Does relativistic motion always degrade quantum Fisher information?

Xiaobao Liu,¹ Jiliang Jing,^{2,*} Zehua Tian,^{3,4,5,6,†} and Weiping Yao^{1,‡}

¹Department of physics and electrical engineering, Liupanshui Normal University,

Liupanshui 553004, Guizhou, China

²Department of Physics, Key Laboratory of Low Dimensional Quantum Structures and Quantum Control

of Ministry of Education, and Synergetic Innovation Center for Quantum Effects and Applications, Hunan Normal University, Changsha, Hunan 410081, People's Republic of China

³CAS Key Laboratory of Microscale Magnetic Resonance and Department of Modern Physics,

University of Science and Technology of China, Hefei 230026, China

⁴Key Laboratory for Research in Galaxies and Cosmology, Chinese Academy of Science,

96 JinZhai Road, Hefei 230026, Anhui, China

⁵Hefei National Laboratory for Physical Sciences at the Microscale,

University of Science and Technology of China, Hefei 230026, China

⁶Synergetic Innovation Center of Quantum Information and Quantum Physics,

University of Science and Technology of China, Hefei 230026, China

(Received 23 December 2020; revised 5 May 2021; accepted 27 May 2021; published 24 June 2021)

We investigate the ultimate estimation precision, characterized by the quantum Fisher information, of a two-level atom as a detector which is coupled to massless scalar field in the Minkowski vacuum. It has been shown that for an inertial detector moving with a constant velocity, its quantum Fisher information is completely unaffected by the velocity, however, it still decays over time due to the decoherence caused by the interaction between the atom and the field. In addition, for a uniformly accelerated detector (w = 0) moving along spatially straight line, the accelerated motion will reduce the quantum Fisher information in the estimation of state parameters. However, when the detector trajectory is generated by a combination of the linear accelerated motion and a component of the four-velocity $w = dy/d\tau$, we find quite unlike the previous results that, for the nonrelativistic case ($w \ll 1$), the acceleration of the quantum Fisher information, while the four-velocity component will suppress the degradation of the quantum Fisher information, and thus could enhance the precision of parameters estimation. Furthermore, in the case for ultrarelativistic velocities ($w \rightarrow \infty$), although the detector still interacts with the environment, it behaves as if it were a closed system as a consequence of relativity correction associated to the velocity, and the quantum Fisher information in this case can be shielded from the effect of the external environment, and thus from the relativistic motion.

DOI: 10.1103/PhysRevD.103.125025

I. INTRODUCTION

Quantum Fisher information (QFI) [1-3] has attracted much interest since it is not only of great significance in quantum estimation theory and quantum information theory [4-12], but also strongly related to rapid progress in quantum-enhanced metrology [13-15]. Indeed, in the field of quantum metrology, the QFI acted as a crucial measure of information content of quantum state, which has already played a significant role in quantum statistical inference for its inextricable relationship with Cramér-Rao

tianzh@ustc.edu.cn

inequality, namely: the lower bound of the estimation error is characterized by the Cramér-Rao bound which is inversely proportional to QFI [16-18]. With different models of the probe systems and different parameters to be estimated, QFI has been applied in various quantum information processing tasks, such as measurements of non-Markovianity [19], entanglement detection [20], qubit thermometry [21,22], as well as relativistic parameters estimation [23-33], and so on. However, a realistic quantum system will unavoidably suffer from the quantum decoherence, due to the interaction between the system and its surrounding environment, which results in the QFI attenuation and thus the estimation precision degradation [34-40]. Moreover, there has been extensive works to investigate the degradation of the QFI caused by the effects of relativistic motion [41–43], or the curvature of curved

^{*}Corresponding author.

jljing@hunn.edu.cn Corresponding author.

[‡]Corresponding author. yao11a@126.com

spacetime [44,45]. In this regard, how to inhibit the attenuation of QFI becomes the key problem to be solved.

On the other hand, there exists a series of papers to investigate how the motion of detector affects the detector's dynamics and the information that encoded in the quantum state [46–55]. For instance, it has been shown in Ref. [56] that the parameters estimation decreases exponentially with time for an inertial atom coupled with an electromagnetic field. Another examples, when the detector is accelerated with a constant acceleration, have given manifest evidence that the accelerated motion will reduce the precision of parameters estimation due to the Unruh effect which was undertaken in Refs. [42,43]. Similar results for the quantum coherence were found in Refs. [57-59]. In such references, the quantum information, such as QFI and quantum coherence, was considered in the framework of open quantum systems will be degraded when the detector is in inertial motion or moves along a straight line with a constant acceleration. Currently a special case of a stationary trajectory was considered in Ref. [60], in which the detector moves along with constant independent magnitudes of both the four-acceleration and of a timelike proper time derivative of the four-acceleration. In that work, the four-velocity component inhibit the increase of the response function of the Unruh-DeWitt detector in the nonrelativistic limit; Meanwhile, in the ultrarelativistic limit, the response function is completely suppressed. Consequently, in contrast to the individual linear accelerated motion and movement with constant velocity, if we consider the relativistic motion of Unruh-DeWitt detector is a superposition of both a linear accelerated motion and a component of the four-velocity as done in Ref. [60], whether can suppress the relativistic motion/bath induced degradation of quantum information encoded in quantum system?

In this paper, we will study the performance of QFI regarding the estimation of parameters for a two-level system as the detector coupled to massless scalar field. Here we consider this detector moving in Minkowski spacetime along an unbounded spatial trajectory in a twodimensional spatial plane with constant independent magnitudes of both the four-acceleration and of a timelike proper time derivative of the four-acceleration, which was shown in Ref. [60]. In such a reference, in a Fermi-Walker frame moving with the detector, the direction of the acceleration rotates at a constant rate around a great circle. Our analytical results demonstrate that in the nonrelativistic limit, the four-velocity component will suppress the attenuation of the QFI, which implies that the precision of parameters estimation of the Unruh-DeWitt detector moving along such a trajectory can be enhanced. What is more, in the ultra-relativistic limit, the QFI may even be shield from the effects of the external environment and detector's motion, as if this detector were a closed system.

Our paper is structured as follows. In Sec. II, we introduce the QFI and the dynamic evolution of detector coupling with massless scalar field. In Sec. III, we study the QFI in the parameters estimation in two situations: for nonrelativistic and ultrarelativistic velocities. Finally, in Sec. IV, we give our conclusions and discussions.

