Topology of black hole thermodynamics in Gauss-Bonnet gravity

Pavan Kumar Yerra^{®*} and Chandrasekhar Bhamidipati^{®†}

School of Basic Sciences, Indian Institute of Technology Bhubaneswar, Bhubaneswar, Odisha, 752050, India

(Received 6 March 2022; accepted 11 May 2022; published 25 May 2022)

Thermodynamics of black holes in anti-de Sitter (AdS) spacetimes typically contain critical points in the phase diagram, some of which correspond to the first order transition ending in a second order one. Following the recent proposal in [Phys. Rev. D 105, 104003 (2022)] on using Duan's ϕ -mapping theory, we classify the critical points of six dimensional charged Gauss-Bonnet black holes in AdS spacetime. We find that the higher derivative curvature terms in the form of Gauss-Bonnet gravity do not change the topological class of critical points in charged black holes in AdS, unlike the case of Born-Infeld corrections noted earlier. The connection between the topological nature of critical points and existence of first order phase transitions breaks down in a certain parameter regime. A resolution is proposed by treating the novel and conventional critical points as phase creation and phase annihilation points, respectively. Examples are provided to support the proposal.

DOI: 10.1103/PhysRevD.105.104053

I. INTRODUCTION

Black holes provide a rich arena to explore the effects of strong gravity [1]. Black hole thermodynamics in particular gives us deep insights to understand the nature of degrees of freedom of gravity [2]. Classically, black holes can be assigned an entropy S which is a constant times the surface area of the even horizon A [3], and this leads to the identification of laws of mechanics with the corresponding laws of thermodynamics [4]. The formal analogy with thermodynamics comes to life only after invoking quantum effects where the temperature T is identified with surface gravity κ [5]. Black hole mass M on the other hand is related to internal energy U. Hawking and Page discovered a remarkable transition which happens between a selfgravitating hot radiation and Schwarzschild black hole, albeit in asymptotically anti-de Sitter (AdS) spacetime [6]. The Hawking-Page (HP) transition of course has an alternate interpretation as a confinement/deconfinement transition in the dual conformal field theory (CFT) [7–9]. In the case of charged black holes in AdS, there exists a first-order phase transition between the small and large black holes analogous to liquid-gas phase transitions of van der Waals (VdW) fluid [10]. The situation becomes interesting in the black hole chemistry paradigm [11-16], where the cosmological constant Λ is assumed to be giving rise to a pressure $P = -\Lambda/8\pi$, with its conjugate thermodynamic volume denoted as V. One also realizes that the erstwhile first law

$$dM = TdS + \Phi dQ, \tag{1.1}$$

pk11@iitbbs.ac.in

can now include the traditional work terms appearing in standard thermodynamics, when written as [12]

$$dM = TdS + VdP + \Phi dQ, \tag{1.2}$$

where Q is the charge and Φ the conjugate potential, with M now reinterpreted as enthalpy H. This novel extended thermodynamics ensures for the charged AdS black holes that, the small-large black hole transition has an exact map to the liquid-gas type phase transition, including the presence of a critical region where first order phase transition terminates in a second order one [10,11,17]. Criticality in phase transitions, especially in black holes is an interesting topic on its own due to several universal phenomena which occur together with the scaling of thermodynamic quantities. Very recently, a remarkable idea was put forward in [18], where a topological approach following Duan's ϕ -mapping theory was used to classify the nature of critical points for charged and Born-Infeld black holes.

It is interesting as well as important to test whether the proposal [18] is valid just for VdW's liquid-gas type transitions or extendable to more general situations when there are multiple critical points. In fact, apart from the small black hole (SBH) to large black hole (LBH) transition, there are further intriguing phenomena akin to day to day thermodynamic systems, such as reentrant phase transitions, multiple solid-liquid-gas type transitions which occur in a variety of black holes, especially in six or larger dimensions [19]. The black hole systems involve Gauss-Bonnet, Born-Infeld and other higher derivative curvature terms added to the Einstein action, in addition to multiple rotating Kerr-AdS black hole systems. The latter in fact also

chandrasekhar@iitbbs.ac.in

display tricritical points in certain range of parameters, including the interesting superfluid phenomena [19–30]. While there is an independent understanding emerging for several of these phase transitions and critical behavior from a thermodynamics point of view, it is important to look for other possible methods to make inroads into the classification of such transitions.

With the above motivations, we look for a tractable system which presents us with multiple critical points in the phase diagram of black holes, i.e., the Gauss-Bonnet (GB) theory in six dimensions [31-33]. Higher derivative curvature terms from gravity sector, such as Gauss-Bonnet and Lovelock terms are important, in the discussions of semiclassical quantum gravity, and from the point of view of low-energy effective action of superstring theories. The particular interest in these higher curvature terms comes from the fact that they result in field equations which contain no more than second derivatives of the metric for dimensions higher than five, topological in four and vanishing identically for dimension less than three. Other Lovelock terms behave similarly with respect to their critical dimension and thus have the advantage of absence of ghost-like modes. In the context of AdS/CFT correspondence, Gauss-Bonnet terms can be understood as providing corrections in the large N expansion of boundary gauge theories, in the strong coupling limit. In the past, such terms have given crucial contributions to the viscosity to entropy ratio and novel bounds [34]. Although, in general in the Gauss-Bonnet theory, the small to large black hole type phase transition exists in any number of dimensions, the six dimensional case is special. This is because there could either be one, two or even three critical points with presence of multiple swallow tail behavior, notwithstanding the narrow range of parameters. This still gives us a rich range of parameters to explore the phase diagram and put the proposals in [18] to test. Another motivation is the surprising result pointed out in [18], that the charged black holes and Born-Infeld black holes are thermodynamically classified into two different topological classes. This raises the question as to whether the higher derivative curvature corrections in gravity sector, such as, the Gauss-Bonnet etc., would change the topological nature of critical points of black holes too.

