Resummation of quantum radiation reaction and induced polarization

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In a previous paper we proposed a new method based on resummations for studying radiation reaction of an electron in a plane-wave electromagnetic field. In this paper we use this method to study the electron momentum expectation value for a circularly polarized monochromatic field with $a_0 = 1$, for which standard locally constant-field methods cannot be used. We also find that radiation reaction has a significant effect on the induced polarization, as compared to the results without radiation reaction, i.e., the Sokolov-Ternov formula for a constant field, or the zero result for a circularly monochromatic field. We also study the Abraham-Lorentz-Dirac equation using Borel-Padé resummations.

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I. INTRODUCTION

Radiation reaction (RR) [1,2], i.e., the difference between the actual trajectory of a charge and the one predicted by the Lorentz force equation (with only the background field, no self-field), is becoming an important effect in high-intensity laser experiments [3,4] (see also [5]). RR is an old topic which continues to be the subject of theoretical studies [6–13]. Even in classical electrodynamics there is still some debate on which equation is correct. Perhaps more relevant for upcoming laser experiments is how to obtain quantum RR with spin etc. and e.g., how to go beyond the standard locally constant-field (LCF) approximation. In [14] we developed a new method for calculating the momentum expectation value of an electron in a plane-wave background field. It is fully quantum, includes both real photon emissions and loops and can be used for arbitrary spin (see e.g., [15-17] for other studies of spin-dependent RR). This method is based on the use of first-order Mueller matrices as building blocks of higherorder probabilities [18,19]. In [14] we focused for simplicity on a constant field, but this method is not restricted to LCF or $a_0 \gg 1$. It works as long as the field is sufficiently long. In this paper we will consider both a constant field as well as a circularly polarized monochromatic field with $a_0 = 1$. The circular case provides a contrasting example compared to the constant field case, which we will now explain.

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If an initially unpolarized electron is in the background field for a sufficiently long time, then it can become polarized. If one neglects RR and considers a constant field, the polarization is given by the famous Sokolov-Ternov result. This induced polarization is along the magnetic field direction. The Sokolov-Ternov effect was originally considered for electrons in storage-rings. However, one has a similar effect in other fields as well [20–22], in particular plane waves. However, if one just lets an electron stay in the laser field then RR will eventually lead to a considerable change in the momentum. This is not included in the Sokolov-Ternov result. The method in [14] is not restricted to only the momentum expectation value; here we will use these methods to study how RR affects the induced polarization.

For a linearly polarized field and in the LCF regime, it has been shown in [20,23] that the induced polarization tends to average out due to the fact that the magnetic field changes direction every half cycle. There are ways [20,23] to compensate for this by choosing some asymmetric field. One might wonder, though, if a circularly polarized field could be used instead to generate polarized particles. If one only considers the contribution from real photon emission then one would conclude that a circularly polarized field would be able to make the particles polarized along the propagation direction of the laser. However, at first order in α , this is exactly canceled by the cross term between the zeroth order amplitude and the first-order loop [19,24,25] (see though [26] for dependence on collision angle). Hence, if one neglects RR then there is no induced polarization for a circularly polarized field. Here we will use the methods of [14] to show that there is in fact a nonzero induced polarization if one takes RR into account.

Polarization effects of particles in a circularly polarized field in LCF have been studied in [27,28]. There is a crucial difference here. For a circular field in LCF, one needs to

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start with polarized particles in the initial state to have longitudinally polarized particles in the final state. Here we are instead interested in a circular field with $a_0 \sim 1$ in which, as we will show, even an initially unpolarized particle may become longitudinally polarized. In this regime, spin up and down the longitudinal direction plays essentially the same role as spin up and down the magnetic field direction for a linearly polarized field.

This paper is organized as follows. In Sec. III we study how RR affects the Sokolov-Ternov polarization for a constant field. In Sec. IV we show the details of how to use the general methods in [14] for a circularly polarized monochromatic field with $a_0=1$ (i.e., beyond LCF), and we apply these methods to both the momentum expectation value and the induced polarization. In order to gain some insights into how the approximation in [14] relates to the full QED result (which cannot be calculated), in Sec. V we compare its leading classical limit, which we in [14] showed is equal to solution of the Landau-Lifshitz (LL) equation [29,30], with a resummation of the Abraham-Lorentz-Dirac (LAD) equation.

II. NOTATION AND BASIC INGREDIENTS

We use units with the electron mass $m_e=1$ (and $c=\hbar=1$). A factor of the electron charge is absorbed into the laser field, $eF_{\mu\nu}\to F_{\mu\nu}$, so e only appear explicitly in $\alpha=e^2/(4\pi)$. We use lightfront coordinates $v^\pm=2v_\mp=v^0\pm v^3$ and $v_\perp=\{v_1,v_2\}$. The vector potential of the laser has $a_\pm=0$ and $a_\perp(\phi)$ is a function of lightfront time $\phi=kx=k_+x^+=\omega(t+z)$. ω is a characteristic laser frequency, E is the field strength, and $a_0:=E/\omega$. p_μ is the initial electron momentum, $b_0:=kp$ and $\chi:=a_0b_0$.

In [14] we studied the expectation value of the lightfront longitudinal momentum $\langle kP \rangle$, where P_{μ} is the final electron momentum. The starting point is the α expansion $\langle kP \rangle = \sum_{n=0}^{\infty} \langle kP \rangle^{(n)}, \langle kP \rangle^{(n)} = \mathcal{O}(\alpha^n)$. The spin dependence can be expressed in terms of Stokes vectors N and strong-field-QED Mueller matrices M, $\langle kP \rangle^{(n)} =$ $(1/2)\mathbf{N}_0 \cdot \mathbf{M}^{(n)} \cdot \mathbf{N}_f$, where \mathbf{N}_0 and \mathbf{N}_f are the initial and final Stokes vector of the electron and the Mueller matrix $\mathbf{M}^{(n)}$ is given by a sum of all the relevant $\mathcal{O}(\alpha^n)$ diagrams. For large a_0 and/or a long pulse, the leading-order terms can be expressed as products of $\mathcal{O}(\alpha)$ Mueller matrices [18,19]. These are obtained from the probabilities $\mathbb{P}_{e^-\to e^-+\gamma}=(1/2)\mathbf{N}_0\cdot\mathbf{M}^{\mathrm{C}}\cdot\mathbf{N}_f+\mathcal{O}(\alpha^2)\quad\text{and}\quad\mathbb{P}_{e^-\to e^-}=$ $(1/2)\mathbf{N}_0\cdot(\mathbf{1}+\mathbf{M}^{\mathrm{L}})\cdot\mathbf{N}_f+\mathcal{O}(\alpha^2)$. $\mathbf{M}^{(n)}$ is given by a sum of 2^n diagrams with $0 \le m \le n$ Compton scattering steps \mathbf{M}^{C} and n-m loop steps \mathbf{M}^{L} . For example, at $\mathcal{O}(\alpha^2)$ we have $\mathbf{M}^{C} \cdot \mathbf{M}^{C}$, $\mathbf{M}^{C} \cdot \mathbf{M}^{L}$, $\mathbf{M}^{L} \cdot \mathbf{M}^{C}$ and $\mathbf{M}^{L} \cdot \mathbf{M}^{L}$, with corrections that are small for large a_0 or a long pulse.

These matrix products are lightfront time ordered: The most general expressions for $\mathbf{M}^{\mathbf{C}}$ and $\mathbf{M}^{\mathbf{L}}$ involve two lightfront-time integration variables ϕ and ϕ' . We change

variables to $\sigma = (1/2)(\phi + \phi')$ and $\theta = \phi - \phi'$ and then to leading order we have a time ordering of the σ variables while the θ integrals can be performed for each $\mathbf{M}^{\mathrm{C,L}}$ separately [31].

The integrals over the transverse momentum components also factorize [32], so only the longitudinal momentum plays a nontrivial role here. The longitudinal momentum changes at Compton steps, so all the subsequent steps depend on how much of the electron longitudinal momentum is carried away by the emitted photon. The loop steps have independent longitudinal momentum integrals, but it is anyway convenient to pull out this integral in order to combine \mathbf{M}^{C} and \mathbf{M}^{L} .

