# Phenomenology of $J^{P C}=3^{--}$tensor mesons 

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(Received 19 January 2021; accepted 15 April 2021; published 28 May 2021)


#### Abstract

We study the strong and radiative decays of the antiquark-quark ground state $J^{P C}=3^{--}$ $\left(n^{2 S+1} L_{J}=1^{3} D_{3}\right)$ nonet $\left\{\rho_{3}(1690), K_{3}^{*}(1780), \phi_{3}(1850), \omega_{3}(1670)\right\}$ in the framework of an effective quantum field theory approach, based on the $S U_{\mathrm{V}}(3)$-flavor symmetry. The effective model is fitted to experimental data listed by the Particle Data Group. We predict numerous experimentally unknown decay widths and branching ratios. An overall agreement of theory (fit and predictions) with experimental data confirms the $\bar{q} q$ nature of the states and qualitatively validates the effective approach. Naturally, experimental clarification as well as advanced theoretical description is needed for trustworthy quantitative predictions, which is observed from some of the decay channels. Besides conventional spin- 3 mesons, theoretical predictions for ratios of strong and radiative decays of a hypothetical glueball state $G_{3}(4200)$ with $J^{P C}=3^{--}$are also presented.


DOI: 10.1103/PhysRevD.103.096027

## I. INTRODUCTION

The spectroscopy and the phenomenological description of conventional mesons is important for many reasons. It allows to test quantum chromodynamics (QCD) in the nonperturbative regime. Furthermore, it aims to the correct and systematic understanding and assignment of experimentally measured states, resonances, and mesons [1]. It is also fundamental for the search of mesons that go beyond the antiquark-quark $(\bar{q} q)$ picture, such as glueballs, hybrids, and multiquark states. Namely, both in the light-quark and heavyquark sectors, it is possible to search for "exotics" only, if the conventional $\bar{q} q$ mesons are fully under control [2].

Consequently, our work is just an incremental contribution to this extensive project in high-energy physics. In our work, we provide a (qualitative) analysis of the
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phenomenology of the spin-3 mesons $\rho_{3}(1690), K_{3}^{*}(1780)$, $\phi_{3}(1850), \omega_{3}(1670)$ with $J^{P C}=3^{--}\left(n^{2 S+1} L_{J}=1^{3} D_{3}\right)$ as well as the potential spin-3 tensor glueball $G_{3}(4200)$ [3-5]. This involves a fit of our effective model to experimental data as well as predictions for decay widths and ratios of the decays of the spin-3 mesons. Although we provide explicit values for the decay widths and branching ratios, these should be considered only as first estimates and mainly interpreted as qualitative results. Hence, this study may for example help in identifying dominant decay channels etc.

## A. Contextualization

In the low-energy sector of QCD, conventional mesons contain the light up $(u)$, down $(d)$, and strange $(s)$ quarks, hence they can be grouped into nonets: $\bar{u} d, \bar{u} s$, etc. Nonets are classified by the quantum numbers $J^{P C}$, where $J$ is the absolute value of the total spin of the meson ( $|L-S| \leq J \leq|L+S|$, where $L$ is the angular momentum and $S$ the spin of the system), $P$ is the sign-change under parity transformations, and $C$ the sign-change under charge conjugation. For $\bar{q} q$ states, besides $J^{P C}$, also the older spectroscopic notation $n^{2 S+1} L_{J}$, with $L=S, P, D, \ldots$ and with $n=1,2, \ldots$, which is the radial quantum number, is used. In the following, we shall restrict to the radial ground
state nonets with $n=1$. For a $\bar{q} q$ system, where $P=(-1)^{L+1}$ and $C=(-1)^{L+S}$, notice that $L$ and $S$ are, strictly speaking, not well-defined quantum numbers in a relativistic setup; hence, the mixing of states with the same $J^{P C}$ but with different $L$ and $S$ is possible-even if it is oftentimes a subordinate effect. However, this nomenclature for the states is still well upheld also in presence of mixing for practical purposes.

In the low-energy regime of QCD the "classical" symmetries of QCD at high-energy scales still play a crucial role-more precisely the spontaneous, anomalous, or explicit breaking of these symmetries. Quickly recapitulated, the situation is as follows: gluons are "democratic": they interact equally strongly with rightand left-handed quarks implying that in the so-called chiral limit (i.e., assuming negligibly small bare quark masses $m_{u}$, $m_{d}$, and $m_{s}$ ) the QCD Lagrangian is invariant under $U_{\mathrm{L}}(3) \times U_{\mathrm{R}}(3)$ transformations. However, this symmetry of the classical action is broken due to quantum effects, which is called the $U_{\mathrm{A}}(1)$ anomaly [6-9]. Still, the baryon number is conserved, which corresponds to the unaffected $U_{\mathrm{V}}(1)$ symmetry. The remaining $S U_{\mathrm{L}}(3) \times S U_{\mathrm{R}}(3)$ symmetry is spontaneously broken at low energies into $S U_{\mathrm{V}}(3)$ $[10,11]$. The $S U_{\mathrm{V}}(3)$ amounts to a rotation in the flavor space spanned by $u, d$, and $s$ and is consequently called "flavor symmetry" or "vector symmetry." Hence, at the composite mesonic level, chiral partner-mesons, which are linked through a chiral transformation, are not degenerate anymore. After all, the residual realization of an approximate $S U_{\mathrm{V}}(3)$, which is only explicitly broken by the bare quark masses, is the reason why nonets of $\bar{q} q$ states can be used to classify the low-energy spectrum of QCD [12-14].

Typical examples of very well known mesonic nonets are the pseudoscalar states,

$$
\begin{equation*}
\left\{\pi, K, \eta^{\prime}(958), \eta\right\} \tag{1}
\end{equation*}
$$

with $J^{P C}=0^{-+}\left(n^{2 S+1} L_{J}=1{ }^{1} S_{0}\right)$, the vector mesons

$$
\begin{equation*}
\left\{\rho(770), K^{*}(892), \phi(1020), \omega(782)\right\} \tag{2}
\end{equation*}
$$

with $J^{P C}=1^{--} \quad\left(n^{2 S+1} L_{J}=1^{3} S_{1}\right)$, the pseudovector mesons,

$$
\begin{equation*}
\left\{b_{1}(1235), K_{1, B}, h_{1}(1415), h_{1}(1170)\right\} \tag{3}
\end{equation*}
$$

with $J^{P C}=1^{+-}\left(n^{2 S+1} L_{J}=1{ }^{1} P_{1}\right)$, the axial-vector meson,

$$
\begin{equation*}
\left\{a_{1}(1260), K_{1, A}, f_{1}(1420), f_{1}(1285)\right\}, \tag{4}
\end{equation*}
$$

with $J^{P C}=1^{++}\left(n^{2 S+1} L_{J}=1^{3} P_{1}\right)$, and the well-established tensor mesons,

$$
\begin{equation*}
\left\{a_{2}(1320), K_{2}^{*}(1430), f_{2}^{\prime}(1525), f_{2}(1270)\right\} \tag{5}
\end{equation*}
$$

with $J^{P C}=2^{++}\left(n^{2 S+1} L_{J}=1^{3} P_{2}\right)$, compare with Ref. [15] or the quark model review by the Particle Data Group (PDG) [1].

The aim of this work is to study the $\bar{q} q$ nonet with quantum numbers $J^{P C}=3^{--}$, resulting from $L=2$ and $S=1$, therefore $n^{2 S+1} L_{J}=1^{3} D_{3}$ states, in a relativistic effective quantum field theory (QFT) model based on $S U_{\mathrm{V}}(3)$-flavor symmetry. The mesons belonging to this nonet, corresponding to the lightest states with the proper quantum numbers in the PDG [1], are

$$
\begin{equation*}
\left\{\rho_{3}(1690), K_{3}^{*}(1780), \phi_{3}(1850), \omega_{3}(1670)\right\} \tag{6}
\end{equation*}
$$

The study of these states is very interesting for a series of reasons (for previous studies on related subjects see also Refs. [15-22] and references therein):
(1) It is the only well-known mesonic ground-state nonet with $J>2$, which is confirmed by several experiments [1].
(2) Many decay channels are known [1], which allows for further theoretical and experimental tests of the assignment.
(3) Predictions for not-yet measured strong and radiative decay rates are possible.
(4) The spin-3 ground-state nonet (6) is also measured in lattice QCD calculations [22,23].
This nonet is therefore tailor made for an effective QFT study of decays. In particular, we shall answer whether and to what extend the $\bar{q} q$ assignment works, we can test validity of flavor symmetry for various decay channels, make various "postdictions" and-most interestinglypredictions for many decay channels. In view of the ongoing experimental efforts in hadron physics at GlueX [24-26] and CLAS12 [27] at Jefferson Lab, at COMPASS and LHCb at CERN [28,29], at BESIII in Beijing [30,31], and at the future PANDA experiment [32] at the GSI facility, we consider a revival of interest on such resonances valuable.

As an additional application of our approach, we also study some decays of the $J^{P C}=3^{--}$glueball state. This can be done by a simple modification of the action of spin-3 mesons and by using the mass of approximately 4.2 GeV found in lattice simulations [3-5] (in the quenched approximation) as an input. Due to the flavor-singlet nature of the glueball, only few decay channels are possible.

## B. Method

From a technical point of view, we construct $S U_{\mathrm{V}}(3)$ invariant effective actions/Lagrangians that involve the mesonic nonet (6) as well as its various decay products consisting in the well-established $\bar{q} q$ nonets that were previously introduced. From a (functional) renormalization group (FRG) perspective, the effective actions/Lagrangians can be interpreted as residual infrared (IR) effective actions with coupling constants that already involve all quantum
effects from higher energy scales: the ultraviolet (UV). These coupling constants are determined via fits to experimental data, rather than by ab initio QCD calculations. All calculations for the decays are therefore performed at tree level. This approach was already implemented in earlier works for studies of various light mesonic nonets, such as tensor mesons [33], axial-vector, and pseudovector mesons [34], the scalar mesons [35,36], the pseudotensor mesons [37], and the orbitally and radially excited vector mesons [38].

## C. Organization and structure

We organize the paper as follows: in Sec. II, after a discussion of our low-energy effective model, we present quark-anti-quark nonets, their transformation rules, and the corresponding Lagrangian of the model. It consists of seven interaction terms. In Sec. III we show the results for the decay widths and branching ratios of spin-3 mesons, while in Sec. IV we list some branching ratios for an hypothetical heavy $3^{--}$glueball. A discussion of the validity of the employed effective interaction terms is reported in Sec. V and conclusions can be found in Sec. VI. Many technical aspects concerning the QFT treatment of $J=3$ fields can be found in the numerous and detailed appendices.

## II. EFFECTIVE MODEL FOR $J^{P C}=3^{--}$MESONS

As mentioned in the introduction, in the chiral limit $\left(m_{u, d, s}=0\right)$ the $S U_{\mathrm{L}}(3) \times S U_{\mathrm{R}}(3) \times U_{\mathrm{V}}(1)$ symmetry of QCD is spontaneously broken into flavor symmetry $S U_{\mathrm{V}}(3)$ and baryon-number conservation-the $U_{\mathrm{V}}(1)$ symmetry. One of the first successful models that describes this process using four-Fermi interactions is given by Refs. [39,40]. At the level of confined light hadrons, this symmetry is evident: quark-antiquark mesons are clearly grouped into nonets, some of which were already listed in the Introduction [12-14]. One possibility to describe lowenergy QCD makes use of effective low-energy hadronic models. If the degrees of freedoms are only hadrons, then confinement and color neutrality is automatically built in. Typically, these models are defined by a proper action that mimics the chiral symmetry of QCD and its spontaneous and explicit breaking.

The chiral symmetry is for example the guiding principle of chiral perturbation theory (CPT) [41-51], in which chiral symmetry is nonlinearly realized, as well as for linear $\sigma$ models [52-57], in which it is linearly realized. In effective models the nonets usually appear in pairs of chiral partners. The breaking of chiral symmetry generates a mass difference between them. This dynamics can also be studied at nonzero temperatures and/or nonzero baryon or isospin chemical potentials. Within the last years, the Functional Renormalization Group [58-63], turned out to be one particularly interesting framework to describe chiral symmetry breaking (and phase transitions) in these kinds of models, see for example Refs. [64-75].

Hadronic models in which only the residual flavor symmetry $S U_{\mathrm{V}}(3)$ is explicitly conserved have also been constructed for a variety of nonets in a series of publications by one of the authors and collaborators [33-38]. These models can be interpreted as the effective emerging terms of chiral models after spontaneous symmetry breaking is worked out. When the effective action is written down, only flavor symmetry is retained and an expansion in dominant and subdominant terms in $1 / N_{\mathrm{c}}$ is carried out. Moreover, terms that break explicitly flavor symmetry, either because of the underlying breaking due to nonzero and unequal quark masses ( $m_{u} \approx 2 \mathrm{MeV}, m_{d} \approx 5 \mathrm{MeV}$, and $m_{s} \approx 93 \mathrm{MeV}$ ) or due to the chiral anomaly, can be included.

In this work, we study the decays of the $\bar{q} q$ ground-state $J^{P C}=3^{--}$nonet by constructing such a model. We shall consider only the dominant terms in the large- $N_{\mathrm{c}}$ expansion [76-78] and neglect flavor-symmetry breaking corrections, since the present level of data accuracy does not allow for their investigations.

## A. Particle content

## 1. Symmetries

Mesonic nonets that transform under the adjoint representation of the approximate flavor symmetry $S U_{\mathrm{V}}(3)$ are the main ingredients of the current work. Transformation rules under parity, charge conjugation and flavor transformations of the mesonic nonets of our model are summarized ${ }^{1}$ in the Table I. On the other hand, the mesonic nonets are in direct correspondence to the physical states, which is listed in Table II.

Each $\bar{q} q$ nonet can be assigned to a certain microscopic antiquark-quark current $N_{i j} \equiv\left(\bar{q}_{j} \Gamma q_{i}\right) / \sqrt{2}$, where $\Gamma$ is a combination of Dirac matrices and derivatives, which-in the nonrelativistic limit-reduces to the corresponding nonrelativistic configuration with the desired $L$ and $S$.

## 2. Mesonic nonets in the model

Pseudoscalar mesons.-The first nonet of Table I describes the matrix of pseudoscalar mesons $P$ with quantum numbers $L=S=0$, leading to $J^{P C}=0^{-+}$ $\left(n^{2 S+1} L_{J}=1{ }^{1} S_{0}\right)$. The elements are $P_{i j} \equiv\left(\bar{q}_{j} i \gamma^{5} q_{i}\right) / \sqrt{2}$ and the mesons are $\left\{\pi, K, \eta^{\prime}(958), \eta\right\}$, where the first entry represents the isospin $I=1$ triplet (the three pions), the second entry the two $I=1 / 2$ isodoublets (the four kaons), and the last two entries the two isoscalar states (the etas). Pseudoscalars form the basis of almost all low-energy effective hadronic models/limits or theories of QCD, e.g., $C P T$ [41-50] and $\sigma$ models [52-55,57] (for an link between $C P T$ and $\sigma$ models, see Refs. [68,69,71,82]). Explicitly, the matrix $P$ reads

[^0]TABLE I. Transformation properties of the pseudoscalar $P$ (7), the vector $V_{1}(9)$, the pseudovector $B_{1}(11)$, the axial-vector $A_{1}$ (15), the rank-2 tensor $A_{2}$ (13), and rank-3 tensor $W_{3}$ (20) nonets under parity transformations $P$, charge conjugation $C$, and $S U_{\mathrm{V}}(3)$-flavor transformations $U$. Notice the position of the Lorentz indices for parity transformations, since spatial and timelike indices do not transform identically. We use the Minkowski metric in the $\left(\eta_{\mu \nu}\right)=\operatorname{diag}(+1,-1,-1,-1)$ convention.

| Nonet | Parity | Charge <br> conjugation | Flavor |
| :--- | :---: | :---: | :---: |
| $J^{P C}=F_{J}^{\cdots}$ | $P$ | $C$ | $S U_{\mathrm{V}}(3)$ |
| $0^{-+}=P$ | $-P(t,-\vec{x})$ | $(P)^{T}$ | $U P U^{\dagger}$ |
| $1^{--}=V_{1}^{\mu}$ | $V_{1, \mu}(t,-\vec{x})$ | $-\left(V_{1}^{\mu}\right)^{T}$ | $U V_{1}^{\mu} U^{\dagger}$ |
| $1^{+-}=B_{1}^{\mu}$ | $-B_{1, \mu}(t,-\vec{x})$ | $-\left(B_{\mu}^{\mu}\right)^{T}$ | $U B_{1}^{\mu} U^{\dagger}$ |
| $1^{++}=A_{1}^{\mu}$ | $-A_{1, \mu}(t,-\vec{x})$ | $\left(A_{1}^{\mu}\right)^{T}$ | $U A_{1}^{\mu} U^{\dagger}$ |
| $2^{++}=A_{2}^{\mu \nu}$ | $A_{2, \mu \nu}(t,-\vec{x})$ | $\left(A_{2}^{\mu \nu}\right)^{T}$ | $U A_{2}^{\mu \nu} U^{\dagger}$ |
| $3^{--}=W_{3}^{\mu \nu \rho}$ | $W_{3, \mu \nu \rho}(t,-\vec{x})$ | $-\left(W_{3}^{\mu \nu \rho}\right)^{T}$ | $U W_{3}^{\mu \nu \rho} U^{\dagger}$ |

$$
P=\frac{1}{\sqrt{2}}\left(\begin{array}{ccc}
\frac{\eta_{N}+\pi^{0}}{\sqrt{2}} & \pi^{+} & K^{+}  \tag{7}\\
\pi^{-} & \frac{\eta_{N}-\pi^{0}}{\sqrt{2}} & K^{0} \\
K^{-} & \bar{K}^{0} & \eta_{S}
\end{array}\right)
$$

where $\eta_{N} \equiv\left(\bar{u} \mathrm{i} \gamma^{5} u+\bar{d} \mathrm{i} \gamma^{5} d\right) / \sqrt{2}$ stands for the purely nonstrange state and $\eta_{S} \equiv \bar{s} i \gamma^{5} s$ stands for the pure strange state. For the pions and kaons, the physical states are directly assigned to the fields in the model. In the isoscalar sector, physical and model fields are related by mixing ${ }^{2}$

$$
\binom{\eta}{\eta^{\prime}(958)}=\left(\begin{array}{cc}
\cos \beta_{p} & \sin \beta_{p}  \tag{8}\\
-\sin \beta_{p} & \cos \beta_{p}
\end{array}\right)\binom{\eta_{N}}{\eta_{S}}
$$

Here we shall use $\beta_{p}=-43.4^{\circ}$ obtained in Ref. [80]. Note that the rather large mixing angle and the unexpectedly high mass of the $\eta^{\prime}(958)$ result from the chiral (or axial) anomaly $U_{\mathrm{A}}(1)[83,84]$. According to a classifications introduced by two of the authors of the present work in Ref. [81], pseudoscalar mesons, together with their chiral partners (the scalars), belong to what we call a "heterochiral multiplet," which allows for the constructions of chirally anomalous mass and interaction terms.

