some previous work in the literature on the effects of new physics in angular observables of semileptonic B-meson decays can be found in [23–28].

A nonzero $A_{\rm FB}$ is present in both the muon and electron channels in the SM due to interference between different helicity amplitudes of the virtual W boson. However, in a Δ -type observable,¹ $\Delta A_{\rm FB} = A_{\rm FB}^{\mu} - A_{\rm FB}^{e}$, where one considers the difference between the muon and electron channels, the SM contributions approximately cancel, except for a small residual effect due to the dependence on the muon mass close to its threshold. Furthermore, we find that the observable $\Delta A_{\rm FB}$ has reduced sensitivity to hadronic uncertainties in form factors. Therefore, any

¹Such observables were first proposed in Ref. [29] for angular analyses study on $B \to K^* \ell \ell$ decay.

deviation from the SM prediction for $\Delta A_{\rm FB}$ is likely due to NP effects. Recently, using the tables of Belle data from Ref. [15], an anomaly in $\Delta A_{\rm FB}$ was reported in Ref. [30]. This could be a signature of LFUV NP [30–32].

LFUV NP in the electron and muon sectors is tightly constrained by the measurement of the ratio of rates $R_{D^{(*)}}^{\mu e} \equiv \mathcal{B}(\bar{B} \to D^{(*)}\mu^-\bar{\nu}_{\mu})/\mathcal{B}(\bar{B} \to D^{(*)}e^-\bar{\nu}_{e})$ which is 1.04 ± 0.05 [14]. We restrict ourselves to NP scenarios in which a deviation of at most 3% from unity is allowed, which could be tested in the future. Even if the effects of LFUV NP are small in the ratios of decay rates, larger effects may be visible in the angular distributions.

In this paper, we discuss various solutions to explain the $\Delta A_{\rm FB}$ anomaly. The framework we use is based on a Monte Carlo generator to simulate a realistic experimental environment. Hence, in this work, we describe a newly developed Monte Carlo (MC) event generator tool [33] to allow simulations of the NP signatures in $B \to D^* \ell \nu$ arising due to the interference between SM and NP amplitudes. We employ our MC tool primarily to study semileptonic decays with a muon and electron in the final state. We assume that the electron decay mode is well described by the SM, but allow for NP contributions in the muon mode. Using this MC tool we generate results for three distinct scenarios with different NP couplings that are consistent with current data and can explain the $\Delta A_{\rm FB}$ anomaly, while remaining consistent with other constraints. Furthermore, using MC simulations we demonstrate that PHYS. REV. D 107, 015011 (2023)

 Δ -type observables, such as $\Delta A_{\rm FB}$ and ΔS_5 , eliminate most QCD uncertainties from form factors and allow for clean measurements of NP. We introduce correlated observables that improve the sensitivity to NP. We also discuss prospects for improved observables sensitive to NP couplings with the expected 50 ab⁻¹ of Belle II data, which seems to be ideally suited for this class of measurements. These measurements may also be possible at LHCb and other hadron collider experiments. We provide both integrated observables, for the benefit of current experimental analyses, and distributions of the observables as a function of q^2 . We also suggest experimental requirements on q^2 and on laboratory lepton momenta to optimize sensitivity to NP and reduce systematics.

The layout of the remainder of this article is as follows. In Sec. II, we discuss the theoretical basis of the full angular distribution for $\bar{B} \rightarrow D^* \ell^- \bar{\nu}$ in an effective theory framework. In Secs. III–V, we present the implementation of our NP MC tool, the signatures of and sensitivity to NP, respectively; finally, we conclude in Sec. VI.

II. THEORY

In the study of NP in charged-current semileptonic *B* decays it is useful to adopt an EFT framework. In an EFT description of the $b \rightarrow c\ell^-\bar{\nu}$ decays, one writes down all possible dimension-six four-quark operators at the scale of the *b*-quark mass. The effective Hamiltonian that describes SM and NP effects can be expressed as

$$\mathcal{H}_{\text{eff}} = \frac{G_F V_{cb}}{\sqrt{2}} \{ [(1+g_L)\bar{c}\gamma_{\alpha}(1-\gamma_5)b + g_R\bar{c}\gamma_{\alpha}(1+\gamma_5)b]\bar{\mu}\gamma^{\alpha}(1-\gamma_5)\nu_{\mu} + [g_S\bar{c}b + g_P\bar{c}\gamma_5b]\bar{\mu}(1-\gamma_5)\nu_{\mu} + g_T\bar{c}\sigma^{\alpha\beta}(1-\gamma_5)b\bar{\mu}\sigma_{\alpha\beta}(1-\gamma_5)\nu_{\mu} \} + \text{H.c.},$$
(1)

where the factors g_X , X = L, R, S, P, and T, are coupling constants that describe NP effects. As indicated earlier, we have only included LH neutrinos in this EFT, however, we have allowed for both LH and RH NP couplings.

Based on the effective Hamiltonian of Eq. (1), one can express the decay amplitude for the process $\bar{B} \to D^* (\to D\pi) \ell \bar{\nu}$ as [20,34],

$$\mathcal{M} = \frac{4G_F V_{cb}}{\sqrt{2}} \{ \langle D\pi | \bar{c}\gamma^{\mu} [(1+g_L)P_L + g_R P_R] b | \bar{B} \rangle (\bar{\ell}\gamma_{\mu} P_L \nu) + \langle D\pi | \bar{c}(g_{S_L} P_L + g_{S_R} P_R) b | \bar{B} \rangle (\bar{\ell} P_L \nu) + g_T \langle D\pi | \bar{c}\sigma^{\mu\nu} b | \bar{B} \rangle (\bar{\ell}\sigma_{\mu\nu} P_L \nu) \},$$
(2)

where $P_{R,L} = (1 \pm \gamma_5)/2$. This decay amplitude contains several hadronic matrix elements that describe the $\bar{B} \to D^* \to D\pi$ transitions through LH and RH scalar and vector currents, as well as a tensor current. The $D^* \to D\pi$ decay is mediated solely by the strong force, so that

$$\langle D\pi | D^*(k,\epsilon) \rangle = \epsilon \cdot (p_D - p_\pi),$$
(3)

where $p_{D(\pi)}$ is the four-momentum of the $D(\pi)$, $k = p_D + p_{\pi}$ is the four-momentum of the D^* and ϵ is its polarization. Note that these satisfy the on-shell condition $k \cdot \epsilon = 0$.

The remaining parts of the hadronic matrix elements that appear in Eq. (2) are (see, for example, [35])

$$\langle D^*(k,\epsilon)|\bar{c}\gamma_{\mu}b|\bar{B}(p)\rangle = -i\varepsilon_{\mu\nu\rho\sigma}\epsilon^{*\nu}p^{\rho}k^{\sigma}\frac{2V(q^2)}{m_B+m_{D^*}},\tag{4}$$

$$\langle D^*(k,\epsilon) | \bar{c}\gamma_{\mu}\gamma^5 b | \bar{B}(p) \rangle = \epsilon^*_{\mu} (m_B + m_{D^*}) A_1(q^2) - (p+k)_{\mu} (\epsilon^* \cdot q) \frac{A_2(q^2)}{m_B + m_{D^*}} - q_{\mu} (\epsilon^* \cdot q) \frac{2m_{D^*}}{q^2} [A_3(q^2) - A_0(q^2)],$$

$$(5)$$

$$\langle D^*(k,\epsilon)|\bar{c}\gamma^5 b|\bar{B}(p)\rangle = -(\epsilon^* \cdot q)\frac{2m_{D^*}}{m_b + m_c}A_0(q^2),\tag{6}$$

$$\langle D^*(k,\epsilon) | \bar{c}\sigma_{\mu\nu} b | \bar{B}(p) \rangle = \varepsilon_{\mu\nu\rho\sigma} \bigg\{ -\epsilon^{\rho*} (p+k)^{\sigma} T_1(q^2) + \epsilon^{\rho*} q^{\sigma} \frac{m_B^2 - m_{D^*}^2}{q^2} [T_1(q^2) - T_2(q^2)] \\ + 2 \frac{\epsilon^* \cdot q}{q^2} p^{\rho} k^{\sigma} \bigg[T_1(q^2) - T_2(q^2) - \frac{q^2}{m_B^2 - m_{D^*}^2} T_3(q^2) \bigg] \bigg\},$$

$$(7)$$

where *p* is the four-momentum of the *B* meson, *q* represents the four-momentum of the lepton-neutrino pair, while $m_{B(D^*)}$ represents the mass of the $B(D^*)$ meson. Here, *V*, A_0 , A_1 , A_2 , A_3 , T_1 , T_2 , and T_3 are the relevant form factors for a $\overline{B} \rightarrow V$ transition. The BGL [36], CLN [37], and HQET [38] parametrizations for these form factors are given in Appendix B. For the Levi-Civita tensor, $\varepsilon_{\mu\nu\rho\sigma}$, we use the convention $\varepsilon_{0123} = +1$.

For easy comparison with similar literature in the field, below we present an alternative notation and its connection to the notation used in this article. Following the presentation in Ref. [30], the effective Lagrangian that describes $b \rightarrow c\ell^-\bar{\nu}$ transitions can be written as

$$\mathcal{L} = -\frac{4G_F}{\sqrt{2}} \sum_i C_i \mathcal{O}_i + \text{H.c.}, \qquad (8)$$

where $i = V_L$, V_R , S_L , S_R , and T, and C_i represents the Wilson coefficient (WC) corresponding to the operator O_i . Note the negative sign added to this Lagrangian in order to obtain the correct sign for the SM term [see, for example, Eq. (20.90) in [39] with errata in [40]]. The WCs can be easily converted into the NP coupling constants that appear in Eq. (1) as follows.

$$C_{V_L} = 1 + g_L, \qquad C_{V_R} = g_R, \qquad C_{S_L} = g_S - g_P,$$

 $C_{S_R} = g_S + g_P, \qquad C_T = g_T.$ (9)

Note that only C_{V_L} has both SM and NP parts while all other WCs are NP only. Furthermore, for a $\bar{B} \rightarrow V$ transition, where V denotes a vector meson, the scalar matrix element $\langle V | \bar{q} b | B \rangle = 0$. As a consequence, the following condition must be imposed,

$$C_{S_R} + C_{S_L} = 2g_S = 0. \tag{10}$$

Thus, there are only four independent NP parameters that can be used to describe the decay $\bar{B} \rightarrow D^* \ell^- \bar{\nu}$ process, namely, g_L , g_R , g_P , and g_T . We will use the g_i parameters to describe the results and plots presented in this article.