Throughout the whole paper we employ natural units $c = \hbar = 1$. Relevant constants are restored when needed for the sake of clarity.

II. QUANTUM FISHER INFORMATION AND DYNAMIC EVOLUTION OF A TWO-LEVEL SYSTEM COUPLED WITH SCALAR FIELDS

In quantum metrology, any given quantum state $\rho(X)$ characterized by the unknown parameter *X* can be inferred from a set of measurements on the state. The measurements usually modeled mathematically by a set of positive operator-valued measures (POVM), whose elements, $\{\Pi_i\}$, saturate to $\sum_i \Pi_i \Pi_i^{\dagger} = 1$. Through the optimization of the measurements and the estimator, an ultimate bound to precision of the unknown parameter estimation satisfies the quantum Cramér-Rao inequality [16–18]

$$\operatorname{Var}(\mathbf{X}) \ge \frac{1}{\mathrm{MF}_{\mathbf{X}}},$$
 (1)

where *M* represents the number of measurements, and $F_X = \text{Tr}[\rho(X)L^2]$ is the QFI. Here, *L* denotes the symmetric logarithmic derivative satisfying the partial differential equation

$$\frac{\partial \rho(X)}{\partial X} = \frac{L\rho(X) + \rho(X)L}{2}.$$
 (2)

For a two-level quantum system, the reduced density matrix of the system can be expressed in the Bloch sphere representation as

$$\rho(\tau) = \frac{1}{2} (\mathbf{I} + \boldsymbol{\omega}(\tau) \cdot \boldsymbol{\sigma}), \qquad (3)$$

where $\boldsymbol{\omega} = (\omega_1, \omega_2, \omega_3)$ represents the Bloch vector, and $\boldsymbol{\sigma} = (\sigma_1, \sigma_2, \sigma_3)$ denotes the Pauli matrices. As shown in Ref. [40], F_X can be described in the simple form as follows:

$$F_X = \begin{cases} |\partial_X \boldsymbol{\omega}|^2 + \frac{(\boldsymbol{\omega}\partial_X \boldsymbol{\omega})^2}{1 - |\boldsymbol{\omega}|^2}, & |\boldsymbol{\omega}| < 1, \\ |\partial_X \boldsymbol{\omega}|^2, & |\boldsymbol{\omega}| = 1. \end{cases}$$
(4)

In quantum sense, any system should be regarded as an open system due to the interaction between the system and its surrounding environments. Therefore, let us study a two-level detector interacting with massless scalar field and, in this regard, the total Hamiltonian of the detectorfield system can be described as

$$H = H_s + H_{\Phi(x)} + H_I, \tag{5}$$

where $H_s = \frac{1}{2}\omega_0\sigma_z$ is the Hamiltonian of the detector with ω_0 and σ_z being the energy-level spacing of the atom and Pauli matrix, respectively, $H_{\Phi(x)}$ represents the Hamiltonian of scalar field and H_I denotes their interaction Hamiltonian. We assume that the coupling between the detector and the massless scalar field is of the form,

$$H_I = \mu(\sigma_+ + \sigma_-)\Phi(x(\tau)), \tag{6}$$

where μ is the coupling constant that we assume to be small, σ_+ (σ_-) is the rasing (lowering) operator of the detector, and $\Phi(x(\tau))$ corresponds to the scalar field operator with τ being the detector's proper time. Note that this interaction is the analogy to the electric dipole interaction. Specifically, $H_I = \mu' \hat{m} \Phi(x(\tau)) = \sum_{i,j=0}^{1} |i\rangle \times$ $\langle i|\mu' \hat{m}|j\rangle \langle j| \Phi(x(\tau)) = \mu(\sigma_+ + \sigma_-) \Phi(x(\tau))$, where we have used $\sum_i |i\rangle \langle i| = I$ and $\langle i|\mu \hat{m}|i\rangle = 0$.

We assume the scalar field in vacuum state $|0\rangle$, which is defined by $a_{\mathbf{k}}|0\rangle = 0$ for all **k**. At the beginning, the total density matrix of the detector-field system can be written as $\rho_{\text{tot}} = \rho_s(0) \otimes |0\rangle \langle 0|$, in which $\rho_s(0)$ is the initial reduced density matrix of the detector. For the whole system, its equation of motion in the interaction picture is given by

$$\frac{\partial \rho_{\text{tot}}(\tau)}{\partial \tau} = -i[H_I(\tau), \rho_{\text{tot}}(\tau)].$$
(7)

We have $\rho_{\text{tot}}(\tau) = \rho_{\text{tot}}(0) - i \int_0^{\tau} ds [H_I(s), \rho_{\text{tot}}(s)]$. Substitute this solution back into Eq. (7) and take the partial trace:

$$\frac{\partial \rho_s(\tau)}{\partial \tau} = -\int_0^\tau ds T r_B[H_I(\tau), [H_I(s), \rho_{\text{tot}}(s)]], \quad (8)$$

where we have used $Tr_B[H_I(\tau), \rho_{tot}(0)] = 0$. We now make our first approximation. For a sufficiently large bath that is in particular much larger than the system, it is reasonable to assume that while the system undergoes nontrivial evolution, the bath remains unaffected, and hence that the state of the composite system at any time is uncorrelated, i.e.,

$$\rho_{\rm tot}(s) \approx \rho_s(s) \otimes \rho_B. \tag{9}$$

This is the so called Born approximation [61]. Let us change the variables to $\tau' = \tau - s$, so that $\int_0^{\tau} ds = -\int_{\tau}^0 d\tau' = \int_0^{\tau} d\tau'$, and Eq. (8) can be rewritten as

$$\frac{\partial \rho_s(\tau)}{\partial \tau} = -\int_0^\tau d\tau' Tr_B[H_I(\tau), [H_I(\tau - \tau'), \rho_s(\tau - \tau') \otimes \rho_B]].$$
(10)

We then introduce the so-called Markov approximation [61]. It states that the bath has a very short correlation time τ_B . If $\tau \gg \tau_B$, we can replace $\rho_s(\tau - \tau')$ by $\rho_s(\tau)$, since the short "memory" of the bath correlation function causes it to keep track of events only within the short period $[0, \tau_B]$. Under this approximation, we have

$$\frac{\partial \rho_s(\tau)}{\partial \tau} = -\int_0^\tau d\tau' Tr_B[H_I(\tau), [H_I(\tau - \tau'), \rho_s(\tau) \otimes \rho_B]].$$
(11)