Vector mesons.-Next, the second entry in Table I refers to the $J^{P C}=1^{--}\left(n^{2 S+1} L_{J}=1^{3} S_{1}\right)$ nonet with $L=0$ and $S=1$. These are the very well-known vector states

[^1]TABLE II. Assignment of physical resonances to $\bar{q} q$ states in the model.
$\left.\begin{array}{cccc}\hline \hline & \begin{array}{c}\text { Physical } \\ \text { resonances }\end{array} & \text { Nonet } \bar{q} q \text { states }\end{array} \quad \begin{array}{c}\text { Mixing } \\ \text { angles }\end{array}\right]$
${ }^{\dagger}$ The isospin $1 / 2$ states in the pseudovector and axial-vector sectors are strong mixtures of $K_{1}(1270)$ and $K_{1}(1400)$, see Ref. [1] and references therein.
$\left\{\rho(770), K^{*}(892), \omega(782), \phi(1020)\right\}$, which are for example also included in certain extensions of $C P T$ [47-50] as well as in enlarged (and realistic) versions of hadronic models [52-56]. The matrix $V_{1}^{\mu}$ with elements $V_{1, i j}^{\mu}=$ $\left(\bar{q}_{j} \gamma^{\mu} q_{i}\right) / \sqrt{2}$ has the form

$$
V_{1}^{\mu}=\frac{1}{\sqrt{2}}\left(\begin{array}{ccc}
\frac{\omega_{1, N}^{\mu}+\rho_{1}^{0 \mu}}{\sqrt{2}} & \rho_{1}^{+\mu} & K_{1}^{*+\mu}  \tag{9}\\
\rho_{1}^{-\mu} & \frac{\omega_{1, N}^{\mu}-\rho_{1}^{0 \mu}}{\sqrt{2}} & K_{1}^{* 0 \mu} \\
K_{1}^{*-\mu} & \bar{K}_{1}^{* 0 \mu} & \omega_{1, S}^{\mu}
\end{array}\right)
$$

where $\omega_{N}$ and $\omega_{S}$ are purely nonstrange and strange states, respectively. Similarly as before, the physical fields arise upon mixing

$$
\binom{\omega(782)}{\phi(1020)}=\left(\begin{array}{cc}
\cos \beta_{v_{1}} & \sin \beta_{v_{1}}  \tag{10}\\
-\sin \beta_{v_{1}} & \cos \beta_{v_{1}}
\end{array}\right)\binom{\omega_{1, N}}{\omega_{1, S}}
$$

where the very small isoscalar-vector mixing angle $\beta_{v_{1}}=$ $-3.9^{\circ}$ is taken from the PDG [1]. Hence, the physical states $\omega(782)$ and $\phi(1020)$ are dominated by nonstrange and strange components, respectively. This is in agreement with the "homochiral" nature of these states [81]. In fact vector and axial-vector mesons form a "homochiral multiplet," for which the effect of the chiral anomaly is of subleading order and can safely be ignored.

Pseudovector mesons.-Next, we consider $L=1$. The choice $L=1$ and $S=0$, leading to $J^{P C}=1^{+-}$ $\left(n^{2 S+1} L_{J}=1^{1} P_{1}\right)$, contains the established pseudovector mesons $\left\{b_{1}(1235), K_{1, B}, h_{1}(1170), h_{1}(1415)\right\}$. The corresponding nonet with elements $B_{1, i j}^{\mu}=\left(\bar{q}_{j} \gamma^{5} \partial^{\mu} q_{i}\right) / \sqrt{2}$ is

$$
B_{1}^{\mu}=\frac{1}{\sqrt{2}}\left(\begin{array}{ccc}
\frac{h_{1, N}^{\mu}+b_{1}^{0 \mu}}{\sqrt{2}} & b_{1}^{+\mu} & K_{1, B}^{+\mu}  \tag{11}\\
b_{1}^{-\mu} & \frac{h_{1, N}^{\mu}-b_{1}^{0 \mu}}{\sqrt{2}} & K_{1, B}^{0 \mu} \\
K_{1, B}^{-\mu} & \bar{K}_{1, B}^{0 \mu} & h_{1, S}^{\mu}
\end{array}\right)
$$

The mixing angle

$$
\binom{h_{1}(1170)}{h_{1}(1415)}=\left(\begin{array}{cc}
\cos \beta_{b_{1}} & \sin \beta_{b_{1}}  \tag{12}\\
-\sin \beta_{b_{1}} & \cos \beta_{b_{1}}
\end{array}\right)\binom{h_{1, N}}{h_{1, S}}
$$

is not known. Quite interestingly, since this nonet belongs to a heterochiral multiplet (the chiral partners are the orbitally excited vector mesons) the mixing angle could be non-negligible, just as for pseudoscalar mesons. Here, we shall consider two scenarios for our model calculations: in the first, the mixing angle $\beta_{b_{1}}$ is set to zero; in the second we consider a large mixing similar to both the pseudoscalar sector, $\beta_{b_{1}} \approx-40^{\circ}$, and to what two of the authors found for the pseudotensor meson nonet in Ref. [37] and elaborated on in Ref. [81].

Tensor mesons.-For $L=S=1$ three nonets are possible. The well-known $J^{P C}=2^{++} \quad\left(n^{2 S+1} L_{J}=1^{3} P_{2}\right)$ tensor states $\left\{a_{2}(1320), K_{2}^{*}(1430), f_{2}(1270), f_{2}^{\prime}(1525)\right\}$, with elements $A_{2, i j}^{\mu}=\left[\bar{q}_{j}\left(\mathrm{i} \gamma^{\mu} \partial^{\nu}+\ldots\right) q_{i}\right] / \sqrt{2}$, represent an almost ideal nonet of quark-antiquark states

$$
A_{2}^{\mu \nu}=\frac{1}{\sqrt{2}}\left(\begin{array}{ccc}
\frac{f_{2, N}^{\mu \nu}+a_{2}^{0 \mu \nu}}{\sqrt{2}} & a_{2}^{+\mu \nu} & K_{2}^{*+\mu \nu}  \tag{13}\\
a_{2}^{-\mu \nu} & \frac{f_{2, N}^{\mu \nu}-a_{2}^{0 \mu \nu}}{\sqrt{2}} & K_{2}^{* 0 \mu \nu} \\
K_{2}^{*-\mu \nu} & \bar{K}_{2}^{* 0 \mu \nu} & f_{2, S}^{\mu \nu}
\end{array}\right) .
$$

The physical isoscalar-tensor states are

$$
\binom{f_{2}(1270)}{f_{2}^{\prime}(1525)}=\left(\begin{array}{cc}
\cos \beta_{a_{2}} & \sin \beta_{a_{2}}  \tag{14}\\
-\sin \beta_{a_{2}} & \cos \beta_{a_{2}}
\end{array}\right)\binom{f_{2, N}}{f_{2, S}}
$$

where $\beta_{a_{2}}=5.7^{\circ}$ is the small mixing angle reported in the PDG [1], in agreement with the fact that tensor mesons belong to a homochiral multiplet. The decays of tensor mesons were studied in great detail in Refs. $[33,85]$ and fit very well into the $\bar{q} q$ picture [15].

Axial-vector mesons.-The choice $L=S=1$ allows also for the $J^{P C}=1^{++}\left(n^{2 S+1} L_{J}=1^{3} P_{1}\right)$ axial-vector nonet $A_{1}$, which contains the resonances $\left\{a_{1}(1260)\right.$, $\left.K_{1, A}, f_{1}(1285), f_{1}(1420)\right\}$, and which are linked to the vector mesons mentioned above by chiral transformations, see, e.g., Refs. [86,87] (building a homochiral multiplet). The nonet matrix, whose microscopic currents are $A_{1, i j}^{\mu} \equiv \bar{q}_{j} \gamma^{5} \gamma^{\mu} q_{i} / \sqrt{2}$, reads

$$
A_{1}^{\mu}=\frac{1}{\sqrt{2}}\left(\begin{array}{ccc}
\frac{f_{1, N}^{\mu}+a_{1}^{0 \mu}}{\sqrt{2}} & a_{1}^{+\mu} & K_{1, A}^{+\mu}  \tag{15}\\
a_{1}^{-\mu} & \frac{f_{1, N}^{\mu}-a_{1}^{0 \mu}}{\sqrt{2}} & K_{1, A}^{0 \mu} \\
K_{1, A}^{-\mu} & \bar{K}_{1, A}^{0 \mu} & f_{1, S}^{\mu}
\end{array}\right)
$$

For the isoscalar sector we find,

$$
\binom{f_{1}(1285)}{f_{1}(1420)}=\left(\begin{array}{cc}
\cos \beta_{a_{1}} & \sin \beta_{a_{1}}  \tag{16}\\
-\sin \beta_{a_{1}} & \cos \beta_{a_{1}}
\end{array}\right)\binom{f_{1, N}}{f_{1, S}}
$$

The mixing angle $\beta_{a_{1}}$ is expected to be small, as the homochiral nature of the multiplet [81] and the decay properties [34] suggest. We set $\beta_{a_{1}}=0$ for simplicity (an anyhow small mixing angle would not affect our results very much).

It is important to note at this point that the kaonic states $K_{1, A}$ and $K_{1, B}$ mix. For a study of the axial-vector and pseudovector nonets, with focus on the $K_{1}(1270) / K_{1}(1400)$ system, using an approach similar to the one used in this work, see Ref. [34]. The main result concerning the kaonic mixing reads

$$
\binom{K_{1}(1270)}{K_{1}(1400)}=\left(\begin{array}{cc}
\cos \varphi_{K} & -\mathrm{i} \sin \varphi_{K}  \tag{17}\\
-\mathrm{i} \sin \varphi_{K} & \cos \varphi_{K}
\end{array}\right)\binom{K_{1, A}}{K_{1, B}}
$$

Equivalently, in terms of kets and the angle $\theta_{K}=\varphi_{K}+90^{\circ}$ typically used in the literature [86-88]:
$\binom{\left|K_{1}^{+}(1270)\right\rangle}{\left|K_{1}^{+}(1400)\right\rangle}=\left(\begin{array}{cc}\sin \theta_{K} & \cos \theta_{K} \\ \cos \theta_{K} & -\sin \theta_{K}\end{array}\right)\binom{\left|K_{1, A}^{+}\right\rangle}{\left|K_{1, B}^{+}\right\rangle}$.
The final result reads $\varphi_{K}=(56.4 \pm 4.3)^{\circ}$, or equivalently $\theta_{K}=(-33.6 \pm 4.3)^{\circ}$; hence $K_{1}(1270)$ is predominantly $K_{1, B}$, and $K_{1}(1400)$ is predominantly $K_{1, A}$, yet the mixing is still large.

Scalar mesons.-For completeness, we recall that $L=S=1$ yields also the scalar states $J^{P C}=0^{++}$ $\left(n^{2 S+1} L_{J}=1^{3} P_{0}\right)$, corresponding to the chiral partners of the pseudoscalar mesons. The scalars are subject to an
ongoing controversial debate on the correct assignment of measured states. The identification is still uncertain, but the set

$$
\begin{equation*}
\left\{a_{0}(1450), K_{0}^{*}(1430), f_{0}(1710), f_{0}(1370)\right\} \tag{19}
\end{equation*}
$$

seems to be favored (for the description of mixing with, e.g., the long-searched scalar glueball, see, e.g., Refs. [35,36,89-94] and references therein). The reason why this nonet is so peculiar is due to the fact that it has the quantum numbers of the QCD vacuum, hence condensation of the scalar-isoscalar fields should take place. The corresponding nonet $S$, whose microscopic currents are $S_{i j} \equiv\left(\bar{q}_{j} q_{i}\right) / \sqrt{2}$, is not included in this work, since no decays of spin-3 states into scalars can be realized in our model approach, see Sec. II B 3.

Moreover, light scalar four-quark states (in the form and combination of tetraquark/molecular/companion poles) below 1 GeV also exist [1]. Nevertheless, they are also not included our work.

Higher-spin mesons.-Besides the peculiar case of the scalar nonet, all other nonets mentioned above are well established. Yet, what about nonets with $L=2$ ? For $L=2$ four nonets can be constructed, one of which is the main subject of this work. Thus, even if, besides spin-3 fields, they do not enter into the Lagrangian of this work, since they are too heavy, they allow us to put the nonet of mesons with $J=3$ into the correct physical framework. Hence, we briefly discuss them.

The easiest choice is $L=2$ and $S=0$, which implies $J^{P C}=2^{-+}\left(n^{2 S+1} L_{J}=1^{1} D_{2}\right)$ : these so-called pseudotensor states with currents $\left[\bar{q}_{j}\left(\mathrm{i} \gamma^{5} \partial^{\mu} \partial^{\nu}+\ldots\right) q_{i}\right] / \sqrt{2}$ are assigned to $\left\{\pi_{2}(1670), K_{2}(1770), \eta_{2}^{\prime}(1870), \eta_{2}(1645)\right\}$ [37,95-97]. Moreover, as recently described in Ref. [81], the chiral anomaly may also be important for this nonet, explaining a potential large mixing between strange and nonstrange components, leading to the isoscalar states $\eta_{2}(1650)$ and $\eta_{2}^{\prime}(1870)$. The details of the need for a large mixing angle can be found in Ref. [37], in which strong decays of these mesons are studied in a similar relativistic effective model.

For $L=2$ and $S=1$ one can construct the orbitally excited vector $J^{P C}=1^{--} \quad\left(n^{2 S+1} L_{J}=1^{3} D_{1}\right) \quad$ states $\left\{\rho(1700), K^{*}(1680), \phi(? ? ? ?), \omega(1650)\right\}$, which are the (hetero)chiral partners of the pseudovector mesons $[79,81]$. Microscopically, the elements are $\left(\bar{q}_{j} \mathrm{i}^{\mu} q_{i}\right) / \sqrt{2}$. The predominately strange member of the nonet $\phi(? ? ? ?)$ could not be identified yet. For predictions of a possible $\phi(1930)$ meson, see Ref. [38] and references therein, in which a flavor-symmetry-based QFT model is employed for two nonets of excited vector mesons.

The quantum numbers $L=2$ and $S=1$ may also lead to a axial tensor nonet with $J^{P C}=2^{--}\left(n^{2 S+1} L_{J}=1^{3} D_{2}\right)$ with currents $\left[\bar{q}_{j}\left(\gamma^{5} \gamma^{\mu} \partial^{\nu}+\ldots\right) q_{i}\right] / \sqrt{2}$, which is however
poorly understood. In fact, only the state $K_{2}(1820)$ is identified and listed by the PDG [1]. This nonet builds together with the $2^{++}$tensor mesons a homochiral multiplet.

The latter two nonets are omitted in this work, because their members are too heavy.

Spin-3 mesons.-Finally and most importantly, the choice $L=2$ and $S=1$ can also lead the nonet of $J^{P C}=3^{--}\left(n^{2 S+1} L_{J}=1^{3} D_{3}\right)$ states $\left\{\rho_{3}(1690), K_{3}^{*}(1780)\right.$, $\left.\phi_{3}(1850), \omega_{3}(1670)\right\}$. The nonet with elements $W_{i j}^{\mu \nu \rho}=$ $\left[\bar{q}_{j}\left(\gamma^{\mu} \partial^{\nu} \partial^{\rho}+\ldots\right) q_{i}\right] / \sqrt{2}$ reads

$$
W_{3}^{\mu \nu \rho}=\frac{1}{\sqrt{2}}\left(\begin{array}{ccc}
\frac{\omega_{3, N}^{\mu \mu \rho}+\rho_{3}^{0 \mu \nu \rho}}{\sqrt{2}} & \rho_{3}^{+\mu \nu \rho} & K_{3}^{+\mu \nu \rho}  \tag{20}\\
\rho_{3}^{-\mu \nu \rho} & \frac{\omega_{3, N}^{\mu \nu \rho}-\rho_{3}^{0 \mu \nu}}{\sqrt{2}} & K_{3}^{0 \mu \nu \rho} \\
K_{3}^{-\mu \nu \rho} & \bar{K}_{3}^{0 \mu \nu \rho} & \omega_{3, S}^{\mu \nu \rho}
\end{array}\right)
$$

In this case, the mixing angle

$$
\binom{\omega_{3}(1670)}{\phi_{3}(1850)}=\left(\begin{array}{cc}
\cos \beta_{w_{3}} & \sin \beta_{w_{3}}  \tag{21}\\
-\sin \beta_{w_{3}} & \cos \beta_{w_{3}}
\end{array}\right)\binom{\omega_{3, N}}{\omega_{3, S}}
$$

is $\beta_{w_{3}}=3.5^{\circ}$ [1]. The small mixing is also in agreement with the homochiral nature of the corresponding chiral multiplet in which this nonet is embedded [81]. In the next section we couple this nonet to the other nonets listed in Table I.

## B. Effective mesonic interactions

In this section, we present the effective mesonic interactions and discuss the derivation of the tree-level decay widths from our model.

## 1. The effective action

Using the nonets introduced in the previous subsection, we construct the effective Lagrangian describing the strong decays of spin-3 tensor mesons as follows:

$$
\begin{align*}
\mathcal{L}_{W, \text { total }}= & \mathcal{L}_{\text {mass }}+\mathcal{L}_{\text {kin }}+\mathcal{L}_{w_{3} p p}+\mathcal{L}_{w_{3} v_{1} p} \\
& +\mathcal{L}_{w_{3} a_{2} p}+\mathcal{L}_{w_{3} a_{1} p}+\mathcal{L}_{w_{3} b_{1} p}+\mathcal{L}_{w_{3} v_{1} v_{1}} \tag{22}
\end{align*}
$$

where $\mathcal{L}_{\text {kin }}=\frac{1}{2} \operatorname{tr}\left(\partial_{\mu} P\right)^{2}+\cdots$ contains the usual kinetic terms and $\mathcal{L}_{\text {mass }}$ contains all the quadratic mass terms describing the masses of all relevant nonets. In this work, all masses are taken from the PDG [1] and assumed to be exact, while some mixing angles are derived from the masses following the quark model review of the PDG [1], see also Appendix A. The other terms describing the decays are listed explicitly in Table III. The quantities $g_{w_{3} \circ \circ}$ are the coupling constants, which are fitted to experimental results, $\varepsilon^{\mu \nu \rho \sigma}$ is the antisymmetric Levi-Civita pseudotensor, $[0, \circ]_{-}$

TABLE III. Effective relativistic interaction terms describing the strong decays of mesons with $J^{P C}=3^{--}$.

| Decay mode | Interaction Lagrangians |
| :---: | :---: |
| $3^{--} \rightarrow 0^{-+}+0^{-+}$ | $\mathcal{L}_{w_{3} p p}=g_{w_{3} p p} \operatorname{tr}\left[W_{3}^{\mu \nu \rho}\left[P,\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} P\right)\right]_{-}\right]$ |
| $3^{--} \rightarrow 0^{-+}+1^{--}$ | $\mathcal{L}_{w_{3} v_{1} p}=g_{w_{3} v_{1} p} \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[W_{3, \mu \alpha \beta}\left\{\left(\partial_{\nu} V_{1, \rho}\right),\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right]$ |
| $3^{--} \rightarrow 0^{-+}+2^{++}$ | $\mathcal{L}_{w_{3} a_{2} p}=g_{w_{3} a_{2} p} \varepsilon_{\mu \nu \rho \sigma} \operatorname{tr}\left[W_{3}{ }^{\mu}{ }_{\alpha \beta}\left[\left(\partial^{\nu} A_{2}^{\rho \alpha}\right),\left(\partial^{\sigma} \partial^{\beta} P\right)\right]_{-}\right]$ |
| $3^{--} \rightarrow 0^{-+}+1^{+-}$ | $\mathcal{L}_{w_{3} b_{1} p}=g_{w_{3} b_{1} p} \operatorname{tr}\left[W_{3}^{\mu \nu \rho}\left\{B_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right\}_{+}\right]$ |
| $3^{--} \rightarrow 0^{-+}+1^{++}$ | $\mathcal{L}_{w_{3} a_{1} p}=g_{w_{3} a_{1} p} \operatorname{tr}\left[W_{3}^{\mu \nu \rho}\left[A_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right]_{-}\right]$ |
| $3^{--} \rightarrow 1^{--}+1^{--}$ | $\mathcal{L}_{w_{3} v_{1} v_{1}}=g_{w_{3} v_{1} v_{1}} \operatorname{tr}\left[W_{3}^{\mu \nu \rho}\left[\left(\partial_{\mu} V_{1, \nu}\right), V_{1, \rho}\right]_{-}\right]$ |

stands for the commutator, and $\{\circ, \circ\}_{+}$for the anticommutator.