One can now express the differential decay distribution for $\overline{B} \to D^*(\to D\pi)\ell^-\overline{\nu}$ as a function of four kinematic variables— q^2 and three helicity angles θ^*, θ_ℓ , and χ (see Fig. 1 for a schematic diagram defining these angles)—in the following form.

$$\frac{d^{4}\Gamma}{dq^{2}d\cos\theta^{*}d\cos\theta_{\ell}d\chi} = \frac{9}{32\pi} \left[(I_{1}^{s}\sin^{2}\theta^{*} + I_{1}^{c}\cos^{2}\theta^{*}) + (I_{2}^{s}\sin^{2}\theta^{*} + I_{2}^{c}\cos^{2}\theta^{*})\cos 2\theta_{\ell} \right]$$
$$+ I_{3}\sin^{2}\theta^{*}\sin^{2}\theta_{\ell}\cos 2\chi + I_{4}\sin 2\theta^{*}\sin 2\theta_{\ell}\cos \chi + I_{5}\sin 2\theta^{*}\sin \theta_{\ell}\cos \chi$$
$$+ (I_{6}^{c}\cos^{2}\theta^{*} + I_{6}^{s}\sin^{2}\theta^{*})\cos \theta_{\ell} + I_{7}\sin 2\theta^{*}\sin \theta_{\ell}\sin \chi$$
$$+ I_{8}\sin 2\theta^{*}\sin 2\theta_{\ell}\sin \chi + I_{9}\sin^{2}\theta^{*}\sin^{2}\theta_{\ell}\sin 2\chi \right],$$
(11)

where the 12 coefficients $I_i^{(s,c)}(q^2)$ (i = 1, ..., 9) can be expressed in terms of eight helicity amplitudes that in turn depend on the NP parameters g_L , g_R , g_P , and g_T . For brevity, the exact dependence of the coefficient functions, $I_i^{(s,c)}$ is given in Appendix A. The distribution

for the *CP*-conjugate process is obtained with the following transformation, $\theta_l \rightarrow \pi - \theta_l$ and $\chi \rightarrow \pi + \chi$. The various helicity amplitudes transform as $\mathcal{A}_{SP} \rightarrow -\bar{\mathcal{A}}_{SP}$, $\mathcal{A}_t \rightarrow -\bar{\mathcal{A}}_t$, $\mathcal{A}_{0(,T)} \rightarrow \bar{\mathcal{A}}_{0(,T)}$, $\mathcal{A}_{||(,T)} \rightarrow \bar{\mathcal{A}}_{||(,T)}$, $\mathcal{A}_{\perp(,T)} \rightarrow -\bar{\mathcal{A}}_{\perp(,T)}(\mathcal{A}_{\pm(,T)} \rightarrow \bar{\mathcal{A}}_{\mp(,T)})$ leading to the angular coefficients transformations $I_{1,2,3,4,7}^{(a)} \rightarrow \bar{I}_{1,2,3,4,7}^{(a)}$ and $I_{5,6,8,9}^{(a)} \rightarrow -\bar{I}_{5,6,8,9}^{(a)}$.² Note that if one writes $\mathcal{A} = |A|e^{i\phi+i\delta}$, then $\bar{\mathcal{A}} = |A|e^{-i\phi+i\delta}$, where ϕ is the *CP*-violating weak phase and δ is the *CP*-conserving strong phase.

The full phase space for the $\bar{B} \rightarrow D^* \ell'^- \bar{\nu}$ decay is obtained by varying the kinematic variables over their allowed ranges which are as follows: $m_{\ell}^2 \leq q^2 \leq m_B^2 - m_{D^*}^2$, $0 \leq \theta_{D^*,\ell} \leq \pi$, and $0 \leq \chi \leq 2\pi$. One can now construct several observables by integrating the distribution of Eq. (11) over one or more of these kinematic variables. The first of these is the differential decay distribution as a function of q^2 , constructed by integrating over the full range of allowed values for all three helicity angles.

$$\frac{d\Gamma}{dq^2} = \frac{1}{4} \left[3I_1^c - I_2^c + 2(3I_1^s - I_2^s) \right].$$
(12)

Next, one can construct double-differential decay distributions as functions of q^2 and one other angle variable at a time, obtained by integrating over the other two angles.

$$\frac{d^2\Gamma}{dq^2 d\cos\theta^*} = \frac{3}{4} \frac{d\Gamma}{dq^2} [2F_L^{D^*}(q^2)\cos^2\theta^* + F_T^{D^*}(q^2)\sin^2\theta^*],$$
(13)

$$\frac{d^2\Gamma}{dq^2d\cos\theta_\ell} = \frac{d\Gamma}{dq^2} \left(\frac{1}{2} + A_{\rm FB}\cos\theta_\ell + \frac{1 - 3\tilde{F}_L^\ell}{4}\frac{3\cos^2\theta_\ell - 1}{2}\right), \quad (14)$$
$$\frac{d^2\Gamma}{dq^2d\chi} = \frac{1}{2\pi}\frac{d\Gamma}{dq^2}(1 + S_3\cos2\chi + S_9\sin2\chi), \quad (15)$$

where $F_{L(T)}^{D^*}(q^2)$ is the longitudinal (transverse) polarization of the D^* , $A_{\rm FB}$ is the charged-lepton forward-backward asymmetry, and S_9 is a triple-product asymmetry. The coefficient functions that appear in Eq. (15) can be expressed in terms of the angular coefficients, $I_i^{(s,c)}$, as follows.

$$F_L^{D^*}(q^2) = 1 - F_T^{D^*}(q^2) = \frac{3I_1^c - I_2^c}{3I_1^c - I_2^c + 2(3I_1^s - I_2^s)},$$
 (16)

A

$$\Lambda_{\rm FB}(q^2) = \frac{3}{2} \frac{2I_6^s + I_6^c}{3I_1^c - I_2^c + 2(3I_1^s - I_2^s)},\tag{17}$$

$$\tilde{F}_{L}^{\ell}(q^{2}) = \frac{I_{1}^{c} - 3I_{2}^{c} + 2(I_{1}^{s} - 3I_{2}^{s})}{3I_{1}^{c} - I_{2}^{c} + 2(3I_{1}^{s} - I_{2}^{s})},$$
(18)

$$S_3(q^2) = \frac{4I_3}{3I_1^c - I_2^c + 2(3I_1^s - I_2^s)},$$
 (19)

$$S_9(q^2) = \frac{4I_9}{3I_1^c - I_2^c + 2(3I_1^s - I_2^s)}.$$
 (20)

Note that there are additional observables that can be extracted from data by performing asymmetric integrals over more than one angles. We discuss some such observables in Sec. IV.

III. NEW-PHYSICS IMPLEMENTATION IN EVTGEN

We implement the preceding discussion in the EvtGen MC simulation framework as the new BTODSTARLNUNP decay model. This NP generator, BTODSTARLNUNP, can run either in a standalone mode or be integrated into a software framework of a *B*-physics experiment. The model includes SM contributions, various NP parameters as well as their interference. The model takes the NP parameters $\delta C_{V_L} \equiv g_L, C_{V_R}, C_{S_L}, C_{S_R}$, and C_T as inputs. The user specifies the NP parameters keeping in mind that the scalar coefficients (C_{S_t}, C_{S_R}) are related to each other by Eq. (10). Each of these parameters can take complex values as inputs and are entered in the user decay file. The default value for each parameter has been set to zero so that when no value is specified for these parameters the code returns SM results. Below we present an example of a user decay file to illustrate the usage of the NP MC generator.

first argument is cartesian(0) or polar(1) representation of NP coefficients which

are three consecutive numbers {id, Re
(C), Im(C) } or {coeff id, |C|, Arg(C) }

id==0 \delta C_VL-left-handed vector
coefficient change from SM

id==1 C_VR-right-handed vector coefficient ## id==2 C_SL-left-handed scalar coefficient

id==3 C_SR-right-handed scalar
coefficient

id==4 C_T-tensor coefficient
Decay B0

²Our convention is similar to the LHCb convention for the $B^0 \to K^{*0}\ell^+\ell^-$ decay where θ_ℓ is defined as the angle between $K^{*0}(\bar{K}^{*0})$ and $\mu^+(\mu^-)$ for the $B^0(\bar{B}^0)$ decay leading to the transformations $I_{1,2,3,4,5,6}^{(a)} \to \bar{I}_{1,2,3,4,5,6}^{(a)}$ and $I_{7,8,9} \to -\bar{I}_{7,8,9}$ for *CP* conjugation with $\chi \to 2\pi - \chi$ [41]. Alternatively, when θ_ℓ is defined as the angle between $K^{*0}(\bar{K}^{*0})$ and the lepton ℓ^- for the $B^0(\bar{B}^0)$ decay while χ is the angle between the $K^{\pm}\pi^{\mp}$ and the $\ell^+\ell^-$ planes in both cases, the angular coefficients transform as $I_{1,2,3,4,7}^{(a)} \to \bar{I}_{1,2,3,4,7}^{(a)}$ and $I_{5,6,8,9}^{(a)} \to -\bar{I}_{5,6,8,9}^{(a)}$ for the *CP*-conjugate process with $\theta_\ell \to \theta_\ell - \pi$ and $\chi \to -\chi$ [42,43]. Note that, in all of these conventions, including ours, the $\frac{d^4(\Gamma+\bar{\Gamma})}{dq^2d\cos\theta' d\cos\theta_\ell d\chi}$ distribution for the untagged decay retains the contribution from the "true" *CP*-violating terms [44,45].

```
## B0 -> D*- e+ nu_e is generated with the
Standard Model only
1 D*- e+ nu_e BTODSTARLNUNP;
Enddecay
Decay anti-B0
## anti-B0 -> D*+ mu- anti-nu_mu is gen-
erated with the addition of New Physics
1 D*+ mu- anti-nu_mu BTODSTARLNUNP 0 0
0.06 0 1 0.075 0 2 0 -0.2 3 0 0.2;
Enddecay
End
```

To generate NP the user inputs several arguments in the user decay file. The first of these specifies whether the remaining arguments are to be entered in Cartesian (0) or polar (1) coordinate system. Next, the user enters sets of three values. The first specifies the type of NP coupling $(\delta C_{V_L}, C_{V_R}, C_{S_L}, C_{S_R}, \text{ and } C_T)$, while the second and third represent the real and imaginary parts in Cartesian coordinates, or magnitude and complex phase in polar coordinates. In the above example we have shown how the user can generate events for the SM as well as for a specific

NP scenario which in our case is NP scenario 2. A complete version of the NP MC tool with an implementation of the BTODSTARLNUNP decay model can be found in Ref. [46].