The rest of the paper is organized as follows. In Sec. II, we closely follow [18,31–33,35,36] and collect salient aspects

of thermodynamics of Gauss-Bonnet AdS black holes in six dimensions and present the topological approach, necessary for performing the calculations in Sec. III. Section III contains the main results of calculations on topological classification of critical points of GB black holes in AdS in various charge ranges. Section IV, contains a discussion on the nature of critical points, were we identify a parameter range where the classification of critical points as conventional or novel, proposed in [18] breaks down. We end this section by giving a new proposal on classifying the critical points as phase creation and phase annihilation, to resolve the ambiguity. Remarks and conclusions are given in Sec. V. Appendix A contains the calculation of topological charge for black holes in six dimensions when the Gauss-Bonnet parameter α is set to zero, which shows that our results are in conformity with the earlier solutions [18] in the Einstein-Maxwell system. In Appendix B, we reason that our novel proposal for understanding topology of critical points presented in Sec. IV, is consistent and works for the case of black holes in third order Lovelock gravity as well.

II. THERMODYNAMICS OF GAUSS-BONNET BLACK HOLES AND TOPOLOGY

We start with the thermodynamics of charged-Gauss-Bonnet (GB) black holes in AdS. As mentioned earlier, this system in six dimensions exhibits a rich phase structure compared to its lower and higher dimensional counterparts [31–33]. The corresponding thermodynamic quantities, i.e., the temperature *T*, entropy *S*, specific heat C_P , and the Gibbs free energy *G* in the extended phase space (where we treat the cosmological constant Λ as the pressure *P* and its conjugate quantity as thermodynamic volume *V* [12]), are given by [31,33]:

$$T = \frac{8\pi P r_h^8 + 6r_h^6 + 2\alpha r_h^4 - Q^2}{8\pi r_h^5 (r_h^2 + 2\alpha)},$$
 (2.1)

$$S = \frac{r_h^4}{4} \left(1 + \frac{4\alpha}{r_h^2} \right), \tag{2.2}$$

$$C_P = \frac{(r_h^3 + 2\alpha r_h)^2 (8\pi P r_h^8 + 6r_h^6 + 2\alpha r_h^4 - Q^2)}{Q^2 (7r_h^2 + 10\alpha) + 2r_h^4 (4\pi P r_h^6 + 3r_h^4 (8\pi P\alpha - 1) + 3\alpha r_h^2 - 2\alpha^2)},$$
(2.3)

$$G = \frac{5Q^2(7r_h^2 + 20) - 6r_h^4(4\pi Pr_h^6 + r_h^4(48\pi P - 5) + 5r_h^2 - 20)}{480\pi r_h^3(r_h^2 + 2)},$$
(2.4)

where, r_h is the horizon radius, Q is the charge of the black hole and α is the Gauss-Bonnet (GB) coupling parameter.



FIG. 1. Behavior of critical pressure P_c with charge Q. The charges at the points A, B, C, and D are $Q_A = 0.1914$, $Q_B = 0.2018$, $Q_C = 0.1705$, and $Q_D = 0.1946$ respectively. We have three critical points for $Q < Q_B$, and one critical point for $Q > Q_B$. The line segment CD denotes the triple points. (Here, we set $\alpha = 1$).

A. Phase structure

The equation of state (obtained by rewriting Eq. (2.1) in terms of pressure) exhibits various phase structures depending on the charge parameter Q [31,33]. This can be seen from the behavior of critical¹ pressure P_c with charge Q, as shown in Fig. 1. Behavior of critical pressure P_c shows the existence of three critical points for $Q < Q_B$, two critical points for $Q = Q_B$, and one critical point at P_c shows the However, for case $Q = Q_B$, the critical point at B in Fig. 1 is not a true critical point, as it is not an inflection point but an undulation point. Thus, we have only one critical point for $Q = Q_B$.

Further, in the case $Q \ge Q_B$, the critical points are related to the van der Waals type small/large black hole phase transitions, whereas in the case $Q < Q_B$, the critical points are related to a rich phase structure involving the triple points (where, the three phases, i.e., small/intermediate/large black hole phases, can coexist). In fact, all the critical points may not globally minimize the Gibbs free energy, and thus may not appear in the phase diagram. The nature of the critical points actually depends on the fixed charge ranges, viz. $Q < Q_C, Q_C < Q < Q_D, Q_D < Q < Q_B$, and $Q_B < Q$. For the detailed discussion of phase structures related to these critical and triple points, one can refer [31,33]. Our aim in the next subsection, is to present the concept of topology in thermodynamics, which can be used to study properties associated with these critical points that belong to different fixed charge ranges.

B. Topology of thermodynamical functions

Recently, it is shown in [18] that, the critical points in the phase diagram of a thermodynamic system (black holes in particular) could be classified into conventional and novel critical type, based on the topological charges they carry. In the extended thermodynamic framework, typically the temperature T of a thermodynamic system is given as a function of the entropy S, pressure P, and other parameters x^i , i.e.,

$$T = T(S, P, x^i). \tag{2.5}$$

It is well known that the critical point, where the second order transition happens in the phase diagram, can be obtained by solving the condition for stationary point of inflection, i.e.,

$$(\partial_S T)_{P,x^i} = 0, \qquad (\partial_{S,S} T)_{P,x^i} = 0.$$
 (2.6)