All the interaction Lagrangians in Table III are invariant under CPT, Poincaré, and flavor transformations listed in Table I. We considered only couplings involving a minimal number of derivatives within each single interaction term. This strategy turned to be successful in previous works, e.g., Refs. [33-38]. ${ }^{3}$ Later on, in Sec. V, we shall also argue how to justify this strategy using insights from the extended linear sigma model (ELSM) and the Functional Renormalization Group (FRG) approaches to low energy QCD. The explicit form of the Lagrangians (after having carried out the traces) in Table III are reported in Appendix B. Each Lagrangian represents the dominant contribution in the large- $N_{\mathrm{c}}$ expansion in the given channel, hence each coupling constant scales as $1 / \sqrt{N_{\mathrm{c}}}$.

## 2. Decay width

The tree-level decay widths have the following general form $[1,10]$,

$$
\begin{align*}
& \Gamma_{W_{3} \rightarrow A+B}\left(m_{w_{3}}, m_{a}, m_{b}\right) \\
& \quad=\frac{\left|\vec{k}_{a, b}\right|}{8 \pi m_{w_{3}}^{2}}|\mathcal{M}|^{2} \kappa_{i} \Theta\left(m_{w_{3}}-m_{a}-m_{b}\right), \tag{23}
\end{align*}
$$

where $m_{w_{3}}$ is the mass of a (decaying) spin-3 particle, while $m_{a}$ and $m_{b}$ are the masses of the decay products " $A$ " and " $B, " \Theta(x)$ denotes the Heaviside step function, and the modulus of the outgoing particles momentum has the following analytic expression:

$$
\begin{equation*}
\left|\vec{k}_{a, b}\right| \equiv \frac{1}{2 m_{w_{3}}} \sqrt{\left(m_{w_{3}}^{2}-m_{a}^{2}-m_{b}^{2}\right)^{2}-4 m_{a}^{2} m_{b}^{2}} \tag{24}
\end{equation*}
$$

We obtain the factors $\kappa_{i}$ in Eq. (23) for the $i$ th decay channel from the explicit forms of the Lagrangians in the tables in Appendix B by considering the square of the coefficients for a given decay channel as well as eventual sum over members of the same isospin multiplet. The $\kappa_{i}$ for

[^2]TABLE IV. Decay amplitudes for different decay modes.

| Decay mode | $\frac{1}{7}\|\mathcal{M}\|^{2}$ |
| :---: | :---: |
| $3^{--} \rightarrow 0^{-+}+0^{-+}$ | $g_{w_{3} p p}^{2} \frac{2}{35}\left\|\vec{k}_{p^{(1)}, p^{(2)}}\right\|^{6}$ |
| $3^{--} \rightarrow 0^{-+}+1^{--}$ | $g_{w_{3} v_{1} p}^{2} \frac{8}{105}\left\|\vec{k}_{v_{1}, p}\right\|^{6} m_{w_{3}}^{2}$ |
| $3^{--} \rightarrow 0^{-+}+2^{++}$ | $g_{w_{3} a_{2} p}^{2} \frac{2}{105}\left\|\vec{k}_{a_{2}, p}\right\|^{4} \frac{m_{w_{3}}^{2}}{m_{a_{2}}^{2}}\left(2\left\|\vec{k}_{a_{2}, p}\right\|^{2}+7 m_{a_{2}}^{2}\right)$ |
| $3^{--} \rightarrow 0^{-+}+1^{+-}$ | $g_{w_{3} b_{1} p}^{2} \frac{2}{105}\left\|\vec{k}_{b_{1}, p}\right\|^{4}\left(7+3 \frac{\left\|\vec{k}_{b_{1}, p}\right\|^{2}}{m_{b_{1}}^{2}}\right)$ |
| $3^{--} \rightarrow 0^{-+}+1^{++}$ | $g_{w_{3} a_{1} p}^{2} \frac{2}{105}\left\|\vec{k}_{a_{1}, p}\right\|^{4}\left(7+3 \frac{\mid \vec{k}_{a_{1}, p}^{2}}{m_{a_{1}}^{2}}\right)$ |
| $3^{--} \rightarrow 1^{--}+1^{--}$ | $\begin{aligned} & g_{w_{3} v_{1} v_{1}}^{2} \frac{1}{105}\left(m_{v_{1}^{(1)}}^{2} m_{v_{1}^{(2)}}^{2}\right)^{-1}\left\|\vec{k}_{v_{1}^{(1)}, v_{2}^{(2)}}\right\|^{2}\left[6\left\|\vec{k}_{v_{1}^{(1)}, v_{1}^{(2)}}\right\|^{4}\right. \\ & \quad+35 m_{v_{1}^{(1)}}^{2} m_{v_{1}^{(2)}}^{2}+14 \mid \vec{k}_{v_{1}^{(1)}, v_{1}^{(2)}}^{(2}\left(\left.\right\|^{2}\left(m_{v_{1}^{(1)}}^{2}+m_{v_{1}^{(2)}}^{2}\right)\right] \end{aligned}$ |

all interaction Lagrangians and all relevant channels are presented in Appendix C. The decay amplitudes $|\mathcal{M}|^{2}$ are derived via Feynman rules under the use of the polarization vectors and tensors as well as their corresponding completeness relations in Appendix D. The results for $\frac{1}{7}|\mathcal{M}|^{2}$ are listed in ${ }^{4}$ Table IV.

## 3. Remarks on the exclusion of scalar mesons

As a final side remark, we note that we cannot couple the spin-3 nonet (20) to scalar mesons (19). In fact, if we try to couple them to scalars and pseudoscalars, the only possible $C P T$ - and flavor-invariant interaction term,

$$
\begin{equation*}
\varepsilon_{\mu \nu \rho \sigma} \operatorname{tr}\left(\partial^{\mu} W_{3}{ }^{\nu}{ }_{\alpha \beta}\left[\left(\partial^{\rho} S\right),\left(\partial^{\sigma} \partial^{\alpha} \partial^{\beta} P\right)\right]_{-}\right)=0 \tag{25}
\end{equation*}
$$

vanishes identically since it involves the contraction of antisymmetric and symmetric tensors. Correspondingly, scalar mesons do not enter our study, which agrees with the experimental results reported in the PDG [1]. This is also of direct advantage for our work, since we do not have to deal with the identification of the scalar states, which, as already mentioned, is a long-standing and yet unsolved issue of low-energy QCD [89-94].

[^3]
## 4. Suppression of next to leading order terms

The Lagrangian terms considered in Table III are large- $N_{c}$ (and thus Okubo-Zweig-Iizuka) dominant and symmetric under $U_{\mathrm{V}}(3)$. As we shall discuss later on, this approximation is expected to be acceptable in view of the
precision of the experimental data presently available. Yet, by considering the $W_{3} B_{1} P$ interaction as an example, the Lagrangian $\mathcal{L}_{w_{3} b_{1} p}$ can be regarded as the first term of an expansion in large- $N_{\mathrm{c}}$ and/or symmetry breaking terms that takes the form:

$$
\begin{align*}
\mathcal{L}_{w_{3} b_{1} p}^{\text {full }}= & g_{w_{3} b_{1} p} \operatorname{tr}\left[W^{\mu \nu \rho}\left\{B_{1 \mu},\left(\partial_{\mu} \partial_{\rho} P\right)\right\}_{+}\right]+g_{w_{3} b_{1} p}^{(2)} \operatorname{tr}\left[W^{\mu \nu \rho}\right] \operatorname{tr}\left[B_{1 \mu}\left(\partial_{\mu} \partial_{\rho} P\right)\right] \\
& +g_{w_{3} b_{1} p}^{(3)} \operatorname{tr}\left[W^{\mu \nu \rho}\right] \operatorname{tr}\left[B_{1 \mu}\right] \operatorname{tr}\left[\partial_{\mu} \partial_{\rho} P\right]+g_{w_{3} b_{1} p}^{(4)} \operatorname{tr}\left[\hat{\delta} W^{\mu \nu \rho}\left\{B_{1 \mu},\left(\partial_{\mu} \partial_{\rho} P\right)\right\}_{+}\right] \\
& +g_{w_{3} b_{1} p}^{(5)} \operatorname{tr}\left[\hat{\delta} W^{\mu \nu \rho}\right] \operatorname{tr}\left[B_{1 \mu}\left(\partial_{\mu} \partial_{\rho} P\right)\right]+g_{w_{3} b_{1} p}^{(6)} \operatorname{tr}\left[W^{\mu \nu \rho}\right] \operatorname{tr}\left[\hat{\delta} B_{1 \mu}\left(\partial_{\mu} \partial_{\rho} P\right)\right]+\ldots . \tag{26}
\end{align*}
$$

The first term, reported in Table III, is a flavor-symmetric term and scales as $g_{w_{3} b_{1} p} \propto N_{\mathrm{c}}^{-1 / 2}$, hence it is the dominant term in a large- $N_{\mathrm{c}}$ expansion. The second and the third term are also flavor symmetric but scale as $g_{w_{3} b_{1} p}^{(2)} \propto N_{c^{-\frac{3}{2}}}^{-}$and $g_{w_{3} b_{1} p}^{(3)} \propto N_{\mathrm{c}^{-}}^{-\frac{5}{2}}$, respectively. They involve gluon exchanges and are therefore suppressed. The fourth term is not large$N_{\mathrm{c}}$ suppressed, but it breaks flavor symmetry via the matrix

$$
\begin{equation*}
\hat{\delta}=\operatorname{diag}\left\{0, \delta_{d}, \delta_{s}\right\} \tag{27}
\end{equation*}
$$

Isospin violation is proportional to $\delta_{d} \propto m_{d}-m_{u}$ and is expected to be very small, yet the breaking due to the $s$ quark $\delta_{s} \propto m_{s}-m_{u}$ can be non-negligible following Refs. [89,90], it could be along the order of $g_{w_{3} b_{1} p}^{(4)} \simeq$ $g_{w_{3} b_{1} p} \operatorname{diag}\{0,0,0.1-0.2\}$, but the actual value for $J=3$ mesons should be determined by an independent fit to data. It is however expected to be sufficiently small to be neglected in this work. At this stage, the first large $-N_{c}$ correction $g_{w_{3} b_{1} p}^{(2)}$ and the first flavor-symmetric correction $g_{w_{3} b_{1} p}^{(4)}$ are expected to be of the same intensity and should be the first to be included. Further terms $g_{w_{3} b_{1} p}^{(n \geq 5)}$ are both large $-N_{c}$ subdominant and flavor suppressed, and thus are regarded to be very small.

The same analysis can be carried out for all the interaction terms, yet it is interesting to observe that the second and the third term would vanish whenever the commutator is present. (Flavor-symmetry violation is expected to be the main next-to-leading-order contribution for those interaction terms.)

## III. PHENOMENOLOGY OF THE $J^{P C}=3^{--}$NONET

In this section we present our results for decay rates and branching ratios of the $J^{P C}=3^{--}$mesons. In each subsection and for each interaction Lagrangian term, we compare the experimental data to our theoretical results.

We recall that the total decay widths of the $J^{P C}=3^{--}$ mesons under consideration are [1]

$$
\begin{gather*}
\Gamma_{\rho_{3}(1690)}^{\mathrm{tot}}=(161 \pm 10) \mathrm{MeV},  \tag{28}\\
\Gamma_{K_{3}^{*}(1780)}^{\mathrm{tot}}=(159 \pm 21) \mathrm{MeV},  \tag{29}\\
\Gamma_{\omega_{3}(1670)}^{\mathrm{tot}}=(168 \pm 10) \mathrm{MeV},  \tag{30}\\
\Gamma_{\phi_{3}(1850)}^{\mathrm{tot}}=\left(87_{-23}^{+28}\right) \mathrm{MeV} . \tag{31}
\end{gather*}
$$

We shall verify that the sum of all the single decay channels from our theoretical calculations never overshoots these values.

For what concerns the accuracy of our results, whenever possible we determine the coupling and its error via a simple fit of experimental decay width and ratios. We ignore the experimental uncertainties for all masses of the particles of the model and assume the masses to be exact, because their errors are small and they are of minor importance for the overall errors. The experimental errors for the decay widths and branching ratios as well as the systematic errors in the model are much larger, which justifies this approximation. Yet, the quoted errors represents a lower(!) limit of the actual error of this work, since other indeterminacy features are inherent in the approximations, that lead us to the effective action, where only those terms are included, which are expected to be dominant in terms of large- $N_{c}$, flavor, and momentumspace expansions. Moreover, the decay widths are calculated at tree level. Since the width/mass ratio for the decaying resonances is rather small, contributions due to loops are expected to be negligible [98,99], if seen from a perturbative perspective. Otherwise, if the interaction terms are interpreted as effective couplings in a full quantum effective IR action, then calculations have to be performed at tree level anyhow.

## A. Decay process $\boldsymbol{W}_{3} \rightarrow \boldsymbol{P}+\boldsymbol{P}$

The effective interaction term describing the decay of spin-3 tensor mesons into two pseudoscalar mesons has the following form

$$
\begin{equation*}
\mathcal{L}_{w_{3} p p}=g_{w_{3} p p} \operatorname{tr}\left[W_{3}^{\mu \nu \rho}\left[P,\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} P\right)\right]_{-}\right] \tag{32}
\end{equation*}
$$

Its extended version is listed in Appendix B in Eq. (B1). Correspondingly, the tree-level decay rate is

$$
\begin{align*}
& \Gamma_{W_{3} \rightarrow P^{(1)}+P^{(2)}}\left(m_{w_{3}}, m_{p^{(1)}}, m_{p^{(2)}}\right) \\
& \quad=g_{w_{3} p p}^{2} \frac{\left|\vec{k}_{p^{(1)}, p^{(2)}}\right|^{7}}{140 \pi m_{w_{3}}^{2}} \kappa_{i} \Theta\left(m_{w_{3}}-m_{p^{(1)}}-m_{p^{(2)}}\right) \tag{33}
\end{align*}
$$

compare Eq. (D11). The factors $\kappa_{i}$ are reported in Table XIX. In order to determine the coupling constant $g_{w_{3} p p}$, we use the above formula and the following experimental data:
(1) For $\rho_{3}(1690) \rightarrow \pi \pi$ one finds $\kappa_{1}=1$. Using the experimental result

$$
\begin{equation*}
\Gamma_{\rho_{3}(1690) \rightarrow \pi \pi}^{\exp }=(38.0 \pm 3.2) \mathrm{MeV} \tag{34}
\end{equation*}
$$

one obtains a first determination of the coupling constant squared and its error. We shall denote them as $\tilde{g}_{1}^{2}$ and $\Delta \tilde{g}_{1}^{2}$.
(2) The experimental value

$$
\begin{equation*}
\Gamma_{\rho_{3}(1690) \rightarrow \bar{K} K}^{\exp }=(2.54 \pm 0.45) \mathrm{MeV} \tag{35}
\end{equation*}
$$

together with $\kappa_{2}=2\left(\frac{1}{2}\right)^{2}$ yields an independent determinations of $\tilde{g}_{2}^{2}$ and $\Delta \tilde{g}_{2}^{2}$.
(3) From

$$
\begin{equation*}
\Gamma_{K_{3}^{p}(1780) \rightarrow \pi \bar{K}}^{\exp }=(29.9 \pm 4.3) \mathrm{MeV} \tag{36}
\end{equation*}
$$

and $\kappa_{3}=\left(\frac{1}{2}\right)^{2}+\left(\frac{\sqrt{2}}{2}\right)^{2}$ we get $\tilde{g}_{3}^{2}$ and $\Delta \tilde{g}_{3}^{2}$.
(4) Finally, from

$$
\begin{equation*}
\Gamma_{K_{3}^{*}(1780) \rightarrow \bar{K} \eta}^{\exp }=(48 \pm 22) \mathrm{MeV} \tag{37}
\end{equation*}
$$

and $\kappa_{4}=\left[\frac{1}{2}\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)\right]^{2}$ we obtain $\tilde{g}_{4}$ and $\Delta \tilde{g}_{4}$.
The coupling constant and its indeterminacy are evaluated using a simple $\chi^{2}$ approach, ${ }^{5}$

$$
\begin{equation*}
g_{w_{3} p p}^{2}=\frac{\sum_{i=1}^{4} \frac{\tilde{g}_{i}}{\Delta \tilde{g}_{i}^{2}}}{\sum_{j=1}^{4} \frac{1}{\Delta \tilde{g}_{j}^{2}}}, \quad \Delta g_{w_{3} p p}^{2}=\sqrt{\frac{1}{\sum_{j=1}^{4} \frac{1}{\Delta g_{j}^{2}}}} \tag{38}
\end{equation*}
$$

which results in

$$
\begin{equation*}
g_{w_{3} p p}^{2}=(1.5 \pm 0.1) \times 10^{-10} \mathrm{MeV}^{-4} \tag{39}
\end{equation*}
$$

[^4]The comparison of theoretical and experimental results, which is obtained by using this value for the coupling constant, is reported in Table V. A good overall agreement is obtained, but there is also a sizable mismatch: the experimental value for $K_{3}^{*}(1780) \rightarrow \bar{K} \eta$ is much larger than our theoretical prediction. Still, the experimental error is large and a better experimental determination would be interesting. Moreover, a noteworthy prediction concerning $\phi_{3}(1850) \rightarrow \bar{K} K$ is obtained. From the theoretical large prediction, we conclude that an experimental determination should be feasible.