IV. SIGNATURES OF NEW PHYSICS

The ratios of branching fractions as well as the differential q^2 distributions have limited sensitivity to NP for $b \rightarrow c\ell\nu$, $\ell = e$, μ , which receive tree-level contributions in the SM and are hence unsuppressed. In contrast, angular observables have much better sensitivity to the interference between SM and NP. The optimal sensitivity to NP can be obtained by studying these angular observables as functions of q^2 . We will examine four angular asymmetries as functions of q^2 to make predictions for our NP scenarios, $A_{\rm FB}$, S_3 , S_5 , and S_7 . $A_{\rm FB}$ and S_3 are previously defined in Sec. II, while S_5 and S_7 are the coefficients of sin $\theta_{\ell} \sin 2\theta^* \cos \chi$ and sin $\theta_{\ell} \sin 2\theta^* \sin \chi$, respectively. These asymmetries can be constructed from the full angular distribution of Eq. (11) through asymmetric integrals shown below.

$$A_{FB}(q^2) = \left(\frac{d\Gamma}{dq^2}\right)^{-1} \left[\int_0^1 - \int_{-1}^0\right] d\cos\theta_\ell \frac{d^2\Gamma}{d\cos\theta_\ell dq^2},\tag{21}$$

$$S_{3}(q^{2}) = \left(\frac{d\Gamma}{dq^{2}}\right)^{-1} \left[\int_{0}^{\pi/4} - \int_{\pi/4}^{\pi/2} - \int_{\pi/2}^{3\pi/4} + \int_{3\pi/4}^{\pi} + \int_{\pi}^{5\pi/4} - \int_{5\pi/4}^{3\pi/2} - \int_{3\pi/2}^{7\pi/4} + \int_{7\pi/4}^{2\pi} \right] d\chi \frac{d^{2}\Gamma}{dq^{2}d\chi},$$
(22)

$$S_5(q^2) = \left(\frac{d\Gamma}{dq^2}\right)^{-1} \left[\int_0^{\pi/2} -\int_{\pi/2}^{\pi} -\int_{\pi}^{3\pi/2} +\int_{3\pi/2}^{2\pi} d\chi \left[\int_0^1 -\int_{-1}^0\right] d\cos\theta^* \frac{d^3\Gamma}{dq^2 d\cos\theta^* d\chi},$$
(23)

$$S_{7}(q^{2}) = \left(\frac{d\Gamma}{dq^{2}}\right)^{-1} \left[\int_{0}^{\pi} - \int_{\pi}^{2\pi} d\chi \left[\int_{0}^{1} - \int_{-1}^{0} d\cos\theta^{*} \frac{d^{3}\Gamma}{dq^{2}d\cos\theta^{*}d\chi}\right]$$
(24)

To extract these asymmetries from data, we calculate the integrals in Eqs. (21)–(24) from binned distributions of the appropriate angular variables. For example, consider S_5 . This distribution involves asymmetric integrals over both $\cos \theta^*$ and χ . For a given bin of q^2 , we first divide the events into χ bins of size $\pi/2$. In each of these bins, we then divide the events into $\cos \theta^*$ bins of size 1. This gives us 8 bins corresponding to the various terms of Eq. (23), which we will label N_i with i = 1, 2, ..., 8. To find the value of S_5 for a given q^2 bin, we then combine the N_i 's in the same way as the integrals in Eq. (23), normalized by $\sum_{i=1}^{8} N_i$.

When generating our predictions, we used $\Delta A_{\text{FB}} = A_{\text{FB}}(B \rightarrow D^* \mu \nu) - A_{\text{FB}}(B \rightarrow D^* e \nu)$, $\Delta S_3 = S_3(B \rightarrow D^* \mu \nu) - S_3(B \rightarrow D^* e \nu)$, and $\Delta S_5 = S_5(B \rightarrow D^* \mu \nu) - S_5(B \rightarrow D^* e \nu)$, where the electron mode has been generated with the SM

only while the muon mode contains both SM and NP contributions. These are Δ -type observables as defined above, which eliminate most of the QCD uncertainties in the form factors, allowing for a clean measurement of LFUV NP. The asymmetry S_7 is always zero in the SM, and therefore was not recast into the form of a Δ observable. The NP dependences of $A_{\rm FB}$, S_3 , S_5 , and S_7 are given in Table II. Note that these dependencies have different weights, which are dependent on q^2 . For all theory plots presented here, we have only used uncorrelated central values of the form-factor parameters as listed in Tables VI and VII. We verify that the Δ variables have minimal dependence on form factors. As a test, we consider BGL [36], CLN [37], and HQET [38] form-factor parametrizations. There are also other formfactor models [47,48]. Unless otherwise stated, we use

Observable	Angular function	NP dependence	m_{ℓ} suppression order
A _{FB}	$\cos heta_\ell$	$\operatorname{Re}[g_T g_P^*]$	$\mathcal{O}(1)$
		$ \begin{array}{c} \operatorname{Re}[(1+g_{L}-g_{R})(1+g_{L}+g_{R})^{*}] \\ \operatorname{Re}[(1+g_{L}-g_{R})g_{P}^{*}] \\ \operatorname{Re}[g_{T}(1+g_{L}-g_{R})^{*}] \\ \end{array} $	${\cal O}(m_\ell/\sqrt{q^2})$
		$\frac{\text{Re}[g_T(1+g_L+g_R)^*]}{ 1+g_L-g_R ^2}$	${\cal O}(m_\ell^2/q^2)$
<i>S</i> ₃	$\sin^2\theta^*\sin^2\theta_{\ell'}\cos 2\chi$	$egin{array}{ll} 1+g_L+g_R ^2 \ 1+g_L-g_R ^2 \ g_T ^2 \end{array}$	$\mathcal{O}(1), \mathcal{O}(m_{\ell}^2/q^2)$
<i>S</i> ₅	$\sin 2\theta^* \sin \theta_{\mathscr{C}} \cos \chi$	$rac{ ext{Re}[g_T g_P^*]}{ 1+g_L-g_R ^2} \ ext{Re}[(1+g_L-g_R)g_P^*] \ ext{Re}[g_T(1+g_L-g_R)^*] \ ext{Re}[g_T(1+g_L-g_R)^*] \ ext{Re}[g_T(1+g_L+g_R)^*] \ ext{Re}[g_T(1+g_R)^*] \ ext{Re}[g$	$\mathcal{O}(1) \ \mathcal{O}(1), \mathcal{O}(m_{\ell}^2/q^2) \ \mathcal{O}(m_{\ell}/\sqrt{q^2})$
S_	$\sin 2\theta^* \sin \theta_2 \sin \eta$	$ g_T ^2$ Im $[a, a^*]$	$\mathcal{O}(m_{\ell}^2/q^2)$
57	$\sin 2\theta \sin \theta_{\ell} \sin \chi$	$Im[gPgT] Im[(1 + g_L + g_R)g_P^*] Im[(1 + g_L - g_R)g_T^*]$	$\mathcal{O}(n)$ $\mathcal{O}(m_\ell/\sqrt{q^2})$
		$Im[(1 + g_L - g_R)(1 + g_L + g_R)^*]$	${\cal O}(m_\ell^2/q^2)$

TABLE II. Angular functions corresponding to angular observables A_{FB} , S_3 , S_5 , and S_7 alongside NP parameters that contribute to each. The dependence on NP parameters has been separated into different orders of $m_{\ell}/\sqrt{q^2}$.

the CLN parametrization of the hadronic form factors as the default in our plots.

V. NEW-PHYSICS SENSITIVITY AND RESULTS

The q^2 distribution alone has little sensitivity to NP, as shown in Fig. 2. On the other hand, angular asymmetries as functions of q^2 are quite sensitive to NP couplings. In particular, the angular asymmetries in the angle θ_{ℓ} and χ can be promising probes of NP as shown in Fig. 2. In this figure, we have used the CLN parametrization to test that our Monte Carlo generator correctly implements the theoretical expressions. However, the angular asymmetries remain quite sensitive to form-factor uncertainties. As an example, the uncertainty in the predictions for $A_{\rm FB}^{\mu}$ in the SM with four different form-factor parametrizations is shown below. To address this issue we consider differences between angular asymmetries in the muon and electron channels using Δ observables. Later in this section, using $\Delta A_{\rm FB}$ as an example, we show that the predictions for the Δ observables are robust against formfactor uncertainties using the same four form-factor parametrizations. In the SM the form-factor uncertainties cancel effectively in the Δ observables while with NP the cancellation is slightly less effective as the NP violates lepton universality.

From our initial scan, we cannot reproduce the experimental- ΔA_{FB} anomaly with a single NP coupling. Instead, we consider scenarios with several NP couplings. In order

to match $\Delta A_{\rm FB}$ from Ref. [30], we require a g_R NP coupling. In order to maintain the LFU BR constraint we also need to add a q_L NP coupling that is comparable to g_R . In addition, it is also possible to include a g_P contribution, but in order to satisfy the constraints it must be imaginary. We also found that negative or complex values for g_L and g_R are ruled out by these constraints. Figure 3 shows the region of parameter space in the $g_L - g_R$ plane that is excluded by $\frac{\mathcal{B}(B \to D^* \mu \nu)}{\mathcal{B}(B \to D^* e \nu)} = 1.00 \pm 0.03(0.06)$ in red and the region in blue excludes $\Delta A_{\rm FB} = 0.0349 \pm$ 0.0089(0.0178) when the error is taken in the 68%(95%)C.L. Further, we observe that an additional nonzero imaginary pseudoscalar interaction strength produces an upward shift in the allowed region of g_R while g_L remains almost the same as shown in the right plot of Fig. 3. In this section we provide results corresponding to the three distinct NP scenarios indicated in Table III chosen with the above constraints.

