Systematic study of hadronic excitation energy using the Schottky anomaly

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We extract the excitation energy scales of the hadron spectra in a less model-dependent method using the Schottky anomaly. The Schottky anomaly is a thermodynamical phenomenon that the specific heat of a system consisting of a finite number of energy levels has a peak at finite temperature due to the energy gaps. Using the masses of all hadrons that are experimentally established, we obtain the excitation energy scales of the hadron spectra and investigate their flavor dependence.

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I. INTRODUCTION

Understanding the hadron structure is one of the important subjects in hadron physics. In particular, the quest of effective constituents of hadrons can be a clue to investigate the hadron structure. For instance, the constituent quark explains the fine structure of the quarkonia [1,2] and the static properties, such as magnetic moment, of the light hadrons [3,4], even though it is not one of the fundamental elements of quantum chromodynamics. To identify effective degrees of freedom in composite objects, it is good to investigate excitation modes. Recently it has been pointed out in Refs. [5,6] that if one takes a diquark picture for the Λ_c baryon the diquark-quark confinement potential should be weaker than the quark-antiquark potential obtained in heavy quarkonia in order to reproduce the Λ_c excitation spectrum.

In this work, we visualize the excitation energy scales in various hadronic spectra in a less model-dependent way by examining specific heats obtained by hadronic mass spectra. This method is used widely in different areas. Reference [7] first applied for hadron physics to extract the effective degrees of freedom of the constituent of a hadron. This is based on the fact that the specific heat of a system is closely related to dynamical degrees of freedom in thermodynamics. This idea was also applied for nuclear systems to investigate their collective degrees of freedom [8]. Recently Ref. [9] investigated hadronic specific heats calculated by observed heavy meson spectra in order to extract different energy scales underlying the internal dynamics of hadron resonances for the identification of exotic hadrons. By following Ref. [9], we extend the study

to all of the observed hadrons, and investigate the hadron structure in a model independent way by comparing the visualized excitation energy scales systematically. In the hadron spectra, excitations in the confinement potential are major modes, while excitations induced by spin dependent forces, such as spin-spin and spin-orbit interactions, are minor and considered as fine structure. In this paper we try to extract such major excitation modes in order to investigate the nature of the confinement force.

The paper is organized as follows. In Sec. II, we explain the characteristics of the Schottky anomaly for two-state and *N*-state systems, and apply it to a hadronic model described with the Coulomb-plus-linear potential. In Sec. III, we perform calculations to visualize the excitation energy scales of hadrons using the Schottky anomaly, and discuss their flavor dependence. Finally, Sec. IV is devoted to a summary of this work and future prospects.

II. SCHOTTKY ANOMALY

The Schottky anomaly is a phenomenon in which the specific heat has a peak called the Schottky peak at a finite temperature. The temperature dependence of the specific heat is determined by excitation energies of the system when the number of energy levels in the system is finite. Now we focus on low excited states because observed hadrons are in low-lying states. In this section, we discuss the Schottky anomaly using some examples. First, as one of the simplest examples, we will discuss the two-state system. Next, we investigate the characteristics of the Schottky anomaly using more general systems. Finally, we will discuss the Schottky anomaly using some in a model close to observed hadronic systems.

A. Two-state system

Let us consider a two-state system with the ground state energy E_0 and the energy gap ΔE . The canonical partition function of the system is obtained as

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$$Z(\beta) = e^{-\beta E_0} (1 + e^{-\beta \Delta E}), \qquad (1)$$

where β is the inverse temperature. The expectation value of the energy of the system can be written as

$$\langle E \rangle = -\frac{1}{Z} \frac{\partial Z}{\partial \beta},\tag{2}$$

using the partition function Z. With this energy expectation value the specific heat of the system is written as

$$C = \frac{\partial}{\partial T} \langle E \rangle = k_B \beta^2 \frac{\partial^2}{\partial \beta^2} \ln Z, \qquad (3)$$

for $\beta = 1/k_B T$. Hereafter, we take unit of $k_B = 1$, and in this unit the specific heat is a dimensionless quantity, and the temperature *T* has the dimension of energy. From Eqs. (1) and (3), the specific heat of the two-state system is obtained as

$$C = (\beta \Delta E)^2 \left(2 \cosh \frac{\beta \Delta E}{2} \right)^{-2}.$$
 (4)

As can be seen in Eq. (4), the specific heat of the two-state system is given as a function of $\beta\Delta E$ and is independent on the energy of the ground state. The limiting values of the specific heat at sufficiently high and low temperatures are both 0. Therefore, the specific heat has a maximum value at a finite temperature. Figure 1 plots the specific heat C in the two-level system against $\beta\Delta E$. The specific heat takes a maximum value when $\beta\Delta E \simeq 2.4$ and converges zero at sufficiently high and low temperature. This behavior is called the Schottky anomaly, and this type of the specific heat is called the Schottky-type specific heat. If one observes the peak temperature of the specific heat, T_{peak} , one can extract the excitation energy by using $\Delta E = 2.4T_{\text{peak}}$.

The qualitative interpretation is as follows. For a two-level system with energies E_0 and E_1 , the probability to take the ground state $E = E_0$ at sufficiently low temperature is unity,

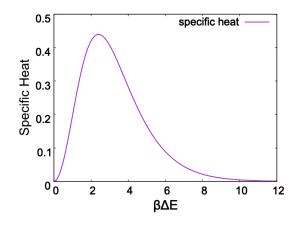


FIG. 1. Specific heat of the two-state system as a function of $\beta \Delta E$. The Schottky peak is seen at $\beta \Delta E \approx 2.4$.

Prob $(E = E_0) = 1$. If the temperature increases, the probability to take the excited state, Prob $(E = E_1)$, becomes large. For the transition from E_0 to E_1 the system receives heat from the exterior to excite the ground state with the energy gap ΔE , and consequently the specific heat gets large. If the temperature raises further, the probabilities for taking these two states become equal, Prob $(E = E_0) =$ Prob $(E = E_1)$, and the system does not need to receive heat from the outside. In this way, the specific heat increases and takes a peak at certain finite temperature corresponding to the energy gap ΔE .

B. *N*-state systems

Let us consider the specific heat of a system with several states and discuss the effect of the multiple energy gaps in the specific heat. The energy of the *k*th state is given by

$$E_k = E_0 + \Delta E_k \quad (k = 0, ..., N - 1), \tag{5}$$

and ΔE_k is the excitation energy of the *k*th state from the ground state and $\Delta E_0 = 0$.

Now, in proceeding with the discussion here, we show that the specific heat of the system generally does not depend on the value of the ground state energy, but only on the excitation energies. Using Eqs. (3) and (5) the partition function of the system can be written as

$$Z(\beta) = \sum_{k=0}^{N-1} e^{-\beta E_k} = e^{-\beta E_0} \sum_{k=0}^{N-1} e^{-\beta \Delta E_k},$$
(6)

From Eq. (2), the expected value of energy is

$$\langle E \rangle = E_0 - \frac{1}{z} \frac{\partial z}{\partial \beta}, \qquad (7)$$

where a reduced partition function z is defined by

$$z(\beta) \equiv \sum_{k=0}^{N-1} e^{-\beta \Delta E_k}.$$
(8)

The specific heat is obtained by differentiating both sides of Eq. (7) with temperature. At this time, the first term of the right-hand side of Eq. (7) does not contribute to the specific heat because it is a constant. The second term does not depend on the ground state energy E_0 . Therefore, the specific heat of the system is determined only by the excitation energies, and independent of the ground state energy.

We first consider the case for N = 3, in which there are two excitation energies, ΔE_1 and ΔE_2 . In this system, the reduced partition function z is given as

$$z(\beta) = 1 + e^{-\beta \Delta E_1} + e^{-\beta \Delta E_2}.$$
 (9)

Let us fix the second excitation energy as $\Delta E_2 = 400 \text{ MeV}$. We compare the specific heats calculated with $\Delta E_1 =$ 50 MeV and $\Delta E_1 = 200$ MeV. In the former case, two excitation energies have different magnitudes, while in the latter they are comparable. In Fig. 2, we plot the specific heats for $\Delta E_1 = 50$ MeV (a) and $\Delta E_1 = 200$ MeV (b). The solid lines show the specific heats of the three-state systems. For comparison, we also plot the specific heats of the two-state systems with ΔE_1 and ΔE_2 as the dashed and dashed-dotted lines, respectively. In Fig. 2(a) there are two separated peaks. Comparing it with the specific heat for the two-state systems, one finds that the peak at lower temperature corresponds to the smaller excitation energy ΔE_1 , while the peak at higher temperature to the larger excitation energy ΔE_2 . It is also seen in the figure that the height of the peak at the lower temperature is higher than that of the other. This implies that the specific heat of the threestate system is dominated by the lower excitation mode. Figure 2(b) shows the specific heat of the three-state system with $\Delta E_1 = 200$ MeV. Here one finds only one peak. It is interesting to note that the peak position is just between the

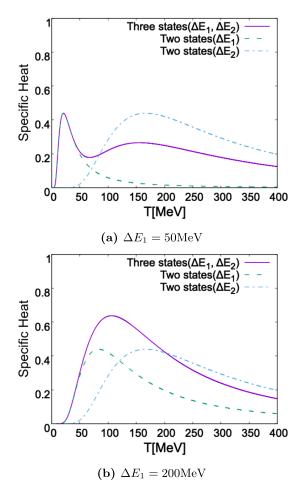


FIG. 2. The solid line in each figure shows specific heat when there are two excitation energies. Dashed lines and dashed-dotted line shows the specific heat when $\Delta E = \Delta E_1$, ΔE_2 (=400 MeV) with one excitation energy.

temperatures corresponding to the excitation energies ΔE_1 and ΔE_2 . Thus, when two excitation energies have similar values, which means that two levels have a similar energy, the specific heat has one peak at a temperature corresponding to an energy between two excitation energies. In this way we can extract an isolated excitation energy scale using the Schottky anomaly without looking at fine energy gaps.