The Stokes vectors are 4 dimensional, $\mathbf{N} = \{1, \mathbf{n}\}$, where \mathbf{n} is a 3D vector pointing in the average spin direction. Consequently, the Mueller matrices are in general 4×4 . However, in some cases some of the elements might be irrelevant. For the two cases we will consider here, a constant field and spin along the magnetic field direction or a circularly polarized monochromatic field and spin along the laser propagation direction, we only have two relevant elements.

Although this approach works for general nonconstant fields, here and in [14] we focus on either a constant field or a monochromatic field. In [14] we showed that for a constant field we have

$$\mathbf{M}^{(n)} = \int_0^1 \frac{\mathrm{d}q}{n\chi} (\mathbf{M}^{\mathrm{C}} \cdot \mathbf{M}^{(n-1)}(\chi(1-q)) + \mathbf{M}^{\mathrm{L}} \cdot \mathbf{M}^{(n-1)}(\chi)), \tag{1}$$

where the $\mathcal{O}(\alpha)$ building blocks are expressed in terms of Airy functions and $\operatorname{Ai}_1(\xi) = \int_{\xi}^{\infty} \mathrm{d}x \operatorname{Ai}(x)$, and if we restrict to an initial electron that is either unpolarized or polarized up or down the magnetic field direction then the 4×4 Mueller matrices reduce to the following 2×2 matrices

$$\mathbf{M}^{C} = \begin{pmatrix} -\mathrm{Ai}_{1}(\xi) - \kappa \frac{\mathrm{Ai}'(\xi)}{\xi} & \frac{q}{s_{1}} \frac{\mathrm{Ai}(\xi)}{\sqrt{\xi}} \\ q \frac{\mathrm{Ai}(\xi)}{\sqrt{\xi}} & -\mathrm{Ai}_{1}(\xi) - 2 \frac{\mathrm{Ai}'(\xi)}{\xi} \end{pmatrix}, \quad (2)$$

$$\mathbf{M}^{L} = \begin{pmatrix} \operatorname{Ai}_{1}(\xi) + \kappa \frac{\operatorname{Ai}'(\xi)}{\xi} & -q \frac{\operatorname{Ai}(\xi)}{\sqrt{\xi}} \\ -q \frac{\operatorname{Ai}(\xi)}{\sqrt{\xi}} & \operatorname{Ai}_{1}(\xi) + \kappa \frac{\operatorname{Ai}'(\xi)}{\xi} \end{pmatrix}, \quad (3)$$

 $\xi=(r/\chi)^{2/3}, r=(1/s_1)-1, \kappa=(1/s_1)+s_1$ and $s_1=1-q$. In [14] we considered the expectation value of the longitudinal momentum $\langle kP \rangle$ and then the first term is given by $\mathbf{M}^{(0)}=b_0\mathbf{1}$.

We can use (1) to first calculate the orders separately and then resum the α expansion afterwards. Each $\mathbf{M}^{L,C}$ comes with a factor of $a_0\Delta\phi$, where $\Delta\phi$ is the length of the field, so at $\mathcal{O}(\alpha^n)$ we have an overall factor of $(\alpha a_0\Delta\phi)^n$ times functions of χ . It is therefore natural to absorb $a_0 \Delta \phi$ into the following effective expansion parameter

constant field:
$$T = \alpha a_0 \Delta \phi$$
. (4)

Even though $\alpha \ll 1$, T can be large for a sufficiently strong field (large a_0) or long pulse. Higher orders can therefore be important. Fortunately, as we showed in [14] and as we will see further examples of here, we can resum these series with Padé approximants using only quite few terms. The α expansions of various strong-field-QED phenomena have recently been studied in [13,33–36].

Or we can resum (1) right from the start, which gives $\mathbf{M} = \sum_{n=0}^{\infty} T^n \mathbf{M}^{(n)}$ as the solution to an integro-differential equation

$$\frac{\partial \mathbf{M}}{\partial T} = \int_0^1 \frac{\mathrm{d}q}{\chi} (\mathbf{M}^{\mathrm{C}} \cdot \mathbf{M}(\chi[1-q]) + \mathbf{M}^{\mathrm{L}} \cdot \mathbf{M}(\chi)), \quad (5)$$

where $M(T = 0) = M^{(0)}$.

III. RR AND THE SOKOLOV-TERNOV EFFECT

In [14] we focused on $\langle kP \rangle$, but this formalism is more general. In this section we will consider the probability of transition between Stokes vector \mathbf{N}_i and \mathbf{N}_f , for which we have $\mathbb{P} = (1/2)\mathbf{N}_i \cdot \mathbf{M} \cdot \mathbf{N}_f$, where \mathbf{M} and $\mathbf{M}^{(n)}$ is obtained from the same (1) but with $\mathbf{M}^{(0)} = \mathbf{1}$ instead of $\mathbf{M}^{(0)} = b_0 \mathbf{1}$.

A. Standard Sokolov-Ternov, no RR

If we neglect RR then we should obtain the standard Sokolov-Ternov result. In our approach, we turn off RR by replacing $\chi(1-q) \to \chi$ in the first term in (1). The first-order Mueller matrix is given by

$$\mathbf{M}_1 = \begin{pmatrix} 0 & \chi^2 \\ 0 & -\frac{5\sqrt{3}}{8}\chi^2 \end{pmatrix} \tag{6}$$

and with no RR it is straightforward to sum up all orders, which gives

$$\mathbb{P} = \frac{1}{2} \mathbf{N}_0 \cdot \begin{pmatrix} 1 & \frac{8}{5\sqrt{3}} (1 - \exp\left[-\frac{5\sqrt{3}}{8} \chi^2 T\right]) \\ 0 & \exp\left[-\frac{5\sqrt{3}}{8} \chi^2 T\right] \end{pmatrix} \cdot \mathbf{N}_f, \quad (7)$$

which agrees with the results in [37,38]. For a very long pulse, $T \gg 1/\chi^2$ and with an initially unpolarized state,

$$\mathbb{P} = \frac{1}{2} \left\{ 1, \frac{8}{5\sqrt{3}} \right\} \cdot \mathbf{N}_f, \tag{8}$$

so the degree of the induced polarization is $8/(5\sqrt{3}) \approx 0.92$. Note that this is just a constant, in particular, independent of

the momentum χ . However, the time scale for this is $T \sim 1/\chi^2$, which is slower than the time scale for RR, which causes the longitudinal momentum to decrease as $\chi(T) = \chi/(1 + [2/3]\chi T)$. Since for large T the combination $\chi^2(T)T \sim 1/T$ is actually small rather than large, this suggests that the result with RR included will be very different from (7). Note that, in experiments where one can use the standard Sokolov-Ternov result, the reason why one can neglect RR is because in such experiments the particles are re-accelerated in order to compensate for RR. But in the following we are instead interested in electrons that are sent into a laser field without anything to counteract RR, and we will now show that the result will be very different (even for $\chi \ll 1$).

B. RR included, low energy

To see the effects of RR, we consider first the leading order in $\chi \ll 1$. Inserting the ansatz

$$\mathbf{M}_n = \chi^{n+1} \begin{pmatrix} 0 & B_n \\ 0 & A_n \end{pmatrix} \tag{9}$$

for $n \ge 1$ into (1), we find $B_n = -(2/3)B_{n-1}$ and $A_n = -(2/3)A_{n-1}$. Thus we have a simple geometric series and we find

$$\mathbb{P} = \frac{1}{2} \mathbf{N}_0 \cdot \left[\mathbf{1} + \frac{T\chi^2}{1 + \frac{2}{3}T\chi} \begin{pmatrix} 0 & 1\\ 0 & -\frac{5\sqrt{3}}{8} \end{pmatrix} \right] \cdot \mathbf{N}_f. \tag{10}$$

Note that this varies on the same time scale as the momentum expectation value, i.e., $T \sim 1/\chi$. For an asymptotically long pulse, $T \gg 1/\chi$, and for an initially unpolarized electron we have

$$\mathbb{P} = \frac{1}{2} \left\{ 1, \frac{3\chi}{2} \right\} \cdot \mathbf{N}_f. \tag{11}$$

This is quite different from the standard Sokolov-Ternov result (8): The asymptotic result is reached much faster $(T \gg 1/\chi)$, but is a much lower degree of polarization for $\chi \ll 1$. However, (11) suggests that we might be able to compensate for the negative effect of RR by increasing χ . Since the degree of polarization cannot be larger than 1, it is clear that the leading order (11) cannot be used for larger χ . So, we now turn to a numerical treatment.