To optimize sensitivity, it is important to measure the Δ observables as functions of q^2 . Using the benchmark scenarios above, we show in Fig. 4 the predictions for the Δ observables. As discussed earlier these observables are sensitive to NP couplings and have much reduced dependence on form-factor uncertainties. In the figure, the SM expectations for these quantities are shown using solid black curves. In addition to the two Δ observables, $\Delta A_{\rm FB}$ and ΔS_5 , Fig. 4 also shows the q^2 dependence of the observable ΔS_3 and S_7 . S_7 represents an angular



FIG. 2. Distribution of $\overline{B} \to D^* \ell^- \overline{\nu}$ events as functions of (clockwise from top left) q^2 , $\cos \theta^*$, χ , and $\cos \theta_{\ell}$. Theory predictions are shown for the SM (solid black curve) and for NP scenario 2 (dashed red curve). EvtGen data are shown for NP scenario 2 (solid red histogram). Each plot is fully integrated over three of the four kinematic variables. The q^2 range is divided into 23 equal bins, to reflect the expected resolution of experimental measurements. The angular bins are chosen to be sufficiently fine to compare MC data to the theory. The $\cos \theta$ ranges are divided into 15 equal bins, and the χ range, being twice as large as the θ ranges, is divided into twice as many bins.

asymmetry in sin χ , where χ is the azimuthal angle between the decay planes. This is a *CP*-odd triple-product asymmetry, which is predicted to be identically zero in the SM for any q^2 . We find that NP scenarios with an imaginary g_P are able to produce a small nonzero signal in the q^2 distribution of S_7 as shown in Fig. 4.

The observable S_3 is the coefficient of $\cos 2\chi$ term in the angular distribution and can be extracted using the



FIG. 3. Allowed parameter space in g_L and g_R , with $g_P = 0$ and 0.6*i*. The two constraints used are that the branching ratio of the muon and electron modes must be unity within 3%, and ΔA_{FB} must be consistent with the value found in Ref. [30]. Nonzero values of g_P produce similar plots, with the allowed region in g_R shifting upwards. This exercise also showed that imaginary values of g_L and g_R are not consistent with these constraints.

TABLE III. Values of NP coefficients for three distinct NP scenarios considered in this paper and used for generating the results presented in this section.

	g_L	g_R	g_P
Scenario 1:	0.06	0.075	0.2 <i>i</i>
Scenario 2:	0.08	0.090	0.6 <i>i</i>
Scenario 3:	0.07	0.075	0

asymmetric integral defined in Eq. (22). Although ΔS_3 is close to zero in the SM, NP can produce a nonzero ΔS_3 in the q^2 range as shown in the lower left plot of Fig. 4. In Fig. 5, using ΔA_{FB} as an example, we show that the predictions for the Δ observables are largely independent of form-factor parametrization and the uncertainties of the form-factor parameters.

Note that due to lepton mass and helicity effects, ΔA_{FB} is negative in the low q^2 region even in the SM. In fact, at the lower momentum transfer threshold, i.e., in the limit $q^2 \rightarrow m_{\ell}^2$, the forward-backward asymmetry $A_{\text{FB}}^{\ell} \rightarrow -1$ which is seen as a large dip in the q^2 distribution as shown in Fig. 5. Hence, for the best experimental sensitivity to NP,

we advocate a necessary low q^2 cut of 1.14 GeV² on such observables in order to predict them unambiguously.

In addition, in order to improve systematic uncertainties from lepton identification efficiencies, we recommend using the same laboratory momentum cutoff for both $\ell = e$ and μ channels (see, for example, [49]). In order to define the detector acceptance we will represent the magnitude of the transverse momentum of particle x in the lab frame by $|\vec{p}_{T,x}|$ and the ratio of the *z* component of the momentum over the total momentum as $\cos \alpha$. We use the Belle II acceptances of $|\vec{p}_{T,\ell}| > 0.8$ GeV for the lepton momenta, $|\vec{p}_{T,\pi}| > 0.1$ GeV for the slow pion momenta, and $-0.866 < \cos \alpha < 0.956$ for all final state particles. The theoretical predictions and uncertainties for these observables integrated over the range of $q^2 \in$ $[1.14 \text{ GeV}^2, (m_B - m_{D^*})^2]$ using the BGL parametrization are displayed in Table IV both for the SM and the specific NP scenarios listed in Table III. One can see that the theoretical uncertainties are less than $\sim 5\%$ for both the SM and NP predictions of all integrated observables except $\langle \Delta S_3 \rangle$ which has a ~15% uncertainty. We also show the variation of the expected statistical uncertainties as a function of the total integrated luminosity for present



FIG. 4. ΔA_{FB} , ΔS_5 , ΔS_3 , and S_7 plotted as functions of q^2 for different values of NP coefficients. Here we have used the CLN parameterizations of the form factors. The NP parameters were chosen so that the ratio of semileptonic branching fractions is constrained to be within 3% of unity, as well as the ΔA_{FB} for the full q^2 range is within the interval 0.0349 \pm 0.0089. EvtGen data for NP scenario 2 ($g_L = 0.08$, $g_R = 0.09$, $g_P = 0.6i$) generated with 10⁷ events (anticipated Belle II statistics) are shown as points with error bars. Theory curves are presented for all three NP scenarios: Scenario 1 is dot-dashed blue, scenario 2 is dashed red, and scenario 3 is dotted blue.



FIG. 5. A_{FB}^{μ} (left plot) in the SM and $\Delta A_{FB} = A_{FB}^{\mu} - A_{FB}^{e}$ (right plot) for different form-factor parametrizations. The left plot shows the SM predictions for various form-factor parametrizations, while the right plot demonstrates the effects of form-factor uncertainties on ΔA_{FB} in NP scenario 1 ($g_L = 0.06$, $g_R = 0.075$, and $g_P = 0.2i$). The solid black curve in the right plot represents the SM prediction for both CLN and HQET (2/1/0) parametrizations. Note that the vertical scale of the right plot is approximately a factor of 10 smaller than that of the left plot. Note also the large negative value at the low q^2 limit. A cutoff of 1.14 GeV² is chosen to avoid this. Note that for the SM and NP 3, $\langle S_7 \rangle$ is exactly zero and are not distinguishable.

and future experimental datasets in Fig. 6 using MC simulations.

Initially, experiments will measure integrated Δ observables. As statistics improve, they will proceed to coarsebinned measurements, as shown, for example, in Fig. 7. At high statistics, unbinned fits to angular observables will be performed, as shown, for example, in Fig. 4.

Furthermore, from Fig. 4 we see that NP couplings produce correlated signatures of deviations from the SM in multiple Δ observables, such as ΔA_{FB} and ΔS_5 . As shown in Fig. 8, the size of the effect on Δ observables is determined primarily by g_R . In this plot, we have varied the NP parameter g_L between 0 and 0.2 for fixed values of g_R . In the presence of NP there are strong correlations between the Δ observables ΔA_{FB} , ΔS_5 , and ΔS_3 . Therefore, if an experimental signal in ΔA_{FB} is observed, it should be accompanied by an observation of nonzero ΔS_5 and ΔS_3 . Conversely, if a nonzero ΔS_5 is observed, there must also be a nonzero ΔA_{FB} . In the absence of a tensor coupling, a correlation with ΔS_3 is also required.

TABLE IV. Theoretical predictions of integrated ΔA_{FB} , ΔS_3 , ΔS_5 , and S_7 for SM and each NP scenario using the BGL formfactor parametrization with estimated theoretical uncertainties. Note that for SM and NP 3, $\langle S_7 \rangle$ is exactly zero as all associated couplings are real.

	$\langle \Delta A_{ m FB} angle \%$	$\langle \Delta S_3 \rangle ~\%$	$\langle \Delta S_5 angle \ \%$	$\langle S_7 \rangle \times 10^{-3}$
SM:	-0.23 ± 0.02	$0.052^{+0.004}_{-0.002}$	0.044 ± 0.005	0
NP 1:	2.7 ± 0.1	$0.87^{+0.12}_{-0.07}$	$2.21_{-0.09}^{+0.08}$	$0.56^{+0.03}_{-0.04}$
NP 2:	2.8 ± 0.1	$1.27^{+0.13}_{-0.09}$	$2.25_{-0.10}^{+0.08}$	$1.69^{+0.09}_{-0.10}$
NP 3:	2.8 ± 0.1	$0.83_{-0.04}^{+0.12}$	$2.24_{-0.09}^{+0.08}$	0

For the benchmark scenarios described above, we have also checked the constraints from the longitudinal polarization fraction of the D^* meson, F_L , and another angular observable \tilde{F}_L , which are proportional to the coefficients of the $\cos^2 \theta^*$ and $\cos^2 \theta_{\ell}$ terms in the angular distribution, respectively. These quantities were extracted for the first time by [30] using the binned *CP*-averaged differential decay distribution data provided by Belle [10]. They obtain a *CP*-averaged SM prediction for the integrated $\langle \Delta F_L \rangle$ and $\langle \Delta \tilde{F}_L \rangle$ to be $(5.43 \pm 0.36) \times 10^{-4}$ and $(-5.20 \pm 0.30) \times$ 10^{-3} respectively. By fitting the data, they also report $\langle \Delta F_L \rangle^{exp} = -0.0065 \pm 0.0059$ and $\langle \Delta \tilde{F}_L \rangle^{exp} = -0.0107 \pm$ 0.0142. We have verified that our benchmark values satisfy these experimental bounds within a 1σ confidence interval.



FIG. 6. Expected statistical uncertainties for the four observables at 1, 5, 50, and 250 ab^{-1} of Belle II data. These expected uncertainties were found using the BTODSTARLNUNP MC simulation.



FIG. 7. Coarse-binned distributions of ΔA_{FB} and ΔS_5 versus q^2 . The horizontal axis spans the allowed range for q^2 which has been divided into three bins. The vertical lines at 4 and 8 GeV² indicate the other edges of these bins. The central values are calculated from theory, and the error bars indicate statistical uncertainties taken from MC simulation with an integrated luminosity of 50 ab⁻¹. The NP1 and NP3 predictions have been offset from the center of each bin for clarity.