Now, the interesting suggestion put forward in [18] is to construct a scalar thermodynamic function as

$$\Phi = \frac{1}{\sin\theta} T(S, x^i). \tag{2.7}$$

In practice, this is found by eliminating one of the variables in Eq. (2.5), through the first of the conditions in Eq. (2.6), and adding an additional factor of $1/\sin\theta$ for the ease of analysis. The set up of Duan's ϕ -mapping theory proceeds by constructing a new vector field $\phi = (\phi^S, \phi^{\theta})$, where $\phi^S = (\partial_S \Phi)_{\theta, x^i}$ and $\phi^{\theta} = (\partial_{\theta} \Phi)_{S, x^i}$. ϕ has an important property that, its zero points always lie at $\theta = \pi/2$, and can be identified with the presence of the critical points of the thermodynamic system. The horizontal lines at $\theta = 0$ and π , act as the boundaries of the parameter space, where the vector field ϕ is perpendicular to these lines. An important property of the above construction is the presence of a topological current j^{μ} , whose nonzero contribution only comes from the zero points of the vector field ϕ^a , i.e., $\phi^a(x^i) = 0$. Let there be N solutions of ϕ^a , whose *i*th solution is denoted as $\vec{x} = \vec{z_i}$. $Q_t = \int_{\Sigma} j^0 d^2 x$ thus gives the corresponding charge, which can be computed as [18,35,36]

$$Q_{t} = \int_{\Sigma} \Sigma_{i=1}^{N} \beta_{i} \eta_{i} \delta^{2}(\vec{x} - \vec{z_{i}}) d^{2}x$$
$$= \Sigma_{i=1}^{N} \beta_{i} \eta_{i} = \Sigma_{i=1}^{N} w_{i}.$$
(2.8)

Here, β_i is the positive integer (Hopf index) measuring the number of loops that ϕ^a makes around the *i*th zero point of ϕ . $\eta_i = \operatorname{sign}(J^0(\phi/x)_{z_i}) = \pm 1$ is called the Brouwer degree and w_i is the winding number for *i*th zero point of ϕ . Since, Q_t is nonzero only at the zero points of ϕ , one can assign a topological charge (given by the winding number) for each critical point, where the vector field ϕ is zero. Since, the Brouwer degree η_i can be positive or negative, the critical points were proposed to be further divided into two different topological classes, i.e., the conventional (where $\eta_i = -1$) and the novel (where $\eta_i = +1$) [18]. Further proposal is that, allowing Σ to span the entire parameter space for a given

¹The expression for P_c can be find in [31,33].

thermodynamic system, the topological properties of different thermodynamic systems can be divided into different classes from thermodynamics point of view.

III. TOPOLOGY OF CRITICAL POINTS IN SIX DIMENSIONAL CHARGED GAUSS-BONNET BLACK HOLES IN AdS

First we obtain the general expression for topological charge which can be used to analyze the critical points of black holes in six dimensional charged-Gauss-Bonnet theories in AdS spacetime.

A. Topological charge

Starting from the temperature² T in Eq. (2.1) and following the discussion in Sec. II B, we now compute the thermodynamic function Φ , as

$$\Phi = \frac{1}{\sin\theta} T(r_h, Q, \alpha), \qquad (3.1)$$

$$=\frac{1}{\sin\theta}\frac{(3r_{h}^{6}+2r_{h}^{4}\alpha-2Q^{2})}{2\pi r_{h}^{5}(r_{h}^{2}+6\alpha)}.$$
(3.2)

The vector field $\phi = (\phi^r, \phi^\theta)$ is obtained to be

$$\phi^{r} = \partial_{r_{h}} \Phi = \frac{\csc\theta(2Q^{2}(7r_{h}^{2} + 30\alpha) - 3r_{h}^{4}(r_{h}^{2} - 2\alpha)^{2})}{2\pi r_{h}^{6}(r_{h}^{2} + 6\alpha)^{2}}, \quad (3.3)$$

$$\phi^{\theta} = \partial_{\theta} \Phi = -\frac{\cot\theta\,\csc\theta(3r_h^6 + 2r_h^4\alpha - 2Q^2)}{2\pi r_h^5(r_h^2 + 6\alpha)}.\tag{3.4}$$

The normalized vector field is thus $n = \left(\frac{\phi^r}{||\phi||}, \frac{\phi^{\theta}}{||\phi||}\right)$. In order to calculate the topological charge associated with a critical point (where $\phi = 0$), one is required to find its winding number w_i . We know from the topology that if a given contour encloses a critical point then its winding number (i.e., the topological charge) is nonzero, otherwise it is zero [18,37–40].

In the orthogonal $\theta - r$ plane, we consider a contour *C*, that is piece-wise smooth and positive oriented. Let the contour *C*, for simplicity, be an ellipse centered at $(r_0, \frac{\pi}{2})$, parametrized by the angle $\vartheta \in (0, 2\pi)$ as [18,39]:

$$\begin{cases} r = a \cos \vartheta + r_0, \\ \theta = b \sin \vartheta + \frac{\pi}{2}. \end{cases}$$
(3.5)

Then, following [18,37–39], one can compute the topological charge (i.e., winding number) by measuring the deflection $\Omega(\vartheta)$ of the vector field ϕ along the given contour as

TABLE I. Parametric coefficients of contours.

Case		C_1	C_2	<i>C</i> ₃	C_4	<i>C</i> ₅
Case-1	a	0.07	0.07			
	b	0.4	0.4			
	r_0	0.69	0.9			
Case-2	а	0.15	0.15	0.15	0.15	0.7
	b	0.2	0.2	0.2	0.2	0.5
	r_0	1.2	1.55	0.67	2.2	1.1
Case-3	а	0.15	0.15	0.15	0.6	
	b	0.4	0.4	0.4	0.7	
	r_0	1.15	1.55	0.8	1.2	
Case-4	а	0.1	0.1	0.1	0.6	
	b	0.4	0.4	0.4	0.8	
	r_0	1.1	0.87	1.55	1.2	
Case-5	а	0.2	0.2			
	b	0.4	0.4			
	r_0	1.61	2.2			

$$Q_t = \frac{1}{2\pi} \Omega(2\pi), \qquad (3.6)$$

where,

$$\Omega(\vartheta) = \int_0^\vartheta \epsilon_{ab} n^a \partial_\vartheta n^b d\vartheta.$$
 (3.7)

In the following subsections, we compute the topological charges corresponding to the critical points of various fixed charge ranges given in Tables I and II. Case 1 in the Table I below corresponds to the Einstein-Maxwell case with a single critical point, where the Gauss-Bonnet parameter $\alpha = 0$. The calculation of topological charge for this case is given in Appendix A and shows that the classification of critical points matches the earlier considerations in lower dimensions [18].