C. RR included, higher energy

The orders \mathbf{M}_n can be calculated separately either by resummation of the χ expansion or with numerical interpolation functions exactly in the same way as in [14]. The only difference is that in [14] we considered the expectation value of the longitudinal momentum, with $\mathbf{M}_0 = b_0 \mathbf{1}$, while now we have $\mathbf{N}_f = \{0,1\}$. However, in [14] we used the "standard" Borel transform to resum the χ expansions, while here we are going to use a slight modification.

We begin by noting that the first order can be expressed as a Mellin integral

$$\mathbf{M}_{1} = \int \frac{\mathrm{d}S}{2\pi i} \frac{\pi}{4} \frac{3^{S/2}}{\chi^{S}} \frac{S(1+S)}{\sin[\pi S]} \begin{pmatrix} 0 & \frac{\Gamma[1-\frac{3S}{2}]}{\Gamma[1-\frac{S}{2}]} \\ 0 & -\frac{\sqrt{3}\Gamma[\frac{1}{2}-\frac{3S}{2}]}{\Gamma[\frac{1}{2}-\frac{S}{2}]} \end{pmatrix}, \quad (12)$$

where the integration contour goes through the real axis at -2 < ReS < 1/3. The $\chi < 1$ expansion is obtained directly from the residues of the poles at S = -n, n = 2, 3, ... For large n we have for the coefficient of χ^m

$$\mathbf{M}_{1}^{(m)} \sim \frac{1}{4} \sqrt{\frac{3}{2\pi}} \left(-\frac{3}{2} \right)^{m} \Gamma \left[m + \frac{5}{2} \right] \begin{pmatrix} 0 & 1 + \mathcal{O}(1/m) \\ 0 & -1 + \mathcal{O}(1/m) \end{pmatrix}.$$
(13)

Of course, for this order we already have an exact expression, e.g., (12), but (13) will help us to guess a convenient resummation for \mathbf{M}_n , n > 1. We write \mathbf{M}_n as in (9) but where A_n and B_n are now functions of χ . We focus on B_n . Its expansion

$$B_n = \sum_{m=0}^{\infty} B_n^{(m)} \chi^m \tag{14}$$

starts with the constants $B_n^{(0)}$ obtained in Sec. III B. In contrast to the n = 1 case, for n > 1 we only have the first ~ 100 terms in χ . From (13) we take

$$(-c)^m \Gamma[m+b_n], \tag{15}$$

as an ansatz for the large-order scaling of $B_n^{(m)}$.

In many applications one finds Borel transforms with poles at t > 0, and then there is an ambiguity in the Laplace integral because one can avoid the pole by integrating either above or below it. This may actually be an interesting property because one can then study how this ambiguity is resolved due to resurgence, which relates different sectors of trans-series (see [39] for a recent review). However, all the asymptotic series in this paper have alternating signs at large orders and we have found no poles for the resummed Borel transforms for t > 0.

In this case, we take c = 3/2 and determine b numerically. We find $b_n = 3 + (3/2)n$ for $n \ge 1$. Note that the value of b_n does not need to be perfect (in fact, $b_n = 3 + (3/2)n$ might be only an approximation), but we see that at larger n b_n is considerably larger than $b_1 = 9/2$, which one might expect to be significant. Thus, instead of the standard Borel transform, we use the Borel-Leroy transform (the factor of c^m is just a variable rescaling)

$$BL[B_n](t) = \sum_{m=0}^{\infty} \frac{B_n^{(m)}}{c^m \Gamma[m+b_n]} t^m.$$
 (16)

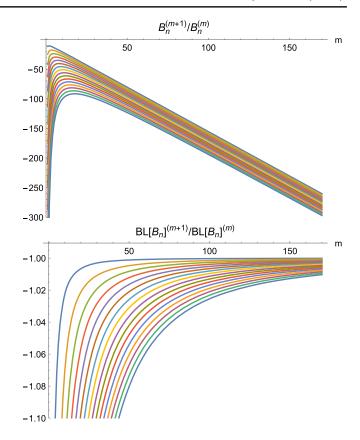


FIG. 1. Ratios of neighboring coefficients for the expansion in (14) and its Borel-Leroy transform (16).

Figure 1 shows that this transforms the asymptotic series in (14) into a convergent one. This can now be resummed with a Padé approximant [40–49]. We have used $[40/(40 + \text{ceil}[b_n] + 1)]$. In contrast to the Padé resummation of the α expansion, where we know or have a good idea of what i and j in [i/j] should be for the fastest convergence, in this case we just chose [40/(40 + $ceil[b_n] + 1)$ because it turned out that having j considerably larger than i gives faster convergence (especially at higher orders in α). That we have to change i - j as nincreases might be understood as a consequence of fact that (9) has an overall factor of χ^{n+1} and so, although we have not calculated the $\chi \gg 1$ scaling of \mathbf{M}_n yet, it seems clear that the $\chi \gg 1$ limit of \mathbf{M}_n is very different from χ^{n+1} , which means that the resummed B_n has to go faster and faster to zero as γ increases. Thus, there is still considerable room for finding more optimized resummations that would give the same results but with much fewer terms.

¹In some cases one may have to choose slightly different Padé orders due to the appearance of spurious poles. That these poles are indeed spurious and not genuine is confirmed by the fact that they disappear by choosing a different order. In contrast, a genuine pole, such as the one closest to the origin at t < 0 which determines the radius of convergence, can be seen even if one chooses different orders. For more about spurious poles see [46,50].

One option would be the new resummation method in [51], which would allow us to incorporate the large- χ scaling in the resummation. The final result is obtained from a Laplace integral

$$B_n^{\text{(resum)}}(\chi) = \int_0^\infty \frac{\mathrm{d}t}{x} \left(\frac{t}{x}\right)^{b_n - 1} e^{-t/x} PBL[B_n](t), \quad (17)$$

where $x = c\chi$. Now we have the first ~10 to 20 terms in the α expansion.

The next step is to resum the α expansion. For this we can use Padé approximants $\lfloor n/m \rfloor$ directly. For n > m the resummation would diverge as $T \to \infty$, which is unphysical. For n < m the resummation would go to zero as $T \to \infty$, but we already know that in the classical limit the resummation should be nonzero at $T \to \infty$, and in the quantum regime we can confirm by solving the integrodifferential equation that the result should not go to zero. Thus, to have a nonzero (and finite) limit as $T \to \infty$ we resum this expansion with diagonal Padé approximants, [n/n]. The result is shown in Fig. 2. As in [14], we again find a fast convergence of the Padé approximants. For this particular value of χ , we are close to the asymptotic limit at T = 100 and for T < 100 we see that the resummation has already converged to a good precision with Padé[4/4], which is obtained with the first eight terms in the α expansion. The leading order (11) gives a rather large overestimation at $\gamma = 0.3$, but Fig. 2 shows that the exact result still gives a significant degree of polarization.