Next we demonstrate the effect on the Schottky peak by higher excited states. For this purpose, we consider N-state systems with an equal energy gap, $\Delta E_k = k \Delta E$ with $\Delta E = 400$ MeV. Figure 3 shows the specific heats for a 2-state system (solid line), a 3-state system (dashed line), and a 4-state system (dotted-dashed line). The peak temperatures can be read as 167 MeV, 213 MeV, and 256 MeV for the 2-state, 3-state, and 4-state systems, respectively. The excitation energy scale of these systems is 400 MeV. This can be extracted in the 2-state system by reading the temperature of the peak position and multiplying it by 2.4. In the 3-state system, the Schottky peaks by these two excitation modes are overlapped and only one peak appears and the peak position is shifted about 20% higher. For the 4-state system, similarly, the peak position is shifted about 20% higher than that of the 3-state system. In this way, if there is a series of excitation modes, the effect of higher excited states on the peak position is seen in a 20% shift of the peak position. This tells us that the estimate of the energy scale by the Schottky anomaly has potentially about 20% error.

It is interesting to point out that the specific heat of the *N*-state systems with an equal energy gap can be expressed as a function of dimensionless parameter $x = \beta \Delta E$ as

$$C = x^2 \frac{\sum_k \sum_l k^2 e^{-(k+l)x} - (\sum_k k e^{-kx})^2}{(\sum_k e^{-kx})^2}.$$
 (10)

Therefore, the peak positions of the specific heat are determined by x independently of the explicit value of

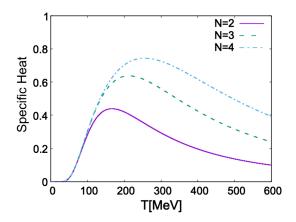


FIG. 3. Specific heats of the *N*-state systems with an equal energy gap $\Delta E = 400$ MeV. The solid, dotted, and dashed-dotted lines denote the specific heats of the 2-state, 3-state, and 4-state systems, respectively.

 ΔE . Because the temperature is scaled as $T = x\Delta E$, the estimation error discussed above is independent of the explicit value of ΔE .

C. Analysis of calculated data

In this subsection, we consider a more realistic spectrum using a potential model for the heavy quarkonia and investigate the behavior of the specific heat. The charmonium spectrum can be reproduced well by a central potential given by

$$V(r) = -\frac{4\alpha_s}{3r} + kr + V_0,$$
 (11)

with spin dependent forces as long as the $\overline{D}D$ threshold does not open [2]. Here, α_s , k, and V_0 are the potential parameters to be fitted by experimental observables. Here we use a parametrization given in Ref. [5] as an example. We calculate the energy eigenvalue of charmonia using only the central potential. The eigenvalue can be obtained by solving the eigenvalue equation

$$-\frac{1}{2\mu}\frac{d^2\chi_{\ell}(r)}{dr^2} + \left[V(r) + \frac{\ell(\ell+1)}{2\mu r^2}\right]\chi_{\ell}(r) = E_{n,\ell}\chi_{\ell}(r), \quad (12)$$

where μ is the reduced mass of the charm and anti-charm quarks. The charm quark mass is set as 1.5 GeV. Table I shows the energy eigenvalues obtained in the calculation. The explicit values of the parameters that we use are k = 0.9 GeV fm⁻¹ and $\alpha_s = 0.4$. Parameter V_0 is adjusted so that the ground state energy is to be zero. We also calculate the charmonium spectrum using k = 0.4 GeV fm⁻¹ in order to check whether the Schottky peak can distinguish the strength of the string tension. The partition function of this system can be written as

TABLE I. Calculated energy eigenvalue of charmonium for k = 0.9, 0.4 GeV fm⁻¹ in units of MeV, using $\alpha_s = 0.4$, $\mu = 0.75$ GeV.

State	$k = 0.9 \text{ GeV} \cdot \text{fm}^{-1}$	$k = 0.4 \text{ GeV} \cdot \text{fm}^{-1}$
1 <i>S</i>	0	0
1 <i>P</i>	430	287
2S	604	384
1 <i>D</i>	723	468
2P	890	564
1F	970	617
35	1043	648
2D	1125	707
1G	1191	709
3 <i>P</i>	1276	793
2F	1335	833
1 <i>H</i>	1395	870

$$Z(\beta) = \sum_{n,\ell} (2\ell + 1)e^{-\beta E_{n,\ell}}.$$
 (13)

Here, the coefficient is the degeneracy due to the magnetic quantum number. The specific heats calculated from these spectra are shown in Fig. 4. We summarize the values of the peak temperature T_{peak} in Table II. Comparing the results for different values of k, one can see that the Schottky peak can distinguish the size of the potential parameter.

In Sec. III, we will calculate the specific heats of the observed hadron spectra to obtain the excitation energies. There we consider the observed hadron spectra, which have up to the first or second excited states. We obtain the excitation energy scale by using $\Delta E = 2.4T_{\text{peak}}$ following the two-state system. Here, we verify the validity of this argument by performing a similar calculation using the spectrum obtained by the model calculation with $k = 0.9 \text{ GeV} \cdot \text{fm}^{-1}$. First, we obtain $\Delta E = 362 \text{ MeV}$ as a result of calculating the excitation energy scale using the states up to 1P. This is about 20% different from 430 MeV that is the energy gap of in the ground state and the 1P states.

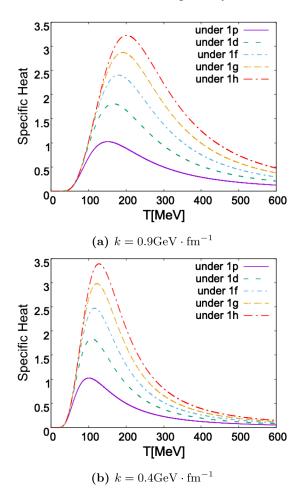


FIG. 4. Specific heat of the charmonium system calculated by a quark model with potential (11).

TABLE II. Peak temperature of the specific heats in units of MeV.

	k = 0.9	$k = 0.9 \text{ GeV} \cdot \text{fm}^{-1}$		$GeV \cdot fm^{-1}$
States	$T_{\rm peak}$	$2.4T_{\text{peak}}$	T_{peak}	$2.4T_{\text{peak}}$
Up to 1P	151	362	101	242
Up to $1D$	166	398	109	262
Up to $1F$	180	432	116	278
Up to $1G$	191	458	122	293
Up to $1H$	202	485	128	307

due to the orbital angular momentum $\ell = 1$. In the application for the observed spectra, we extract the excitation energies without specifying the angular momentum of the excited states by using $\Delta E = 2.4T_{\text{peak}}$. Thus this may underestimate the excitation energy about 20% and we accept this difference as uncertainty of the estimation. We also discuss the extraction uncertainty coming from the number of the observed states. Let us take the potential (11)again as an example. If we consider the excited states up to the 1D state of the potential (11), the peak temperature for this spectrum is about 10% larger than that for the spectrum up to the 1P state according to Table II. Similarly the peak temperature for the spectrum up to the 1F state is about 10% larger than that up to 1D state. Considering the fact that only some lower excited states have been observed, we evaluate the uncertainty of the estimation of the excitation energies to be 20%.

III. APPLICATION TO HADRON

In this chapter, we apply the Schottky anomaly for the observed hadron spectra and visualize a typical hadron excitation energy scale for each hadron spectrum. Furthermore, we investigate the hadron structure systematically by comparing the excitation energy scales of each spectrum. We note that the specific heat and temperature treated here do not represent the thermodynamic feature of hadrons but are merely mathematical parameters. For a hadronic system, the canonical partition function of the system is given as

$$Z_{\text{hadron}}(\beta) = \sum_{i} (2J_i + 1)(2I_i + 1)e^{-\beta E_i}, \quad (14)$$

using observed hadron mass E_i with the degeneracy of the total angular momentum J and the isospin I. Here we use the hadron masses without their decay width. The specific heat of the system is expressed with the partition function as

$$C_{\text{hadron}} = \beta^2 \frac{\partial^2}{\partial \beta^2} \ln Z_{\text{hadron}}.$$
 (15)

By plotting C_{hadron} against temperature we read the Schottky peak temperature T_{peak} and then extract excited energy scales of the system by

$$\Delta E = 2.4T_{\text{peak}}.\tag{16}$$

Here we presume that the observed hadron spectra have up to first or second excited states. It is important that the first excited states should be observed enough to obtain reliable results.