FIG. 8. Correlations between $\langle \Delta A_{FB} \rangle$, $\langle \Delta S_3 \rangle$, and $\langle \Delta S_5 \rangle$ in NP scenarios. For each point, g_L is varied between 0 and 0.2 (light to dark in the color scale as depicted in the bar legend; applies for each value of g_R), with $g_R = 0$, 0.1, or 0.2, which are representative values in the allowed range, and $g_P = 0$. All points for which only g_L is nonzero return the SM values of the three observables.

VI. CONCLUSIONS

Motivated by the $\Delta A_{\rm FB}$ anomaly in $\bar{B} \to D^{*+} \mu^- \bar{\nu}$ decay, which could be a sign of physics beyond the Standard Model [30], we have developed a new Monte Carlo new physics (NP) generator tool for $B \to D^* \ell \nu_\ell$ with $\ell = e, \mu, \tau$ in the EvtGen framework [46]. The full theoretical description for the effective basis we use to parametrize NP as well as the different angular asymmetries has been comprehensively discussed in this article. We used this tool to examine signatures of NP, which are consistent with current data and with the hints of NP in $B \rightarrow D^* \mu \nu_{\mu}$ assuming that the decay $B \to D^* e \nu_e$ is well described by the SM. We found that the angular asymmetries, A_{FB} , S_5 , S_3 , and S_7 , which can be extracted from the fully reconstructed angular distribution, are sensitive to new physics. With current experimental constraints, we show the part of the q_i NP parameter space that is still allowed (see Fig. 3).

We introduce the Δ observables, which are obtained by taking the differences between the observables for the muon and the electron modes, in order to avoid theory uncertainties due to form factors, which might obscure signals of NP. We suggest experimental requirements on q^2 and lepton momenta in order to increase sensitivity to NP and reduce systematics. We identify $\Delta A_{\rm FB}$ and ΔS_5 as the most powerful probes of NP with little sensitivity to form-factor uncertainties; this is shown in Fig. 5. We also observe that correlated signatures of NP in multiple observables such as $\Delta A_{\rm FB}$ and ΔS_5 are required to confirm the presence of NP (see Fig. 4.) Therefore, if a NP signal for $\Delta A_{\rm FB}$ is observed in future experiments, it must be accompanied by a corresponding signal in ΔS_5 both in the integrated variable and the q^2 distribution. We calculate integrated observables and plot coarse binned expectations for $\Delta A_{\rm FB}$ and ΔS_5 , as well as correlations between the two. The NP signatures described here are ideally suited for Belle II at 1, 5, 50, and 250 ab^{-1} and might also be explored at hadron collider experiments.

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Note added.—This paper is an improved version of Ref. [33], which was submitted to the U.S. Community Summer Study on the Future of Particle Physics (Snowmass 2021) but will not appear in the final proceedings. Improvements include calculations of correlations between several observables and discussions of prospects for NP-sensitive observables with several benchmark values of Belle II integrated luminosity.

APPENDIX A: ANGULAR COEFFICIENTS

The angular distribution of $\overline{B} \to D^* \ell^- \overline{\nu}$ presented in Eq. (11) contains 12 coefficients labeled $I_i^{(s,c)}$ with i = 1, ..., 9. The full list of angular coefficients are presented below as functions of eight helicity amplitudes, $\mathcal{A}_{SP}, \mathcal{A}_t, \mathcal{A}_0, \mathcal{A}_{||}, \mathcal{A}_{\perp}, \mathcal{A}_{0T}, \mathcal{A}_{||,T}$, and $\mathcal{A}_{\perp,T}$. These helicity amplitudes depend on hadronic form factors as well as NP coefficients. The form of the eight helicity amplitudes are given in Appendix B.

$$I_{i}^{(s,c)} = \frac{G_{F}^{2}|V_{cb}|^{2}(q^{2}-m_{\ell}^{2})^{2}|p_{D^{*}}|}{192\pi^{3}m_{B}^{2}q^{2}}\mathcal{B}(D^{*}\to D\pi)\tilde{I}_{i}^{(s,c)},\tag{A1}$$

$$\tilde{I}_{1}^{c} = 4\left(|\mathcal{A}_{SP}|^{2} + \frac{m_{\ell}^{2}}{q^{2}}|\mathcal{A}_{t}|^{2}\right) + 2\left(1 + \frac{m_{\ell}^{2}}{q^{2}}\right)(|\mathcal{A}_{0}|^{2} + 16|\mathcal{A}_{0,T}|^{2}) + 8\frac{m_{\ell}}{\sqrt{q^{2}}}\{\operatorname{Re}[\mathcal{A}_{t}\mathcal{A}_{SP}^{*}] - 4\operatorname{Re}[\mathcal{A}_{0}\mathcal{A}_{0,T}^{*}]\}, \quad (A2)$$

$$\widetilde{I}_{1}^{s} = \left\{ \frac{3}{2} (|\mathcal{A}_{||}|^{2} + |\mathcal{A}_{\perp}|^{2}) + 8(|\mathcal{A}_{||,T}|^{2} + |\mathcal{A}_{\perp,T}|^{2}) \right\} - 16 \frac{m_{\ell}}{\sqrt{q^{2}}} \{ \operatorname{Re}[\mathcal{A}_{||}\mathcal{A}_{||,T}^{*}] + \operatorname{Re}[\mathcal{A}_{\perp}\mathcal{A}_{\perp,T}^{*}] \} \\
+ \frac{m_{\ell}^{2}}{q^{2}} \left\{ \frac{1}{2} (|\mathcal{A}_{||}|^{2} + |\mathcal{A}_{\perp}|^{2}) + 24(|\mathcal{A}_{||,T}|^{2} + |\mathcal{A}_{\perp,T}|^{2}) \right\},$$
(A3)

$$\check{T}_{2}^{c} = -2\left(1 - \frac{m_{\ell}^{2}}{q^{2}}\right)\{|\mathcal{A}_{0}|^{2} - 16|\mathcal{A}_{0,T}|^{2}\},\tag{A4}$$

$$\tilde{I}_{2}^{s} = \frac{1}{2} \left(1 - \frac{m_{\ell}^{2}}{q^{2}} \right) \{ (|\mathcal{A}_{||}|^{2} + |\mathcal{A}_{\perp}|^{2}) - 16(|\mathcal{A}_{||,T}|^{2} + |\mathcal{A}_{\perp,T}|^{2}) \},$$
(A5)

$$\tilde{I}_{3} = -\left(1 - \frac{m_{\ell}^{2}}{q^{2}}\right) \{ (|\mathcal{A}_{||}|^{2} - |\mathcal{A}_{\perp}|^{2}) - 16(|\mathcal{A}_{||,T}|^{2} - |\mathcal{A}_{\perp,T}|^{2}) \},\tag{A6}$$

$$\tilde{I}_{4} = \sqrt{2} \left(1 - \frac{m_{\ell}^{2}}{q^{2}} \right) \{ 16 \operatorname{Re}[\mathcal{A}_{0,T}\mathcal{A}_{||,T}^{*}] - \operatorname{Re}[\mathcal{A}_{0}\mathcal{A}_{||}^{*}] \},$$
(A7)

$$\tilde{I}_{5} = 2\sqrt{2} \bigg\{ (\operatorname{Re}[\mathcal{A}_{0}\mathcal{A}_{\perp}^{*}] + 4\operatorname{Re}[\mathcal{A}_{\parallel,T}\mathcal{A}_{SP}^{*}]) + \frac{m_{\ell}^{2}}{q^{2}} (16\operatorname{Re}[\mathcal{A}_{0,T}\mathcal{A}_{\perp,T}^{*}] - \operatorname{Re}[\mathcal{A}_{\parallel}\mathcal{A}_{r}^{*}]) \\ + \frac{m_{\ell}}{\sqrt{q^{2}}} (4\operatorname{Re}[\mathcal{A}_{\parallel,T}\mathcal{A}_{t}^{*}] - 4\operatorname{Re}[\mathcal{A}_{0}\mathcal{A}_{\perp,T}^{*}] - 4\operatorname{Re}[\mathcal{A}_{0,T}\mathcal{A}_{\perp}^{*}] - \operatorname{Re}[\mathcal{A}_{\parallel}\mathcal{A}_{SP}^{*}]) \bigg\},$$
(A8)

$$\tilde{I}_{6}^{c} = 32 \operatorname{Re}[\mathcal{A}_{0,T}\mathcal{A}_{SP}^{*}] + \frac{m_{\ell}}{\sqrt{q^{2}}} \{ 32 \operatorname{Re}[\mathcal{A}_{0,T}\mathcal{A}_{t}^{*}] - 8 \operatorname{Re}[\mathcal{A}_{0}\mathcal{A}_{SP}^{*}] \} - 8 \frac{m_{\ell}^{2}}{q^{2}} \operatorname{Re}[\mathcal{A}_{0}\mathcal{A}_{t}^{*}],$$
(A9)

$$\tilde{I}_{6}^{s} = -4\operatorname{Re}[\mathcal{A}_{\parallel}\mathcal{A}_{\perp}^{*}] + 16\frac{m_{\ell}}{\sqrt{q^{2}}}\{\operatorname{Re}[\mathcal{A}_{\parallel}\mathcal{A}_{\perp,T}^{*}] + \operatorname{Re}[\mathcal{A}_{\parallel,T}\mathcal{A}_{\perp}^{*}]\} - 64\frac{m_{\ell}^{2}}{q^{2}}\operatorname{Re}[\mathcal{A}_{\parallel,T}\mathcal{A}_{\perp,T}^{*}],$$
(A10)

$$\tilde{I}_{7} = -8\sqrt{2} \mathrm{Im}[\mathcal{A}_{SP}\mathcal{A}_{\perp,T}^{*}] - 2\sqrt{2} \mathrm{Im}[\mathcal{A}_{0}\mathcal{A}_{||}^{*}] + 2\sqrt{2} \frac{m_{\ell}^{2}}{q^{2}} \mathrm{Im}[\mathcal{A}_{t}\mathcal{A}_{\perp}^{*}]
+ 2\sqrt{2} \frac{m_{\ell}}{\sqrt{q^{2}}} \{4 \mathrm{Im}[\mathcal{A}_{0}\mathcal{A}_{||,T}^{*}] - 4 \mathrm{Im}[\mathcal{A}_{||}\mathcal{A}_{0,T}^{*}] - 4 \mathrm{Im}[\mathcal{A}_{t}\mathcal{A}_{\perp,T}^{*}] - \mathrm{Im}[\mathcal{A}_{\perp}\mathcal{A}_{SP}^{*}]\},$$
(A11)

$$\tilde{I}_8 = -\sqrt{2} \left(1 - \frac{m_\ell^2}{q^2} \right) \operatorname{Im}[\mathcal{A}_\perp \mathcal{A}_0^*], \tag{A12}$$

$$\tilde{I}_9 = 2\left(1 - \frac{m_\ell^2}{q^2}\right) \operatorname{Im}[\mathcal{A}_{||}\mathcal{A}_{\perp}^*],\tag{A13}$$

where $|p_{D^*}| = \sqrt{\lambda(m_B^2, m_{D^*}^2, q^2)}/(2m_B)$ represents the magnitude of the D^* 3-momentum, and $\lambda(a, b, c) = a^2 + b^2 + c^2 - 2ab - 2bc - 2ca$.