TABLE II. Critical values.

		Critical point (CP)			
Case		CP_1	CP_2	CP ₃	
Case-2	r_c	1.21313	1.53448	0.676405	
	P_c	0.0195571	0.0197829	0.0321348	
	T_c	0.11228	0.112424	0.115055	
Case-3	r_c	1.13426	1.55453	0.792783	
	P_c	0.0192661	0.0197442	0.021548	
	T_c	0.11208	0.112381	0.11271	
Case-4	r_c	1.06666	0.87761	1.5642	
	P_c	0.0190022	0.0193334	0.0197237	
	T_c	0.111932	0.11203	0.11236	
Case-5	r_c P_c T_c	1.62636 0.0195648 0.11217	 	 	

²Here, for a given α , entropy $S = S(r_h)$ from Eq. (2.2).



FIG. 2. For case-2: (a) The blue arrows represent the vector field *n* on a portion of the $\theta - r$ plane. The critical points CP₁, CP₂, and CP₃ are located at $(r, \theta) = (1.21, \frac{\pi}{2}), (1.53, \frac{\pi}{2})$, and $(0.67, \frac{\pi}{2})$ marked with black dots, and they are enclosed with the contours C_1 , C_2 and C_3 , respectively. The contour C_4 does not enclose any critical point, while the contour C_5 encloses all the three critical points. (b) Ω vs ϑ for contours C_1 (green curve), C_2 (yellow curve), C_3 (red curve), C_4 (brown curve), and C_5 (gray curve).

B. Case 2: $Q < Q_C$

For simplicity, now onward we set the Gauss-Bonnet parameter $\alpha = 1$. In this case, we have three critical points (see Table II for the critical values at Q = 0.15) related to rich phase structure of small/large black hole phase transitions and the Gibbs free energy also exhibits the double swallow tail behavior [31,33]. For this case, the vector field *n* is plotted in Fig. 2(a), where it clearly shows the presence of three critical points. We construct five contours C_1 , C_2 , C_3 , C_4 , and C_5 , such that the contours C_1 , C_2 , and C_3 enclose the critical points CP₁, CP₂, and CP₃, respectively, while the contour C_4 does not enclose any critical point. The contour C_5 encloses all the three critical points (see the Table I for parametric coefficients of the contours.).

The behavior of the deflection angle $\Omega(\vartheta)$ for these contours is as shown in Fig. 2(b), which gives $\Omega(2\pi) = 2\pi, -2\pi, -2\pi, 0$, and -2π , for the contours C_1, C_2, C_3, C_4 , and C_5 , respectively. Then, the topological charges associated with the critical points are given by $Q_t|_{CP_1} = +1$, $Q_t|_{CP_2} = -1$, and $Q_t|_{CP_3} = -1$. Again the topological charge for the contour C_4 is zero as it does not enclose the critical point, whereas the topological charge for the contour C_5 is -1, as it encloses all the three critical points and since its topological charge is additive.

Again, following the proposal in [18], CP₁ is now a novel critical point, while CP₂, and CP₃, are conventional critical points. Therefore, in this case, the total topological charge of the system is $Q_t = Q_t|_{CP_1} + Q_t|_{CP_2} + Q_t|_{CP_3} = -1$.

C. Case 3: $Q_C < Q < Q_D$

In this case, we have three critical points which give rise to a nice phase structure consisting of small/intermediate/large black hole phase transitions, including the appearance of triple³ points [31,33]. The vector field *n* plotted in Fig. 3(a) shows the three critical points (see Table II for critical values at charge Q = 0.18).

We construct four contours C_1 , C_2 , C_3 , and C_4 , such that the contours C_1 , C_2 , and C_3 enclose the critical points CP₁, CP₂, and CP₃, respectively, while the contour C_4 encloses all the three critical points. The behavior of the deflection angle $\Omega(\vartheta)$ along these contours is shown in Fig. 3(b), from where we obtain that $\Omega(2\pi) = 2\pi, -2\pi, -2\pi, -2\pi$, for the contours C_1 , C_2 , C_3 , and C_4 , respectively. Then, the topological charges corresponding to the critical points are $Q_t|_{CP_1} = +1$, $Q_t|_{CP_2} = -1$, and $Q_t|_{CP_3} = -1$. Thus, the critical point CP₁ is a novel one, while the critical points CP₂, and CP₃, are conventional critical points.

Therefore, in this case also, the total topological charge of the system (given by the contour C_4) is $Q_t = -1$.

D. Case 4: $Q_D < Q < Q_B$

In this case also we have three critical points with the existence of small/large black hole phase transitions [31,33]. The vector field n, showing these critical points (see Table II for critical values at charge Q = 0.195), is plotted in Fig. 4(a).

³Here, we do not pursue the case of triple points.



FIG. 3. For case-3: (a) The blue arrows represent the vector field *n* on a portion of the $\theta - r$ plane. The critical points CP₁, CP₂, and CP₃ are located at $(r, \theta) = (1.13, \frac{\pi}{2}), (1.55, \frac{\pi}{2}), \text{ and } (0.79, \frac{\pi}{2})$ marked with black dots, and they are enclosed with the contours C_1, C_2 , and C_3 , respectively. The contour C_4 encloses all the three critical points. (b) Ω vs ϑ for contours C_1 (green curve), C_2 (yellow curve), C_3 (red curve), and C_4 (gray curve).