For the calculation of the specific heats we use the hadron masses shown by the Particle Data Group [10] and pick up the hadrons of which total angular moment and isospin have been determined and of which existence is sufficiently confirmed experimentally with • for mesons and with either symbol **** or *** for baryons in the summary tables. We have summarized the list of states used in the calculations in this chapter in the Appendix.

A. Heavy mesons

In this section, we calculate the excitation energy scales of heavy mesons using the Schottky anomaly. First we perform calculations for heavy qurkonia and then for open charm and open bottom states. Finally we discuss the flavor dependence of heavy meson energy scales from these results. Reference [9] calculated the specific heat of the heavy mesons to search for exotic mesons.

1. Heavy quarkonia

In the calculation we use the charmonia listed in Table IX. We calculate the specific heat using these charmonium states with the spin degeneracy. Figure 5 shows the results of calculating the specific heat of the charmonia as a function of 2.4T. The solid line in Fig. 5 shows the specific heat calculated with all of the charmonia

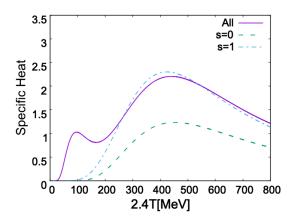


FIG. 5. Specific heat of charmonium. The solid line shows the specific heat of all charmonium system. The dashed and dashed-dotted lines denote the specific heats calculated by using charmonia with s = 0 and s = 1 separately, respectively. The separation is done in a way written in the text.

in the list. We find that the solid line has two Schottky peaks. From these peaks we read the excitation energies $\Delta E = 106$ MeV and 441 MeV for the lower and higher peaks, respectively. As will be described later, we consider that the peak on the low temperature corresponds to the hyperfine structure due to the spin-spin interaction and the peak on the high temperature corresponds to the orbital excitation due to the central potential.

In order to understand the nature of the excitation energies, we separately plot the specific heats for the charmonia by the spin configuration of the quarks, s = 0and s = 1. If the charmonia are assumed to be $c\bar{c}$ states, the spin configuration can be specified by parity *P* and charge conjugation *C* as follows. The parity of the meson can be written as

$$P = (-1)^{\ell+1},\tag{17}$$

using the orbital angular momentum ℓ . This is because orbital motion gives parity $(-1)^{\ell}$ and $q\bar{q}$ has negative internal parity in addition. Next for *C* parity, because taking charge-conjugation of $q\bar{q}$ is same as the replacement of quark and antiquark, the charge conjugate *C* is found as

$$C = (-1)(-1)^{\ell}(-1)^{s+1},$$
(18)

for the orbital angular momentum ℓ and the spin *s*. With Eqs. (17) and (18) one finds

$$C = P(-1)^{s+1}. (19)$$

This implies that P = C for s = 1 and P = -C for s = 0.

In Fig. 5 the results of the specific heats for s = 0 and s = 1 are shown as the short dashed line and dash-dotted line, respectively. We see that the lower peak disappears in both plots and find the excitation energies $\Delta E = 453$ MeV for s = 0 and 424 MeV for s = 1. This means that the lower peak seen in the full spectrum should correspond to the hyperfine splitting induced by the spin-spin interaction between quark and antiquark, and that the higher peak in the full spectrum is common for both spin configuration.

Let us move to the calculation for bottomonium. Figure 6 shows the results of the specific heat for bottomonium. In the calculation we use the bottomonia shown in Table X. The solid line in Fig. 6 shows the specific heat calculated with all of the bottomonia in the list. Again, we obtain two excitation energy scales, and the values of each energy scale are 51 MeV and 445 MeV. Similarly to charmonium, we plot separately the specific heats for S = 0 and S = 1 in Fig. 6 as dashed and dot-dashed lines, respectively. It is considered that the peak on the low temperature corresponds to the hyperfine structure and the peak on the high temperature corresponds to the orbital excitation, as in the case of charmonium. The energy scales of orbital excitation are the nergy scales of the orbital excitation are

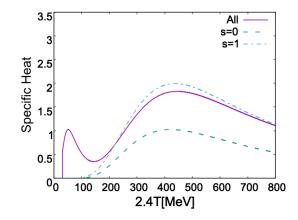


FIG. 6. Same as in Fig. 5 but for bottomonium.

insensitive to the flavor for the heavy quarkonia. On the other hand, the energy scale of the hyperfine structure of bottomonium is about half that of charmonium. This can be explained by the fact that the magnitude of splitting due to spin-spin interaction can be written as,

$$\Delta E_{\rm hf} = \frac{32\pi}{9} \frac{\alpha_s}{m_a m_{\bar{a}}} |\Psi(0)|^2, \qquad (20)$$

using quark model [11], and the splitting decreases as the quark mass increases. The excitation energy scales obtained here are summarized in Table III.

2. Open charm and open bottom mesons

Here we consider open charm and open bottom mesons. In the calculation we use the charmed mesons and bottomed mesons shown in Tables XI and XII, respectively.

First, we calculate the specific heat of the open charm system. Figure 7(a) shows the specific heat calculated with all of the open charm mesons. There are two peaks again and the corresponding excitation energies are read as 108 MeV and 344 MeV for lower and higher peaks, respectively. We will see that the low temperature peak is due to the mass difference between the *s* quark and the *u* or *d* quark, or hyperfine structure.

We plot separately the D_s and D mesons in Figs. 7(b) and 7(c) respectively. The solid lines in both figures show the

TABLE III. Excitation energy scales of the heavy mesons obtained by the Schottky anomaly in units of MeV.

	$c\bar{c}$	C = 1, S = 1	C = 1, S = 0
All	$ \begin{array}{r} 106, 441 \\ s = 0 s = 1 \\ 453 424 \\ b\bar{b} \end{array} $	143, 338 D_s D_s^* 391 404 B' = 1, S = 1	119, 353 D D^* 412 384 B' = 1, S = 0
All	51, 445 s = 0 $s = 1446 437$	$\begin{array}{ccc} 40, \ 386 \\ B_s & B_s^* \\ 390 & 395 \end{array}$	39, 378 <i>B B</i> * 378 393

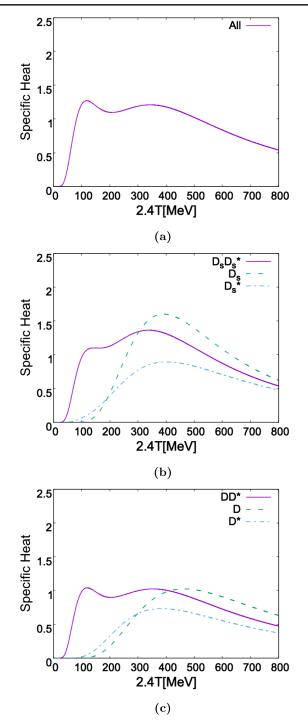


FIG. 7. Specific heats of the open charm systems. (a) Specific heat calculated with all the open charm states given Table IX. (b) Specific heats of the open charm states with strangeness. The solid line stands for the specific heat of all the strange charm mesons, while the dashed and dashed-dotted lines show the specific heats of the D_s and D_s^* mesons, respectively. (c) Specific heats of the nonstrange open charm states. The solid line stands for the specific heat of all the nonstrange charm mesons, while the dashed and dashed-dotted lines show the specific heat of all the nonstrange charm mesons, while the dashed and dashed-dotted lines show the specific heats of the D and D^* mesons, respectively.

specific heat calculated with all of the D_s and D mesons. Similar to the charmonium, two peaks appear for both cases, and the values of the energy scale are 119 MeV and 353 MeV for strange open charm and 143 MeV and 338 MeV for nonstrange open charm.

In order to see the origins of these peaks, we separately calculate the specific heats of the D_s and D_s^* mesons and the D and D^* mesons, in which the D_s and D states have $J^p = 0^-, 1^+, 2^-, \dots$, while D^*_s and D^* have $J^p =$ $0^+, 1^-, 2^+, \dots$ Although this is not a classification of the quark spin configuration, we can resolve the splitting of the ground states induced by the spin-spin interaction. The separated specific heats are plotted in Figs. 7(b) and 7(c) as dashed lines for D_s and D and dashed-dotted lines for D_s^* and D^* . The figures show only one peak appears at higher temperature for each line. This implies that the peaks appearing at lower temperature in the solid lines of Figs. 7(b) and 7(c) are considered to be the energy scale of the hyperfine splitting as seen in quarkonia. It is notable that the energy scales of the hyperfine splitting read 143 MeV for D_s and 119 MeV for D. This is opposite to the quark mass dependence of the hyperfine splitting Eq. (20). This would implies that the overlap of the wave function $|\Psi(0)|^2$ for D_s be larger than that for D. Regarding the peaks at higher temperature, the energy scales read 391 MeV for D_s , 402 MeV for D_s^* , 464 MeV for D, and 384 MeV for D^* from the dashed and dashed-dotted lines in Figs. 7(b) and 7(c). These are larger than the energy scales seen at higher temperature in the solid lines. That is, the excitation energy scale found in the separated specific heats for $D - D^*(D_s - D_s^*)$ is larger than that of the higher peak in the unseparated specific heat. According to the discussion in Sec. II B, it is expected that there is some energy scale corresponding to the difference between $D(D_s)$ and $D^*(D^*_s)$ and it is seen as one peak together with the orbital excitation in the unseparated specific heat. For the origin of the energy scale, we need detailed model calculations.