APPENDIX B: HELICITY AMPLITUDES AND FORM FACTORS

The 12 angular coefficients needed to construct the full angular distribution of Eq. (11) were presented in Appendix A. These angular coefficients depend on eight helicity amplitudes that can be further expressed in terms of NP coefficients (g_P , g_L , g_R , and g_T) and hadronic form factors. We list the helicity amplitudes below [35,50].

$$\mathcal{A}_{SP} = -g_P \frac{\sqrt{\lambda(m_B^2, m_{D^*}^2, q^2)}}{m_b + m_c} A_0(q^2), \tag{B1}$$

$$\mathcal{A}_{0} = -\frac{(1+g_{L}-g_{R})(m_{B}+m_{D^{*}})}{2m_{D^{*}}\sqrt{q^{2}}} \bigg[(m_{B}^{2}-m_{D^{*}}^{2}-q^{2})A_{1}(q^{2}) - \frac{\lambda(m_{B}^{2},m_{D^{*}}^{2},q^{2})}{(m_{B}+m_{D^{*}})^{2}}A_{2}(q^{2}) \bigg],$$
(B2)

$$\mathcal{A}_{t} = -(1 + g_{L} - g_{R}) \frac{\sqrt{\lambda(m_{B}^{2}, m_{D^{*}}^{2}, q^{2})}}{\sqrt{q^{2}}} A_{0}(q^{2}),$$
(B3)

$$\mathcal{A}_{\pm} = (1 + g_L - g_R)(m_B + m_{D^*})A_1(q^2) \mp (1 + g_L + g_R)\frac{\sqrt{\lambda(m_B^2, m_{D^*}^2, q^2)}}{m_B + m_{D^*}}V(q^2), \tag{B4}$$

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$$\mathcal{A}_{0,T} = \frac{g_T}{2m_{D^*}(m_B^2 - m_{D^*}^2)} ((m_B^2 - m_{D^*}^2)(m_B^2 + 3m_{D^*}^2 - q^2)T_2(q^2) - \lambda(m_B^2, m_{D^*}^2, q^2)T_3(q^2)), \tag{B5}$$

$$\mathcal{A}_{\pm,T} = g_T \frac{\sqrt{\lambda(m_B^2, m_{D^*}^2, q^2)} T_1(q^2) \pm (m_B^2 - m_{D^*}^2) T_2(q^2)}{\sqrt{q^2}}.$$
 (B6)

The angular coefficients requiring vector and/or tensor type contributions may also require the amplitudes to be expressed in the transversity basis as follows.

$$\mathcal{A}_{\parallel,T} = (\mathcal{A}_{+(,T)} + \mathcal{A}_{-(,T)})/\sqrt{2},\tag{B7}$$

$$\mathcal{A}_{\perp,T} = (\mathcal{A}_{+(,T)} - \mathcal{A}_{-(,T)})/\sqrt{2}.$$
 (B8)

The above helicity amplitudes depend on the seven hadronic form factors listed below.

$$V(q^2) = \frac{m_B + m_{D^*}}{2\sqrt{m_B m_{D^*}}} h_V(w(q^2)), \tag{B9}$$

$$A_1(q^2) = \frac{(m_B + m_{D^*})^2 - q^2}{2\sqrt{m_B m_{D^*}}(m_B + m_{D^*})} h_{A_1}(w(q^2)), \tag{B10}$$

$$A_2(q^2) = \frac{m_B + m_{D^*}}{2\sqrt{m_B m_{D^*}}} \left[h_{A_3}(w(q^2)) + \frac{m_{D^*}}{m_B} h_{A_2}(w(q^2)) \right],$$
(B11)

$$A_{0}(q^{2}) = \frac{1}{2\sqrt{m_{B}m_{D^{*}}}} \left[\frac{(m_{B} + m_{D^{*}})^{2} - q^{2}}{2m_{D^{*}}} h_{A_{1}}(w(q^{2})) - \frac{m_{B}^{2} - m_{D^{*}}^{2} + q^{2}}{2m_{B}} h_{A_{2}}(w(q^{2})) - \frac{m_{B}^{2} - m_{D^{*}}^{2} - q^{2}}{2m_{D^{*}}} h_{A_{3}}(w(q^{2})) \right], \quad (B12)$$

$$T_1(q^2) = \frac{1}{2\sqrt{m_B m_{D^*}}} [(m_B + m_{D^*})h_{T_1}(w(q^2)) - (m_B - m_{D^*})h_{T_2}(w(q^2))],$$
(B13)

$$T_2(q^2) = \frac{1}{2\sqrt{m_B m_{D^*}}} \left[\frac{(m_B + m_{D^*})^2 - q^2}{m_B + m_{D^*}} h_{T_1}(w(q^2)) - \frac{(m_B - m_{D^*})^2 - q^2}{m_B - m_{D^*}} h_{T_2}(w(q^2)) \right], \tag{B14}$$

$$T_{3}(q^{2}) = \frac{1}{2\sqrt{m_{B}m_{D^{*}}}} \left[(m_{B} - m_{D^{*}})h_{T_{1}}(w(q^{2})) - (m_{B} + m_{D^{*}})h_{T_{2}}(w(q^{2})) - 2\frac{m_{B}^{2} - m_{D^{*}}^{2}}{m_{B}}h_{T_{3}}(w(q^{2})) \right],$$
(B15)

where the recoil angle, $w(q^2)$ can be expressed as is $w(q^2) = (m_B^2 + m_{D^*}^2 - q^2)/2m_Bm_{D^*}$. The above expressions depend on several lepton and meson masses that are used as input parameters. In our calculations we use the values of meson and lepton masses given in Table V. We

TABLE V. Input values used for meson and lepton masses taken from the Particle Data Group [51]. Numbers in parentheses represent the errors in the last two digits.

Masses	Value (MeV)
m_{B^0}	5279.63(20)
$m_{D^{*+}}$	2010.26(05)
m_e	0.5109989461(31)
m_{μ}	105.6583745(24)

have also used the following values for the quark masses, $m_b = \text{GeV}$ and $m_c = \text{GeV}$.

Note that the above form factors still depend on several additional functions of q^2 , namely, h_V , h_{A_1} , h_{A_2} , h_{A_3} , h_{T_1} , h_{T_2} , h_{T_3} , R_1 , R_2 , and R_3 . There are several ways of parametrizing these functions using heavy quark effective theory (HQET). Two such parametrizations are presented in Appendix C.

APPENDIX C: PARAMETRIZATIONS OF THE HADRONIC FORM FACTORS

The hadronic form factors described in Appendix B depend on several form factors that appear as functions of q^2 in HQET. At present there are several ways of parametrizing these functions. Although each parametrization

gives a slightly different value for the underlying function, a conclusive identification of the best way to parametrize these functions still eludes us. This problem adds to the theoretical uncertainties associated with the determinations of some of the NP observables discussed in this article.

A commonly used parametrization for the HQET form factors, first presented by Caprini, Lellouch, and Neubert (CLN) in Ref. [37] is given below.

$$h_V(w) = R_1(w)h_{A_1}(w),$$
 (C1)

$$h_{A_2}(w) = \frac{R_2(w) - R_3(w)}{2r_{D^*}} h_{A_1}(w), \qquad (C2)$$

$$h_{A_3}(w) = \frac{R_2(w) + R_3(w)}{2} h_{A_1}(w), \qquad (C3)$$

$$h_{T_1}(w) = \frac{1}{2(1+r_{D^*}^2 - 2r_{D^*}w)} \\ \times \left[\frac{m_b - m_c}{m_B - m_{D^*}}(1-r_{D^*})^2(w+1)h_{A_1}(w) - \frac{m_b + m_c}{m_B + m_{D^*}}(1+r_{D^*})^2(w-1)h_V(w)\right], \quad (C4)$$

$$h_{T_2}(w) = \frac{(1 - r_{D^*}^2)(w+1)}{2(1 + r_{D^*}^2 - 2r_{D^*}w)} \left[\frac{m_b - m_c}{m_B - m_{D^*}}h_{A_1}(w) - \frac{m_b + m_c}{m_B + m_{D^*}}h_V(w)\right],$$
(C5)

$$h_{T_{3}}(w) = -\frac{1}{2(1+r_{D^{*}})(1+r_{D^{*}}^{2}-2r_{D^{*}}w)} \\ \times \left[2\frac{m_{b}-m_{c}}{m_{B}-m_{D^{*}}}r_{D^{*}}(w+1)h_{A_{1}}(w) + \frac{m_{b}-m_{c}}{m_{B}-m_{D^{*}}}(1+r_{D^{*}}^{2}-2r_{D^{*}}w)(h_{A_{3}}(w) - r_{D^{*}}h_{A_{2}}(w)) - \frac{m_{b}+m_{c}}{m_{B}+m_{D^{*}}}(1+r_{D^{*}})^{2}h_{V}(w)\right],$$
(C6)

where $r_{D^*} = m_{D^*}/m_B$ and the *w* dependencies are expressed as

$$h_{A_1}(w) = h_{A_1}(1)[1 - 8\rho_{D^*}^2 z + (53\rho_{D^*}^2 - 15)z^2 - (231\rho_{D^*}^2 - 91)z^3],$$
(C7)

TABLE VI. Input values of parameters needed for the CLN parametrization of form factors used here were taken from [35].