Moreover, for the same reason (correlation function negligible for $\tau' \gg \tau_B$ [61]) we can extend the upper limit of the integral to infinity without changing the value of the integral. Therefore,

$$\frac{\partial \rho_s(\tau)}{\partial \tau} = -\int_0^\infty d\tau' Tr_B[H_I(\tau), [H_I(\tau - \tau'), \rho_s(\tau) \otimes \rho_B]].$$
(12)

Substitute the interaction Hamiltonian $H_I(\tau) = \mu(\sigma_+ e^{i\omega_0\tau} + \sigma_- e^{-i\omega_0\tau})\Phi(x(\tau))$ into Eq. (12) and after long but straightforward calculations, we can derive finally the master equation in the Kossakowski-Lindblad form [62–64]

$$\frac{\partial \rho_s(\tau)}{\partial \tau} = -i[H_{\text{eff}}, \rho_s(\tau)] + \sum_{j=1}^3 [2L_j \rho_s L_j^{\dagger} - L_j^{\dagger} L_j \rho_s - \rho_s L_j^{\dagger} L_j],$$
(13)

where

$$H_{\rm eff} = \frac{1}{2} \Omega \sigma_z = \frac{1}{2} \{ \omega_0 + \mu^2 {\rm Im}(\Gamma_+ + \Gamma_-) \} \sigma_z \quad (14)$$

is the effective Hamiltonian in which Ω denotes the effective energy level-spacing of the detector with a correction term $\mu^2 \text{Im}(\Gamma_+ + \Gamma_-)$ being the Lamb shift. Note that the Lamb shift can be neglected because it is far less than ω_0 , i.e., $\Omega \approx \omega_0$. We have defined

$$\begin{split} \gamma_{\pm} &= 2\mu^2 \mathrm{Re}\Gamma_{\pm} = \mu^2 \int_{-\infty}^{+\infty} e^{\mp i\omega_0 \bigtriangleup \tau} G^+(\bigtriangleup \tau - i\epsilon) d\bigtriangleup \tau, \\ \gamma_z &= 0, \\ L_1 &= \sqrt{\frac{\gamma_-}{2}} \sigma_-, \qquad L_2 = \sqrt{\frac{\gamma_+}{2}} \sigma_+, \qquad L_3 = \sqrt{\frac{\gamma_z}{2}} \sigma_z, \quad (15) \end{split}$$

where $\triangle \tau = \tau - \tau'$. Here, $G^+(\triangle \tau)$ is given by $G^+(x - x') = \langle 0 | \Phi(x(\tau)) \Phi(x(\tau')) | 0 \rangle$ being the two-point correlation function, which for massless scalar field reads [65]

$$G^{+}(x - x') = \frac{1}{4\pi^{2}[(x - x')^{2} + (y - y')^{2} + (z - z')^{2} - (t - t' - i\epsilon)^{2}]},$$
(16)

where ϵ is an infinitesimal constant.

We take the initial state of the detector as

$$|\psi(0)
angle = \sinrac{ heta}{2}|0
angle + e^{-i\phi}\cosrac{ heta}{2}|1
angle,$$
 (17)

where θ and ϕ denote the initial weight parameter and phase parameter, and $|0\rangle$, $|1\rangle$ are the ground state and excited state of the detector, respectively. By substituting the density matrix $\rho(\tau)$ Eq. (3) into the master equation (13), the time dependent state parameters $\omega(\tau)$ in terms of the proper time τ , after a series of calculations, are found to be

$$\omega_{1}(\tau) = \omega_{1}(0) \cos(\Omega\tau) e^{-\frac{1}{2}A\tau} - \omega_{2}(0) \sin(\Omega\tau) e^{-\frac{1}{2}A\tau},
\omega_{2}(\tau) = \omega_{1}(0) \sin(\Omega\tau) e^{-\frac{1}{2}A\tau} + \omega_{2}(0) \cos(\Omega\tau) e^{-\frac{1}{2}A\tau},
\omega_{3}(\tau) = \omega_{3}(0) e^{-A\tau} + \frac{B}{A} (1 - e^{-A\tau}),$$
(18)

where $A = \gamma_+ + \gamma_-$, $B = \gamma_+ - \gamma_-$, and $\lim_{\tau \to 0} \omega_i(\tau) = \omega_i(0)$. Substituting the initial state (17) into Eq. (18), the general analytic solution of the evolution of two-level system then can be written as

$$\omega_{1}(\tau) = \sin\theta\cos(\Omega\tau + \phi)e^{-\frac{1}{2}A\tau},$$

$$\omega_{2}(\tau) = \sin\theta\sin(\Omega\tau + \phi)e^{-\frac{1}{2}A\tau},$$

$$\omega_{3}(\tau) = \cos\theta e^{-A\tau} + \frac{B}{A}(1 - e^{-A\tau}).$$
(19)

III. RELATIVISTIC MOTION AFFECTS ON PARAMETERS ESTIMATION

Now let us first calculate the QFI for the parameter estimation of the detector, moving along a spatially straight line with constant four-velocity component w, whose spacetime coordinates are given by [60]

$$t(\tau) = \sqrt{1 + w^2}\tau, \quad x(\tau) = 0, \quad y(\tau) = w\tau, \quad z(\tau) = 0.$$
(20)

Submitting Eqs. (16) and (20) into Eq. (15), the A and B can be calculated easily as

$$A = \gamma_0, \qquad B = -\gamma_0, \tag{21}$$

where $\gamma_0 = \frac{\mu^2 \omega_0}{2\pi}$ is the spontaneous emission rate. Thus, the Bloch vector of detector's state evolves with proper time can be obtained as

$$\omega_{1}(\tau) = \sin\theta\cos(\Omega\tau + \phi)e^{-\frac{1}{2}\gamma_{0}\tau},$$

$$\omega_{2}(\tau) = \sin\theta\sin(\Omega\tau + \phi)e^{-\frac{1}{2}\gamma_{0}\tau},$$

$$\omega_{3}(\tau) = \cos\theta e^{-\gamma_{0}\tau} - (1 - e^{-\gamma_{0}\tau}).$$
(22)

As a result, the QFI of the initial weight θ and phase parameter ϕ become $F_{\theta} = e^{-\gamma_0 \tau}$ and $F_{\phi} = \sin^2 \theta e^{-\gamma_0 \tau}$. It implies that the QFI of both weight and phase parameters decreases exponentially with time, due to the decoherence caused by the interaction between the detector and the massless scalar field. However, the QFI is completely unaffected by the four-velocity component.