Next, we calculate the specific heat of the open bottom system. Figure 8(a) shows the specific heat calculated with all of the open bottom mesons and there are two peaks again and the corresponding excitation energies are read as 41 MeV and 367 MeV for lower and higher peaks, respectively.

We plot separately the B_s and B mesons in Figs. 8(b) and 8(c) respectively and the solid lines in both figures show the specific heat calculated with all of the B_s and B mesons. Similar to open charm, two peaks appear for both case, and the energy scale values are 40 MeV and 386 MeV for strange open bottom and 39 MeV and 378 MeV for nonstrange open bottom.

In order to see the origins of these peaks, we separately calculate the specific heats of the B_s and B_s^* mesons and the B and B^* mesons. The separated specific heats are plotted

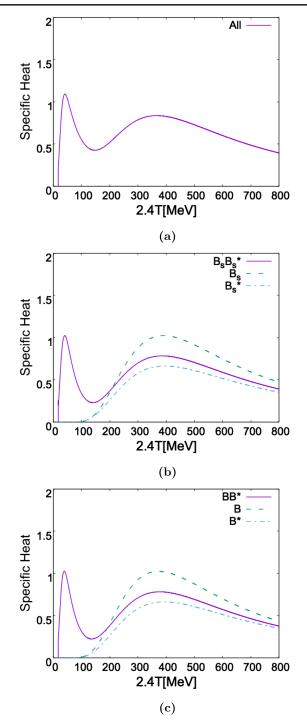


FIG. 8. Same as in Fig. 7 but for the open bottom systems.

in Figs. 8(b) and 8(c) as dashed lines for B_s and B and dashed-dotted lines for B_s^* and B^* . The figures show only one peak appears at higher temperature for each line. This implies that the peaks appearing at lower temperature in the solid lines of Figs. 8(b) and 8(c) are considered to be the energy scale of the hyperfine splitting as seen in open charms. Regarding the peaks at higher temperature, the energy scales read 390 MeV for B_s , 395 MeV for B_s^* , 378 MeV for B, and 393 MeV for B^* from the dashed and

dashed-dotted lines in Figs. 8(b) and 8(c). These energy scales are a bit larger than those seen at higher temperature in the solid lines. We do not find a strong indication for the extra energy scale found in the open charm systems. This would imply that the energy scale corresponding to the difference of $B - B^*(B_s - B_s^*)$ should be small. Alternatively, the difference of the energy scale between the separated and unseparated specific heats might be caused by lack of open bottom states used in the calculation(see Sec. II C).

3. Summary of heavy mesons

We summarize the excitation energy scale of charmonium, bottomonium, open charm and open bottom system in Table III. In the heavy meson systems, we have found two energy scales for hyperfine structure by spin-spin interaction and the orbital excitation by central potential. Comparing the results of charmonium and bottomonium, we find that the energy scale of the orbital excitation is about 440 MeV, and it is insensitive to the flavor of the heavy mesons. On the other hand, the energy scale of hyperfine structure becomes smaller as the quark mass increases. For the open charm and bottom systems we obtain a similar result of the quarkonia, that is, the two peaks are found in the specific heats for all the open heavy mesons; the higher peaks correspond to the orbital excitations, while the lower peaks to the hyperfine splitting. After separating the open heavy mesons in terms of strangeness and spin, we find that the excitation energies corresponding to the orbital motion are about 400 MeV independently of their flavor.

B. Light mesons

In this section, we calculate the specific heat for light mesons. First of all, let us start with the calculation of the specific heat for all the light mesons summarized in Table XIII. For the light mesons, there are several states out of regularity such as Nambu-Goldstone bosons and exotic candidates. These irregular mesons could spoil the systematic extraction of the excitation energies. In Fig. 9(a)we show the specific heat obtained by all the light mesons in the solid line. The plot shows that there is only one peak structure at 554 MeV. This energy scale is much larger than the global excitation energy seen in the heavy mesons. We expect that the Schottky peak in Fig. 9(a) contains several excitation energies. To resolve these energies, we first subtract the Nambu-Goldstone bosons $(\pi, K, \text{ and } \eta)$ from the spectrum and calculate the specific heat without them. As seen in the dashed line of Fig. 9(a), there are two peaks at 178 MeV and 393 MeV. Next we remove the exotic candidates, $f_0(500)$ and $K_0^*(700)$, from the calculation. The specific heat is shown in the dashed-dotted line of Fig. 9(a). This plot shows two excitation energies at 13 MeV and 432 MeV. The lower peak comes from the energy difference between $\rho(770)$ and $\omega(782)$.

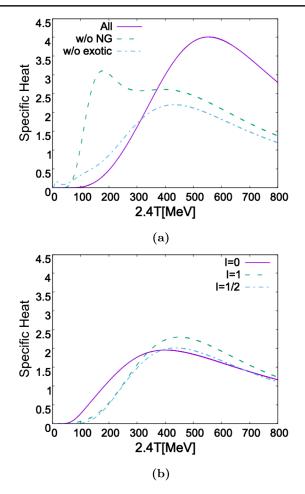


FIG. 9. (a) Specific heats for the light mesons. The solid line denotes the specific heat calculated with all the light mesons given in Table XI. The dashed line stands for the specific heat calculated without the Nambu-Goldstone bosons, π , K, η , while the dashed-dotted line shows the one without the Nambu-Goldstone bosons and the scalar exotics, $f_0(500)$ and $K_0^*(700)$. (b) Specific heats of the light mesons for each isospin. The specific heats are obtained without the Nambu-Goldstone bosons and the scalar exotics. The solid, dashed, and dashed-dotted lines show the specific heats of the light mesons with I = 0, I = 1, and I = 1/2, respectively.

The position of the higher peak is closed to that of the dashed line. Thus, we regard that the higher excitation energy is a typical excitation energy scale of the light mesons. It is interesting that this energy scale is very similar with what we find in heavy mesons.

Next we decompose the light mesons in terms of their isospin. Again we remove the Nambu-Goldstone bosons and the two scalar mesons from the calculation. The solid line in Fig. 9(b) shows the specific heat of the I = 0 states and the peak position is read as 401 MeV. Similarly the specific heats of the I = 1 and I = 1/2 are shown in the dashed and dashed-dotted lines of Fig. 9(b) respectively, and the excitation energies are read as 450 MeV and 438 MeV. These excitation energy scales are similar to that we have obtained using the light mesons before the

TABLE IV. Excitation energies of the light mesons obtained by the Schottky anomaly in units of MeV.

	All	w/o π, K, η	w/o, $\pi, K, \eta, f_0(500), K_0^*(700)$
All	554	178, 393	13, 432
I = 0			401
I = 1			450
I = 1/2			438

isospin decomposition. From these results excluding the Nambu-Goldstone bosons and the scalar mesons, the typical excitation energy scale of the light mesons is estimated to be about 400 to 450 MeV. We summarize the excitation energy scales of the light meson systems in Table IV. Considering this result together with those of the heavy mesons (Table III), we find flavor independence in the excitation energy scale of the meson systems.

C. Baryons

In this section, we calculate the Schottky peaks of baryon. First, we examine the light baryons, and next we consider the heavy baryons. Finally we discuss the flavor dependence of the baryon energy scales from these results.

1. Light baryons

We calculate the specific heat using all of the light baryons listed in Tables XIV and XV. In Fig. 10 we show the specific heats calculated with all the light baryons and find only one peak at 324 MeV. In this peak, several excitation energies are expected to be contributed because the hyperfine splitting can be larger than heavy quark systems according to Eq. (20).

To resolve the compound excitations, we decompose the light baryons in terms of their strangeness. In Fig. 11, we show the specific heat for the *N* and Δ baryons in the solid line. Only one peak is seen at 384 MeV. Further, we decompose the spectrum into the *N* and Δ sectors in order to resolve possible spin-spin interactions seen in the ground

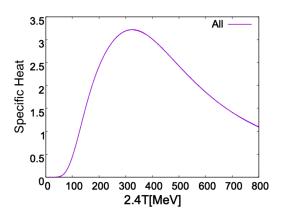


FIG. 10. Specific heat of light baryon.