## B. Decay process $W_{3} \rightarrow V_{1}+\boldsymbol{P}$

The interaction Lagrangian for the vector and pseudoscalar decay mode reads

$$
\begin{align*}
& \mathcal{L}_{w_{3} v_{1} p} \\
& \quad=g_{w_{3} v_{1} p} \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[W_{3, \mu \alpha \beta}\left\{\left(\partial_{\nu} V_{1, \rho}\right),\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right] . \tag{40}
\end{align*}
$$

An extended version is given by Eq. (B2) in Appendix B. In this case, the tree-level decay rate formula has the form

$$
\begin{align*}
& \Gamma_{W_{3} \rightarrow V_{1}+P}\left(m_{w_{3}}, m_{v_{1}}, m_{p}\right) \\
& \quad=g_{w_{3} v_{1} p}^{2} \frac{\left|\vec{k}_{v_{1}, p}\right|^{7}}{105 \pi} \kappa_{i} \Theta\left(m_{w_{3}}-m_{v_{1}}-m_{p}\right) \tag{41}
\end{align*}
$$

see also Eq. (D14), where the factors $\kappa_{i}$ are reported in Table XXX. In order to define the coupling constant we proceed as in the pseudoscalar-pseudoscalar case. We use

$$
\begin{aligned}
\Gamma_{\rho_{3}(1690) \rightarrow \omega(782) \pi}^{\exp } & =(25.8 \pm 9.8) \mathrm{MeV}, \\
\Gamma_{K_{3}^{*}(1780) \rightarrow \rho(770) K}^{\exp } & =(49.3 \pm 15.7) \mathrm{MeV}, \\
\Gamma_{K_{3}^{*}(1780) \rightarrow \bar{K}^{*}(892) \pi}^{\exp } & =(31.8 \pm 9.0) \mathrm{MeV}
\end{aligned}
$$

The coupling constant and its error are

$$
\begin{equation*}
g_{w_{3} v_{1} p}^{2}=\frac{\sum_{i=1}^{3} \frac{\tilde{g}_{i}}{\Delta \tilde{g}_{i}^{2}}}{\sum_{j=1}^{3} \frac{1}{\Delta \tilde{g}_{j}^{2}}}, \quad \Delta g_{w_{3} v_{1} p}^{2}=\sqrt{\frac{1}{\sum_{j=1}^{3} \frac{1}{\Delta g_{j}^{2}}}} \tag{42}
\end{equation*}
$$

TABLE V. Decays of $J^{P C}=3^{--}$mesons into two pseudoscalars. Experimental data is taken from Ref. [1].

| Decay process | Theory $\Gamma / \mathrm{MeV}$ | Experiment $\Gamma / \mathrm{MeV}$ |
| :--- | :---: | :---: |
| $\rho_{3}(1690) \rightarrow \pi \pi$ | $32.7 \pm 2.3$ | $38.0 \pm 3.2$ |
| $\rho_{3}(1690) \rightarrow \bar{K} K$ | $4.0 \pm 0.3$ | $2.54 \pm 0.45$ |
| $K_{3}^{*}(1780) \rightarrow \pi \bar{K}$ | $18.5 \pm 1.3$ | $29.9 \pm 4.3$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K} \eta$ | $7.4 \pm 0.5$ | $48 \pm 22$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K} \eta^{\prime}(958)$ | $0.021 \pm 0.001$ |  |
| $\omega_{3}(1670) \rightarrow \bar{K} K$ | $3.0 \pm 0.2$ |  |
| $\phi_{3}(1850) \rightarrow \bar{K} K$ | $18.8 \pm 1.3$ | Seen |

hence

$$
\begin{equation*}
g_{w_{3} v_{1} p}^{2}=(9.2 \pm 1.9) \times 10^{-16} \mathrm{MeV}^{-6} . \tag{43}
\end{equation*}
$$

This value leads to the results listed in Table VI. We observe that an acceptable agreement is reached, although the $\Gamma_{K_{3}^{*}(1780) \rightarrow \rho(770) K}$ mode is theoretically underestimated (the experimental error is nevertheless large). Quite remarkably, the two theoretically sizable and dominant decays $\omega_{3}(1670) \rightarrow \rho(770) K$ and $\phi_{3}(1850) \rightarrow K^{*}(892) K$ have been indeed seen in experiments, though they could not be quantified. Their future determination would represent a test of our approach-at least on a qualitative level.

In addition, we present the following ratio

$$
\begin{equation*}
\frac{\Gamma_{\phi_{3}(1850) \rightarrow \bar{K}^{*}(892) K}}{\Gamma_{\phi_{3}(1850) \rightarrow \bar{K} K}}=1.9 \pm 0.4, \tag{44}
\end{equation*}
$$

which does not totally contradict the PDG [1] average taken from Ref. [100],

$$
\begin{equation*}
\frac{\Gamma_{\phi_{3}(1850) \rightarrow \bar{K}^{*}(892) K}^{\exp }}{\Gamma_{\phi_{3}(1850) \rightarrow \bar{K} K}^{\exp ^{2}}}=0.55_{-0.45}^{+0.85} . \tag{45}
\end{equation*}
$$

Interestingly Ref. [101] even reports a slightly larger ratio of

$$
\begin{equation*}
\frac{\Gamma_{\phi_{3}(1850) \rightarrow \bar{K}^{*}(892) K}^{\exp }}{\Gamma_{\phi_{3}(1850) \rightarrow \bar{K} K}^{\exp }}=0.8 \pm 0.4 . \tag{46}
\end{equation*}
$$

TABLE VI. Decays of $J^{P C}=3^{--}$mesons into a pseudoscalarvector pair. Experimental data taken from Ref. [1].

| Decay process | Theory $\Gamma / \mathrm{MeV}$ | Experiment <br> $\Gamma / \mathrm{MeV}$ |
| :--- | :---: | :---: |
| $\rho_{3}(1690) \rightarrow \rho(770) \eta$ | $3.8 \pm 0.8$ | Seen |
| $\rho_{3}(1690) \rightarrow \bar{K}^{*}(892) K$ | $3.4 \pm 0.7$ |  |
| $\rho_{3}(1690) \rightarrow \omega(782) \pi$ | $35.8 \pm 7.4$ | $25.8 \pm 9.8$ |
| $\rho_{3}(1690) \rightarrow \phi(1020) \pi$ | $0.036 \pm 0.007$ |  |
| $K_{3}^{*}(1780) \rightarrow \rho(770) K$ | $16.8 \pm 3.5$ | $49.3 \pm 15.7$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}^{*}(892) \pi$ | $27.2 \pm 5.6$ | $31.8 \pm 9.0$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}^{*}(892) \eta$ | $0.09 \pm 0.02$ |  |
| $K_{3}^{*}(1780) \rightarrow \omega(782) \bar{K}$ | $4.3 \pm 0.9$ |  |
| $K_{3}^{*}(1780) \rightarrow \phi(1020) \bar{K}$ | $1.2 \pm 0.3$ |  |
| $\omega_{3}(1670) \rightarrow \rho(770) \pi$ | $97 \pm 20$ | Seen |
| $\omega_{3}(1670) \rightarrow \bar{K}^{*}(892) K$ | $2.9 \pm 0.6$ |  |
| $\omega_{3}(1670) \rightarrow \omega(782) \eta$ | $2.8 \pm 0.6$ |  |
| $\omega_{3}(1670) \rightarrow \phi(1020) \eta$ | $(7.6 \pm 1.6) \times 10^{-6}$ |  |
| $\phi_{3}(1850) \rightarrow \rho(770) \pi$ | $1.1 \pm 0.2$ |  |
| $\phi_{3}(1850) \rightarrow \bar{K}^{*}(892) K$ | $35.5 \pm 7.3$ | Seen |
| $\phi_{3}(1850) \rightarrow \omega(782) \eta$ | $0.015 \pm 0.003$ |  |
| $\phi_{3}(1850) \rightarrow \omega(782) \eta^{\prime}(958)$ | $0.003 \pm 0.001$ |  |
| $\phi_{3}(1850) \rightarrow \phi(1020) \eta$ | $3.8 \pm 0.8$ |  |

Furthermore, we would like to mention a recent lattice QCD study [22], which also confirms our overall predictions on dominant and less dominant vector-pseudoscalar decay channels. Reference [22] predicts (without providing explicit errors due to large uncertainties),

$$
\begin{align*}
& \Gamma_{\rho_{3}(1690) \rightarrow \omega(782) \pi}^{\mathrm{J}, \mathrm{D}}=22 \mathrm{MeV},  \tag{47}\\
& \Gamma_{\rho_{3}(1690) \rightarrow \bar{K}^{*}(892) K}^{\mathrm{J}, \mathrm{D}}=2 \mathrm{MeV} \text {, }  \tag{48}\\
& \Gamma_{\omega_{3}(1670) \rightarrow \rho(770) \pi}^{, \mathrm{D}}=62 \mathrm{MeV},  \tag{49}\\
& \Gamma_{\omega_{3}(1670) \rightarrow \bar{K}^{*}(892) K}^{\mathrm{J}}=2 \mathrm{MeV},  \tag{50}\\
& \Gamma_{\omega_{3}(1670) \rightarrow \omega(782) \eta}^{\mathrm{J}}=1 \mathrm{MeV},  \tag{51}\\
& \Gamma_{\phi_{3}(1850) \rightarrow \bar{K}^{*}(892) K}^{\mathrm{J}, \mathrm{D}}=20 \mathrm{MeV} \text {, }  \tag{52}\\
& \Gamma_{\phi_{3}(1850) \rightarrow \phi(1020)_{\eta}}^{\mathrm{J}}=3 \mathrm{MeV} . \tag{53}
\end{align*}
$$

It is quite remarkable that our (quite simple) model is able to predict at least qualitatively rather similar results to such an advanced and comprehensive lattice QCD study. This leads us to the conclusion, that even for high-spin mesons with masses above 1 GeV like the conventional $3^{--}$-nonet chiral symmetry (breaking) is the decisive guiding principle for their phenomenology.

## C. Decay process $W_{3} \rightarrow \gamma+P$

As a next step, we also present the results for the radioactive decays $W_{3} \rightarrow \gamma P$, where $\gamma$ represents the photon. These can be obtained by using vector meson dominance [102-104], which takes into account of the photon-vector-meson mixing through the shift

$$
\begin{equation*}
V_{\mu \nu} \mapsto V_{\mu \nu}+\frac{e Q}{g_{\rho}} F_{\mu \nu}, \tag{54}
\end{equation*}
$$

where $V_{\mu \nu} \equiv \partial_{\mu} V_{\nu}-\partial_{\nu} V_{\mu}$ and $Q=\operatorname{diag}\left(\frac{2}{3},-\frac{1}{3},-\frac{1}{3}\right)$ is the charge quark matrix, which includes the charges of the up, down, and strange quarks. The electromagnetic field tensor is denoted as $F_{\mu \nu}=\partial_{\mu} a_{\nu}-\partial_{\nu} a_{\mu}$, with $a_{\mu}$ being the photon field, while $e=\sqrt{4 \pi \alpha}$ is the electric coupling constant, and $g_{\rho} \simeq 5.5$ parametrizes the photon-vector-meson transition. By applying the shift of Eq. (54) into Eq. (40) we obtain the Lagrangian for radioactive decays

$$
\begin{align*}
& \mathcal{L}_{w_{3} \gamma p} \\
& \quad=g_{w_{3} v_{1}} \frac{e}{g_{\rho}} \varepsilon^{\mu \nu \rho \sigma}\left(\partial_{\nu} a_{\rho}\right) \operatorname{tr}\left[W_{3, \mu \alpha \beta}\left\{Q,\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right] . \tag{55}
\end{align*}
$$

The extended version of this Lagrangian is provided by Eq. (B3) in Appendix B. For the tree-level decay rate formula, we obtain via Eq. (D17),

$$
\begin{equation*}
\Gamma_{W_{3} \rightarrow \gamma+P}\left(m_{w_{3}}, m_{p}\right)=g_{w_{3} v_{1} p}^{2}\left(\frac{e}{g_{\rho}}\right)^{2} \frac{\left|\vec{k}_{\gamma, p}\right|^{7}}{105 \pi} \kappa_{i}^{\gamma} \tag{56}
\end{equation*}
$$

where one of the masses $m_{a}$ and $m_{b}$ in Eq. (24) is set to zero, because the photon is massless. The factors $\kappa_{i}^{\gamma}$ are listed in Table XX. (The Heaviside function is not needed, because all pseudoscalars are lighter than the conventional spin-3 mesons and the photons can be arbitrarily soft.)

Various predictions for the radiative decays $W_{3} \rightarrow \gamma P$ are calculated and presented in Table VII. Because of the rather large errors in the coupling constant (43), we round all results to integers in units of keV . Quite large radiative decay channels are $\omega_{3}(1670) \rightarrow \gamma \pi^{0}$ as well as $K_{3}^{0}(1780) \rightarrow \gamma K^{0}$.

In general, these processes imply that a photoproduction of mesons with spin $J^{P C}=3^{--}$can take place at the ongoing GlueX [24-26] and CLAS12 [27] experiments at Jefferson Lab.

## D. Decay process $\boldsymbol{W}_{3} \rightarrow \boldsymbol{A}_{2}+\boldsymbol{P}$

The interaction Lagrangian describing the decay of the spin-3 tensor mesons into a spin-2 tensor and a pseudoscalar mesons has the following form

$$
\begin{equation*}
\mathcal{L}_{w_{3} a_{2} p}=g_{w_{3} a_{2} p} \varepsilon_{\mu \nu \rho \sigma} \operatorname{tr}\left[W_{3}{ }^{\mu}{ }_{\alpha \beta}\left[\left(\partial^{\nu} A_{2}^{\rho \alpha}\right),\left(\partial^{\sigma} \partial^{\beta} P\right)\right]_{-}\right] . \tag{57}
\end{equation*}
$$

The extended version is provided in Appendix B in Eq. (B4). Using Eq. (D20) the corresponding decay formula is

$$
\begin{align*}
& \Gamma_{W_{3} \rightarrow A_{2}+P}\left(m_{w_{3}}, m_{a_{2}}, m_{p}\right) \\
&=g_{w_{3} a_{2} p}^{2} \frac{\left|\vec{k}_{a_{2}, p}\right|^{5}}{420 \pi m_{a_{2}}^{2}}\left(2\left|\vec{k}_{a_{2}, p}\right|^{2}+7 m_{a_{2}}^{2}\right) \kappa_{i} \\
& \quad \times \Theta\left(m_{w_{3}}-m_{a_{2}}-m_{p}\right), \tag{58}
\end{align*}
$$

where the $\kappa_{i}$ are listed in Table XXI. For this channel, there are no experimental reported branching ratios that allow for a

TABLE VII. Theoretical predictions for the radiative decays $W_{3} \rightarrow \gamma P$.

| Decay process | Theory $\Gamma / \mathrm{keV}$ |
| :--- | :---: |
| $\rho_{3}^{ \pm / 0}(1690) \rightarrow \gamma \pi^{ \pm / 0}$ | $69 \pm 14$ |
| $\rho_{3}^{0}(1690) \rightarrow \gamma \eta$ | $157 \pm 32$ |
| $\rho_{3}^{0}(1690) \rightarrow \gamma \eta^{\prime}(958)$ | $20 \pm 4$ |
| $K_{3}^{ \pm}(1780) \rightarrow \gamma K^{ \pm}$ | $58 \pm 12$ |
| $K_{3}^{0}(1780) \rightarrow \gamma K^{0}$ | $231 \pm 48$ |
| $\omega_{3}(1670) \rightarrow \gamma \pi^{0}$ | $560 \pm 120$ |
| $\omega_{3}(1670) \rightarrow \gamma \eta$ | $19 \pm 4$ |
| $\omega_{3}(1670) \rightarrow \gamma \eta^{\prime}(958)$ | $1.4 \pm 0.3$ |
| $\phi_{3}(1850) \rightarrow \gamma \pi^{0}$ | $4 \pm 1$ |
| $\phi_{3}(1850) \rightarrow \gamma \eta$ | $129 \pm 26$ |
| $\phi_{3}(1850) \rightarrow \gamma \eta^{\prime}(958)$ | $35 \pm 7$ |

direct determination of the coupling constant. Yet, we can use the experimental ratio listed by the PDG [1] and taken from Ref. [105],

$$
\begin{equation*}
\frac{\Gamma_{\rho_{3}(1690) \rightarrow a_{2}(1320) \pi}^{\exp }}{\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \eta}^{\exp }}=5.5 \pm 2.0 \tag{59}
\end{equation*}
$$

together with the previously determined theoretical value $\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \eta}=(3.8 \pm 0.8) \mathrm{MeV}$ reported in Table VI. The corresponding value for the coupling constant is

$$
\begin{equation*}
g_{w_{3} a_{2} p}^{2}=(2.8 \pm 1.2) \times 10^{-9} \mathrm{MeV}^{-4} \tag{60}
\end{equation*}
$$

Once the coupling constant is fixed, we get the results for the decay rates reported in Table VIII. The decay width $\Gamma_{K_{3}^{*}(1780) \rightarrow \bar{K}_{2}^{*}(1430) \pi}$ is safely smaller the experimental upper limit $[1,106]$. Moreover, the quite large mode $\rho_{3}(1690) \rightarrow$ $a_{2}(1320) \pi$ is seen in experiments, but also in this case no branching ratio is listed in the PDG [1].