However, in this paper we may wonder how the detector motion generated by a combination of the linear accelerated motion and a component of the four-velocity affects on the performance of QFI of parameters estimation. Now we consider the detector moving in flat spacetime along an unbounded spatial trajectory in a two-dimensional spatial plane with a constant square of magnitude of four-acceleration $a_{\mu}a^{\mu} = a^2$, and constant magnitudes of a timelike proper time derivative of four-acceleration $(da_{\mu}/d\tau)(da^{\mu}/d\tau)$, which has a constant component of the four-velocity $w = dy/d\tau = \text{const.}$ Here, $a^{\mu} = d^2 x^{\mu}/d\tau^2$. As a result, the spacetime coordinates of detector are described by [60,66]

$$t(\tau) = \frac{a}{\alpha^2} \sinh \alpha \tau, \qquad x(\tau) = \frac{a}{\alpha^2} \cosh \alpha \tau,$$

$$y(\tau) = w\tau, \qquad z(\tau) = 0, \qquad (23)$$

where $\alpha = \frac{a}{\sqrt{1+w^2}} > 0$. Applying the trajectory of detector (23) into Eq. (16), the two-point correlation function for the massless scalar field is given by

$$G^{+}(x-x') = -\frac{\alpha^4}{16\pi^2} \left[\sinh^2\left(\frac{\alpha\Delta\tau}{2} - \frac{i\epsilon\alpha^2}{a}\right) - \frac{w^2\alpha^4}{4a^2}\Delta\tau^2 \right]^{-1}.$$
(24)

Note that for w = 0, we have $\alpha = a$ and the two-point correlation function in Eq. (24) recovers to that of a detector moving along a spatially straight line along the *x* direction with constant magnitude of the four-acceleration (spatially one-dimensional) [27,28], as expected, which is

$$G^{+(0)}(\Delta\tau) = -\frac{a^2}{16\pi^2} \left[\sinh^2 \left(\frac{a\Delta\tau}{2} - i\epsilon a \right) \right]^{-1}.$$
 (25)

In the following, we are interested in investigating the QFI of this detector moving along such a trajectory in Eq. (23) in two situations: in the nonrelativistic and ultrarelativistic limit, respectively.

A. In the case for nonrelativistic velocities

Let us consider the case for the Unruh-DeWitt detector moving along such a trajectory in Eq. (23) in the

PHYS. REV. D 103, 125025 (2021)

nonrelativistic limit, i.e., $w \ll 1$, the two-point correlation function for the massless scalar field in Eq. (24) up to second order in w^2 which is

$$G^{+}(x - x')$$

$$= (1 - 2w^{2})G^{+(0)}(\Delta\tau)$$

$$- \frac{a^{2}}{16\pi^{2}} \left[\sinh(a\Delta\tau - 2i\epsilon a) \left(\frac{a\Delta\tau}{4} - i\epsilon a \right) + \frac{a^{2}\Delta\tau^{2}}{4} \right]$$

$$\times \left[\sinh^{4} \left(\frac{a\Delta\tau}{2} - \frac{i\epsilon\alpha^{2}}{a} \right) \right]^{-1} w^{2}.$$
(26)

Applying Eq. (26) to Eq. (15), and through the contour integral, we can obtain

$$\gamma_{\pm} = 2\mu^2 [\mathcal{G}^0(\pm\omega_0) + \mathcal{G}^1(\pm\omega_0)w^2]$$
(27)

with

$$\mathcal{G}^{0}(\omega_{0}) = \frac{\omega_{0}}{2\pi} \left(\frac{1}{e^{2\pi\omega_{0}/a} - 1} \right),$$

$$\mathcal{G}^{0}(-\omega_{0}) = \frac{\omega_{0}}{2\pi} \left(\frac{e^{2\pi\omega_{0}/a}}{e^{2\pi\omega_{0}/a} - 1} \right),$$

$$\mathcal{G}^{1}(\pm\omega_{0}) = -f(a),$$
(28)

where $f(a) = \frac{ae^{2\pi\omega_0/a}}{6[e^{2\pi\omega_0/a}-1]^2} \left[2 + \frac{9\omega_0^2}{a^2} - \frac{2\pi\omega_0}{a} \left(1 + \frac{\omega_0^2}{a^2}\right) \frac{e^{2\pi\omega_0/a}+1}{e^{2\pi\omega_0/a}-1}\right]$. Therefore, we have *A* and *B* in Eq. (19) which are

$$A = \gamma_0 \left[\frac{e^{2\pi\omega_0/a} + 1}{e^{2\pi\omega_0/a} - 1} - \frac{4\pi}{\omega_0} f(a) w^2 \right],$$

$$B = -\gamma_0.$$
(29)

1. Relativistic motion affects on the precision in the estimation of phase parameter ϕ

We discuss the relativistic motion of the detector in Eq. (23) how to affect the precision in the estimation of phase parameter ϕ in the nonrelativistic limit. For the sake of simplicity, in this paper we will work with dimensionless quantities by rescaling time τ and four-acceleration *a*

$$\tilde{\tau} \equiv \gamma_0 \tau, \qquad \tilde{a} \equiv \frac{a}{\omega_0}.$$
 (30)

Substituting Eqs. (19) and (29) into Eq. (4), one can easily obtain the detailed formula of the QFI with respect to ϕ as

$$F_{\phi} = \sin^2 \theta e^{-h(\tilde{a})\tilde{\tau}},\tag{31}$$

where $h(\tilde{a}) = \frac{e^{2\pi/\tilde{a}}+1}{e^{2\pi/\tilde{a}}-1} - 4\pi f(\tilde{a})w^2$. It is worth mentioning that $\gamma = h(\tilde{a})\gamma_0$ represents the decay rate for a two-level detector moving the trajectory (23) with a component of the four-velocity $w = dy/d\tau = \text{const.}$ Interesting, we notice



FIG. 1. (a) The QFI F_{ϕ} as a function of the initial weight parameter θ with different effective time $\tau = 1$ (solid line), $\tau = 2$ (dashed line), $\tau = 3$ (dot-dashed line); (b) The QFI F_{ϕ} as a function of the effective time τ with $\theta = \pi/2$ (solid line), $\theta = \pi/3$ (dashed line), $\theta = \pi/6$ (dot-dashed line). Here, we take the effective four-acceleration $a = \pi$ and four-velocity component w = 0.01.

that the QFI in Eq. (31) is irrespective of quantum phase ϕ , but depends on the value of initial weight parameter θ , time $\tilde{\tau}$, four-acceleration \tilde{a} and four-velocity component w. Hereafter, for convenience, we continue to term $\tilde{\tau}$ and \tilde{a} as τ and a, respectively, in this paper.