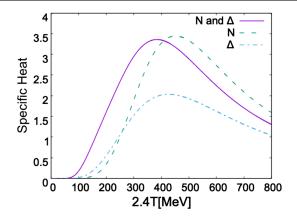


FIG. 11. Specific heats of the light unflavored baryons. The solid line stands for the specific heat calculated with all the unflavored baryons, while the dashed and dashed-dotted lines show those of the nucleons and the Δ baryons, respectively.

states of *N* and Δ . The decomposed specific heats are shown in the dashed line for *N* and the dashed-dotted line for Δ of Fig. 11. We find that the corresponding Schottky peaks are 452 MeV and 427 MeV, respectively. These values are larger than the peak position of the solid line where the *N* and Δ are included into the calculation collectively. This is because the mass difference of *N*(940) and Δ (1232), which is around 270 MeV, is also seen in the Schottky peak of the *N* – Δ spectrum. Although the *N*(940) – Δ (1232) mass difference stems from the hyperfine splitting, one cannot resolve all the effects of the hyperfine splitting in the decomposed specific heats.

Next we consider hyperons with S = -1. The solid line in Fig. 12(a) shows the specific heat for the Λ and Σ spectra and we find two peaks at 66 MeV and 358 MeV. The lower peak is considered to be the mass difference of $\Lambda(1115)$ and $\Sigma(1190)$, because this peak disappears when we consider the specific heat separately for Λ and Σ after performing the isospin decomposition. The separated specific heats for Λ and Σ are plotted as the dashed and dashed-dotted lines in Fig. 12(a) respectively. We find that they have only one Schottky peak located at 372 MeV and 373 MeV, respectively. Since the specific heat of the Σ system has a shoulder on the low temperature, it can be expected that an energy scale other than orbital excitation appears. This may be considered to be spin-spin splitting between $\Sigma(1190)$ and $\Sigma(1385)$. But, we cannot separate this effect by the model independence way. We also plot the specific heat for Λ without the $\Lambda(1405)$ resonance, which is a candidate of hadronic composite states [12-17], in Fig. 12(a) as the dotted line. We find the excitation energy scale to be 392 MeV.

Finally, we will analyze the Ξ system. Unfortunately there are only several states confirmed experimentally. The specific heat is shown in Fig. 12(b), and the excitation energy scale is found to be 326 MeV. Only one peak appears, being considered to be the average energy scale of

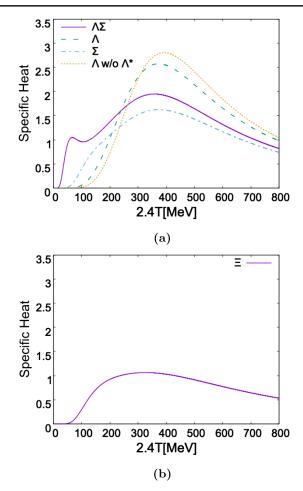


FIG. 12. Specific heats of the hyperons. (a) Specific heats of the hyperons with strangeness S = -1. The solid line shows the specific heats for the Λ and Σ baryons, while the dashed and dashed-dotted lines stand for those for the Λ and Σ baryons, respectively. The dotted line shows the specific heat of Λ without $\Lambda^*(1405)$. (b) Specific heat of the hyperons with strangeness S = -2.

hyperfine structure and orbital excitation like other baryons. We cannot separate the effect of the ground state hyperfine structure by a method with model independence.

2. Heavy baryons

In this section, we investigate the Schottky peaks for heavy baryons listed in Table XVI. In Fig. 13(a) we show the specific heat for the Λ_c and Σ_c baryons in the solid line. Here there are only one peak in the $\Lambda_c - \Sigma_c$ spectrum at 148 MeV. This peak may be contributed mainly by the mass difference between the ground states of Λ_c and Σ_c together with their internal excitations. In order to see the internal excitations in the Λ_c baryon, we show separately the specific heat of the Λ_c baryon in the dashed line of Fig. 13(a) and find a Schottky peak at 329 MeV. For Σ_c , we do not have enough states to investigate the internal excitation energy.

Figure 13(b) shows the results of the calculation of the specific heats for the Ξ_c system, which contains one charm

TABLE V. Excitation energy scales of the baryons obtained by the Schottky anomaly in units of MeV. Symbol S, C, and B' stand for strangeness, charm and bottom quantum numbers, respectively. The value in the bracket is the excitation energy scale of the Λ system without $\Lambda(1405)$.

Light baryons	S = 0	S = -1	S = -2
All	384	66, 358	326
$I = \frac{1}{2}, 0$	452	372 (392)	
$I = \frac{5}{2}, 1$	427	373	
Heavy baryons	C = 1, S = 0	C = 1, S = -1	B' = -1, S = 0
All	148	154	143
I = 0	329		290

quark and one strange quark. The energy scale is found to be 154 MeV. This peak may be considered to be a mixture of orbital excitation and hyperfine splitting as with other baryons.

Finally, the same analysis is performed for the $\Lambda_b - \Sigma_b$ system as the charmed baryons. The specific heat is shown in Fig. 13(c) as the solid line and we find the energy scale to be 143 MeV. Again, this may correspond to the mass difference between the ground states of Λ_b and Σ_b . We also show the specific heat calculated with only the Λ_b states in Fig. 13(c) as the dashed line. This plot shows that the energy scale is 290 MeV.

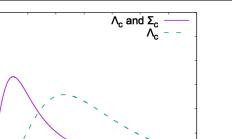
3. Discussion for the baryon system

Table V summarizes the baryon excitation energies extracted by the Schottky anomaly. As seen above, it is hard to separate out the hyperfine splitting in the baryon excitations in a model independent way. In order to investigate the flavor dependence of the energy scale of the baryons systematically, we compare the energy scales of N, Λ , Λ_c , and Λ_b . The Λ , Λ_c , and Λ_b baryons have isospin I = 0. This means that the *ud* system in the ground states of Λ s has the spin zero and the spin-spin force may not contribute to the mass spectrum. The ratios of the excitation energy to the nucleon one is found to be 1.00, 0.82, 0.73, 0.64 for N, Λ , Λ_c , and Λ_b , respectively.

In order to extract the systematics of the excitation energies among the N, Λ , Λ_c , and Λ_b baryons, we investigate the scaling law of the excitation energy. We assume that the excitation energy can be described by the power of a mass scale of the system:

$$\Delta E(M) = \frac{a}{(M/\Lambda)^b},\tag{21}$$

where a and b are constant parameters, M is a mass of interest and Λ is a normalization of the mass scale. Here we



3.5

3

2.5

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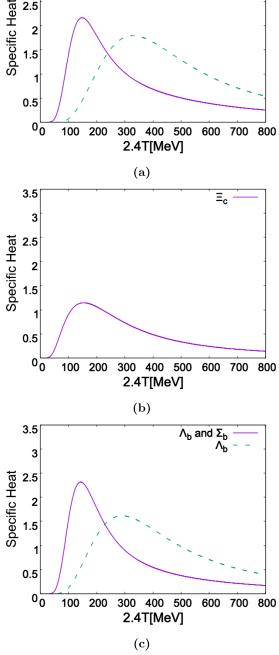


FIG. 13. Specific heats of the heavy baryons. (a) Specific heats of the charmed baryons with strangeness S = 0. The solid line shows the specific heats for the Λ_c and Σ_c baryons, while the dashed line stands for that for the Λ_c baryons only. (b) Specific heat of the Ξ_c baryon. (c) Specific heats of the bottomed baryons with strangeness S = 0. The solid line shows the specific heats for the Λ_b and Σ_b baryons, while the dashed line stands for that for the Λ_b baryons only.

set $\Lambda = 1$ MeV. We try several models to investigate the systematics of the excitation energy.

First, we consider the excitation energy to be a function of the ground state mass M_0 for each system. We fit the excitation energies by Eq. (21) with $M = M_0$ and find out the values of *a* and *b*. The ground state masses used in the fitting are shown in Table VI. The fitted values are a =2034 MeV and b = 0.23. This implies that $\Delta E \propto M_0^{-1/4}$.

Next, we consider ΔE to be a function of a constituent quark mass M_q . These baryons have the up and down quarks in common and one of the quarks are different among them. Therefore, these baryons can be specified by u or d, s, c, and b quarks, respectively. The excitation energy can be regarded as a function of these constituent quarks. We fit the excitation energy by Eq. (21) with $M = M_0$. In Table VI, we show the values of the constituent quark masses used in the fitting and the fitted values of a and b. The fitted value of b tells us that $\Delta E \propto M_a^{-1/6}$.