The critical points CP₁, CP₂, and CP₃, are enclosed by the contours C_1 , C_2 , and C_3 , respectively, and the contour C_4 encloses all the three critical points. For these contours C_1 , C_2 , C_3 , and C_4 , the deflection angle $\Omega(\vartheta)$ behavior is as shown in Fig. 4(b), from where we note that $\Omega(2\pi) = +2\pi, -2\pi, -2\pi, -2\pi,$ and -2π , respectively. Then, the corresponding topological charges of the critical points are $Q_t|_{CP_1} = +1$ (novel), $Q_t|_{CP_2} = -1$ (conventional), and $Q_t|_{CP_3} = -1$ (conventional).



FIG. 4. For case-4: (a) The blue arrows represent the vector field *n* on a portion of the $\theta - r$ plane. The critical points CP₁, CP₂, and CP₃ are located at $(r, \theta) = (1.06, \frac{\pi}{2}), (0.87, \frac{\pi}{2})$, and $(1.56, \frac{\pi}{2})$ marked with black dots, and they are enclosed with the contours C_1 , C_2 , and C_3 , respectively. The contour C_4 encloses all the three critical points. (b) Ω vs ϑ for contours C_1 (green curve), C_2 (red curve), C_3 (yellow curve), and C_4 (gray curve).

The total topological charge of the system is $Q_t = -1$, as in the previous case.

E. Case 5: $Q_B < Q$

In this case, there is only one critical point. The situation is similar to the van der Waals system, where the first order small/large black hole phase transition which terminates at this second order critical point [31,33].

For this case (see the Table II for critical values at charge Q = 0.3), the vector field *n* is plotted in Fig. 5(a), showing the critical point CP₁. We construct two contours C_1 , and C_2 , such that the contour C_1 encloses the critical point, while the contour C_2 does not. The topological charge must be nonzero for the contour C_1 and zero for the contour C_2 . The plot of the deflection angle $\Omega(\vartheta)$ for the contours C_1 , and C_2 , is as shown in Fig. 5(b). The angle $\Omega(\vartheta)$ decreases to reach -2π for the contour C_1 , whereas it decreases first, then increases and finally vanishes for the contour C_2 . Thus, as we expected, the topological charge for the contour C_1 is -1, and for the contour C_2 it is zero.

Therefore, in this case, the topological charge of the critical point CP_1 is $Q_t|_{CP_1} = -1$ (conventional critical point), which is the total topological charge of the system as well.

IV. NATURE OF THE CRITICAL POINTS

According to the proposal in [18], we have two types of critical points from the topological classification. The conventional critical point (for which the topological charge is $Q_t = -1$), and the novel critical point (for which the

topological charge is $Q_t = +1$). A further proposal of [18] is that, the conventional critical point indicates the presence of first-order phase transitions near it, while the novel critical point cannot serve as an indicator of the presence of a firstorder phase transition. However, this notion of the presence or absence of first-order phase transitions based on the classification of the critical point as conventional or novel, does not seem to be true in general. This can be inferred, for example, from our results in case-2 (i.e., Sec. III B).

In case-2, we obtained the topological charges of the three critical points CP₁, CP₂, and CP₃, as $Q_t|_{CP_1} = +1$ (novel), $Q_t|_{CP_2} = -1$ (conventional), and $Q_t|_{CP_3} = -1$ (conventional), respectively. As shown in Fig. 6(a), only the critical point CP₃ appears in the phase diagram, while the other two critical points do not appear, as they do not globally minimize the Gibbs free energy [31,33]. There exist a first-order phase transitions near the critical point CP₃ (conventional), and there are no first-order phase transitions near the critical point CP₁ (novel). However, even though the critical point CP₂ is a conventional one, there is no first-order phase transitions near it, which disaccords with the proposal in [18].

We can resolve this disagreement, if we classify the critical points in the following way. As the pressure increases, the novel critical point is the one from which new phases (stable or unstable) appear, whereas, the conventional point is the one at which the phases disappear, as can be seen from Figs. 6(b) and 7.

From the Fig. 6(b), we have three black hole phases (small, unstable intermediate, and large) when the pressure $P < P_{c1}$. We see the appearance of new phases from the



FIG. 5. For case-5: (a) The blue arrows represent the vector field *n* on a portion of the $\theta - r$ plane. The critical point CP₁ is located at $(r, \theta) = (1.63, \frac{\pi}{2})$ marked with black dot, and is enclosed with the contour C_1 . The contour C_2 does not enclose the critical point. (b) Ω vs ϑ for contours C_1 (red curve), and C_2 (green curve).



FIG. 6. For case-2: (a) Phase diagram showing the first-order phase transitions near the conventional critical point CP₃. (b) *T* as a function of r_h for different pressures (Inset: for $P > P_{c2}$.), showing the appearance of new phases (stable or unstable) near the novel critical point CP₁ (green dot), and disappearance of phases near the conventional critical points CP₂, and CP₃ (red dots). Dashed curves denote unstable black hole branches and solid curves denote stable black hole branches. Pressure of the isobars increases from bottom to top, where the critical pressures are P_{c1} , P_{c2} , P_{c3} with $P_{c1} < P_{c2} < P_{c3}$.

critical point CP₁ (novel) when we just increase the pressure $P > P_{c1}$, and we now have five phases (small, unstable intermediate, stable intermediate, unstable intermediate, large). On further increasing the pressure *P* to P_{c2} ,

two phases among five are disappearing at the critical point CP_2 (conventional), and we are left with three phases only. If we further increase the pressure *P* to P_{c3} , two phases among three are disappearing at the critical point CP_3



FIG. 7. For case-2: Behavior of the Gibbs free energy as a function *T* for different pressures, showing the appearance of new (stable or unstable) phases near the novel critical point CP₁ (with pressure P_{c1}), and disappearance of phases near the conventional critical points CP₂ (with pressure P_{c2}), and CP₃ (with pressure P_{c3}), on increase of pressure. (a) P = 0.0192, (b) P = 0.01968, (c) $P = P_{c2} = 0.01978$, (d) P = 0.025, $P_{c3} = 0.0321$, 0.034.