To show the properties of the precision of the phase parameter estimation, we plot, in Fig. 1, the QFI as the function of the initial state parameter θ (effective time τ) with different effective time τ (initial weight parameter θ). We are interested in finding that the maximal F_{ϕ} can be achieved by taking $\theta = \pi/2$, i.e., by preparing the two-level detector in the *balance-weighted state* which is preferable (see figure 1). From Fig. 1(a), we see that the QFI in the estimation of phase parameter ϕ is symmetric with respect to $\theta = \pi/2$. Moreover, in Fig. 1(b), the QFI F_{ϕ} takes the maximum value when the quantum state is at the beginning ($\tau = 0$), which implies that the precision in the estimation of phase parameter decreases with the effective evolution time, because the decoherence is caused by the interaction between the detector and massless scalar field.

As can be seen in Fig. 2, the QFI in the estimation of phase parameter ϕ is plotted as a function of the effective



FIG. 2. The QFI F_{ϕ} as a function of the effective fouracceleration parameter *a* with three different effective evolution time τ . We take $\theta = \pi/2$ and w = 0.01.

four-acceleration *a*. We observe that as the effective fouracceleration *a* increases, the QFI F_{ϕ} gradually decreases and converges to zero value in the limit of infinite fouracceleration, which is reminiscent of previous results that the quantum entanglement vanishes with infinite acceleration [67]. The reason is that the larger effective fouracceleration results in the larger decay rate of the atom which is modified by the factor $\left[\frac{e^{2\pi/a}+1}{e^{2\pi/a}-1}-4\pi f(a)w^2\right]$ comparing with the spontaneous emission rate, which implies that the precision of phase parameter is an decreasing function of the four-acceleration.

More remarkably, to analyze the four-velocity component for the accelerated Unruh-DeWitt detector in the nonrelativistic limit how to affect the precision in the estimation of phase parameter, we show the QFI F_{ϕ} as a function of the four-velocity component w in Fig. 3. We are interested in noting that the higher the four-velocity component w, the bigger the QFI is, i.e., it is easier to achieve a given precision of phase parameter estimation. As a result, we can infer that in the nonrelativistic limit, the four-velocity component can suppress the degradation of



FIG. 3. QFI in the estimation of phase parameter as a function of the four-velocity component *w*. By preparing the detector in the balance-weighted state $\theta = \pi/2$, the effective time $\tau = 1$ and the effective four-acceleration $a = \pi$.

the QFI, which means that the precision in the estimation of phase parameter ϕ is enhanced when the detector moves along such a stationary trajectory in Eq. (23).

2. Relativistic motion affects on the precision in the estimation of initial weight parameter θ

Then we want to examine the relativistic motion of the detector in Eq. (23) how to affect the precision in the estimation of the initial weight parameter θ in the non-relativistic limit. With the help of Eqs. (4), (19), and (29), we can evaluate the QFI in terms of θ which is

$$F_{\theta} = e^{-h(a)\tau} \left\{ \cos^{2}\theta + \sin^{2}\theta e^{-h(a)\tau} \times \left[1 - \frac{(1 - e^{h(a)\tau})^{2}[1 + h(a)\cos\theta]^{2}}{[h(a)^{2} - 1]e^{h(a)\tau} + [1 + h(a)\cos\theta]^{2}} \right] \right\},$$
(32)

where τ is the effective time, and the factor $h(a) = \frac{e^{2\pi/a}+1}{e^{2\pi/a}-1} - 4\pi f(a)w^2$ with *a* being the effective four-acceleration. Let us note that the QFI F_{θ} only depends on the initial weight



FIG. 4. (a) The QFI F_{θ} as a function of the initial weight parameter θ with fixed values of τ , i.e., $\tau = 1$ (solid line), $\tau = 2$ (dashed line), $\tau = 3$ (dot-dashed line); (b) The QFI F_{θ} as a function of the effective time τ with fixed values of θ , i.e., $\theta = 0$ (solid line), $\theta = \pi/3$ (dashed line), $\theta = \pi/2$ (dot-dashed line). Here, we take the effective four-acceleration $a = \pi$ and fourvelocity component w = 0.01.

parameter θ , effective time τ , effective four-acceleration *a* and four-velocity component *w*, but is independent of the phase parameter ϕ of the detector.

Similarly, to show the behaviors of the precision in the estimation of θ , in Fig. 4, we plot the QFI F_{θ} as a function of the initial weight parameter θ (effective time τ) with different effective time τ (initial weight parameter θ). As we can see from Fig. 4(a), the maximal QFI F_{θ} is obtained by taking $\theta = 0$. That is, the precision in the estimation of initial weight parameter can be achieved by preparing the detector in the excited state. In addition, we find the symmetry of the function of F_{θ} with respect to $\theta = 0$. It is obvious from Fig. 4(b) that the QFI F_{θ} decreases by increasing the value of effective time τ , which means that the precision in the estimation of initial weight parameter reduced by the decoherence of the detector which is caused by the interaction between the detector and field. Moreover, in Fig. 4(b), we note that the maximal value of the QFI is obtained initially, i.e., $F_{\theta} = 1$, which implies that the QFI F_{θ} is immune to the external environment at the beginning $(\tau = 0)$. This result is sharp contrast with the behavior of the QFI in the estimation of phase parameter ϕ shown in Fig. 1(b).

In Fig. 5, we plot the QFI of initial weight parameter in Eq. (32) as a function of the effective four-acceleration parameter *a* with different effective time τ at fixed w = 0.01 for $\theta = 0$. In a similar way, we find that as the effective four-acceleration parameter *a* gets larger values, the QFI F_{θ} decreases and reduces to zero in the limit of infinite four-acceleration, which means that the precision in the estimation of initial weight parameter decreases as the effective four-acceleration increases. This is due to the fact that the larger effective four-acceleration results in a larger decay rate.

To assess the performance of the four-velocity component for the accelerated Unruh-DeWitt detector in the nonrelativistic limit how to influence the precision in the estimation of initial weight parameter, Fig. 6 represents



FIG. 5. The QFI F_{θ} as a function of the effective fouracceleration parameter *a* with the effective time $\tau = 1$ (solid line), $\tau = 2$ (dashed line), $\tau = 3$ (dot-dashed line). We take $\theta = 0$ and w = 0.01.