Finally, we take a diquark picture for these baryons. We regard the *u* and *d* quarks in the baryon to form a diquark and these baryon are considered to be a two-body system of the diquark and a quark. In this picture the excitation energy can be a function of the reduced mass μ of the diquark and the quark. Here we assume the constituent diquark mass as $M_{dq} = 500$ MeV and the constituent quark masses as shown in Table VI. In Fig. 14(a) we show the excitation energy as a function of the reduced mass. Fitting the excitation energy by Eq. (21) with $M = \mu$, we obtain a = 5238 MeV and b = 0.47. This implies that $\Delta E \propto \mu^{-1/2}$. Because this dependence is same as a simple harmonic oscillator, the system can be regarded as a *ud* diquark and quark two-body system with a harmonic oscillator potential. From the fitted value *a* the string constant of the oscillator is found to be $k = 707 \text{ MeV/fm}^2$. We perform the same calculation without $\Lambda(1405)$, and we obtain a = 5730 MeV and b = 0.49. As a result, the tendency of $\Delta E \propto \mu^{-1/2}$ appears to be more probable. In Fig. 14(a) we also plot the excitation energies for A without $\Lambda(1405)$ and fitted excitation energy as the dashed line. This plot shows that the excitation energy of Λ without

TABLE VI. Parameters used for discussion in Sec. III C 3. ΔE is the excitation energy scale obtained by the Schottky anomaly. M_0 is the ground state mass of each baryon. M_q is an assumed values of the constituent quark mass of each baryon. μ is the reduced mass of the *ud* diquark and the constituent quark mass, where the diquark mass is assumed to be $M_{dq} = 500$ MeV.

	$\Delta E(MeV)$	M_0 (MeV)	M_q (MeV)	μ(MeV)
N	452	940	300	188
Λ	372	1116	450	237
Λ_c	329	2286	1500	375
Λ_b	290	5620	4000	444
a(MeV)		2034	1073	5238
b	•••	0.2306	0.160	0.473

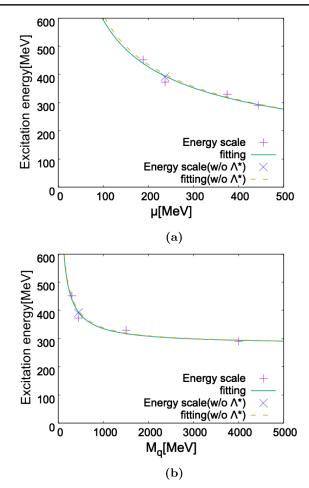


FIG. 14. (a) Excitation energy scales of baryon as a function of the reduced mass of the ud diquark and the constituent quark. The plus symbol points are the excitation energy obtained by the Schottky anomaly, while the solid line denotes the calculation using Eq. (21) with the fitted values a = 5238 MeV and b = 0.47. The diquark mass is assumed to 500 MeV. The constituent quark masses used in the plot are shown in Table VI. The times symbol point is the excitation energy of Λ without $\Lambda^*(1405)$ and the dashed line is the calculation Eq. (21) with a = 5730 MeV and b = 0.49, which is obtained without $\Lambda^*(1405)$. (b) Excitation energy scales of baryon as a function of the constituent quark mass M_q . The solid line is calculated using Eq. (22) with the fitted values $k = 890 \text{ MeV}/\text{fm}^2$ and $M_{dq} = 445$ MeV. The dashed line is the calculation Eq. (22) with $k = 922 \text{ MeV/fm}^2$ and $M_{dq} = 474 \text{ MeV}$, which is obtained without $\Lambda^*(1405)$.

 $\Lambda(1405)$ appears closer to the fitted line and the systematics for the baryon excitation energy gets better than that with $\Lambda(1405)$. This implies that the $\Lambda(1405)$ can be considered as an exotic hadron better than other baryons.

Now assuming that these baryons are composed of a *ud* diquark and a quark with a harmonic oscillator potential, we try to extract the string constant *k* and the *ud* diquark mass M_{dq} from the excited energies. For the harmonic oscillator system the excitation energy is described as

$$\Delta E = \sqrt{\frac{k}{\mu}}, \quad \text{with} \quad \mu = \frac{M_{dq}M_q}{M_{dq} + M_q}. \tag{22}$$

By fitting the excitation energies with Eq. (22), we find that $k = 890 \text{ MeV/fm}^2$ and $M_{dq} = 445 \text{ MeV}$. In Fig 14(b) we show the excitation energies of the baryons extracted by the Schottky anomaly in points as a function of the constituent quark mass M_q together with that obtained using the fitted values of k and M_{dq} in the solid line. We perform the same calculation without $\Lambda(1405)$, and we obtain $k = 992 \text{ MeV/fm}^2$ and $M_{dq} = 474 \text{ MeV}$. In Fig. 14(b) we also show the fitted excitation energy without $\Lambda(1405)$ as the dashed line.

D. Excitation energy scales for positive and negative parity hadrons

The N(1440) nucleon resonance is the first excited state of nucleon with $J^P = \frac{1}{2}^+$. This state is called Roper resonance [18] and the mass cannot be reproduced by simple quark models [19]. The property of the Roper resonance is one of the important subject in hadron physics. It was pointed out that similar lower excited states with positive parity are also found for hyperons [20], and recently their heavy baryon candidates have been proposed in Ref. [21]. It is very interesting that these Roper like excited states are located about 500 MeV higher than their ground state. Here we separately calculate the excitation energies for the positive and negative party baryons.

For the *N*, Δ , Λ , Σ , Λ_c , and Λ_b baryons, we calculate the excitation energy scale of the positive and negative parity states using their specific heats. Figure 15(a) shows the specific heats of the positive parity states. We read the excitation energies from the figure and we summarized the excitation energy scales in Table VII. The obtained excited energies are around 500 MeV less dependently on their flavor. This supports the proposal given in Ref. [21]. In order to see the excitation energy to negative parity states, we calculate the specific heat with the negative parity states together with the ground state. Figure 15(b)shows the specific heats and Table VII summarizes the excitation energies extracted from the specific heats. From the table we find strong flavor dependence of the excited energies to the negative parity states; the excitation energy gets smaller for the heavier baryons. This is consistent with what we have found in Sec. III C 3. Thus, we say that for baryons we have flavor dependence on the orbital excitation energies. For the negative parity Λ baryon, we also calculate the specific heat without $\Lambda(1405)$ and find the excitation energy scale to be 410 MeV, which is larger than that with $\Lambda(1405)$. As a result, the difference of the excitation energy scales between Λ and Σ gets smaller.

For mesons also, we calculate the excitation energy scale of the positive and negative parity states. In order to see the excitation energy to the positive parity states, we calculate

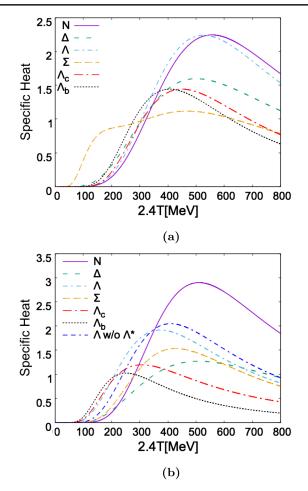


FIG. 15. (a) Specific heat of the positive parity baryons. The solid line, short dashed, short dashed-dotted, long dashed, long dashed-dotted, and dotted lines show specific heat of the nucleon, Δ , Λ , Σ , Λ_c , and Λ_b baryons, respectively. (b) Same as (a) but for negative parity baryons. In addition, the dashed-dashed-dotted line shows specific heat of Λ without $\Lambda^*(1405)$.

the specific heat with the positive parity states together with the ground state, which has a negative parity, for each flavor. Figure 16(a) shows the specific heats of the heavy mesons, while Fig. 16(b) plots those for the light mesons without the Nambu-Goldstone bosons, $f_0(500)$ and $K_0^*(700)$. Table VIII summarizes the excitation energies extracted from the specific heats. We find different excitation energy scales for heavy quarkonia and open heavy states. The former is around 440 MeV, while the latter

TABLE VII. Excitation energy scales of the positive and negative parity baryons obtained by the Schottky anomaly in units of MeV. The value in the bracket is excitation energy scale of Λ system without $\Lambda(1405)$.

	Ν	Δ	Λ	Σ	Λ_c	Λ_b
+ Parity	557	501	527	468	450	407
 Parity 	510	509	373 (410)	427	307	251

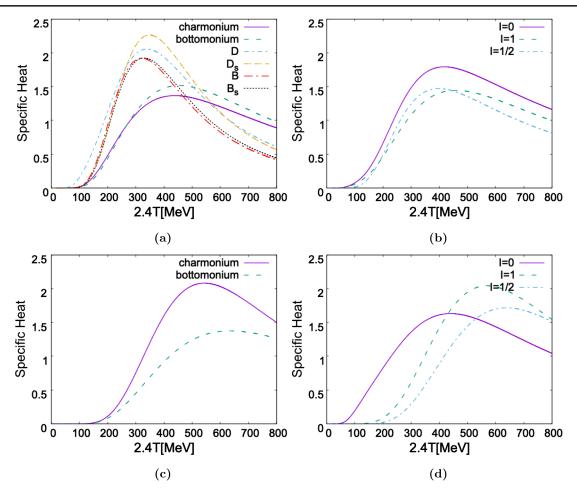


FIG. 16. (a) Specific heat of the positive parity heavy mesons. The solid line, short dashed, short dashed-dotted, long dashed, long dashed-dotted, and dotted lines show specific heat of the charmonium, bottomonium, D, D_s , B, and B_s mesons, respectively. (b)Specific heat of the positive parity light mesons. The solid line, dashed, dashed-dotted lines show specific heat of the I = 1, I = 0, and I = 1/2, mesons, respectively. (c) Specific heat of the negative parity heavy mesons. The solid line, dashed-dotted lines show specific heat of the show specific heat of the negative parity heavy mesons. The solid line, dashed-dotted lines show specific heat of the show specific heat of the negative parity heavy mesons. The solid line, dashed-dotted lines show specific heat of the show specific heat of the negative parity heavy mesons. The solid line, dashed-dotted lines show specific heat of the show specific heat of the negative parity heavy mesons. The solid line, dashed-dotted lines show specific heat of the negative parity heavy mesons. The solid line, dashed-dotted lines show specific heat of the negative parity heavy mesons. The solid line, dashed-dotted lines show specific heat of the charmonium mesons, respectively. (d) Same as (b), but for negative parity light mesons.

around 330 MeV. For the light mesons, the I = 1 states have rather a larger excitation energy (~460 MeV) than those for the I = 0 and I = 1/2 states (~410 MeV). These energies are close each other, but we do not find their systematics. If one could find the origin of the difference of these energy scales, it would be interesting. Figures 16(c) and 16(d) show the specific heats of the negative parity mesons. Unfortunately, for the open heavy states, we do not have enough states to investigate their excitation

TABLE VIII. Excitation energy scales of the positive and negative parity mesons obtained by the Schottky anomaly in units of MeV.