IV. CIRCULARLY POLARIZED FIELD

In the previous section and in [14] we considered a constant field. Now we will consider a circularly polarized monochromatic field. As a constant field can be seen as a first step toward LCF, a monochromatic field can be

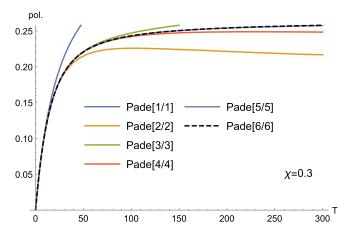


FIG. 2. The polarization induced on an initially unpolarized electron. Pade [n/n] refers to diagonal Padé approximants $(\sum_{i=1}^n c_i T^i)/(1+\sum_{i=1}^n d_i T^i)$. The leading order (11) leads to a large overestimation for this value of χ and is therefore not shown.

seen as the first step toward a locally monochromatic-field (LMF) approximation [19,52,53] (see also [54–57]). While approaches based on LCF only work for large a_0 , LMF works also for smaller a_0 provided the pulse length is sufficiently long. In this section we will therefore focus on $a_0 \sim 1$. The effective expansion parameter is

circular field:
$$T = \alpha \Delta \phi$$
, (18)

for $\mathbf{a}(\phi) = a_0 \{ \sin \phi, \cos \phi, 0 \}$ and $0 < \phi < \Delta \phi$.

A. Definitions

In general the Mueller matrices are 4×4 . For a constant field one can reduce this to 2×2 if one considers initial and final electrons that are either unpolarized or polarized parallel or antiparallel to the magnetic field. For a circularly polarized field (with $a_0 \sim 1$) the general 4×4 matrices can be found in [19]. However, also in this case one can reduce the Mueller matrices to 2×2 , but in this case the special spin direction is parallel/antiparallel to the wave vector of the field. The recursive and integro-differential equations for the circular field are obtained from (1) and (5) by replacing $\chi \to b_0$ and, for initial and final states that are unpolarized or polarized along the wave vector of the field, the first-order Mueller-matrix building blocks are given by [19]

$$\mathbf{M}^{\mathrm{C}} = \begin{pmatrix} \frac{\kappa}{2} \mathcal{J}_{1} - \mathcal{J}_{0} & \frac{q}{2} (1 + \frac{1}{s_{1}}) \mathcal{J}_{2} \\ \frac{q}{2} (1 + \frac{1}{s_{1}}) \mathcal{J}_{2} & -\frac{q^{2}}{2s_{1}} \mathcal{J}_{0} + \frac{\kappa}{2} (\mathcal{J}_{1} - \mathcal{J}_{0}) \end{pmatrix}$$
(19)

and

$$\mathbf{M}^{L} = \begin{pmatrix} -\frac{\kappa}{2}\mathcal{J}_{1} + \mathcal{J}_{0} & -\frac{q}{2}(1 + \frac{1}{s_{1}})\mathcal{J}_{2} \\ -\frac{q}{2}(1 + \frac{1}{s_{1}})\mathcal{J}_{2} & -\frac{\kappa}{2}\mathcal{J}_{1} + \mathcal{J}_{0} \end{pmatrix}$$
(20)

where $s_1 = 1 - q$, $\kappa = (1/s_1) + s_1$, $r = (1/s_1) - 1$,

$$\mathcal{J}_0 = \frac{i}{2\pi} \int \frac{\mathrm{d}\theta}{\theta} \exp\left\{\frac{ir}{2b_0}\Theta\right\} = \sum_{r} J_n^2(z), \quad (21)$$

$$\mathcal{J}_{1} = -2a_{0}^{2}(u)\frac{i}{2\pi} \int \frac{d\theta}{\theta} \sin^{2}\left(\frac{\theta}{2}\right) \exp\left\{\frac{ir}{2b_{0}}\Theta\right\}$$
$$= \frac{a_{0}^{2}(u)}{2} \sum_{n} [J_{n+1}^{2}(z) + J_{n-1}^{2}(z) - 2J_{n}^{2}(z)] \tag{22}$$

and

$$\mathcal{J}_{2} = -\frac{a_{0}^{2}(u)}{2\pi} \int \frac{d\theta}{\theta} \left(\sin \theta - \frac{4}{\theta} \sin^{2} \frac{\theta}{2} \right) \exp \left\{ \frac{ir}{2b_{0}} \Theta \right\}
= \frac{1}{2} \sum_{n} \left(\frac{b_{0}z^{2}}{nr} - a_{0}^{2}(u) \right) [J_{n-1}^{2}(z) - J_{n+1}^{2}(z)],$$
(23)

where the first version of these functions involve the effective mass M [58], which in LMF is given by

$$\Theta = \theta M^2 = \theta \left[1 + a_0^2(u) \left(1 - \operatorname{sinc}^2 \frac{\theta}{2} \right) \right], \quad (24)$$

 $u = \sigma/\mathcal{T}$, $a_0(u) = a_0h(u)$ for a field given by $\mathbf{a}(\phi) = a_0(\phi/\mathcal{T})\{\sin\phi,\cos\phi,0\}$ with pulse length \mathcal{T} and pulse envelope h, $\mathrm{sinc}(x) = \sin(x)/x$, and the singularity at $\theta = 0$ is avoided using a contour equivalent to $\theta \to \theta + i\epsilon$, and the second version is expressed in terms of Bessel functions J_n with argument

$$z = a_0(u) \frac{r}{b_0} \left[\frac{2b_0 n}{r} - 1 - a_0^2(u) \right]^{\frac{1}{2}}, \tag{25}$$

and the sum over n is restricted by

$$n \ge \frac{r}{2b_0} [1 + a_0^2(u)]. \tag{26}$$

While this formalism can be used for e.g., a Gaussian pulse envelope $h(u) = e^{-u^2}$, in this paper we will for simplicity consider a monochromatic field, i.e., $a_0(u) = a_0\theta(u)\theta(1-u)$ and we use $\Delta\phi$ instead of $\mathcal T$ for fields with constant envelope. We change integration variable from q to $\gamma = r/b_0$.

B. Integrals for the b_0 expansion

To obtain the b_0 expansions we need the following integrals

$$\mathcal{I}_{i}^{(n)} = \int_{0}^{\infty} d\gamma \gamma^{n} \mathcal{J}_{i}(\gamma). \tag{27}$$

We have calculated these using the θ -integral rather than the Bessel-function representations. We have

$$\mathcal{I}_0^{(n)} = \int d\theta \frac{i2^n n!}{\pi \theta (-i\Theta)^{1+n}}, \qquad (28)$$

and similarly for $\mathcal{I}_1^{(n)}$ and $\mathcal{I}_2^{(n)}$. With a partial integration one can show that $\mathcal{I}_2^{(n)} = -n\mathcal{I}_0^{(n-1)}$, so we focus on $\mathcal{I}_0^{(n)}$ and $\mathcal{I}_1^{(n)}$. The factorial growth of these integrals is responsible for the factorial growth of the b_0 expansion.

We can obtain a large-order expansion by exponentiating

$$\frac{1}{(-i\Theta)^n} = \exp\left(-n\ln[-i\Theta]\right) \tag{29}$$

and then performing the θ integral with the saddlepoint method. The saddle point, θ_s , is determined by a transcendental equation which can easily be solved numerically. Then one can change integration variable from $\theta = \theta_s + \delta\theta/\sqrt{n}$ to $\delta\theta$, expand the integrand in 1/n and perform the resulting Gaussian integrals (analytically). We find

$$\mathcal{I}_0^{(n)} \approx (1.588...)^n \Gamma \left[n + \frac{1}{2} \right] \times \left(0.34... - \frac{0.21...}{n} + \frac{0.22...}{n^2} - \cdots \right)$$
 (30)

and

$$\mathcal{I}_{1}^{(n)} \approx (1.588...)^{n} \Gamma \left[n + \frac{1}{2} \right] \times \left(0.74... + \frac{0.017...}{n} - \frac{0.019...}{n^{2}} - \cdots \right)$$
 (31)

where the coefficients have been obtained to high precision and have not been rounded for the digits shown above. A direct sum of the first 11 terms, i.e., including the term with $1/n^{10}$, has a precision of $\sim 10^{-11}$ for n=10 and $\sim 10^{-22}$ for n=100.

We still need to obtain $\mathcal{I}_i^{(n)}$ for finite n. We have found that in the process of calculating a large number of terms in the b_0 expansion and several terms in the α expansion, and then resumming these expansions, the precision can be reduced by several orders of magnitude. However, this is not a problem, because we can obtain a very high precision of $\mathcal{I}_i^{(n)}$. For odd n the integrals are simply given by $-i\pi$ times the residue at $\theta=0$, which gives rational numbers. So, for odd n we have exact expressions. For even n we could in principle close the contour in the upper half plane and use the residue theorem for the poles there. However, there are an infinite number of poles, and, although each pole can be obtained quickly and to a high precision using

$$2\pi\left(n+\frac{1}{2}\right)+i\ln\left[2(2\pi)^2\left(n+\frac{1}{2}\right)^2\right]$$
 (32)

as starting point for a numerical root finding, the sum over these poles converges slowly.