FIG. 6. The QFI F_{θ} as a function of the four-velocity component *w*. Here, we take the detector in the exited state $\theta = 0$, the effective time $\tau = 1$ and the effective four-acceleration $a = \pi$.

the QFI F_{θ} as a function of the four-velocity component w. It is worthy noting from Fig. 6 that the QFI is increased as the growth of the four-velocity component w, which indicates that the highest precision in the estimation of initial weight parameter can be obtained for a larger fourvelocity component. Thus, we argue that when the detector follows such trajectory shown in Eq. (23) for nonrelativistic velocities, the quantum estimation of initial weight parameter can be enhanced by the four-velocity component w, i.e., such relativistic motion of detector in the nonrelativistic limit can provide us a better precision.

3. Relativistic motion affects on the precision in the estimation of parameter β

In this section, we want to explore the relativistic motion of the detector in Eq. (23) how to affect the precision in the estimation of parameter $\beta = 2\pi/a$ for the case of nonrelativistic velocities, comparing with the results in Refs. [27] which shown that the uniformly accelerated detector (w = 0) moving along a spatially straight line degrade the QFI. Similarly, substituting Eqs. (19) and (29) into Eq. (4), we can also get the detailed formula of the QFI F_{β} , which does not contain any information about phase parameter ϕ . It is needed to note that the expression is too long to exhibit here.

To clarify what value of initial weight parameter θ could allow better estimation, we plot, in Fig. 7, the QFI F_{β} as a function of the initial weight parameter θ (effective time τ) with different effective time τ (initial weight parameter θ). Here, we fix the parameter $\beta = 10$ which was also considered in Ref. [27]. It is interesting to note that in Fig. 3 of Ref. [27], the behavior of QFI of β for a uniformly accelerated detector (w = 0) is very similar with the behavior of QFI F_{β} in Fig. 7 of this paper. Therefore, we can deduce from the Fig. 7(a) that the QFI F_{β} reaches the maximum value at $\theta = \pi$, which implies that the maximum sensitivity in the predictions for the parameter β can be obtained by initially preparing the detector in its



FIG. 7. (a) The QFI F_{β} as a function of the initial weight parameter θ with different effective time $\tau = 10$ (solid line), $\tau =$ 5 (dashed line), $\tau = 1$ (dot-dashed line); (b) The QFI F_{β} as a function of the effective time τ with $\theta = \pi$ (solid line), $\theta = 2\pi/3$ (dashed line), $\theta = \pi/2$ (dot-dashed line). Here, we take the parameter $\beta = 10$ and four-velocity component w = 0.01.

ground state. Moreover, the symmetry of the QFI F_{β} with respect to $\theta = \pi$ shows in Fig. 7(a). Besides, Fig. 7(b) presents that the QFI F_{β} is a monotonically increasing function of effective time τ during the initial period. However, when the detector evolves for a long enough time, i.e., $\tau \gg \frac{1}{A}$ with 1/A being the time scale for atomic transition, whatever the initial state of detector is prepared in, the QFI achieves the maximum and equals to each other, which represents that the optimal precision in the estimation of the parameter β is completely unaffected by initial preparation of the detector if the effective time is long enough.

Furthermore, in Fig. 8, we plot the QFI of parameter β as a function of the effective time τ with different values of β at fixed w = 0.01 for $\theta = \pi$. We find that the QFI F_{β} saturates at different maximum values for different parameter β in the limit of infinite effective time. However, we can see that when $\beta = 1$, the QFI in the estimation of β will increases for a while and starts to decrease but converges to nonzero value for a long enough time. This is due to the fact that for different values of parameter β , i.e., by taking different values of the effective four-acceleration of detector, the conditions of the detector eventually approaching to



FIG. 8. The QFI F_{β} as a function of the effective time τ with different values of parameter β , i.e., $\beta = 1, 2, 3$. By preparing the detector initially in the ground state $\theta = \pi$ and four-velocity component w = 0.01.

the equilibrium state are different. We can also obtain the same results of Ref. [27], although it was not shown. This implies that for the small value $\beta = 1$, the optimal precision in the estimation of β can be obtained when the detector evolves for a finite time. Besides, we note that the smaller the parameter β , i.e., for larger value of the effective four-acceleration parameter *a*, the higher the QFI F_{β} is. That is, the highest precision in the estimation of parameter β can be obtained for a larger four-acceleration parameter.

Similarly, to analyze the four-velocity component for the accelerated two-level detector in the nonrelativistic limit how to affect the precision in the estimation of parameter β , we plot the QFI F_{β} as a function of the four-velocity component w in Fig. 9. We find that the higher the four-velocity component, the bigger the QFI is, i.e., the easier it is to achieve a given precision in the estimation of parameter β . In this respect, comparing the above analysis with the results in Refs. [27], we find that the relativistic motion of detector moving along such trajectory shown in Eq. (23) can inhibition the degradation of the QFI, which



FIG. 9. The QFI F_{β} as a function of the four-velocity component *w*. Here, we take the detector in the ground state $\theta = \pi$, the effective time $\tau = 1$ and the parameter $\beta = 1$.

implies that the precision in the estimation of parameter β can be enhanced.

B. In the case for ultrarelativistic velocities

Now we consider the detector moving along the trajectory in Eq. (23) for the case of ultrarelativistic velocities, i.e., $w \rightarrow \infty$, the two-point correlation function for the massless scalar field in Eq. (24) which is suppressed as

$$G^{+}(x - x') = -\frac{a^{2}}{16\pi^{2}} \left[\sinh^{2} \left(\frac{a\Delta\tau}{2} - i\epsilon a \right) \right]^{-1} \frac{1}{w^{4}}.$$
 (33)

According to Eqs. (15) and (33), and by invoking the contour integral, the A and B in Eq. (19) can be obtained as

$$A = \gamma_0 \left(\frac{e^{2\pi\omega_0/a} + 1}{e^{2\pi\omega_0/a} - 1} \right) \frac{1}{w^4}, \qquad B = -\gamma_0 \frac{1}{w^4}.$$
 (34)

Thus, when $w \to \infty$, the A and B in Eq. (34) for ultrarelativistic velocities are given by

$$A \to 0, \qquad B \to 0.$$
 (35)

This suggests that the detector evolves with time as a closed system, whose evolution is completely unaffected by the external environment and detector's motion. Because the trajectory of detector in Eq. (23) is modified by the factor $1/\sqrt{1+w^2}$, such trajectory becomes constant in the ultrarelativistic limit. Therefore, submitting Eq. (35) into Eq. (19), the Bloch vector of the state $\rho(\tau)$ with respect to the proper time τ can easily be written as

$$\omega_{1}(\tau) = \sin\theta\cos(\omega_{0}\tau + \phi),$$

$$\omega_{2}(\tau) = \sin\theta\sin(\omega_{0}\tau + \phi),$$

$$\omega_{3}(\tau) = \cos\theta.$$
(36)

With the help of Eqs. (4) and (36), one can obtain the QFI in terms of θ and ϕ as

$$F_{\theta} = 1, \qquad F_{\phi} = \sin^2 \theta, \tag{37}$$

which shows that the QFI is time independent.