Parameter	Value	
$h_{A_1}(1)$	0.908 ± 0.017	
$ ho_{D^*}^2$	1.207 ± 0.026	
$\overline{R_1}(1)$	1.403 ± 0.033	
$R_2(1)$	0.854 ± 0.020	

$$R_1(w) = R_1(1) - 0.12(w - 1) + 0.05(w - 1)^2$$
, (C8)

$$R_2(w) = R_2(1) + 0.11(w - 1) - 0.06(w - 1)^2$$
, (C9)

$$R_3(w) = 1.22 - 0.052(w - 1) + 0.026(w - 1)^2.$$
 (C10)

The parameter z is related to the recoil angle w through $z(w) = (\sqrt{w+1} - \sqrt{2})/(\sqrt{w+1} + \sqrt{2})$. The values of $h_{A_1}(1)$, $R_1(1)$, $R_2(1)$, and $\rho_{D^*}^2$, listed in Table VI, were taken from Ref. [35].

Yet another way of parametrizing the HQET form factors is to express them in terms of the leading Isgur-Wise (IW) function $\xi(w)$ [52] and subleading IW terms, which represents higher order power corrections to the leading IW function as

$$h_X(w) = \xi(w)\hat{h}_X(w),$$
 $(X = V, A_1, A_2, A_3, T_1, T_2, T_3),$
(C11)

where

$$\hat{h}_{X}(w) = \hat{h}_{X,0} + \varepsilon_a \delta \hat{h}_{X,\alpha_s} + \varepsilon_b \delta \hat{h}_{X,m_b} + \varepsilon_c \delta \hat{h}_{X,m_c} + \varepsilon_c^2 \delta \hat{h}_{X,m_c^2}.$$
(C12)

Here, ε_a , ε_b , ε_c denote the expansion coefficients corresponding to the higher order corrections in α_s and $1/m_{b,c}$, respectively, which were worked out by [37,53] using heavy quark symmetry.

The leading term in (C12) is

$$\hat{h}_{X,0} = \begin{cases} 1 & \text{for } X = A_1, A_3, T_1, \\ 0 & \text{for } X = A_2, T_2, T_3. \end{cases}$$
(C13)

The α_s corrections are given as

$$\delta \hat{h}_{V,\alpha_s} = \frac{1}{6z_{cb}(w - w_{cb})} [4z_{cb}(w - w_{cb})\Omega_w(w) + 2(w + 1)((3w - 1)z_{cb} - z_{cb}^2 - 1)r_w(w) - 12z_{cb}(w - w_{cb}) - (z_{cb}^2 - 1)\log z_{cb}] + V(\mu),$$
(C14)

$$\delta \hat{h}_{A_1,\alpha_s} = \frac{1}{6z_{cb}(w - w_{cb})} [4z_{cb}(w - w_{cb})\Omega_w(w) + 2(w - 1)((3w + 1)z_{cb} - z_{cb}^2 - 1)r_w(w) - 12z_{cb}(w - w_{cb}) - (z_{cb}^2 - 1)\log z_{cb}] + V(\mu),$$
(C15)

$$\delta \hat{h}_{A_2,\alpha_s} = \frac{-1}{6z_{cb}^2 (w - w_{cb})^2} [(2 + (2w^2 - 5w - 1)z_{cb} + 2w(2w - 1)z_{cb}^2 + (1 - w)z_{cb}^3)r_w(w) - 2z_{cb}(z_{cb} + 1)(w - w_{cb}) + (z_{cb}^2 - (4w + 2)z_{cb} + 3 + 2w)z_{cb}\log z_{cb}],$$
(C16)

$$\delta \hat{h}_{A_3,\alpha_s} = \delta \hat{h}_{A_1,\alpha_s} + \frac{1}{6z_{cb}(w - w_{cb})^2} [2z_{cb}(z_{cb} + 1)(w_{cb} - w) + (2z_{cb}^3 + z_{cb}^2(2w^2 - 5w - 1) + z_{cb}(4w^2 - 2w) - w + 1)r_w(w) - (z_{cb}^2(2w + 3) - z_{cb}(4w + 2) + 1)\log z_{cb}],$$
(C17)

$$\delta \hat{h}_{T_1,\alpha_s} = \frac{1}{3z_{cb}(w - w_{cb})} [2z_{cb}(w - w_{cb})\Omega_w(w) + 2z_{cb}(w^2 - 1)r_w(w) - 6z_{cb}(w - w_{cb}) + (1 - z_{cb}^2)\log z_{cb}] + T(\mu), \quad (C18)$$

$$\delta \hat{h}_{T_{2},\alpha_{s}} = \frac{w+1}{3z_{cb}(w-w_{cb})} [(1-z_{cb}^{2})r_{w}(w) + 2z_{cb}\log z_{cb}],$$
(C19)

$$\delta \hat{h}_{T_{3},\alpha_{s}} = \frac{1}{3z_{cb}(w - w_{cb})} [(z_{cb}w - 1)r_{w}(w) - z_{cb}\log z_{cb}],$$
(C20)

where

$$z_{cb} = \frac{m_c}{m_b}, \qquad w_{cb} = \frac{1}{2}(z_{cb} + z_{cb}^{-1}), \qquad w_{\pm}(w) = w \pm \sqrt{w^2 - 1},$$
 (C21)

$$r_w(w) = \frac{\log w_+(w)}{\sqrt{w^2 - 1}},$$
(C22)

$$\Omega_{w}(w) = \frac{w}{2\sqrt{w^{2}-1}} [2\text{Li}_{2}(1-w_{-}(w)z_{cb}) - 2\text{Li}_{2}(1-w_{+}(w)z_{cb}) + \text{Li}_{2}(1-w_{+}^{2}(w)) - \text{Li}_{2}(1-w_{-}^{2}(w))] - wr_{w}(w)\log z_{cb} + 1.$$
(C23)

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Here $\text{Li}_2(x) = \int_x^0 dt \log(1-t)/t$ is the dilogarithm function and $V(\mu)$, $T(\mu)$ are scale factors given as

$$V(\mu) = -\frac{2}{3}(wr_w(w) - 1)\log\frac{m_b m_c}{\mu^2}, \qquad (C24)$$

$$T(\mu) = -\frac{1}{3}(2wr_w(w) - 3)\log\frac{m_b m_c}{\mu^2}.$$
 (C25)

In our calculations we choose the scale $\mu = 4.2$ GeV. The $1/m_{b,c}$ corrections in Eq. (C12) are given as

$$\delta \hat{h}_{V,m_b} = \hat{L}_1(w) - \hat{L}_4(w),$$
 (C26)

$$\delta \hat{h}_{V,m_c} = \hat{L}_2(w) - \hat{L}_5(w),$$
 (C27)

$$\delta \hat{h}_{A_1,m_b} = \hat{L}_1(w) - \frac{w-1}{w+1}\hat{L}_4(w),$$
 (C28)

$$\delta \hat{h}_{A_1,m_c} = \hat{L}_2 - \frac{w-1}{w+1} \hat{L}_5(w), \qquad (C29)$$

 $\delta \hat{h}_{A_2,m_b} = 0, \qquad (C30)$

$$\delta \hat{h}_{A_2,m_c} = \hat{L}_3(w) + \hat{L}_6(w),$$
 (C31)

$$\delta \hat{h}_{A_3,m_b} = \hat{L}_1(w) - \hat{L}_4(w),$$
 (C32)

$$\delta \hat{h}_{A_3,m_c} = \hat{L}_2(w) - \hat{L}_3(w) + \hat{L}_6(w) - \hat{L}_5(w), \quad (C33)$$

$$\delta \hat{h}_{T_1,m_b} = \hat{L}_1(w), \tag{C34}$$

$$\delta \hat{h}_{T_1,m_c} = \hat{L}_2(w), \tag{C35}$$

$$\delta \hat{h}_{T_2,m_b} = -\hat{L}_4(w), \tag{C36}$$

$$\delta \hat{h}_{T_2,m_c} = \hat{L}_5(w), \tag{C37}$$

$$\delta \hat{h}_{T_3,m_b} = 0, \qquad (C38)$$

$$\delta \hat{h}_{T_3,m_c} = \frac{1}{2} (\hat{L}_6(w) - \hat{L}_3(w)),$$
 (C39)

where the $\hat{L}(w)$ functions read

$$\hat{L}_1(w) = -4(w-1)\hat{\chi}_2(w) + 12\hat{\chi}_3(w), \quad (C40)$$

$$\hat{L}_2(w) = -4\hat{\chi}_3(w),$$
 (C41)

$$\hat{L}_3(w) = 4\hat{\chi}_2(w),$$
 (C42)

$$\hat{L}_4(w) = 2\eta(w) - 1,$$
 (C43)

$$\hat{L}_5(w) = -1,$$
 (C44)

$$\hat{L}_6(w) = -\frac{2(1+\eta(w))}{w+1}.$$
 (C45)

The corrections of order $1/m_c^2$ are included via the subleading reduced IW functions $\hat{l}_{1-6}(w)$ as [38,54]

$$\delta \hat{h}_{V,m_c^2} = \hat{\ell}_2(w) - \hat{\ell}_5(w),$$
 (C46)

$$\delta \hat{h}_{A_1, m_c^2} = \hat{\ell}_2(w) - \frac{w-1}{w+1} \hat{\ell}_5(w), \qquad (C47)$$

$$\delta \hat{h}_{A_2,m_c^2} = \hat{\ell}_3(w) + \hat{\ell}_6(w),$$
 (C48)

$$\delta \hat{h}_{A_3,m_c^2} = \hat{\ell}_2(w) - \hat{\ell}_3(w) - \hat{\ell}_5(w) + \hat{\ell}_6(w), \quad (C49)$$

$$\delta \hat{h}_{T_1,m_c^2} = \hat{\ell}_2(w), \tag{C50}$$

$$\delta \hat{h}_{T_2,m_c^2} = \hat{\ell}_5(w), \tag{C51}$$

$$\delta \hat{h}_{T_3,m_c^2} = \frac{1}{2} (\hat{\ell}_3(w) - \hat{\ell}_6(w)).$$
 (C52)