## E. Decay process $W_{3} \rightarrow B_{1}+P$

The interaction Lagrangian describing the decay into a pseudovector and pseudoscalar meson is given by

$$
\begin{equation*}
\mathcal{L}_{w_{3} b_{1} p}=g_{w_{3} b_{1} p} \operatorname{tr}\left[W_{3}^{\mu \nu \rho}\left\{B_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right\}_{+}\right] \tag{61}
\end{equation*}
$$

where the extended form is given in Eq. (B5) in Appendix B. The tree-level decay rate takes the form

$$
\begin{align*}
& \Gamma_{W_{3} \rightarrow B_{1}+P}\left(m_{w_{3}}, m_{b_{1}}, m_{p}\right) \\
& =g_{w_{3} b_{1} p}^{2} \frac{\left|\vec{k}_{b_{1}, p}\right|^{5}}{420 \pi m_{w_{3}}^{2}}\left(7+3 \frac{\left|\vec{k}_{b_{1}, p}\right|^{2}}{m_{b_{1}}^{2}}\right) \kappa_{i} \\
& \quad \times \Theta\left(m_{w_{3}}-m_{b_{1}}-m_{p}\right) . \tag{62}
\end{align*}
$$

Here, we use Eq. (D25) and the $\kappa_{i}$ can be found in Table XXII. Also in this case, a direct determination of the coupling constant $g_{w_{3} b_{1} p}$ is not possible due to the lack of experimental information. In order to estimate it, we proceed as it follows.
(1) According to the PDG [1], there exists a lower limit for the ratio

TABLE VIII. Decays of $J^{P C}=3^{--}$mesons into a pseudoscalartensor pair. Experimental data taken from Ref. [1].

| Decay process | Theory <br> $\Gamma / \mathrm{MeV}$ | Experiment <br> $\Gamma / \mathrm{MeV}$ |
| :--- | :---: | :---: |
| $\rho_{3}(1690) \rightarrow a_{2}(1320) \pi$ | $20.9 \pm 8.7$ | Seen |
| $K_{3}^{*}(1780) \rightarrow \bar{K}_{2}^{*}(1430) \pi$ | $5.8 \pm 2.4$ | $<25.4 \pm 3.4$ |
| $K_{3}^{*}(1780) \rightarrow f_{2}(1270) \bar{K}$ | $(5.4 \pm 2.2) \times 10^{-5}$ |  |

$$
\begin{equation*}
\frac{\Gamma_{\omega_{3}(1670) \rightarrow b_{1}(1235) \pi}}{\Gamma_{\omega_{3}(1670) \rightarrow \omega(782) \pi \pi}}>0.75 \tag{63}
\end{equation*}
$$

The PDG [1] extracts this lower bound for the ratio from Ref. [107] and does not list this bound as confirmed data. In Ref. [107] the authors in fact report

$$
\begin{equation*}
\frac{\Gamma_{\omega_{3}(1670) \rightarrow b_{1}(1235) \pi}}{\Gamma_{\omega_{3}(1670) \rightarrow \omega(782) \pi \pi}}=1.0_{-0.25}^{+0.0} . \tag{64}
\end{equation*}
$$

They further argue that the decay mode $\omega_{3}(1670) \rightarrow$ $b_{1}(1235) \pi$ is the dominant contribution to the $\omega(782) \pi \pi$ final state via the subsequent decay $b_{1}(1235) \rightarrow \omega(782) \pi$. They claim that one can assume that

$$
\begin{equation*}
\Gamma_{\omega_{3}(1670) \rightarrow b_{1}(1235) \pi} \approx \Gamma_{\omega_{3}(1670) \rightarrow \omega(782) \pi \pi} \tag{65}
\end{equation*}
$$

and deviations are expected to be smaller than $10 \%$.
(2) Based on this assumption we use the following approximation:

$$
\begin{align*}
& \frac{\Gamma_{\omega_{3}(1670) \rightarrow b_{1}(1235) \pi}}{\Gamma_{\omega_{3}(1670) \rightarrow \rho(770) \pi}} \\
& \stackrel{\mathrm{Eq} .(65)^{\sim}}{\approx} \frac{\Gamma_{\omega_{3}(1670) \rightarrow \omega \pi \pi}}{\Gamma_{\omega_{3}(1670) \rightarrow \rho(770) \pi}} \stackrel{\mathrm{PDG}}{=} 0.71 \pm 0.27 . \tag{66}
\end{align*}
$$

Here, one should mention that the PDG [1] cites Ref. [108] for the second ratio but still excludes it from their confirmed data. Additionally, Ref. [108] even lists a different branching ratio,

$$
\begin{equation*}
\frac{\Gamma_{\omega_{3}(1670) \rightarrow \omega \pi \pi}}{\Gamma_{\omega_{3}(1670) \rightarrow \rho(770) \pi}}=0.47 \pm 0.18 \tag{67}
\end{equation*}
$$

as well as

$$
\begin{equation*}
\frac{\Gamma_{\omega_{3}(1670) \rightarrow b_{1}(1235) \pi}}{\Gamma_{\omega_{3}(1670) \rightarrow \rho(770) \pi}}=0.32 \pm 0.16 \tag{68}
\end{equation*}
$$

It is not clear to the authors of this work, how the PDG [1] extracted their ratio from Ref. [108]. Nevertheless, we will stick to the data reported in the PDG [1] and assume that there was some reasonable reanalysis of the data of Ref. [108]. In this context, it might be worth mentioning that all experimental data concerning the $\omega_{3}(1670)$ is rather old, and its mass and total width could not be determined up to high accuracy at the time when Refs. $[107,108]$ were published. Definitely, there is a need for future experimental investigations of this state.

For what concerns our work, we use the result $\Gamma_{\omega_{3}(1670) \rightarrow \rho(770) \pi}=(97 \pm 20) \mathrm{MeV}$ presented in the previous section as well as the ratio (66). We obtain the following estimate for the coupling constant

$$
\begin{equation*}
g_{w_{3} b_{1} p}^{2} \approx(0.008 \pm 0.003) \mathrm{MeV}^{-2} \tag{69}
\end{equation*}
$$

We are aware that this value and the corresponding results, which are listed in Table IX, are only first rough estimates. Still, they might again help identifying dominant decay channels. Additionally, we consider the following aspects for the determination and interpretation of our results.
(1) We consider $K_{1, B} \approx K_{1}(1270)$, since this is the dominant contribution; then, the decay rate $\Gamma_{\phi_{3}(1850) \rightarrow \bar{K}_{1, B} K}$ is rather small. A disclaimer is in order: due to the large mixing between $K_{1, A}$ and $K_{1, B}$ the results involving the identification $K_{1, B} \approx$ $K_{1}(1270)$ can be only considered as a first approximation. The more correct procedure should be to consider the full mixing in Eq. (17); hence $K_{1, B}$ should be expressed as a superposition of $K_{1}(1270)$ and $K_{1}(1400)$. Yet, this is not an easy task, since the interaction in Eq. (61) alone is not enough. One should also take into account the interaction terms for the decays $W_{3} \rightarrow A_{1} P$ in Eq. (70), which includes $K_{1, A}$, see next section. Then, the decays into $K_{1}(1270)$ and $K_{1}(1400)$ should be calculated by the joint Lagrangians (61) and (70) together with the mixing (17). Furthermore, interference between the two Lagrangians is expected. This calculation is however not possible, since the coupling constant $g_{w_{3} a_{1} p}$ cannot be determined by the present experimental data (see next subsection). Hence, we must limit our study to the dominant assignment $K_{1, B} \approx$ $K_{1}(1270)$ (in this subsection) and $K_{1, A} \approx K_{1}(1400)$ (in the next subsection).

TABLE IX. Decays of $J^{P C}=3^{--}$mesons into a pseudoscalarpseudovector pair obtained by using the mixing angle $\beta_{b_{1}}=0$ and $\beta_{b_{1}}=-40^{\circ}$, respectively. No experimental value is listed in the PDG [1]. The decay channel $\omega_{3}(1670) \rightarrow b_{1}(1235) \pi$ is possibly seen.

|  | Theory for <br> $\beta_{b_{1}}=0 \Gamma / \mathrm{MeV}$ | Theory for <br> $\beta_{b_{1}}=-40^{\circ} \Gamma / \mathrm{MeV}$ |
| :--- | :---: | :---: |
| Decay process | $53 \pm 23$ | $31 \pm 13$ |
| $\rho_{3}(1690) \rightarrow h_{1}(1170) \pi$ | 0 | $0.73 \pm 0.31$ |
| $\rho_{3}(1690) \rightarrow h_{1}(1415) \pi$ | $0.55 \pm 0.24$ | $0.55 \pm 0.24$ |
| $K_{3}^{*}(1780) \rightarrow b_{1}(1235) K$ | $37 \pm 16$ | $37 \pm 16$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}_{1, B} \pi$ | $1.61 \pm 0.70$ | $0.03 \pm 0.01$ |
| $K_{3}^{*}(1780) \rightarrow h_{1}(1170) \bar{K}$ | $69 \pm 30$ | $69 \pm 30$ |
| $\omega_{3}(1670) \rightarrow b_{1}(1235) \pi$ | $1.17 \pm 0.51$ | $1.17 \pm 0.51$ |
| $\phi_{3}(1850) \rightarrow b_{1}(1235) \pi$ | $9.1 \pm 3.9$ | $9.1 \pm 3.9$ |
| $\phi_{3}(1850) \rightarrow \bar{K}_{1, B} K+$ c.c. |  | $1.1 \pm 0.5$ |
| $\phi_{3}(1850) \rightarrow h_{1}(1170) \eta$ | $0.006 \pm 0.003$ |  |

(2) We present the results in Table IX for two values of the unknown mixing angle in the isoscalar sector. In one case, we assume that the mixing angle vanishes, $\beta_{b_{1}} \approx 0$, in the second we study $\beta_{b_{1}} \approx-40^{\circ}$, which is a quite large and negative value similar to the mixing angle for the pseudoscalar isoscalars. Note, the mixing angle $\beta_{b_{1}}$ appears in the decay channels $\rho_{3}(1690) \rightarrow h_{1}(1170) \pi, \quad \rho_{3}(1690) \rightarrow h_{1}(1415) \pi$, $K_{3}^{*}(1780) \rightarrow h_{1}(1170) \bar{K}, \quad$ and $\quad \phi_{3}(1850) \rightarrow$ $h_{1}(1170) \eta$, see Table XXII. The only experimentally possibly seen decay $\omega_{3}(1670) \rightarrow b_{1}(1235) \pi$, see Ref. [108], corresponds to a quite large theoretical partial decay widths and is independent of any assumption on $\beta_{b_{1}}$.
(3) The decay $\rho_{3}(1690) \rightarrow h_{1}(1170) \pi$ is also quite large for both choices of the mixing angle, hence it is a potentially interesting channel for future search. Moreover, it is also an interesting decay channel in order to determine the value of the mixing angle $\beta_{b_{1}}$. The remaining decay channels are pretty small, which might be an explanation for why they could not be observed in experiment.

## F. Decay process $W_{3} \rightarrow A_{1}+\boldsymbol{P}$

The interaction Lagrangian and the decay formula for the decay into an axial vector and a pseudoscalar meson are given by

$$
\begin{equation*}
\mathcal{L}_{w_{3} a_{1} p}=g_{w_{3} a_{1} p} \operatorname{tr}\left[W_{3}^{\mu \nu \rho}\left[A_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right]_{-}\right] \tag{70}
\end{equation*}
$$

and

$$
\begin{align*}
& \Gamma_{W_{3} \rightarrow A_{1}+P}\left(m_{w_{3}}, m_{a_{1}}, m_{p}\right) \\
& \quad=g_{w_{3} a_{1} p}^{2} \frac{\left|\vec{k}_{a_{1}, p}\right|^{5}}{420 \pi m_{w_{3}}^{2}}\left(7+3 \frac{\left|\vec{k}_{a_{1}, p}\right|^{2}}{m_{a_{1}}^{2}}\right) \kappa_{i} \\
& \quad \times \Theta\left(m_{w_{3}}-m_{a_{1}}-m_{p}\right) . \tag{71}
\end{align*}
$$

The extended version of the Lagrangian is given by Eq. (B6) in Appendix B. The $\kappa_{i}$ values for each channel are taken from Table XXIII, and the decay formula was derived by using Eq. (D26). As explained previously, for a first rough estimate, we assume here that $K_{1, A} \approx K_{1}(1400)$.

Since we do not have enough experimental information for obtaining the coupling constant $g_{w_{3} a_{1} p}$, we can only get some theoretical predictions for ratios among different decay channels, which are reported in Table X. Note that the mixing angle $\beta_{a_{1}}$ does not appear in any of the nonzero reported decay channels, as it can be read from Table XXIII. Anyway, the mixing is expected to be small [81].

In the end, it is interesting to mention that there might be an option to size the coupling constant by linking the present model to an underlying chiral model, e.g., the ELSM. This approximately predicts 2 MeV for

TABLE X. Predictions for the nonzero branching ratios of $W_{3} \rightarrow A_{1} P$. We choose as a reference the $a_{1}(1260) \pi$ decay mode.

| Branching ratio | Theory |
| :---: | :---: |
| $\frac{\Gamma_{K_{3}^{*}(1780) \rightarrow K_{1, A}}}{\Gamma_{\rho_{3}(169) \rightarrow a_{1}(1260) \pi}}$ | 0.12 |
| $\frac{\Gamma_{K_{3}^{*}(1780) \rightarrow a_{1}(1260) K}}{\Gamma_{\rho_{3}(1690) \rightarrow a_{1}(1260) \pi}}$ | 0.01 |

$\rho_{3}(1690) \rightarrow a_{1}(1260) \pi$, see Sec . V for an estimate of $g_{w_{3} a_{1} p}$. It seems therefore that the decays of the type $W_{3} \rightarrow A_{1}+P$ are suppressed.

## G. Decay process $W_{3} \rightarrow V_{1}+V_{1}$

The interaction Lagrangian describing the decay into two vector mesons is given by

$$
\begin{equation*}
\mathcal{L}_{w_{3} v_{1} v_{1}}=g_{w_{3} v_{1} v_{1}} \operatorname{tr}\left[W_{3}^{\mu \nu \rho}\left[\left(\partial_{\mu} V_{1, \nu}\right), V_{1, \rho}\right]_{-}\right], \tag{72}
\end{equation*}
$$

see Eq. (B7) in Appendix B for the extended version. The corresponding decay rate reads

$$
\begin{align*}
\Gamma_{W_{3}} \rightarrow V_{1}^{(1)}+V_{1}^{(2)} & \left(m_{w_{3}}, m_{v_{1}^{(1)}}, m_{v_{1}^{(2)}}\right) \\
& =g_{w_{3} v_{1} v_{1}}^{2} \frac{\left|\vec{k}_{v_{1}^{(1)}, v_{1}^{(2)}}\right|^{3}}{840 \pi m_{v_{1}^{(1)}}^{2} m_{v_{1}^{(2)}}^{2} m_{w_{3}}^{2}}\left[6\left|\vec{k}_{v_{1}^{(1)}, v_{1}^{(2)}}\right|^{4}\right. \\
& \left.+35 m_{v_{1}^{(1)}}^{2} m_{v_{1}^{(2)}}^{2}+14\left|\vec{k}_{v_{1}^{(1)}, v_{1}^{(2)}}\right|^{2}\left(m_{v_{1}^{(1)}}^{2}+m_{v_{1}^{(2)}}^{2}\right)\right] \\
& \times \kappa_{i} \Theta\left(m_{w_{3}}-m_{v_{1}^{(1)}}-m_{v_{1}^{(2)}}\right), \tag{73}
\end{align*}
$$

where the $\kappa_{i} \mathrm{~s}$ are listed in Table XXIV. For the derivation of the decay formula, we use Eq. (29). Also in this case, we cannot determine the coupling constant, but we can present the branching ratios reported in Table XI.

Moreover, we can find an upper limit for the coupling constant by considering

$$
\begin{align*}
& \frac{\Gamma_{\rho_{3}(1690) \rightarrow \pi \pi}}{\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \rho(770)}} \\
& \quad \gtrsim \frac{\Gamma_{\rho_{3}(1690) \rightarrow \pi \pi}}{\Gamma_{\rho_{3}(1690) \rightarrow \pi^{ \pm} \pi^{+} \pi^{-} \pi^{0}}} \stackrel{\mathrm{PDG}}{=} 0.35 \pm 0.11, \tag{74}
\end{align*}
$$

TABLE XI. Predictions for the branching ratios of $W_{3} \rightarrow$ $V_{1}+V_{1}$. We choose as a reference the $\rho \rho$ decay mode.

| Branching ratio | Theory |
| :--- | :--- | :--- |
| $\frac{\Gamma_{K_{3}^{*}(1780) \rightarrow \rho(770) K^{*}(892)}}{\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \rho(770)}}$ | 0.49 |
| $\frac{\Gamma_{K_{3}^{*}(1780) \rightarrow K^{*}(892) \omega(782)}}{\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \rho(770)}}$ | 0.17 |
| $\frac{\Gamma_{\phi_{3}(1850) \rightarrow K^{*}(892) K^{*}(892)}}{\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \rho(770)}}$ | 0.35 |

which follows from the fact that the $\rho_{3}(1690) \rightarrow$ $\rho(770) \rho(770)$ mode is part of the $\pi^{ \pm} \pi^{+} \pi^{-} \pi^{0}$ decay mode (and eventually it is one of its dominating contributors). Since the experimental value for $\Gamma_{\rho_{3}(1690) \rightarrow \pi \pi}=(38.0 \pm$ 3.2) MeV is known, we find

$$
\begin{equation*}
g_{w_{3} v_{1} v_{1}}^{2} \lesssim 535 \pm 174 . \tag{75}
\end{equation*}
$$

Based on this information we can predict upper limits for the decay rates in Table XII. The theoretically largest decay is $\rho_{3}(1690) \rightarrow \rho(770) \rho(770)$, which has been experimentally seen. Yet, the decay width of about 108 MeV is surely too large, since using a value close to this upper limit would
imply that the sum of all the decays of the state $\rho_{3}(1690)$ would overshoot the experimental total width of $(161 \pm 10) \mathrm{MeV}$. This is in agreement with the ratio

$$
\begin{equation*}
\frac{\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \rho(770)}}{\Gamma_{\rho_{3}(1690) \rightarrow \pi^{ \pm} \pi^{+} \pi^{-} \pi^{0}}}, \tag{76}
\end{equation*}
$$

which is smaller than 1 for all experimental measurements [109-112], which are listed by the PDG [1].

However, we can then obtain a second, more realistic estimate of these decay channels by summing up the largest decay channels:

$$
\begin{align*}
\Gamma_{\rho_{3}}^{\mathrm{tot}}= & 161 \mathrm{MeV} \approx \\
= & \Gamma_{\rho_{3}(1690) \rightarrow \pi \pi}+\Gamma_{\rho_{3}(1690) \rightarrow \bar{K} K}+\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \eta}+\Gamma_{\rho_{3}(1690) \rightarrow \bar{K}^{*}(892) K} \\
& +\Gamma_{\rho_{3}(1690) \rightarrow \omega(782) \pi}+\Gamma_{\rho_{3}(1690) \rightarrow \phi(1020) \pi}+\Gamma_{\rho_{3}(1690) \rightarrow a_{2}(1320) \pi}+\Gamma_{\rho_{3}(1690) \rightarrow a_{1}(1260) \pi} \\
& +\Gamma_{\rho_{3}(1690) \rightarrow h_{1}(1170) \pi}+\Gamma_{\rho_{3}(1690) \rightarrow h_{1}(1415) \pi}+\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \rho(770)}+\Gamma_{\rho_{3}(1690) \rightarrow \mathrm{other}}, \tag{77}
\end{align*}
$$

where $\Gamma_{\rho_{3} \rightarrow \text { other }}$ refers to all other decay channels not listed above. Anyhow, these should be small. For $\Gamma_{\rho_{3}(1690) \rightarrow h_{1}(1170) \pi}$ we use the smaller value presented in Table IX, assuming a large mixing $\beta_{b_{1}}=-40^{\circ}$. In this way we estimate the more reasonable upper limit

$$
\begin{equation*}
\Gamma_{\rho_{3}(1690) \rightarrow \rho(770) \rho(770)} \lesssim 30 \mathrm{MeV} \tag{78}
\end{equation*}
$$

from which follows

$$
\begin{equation*}
g_{w_{3} v_{1} v_{1}}^{2} \lesssim 148 \tag{79}
\end{equation*}
$$

The corresponding estimates for the other channels can be found in Table XII.