Note that, in contrast to the previous results show that in the nonrelativistic limit the QFI will be enhanced with the increase of the four-velocity component w, the QFI in Eq. (37), for the case of ultrarelativistic velocities, will be never subjected to affected by the environment and remains constant with time, as if the detector were a closed system. More interestingly, it is worth emphasizing that the Unruh-DeWitt detector moving along such trajectory in the ultrarelativistic limit, has the same impact as the results of that a system interacted with environment by the presence of boundaries in certain circumstances [43,56–58], which indicates that the QFI can be shielded from the influence of the environment.

It is worth emphasizing that when the detector is in inertial motion, i.e., when the detector moving with a single constant velocity, although the QFI does not dependent on the detector's velocity, it is still degraded exponentially with the evolution time as a result of the interaction between the detector and field. Therefore, the precision of the parameters estimation is decreased. Moreover, for an uniformly accelerated detector, the decrease of QFI over time would be enhanced by the acceleration, as shown in Refs. [42,43]. However, when the detector moves with a combination of the linear accelerated motion and a component of the four-velocity $w = dy/d\tau$, we find the QFI depends on both the velocity and acceleration, and the velocity will suppress the degradation of OFI compared with the individual acceleration case in the nonrelativistic limit. What is more, in the ultrarelativistic limit, the QFI may be shielded from the effects of the detector's motion, and even remains constant with time as if it were a closed system. This intriguing behaviors come from the composite effect of both velocity and acceleration, and are not valid for the individual accelerated case and individual constantvelocity case.

IV. CONCLUSIONS

In the framework of open quantum systems, we studied the dynamics of the QFI of the parameters estimation for a detector interacted with massless scalar field. For the detector moving along a spatially straight line with a constant velocity, we found that the QFI always will be degraded by the external environment, but unaffected by the velocity. Besides, for the uniformly linear accelerated detector (w = 0), the acceleration will cause the QFI and thus the precision limit of parameter estimation to degrade, as shown in Refs. [42,43].

However, when the detector moving along an unbounded spatial trajectory in a two-dimensional spatial plane with constant independent magnitudes of both the four-acceleration a and also having a component of four-velocity $w = dy/d\tau$ constant, the QFI of this detector in two different situations, in the nonrelativistic and ultrarelativistic limit, have been considered in detail. In the nonrelativistic limit, we can achieve the optimal strategy for the parameters estimation by preparing the proper probe and adjusting the interaction parameters. Moreover, the four-velocity component will suppress the degradation of the QFI. That is, the precision of the parameters estimation can be enhanced by the relativistic motion of the detector following trajectory (23) for nonrelativistic velocities. In the ultrarelativistic limit, counterintuitively, the QFI remains constant with time due to the relativity correction to the four-velocity component, which implies that the precision of the parameters estimation can be completely

ACKNOWLEDGMENTS

This work was supported by the National Natural Science Foundation of China under Grants No. 12065016, No. 11875025, and No. 11905218; Hunan Provincial Natural Science Foundation of China under Grant No. 2018JJ1016; The CAS Key Laboratory for Research in Galaxies and Cosmology, Chinese Academy of

- S. L. Braunstein and C. M. Caves, Phys. Rev. Lett. 72, 3439 (1994).
- [2] S. L. Braunstein, C. M. Caves, and G. J. Milburn, Ann. Phys. (N.Y.) 247, 135 (1996).
- [3] D. Petz and C. Ghinea, in *QP-PQ: Quantum Probab. White Noise Anal.* (World Scientific, Singapore, 2011), Vol. 27, pp. 261–281.
- [4] D. Petz, Linear Algebra Appl. 244, 81 (1996).
- [5] L. Pezz and A. Smerzi, Phys. Rev. Lett. 102, 100401 (2009).
- [6] J. Ma and X. G. Wang, Phys. Rev. A 80, 012318 (2009).
- [7] Á. Rivas and A. Luis, Phys. Rev. Lett. 105, 010403 (2010).
- [8] Z. Sun, J. Ma, X. M. Lu, and X. G. Wang, Phys. Rev. A 82, 022306 (2010).
- [9] R. Chaves, L. Aolita, and A. Acín, Phys. Rev. A 86, 020301
 (R) (2012).
- [10] J. Ma, X. G. Wang, C. P. Sun, and F. Nori, Phys. Rep. 509, 89 (2011).
- [11] P. Hyllus, W. Laskowski, R. Krischek, C. Schwemmer, W. Wieczorek, H. Weinfurter, L. Pezzé, and A. Smerzi, Phys. Rev. A 85, 022321 (2012).
- [12] G. Tóth, Phys. Rev. A 85, 022322 (2012).
- [13] V. Giovannetti, S. Lloyd, and L. Maccone, Science 306, 1330 (2004).
- [14] V. Giovannetti, S. Lloyd, and L. Maccone, Phys. Rev. Lett. 96, 010401 (2006).
- [15] V. Giovannetti, S. Lloyd, and L. Maccone, Nat. Photonics 5, 222 (2011).
- [16] C. W. Helstrom, *Quantum Detection and Estimation Theory* (Academic Press, New York, 1976).
- [17] A. S. Holevo, *Probabilistic and Statistical Aspects of Quantum Theory* (North-Holland, Amsterdam, 1982).
- [18] M. Hübner, Phys. Lett. A 163, 239 (1992); 179, 226 (1993).
- [19] X.-M. Lu, X. Wang, and C. P. Sun, Phys. Rev. A 82, 042103 (2010).
- [20] N. Li and S. Luo, Phys. Rev. A 88, 014301 (2013).
- [21] M. Brunelli, S. Olivares, and M. G. A. Paris, Phys. Rev. A 84, 032105 (2011).
- [22] M. Brunelli, S. Olivares, M. Paternostro, and M. G. A. Paris, Phys. Rev. A 86, 012125 (2012).
- [23] M. Aspachs, G. Adesso, and I. Fuentes, Phys. Rev. Lett. 105, 151301 (2010).