It is actually faster to just perform the integrals directly by integrating along the real axis except for a detour above $\theta=0$ at a distance $|\theta|\sim 1$. With this approach we have obtained the integrals to a precision of $\sim 10^{-40}$ for all the relevant n. However, the first couple of n takes a rather long time to compute. For larger n, on the other hand, the integrals converge very quickly, so for $n\gtrsim 20$ (the largest n, e.g., $n_{\rm max}=130$, needs to be about the same as the maximum order in b_0) we can quickly obtain a precision of e.g., $\sim 10^{-100}$.

We can also obtain such a high precision for lower n with the following approach, where we use a third expansion. This is essentially a perturbative expansion in the amplitude

of the field, a_0 , except that, as seen from (24), it is more natural to expand in $\delta \coloneqq a_0^2/(1+a_0^2)$ instead. Then the expansion of $1/(1-\delta \mathrm{sinc}^2[\theta/2])^{1+n}$ is a binomial series. Each term in the integrand now has the form $(1/\theta^j)\mathrm{sinc}^{2m}[\theta/2]$. We use the binomial theorem to write this as a sum of terms with $e^{i(m-k)\theta}$, where $0 \le k \le 2m$. Now we can use the residue theorem to calculate each term. Only those terms with $m < k \le 2m$ contribute. We find

$$\mathcal{I}_{0}^{(n)} = (1 - \delta)^{1+n} \sum_{m=1}^{\infty} \delta^{m} \sum_{k=m+1}^{2m} \times \frac{(-1)^{k} 2^{1+n} (k-m)^{2m+n+1} (2m)! (m+n)!}{m! k! (2m+n+1)! (2m-k)!}$$
(33)

and

$$\mathcal{I}_{1}^{(n)} = (1 - \delta)^{n} \sum_{m=1}^{\infty} \delta^{m} \sum_{k=m+1}^{2m} \times \frac{(-1)^{k} 2^{n} (k-m)^{2m+n-1} (2m)! (m+n-1)!}{(m-1)! k! (2m+n-1)! (2m-k)!}.$$
(34)

The sums over m have a finite radius of convergence. To see this, we go back a couple of steps and write $\operatorname{sinc}^{2m}[\theta/2] = \exp\{m \ln{(\operatorname{sinc}^m[\theta/2])}\}$, and then the large m behavior can be obtained by rescaling $\theta \to \theta/\sqrt{m}$ and expanding the integrand in 1/m. For even n (recall that odd n are anyway simple) we find

$$\mathcal{I}_0^{(n,m\gg 1)} \approx \frac{(1-\delta)^{1+n} m^{(n+1)/2} (m+n)!}{2 \times 3^{(n+1)/2} m! \Gamma[\frac{3+n}{2}]}$$
(35)

and

$$\mathcal{I}_{1}^{(n,m\gg1)} \approx \frac{(1-\delta)^{n} m^{(n-1)/2} (m+n-1)!}{3^{(n-1)/2} (m-1)! \Gamma[\frac{1+n}{2}]}$$
(36)

or to leading order $\mathcal{I}_0^{(n,m\gg 1)} \sim \mathrm{const.} \times m^{(3n+1)/2}$ and $\mathcal{I}_1^{(n,m\gg 1)} \sim \mathrm{const.} \times m^{(3n-1)/2}$. From $\mathcal{I}_i^{(n,m+1)}/\mathcal{I}_i^{(n,m)} \to 1$ and the ratio test we find that the radius of convergence is $\delta=1$, which is fortunate since we always have $\delta=a_0^2/(1+a_0^2)<1$. Hence, apart from the partial resummation achieved by using δ rather than a_0 , no additional resummation is needed and (33) can therefore be summed directly. (33) allows us to obtain a very high precision quickly. It is much faster than the direct numerical integration for lower n, but slower for large n. Thus, we have used (33) for the first ~ 20 terms and a direct integration for $n \gtrsim 20$, in both cases obtaining a precision

of $\lesssim 10^{-100}$. Summing the first $1 \le m \le 500$ terms in (33) is enough to achieve this precision for $n \sim 20$. Fewer terms are needed for lower n, while more terms are needed for larger n. The precision can be checked either by summing many more terms and checking the relative precision, or by estimating the size of the remainder by summing the large order approximation (35) and (36) from e.g., m = 500 to ∞ in terms of polylogarithms, or by comparing with the direct integration.

C. Momentum expectation value

We consider first the momentum expectation value and to leading order in $\chi \ll 1$. We use the ansatz $\mathbf{N}^{(n)} \approx a_0^{2n} b_0^{n+1} \{A_n + C_n b_0, B_n b_0\}$. We find a recursive formula, $A_n = -(2/3)A_{n-1}$ and

$$B_n = \frac{2}{3} \left(A_{n-1} - \left[1 + \frac{1}{n} \right] B_{n-1} \right) \tag{37}$$

so C_n is not needed, with initial conditions $A_1 = -2/3$ and $B_1 = 2/3$. With Eq. (2.2.7) in [46] we can solve this in terms of the harmonic number H_n ,

$$B_n = \left(-\frac{2}{3}\right)^n (1+n)(1-H_{n+1}). \tag{38}$$

Resumming this gives

$$\sum_{n=1}^{\infty} \{0, 1\} \cdot \mathbf{N}^{(n)} T^n = \frac{b_0^2 \ln\left[1 + \frac{2}{3}a_0^2 b_0 T\right]}{\left[1 + \frac{2}{3}a_0^2 b_0 T\right]^2}.$$
 (39)

Apart from some constants, this is essentially the same function³ as the spin difference for a constant field [14], even though we are considering a different spin direction.

At higher energy we can use the same resummation methods or solve the integro-differential equation as in the constant-field case [14]. For the resummation approach we have used the Borel-Leroy method (16) with c=1.588... [same as in (30)], $b_n=3+(3/2)n$ and $b_n=4+(3/2)n$ for $\langle kP\rangle(\uparrow)+\langle kP\rangle(\downarrow)$ and $\langle kP\rangle(\uparrow)-\langle kP\rangle(\downarrow)$, respectively, and $[40/(40+\text{ceil.}[b_n]+1)]$ or a nearby approximant if there is a spurious pole for some n.

The integro-differential equation has been solved using the midpoint method, i.e., we make an interpolation function for the b_0 dependence of

$$\mathbf{M}(T_{n+1/2}) = \mathbf{M}(T_n) + \frac{dT}{2} \frac{\partial \mathbf{M}(T_n)}{\partial T}$$
(40)

where the derivative is obtained from (the circular version of) (5), and then we use this to compute an interpolation function at the next step

 $^{^2}$ If δ is close to 1 it might be useful to find some resummation. However, then a_0 would have to be large, which might mean that an LCF expansion in $1/a_0$ would anyway be more convenient.

³Note that T includes a factor of a_0 in the constant field case (4), but not in the circular case (18).

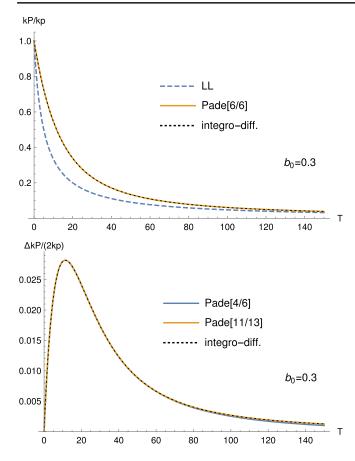


FIG. 3. Momentum expectation value for a circularly polarized monochromatic field.

$$\mathbf{M}(T_{n+1}) = \mathbf{M}(T_n) + dT \frac{\partial \mathbf{M}(T_{n+1/2})}{\partial T}, \qquad (41)$$

and so on. In all cases in this paper we have used a step size dT = 1/10. This integro-differential approach is much slower than the resummation approach.