FIG. 8. *T* as a function of r_h for different pressures, showing the appearance of new phases (stable or unstable) near the novel critical point CP₁ (green dot), and disappearance of phases near the conventional critical points CP₂, and CP₃ (red dots). Dashed curves denote unstable black hole branches and solid curves denote stable black hole branches. Pressure of the isobars increases from bottom to top, where the critical pressures are P_{c1} , P_{c2} , P_{c3} with $P_{c1} < P_{c2} < P_{c3}$. (a) for case-3, (b) for case-4.

(conventional). This appearance/disappearance of phases near the critical points can be seen clearly from the Gibbs free energy behavior shown in Fig. 7.

For the cases 3 and 4, the appearance/disappearance of phases at the critical points can be seen from the Figs. 8(a), and 8(b), respectively. It can be checked that our novel proposal of the appearance/disappearance of phases from the nature of critical points is also valid for the Born-Infeld case [18] and the third order Lovelock gravity presented in Appendix B.

V. CONCLUSIONS

In this paper, we considered six dimensional charged AdS black holes in Gauss-Bonnet gravity in the extended phase space, where the black holes exhibit a rich phase structure admitting multicritical points, depending on the range of the black hole charge Q. We computed the topological charges corresponding to the critical points by following Duan's topological current ϕ -mapping theory and the proposal in [18]. Our findings are summarized below.

We first considered the case of charged AdS black hole by switching-off the Gauss-Bonnet (GB) coupling α (calculation presented in Appendix A). In this case, we have only one critical point for which the topological charge Q_t is found to be $Q_t|_{CP_1} = -1$, and thus this critical point is a conventional one, following the classification in [18].

Next, we considered the effect of GB coupling on the topological charge. Depending on the range of the black hole charge Q, we now have three critical points for $Q < Q_B$, and one critical point for $Q > Q_B$ (where, $Q_B = 0.2018$ at $\alpha = 1.$), and the corresponding topological charges are given in the Table III. Among the three critical points for $Q < Q_B$, we have one novel critical point and the other two are conventional critical points, while the critical point for $Q > Q_B$ is a conventional one. However, the total topological charge of the system is -1 independent of the range of the black hole charge Q.

Therefore, we find that the topological charge of the charged AdS black hole system is $(Q_t = -1)$ unaltered due to the Gauss-Bonnet coupling, unlike the Born-Infeld coupling [18]. Further, we found that as the pressure

TABLE III. Summary of the results.

Case	Number of critical points	Topological charge	Total topological charge Q_t
Case 1: $\alpha = 0$	1	$Q_t _{CP_1} = -1$	-1
Case 2: $Q = 0.15 < Q_C$	3	$Q_t _{\rm CP_1} = +1$	-1
		$Q_t _{\rm CP_2} = -1$	
		$Q_t _{\rm CP_3} = -1$	
Case 3: $Q_C < Q = 0.18 < Q_D$	3	$Q_t _{\mathrm{CP}_1} = +1$	-1
		$Q_t _{\mathrm{CP}_2} = -1$	
		$Q_t _{\mathrm{CP}_3} = -1$	
Case 4: $Q_D < Q = 0.195 < Q_B$	3	$Q_t _{\operatorname{CP}_1} = +1$	-1
		$Q_t _{\mathrm{CP}_2} = -1$	
		$Q_t _{\mathrm{CP}_3} = -1$	
Case 5: $Q_B < Q = 0.3$	1	$Q_t _{\mathrm{CP}_1} = -1$	-1

increases in phase space, new phases (stable or unstable) appear from the novel critical point, whereas the phases disappear at the conventional critical point. Thus, we name the novel critical point as the phase creation point, and the conventional critical point as the phase annihilation point. This removes some ambiguity in the connection between existence/absence of first order phase transition and classification of critical points based on topology.

First observation is with regards to the parity of critical points. Black holes in two different systems with same parity (odd-odd or even-even type) of the number of critical points, have the total topological charge matching with the parity. Examples to support this view are charged AdS black hole (1 critical point) and charged GB AdS black hole (1 or 3 critical points), where the parity of the number of critical points (odd in both cases) is same irrespective of the sign of topological charge. On the other hand, in the case of charged black holes in AdS (1 critical point) and Born-Infeld AdS black hole (2 critical points), the parity of the number of critical points is different, i.e., odd in the first and even in the same case, respectively. Correspondingly, the total topological charge in these cases matches the parity of number of critical points, i.e., odd (-1) for charged black holes and even (0) with Born-Infeld corrections. This trend continues for black holes in the third order Lovelock theories (Appendix B), where for odd (even) number of critical points the total topological charge is an odd (even) number. Of course, our study is preliminary and more examples are required to settle these claims.

An important result concerns the observations in [18], where for black holes in the Einstein-Maxwell system and their counterparts in the Einstein-Born-Infeld system, the critical points were shown to belong to different topological classes, as their topological charges are different. On the other hand, as we found in this work, that the charged black holes in the Gauss-Bonnet gravity belong to the same topological class as their counterparts in the Einstein-Maxwell system. As shown in Appendix B, for black holes in Lovelock theories of gravity up to third order the topological class of critical points remains unaltered. Naively, it appears that the higher derivative curvature corrections in the gravity sector do not change the topological class of black hole critical points, where as the gauge corrections do. There might be a deeper reason for this which requires further study, possibly by studying black holes in theories with both higher curvature gauge and gravity corrections side by side in the action. It would be interesting to explore this issue further from topology point of view, as it may teach us something new about the phase structure of black holes, which has been missed from standard thermodynamic treatments. There is no doubt that, the above conjectures require more scrutiny in systems having multiple critical points and with other gauge/gravity corrections, possibly involving rotating black hole solutions [33,41–45]. Further, it would be nice to have a mechanism to study the topological properties associated with the triple points in black holes as well [31,33,46].