Following this estimate, we perform the sum of the predicted and sizable decay channels for the other spin-3 mesons, obtaining:

TABLE XII. Upper limits for the decays of $J^{P C}=3^{--}$mesons into a vector-vector pair. Two different estimates are used, see text.

|  | Theory using <br> $(74) \&(75)$ <br> $\Gamma / \mathrm{MeV}$ | Theory using <br> $(77) \&(79)$ <br> $\Gamma / \mathrm{MeV}$ |
| :--- | :---: | :---: |
| Decay process | $\lesssim 108.6 \pm 35.3$ | $\lesssim 30$ |
| $\rho_{3}(1690) \rightarrow \rho(770) \rho(770)$ | $\lesssim 53.6 \pm 17.4$ | $\lesssim 15$ |
| $K_{3}^{*}(1780) \rightarrow \rho(770) \bar{K}^{*}(892)$ | $\lesssim 19.0 \pm 6.2$ | $\lesssim 5$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}^{*}(892) \omega(782)$ | $\lesssim 38.3 \pm 12.5$ | $\lesssim 11$ |
| $\phi_{3}(1850) \rightarrow \bar{K}^{*}(892) K^{*}(892)$ |  |  |

$$
\begin{align*}
& \Gamma_{K_{3}^{*}(1780)}^{\mathrm{sum}} \approx 146 \mathrm{MeV},  \tag{80}\\
& \Gamma_{\omega_{3}(1670)}^{\mathrm{sum}} \approx 175 \mathrm{MeV}  \tag{81}\\
& \Gamma_{\phi_{3}(1850)}^{\mathrm{sum}} \approx 80 \mathrm{MeV} \tag{82}
\end{align*}
$$

which are in agreement with the experimental values from the PDG [1],

$$
\begin{gather*}
\Gamma_{K_{3}^{*}(1780)}^{\mathrm{tot}}=(159 \pm 21) \mathrm{MeV}  \tag{83}\\
\Gamma_{\omega_{3}(1670)}^{\mathrm{tot}}=(168 \pm 10) \mathrm{MeV}  \tag{84}\\
\Gamma_{\phi_{3}(1850)}^{\mathrm{tot}}=\left(87_{-23}^{+28}\right) \mathrm{MeV} \tag{85}
\end{gather*}
$$

These results demonstrate that the model is internally consistent.

## IV. PHENOMENOLOGY OF THE $J^{P C}=3^{--}$GLUEBALL

In this chapter, we study the branching ratios of the decays of a hypothetical glueball with $J^{P C}=3^{--}$. Similar to all other bound states of gluons, this glueball is not yet experimentally detected. However, lattice QCD calculations in the quenched approximation predict its mass of approximately 4.13 GeV [3], 4.33 GeV [4], or 4.20 GeV [5]. For general works on glueballs, see, e.g., Refs. [113116]. For model approaches to the phenomenology of glueballs, which are similar to this one, see for example previous publications by some of the authors [33,37,56,79,117].

TABLE XIII. Strong interaction Lagrangian terms for the hypothetical glueball state $G_{3}(4200)$ [in comparison with Eq. (22)].

| Decay mode | Interaction Lagrangians |
| :--- | :---: |
| $G_{3} \rightarrow V_{1}+P$ | $\mathcal{L}_{g_{3} v_{1} p}=c_{g_{3} v_{1} p} G_{3, \mu \alpha \beta} \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[\left\{\left(\partial_{\nu} V_{1, \rho}\right),\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right]$ |
| $G_{3} \rightarrow V_{1}+A_{1}$ | $\mathcal{L}_{g_{3} v_{1} a_{1}}=c_{g_{3} v_{1} a_{1}} G_{3, \mu \alpha} \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[\left\{\left(\partial_{\nu} V_{1, \rho}\right),\left(\partial^{\alpha} \partial^{\beta} A_{1, \sigma}\right)\right\}_{+}\right]$ |
| $G_{3} \rightarrow B_{1}+P$ | $\left.\mathcal{L}_{g_{3} b_{1} p}=c_{g_{3} b_{1} p} G_{3}^{\mu \nu \rho} \operatorname{tr} \operatorname{tr}\left[\left\{B_{1, \mu}, \partial_{\nu} \partial_{\rho} P\right)\right\}_{+}\right]$ |
| $G_{3} \rightarrow B_{1}+A_{1}$ | $\mathcal{L}_{g_{3} b_{1} a_{1}}=c_{g_{3} b_{1} a_{1}} G_{3}^{\mu \nu \rho} \operatorname{tr} \operatorname{tr}\left[\left\{B_{1, \mu},\left(\partial_{\nu} A_{1, \rho}\right)\right\}_{+}\right]$ |

Before we start our discussion, we explicitly state that all branching ratios should be considered as first indicative results.

## A. The effective model for glueball decays

The Lagrangian describing the decays of this glueball can be obtained by making use of the fact that each glueball is a flavor singlet-an object that is invariant under $S U_{\mathrm{V}}(3)$ transformations. Hence, we can perform the simple replacement

$$
\begin{equation*}
W^{\mu \nu \rho} \rightarrow G_{3}^{\mu \nu \rho} \cdot \mathbb{1}_{3 \times 3} \tag{86}
\end{equation*}
$$

in the effective interaction terms, where $\mathbb{1}_{3 \times 3}$ is the flavor-space-identity matrix and $G_{3}^{\mu \nu \rho}$ is the glueball field [33,117]. As a consequence, the interaction terms that come with a commutator vanish, and the only nonzero contributions to a residual IR-effective action for the $3^{--}$ glueball can be derived from the former interaction terms involving the anticommutator. Thus,

$$
\begin{equation*}
\Gamma_{G_{3}(4200) \rightarrow V_{1}+P} \neq 0, \quad \Gamma_{G_{3}(4200) \rightarrow B_{1}+P} \neq 0 \tag{87}
\end{equation*}
$$

The corresponding Lagrangian terms are reported in Table XIII. In addition, there are also decays of the type $G_{3}(4200) \rightarrow V_{1}+A_{1}$ and $G_{3}(4200) \rightarrow B_{1}+A_{1}$ that were not included in for the lightest conventional mesons with $J^{P C}=3^{--}$because they were kinematically forbidden. They are however possible for the glueball $G_{3}(4200)$ since it is expected to be much heavier than the $3^{--}$nonet (6). Below we present the theoretical predictions for ratios of decays, which can be easily calculated following the same steps of the previous sections. The only ingredients that are needed for a determination of the branching ratios are the masses of the particles, which are for the sake of simplicity again assumed to be exact. For the $3^{--}$glueball mass, we use $m_{g_{3}}=4.2 \mathrm{GeV}$ [5].

## B. Decay process $\boldsymbol{G}_{\mathbf{3}} \rightarrow \boldsymbol{V}_{\mathbf{1}}+\boldsymbol{P}$

The interaction Lagrangian for decaying spin-3 tensor glueball into a vector and a pseudoscalar meson reads

$$
\begin{align*}
& \mathcal{L}_{g_{3} v_{1} p} \\
& \quad=c_{g_{3} v_{1} p} G_{3, \mu \alpha \beta} \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[\left\{\left(\partial_{\nu} V_{1, \rho}\right),\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right] . \tag{88}
\end{align*}
$$

An extended version is given by Eq. (B8) in Appendix B. The corresponding decay formula is analogous to Eq. (41),

$$
\begin{align*}
& \Gamma_{G_{3} \rightarrow V_{1}+P}\left(m_{g_{3}}, m_{v_{1}}, m_{p}\right) \\
& \quad=c_{g_{3} v_{1} p}^{2} \frac{\left|\vec{k}_{v_{1}, p}\right|^{7}}{105 \pi} \kappa_{i} \Theta\left(m_{g_{3}}-m_{v_{1}}-m_{p}\right) \tag{89}
\end{align*}
$$

Here, the $\kappa_{i}$ s can be taken from Table XXV. We use again Eq. (24) for the momenta of the outgoing particles throughout this chapter, while replacing the mass of the conventional spin-3 mesons $m_{w_{3}}$ with the glueball mass $m_{g_{3}}=4200 \mathrm{MeV}$. Since the coupling constant $c_{g_{3} v_{1} p}$ is unknown, we calculate the nonzero ratios in Table XIV. This is an interesting decay channel to search for this glueball. Since the vector mesons further decay in two (or three) pseudoscalar mesons, one should search for final channels with three (or four) pseudoscalar mesons.

## C. Decay process $G_{3} \rightarrow \gamma+\boldsymbol{P}$

Using vector meson dominance via the shift of Eq. (54) and analogous replacements, we can also construct a radioactive decay term for the spin- 3 glueball,

$$
\begin{align*}
& \mathcal{L}_{g_{3} \gamma p} \\
& \quad=c_{g_{3} \gamma p} \frac{e}{g_{\rho}} G_{3, \mu \alpha \beta}\left(\partial_{\nu} a_{\rho}\right) \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[\left\{Q,\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right] . \tag{90}
\end{align*}
$$

An extended version of this Lagrangian can be found in Eq. (B9) in Appendix B. For the decay formula one derives

TABLE XIV. Predictions for the branching ratios of glueball $G_{3}(4200) \rightarrow V_{1}+P$.

|  | Branching ratio | Theory |
| :--- | :--- | :--- |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow \bar{K}^{*}(892) K}}{\Gamma_{G_{3}(4200) \rightarrow \rho(770) \pi}}$ | 1.11 |  |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow \omega(782) \eta}}{\Gamma_{G_{3}(4200) \rightarrow \rho(770) \pi}}$ | 0.17 |  |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow \omega(782) \eta^{\prime}(958)}}{\Gamma_{G_{3}(4200) \rightarrow \rho(770) \pi}}$ | 0.089 |  |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow \phi(1020) \eta}}{\Gamma_{G_{3}(4200) \rightarrow \rho(770) \pi}}$ | 0.098 |  |
|  | $\frac{\Gamma_{G_{3}(4200) \rightarrow \phi(1020) \eta^{\prime}(958)}}{\Gamma_{G_{3}(4200) \rightarrow \rho(770) \pi}}$ | 0.11 |

$$
\Gamma_{G_{3} \rightarrow \gamma+P}\left(m_{g_{3}}, m_{p}\right)=c_{g_{3} v_{1} p}^{2}\left(\frac{e}{g_{\rho}}\right)^{2} \frac{\left|\vec{k}_{\gamma, p}\right|^{7}}{105 \pi} \kappa_{i}^{\gamma}
$$

where the corresponding factors $\kappa_{i}^{\gamma}$ are listed in Table XXVI. Similar to Eq. (56) one of the masses $m_{a}$ and $m_{b}$ in Eq. (24) has to be set to zero, due to the masslessness of the photon.

From these formula, we calculate ratios for the radioactive decays $\Gamma_{G_{3} \rightarrow \gamma P}$ with respect to the strong decay channel $\Gamma_{G_{3}(4200) \rightarrow \rho(770) \pi}$. The results are listed in Table XV.

## D. Decay process $G_{3} \rightarrow V_{1}+A_{1}$

We also present the glueball decays into vectors and axial vectors by considering the Lagrangian in which $G_{3}$ couples to $V_{1} A_{1}$

$$
\begin{align*}
& \mathcal{L}_{g_{3} v_{1} a_{1}} \\
& \quad=c_{g_{3} v_{1} a_{1}} G_{3, \mu \alpha \beta} \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[\left\{\left(\partial_{\nu} V_{1, \rho}\right),\left(\partial^{\alpha} \partial^{\beta} A_{1, \sigma}\right)\right\}_{+}\right] \tag{91}
\end{align*}
$$

The extended version is provided in Appendix B in Eq. (D20). The tree-level decay formula is given by

$$
\begin{align*}
& \Gamma_{G_{3} \rightarrow V_{1}+A_{1}}\left(m_{g_{3}}, m_{v_{1}}, m_{a_{1}}\right) \\
& \quad=c_{g_{3} v_{1} a_{1}}^{2} \frac{\left|\vec{k}_{v_{1}, a_{1}}\right|^{5}}{210 \pi m_{g_{3}}^{2}}\left[\left|\vec{k}_{v_{1}, a_{1}}\right|^{2}\left(3+2 \frac{m_{g_{3}}^{2}}{m_{a_{1}}^{2}}\right)+7 m_{v_{1}}^{2}\right] \\
& \quad \times \kappa_{i} \Theta\left(m_{g_{3}}-m_{v_{1}}-m_{a_{1}}\right) \tag{92}
\end{align*}
$$

where we used Eq. (D30). The coefficients $\kappa_{i}$ are listed in Table XXVII. Results are listed in Table XVI.

## E. Decay process $\boldsymbol{G}_{\mathbf{3}} \rightarrow \boldsymbol{B}_{\mathbf{1}}+\boldsymbol{P}$

The second dominant decay of the $J^{P C}=3^{--}$glueball is the one into a pseudovector and a pseudoscalar meson. The interaction Lagrangian is given by

$$
\begin{equation*}
\mathcal{L}_{g_{3} b_{1} p}=c_{g_{3} b_{1} p} G_{3}^{\mu \nu \rho} \operatorname{tr}\left[\left\{B_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right\}_{+}\right], \tag{93}
\end{equation*}
$$

and its extended version can be found in Eq. (B11) in Appendix B. This implies that the formula for the decay width is identical to Eq. (62),

TABLE XV. Predictions for the branching ratios of glueball $G_{3}(4200) \rightarrow \gamma+P$ with the $G_{3}(4200) \rightarrow \rho \pi$ reference decay mode.

|  | Branching ratio | Theory |
| :--- | :--- | :--- |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow \gamma \pi^{0}}}{\Gamma_{G_{3}(4200) \rightarrow \rho(770) \pi}}$ | $129 \times 10^{-5}$ |  |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow r \eta}}{\Gamma_{G_{3}(4200) \rightarrow \rho(770) \pi}}$ | $37 \times 10^{-5}$ |  |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow r \eta^{\prime}(958)}}{\Gamma_{G_{3}(4200) \rightarrow \rho(770) \pi}}$ | $1 \times 10^{-5}$ |  |

TABLE XVI. Predictions for the branching ratios of glueball $\underline{\underline{G_{3}(4200)} \rightarrow V_{1}+A_{1} .}$

| Branching ratio | Theory |
| :---: | :---: |
| $\Gamma_{G_{3}(4200) \rightarrow K^{*}(892) K_{1, A}}$ | 0.78 |
| $\overline{\Gamma_{C_{3}(4200) \rightarrow \rho(770) a_{1}(1260)}}$ |  |
| $\frac{\Gamma_{C_{3}(4200) \rightarrow \omega(782) f_{1}(1285)}}{\Gamma^{(120}}$ | 0.29 |
| $\overline{\Gamma_{G_{3}(4200) \rightarrow \rho(770) a_{1}(1260)}}$ |  |
| $\frac{\Gamma_{C_{3}(4200) \rightarrow \omega(788) f_{1}(1420)}}{\Gamma_{1}}$ | 0.0009 |
| $\bar{\Gamma}_{\Gamma_{3}(4200) \rightarrow \rho(770) a_{1}(1260)}$ |  |
| $\frac{\Gamma_{C_{3}(4200) \rightarrow \phi(1020) f_{1}(1285)}}{\Gamma_{6}}$ | 0.0011 |
| $\Gamma_{G_{3}(4200) \rightarrow \rightarrow(770) a_{1}(1260)}$ $\Gamma_{G_{3}(220) \rightarrow \phi(120) f_{1}(1220)}$ $\Gamma_{6}(120)$ | 0.16 |
| $\overline{\Gamma_{G_{3}(4200) \rightarrow \rho(770) a_{1}(1260)}}$ |  |

$$
\begin{align*}
& \Gamma_{G_{3} \rightarrow B_{1}+P}\left(m_{g_{3}}, m_{b_{1}}, m_{p}\right) \\
&=c_{g_{3} b_{1} p}^{2} \frac{\left|\vec{k}_{b_{1}, p}\right|^{5}}{420 \pi m_{g_{3}}^{2}}\left(7+3 \frac{\left|\vec{k}_{b_{1}, p}\right|^{2}}{m_{b_{1}}^{2}}\right) \\
& \times \kappa_{i} \Theta\left(m_{g_{3}}-m_{b_{1}}-m_{p}\right) . \tag{94}
\end{align*}
$$

The factors $\kappa_{i}$ are listed in Table XXVIII. Again, we study two possible choices for the mixing angle ( $\beta_{b_{1}} \approx 0^{\circ}$ and $\beta_{b_{1}} \approx-40^{\circ}$ ) and upon using $m_{K_{1, B}} \approx m_{K_{1}(1270)}$, we present the results in Table XVII.