Science (No. 18010203); The Physics Key Discipline of Liupanshui Normal University under Grant No. LPSSYZDXK201801; The cultivation project of Master's degree of Liupanshui Normal University under Grant No. LPSSYSSDPY201704. X. Liu thanks for the Young scientific talents growth project of the department of education of the department of education of Guizhou province under Grant No. QJHKYZ[2019]129; The talent recruitment program of Liupanshui normal university of China under Grant No. LPSSYKYJJ201906.

- [24] M. Ahmadi, D. E. Bruschi, C. Sabín, G. Adesso, and I. Fuentes, Sci. Rep. 4, 4996 (2015).
- [25] M. Ahmadi, D. E. Bruschi, and I. Fuentes, Phys. Rev. D 89, 065028 (2014).
- [26] D. E. Bruschi, A. Datta, R. Ursin, T. C. Ralph, and I. Fuentes, Phys. Rev. D 90, 124001 (2014).
- [27] Z. Tian, J. Wang, J. Jing, and H. Fan, Sci. Rep. 5, 7946 (2015).
- [28] J. Wang, Z. Tian, J. Jing, and H. Fan, Sci. Rep. 4, 7195 (2015).
- [29] S. P. Kish and T. C. Ralph, Phys. Rev. D 93, 105013 (2016).
- [30] J. Kohlrus, D. E. Bruschi, J. Louko, and I. Fuentes, Eur. Phys. J. Quantum Technol. 4, 7 (2017).
- [31] Z. Zhao, Q. Pan, and J. Jing, Phys. Rev. D **101**, 056014 (2020).
- [32] X. Liu, J. Jing, J. Wang, and Z. Tian, Quantum Inf. Process. 19, 26 (2020).
- [33] H. Du and R. B. Mann, J. High Energy Phys. 05 (2021) 112.
- [34] S. F. Huelga, C. Macchiavello, T. Pellizzari, A. K. Ekert, M. B. Plenio, and J. I. Cirac, Phys. Rev. Lett. 79, 3865 (1997).
- [35] M. Rosenkranz and D. Jaksch, Phys. Rev. A 79, 022103 (2009).
- [36] B. M. Escher, R. L. de MatosFilho, and L. Davidovich, Nat. Phys. 7, 406 (2011).
- [37] A. W. Chin, S. F. Huelga, and M. B. Plenio, Phys. Rev. Lett. 109, 233601 (2012).
- [38] R. Chaves, J. B. Brask, M. Markiewicz, J. Kołodyński, and A. Acín, Phys. Rev. Lett. **111**, 120401 (2013).
- [39] J. Ma, Y. X. Huang, X. Wang, and C. P. Sun, Phys. Rev. A 84, 022302 (2011).
- [40] W. Zhong, Z. Sun, J. Ma, X. Wang, and F. Nori, Phys. Rev. A 87, 022337 (2013).
- [41] Y. Yao, X. Xiao, L. Ge, X. G. Wang, and C. P. Sun, Phys. Rev. A 89, 042336 (2014).
- [42] X. Hao and Y. Wu, Ann. Phys. (Amsterdam) 372, 110 (2016).
- [43] Y. Yang, X. Liu, J. Wang, and J. Jing, Quantum Inf. Process. 17, 54 (2018).
- [44] Z. Huang, Eur. Phys. J. Plus 133, 101 (2018).
- [45] Y. Yang, J. Wang, M. Wang, J. Jing, and Zehua Tian, Classical Quantum Gravity 37, 065017 (2020).

- [46] W.G. Unruh and R.M. Wald, Phys. Rev. D 29, 1047 (1984).
- [47] J. I. Korsbakkena and J. M. Leinaas, Phys. Rev. D 70, 084016 (2004).
- [48] L. C. Barbado and M. Visser, Phys. Rev. D 86, 084011 (2012).
- [49] H. Yu and S. Lu, Phys. Rev. D 72, 064022 (2005).
- [50] J. Hu and H. Yu, Phys. Rev. A 85, 032105 (2012).
- [51] J. Hu and H. Yu, Phys. Rev. A 91, 012327 (2015).
- [52] Y. Jin, J. Hu, and H. Yu, Phys. Rev. A 89, 064101 (2014).
- [53] X. Liu, Z. Tian, J. Wang, and J. Jing, Eur. Phys. J. C 78, 665 (2018).
- [54] Y. Yang, J. Jing, and Z. Zhao, Quantum Inf. Process. 18, 120 (2019).
- [55] J. Zhang and H. Yu, Phys. Rev. D 102, 065013 (2020).
- [56] Y. Jin and H. Yu, Phys. Rev. A 91, 022120 (2015).
- [57] X. Liu, Z. Tian, J. Wang, and J. Jing, Ann. Phys. (Amsterdam) 366, 102 (2016).

- [58] X. Liu, Z. Tian, J. Wang, and J. Jing, Quantum Inf. Process. 15, 3677 (2016).
- [59] Z. Huang and W. Zhang, Braz. J. Phys. 49, 161 (2019).
- [60] S. Abdolrahimi, Classical Quantum Gravity **31**, 135009 (2014).
- [61] H.-P. Breuer and F. Petruccione, *The Theory of Open Quantum Systems* (Oxford University Press, Oxford, 2002).
- [62] V. Gorini, A. Kossakowski, and E. C. G. Surdarshan, J. Math. Phys. (N.Y.) 17, 821 (1976).
- [63] G. Lindblad, Commun. Math. Phys. 48, 119 (1976).
- [64] F. Benatti, R. Floreanini, and M. Piani, Phys. Rev. Lett. 91, 070402 (2003).
- [65] N. D. Birrell and P. C. W. Davies, *Quantum Fields in Curved Space* (Cambridge University Press, Cambridge, England, 1982).
- [66] J. R. Letaw, Phys. Rev. D 23, 1709 (1981).
- [67] I. Fuentes-Schuller and R. B. Mann, Phys. Rev. Lett. 95, 120404 (2005).