ACKNOWLEDGMENTS

One of us (C. B.) thanks the DST (SERB), Government of India, for financial support through the Mathematical Research Impact Centric Support (MATRICS) Grant No. MTR/2020/000135. We thank the anonymous referees for helpful suggestions which improved the manuscript.

APPENDIX A: CASE-1 ($\alpha = 0$): TOPOLOGY OF EINSTEIN-MAXWELL SYSTEM IN SIX DIMENSIONS

The aim here is to compute the topological charge Q_t associated with the Einstein-Maxwell system by switchingoff the Gauss-Bonnet parameter (i.e., $\alpha = 0$) in Eq. (2.1) and this will help validate our results when the charge parameter is introduced in Sec. III A. This system has only one critical point, showing the van der Waals type small/large black hole phase transitions, given by [31,47]

$$T_c = \frac{9}{7\pi r_c}, \qquad P_c = \frac{9}{16\pi r_c^2}, \qquad r_c = \left(\frac{14Q^2}{3}\right)^{1/6}.$$
 (A1)

For this case, the vector field *n* is plotted in Fig. 9(a), where it shows the critical point at $(r, \theta) = (r_c, \frac{\pi}{2})$. We then construct two contours C_1 and C_2 such that the contour C_1 encloses the critical point, while the contour C_2 does not (see Table I for parametric coefficients of the contours.).

The deflection angle $\Omega(\vartheta)$ [using Eq. (3.7)] of the vector field ϕ along the contours C_1 and C_2 is plotted in Fig. 9(b). It shows the following features. For contour C_1 , it decreases and reaches $\Omega(2\pi) = -2\pi$, while for the contour C_2 , it decreases first and then increases, before finally vanishing, i.e., $\Omega(2\pi) = 0$. Therefore, the topological charge $Q_t =$ $\frac{1}{2\pi}\Omega(2\pi)$ for the contour C_1 is -1, while for the contour C_2 it is zero. These results are consistent with our expectations that, if a given contour encloses the critical point then its topological charge is nonzero. Therefore, the topological charge corresponding to the critical point CP1 is $Q_t|_{\rm CP_1} = -1$. According to the classification of the critical points from topology proposed in [18], this critical point is a conventional critical point. Since there exists only one critical point, the total topological charge of the Einstein-Maxwell system would be $Q_t = -1$.

APPENDIX B: TOPOLOGY OF CRITICAL POINTS IN THIRD ORDER LOVELOCK GRAVITY

Following the computation of topological charge and analysis in Secs. III and IV, we found that in the case of



FIG. 9. For case-1: (a) The blue arrows represent the vector field *n* on a portion of the $\theta - r$ plane. The critical point CP₁ is located at $(r, \theta) = (r_c, \frac{\pi}{2})$ marked with a black dot. The contour C_1 encloses the critical point, while the contour C_2 does not. (b) Ω vs ϑ for contours C_1 (red curve) and C_2 (green curve). (Here, we set the charge Q = 0.15).

Gauss-Bonnet black holes in AdS, the nature of critical points needs to be understood as phase creation and phase annihilation points. Here, our aim is to give a prefatory computation that the arguments in Sec. IV continue to hold in Lovelock theories of gravity as well. As an example, we consider the third-order Lovelock gravity but generalization to higher order should also be possible. Following [33], it is known that the third-order Lovelock coupling to Einstein-Maxwell system shows the standard van der Waals behavior with one critical point in seven dimensions, and thus its (total) topological charge would be -1 (conventional critical point). However, this system in higher dimensions (≥ 8) exhibits the rich phase structure with multicritical points [33]. In eight dimensions, this system can have up to three critical points (depending on

the charge q and Lovelock coupling parameter α), having the equation of state given by [33]:

$$p = \frac{t}{v} - \frac{15}{2\pi v^2} + \frac{2\alpha t}{v^3} - \frac{9\alpha}{2\pi v^4} + \frac{3t}{v^5} - \frac{3}{2\pi v^6} + \frac{q^2}{v^{12}}, \quad (B1)$$

where, v is the specific volume. The corresponding phase diagrams with one, two, and three critical points (see Table IV for critical values) are as shown in Figs. 10(a), 10(b), and 11, respectively.

In Figures 10(a), 10(b), and 11, the critical points shown in red color correspond to the phase annihilation points and thus their topological charge is -1. Where as, the critical point shown in green color corresponds to the phase creation point and hence its topological charge

			Critical point (CP)		
Parameters (q, α)	Number of	critical points	CP ₁	CP ₂	CP ₃
(0.01,2.6)	2	v_c	2.36519	0.4213	
		p_c	0.090224	2.46059	
		t_c	0.745863	0.72603	
(0.01,2.8)	3	v_c	0.8172	2.0874	0.423
		p_c	0.039	0.08996	2.7776
		t_c	0.7206	0.7304	0.7442
(0.01,3)	1	v_c	0.4241		
		p_c	3.1603		
		t_c	0.7621		

TABLE IV. Critical values for Lovelock gravity.



FIG. 10. For Lovelock gravity: t as a function of v for different pressures. The critical points shown in red color are phase annihilation points. Dashed curves denote unstable black hole branches and solid curves denote stable black hole branches. Pressure of the isobars increases from bottom to top. (a) with one critical point, (b) with two critical points (Inset: for $p > p_{c1}$).



FIG. 11. For Lovelock gravity: *t* as a function of *v* for different pressures. The critical points shown in red (green) color are phase annihilation (phase creation) points. Dashed curves denote unstable black hole branches and solid curves denote stable black hole branches. Pressure of the isobars increases from bottom to top, with three critical points (Insets: for $p > p_{c2}$).

is +1. Therefore, the total topological charge of the Einstein-Maxwell system with Lovelock coupling would be -1(with one or three critical points), or , -2 (with two critical points). Thus, similar to cases of

Gauss-Bonnet theories discussed in Sec. IV, we can advance the argument that the Lovelock coupling also does not alter the topological class of the critical points for black holes in Einstein-Maxwell system.