## F. Decay process $\boldsymbol{G}_{3} \rightarrow \boldsymbol{B}_{1}+\boldsymbol{A}_{1}$

Moreover, we present the results for the glueball into a pseudovector and an axial vector, $G_{3} \rightarrow B_{1}+A_{1}$ in Table XVIII for $\beta_{a_{1}} \approx 0^{\circ}$. Namely, the Lagrangian for this decay can be obtained from Eq. (93) as

$$
\begin{equation*}
\mathcal{L}_{g_{3} b_{1} a_{1}}=c_{g_{3} b_{1} a_{1}} G_{3}^{\mu \nu \rho} \operatorname{tr}\left[\left\{B_{1, \mu},\left(\partial_{\nu} A_{1, \rho}\right)\right\}_{+}\right] . \tag{95}
\end{equation*}
$$

The extended form is presented in Eq. (B12) in Appendix B. The corresponding decay formula is given by

TABLE XVII. Predictions for the branching ratios of glueball $\underline{\underline{G_{3}(4200)} \rightarrow B_{1}+P .}$

| Branching ratio | Theory for $\beta_{b_{1}}=0^{\circ}$ | Theory for $\beta_{b_{1}}=-40^{\circ}$ |
| :---: | :---: | :---: |
| $\Gamma_{G_{3}(4200) \rightarrow K_{1, B}{ }^{\text {K }}}$ | 1.15 | 1.15 |
| $\overline{\Gamma_{G_{3}(4200) \rightarrow b_{1}(1235) \pi}}$ |  |  |
| $\frac{\Gamma_{G_{3}(4200)-h_{1}(1170) n}}{\Gamma^{\prime}}$ | 0.17 | 0.33 |
| $\Gamma_{G_{3}(4200) \rightarrow b_{1}(1235) \pi}$ |  |  |
|  | 0.12 | 0.001 |
| $\Gamma_{G_{3}(4200) \rightarrow b_{1}(1235) \pi}$ |  |  |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow h_{1}(1415) n}}{\Gamma^{\prime}}$ | 0.10 | 0.001 |
| $\overline{\Gamma_{G_{3}(4200) \rightarrow b_{1}(1235) \pi}}$ |  |  |
| $\frac{\Gamma_{G_{3}(4220) \rightarrow h_{1}(1415) h^{\prime}(958)}}{\Gamma^{(120)}}$ | 0.08 | 0.16 |
| $\Gamma_{\mathrm{G}_{3}(4200) \rightarrow b_{1}(1235) \pi}$ |  |  |

TABLE XVIII. Predictions for the branching ratios of glueball $G_{3}(4200) \rightarrow B_{1}+A_{1}$.

| Branching ratio | Theory for $\beta_{b_{1}}=0^{\circ}$ | Theory for $\beta_{b_{1}}=-40^{\circ}$ |
| :---: | :---: | :---: |
| $\Gamma_{G_{3}(4200) \rightarrow K_{1, B} K_{1, A}}$ | 0.96 | 0.96 |
| $\overline{\Gamma_{G_{3}(4200) \rightarrow b_{1}(1235) a_{1}(1260)}}$ |  |  |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow h_{1}(1170) f_{1}(1285)}}{\Gamma}$ | 0.34 | 0.20 |
| $\Gamma_{C_{3}(4200) \rightarrow b_{1}(1235) a_{1}(1260)}$ |  |  |
| $\frac{\Gamma_{C_{3}(4200) \rightarrow h_{1}(1170) f_{1}(1420)}}{\Gamma_{G^{\prime}}(1200)}$ | 0 | 0.11 |
| $\Gamma_{C_{G_{3}}(4200) \rightarrow b_{1}(1235) a_{1}(1260)} \Gamma_{G_{3}(1200) \rightarrow h_{1}(1415) f_{1}(1285)}$ | 0 | 0.092 |
| $\frac{\left.C_{3}(4200) \rightarrow h_{1}(124)\right)_{1}(1285)}{\Gamma_{C_{3}(4200) \rightarrow b_{1}(1235) a_{1}(1260)}}$ | 0 | 0.092 |
| $\frac{\Gamma_{G_{3}(4200) \rightarrow h_{1}(1415) f_{1}(1420)}}{\Gamma_{G_{4}(2200) \rightarrow b_{1}(1235)}(1200)}$ | 0.17 | 0.10 |
| $\underline{\Gamma_{G_{3}(4200) \rightarrow b_{1}(1235) a_{1}(1260)}}$ |  |  |

$$
\begin{align*}
& \Gamma_{G_{3} \rightarrow B_{1}+A_{1}}\left(m_{g_{3}}, m_{b_{1}}, m_{a_{1}}\right) \\
&= c_{g_{3} b_{1} a_{1}}^{2} \frac{\left|\vec{k}_{b_{1}, a_{1}}\right|^{3}}{840 \pi m_{b_{1}}^{2} m_{a_{1}}^{2} m_{g_{3}}^{2}}\left[6\left|\vec{k}_{b_{1}, a_{1}}\right|^{4}\right. \\
&\left.+35 m_{b_{1}}^{2} m_{a_{1}}^{2}+14\left|\vec{k}_{b_{1}, a_{1}}\right|^{2}\left(m_{b_{1}}^{2}+m_{a_{1}}^{2}\right)\right] \\
& \times \kappa_{i} \Theta\left(m_{g_{3}}-m_{b_{1}}-m_{a_{1}}\right) \tag{96}
\end{align*}
$$

where the $\kappa_{i} \mathrm{~s}$ are listed in Table XXIX.

## V. DISCUSSION OF THE CONSTRUCTION OF THE MODEL

In this section we discuss the interpretation of Eq. (22), the interaction Lagrangian $\mathcal{L}_{W, \text { total }}$, which was constructed by keeping only the lowest possible number of derivatives for a given interaction term. We justify this choice by following two different lines of argumentation. First, our effective model naturally emerges if the Lagrangian $\mathcal{L}_{W, \text { total }}$ is interpreted as being part of a more general and complete chiral hadronic model in the vacuum, such as the (extended) linear sigma model [57], which was already mentioned in the Introduction. Second, a (Functional) Renormalization Group perspective also indicates that terms with the lowest derivatives might be retained as the dominant contributions within our effective hadronic model.

The ELSM is built under the assumption of chiral symmetry, thus

$$
U_{\mathrm{L}}(3) \times U_{\mathrm{R}}(3)=S U_{\mathrm{L}}(3) \times S U_{\mathrm{R}}(3) \times U_{\mathrm{L}}(1) \times U_{\mathrm{R}}(1)
$$

and the assumption of scale invariance. Besides the small explicit breaking of both of these symmetries, because of the nonzero bare quark masses, which derive from the electro-weak sector, the symmetries are also broken dynamically. The $S U_{\mathrm{L}}(3) \times S U_{\mathrm{R}}(3)$ breaks down to a $S U_{\mathrm{V}}(3)$ symmetry spontaneously, which was already discussed before. Also the scale invariance breaks down via gluonic quantum fluctuations and the emergence of a gluon condensate, which is also called the "conformal (trace)
anomaly" $[10,118,119]$. Additionally, a second anomalythe chiral anomaly-breaks $U_{\mathrm{A}}(1)$ and explains the heavy mass of $\eta^{\prime}(958)$ as well as the large pseudoscalar mixing angle [84]. As discussed in Ref. [81] and commented in various parts of this work, this anomaly may affect also the masses and the decays of other nonets (those being part of homochiral multiplets). In the chiral limit, and ignoring the chiral anomaly, all the interaction terms of the ELSM are invariant under scale/dilatation transformations. Consequently, they are parametrized by dimensionless coupling constants.

The Lagrangian $\mathcal{L}_{w_{3}, \text { total }}(22)$, which we use in this work, contains various decay constants, see Table III. With the exception of the last entry, all of them are dimensionful. At a first sight, this feature seems in disagreement with the requirement that should descend from a chiral model such as the ELSM $[53,56]$. Yet, this is not the case and a closer inspection also shows why the lowest number of derivatives should be kept. First, let us recall that the spontaneous symmetry breaking implies a peculiar scalar-axial-vector mixing, which is removed by the a shift of the axial-vector nonet $A_{1}$. In the simplified case with vanishing bare quark masses, where $U_{\mathrm{V}}(3)$ is exact (this is enough for our illustrative purposes), it takes the following form [53,120]:

$$
\begin{equation*}
A_{1}^{\mu} \mapsto A_{1}^{\mu}+Z w \partial^{\mu} P \tag{97}
\end{equation*}
$$

where $Z \approx 1.6$ and $w \approx \frac{g_{1} \phi_{N}}{m_{a_{1}}}$, where $g_{1} \approx g_{\rho} \approx 5.5$, $m_{a_{1}} \simeq 1.4 \mathrm{GeV}$, and $\phi_{N} \approx Z f_{\pi} \quad\left(f_{\pi}=92.4 \mathrm{MeV}\right.$ is the pion decay constant).

Now, the ELSM was studied in a variety of frameworks: besides (pseudo)scalar and (axial-)vector mesons [53] and the aforementioned calculation of glueballs [56,79,117], it was also applied to study hybrid mesons [120] and excited scalar mesons [121]. Even if various mesonic nonets were not included in the ELSM yet, it is clear which features such terms should have.

Within this framework, the $W_{3} P P$ Lagrangian,

$$
\begin{equation*}
\mathcal{L}_{w_{3} p p}=g_{w_{3} p p} \operatorname{tr}\left[W^{\mu \nu \rho}\left[P,\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} P\right)\right]_{-}\right], \tag{98}
\end{equation*}
$$

where $g_{w_{3} p p}$ has dimension $E^{-2}$, and the $W_{3} A_{1} P$ Lagrangian,

$$
\begin{equation*}
\mathcal{L}_{w_{3} a_{1} p}=g_{w_{3} a_{1} p} \operatorname{tr}\left[W^{\mu \nu \rho}\left[A_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right]_{-}\right] \tag{99}
\end{equation*}
$$

where $g_{w_{3} a_{1} p}$ has dimension $E^{-1}$, can be seen as the result of the interaction Lagrangian

$$
\begin{equation*}
\mathcal{L}_{w_{3} a_{1} a_{1}}=g_{w_{3} a_{1} a_{1}} \operatorname{tr}\left(W^{\mu \nu \rho}\left[A_{1, \mu}, \partial_{\nu} A_{1, \rho}\right]_{-}\right) \tag{100}
\end{equation*}
$$

involving solely the dimensionless coupling $g_{w_{3} a_{1} a_{1}}$ via the shift of Eq. (97) applied once/twice. In other words, only the Lagrangian $\mathcal{L}_{w_{3} a_{1} a_{1}}$ is part of a generalized ELSM Lagrangian that includes fields with $J=3$ in agreement
with dilatation invariance. Then, upon spontaneous symmetry breaking and mixing/shifts, the Lagrangians $\mathcal{L}_{w_{3} a_{1} p}$ and $\mathcal{L}_{w_{3} p p}$ are a consequence of $\mathcal{L}_{w_{3} a_{1} a_{1}}$, thus explaining how dimensional couplings appear in the model, even if one starts solely with dimensionless ones.

While the decay $W_{3} \rightarrow A_{1} A_{1}$ is not kinematically allowed, the corresponding term explains how the first and the fifth entries of Table III emerge. The simple relations

$$
\begin{equation*}
g_{w_{3} p p} \approx Z^{2} w^{2} g_{w_{3} a_{1} a_{1}} \tag{101}
\end{equation*}
$$

and

$$
\begin{equation*}
g_{w_{3} a_{1} p} \approx Z w g_{w_{3} a_{1} a_{1}} \tag{102}
\end{equation*}
$$

follow. Moreover, the ratio $g_{w_{3} a_{1} p} / g_{w_{3} p p} \approx(Z w)^{-1}$, out of which we may estimate that

$$
\begin{equation*}
g_{w_{3} a_{1} p}^{2} \approx 3 \times 10^{-4} \mathrm{MeV}^{-2} \tag{103}
\end{equation*}
$$

is obtained (it should be however stressed that the present heuristic discussion cannot provide a precise determination of ratios of couplings, but it solely offers a guide to understand the origin of the model). This value leads to $\Gamma_{\rho_{3}(1690) \rightarrow a_{1}(1260) \pi} \approx 2 \mathrm{MeV}$, as previously mentioned.

Next, let us consider the $W_{3} B_{1} P$ Lagrangian,

$$
\begin{equation*}
\mathcal{L}_{w_{3} b_{1} p}=g_{w_{3} b_{1} p} \operatorname{tr}\left(W^{\mu \nu \rho}\left\{B_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right\}_{+}\right), \tag{104}
\end{equation*}
$$

with $g_{w_{3} b_{1} p}$, which can be seen as emerging from

$$
\begin{equation*}
\mathcal{L}_{w_{3} b_{1} a_{1}}=g_{w_{3} b_{1} a_{1}} \operatorname{tr}\left[W^{\mu \nu \rho}\left\{B_{1, \mu},\left(\partial_{\nu} A_{1, \rho}\right)\right\}_{+}\right] \tag{105}
\end{equation*}
$$

with a dimensionless coupling $g_{w_{3} b_{1} a_{1}}$. The $A_{1}$ shift introduced above implies that

$$
\begin{equation*}
g_{w_{3} b_{1} p} \approx Z w g_{w_{3} b_{1} a_{1}} \tag{106}
\end{equation*}
$$

Next, let us study the remaining terms. The $W_{3} V_{1} V_{1}$ Lagrangian is already dilatation invariant, while the two other Lagrangians $\mathcal{L}_{w_{3} v_{1} p}$ and $\mathcal{L}_{w_{3} a_{2} p}$ are not, since they involve the coupling $g_{w_{3} v_{1} p}$ and $g_{w_{3} a_{2} p}$, having dimensions $E^{-3}$ and $E^{-2}$. In this case, one could build analogous Lagrangians $\mathcal{L}_{w_{3} v_{1} a_{1}}$ and $\mathcal{L}_{w_{3} a_{2} a_{1}}$, but their coupling constants $g_{w_{3} v_{1} a_{1}}$ and $g_{w_{3} a_{2} a_{1}}$ still have dimensions $E^{-2}$ and $E^{-1}$. This is however understandable, since these terms involve the Levi-Civita pseudotensor $\varepsilon_{\mu \nu \rho \sigma}$ and hence are a manifestation of the chiral anomaly.

In conclusion, terms being part of the model are (i) dilatation invariant, (ii) can be obtained from dilatation invariant terms of a more general underlying chiral model via the shift of the axial-vector nonet, or (iii) they are linked
to the chiral anomaly. Now, one also understand why only terms with the lowest number of derivatives appear.

For the effective interaction terms of the $3^{--}$glueball, the argument is completely analogous.

Still, we also provide a brief alternative motivation for the construction of the model from a renormalization group perspective. In this framework, we know that the UV limit of all hadronic physics is QCD, where the degrees of freedom are quarks and gluons. However, we want to describe low-energy QCD and the effective hadronic degrees of freedom. By integrating out quantum fluctuations of gluons and quarks via renormalization group transformations from the UV scales to the IR scales, including dynamical hadronization, one finally ends up with an effective hadronic theory, where hadrons are the effective degrees of freedom. While integrating out quantum fluctuations momentum shell after momentum shell, all kinds of effective couplings/vertices, which are in accordance with the UV symmetries of the theory (the symmetries of QCD), will be generated. This process is effectively described as the RG flow of the coupling constants of the theory, where effective couplings of hadrons are initially zero and only generated dynamically during the RG flow. The same applies to the effective fermionic and hadronic quantum fields.

An efficient and modern framework to describe this process, which effectively is an implementation of Wilson's idea of the RG [122-124], is the FRG in its formulation via the exact renormalization group equation [60,61], see for example Refs. [58,59,63,125-127] and Ref. [62] for an comprehensive up-to-date review. For our purpose, it is solely important to note that the FRG is most efficiently formulated on the level of the quantum effective IR action $\Gamma[\Phi]$, which is the generating functional of all 1PI-n-pointcorrelation functions. The FRG allows us to calculate $\Gamma[\Phi]$-at least within certain truncations-from the UV theory, which here is QCD. The interesting part is that the quantum effective IR action contains all kinds of effective vertices, which were generated during the RG flow and contain all information about the higher energy scales. For the calculation of decays etc., however, only these effective couplings are relevant and calculations are performed at tree level because all loop contributions are already contained in the couplings and because we are working with the 1PI- $n$-point-correlation functions. Dynamical symmetry breaking is realized in this framework as a nontrivial vacuum/ground state, which is realized as a minimum of the effective action $\Gamma[\Phi]$. All 1PI-n-point functions must be extracted at this physical point, which yields shifts like in Eq. (97).

In practice, it is almost impossible to really do this calculation from first principles, especially concerning couplings for mesons with higher spin and large masses. Nevertheless, we know from the previous argumentation that a low IR-effective action $\Gamma[\Phi]$ must contain all possible
interaction terms, which are in accordance with the symmetry of the theory, and if shifted to the physical point, then it must contain all symmetries of the vacuum of the theory. This also applies to the effective couplings of our spin-3 mesons. Thus, we can use the following argument: instead of calculating all interaction terms and couplings, which are in accordance with the symmetries of QCD, via an RG flow from first principles (which is anyhow impossible), we just write down all possible effective interaction terms that are invariant under the residual IR vacuum symmetries of QCD and simply fit the effective coupling constants to experimental data. This approach additionally justifies the use of tree-level calculations, because experimental measurements for decay widths etc. already contain all information about the high energy scales of QCD and so does an effective IR action $\Gamma[\Phi]$, from which information is extracted at tree level. We therefore interpret all effective Lagrangians $\mathcal{L}_{w_{3} \circ \circ}$ as being part of an effective IR action $\Gamma[\Phi]$ for QCD .

Last, we argue, that as long as the momentum exchange in processes like three-point-interactions (decays) is small, we can expand an effective IR action in powers of momenta of the effective hadronic fields. In the context of the FRG, this approach is denoted as a "derivative expansion" and turned out as a decent approximation in a lot of low-energy models of QCD [64-75].

All in all, both approaches complement each other and explain that we only retain terms that are of the lowest orders in the derivatives of the fields, that are of leading order in a large $-N_{\mathrm{c}}$ expansion, and that are flavor invariant.

## VI. CONCLUSIONS

In this paper, we have studied the decays of the lightest mesonic $\bar{q} q$ nonet with quantum numbers $J^{P C}=3^{--}$using an effective QFT model based on flavor symmetry. Our model retained only the dominant terms in an large- $N_{c}$ expansion and the lowest possible number of derivatives. By comparing the theoretical results with the current status of experimental data for decay width and some known branching ratios, which are all reported by PDG [1], we conclude that the $\bar{q} q$ assignment works quite well. Still, we remark that there are also other rather successful approaches different from the $\bar{q} q$ picture towards a coherent description of the nature of the $J^{P C}=3^{--}$states, see, e.g., via multi- $\rho(770)$ resonances $[128,129]$. However, also in our work some of the decay channels deserve deeper theoretical and experimental investigation. Additionally, we presented decay ratios of a putative and not yet detected $J^{P C}=3^{--}$glueball by considering decays into the vectorpseudoscalar and pseudovector-pseudoscalar mesonic pairs. In summary, we provide many qualitative predictions of strong and radiative decays of conventional $J^{P C}=3^{--}$ mesons and the $J^{P C}=3^{--}$glueball for future experimental tests.

## ACKNOWLEDGMENTS

S. J. acknowledges financial support through the project Development Accelerator of the Jan Kochanowski University of Kielce, cofinanced by the European Union under the European Social Fund, with Grant No. POWR.03.05. 00-00-Z212 /18. A. K. acknowledges the support of the Deutsche Forschungsgemeinschaft (DFG, German Research Foundation) through the collaborative research center trans-region CRC-TR 211 Stronginteraction matter under extreme conditions: Project No. $315477589 —$ TRR 211. A. K. also acknowledges the support of the Friedrich-Naumann-Foundation for Freedom, the Giersch Foundation, and the Helmholtz Graduate School for Hadron and Ion Research. F. G. acknowledges support from the Polish National Science Centre (NCN) through the OPUS Projects No. 2019/33/B/ST2/00613 and No. 2018/29/ B/ST2/02576. The authors thank J. Braun, F. Divotgey, J. Eser, M. Piotrowska, D. H. Rischke, and M. J. Steil for useful discussions.