	Heavy mesons				Li	ight me	sons		
	сī	$b\bar{b}$	D	D_s	В	B_s	I = 0	I = 1	I = 1/2
+ parity	438	454	340	348	321	329	418	459	398
– parity	544	635	•••	•••	•••	•••	440	569	639

energies. The extracted energies are summarized in Table VIII. Although we find a large flavor dependence in the excitation energies, we could not find significant systematics in the flavor dependence.

IV. SUMMARY

The specific heats of the hadronic systems have been calculated using the hadron masses observed in experiment. We use the Schottky anomaly to extract the typical excitation energy scales of a hadronic system ΔE . The extraction is done by reading the temperature T_{peak} where the specific heat has a peak and by assuming $\Delta E \simeq 2.4T_{\text{peak}}$, which is a general relation in the two-state system.

We have obtained excitation energy scales for hadronic systems. Among the excitation energy scales, we have pinned down that of the hyperfine structure in the heavy meson systems and the excitation energy coming from the global structure of the confinement potential which is seen commonly in all of the hadron spectra. Thus, the latter scale may stem from orbital excitation. In the meson system, the energy scale of the orbital excitation is found to be 400-440 MeV insensitively to flavor dependence. On the other hand, in the baryon systems, we find flavor dependence of the energy scale of orbital excitation. The excitation energies for heavier systems are smaller than those for lighter systems. We have found by comparing the obtained excitation energies of N, Λ , Λ_c , and Λ_b baryons, which have similar isospin structures, that the excitation energy for these baryons can be described as $\Delta E \propto \mu^{-1/2}$, where μ is the reduced mass of the consistent quark and a ud diquark. This implies that these baryons may be described by a constituent quark and a *ud* diquark with a harmonic oscillator potential. We have also found that without the $\Lambda(1405)$ resonance this systematics becomes better. This gives a further support for $\Lambda(1405)$ to be an exotic hadron. We have also examined separately the excitation energies to the positive and negative parity excited states from the ground state for baryons. We have found that the positive parity excitation energy for the baryons are rather flavor independent as pointed out in Ref. [21], while the negative parity excitation energies have flavor dependence that the heavier baryons have smaller excitation energies than the lighter baryons. For the mesons, we have not found significant systematics for the separated excitation energies. As a future prospect, it would be interesting to confirm these conclusions if one could observe further excited states, such as open heavy mesons, Ξ , Ξ_c , and Ξ_b baryons, in which lower excited states are not sufficiently observed.

ACKNOWLEDGMENTS

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APPENDIX: TABLES OF STATES USED FOR CALCULATION OF SPECIFIC HEAT

In this section we list the hadron states used to calculate the specific heat. We use the value in the bracket of the name of the state for its mass.

TABLE IX. Charmonium states used in the calculation. In the left (right) line, the charmonia in which two quarks form spin 0 (1) are listed. (See the text for the details.)

< <i>i</i>		,	
η_c, h_c	J^{PC}	Ψ, χ_c	J^{PC}
$\eta_c(2984)$	0-+	$J/\Psi(3097)$	1
$h_c(3525)$	1^{+-}	$\chi_{c0}(3415)$	0^{++}
$\eta_c(3638)$	0^{-+}	$\chi_{c1}(3511)$	1++
,		$\chi_{c2}(3556)$	2^{++}
		Ψ(3686)	1
		Ψ(3770)	1
		$\Psi_{2}(3823)$	2
		$\Psi_{3}(3842)$	3
		$\chi_{c1}(3872)$	1++
		$\chi_{c2}(3930)$	2^{++}
		$\Psi(4040)$	1
		$\chi_{c1}(4140)$	1++
		$\Psi(4160)$	1
		Ψ(4230)	1
		$\chi_{c1}(4274)$	1++
		Ψ(4360)	1
		Ψ(4415)	1
		Ψ(4660)	1

TABLE X. Bottomonium states used in the calculation. In the left (right) line, the bottomonia in which two quarks form spin 0 (1) are listed. (See the text for the details.)

η_b, h_b	J^{PC}	Ψ, χ_b	J^{PC}
$\eta_b(9399)$	0-+	Υ(9460)	1
$h_b(9899)$	1+-	$\chi_{b0}(9859)$	0^{++}
- ()		$\chi_{b1}(9893)$	1^{++}
		$\chi_{b2}(9912)$	2^{++}
		Ύ(10023)	1
		$\Upsilon_{2}(10164)$	2
		$\chi_{b0}(10233)$	0^{++}
		$\chi_{b1}(10255)$	1^{++}
		$\chi_{b2}(10269)$	2^{++}
		Ύ(10355)	1
		$\chi_{b1}(10513)$	1^{++}
		$\chi_{b2}(10524)$	1^{++}
		Ύ(10579)	1
		$\Upsilon(10860)$	1
		Υ(11020)	1

TABLE XI. Open charm states used in the calculation.

D, D^*	J^P	D_s, D_s^*	J^P
D(1870)	0-	$D_{s}(1968)$	0-
D(2420)	1^{+}	$D_{s1}(2460)$	1^{+}
		$D_{s1}(2536)$	1^{+}
$D^{*}(2007)$	1-	$D_s^*(2112)$	1-
$D_0^*(2300)$	0^+	$D_{s0}^{*}(2317)$	0^+
$D_2^*(2460)$	2^{+}	$D_{s2}^{*}(2573)$	2^{+}
		$D_{s1}^{*}(2700)$	1-

TABLE XII. Open bottom states used in the calculation.

B, B^*	J^P	B_s, B_s^*	J^P
B(5279)	0-	$B_{s}(5367)$	0-
B(5721)	1^{+}	$B_{s1}(5830)$	1^{+}
$B^{*}(5325)$	1-	$B_{s}^{*}(5425)$	1-
$B_2^*(5747)$	2^{+}	$B_{s2}^{*}(5840)$	2^{+}

TABLE XIII. Light meson states used in the calculation.

I = 1		I = 0		$I = \frac{1}{2}$	
π, ρ, a, b	J^{PC}	$\overline{\eta,\eta',\omega,\phi,h,h',f,f'}$	J^{PC}	K, K^*	J^P
$\pi(135)$	0^{-+}	$f_0(500)$	0^{++}	K(498)	0-
$\rho(770)$	1	$\eta(547)$	0^{-+}	$K_0^*(700)$	0^+
$a_0(980)$	0^{++}	$\omega(782)$	1	$K^{*}(892)$	1-
$b_1(1235)$	1^{+-}	$\eta'(958)$	0^{-+}	$K_1(1270)$	1^{+}
$a_1(1260)$	1^{++}	$f_0(980)$	0^{++}	$K_1(1400)$	1^{+}
$\pi(1300)$	0^{-+}	$\phi(1020)$	1	$K^{*}(1410)$	0^+
$a_2(1320)$	2^{++}	$h_1(1170)$	1^{+-}	$K_0^*(1430)$	0^+
$\pi_1(1400)$	1^{-+}	$f_2(1270)$	2^{++}	$K_{2}^{*}(1430)$	2^{+}
$a_0(1450)$	0^{++}	$f_1(1285)$	1^{++}	$K^{*}(1680)$	1-
$\rho(1450)$	1	$\eta(1295)$	0^{-+}	$K_2(1770)$	2-
$\pi_1(1600)$	1^{-+}	$f_0(1370)$	0^{++}	$K_{3}^{*}(1780)$	3-
$a_1(1640)$	1^{++}	$\eta(1405)$	0^{-+}	$K_2(1820)$	2-
$\pi_2(1670)$	2^{-+}	$h_1(1415)$	1^{+-}	$\tilde{K_{4}^{*}(2045)}$	4^{+}
$\rho_{3}(1690)$	3	$f_1(1420)$	1^{++}		
$\rho(1700)$	3 1	$\omega(1420)$	1		
$a_2(1700)$	2^{++}	$\eta(1475)$	0^{-+}		
$\pi(1800)$	0^{-+}	$f_0(1500)$	0^{++}		
$\pi_2(1880)$	2^{-+}	$f'_2(1525)$	2^{++}		
$a_4(1970)$	4^{++}	$\eta_2(1645)$	2^{-+}		
1()		$\omega(1650)$	1		
		$\omega_3(1670)$	3		
		$\phi(1680)$	1		
		$f_0(1710)$	0^{++}		
		$\phi_3(1850)$	1		
		$\eta_2(1870)$	2^{-+}		
		$f_2(1950)$	2^{++}		
		$f_2(2010)$	2^{++}		
		$f_4(2050)$	4^{++}		
		$\phi(2170)$	1		
		$f_2(2300)$	2^{++}		
		$f_2(2340)$	2^{++}		

TABLE XIV. Light unflavored baryon states used in the calculation.