- [1] B. P. Abbott *et al.* (LIGO Scientific and Virgo Collaborations), Phys. Rev. Lett. **118**, 221101 (2017).
- [2] A. Sen, Gen. Relativ. Gravit. 40, 2249 (2008).
- [3] J. D. Bekenstein, Phys. Rev. D 7, 2333 (1973).
- [4] J. M. Bardeen, B. Carter, and S. W. Hawking, Commun. Math. Phys. 31, 161 (1973).
- [5] S. W. Hawking, Commun. Math. Phys. 43, 199 (1975); 46, 206(E) (1976).
- [6] S. W. Hawking and D. N. Page, Commun. Math. Phys. 87, 577 (1983).
- [7] J. M. Maldacena, Adv. Theor. Math. Phys. 2, 231 (1998).
- [8] S. S. Gubser, I. R. Klebanov, and A. M. Polyakov, Phys. Lett. B 428, 105 (1998).
- [9] E. Witten, Adv. Theor. Math. Phys. 2, 253 (1998).
- [10] A. Chamblin, R. Emparan, C. V. Johnson, and R. C. Myers, Phys. Rev. D 60, 064018 (1999).
- [11] M. M. Caldarelli, G. Cognola, and D. Klemm, Classical Quantum Gravity 17, 399 (2000).
- [12] D. Kastor, S. Ray, and J. Traschen, Classical Quantum Gravity 26, 195011 (2009).
- [13] M. Cvetic, G. W. Gibbons, D. Kubiznak, and C. N. Pope, Phys. Rev. D 84, 024037 (2011).
- [14] B. P. Dolan, Classical Quantum Gravity 28, 235017 (2011).
- [15] A. Karch and B. Robinson, J. High Energy Phys. 12 (2015) 073.
- [16] D. Kubiznak, R. B. Mann, and M. Teo, Classical Quantum Gravity 34, 063001 (2017).
- [17] D. Kubiznak and R. B. Mann, J. High Energy Phys. 07 (2012) 033.
- [18] S.-W. Wei and Y.-X. Liu, Phys. Rev. D 105, 104003 (2022).
- [19] N. Altamirano, D. Kubizňák, R.B. Mann, and Z. Sherkatghanad, Classical Quantum Gravity 31, 042001 (2014).
- [20] J. X. Lu, S. Roy, and Z. Xiao, J. High Energy Phys. 01 (2011) 133.
- [21] J.-y. Shen, R.-G. Cai, B. Wang, and R.-K. Su, Int. J. Mod. Phys. A 22, 11 (2007).
- [22] M. Cvetic, S. Nojiri, and S. D. Odintsov, Nucl. Phys. B628, 295 (2002).
- [23] R.-G. Cai, Phys. Rev. D 65, 084014 (2002).
- [24] N. Altamirano, D. Kubiznak, R. B. Mann, and Z. Sherkatghanad, Galaxies 2, 89 (2014).

- [25] N. Altamirano, D. Kubiznak, and R. B. Mann, Phys. Rev. D 88, 101502 (2013).
- [26] R.-G. Cai, L.-M. Cao, L. Li, and R.-Q. Yang, J. High Energy Phys. 09 (2013) 005.
- [27] G. Kofinas and R. Olea, Phys. Rev. D 74, 084035 (2006).
- [28] S.-W. Wei, Y.-X. Liu, and R. B. Mann, Phys. Rev. Lett. 123, 071103 (2019).
- [29] R. A. Hennigar, R. B. Mann, and E. Tjoa, Phys. Rev. Lett. 118, 021301 (2017).
- [30] R. Banerjee and D. Roychowdhury, Phys. Rev. D 85, 044040 (2012).
- [31] S.-W. Wei and Y.-X. Liu, Phys. Rev. D 90, 044057 (2014).
- [32] S.-W. Wei and Y.-X. Liu, Phys. Rev. D 87, 044014 (2013).
- [33] A. M. Frassino, D. Kubiznak, R. B. Mann, and F. Simovic, J. High Energy Phys. 09 (2014) 080.
- [34] M. Brigante, H. Liu, R. C. Myers, S. Shenker, and S. Yaida, Phys. Rev. D 77, 126006 (2008).
- [35] Y.-S. Duan and M.-L. Ge, Sci. Sin. 9, 1072 (1979).
- [36] Y.-S. Duan, Report No. SLAC-PUB-3301, 1984.
- [37] P. V. P. Cunha and C. A. R. Herdeiro, Phys. Rev. Lett. 124, 181101 (2020).
- [38] H. C. D. L. Junior, J.-Z. Yang, L. C. B. Crispino, P. V. P. Cunha, and C. A. R. Herdeiro, Phys. Rev. D 105, 064070 (2022).
- [39] S.-W. Wei, Phys. Rev. D 102, 064039 (2020).
- [40] P. V. P. Cunha, E. Berti, and C. A. R. Herdeiro, Phys. Rev. Lett. **119**, 251102 (2017).
- [41] S. H. Hendi and M. Momennia, Eur. Phys. J. C 78, 800 (2018).
- [42] M. Zhang, D.-C. Zou, and R.-H. Yue, Adv. High Energy Phys. 2017, 3819246 (2017).
- [43] M. Momennia and S. H. Hendi, Phys. Lett. B 822, 136692 (2021).
- [44] S.-W. Wei, P. Cheng, and Y.-X. Liu, Phys. Rev. D 93, 084015 (2016).
- [45] M.-D. Li, H.-M. Wang, and S.-W. Wei, arXiv:2201.09026 [Phys. Rev. D (to be published)].
- [46] N. Altamirano, D. Kubizňák, R. B. Mann, and Z. Sherkatghanad, Classical Quantum Gravity 31, 042001 (2014).
- [47] S. Gunasekaran, R. B. Mann, and D. Kubiznak, J. High Energy Phys. 11 (2012) 110.