Figure 3 shows that both approaches give the same result. These methods are the same as for the constant-field case. Figure 3 shows that also the results are very similar in shape as in the constant field case.

D. Sokolov-Ternov effect

Without RR, three of the four elements in the sum $M^C + M^L \propto -\hat{k} \; \hat{k}$ cancel, and the sum of all orders gives

$$\mathbb{P} = \frac{1}{2} \mathbf{N}_0 \cdot \begin{pmatrix} 1 & 0 \\ 0 & \exp\left[-\mathcal{T}\hat{M}_1\right] \end{pmatrix} \cdot \mathbf{N}_f. \tag{42}$$

The factor \hat{M}_1 can be obtained directly by integrating $\mathbf{M}^{\mathrm{C}} + \mathbf{M}^{\mathrm{L}}$. But we do not need to obtain it in order to see that without RR the degree of polarization can only be reduced: If we start with an unpolarized electron $\mathbf{N}_0 = \{1,0\}$ then there is no effect at all, while if we start with a polarized electron $\mathbf{N}_0 = \{1,\pm 1\}$ then the degree of

polarization decreases as the pulse length \mathcal{T} increases. We will now show that this changes if RR is included.

To see what happens when RR is included, we consider first the low-energy limit. With the ansatz

$$\mathbf{M}_{n} = a_{0}^{2(n-1)} b_{0}^{n+1} \begin{pmatrix} 0 & B_{n} b_{0} \\ 0 & A_{n} + C_{n} b_{0} \end{pmatrix}$$
(43)

for $n \ge 1$, we find $A_n = (-2/3)^{n-1}A_1$ and

$$B_n = -\frac{2}{3} \left(1 + \frac{1}{n} \right) B_{n-1} - \left(-\frac{2}{3} \right)^{n-1} A_1, \tag{44}$$

while C_n is not needed. We can solve this using again Eq. (2.2.7) in [46]. We find

$$B_n = \left(-\frac{2}{3}\right)^{n-1} (1+n) \left(\frac{3}{2} - H_{1+n}\right) A_1, \tag{45}$$

which is very similar to the LCF case. Summing all orders we find

$$\mathbb{P} = \frac{1}{2} \mathbf{N}_0 \cdot \begin{pmatrix} 1 & B \\ 0 & 1+A \end{pmatrix} \cdot \mathbf{N}_f, \tag{46}$$

where

$$A = -\frac{b_0^2 a_0^2 \mathcal{T}}{1 + \frac{2}{3} b_0 a_0^2 \mathcal{T}} \mathcal{I}$$
(47)

and

$$B = \frac{-b_0^2}{(1 + \frac{2}{3}b_0a_0^2\mathcal{T})^2} \left(b_0a_0^2\mathcal{T}\left[1 + \frac{1}{3}b_0a_0^2\mathcal{T}\right] - \frac{3}{2}\ln\left[1 + \frac{2}{3}b_0a_0^2\mathcal{T}\right]\right)\mathcal{I},\tag{48}$$

where

$$\mathcal{I} = \frac{1}{a_0^2} \int_0^\infty \mathrm{d}\gamma \gamma^2 \mathcal{J}_0(\gamma). \tag{49}$$

For small a_0 we have $\mathcal{I} \approx 2/5$, and $\mathcal{I} \approx 35a_0/(24\sqrt{3})$ for large a_0 . As a check, note that by summing over the final spin we have $\mathbb{P} = \mathbb{N}_0 \cdot \{1,0\} = 1$, which also explains why there are only two nontrivial elements. For an unpolarized initial electron we have $\mathbb{P} = (1/2)\{1,B\} \cdot \mathbb{N}_f$, so B gives an induced polarization. Hence, without RR the degree of polarization can only be reduced, but with RR we find a nonzero induced polarization. For a very long pulse we find

$$A \approx -\frac{3}{2}b_0 \mathcal{I} \quad B \approx -\frac{3}{4}b_0^2 \mathcal{I}. \tag{50}$$

This should be compared with the Sokolov-Ternov effect. For low-energy electrons, $b_0 \ll 1$, the induced polarization

is a small effect. The reason for this is that RR leads to a reduction of the (lightfront) longitudinal momentum.

For larger b_0 we can use the resummation or the integrodifferential approach. We have used the Borel-Leroy method for resumming the b_0 expansion. For B we have (16) with $b_n = 4 + (3/2)n$ for $n \ge 2$, while for A we have $b_n = 3 + (3/2)n$ for $n \ge 1$. For both terms c = 1.588...and we have used $[40/(40 + \text{ceil.}[b_n] + 1)]$.

The results are shown in Fig. 4. At $b_0 = 0.3$ the induced polarization is on a $\sim 1\%$ level. This is not a large effect compared to the constant-field case, but it is non-negligible, and it is a clear signal of RR since without RR there would be no induced polarization. One might have expected such a signal, because, although the off-diagonal terms vanish in $\mathbf{M}^{C} + \mathbf{M}^{L} \propto -\hat{\mathbf{k}} \hat{\mathbf{k}}$ (which, recall, is the reason for the absence of an induced polarization without RR), RR breaks this cancellation because all the steps that follow a photon emission step \mathbf{M}^{C} depend on the longitudinal momentum that was emitted in that step, but subsequent steps do not depend on the photon momentum in the loop \mathbf{M}^{L} .

E. Circular perturbative

In this section we will expand each order in α to leading order in $a_0 \ll 1$. In this regime the \mathcal{J}_i functions become very simple,

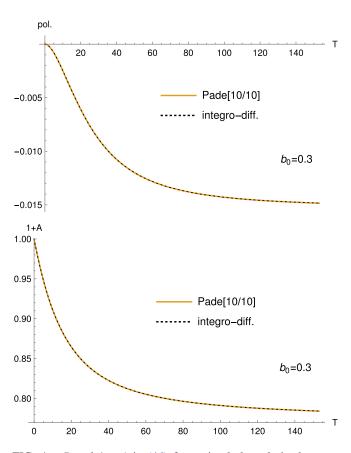


FIG. 4. B and 1 + A in (46) for a circularly polarized monochromatic field.

$$\mathcal{J}_0 = a_0^2 \frac{\gamma}{4} (2 - \gamma) \quad \mathcal{J}_1 = \frac{a_0^2}{2} \quad \mathcal{J}_2 = a_0^2 \frac{1 - \gamma}{2}$$
 (51)

and $\mathcal{J}_i = 0$ for $\gamma > 2$. Since \mathbf{M}^{C} and \mathbf{M}^{L} are proportional to a_0^2 it is natural to define the following effective expansion parameter

circular field, perturbative:
$$T = \alpha a_0^2 \Delta \phi$$
. (52)

Since $a_0^2 \ll 1$, the pulse length would have to be much longer in order to have a not small T. This makes perhaps this regime less interesting from an experimental point of view. It can also be challenging for calculations, because small a_0 in general makes the corrections to the incoherent product approximation more important, so the pulse length has to be extra long to compensate for this. However, there is an interesting difference compared to the other cases: The b_0 expansions are no longer asymptotic, but have a finite radius of convergence. We can therefore resum both the b_0 and the α expansions with Padé approximants. In Fig. 5 we have resummed the b_0 expansion with $\lfloor 46/52 \rfloor$ and $\lfloor 46/51 \rfloor$ for $\langle kP \rangle (\uparrow) + \langle kP \rangle (\downarrow)$ and $\langle kP \rangle (\uparrow) - \langle kP \rangle (\downarrow)$, respectively.

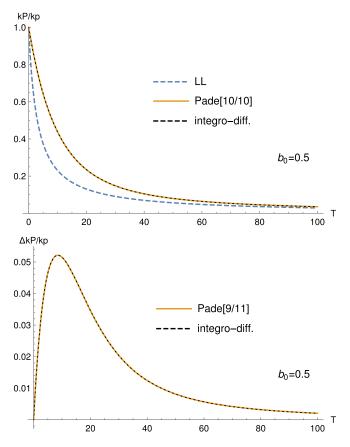


FIG. 5. Momentum expectation value to leading order in $a_0 \ll 1$.