\mathbf{r}^{D}	
$N(I=1/2) J^P J^P$	$\Delta(I=3/2) \qquad \qquad J^P$
$N(940)$ $\frac{1}{2}^+$	$\Delta(1232) \frac{3}{2}$
$N(1440)$ $\frac{1}{2}+$	$\Delta(1600)$ $\frac{2}{3}$
$N(1520)$ $\frac{3}{2}$	$\Delta(1620)$ $\frac{1}{2}$
$N(1535)$ $\frac{1}{2}$	$\Delta(1700)$ $\frac{3}{2}$
$N(1650)$ $\frac{1}{2}$	$\Delta(1900)$ $\frac{1}{2}$
$N(1675)$ $\frac{5}{2}$	$\Delta(1905) = \frac{5}{2}$
$N(1680)$ $\frac{5}{2}+$	$\Delta(1910)$ $\frac{1}{2}^{+}$
$N(1700)$ $\frac{3}{2}$	$\Delta(1920)$ $\frac{3}{2}$
$N(1710)$ $\frac{1}{2}+$	$\Delta(1930)$ $\frac{5}{2}$
$N(1720)$ $\frac{3}{2}+$	$\begin{array}{ccc} \Delta(1232) & \frac{3}{2} + \\ \Delta(1600) & \frac{3}{2} + \\ \Delta(1620) & \frac{1}{2} - \\ \Delta(1700) & \frac{3}{2} - \\ \Delta(1900) & \frac{1}{2} - \\ \Delta(1905) & \frac{5}{2} + \\ \Delta(1910) & \frac{1}{2} + \\ \Delta(1920) & \frac{3}{2} + \\ \Delta(1930) & \frac{5}{2} - \\ \Delta(1950) & \frac{7}{2} + \\ \Delta(2200) & \frac{7}{2} - \\ \Delta(2420) & \frac{11}{2} + \\ \Delta(2420) & \frac{11}{2} + \\ \Delta(1920) & \frac{11}{2} - \\ \Delta$
$N(1875)$ $\frac{5}{2}$	$\Delta(2200)$ $\frac{7}{2}$
$N(1880)$ $\frac{1}{2}^+$	$\Delta(2420)$ $\frac{11}{2}$
$N(1895)$ $\frac{1}{2}$	2
$N(1900)$ $\frac{5}{2}+$	
$N(2060)$ $\frac{5}{2}$	
$N(2100)$ $\frac{1}{2}^+$	
$N(2120)$ $\frac{3}{2}$	
$N(2190)$ $\frac{7}{2}$	
$N(2220)$ $\frac{5}{2}+$	
$N(2250)$ $\frac{5}{2}$	
$\begin{array}{c ccccc} N(940) & \frac{1}{2}^+ \\ N(1440) & \frac{1}{2}^+ \\ N(1520) & \frac{3}{2}^- \\ N(1535) & \frac{1}{2}^- \\ N(1650) & \frac{1}{2}^- \\ N(1650) & \frac{5}{2}^- \\ N(1680) & \frac{5}{2}^+ \\ N(1680) & \frac{5}{2}^+ \\ N(1700) & \frac{3}{2}^- \\ N(1710) & \frac{1}{2}^+ \\ N(1720) & \frac{3}{2}^+ \\ N(1875) & \frac{3}{2}^- \\ N(1880) & \frac{1}{2}^+ \\ N(1895) & \frac{1}{2}^- \\ N(1895) & \frac{1}{2}^- \\ N(1895) & \frac{1}{2}^- \\ N(1900) & \frac{3}{2}^+ \\ N(2060) & \frac{3}{2}^- \\ N(2100) & \frac{1}{2}^+ \\ N(2120) & \frac{3}{2}^- \\ N(2120) & \frac{3}{2}^- \\ N(2120) & \frac{3}{2}^- \\ N(2120) & \frac{3}{2}^- \\ N(2220) & \frac{9}{2}^+ \\ N(2250) & \frac{9}{2}^- \\ N(2600) & \frac{11}{2}^- \\ \end{array}$	

TABLE XV. Light strange baryon states used in the calculation.

$\Lambda(I=0)$	J^P	$\Sigma(I=1)$	J^P	$\Xi(I=1/2), \Omega(I=0)$	J^P
Λ(1115)	$\frac{1}{2}^{+}$	$\Sigma(1190)$	$\frac{1}{2}^{+}$	Ξ (1315)	$\frac{1}{2}^{+}$
$\Lambda(1405)$	$\frac{\tilde{1}}{2}$	$\Sigma(1385)$	$\frac{1}{2} + \frac{1}{2} + \frac{1}{2} - \frac{1}$	Ξ(1530)	$\frac{\frac{1}{2}}{\frac{3}{2}+}$ $\frac{1}{2}-$ $\frac{5}{2}$ $\frac{3}{2}+$
$\Lambda(1520)$	$\frac{\overline{3}}{2}$	$\Sigma(1660)$	$\frac{\tilde{1}}{2}$ +	$\Xi(1820)$	$\frac{\tilde{1}}{2}$
$\Lambda(1600)$	$\frac{\tilde{1}}{2}$ +	$\Sigma(1670)$	$\frac{\overline{3}}{2}$	$\Xi(2030)$	$\frac{\tilde{5}}{2}$?
$\Lambda(1670)$	$\frac{\overline{1}}{2}$	$\Sigma(1750)$	$\frac{\overline{1}}{2}$	$\Omega(1672)$	$\frac{\tilde{3}}{2}$ +
$\Lambda(1690)$	$\frac{\overline{3}}{2}$	$\Sigma(1775)$	$\frac{5}{2}$		2
$\Lambda(1800)$	$\frac{\tilde{1}}{2}$	$\Sigma(1910)$	$\frac{\overline{3}}{2}$		
$\Lambda(1810)$	$\frac{\overline{1}}{2}^+$	$\Sigma(1915)$	$\frac{\bar{5}}{2}$ +		
$\Lambda(1820)$	$\frac{\overline{5}}{2}$ +	$\Sigma(2030)$	$\frac{\tilde{7}}{2}$ +		
$\Lambda(1830)$	$\frac{\overline{5}}{2}$		2		
$\Lambda(1890)$	$\frac{\tilde{3}}{2}$ +				
$\Lambda(2100)$	$\frac{\overline{7}}{2}$				
$\Lambda(2110)$	$\frac{5}{2}$ +				
Λ(2350)	$\frac{1}{2} \frac{1}{2} \frac{1}$				

$\overline{\Lambda_c, \Sigma_c}$	J^P	Ξ_c	J^P	Λ_b, Σ_b	J^P
$\Lambda_c(2286)$	$\frac{1}{2}$ +	$\Xi_c(2470)$	$\frac{1}{2}$ +	$\Lambda_b(5620)$	$\frac{1}{2}$ +
$\Lambda_c(2595)$	$\frac{1}{2}$	$\Xi_{c}(2790)$	$\frac{1}{2}$	$\Lambda_b(5912)$	$\frac{1}{2}$
$\Lambda_c(2625)$	$\frac{2}{3}$ -	$\Xi_{c}(2815)$	$\frac{\frac{2}{3}}{2}$	$\Lambda_b(5920)$	$\frac{\frac{2}{3}}{2}$
$\Lambda_c(2860)$	$\frac{2}{3}+$		2	$\Lambda_b(6146)$	$\frac{3}{2}$ +
$\Lambda_c(2880)$	$\frac{\frac{2}{5}}{2}$ +			$\Lambda_b(6152)$	$\frac{\frac{2}{5}}{2}$ +
$\Lambda_c(2940)$	$\frac{\frac{2}{3}}{2}$				2
$\Sigma_{c}(2455)$	$\frac{1}{2}$ +	$\Xi_{c}^{\prime}(2580)$	$\frac{1}{2}$ +	$\Sigma_{b}(5810)$	$\frac{1}{2}^{+}$
$\Sigma_c(2520)$	$\frac{\frac{2}{3}}{2}$ +	$\Xi_{c}^{\prime}(2645)$	$\frac{\frac{2}{3}}{\frac{2}{2}}$ +	$\Sigma_b(5830)$	$\frac{\frac{2}{3}}{\frac{2}{2}}$ +

TABLE XVI.	Heavy baryon	states used i	n the calculation.

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