The IW functions are expressed, in general, as expansions about w = 1 as

$$f(w) = \sum_{n=0}^{\infty} \frac{f^{(n)}}{n!} (w-1)^n,$$
 (C53)

with $f = \xi, \eta, \hat{\chi}_2, \hat{\chi}_3$ and $\hat{\ell}_i$. One can further relate the kinematic variable *w* with the expansion variable *z* as

$$w(z) = 2\left(\frac{1+z}{1-z}\right)^2 - 1.$$
 (C54)

One can then expand the IW functions up to any order in z as

$$f(w) = f^{(0)} + 8f^{(1)}z + 16(f^{(1)} + 2f^{(2)})z^2 + \frac{8}{3}(9f^{(1)} + 48f^{(2)} + 32f^{(3)})z^3 + \cdots$$
(higher orders). (C55)

The authors of Ref. [55] have performed a simultaneous fit of the HQET parameters and the CKM element V_{cb} by considering an expansion of the IW functions up to order NNLO (3/2/1) and NNLO (2/1/0), where

NNLO(3/2/1):
$$\xi(w)$$
 up to z^3 , $\hat{\chi}_{2,3}(w)$, $\eta(w)$ up to order z^2 and $\hat{\ell}_i$ up to order z , (C56)

NNLO(2/1/0):
$$\xi(w)$$
 up to z^2 , $\hat{\chi}_{2,3}(w)$, $\eta(w)$ up to order z and $\hat{\ell}_i$ up to order z^0 . (C57)

The fitted value of the parameters for the above two scenarios from Ref. [55] are given in Table VII.

The other alternate way of parametrizing the form factors is due to Boyd, Grinstein, and Lebed (BGL) [36]. Both the CLN and BGL form-factor coefficients are constrained from the same dispersive bounds. However, unlike CLN, they do not employ HQET relations to reduce the number of form-factor parameters and are hence, more general. The form factors $\mathcal{F}_i \equiv \{f, g, \mathcal{F}_1, \mathcal{F}_2\}$ are expressed as series expansions in z as

$$\mathcal{F}_{i}(z) = \frac{1}{P_{i}(z)\phi_{i}(z)} \sum_{j=0}^{N} a_{j}^{\mathcal{F}_{i}} z^{j}, \qquad (C58)$$

where *z* is related to the recoil angle *w* as in Eq. (C54) and $P_i(z) = \prod_p \frac{z-z_p}{1-zz_p}$ are called the Blaschke factors that help eliminate poles at $z = z_p$ at the B_c resonances given by

$$z_p = \frac{\sqrt{t_+ - M_p^2} - \sqrt{t_+ - t_-}}{\sqrt{t_+ - M_p^2} + \sqrt{t_+ - t_-}}; \qquad t_\pm = (m_B \pm m_{D^*})^2.$$
(C59)

The pole mass (M_p) for the different types of resonances are listed in Table VIII. The outer functions ϕ_i are given as

$$\phi_f = \frac{4r_{D^*}}{m_B^2} \sqrt{\frac{n_I}{6\pi\chi_{1^+}^T(0)}} \frac{(1+z)(1-z)^{3/2}}{[(1+r_{D^*})(1-z) + 2\sqrt{r_{D^*}}(1+z)]^4},$$
(C60)

$$\phi_g = 16r_{D^*}^2 \sqrt{\frac{n_I}{3\pi\tilde{\chi}_{1^-}^T(0)}} \frac{(1+z)^2(1-z)^{-1/2}}{[(1+r_{D^*})(1-z) + 2\sqrt{r}(1+z)]^4},$$
(C61)

TABLE VII. Values of input parameters needed for the HQET (3/2/1) and HQET (2/1/0) parametrizations of the hadronic form factors taken from [55].

Parameter	HQET (3/2/1)	HQET (2/1/0)	
$\frac{\xi^{(0)}}{\xi^{(0)}}$	1	1	
$\xi^{(1)}$	-0.93 ± 0.10	-1.10 ± 0.04	
$\xi^{(2)}$	$+1.35\pm0.26$	$+1.57\pm0.10$	
$\xi^{(3)}$	-2.67 ± 0.75		
$\hat{\chi}_{2}^{(0)}$	-0.05 ± 0.02	-0.06 ± 0.02	
$\hat{\chi}_2^{(1)}$	$+0.01\pm0.02$	-0.06 ± 0.02	
$\hat{\chi}_{2}^{(2)}$	-0.01 ± 0.02		
$\hat{\chi}_{3}^{(0)}$	0	0	
$\hat{\chi}_{3}^{(1)}$	-0.05 ± 0.02	-0.03 ± 0.01	
$\hat{\chi}_3^{(2)}$	-0.03 ± 0.03		
$\eta^{(0)}$	$+0.74\pm0.11$	$+0.38\pm0.06$	
$\eta^{(1)}$	$+0.05\pm0.03$	$+0.08\pm0.03$	
$\eta^{(2)}$	-0.05 ± 0.05		
$\tilde{\mathscr{C}}_1^{(0)}$	$+0.09\pm0.18$	$+0.50\pm0.16$	
$\tilde{\ell}_1^{(1)}$	$+1.20\pm2.09$		
$\tilde{\ell}_2^{(0)}$	-2.29 ± 0.33	-2.16 ± 0.29	
$\tilde{\ell}_2^{(1)}$	-3.66 ± 1.56		
$\tilde{\ell}_{3}^{(0)}$	-1.90 ± 12.4	-1.14 ± 2.34	
$\tilde{\ell}_{3}^{(1)}$	$+3.91\pm4.35$		
$\tilde{\ell}_4^{(0)}$	-2.56 ± 0.94	$+0.82\pm0.47$	
$\tilde{\ell}_4^{(1)}$	$+1.78\pm0.93$		
$\tilde{\ell}_5^{(0)}$	$+3.96\pm1.17$	$+1.39\pm0.43$	
$\tilde{\ell}_5^{(1)}$	$+2.10\pm1.47$		
$\tilde{\ell}_6^{(0)}$	$+4.96\pm5.76$	$+0.17\pm1.15$	
$\tilde{\ell}_{6}^{(1)}$	$+5.08\pm2.97$		

$$\phi_{\mathcal{F}_{1}} = \frac{4r_{D^{*}}}{m_{B}^{3}} \sqrt{\frac{n_{I}}{6\pi\chi_{1^{+}}^{T}(0)}} \times \frac{(1+z)(1-z)^{5/2}}{\left[(1+r_{D^{*}})(1-z) + 2\sqrt{r_{D^{*}}}(1+z)\right]^{5}}, \quad (C62)$$

$$\begin{split} \phi_{\mathcal{F}_2} &= 8\sqrt{2}r_{D^*}^2 \sqrt{\frac{n_I}{\pi \tilde{\chi}_{1^+}^{L}(0)}} \\ &\times \frac{(1+z)^2(1-z)^{-1/2}}{\left[(1+r_{D^*})(1-z)+2\sqrt{r_{D^*}}(1+z)\right]^4}. \end{split} \tag{C63}$$

The various relevant inputs for computing the outer functions are listed in Table IX. The form-factor coefficients $a_i^{\mathcal{F}_i}$ satisfy the weak unitarity constraints given by

TABLE VIII. The pole masses corresponding to different types of B_c resonances as listed in [57].

Form factor	Туре	Pole masses M_p (GeV)
g	1-	6.329,6.920,7.020
f, \mathcal{F}_1	1^{+}	6.739,6.750,7.145,7.150
\mathcal{F}_2	0-	6.275,6.842,7.250

TABLE IX. Relevant inputs for the outer functions taken from [57].

Form factor	Туре	
$egin{aligned} &n_I\ &\chi_{1^+}^T(0)~{ m GeV^{-2}}\ & ilde{\chi}_{1^-}^T(0)~{ m GeV^{-2}}\ & ilde{\chi}_{1^-}^L(0)~{ m GeV^{-2}}\ & ilde{\chi}_{1^+}^L(0) \end{aligned}$	$\begin{array}{c} 2.6\\ 3.894\times 10^{-4}\\ 5.131\times 10^{-4}\\ 1.9421\times 10^{-2} \end{array}$	

$$\sum_{j=0}^{N} (a_j^g)^2 < 1, \qquad \sum_{j=0}^{N} (a_j^f)^2 + (a_j^{\mathcal{F}_1})^2 < 1,$$
$$\sum_{j=0}^{N} (a_j^{\mathcal{F}_2})^2 < 1.$$
(C64)

In addition to this, they are also subject to two kinematic constraints, one each at zero and maximum recoil, respectively, given by

$$\mathcal{F}_1(1) = m_B(1 - r_{D^*})f(1), \tag{C65}$$

$$\mathcal{F}_{2}(w_{\max}) = \frac{1 + r_{D^{*}}}{m_{B}^{2}(1 + w_{\max})(1 - r_{D^{*}})r_{D^{*}}} \mathcal{F}_{1}(w_{\max}).$$
(C66)

In our analysis, we consider the fitted values of the formfactor parameters from [12]. Lastly, for completion, we would like to list the relations between the BGL form factors and the hadronic form factors [56]:

$$g = \frac{2}{m_B + m_{D^*}} V,$$
 (C67)

$$f = (m_B + m_{D^*})A_1, (C68)$$

$$\mathcal{F}_{1} = m_{B}(m_{B} + m_{D^{*}})(w - r_{D^{*}})A_{1} - \frac{2m_{B}m_{D^{*}}(w^{2} - 1)}{1 + r_{D^{*}}}A_{2},$$
(C69)

$$\mathcal{F}_2 = 2A_0. \tag{C70}$$

The form-factor dependences on q^2 for the various types of parametrizations are shown in Fig. 9.



FIG. 9. Form-factor dependence on q^2 for three different form-factor parametrizations. The shaded band show the region with the 1σ upper and lower limits of the form-factor parameters listed in Tables VI and VII are considered without any correlation. For the HQET form factors, we show only the 2/1/0 scenario following the analysis presented in Ref. [55]. Here, $\tilde{T}_3(q^2)$ is defined as $\tilde{T}_3(q^2) = T_3(q^2)q^2/(m_B^2 - m_D^{*2})$.

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