V. RESUMMATION OF LAD

Both the classical LL and our quantum approximation give series in α which can be resummed with Padé approximants. However, we know/suspect that the full QED expansion is asymptotic. If this is the case then LL cannot be the exact (classical) equation, but LAD might be since it leads to asymptotic series in α . In [7] it was shown how the nonrelativistic LAD can be resummed with the Borel method. In this section we will consider the full LAD expansion. Since we cannot obtain a compact/closed form expression for each order, we will restrict to a constant-crossed field for simplicity, $f_{\mu\nu} = k_{\mu}a'_{\nu} - a'_{\mu}k_{\nu}$ with a' constant.

LAD is given by

$$\ddot{x}^{\mu} = f^{\mu\nu}\dot{x}_{\nu} + \frac{2}{3}\alpha(\ddot{x}^{\mu} + \ddot{x}^{2}\dot{x}^{\mu}). \tag{53}$$

Since we consider a constant field, only derivatives of the position appear. We denote $P^{\mu} = \dot{x}$. We write the solution as

$$P^{\mu}(\tau) = \left(g_0 \delta^{\mu}_{\nu} + g_1 \frac{f^{\mu}_{\nu}}{\chi_0} + g_2 \frac{(f^2)^{\mu}_{\nu}}{\chi_0^2}\right) P^{(0)\nu}, \quad (54)$$

where $P^{(0)}$ is the momentum at $\tau=0$, $\chi_0=\sqrt{P^{(0)}f^2P^{(0)}}$ and $g_i(\tau)$ are functions to be determined. From the on-shell condition $P^2=1$ and $f^3=0$ we find

$$g_2 = \frac{1 + g_1^2 - g_0^2}{2g_0}. (55)$$

Instead of proper time τ , it is natural to use a rescaled lightfront time

$$u(\tau) = \frac{\chi_0}{kP^{(0)}} kx(\tau) \quad \frac{\mathrm{d}}{\mathrm{d}\tau} = \chi_0 g_0 \frac{\mathrm{d}}{\mathrm{d}u}.$$
 (56)

We also define

$$\delta = \frac{2}{3}\alpha\chi_0. \tag{57}$$

Projecting LAD with $P^{(0)}f^2$ and $P^{(0)}f$ gives us two equations for the two unknown functions

$$g'_0(u) = \delta(\partial_u[g_0g'_0] + P'^2g_0^2)$$

$$g'_1(u) = 1 + \delta(\partial_u[g_0g'_1] + P'^2g_0g_1),$$
(58)

where F' = dF/du and

$$P^{\prime 2} = g_0^{\prime 2} + 2g_0^{\prime}g_2^{\prime} - g_1^{\prime 2}. \tag{59}$$

We will first attempt to solve (58) by a direct expansion in δ . The zeroth order is given by

$$g_0 = 1 + \mathcal{O}(\delta) \quad g_1 = u + \mathcal{O}(\delta). \tag{60}$$

Substituting $g_i = \sum_{n=0}^N \delta^n g_i^{(n)}(u) + \mathcal{O}(\delta^{N+1})$ into the right-hand sides of (58) and integrating over u gives $g_i^{(N+1)}(u)$. For the first couple of orders we find

$$g_{0} = 1 - u\delta + u^{2}\delta^{2} + (6u - u^{3})\delta^{3} + (-18u^{2} + u^{4})\delta^{4}$$

$$+ (-80u + 36u^{3} - u^{5})\delta^{5} + (392u^{2} - 60u^{4} + u^{6})\delta^{6}$$

$$+ \left(1520u - \frac{3524u^{3}}{3} + 90u^{5} - u^{7}\right)\delta^{7} + \mathcal{O}(\delta^{8})$$
(61)

$$g_{1} = u - \frac{u^{2}\delta}{2} + \left(-2u + \frac{u^{3}}{2}\right)\delta^{2} + \left(6u^{2} - \frac{u^{4}}{2}\right)\delta^{3} + \left(20u - \frac{41u^{3}}{3} + \frac{u^{5}}{2}\right)\delta^{4} + \mathcal{O}(\delta^{5}).$$
 (62)

Each order is a polynomial in u and the first tens of orders can be obtained quickly. However, the number of terms increases as we go to higher orders and eventually the calculation becomes very slow. We stopped at $\mathcal{O}(\delta^{77})$. To study the (non)-convergence properties of a series one can usually plot the ratios of neighboring coefficients. However, here it turns out to be easier to study the even and odd orders separately. In Fig. 6 we plot the ratios of the odd-order coefficients for a small, moderately large and large value of u. For small u we see a typical line for an asymptotic series with factorially growing coefficients with alternating sign. Such a series can be dealt with using the Borel-Padé method. For the moderately large u we have a very different plot with no discernible pattern up to the maximum order calculated. In the large u plot we again see a simple pattern, but this time in the first ~ 20 orders rather than as an asymptotic limit. This almost constant line is what one expects from a function that is close to LL, which has a finite radius of convergence. We can understand this from (61): After a long time in the field, $u \gg 1$, one may expect the highest power in u at each order in δ to give a good approximation, i.e.,

$$g_0 \to \sum_{n=0}^{\infty} (-\delta u)^n = \frac{1}{1 + \delta u} = g_0^{LL}.$$
 (63)

Hence, for a sufficiently long pulse we expect the solution to LAD to converge to the LL solution (this has been shown in [8,9]), where the in general asymptotic series can be approximated by a convergent series. For smaller u, on the other hand, we have checked that the δ expansion can be resummed with the Borel-Padé method. However, as u increases this resummation seems to eventually break down (given the fixed number of terms we have calculated). One might try to find different resummations or a way to obtain

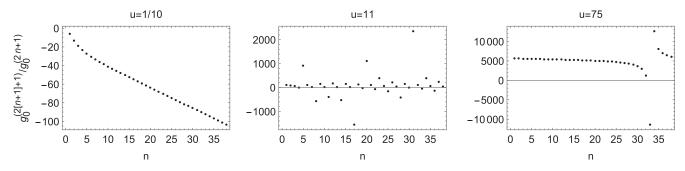


FIG. 6. Ratios of neighboring odd-order coefficients for g_0 , $g_0^{(2[n+1]+1)}/g_0^{(2n+1)}$, with $u = \{1/10, 11, 75\}$ from left to right. The corresponding plots for the even orders, $g_0^{(2[n+1])}/g_0^{(2n)}$, have a similar behavior.

higher orders. However, we have used a different approach instead, which seems to work for any u.

In this approach we begin by noting that, e.g., from (63), we expect to have a considerable change in the momentum for $\delta u \sim 1$. We therefore change variable to $v = \delta u$ and treat v as independent of δ . We also define

$$h_0(v) = g_0(u = v/\delta)$$
 $h_1(v) = \delta g_1(u = v/\delta).$ (64)

We expand the solution as

$$h_i(v) = \sum_{n=0}^{\infty} h_i^{(n)}(v)\delta^{2n}.$$
 (65)

Note that this is not the same as repeatedly substituting LAD into itself to remove higher-order derivatives, an approach that has recently been studied in [12]. To zeroth order in δ , the solution is now given by LL

$$h_0 = \frac{1}{1+v} + \mathcal{O}(\delta^2) = h_0^{LL} + \mathcal{O}(\delta^2)$$
 (66)

and

$$h_1 = v - \frac{v^2}{2(1+v)} + \mathcal{O}(\delta^2) = h_1^{LL} + \mathcal{O}(\delta^2).$$
 (67)

In this approach, each order is not just a polynomial in v. Fortunately, though, we can still find simple explicit expressions for each order. The first couple of orders are given by

$$h_0 = \frac{1}{v+1} + \frac{6v}{(v+1)^3} \delta^2$$

$$-\frac{4v(60+6v+11v^2)}{3(1+v)^5} \delta^4 + \mathcal{O}(\delta^6)$$
 (68)

and

$$h_{1} = \frac{v(2+v)}{2(1+v)} + \left(\frac{4\log(1+v)}{1+v} - \frac{3v(2+2v+v^{2})}{(1+v)^{3}}\right)\delta^{2} + 2v\left(\frac{12\log(1+v)}{(1+v)^{3}} + \frac{30-48v+4v^{2}+28v^{3}+11v^{4}}{3(1+v)^{5}}\right)\delta^{4} + \mathcal{O}(\delta^{6})$$
(69)

Note that using v instead of u gives series in δ^2 rather than δ . We have obtained the first 34 terms, i.e., up to $\mathcal{O}(\delta^{66})$. The expressions become very long, we just note that in the limit $v \to \infty$ all higher orders scale as 1/v compared to the zeroth (LL) order. If we go back to the original variable u, then each function $h_i^{(n)}(v=\delta u)$ provides a partial resummation of the δ expansion from the first approach. In particular, the exact solution to LL is already included. Consequently, the ratios of neighboring coefficients, $h_0^{(n+1)}(v)/h_0^{(n)}(v)$ shown in Fig. 7, now have a simple behavior (typical for alternating and factorially divergent series) for both small and large v, in contrast to Fig. 6.

This means that for an arbitrary value of v we can resum the δ expansion using the Borel-Padé method. We have used [15/15] approximants. The results are shown in Fig. 8 and 9 for $\delta=1/3$. The reason for choosing this rather large value of δ is to be able to clearly distinguish LAD and LL.

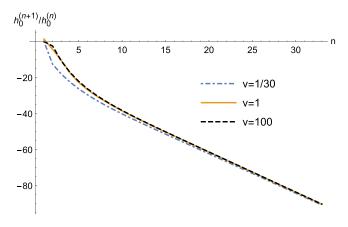


FIG. 7. Ratios of neighboring coefficients for $h_0(v)$, $h_0^{(n+1)}(v)/h_0^{(n)}(v)$.

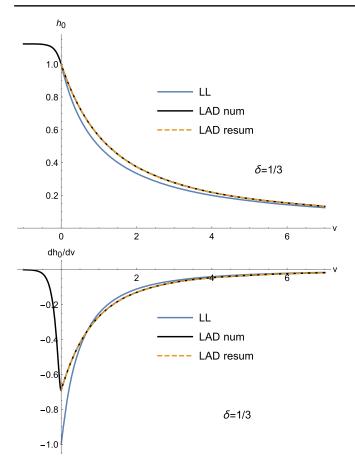


FIG. 8. Resummation of LAD using Borel-Padé[15/15]. The numerical results have been obtained by integrating LAD backwards with the resummation result at v=15 as final conditions.

For smaller δ it is difficult to see the difference because it is given by a function of δ^2 . After a very long time in the field, $v \gg 1$ or $u \gg 1/\delta$, the result converges to LL. We have also compared the resummation with a numerical solution to LAD. Since LAD is numerically unstable when integrated forwards in time, we have integrated LAD backwards in time [59,60] with the resummed result $h_i^{\text{resum}}(v_f \gg 1)$ at

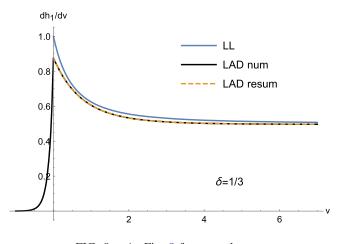


FIG. 9. As Fig. 8 for second term.

one instant of time v_f as final conditions. As shown in Fig. 8 and 9, we find excellent agreement between the numerical solution and the resummation. The considered value of δ is quite large, but since the resummation is based on a $\delta \ll 1$ expansion, it would be even more precise for smaller δ . For the numerical solution we have continued passed v < 0 assuming that the field is zero for v < 0, to illustrate the fact that the solution has pre-acceleration. This means that $P_{\mu}^{(0)}$ is strictly speaking not the (exact) initial momentum even if the field is zero at u < 0. Thus, the δ series gives a pre-acceleration solution without runaway.

To obtain the runaway solutions one would start with a trans-series-like ansatz; see e.g., the recent papers [61,62] for trans-series solutions to differential equations. We first write the solution as

$$g_i(u) = g_i^{(0)}(u) + g_i^{(1)}(u)$$
(70)

where $g_i^{(0)}(u)$ is the solution found above by resumming the perturbative series in δ . As mentioned, $g_i^{(0)}(u)$ is free from runaways. We expand the nonlinear equations (58) to linear order in the perturbation $g_i^{(1)}$. This gives two homogeneous differential equations which we can solve with the ansatz

$$g_i^{(1)}(u) = e^{u/\delta} \sum_{n=0}^{\infty} g_i^{(1,n)}(u) \delta^n.$$
 (71)

To leading order we find

$$\begin{pmatrix} g_0^{(1)} \\ g_1^{(1)} \end{pmatrix} = \operatorname{Re}c \begin{pmatrix} 1 + \mathcal{O}(\delta) \\ i + u + \mathcal{O}(\delta) \end{pmatrix} \exp\left\{ \frac{u}{\delta} + \frac{u^2}{2} + 2iu \right\}, \tag{72}$$

where c is an arbitrary, small complex number. The resummation approach for $g_i^{(0)}(\boldsymbol{u})$ selects one particular value of the initial acceleration with the initial momentum as input. The arbitrary complex constant in (72) gives the remaining solutions to LAD. It is straightforward to calculate, at least the first ~ 20 , terms in the δ expansion in the preexponential part of (72) using the expansion of $g_i^{(0)}(u)$ which we have already obtained. However, we can already see from (72) that any nonzero value of c will lead to runaway solutions. Note that both LL and the preacceleration solution of LAD varies significantly on time scales $u \gtrsim 1/\delta$. But (72), on the other hand, leads to an exponential difference already at $u \sim 1$. Hence, if we start with a small perturbation at u = 0, then at $u \sim 1$ we already have a much larger $g_i^{(1)}$. Thus, we can see explicitly the reason why LAD is numerically unstable when integrated forwards in time. To obtain the full runaway solutions one would have to include additional terms $g_i^{(n)}$ with $e^{nu/\delta}$, $n \ge 1$. However,

since the runaway solutions are anyway unphysical, we will stop here.

VI. CONCLUSIONS

We have studied RR using our recently proposed resummation methods. In [14] we studied the electron momentum expectation value for a constant-crossed field and with the initial electron either unpolarized or polarized along the magnetic field. Here we have considered the momentum for a circularly polarized monochromatic field with $a_0 = 1$ and the initial electron unpolarized or polarized along the laser propagation direction. We have found that, as a function of an appropriate effective pulse length parameter (proportional to α), both the spin-averaged momentum $P_-(\uparrow) + P_-(\downarrow)$ and the difference $P_-(\uparrow) - P_-(\downarrow)$ look similar to the constant-field case [14], even though both the field and spin direction are different.

In both cases, we have also studied the polarization of the final electron. For a constant field and without RR, the polarization is given by the Sokolov-Ternov result, which is ≈ 0.92 asymptotically. Here we find that with RR there is still a nonzero degree of polarization asymptotically, but this asymptotic value is reached on a much shorter time scale and

is significantly lower. For a circularly polarized monochromatic field and without RR, there is no induced polarization to leading order [19,24,25], because the loop cancels the contribution from photon emission. However, here we find that with RR there is a nonzero degree of polarization.

Finally, in order to gain some insights into what one might find if one could find a way to include the terms that have been neglected in the incoherent product approach, we have shown how to use resummation methods for LAD.

ACKNOWLEDGMENTS

On the day that this manuscript was submitted, [63] appeared on the arXiv. The subject of [63] is very similar to our section about resummation of LAD. Our Fig. 8 is very similar to their Fig. 7 (even the plot style happens to be similar). The main difference is how LAD has been resummed. [63] resummed terms appearing in the differential equation, while here we have resummed the solution directly. We have checked, though, that the expansions for $g_0(u=0)$ and $g_1(u=0)$ agree with Eq. (7) in [63]. I thank Robin Ekman, Tom Heinzl and Anton Ilderton for discussions about this (after the appearance of [63]).

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