## APPENDIX A: MASSES AND STRANGENONSTRANGE MIXING ANGLES

For a certain nonet, we denote as $\vec{a}$ the triplet of fields describing the isospin $I=1$ elements $\bar{d} u, \bar{u} d, \frac{1}{\sqrt{2}}(\bar{u} u-\bar{d} d)$, with $K^{ \pm}$the complex kaonic fields representing the doublet $\bar{s} u$ and $\bar{u} s$, and $\bar{K}^{0}$ the analogous fields $\bar{d} s$ and $\bar{s} d$ for the neutral kaonic elements, and as $f_{8}$ and $f_{0}$ the octet and the singlet elements, whose quark-antiquark configurations are $\frac{1}{\sqrt{6}}(\bar{u} u+\bar{d} d-2 \bar{s} s)$ and $\frac{1}{\sqrt{3}}(\bar{u} u+\bar{d} d+\bar{s} s)$, respectively. The corresponding $S U_{\mathrm{V}}(3)$-invariant mass term reads (omitting possible Lorentz indices for simplicity)

$$
\begin{align*}
& \frac{1}{2} m_{a}^{2} \vec{a}^{2}+m_{K}^{2}\left(K^{+} K^{-}+\bar{K}^{0} K^{0}\right) \\
& \quad+\frac{1}{2} m_{8}^{2} f_{8}^{2}+\frac{1}{2} m_{0}^{2} f_{0}^{2}+m_{08} f_{0} f_{8} \tag{A1}
\end{align*}
$$

where also a singlet-octet mixing parameter $m_{08}$ has been introduced. The masses of the fields $\vec{a}, \bar{K}^{0 / \pm}$, and $f_{8}$ are expressed as

$$
\begin{gather*}
m_{a}^{2}=m_{\mathrm{nonet}}^{2}+2 \delta_{n}^{2}  \tag{A2}\\
m_{K}^{2}=m_{\mathrm{nonet}}^{2}+\delta_{n}^{2}+\delta_{s}^{2},  \tag{A3}\\
m_{8}^{2}=m_{\mathrm{nonet}}^{2}+\frac{2}{3} \delta_{n}^{2}+\frac{4}{3} \delta_{s}^{2}, \tag{A4}
\end{gather*}
$$

where $m_{\text {nonet }}$ is the mass of the nonet in the chiral limit as a consequence of spontaneous symmetry breaking (see, e.g., Ref. [53]), and $\delta_{n}^{2}$ represents the (almost negligible) contribution of the bare nonstrange quarks $u$ and $d$ and $\delta_{s}^{2}$ the (on hadronic scales) small but nonnegligible contribution of the bare quark $s$. (Besides the pseudoscalar
nonet, one simply approximates $\delta_{n} \propto m_{u} \simeq m_{d}$ and $\delta_{s} \propto m_{s}$. ) From the expressions above it follows that

$$
\begin{align*}
m_{8}^{2} & =\frac{4}{3} m_{K}^{2}-\frac{1}{3} m_{a}^{2} \\
& =m_{\text {nonet }}^{2}+\frac{2}{3} \delta_{n}^{2}+\frac{4}{3} \delta_{s}^{2} . \tag{A5}
\end{align*}
$$

The mass of the singlet member of the nonet may contain an additional unknown contribution (due to the anomaly and/or conversion to gluons), parametrized by

$$
\begin{align*}
m_{0}^{2} & =m_{\text {nonet }}^{2}+\frac{4}{3} \delta_{n}^{2}+\frac{2}{3} \delta_{s}^{2}+\alpha \\
& =\frac{2}{3} m_{K}^{2}+\frac{1}{3} m_{a}^{2}+\alpha \tag{A6}
\end{align*}
$$

Next, we diagonalize the system in the isoscalar sector. We then introduce the physical fields $f$ and $f^{\prime}$ as

$$
\binom{-f^{\prime}}{f}=\left(\begin{array}{cc}
\cos \beta_{\mathrm{PDG}} & -\sin \beta_{\mathrm{PDG}}  \tag{A7}\\
\sin \beta_{\mathrm{PDG}} & \cos \beta_{\mathrm{PDG}}
\end{array}\right)\binom{f_{8}}{f_{0}}
$$

where $\beta_{\mathrm{PDG}}$ is the corresponding mixing angle,

$$
\begin{equation*}
\tan \left(2 \beta_{\mathrm{PDG}}\right)=\frac{2 m_{08}}{m_{0}^{2}-m_{8}^{2}} \tag{A8}
\end{equation*}
$$

which is obtained via diagonalization of the mass matrix together with the relation,

$$
\begin{align*}
& m_{8}^{2}=m_{f^{\prime}}^{2} \cos ^{2} \beta_{\mathrm{PDG}}+m_{f}^{2} \sin ^{2} \beta_{\mathrm{PDG}}  \tag{A9}\\
& m_{0}^{2}=m_{f}^{2} \cos ^{2} \beta_{\mathrm{PDG}}+m_{f^{\prime}}^{2} \sin ^{2} \beta_{\mathrm{PDG}} \tag{A10}
\end{align*}
$$

In Eq. (A7) the extra minus sign in front of $f^{\prime}$ is introduced for later convenience. Using Eq. (A9) one gets

$$
\begin{align*}
\frac{3 m_{8}^{2}-3 m_{f^{\prime}}^{2}}{-3 m_{8}^{2}+3 m_{f}^{2}} & =\frac{3 m_{f^{\prime}}^{2} \cos ^{2} \beta_{\mathrm{PDG}}+3 m_{f}^{2} \sin ^{2} \beta_{\mathrm{PDG}}-3 m_{f^{\prime}}^{2}}{-3 m_{f^{\prime}}^{2} \cos ^{2} \beta_{\mathrm{PDG}}-3 m_{f}^{2} \sin ^{2} \beta_{\mathrm{PDG}}+3 m_{f}^{2}} \\
& =\tan ^{2} \beta_{\mathrm{PDG}} \tag{A11}
\end{align*}
$$

but on the other hand using the above expressions for the masses the lhs can be written as

$$
\begin{equation*}
\frac{3 m_{8}^{2}-3 m_{f^{\prime}}^{2}}{-3 m_{8}^{2}+3 m_{f}^{2}}=\frac{4 m_{K}^{2}-m_{a}^{2}-3 m_{f^{\prime}}^{2}}{-4 m_{K}^{2}+m_{a}+3 m_{f}^{2}} \tag{A12}
\end{equation*}
$$

leading to expressions used in the PDG [1] to calculate the mixing angle in the singlet-octet basis:

$$
\begin{equation*}
\tan ^{2} \beta_{\mathrm{PDG}}=\frac{4 m_{k}^{2}-m_{a}^{2}-3 m_{f^{\prime}}^{2}}{-4 m_{K}^{2}+m_{a}+3 m_{f}^{2}} \tag{A13}
\end{equation*}
$$

In the last step we relate this to the (for our purposes) more convenient nonstrange-strange basis $f_{N}$ and $f_{S}$ that corresponds to the configurations $\frac{1}{\sqrt{2}}(\bar{u} u+\bar{d} d)$ and $\bar{s} s$. Via

$$
\binom{f_{8}}{f_{0}}=\frac{1}{\sqrt{3}}\left(\begin{array}{cc}
-\sqrt{2} & 1  \tag{A14}\\
1 & \sqrt{2}
\end{array}\right)\binom{f_{S}}{f_{N}}
$$

we introduce the strange-nonstrange mixing $\beta$ as

$$
\begin{align*}
\binom{-f^{\prime}}{f} & =\frac{1}{\sqrt{3}}\left(\begin{array}{cc}
\cos \beta_{\mathrm{PDG}} & -\sin \beta_{\mathrm{PDG}} \\
\sin \beta_{\mathrm{PDG}} & \cos \beta_{\mathrm{PDG}}
\end{array}\right)\left(\begin{array}{cc}
-\sqrt{2} & 1 \\
1 & \sqrt{2}
\end{array}\right)\binom{f_{S}}{f_{N}} \\
& =\left(\begin{array}{cc}
-\cos \beta & \sin \beta \\
\sin \beta & \cos \beta
\end{array}\right)\binom{f_{S}}{f_{N}}, \tag{A15}
\end{align*}
$$

or, as reported in the main text, as

$$
\binom{f}{f^{\prime}}=\left(\begin{array}{cc}
\cos \beta & \sin \beta  \tag{A16}\\
-\sin \beta & \cos \beta
\end{array}\right)\binom{f_{N}}{f_{S}}
$$

The strange-nonstrange mixing angle $\beta$ and the angle $\beta_{\mathrm{PDG}}$ can be obtained by Eq. (A15) as

$$
\begin{align*}
& \sqrt{\frac{1}{3}} \cos \beta_{\mathrm{PDG}}-\sqrt{\frac{2}{3}} \sin \beta_{\mathrm{PDG}} \\
& \quad=\sin \beta \equiv \sin \left(-\beta_{\mathrm{PDG}}+\beta_{0}\right) \tag{A17}
\end{align*}
$$

thus

$$
\begin{equation*}
\beta=-\beta_{\mathrm{PDG}}+\beta_{0} \tag{A18}
\end{equation*}
$$

with

$$
\begin{equation*}
\beta_{0}=\arccos \sqrt{\frac{2}{3}} \simeq 35.3^{\circ} \tag{A19}
\end{equation*}
$$

The mixing angle $\beta_{\text {PDG }}$ (and hence $\beta$ ) has been calculated in the PDG for certain well-known mesonic nonets (pseudoscalar, vector, as well as $J^{P C}=2^{++}$and $J^{P C}=3^{--}$). While for the pseudoscalar case we employ the result of Ref. [80] (where not only masses but also decays are used to get $\beta$ ), in the remaining three cases we used the value $\beta_{\text {PDG }}$ to determine the strange-nonstrange mixing angle. With the exception of the peculiar pseudoscalar nonet, where the axial anomaly is large, the other values of $\beta$ calculated from $\beta_{\text {PDG }}$ are quite small-in agreement with the theoretical expectations presented in Ref. [81]. As a consequence, the parameter $\alpha$ of Eq. (A6) is negligible. In particular, for $J=3$ the numerical value $\beta=\beta_{w_{3}}=3.5^{\circ}$ has been used throughout our calculations. When future experimental data will be more accurate, one could also include $\beta_{w_{3}}$ as a fit parameter in an improved model. In this respect, better experimental data for radiative decays would be useful.

## APPENDIX B: EXTENDED FORM OF THE LAGRANGIANS OF THE MODEL

In this Appendix, we provide the explicit form of the interaction Lagrangians in Tables III and XIII.

$$
\begin{align*}
& \mathcal{L}_{w_{3} p p}=g_{w_{3} p p} \operatorname{tr}\left(W_{3}^{\mu \nu \rho}\left[P,\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} P\right)\right]_{-}\right) \\
& =\frac{g_{w_{3} p p}}{4}\left(+\rho_{3}^{0, \mu \nu \rho}\left\{+\bar{K}^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{0}\right)-K^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)+K^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{-}\right)\right.\right. \\
& \left.-K^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{+}\right)+2\left[\pi^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{-}\right)-\pi^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{+}\right)\right]\right\} \\
& +\rho_{3}^{+, \mu \nu \rho}\left\{\sqrt{2}\left[-K^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{0}\right)+K^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{-}\right)\right]+2\left[\pi^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{0}\right)-\pi^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{-}\right)\right]\right\} \\
& +\rho_{3}^{-, \mu \nu \rho}\left\{\sqrt{2}\left[K^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)-\bar{K}^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{+}\right)\right]+2\left[-\pi^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{0}\right)+\pi^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{+}\right)\right]\right\} \\
& +\bar{K}_{3}^{* 0, \mu \nu \rho}\left\{+K^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{0}\right)-\pi^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{0}\right)+\sqrt{2}\left[-K^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{-}\right)+\pi^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{+}\right)\right]\right. \\
& +K^{0}\left[\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \eta\right)\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.-\left[\eta\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\eta^{\prime}\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right]\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{0}\right)\right\} \\
& +K_{3}^{* 0, \mu \nu \rho}\left\{-\bar{K}^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{0}\right)+\pi^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)+\sqrt{2}\left[K^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{+}-\pi^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{-}\right)\right]\right.\right. \\
& -\bar{K}^{0}\left[\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \eta\right)\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\eta\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\eta^{\prime}\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right]\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)\right\} \\
& +K_{3}^{*+, \mu \nu \rho}\left\{+K^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{0}\right)-\pi^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{-}\right)+\sqrt{2}\left[\bar{K}^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{-}\right)-\pi^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)\right]\right. \\
& -K^{-}\left[\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \eta\right)\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\eta\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\eta^{\prime}\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right]\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{-}\right)\right\} \\
& +K_{3}^{*-, \mu \nu \rho}\left\{-K^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{0}\right)+\pi^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{+}\right)+\sqrt{2}\left[-K^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \pi^{+}\right)+\pi^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{0}\right)\right]\right. \\
& +K^{+}\left[\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \eta\right)\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.-\left[\eta\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\eta^{\prime}\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right]\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{+}\right)\right\} \\
& +\omega_{3}^{\mu \nu \rho}\left\{\left(-\cos \beta_{w_{3}}+\sqrt{2} \sin \beta_{w_{3}}\right) \times\left[\bar{K}^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{0}\right)-K^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)-K^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{-}\right)+K^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{+}\right)\right]\right\} \\
& \left.+\phi_{3}^{\mu \nu \rho}\left\{\left(\sin \beta_{w_{3}}+\sqrt{2} \cos \beta_{w_{3}}\right) \times\left[\bar{K}^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{0}\right)-K^{0}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)-K^{+}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{-}\right)+K^{-}\left(\partial_{\mu} \partial_{\nu} \partial_{\rho} K^{+}\right)\right]\right\}\right) . \tag{B1}
\end{align*}
$$

$$
\begin{aligned}
\mathcal{L}_{w_{3} v_{1} p}= & g_{w_{3} v_{1} p} \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[W_{3, \mu \alpha \beta}\left\{\left(\partial_{\nu} V_{1, \rho}\right),\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right] \\
= & \frac{g_{w_{3} v_{1} p}}{4} \varepsilon^{\mu \nu \rho \sigma}\left(+\rho_{3, \mu \alpha \beta}^{0}\left\{-\left(\partial_{\nu} \bar{K}_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{0}\right)-\left(\partial_{\nu} K_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \bar{K}^{0}\right)+\left(\partial_{\nu} K_{\rho}^{*+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{-}\right)+\left(\partial_{\nu} K_{\rho}^{*-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{+}\right)\right.\right. \\
& \left.+2\left[\left(\partial_{\nu} \omega_{\rho}\right) \cos \beta_{v}-\left(\partial_{\nu} \phi_{\rho}\right) \sin \beta_{v}\right]\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+2\left(\partial_{\nu} \rho_{\rho}^{0}\right)\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right) \cos \beta_{p}-\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right) \sin \beta_{p}\right]\right\} \\
& +\rho_{3, \mu \alpha \beta}^{+}\left\{\sqrt{2}\left[\left(\partial_{\nu} K_{\rho}^{*-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{0}\right)+\left(\partial_{\nu} K_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{-}\right)\right]+2\left[\left(\partial_{\nu} \omega_{\rho}\right) \cos \beta_{v}-\left(\partial_{\nu} \phi_{\rho}\right) \sin \beta_{v}\right]\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{-}\right)\right. \\
& \left.+2\left(\partial_{\nu} \rho_{\rho}^{-}\right)\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right) \cos \beta_{p}-\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right) \sin \beta_{p}\right]\right\}+\rho_{3, \mu \alpha \beta}^{-}\left\{\sqrt{2}\left[\left(\partial_{\nu} K_{\rho}^{*+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \bar{K}^{0}\right)+\left(\partial_{\nu} \bar{K}_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{+}\right)\right]\right. \\
& \left.+2\left[\left(\partial_{\nu} \omega_{\rho}\right) \cos \beta_{v}-\left(\partial_{\nu} \phi_{\rho}\right) \sin \beta_{v}\right]\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{+}\right)+2\left(\partial_{\nu} \rho_{\rho}^{+}\right)\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right) \cos \beta_{p}-\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right) \sin \beta_{p}\right]\right\} \\
& +\bar{K}_{3, \mu \alpha \beta}^{* 0}\left\{-\left(\partial_{\nu} K_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+\sqrt{2}\left(\partial_{\nu} K_{\rho}^{*+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{-}\right)-\left(\partial_{\nu} \rho_{\rho}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{0}\right)+\sqrt{2}\left(\partial_{\nu} \rho_{\rho}^{-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{+}\right)\right. \\
& +\left(\partial_{\nu} K_{\rho}^{* 0}\right)\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right)\left(\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right)\left(-\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\left(\partial_{\nu} \omega_{\rho}\right)\left(\cos \beta_{v}+\sqrt{2} \sin \beta_{v}\right)+\left(\partial_{\nu} \phi_{\rho}\right)\left(-\sin \beta_{v}+\sqrt{2} \cos \beta_{v}\right)\right]\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{0}\right)\right\}
\end{aligned}
$$

$$
\begin{align*}
& +K_{3, \mu \alpha \beta}^{* 0}\left\{-\left(\partial_{\nu} \bar{K}_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+\sqrt{2}\left(\partial_{\nu} K_{\rho}^{*-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{+}\right)-\left(\partial_{\nu} \rho_{\rho}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \bar{K}^{0}\right)+\sqrt{2}\left(\partial_{\nu} \rho_{\rho}^{+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{-}\right)\right. \\
& +\left(\partial_{\nu} \bar{K}_{\rho}^{* 0}\right)\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right)\left(\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right)\left(-\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\left(\partial_{\nu} \omega_{\rho}\right)\left(\cos \beta_{v}+\sqrt{2} \sin \beta_{v}\right)+\left(\partial_{\nu} \phi_{\rho}\right)\left(-\sin \beta_{v}+\sqrt{2} \cos \beta_{v}\right)\right]\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \bar{K}^{0}\right)\right\} \\
& +K_{3, \mu \alpha \beta}^{*+}\left\{+\left(\partial_{\nu} K_{\rho}^{*-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+\sqrt{2}\left(\partial_{\nu} \bar{K}_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{-}\right)-\right. \\
& +\left(\partial_{\nu} \rho_{\rho}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{-}\right)+\sqrt{2}\left(\partial_{\nu} \rho_{\rho}^{-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \bar{K}^{0}\right) \\
& +\left(\partial_{\nu} K_{\rho}^{*-}\right)\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right)\left(\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right)\left(-\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\left(\partial_{\nu} \omega_{\rho}\right)\left(\cos \beta_{v}+\sqrt{2} \sin \beta_{v}\right)+\left(\partial_{\nu} \phi_{\rho}\right)\left(-\sin \beta_{v}+\sqrt{2} \cos \beta_{v}\right)\right]\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{-}\right)\right\} \\
& +K_{3, \mu \alpha \beta}^{*-}\left\{+\left(\partial_{\nu} K_{\rho}^{*+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+\sqrt{2}\left(\partial_{\nu} K_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{+}\right)-\right. \\
& +\left(\partial_{\nu} \rho_{\rho}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{+}\right)+\sqrt{2}\left(\partial_{\nu} \rho_{\rho}^{+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{0}\right) \\
& +\left(\partial_{\nu} K_{\rho}^{*+}\right)\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right)\left(\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right)\left(-\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\left(\partial_{\nu} \omega_{\rho}\right)\left(\cos \beta_{v}+\sqrt{2} \sin \beta_{v}\right)+\left(\partial_{\nu} \phi_{\rho}\right)\left(-\sin \beta_{v}+\sqrt{2} \cos \beta_{v}\right)\right]\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{+}\right)\right\} \\
& +\omega_{3, \mu \alpha \beta}\left\{\left(\cos \beta_{w_{3}}+\sqrt{2} \sin \beta_{w_{3}}\right) \times\left[\left(\partial_{\nu} \bar{K}_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{0}\right)+\left(\partial_{\nu} K_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \bar{K}^{0}\right)\right.\right. \\
& \left.+\left(\partial_{\nu} K_{\rho}^{*+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{-}\right)+\left(\partial_{\nu} K_{\rho}^{*-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{+}\right)\right] \\
& +2 \cos \beta_{w_{3}}\left[\left(\partial_{\nu} \rho_{\rho}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+\left(\partial_{\nu} \rho_{\rho}^{+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{-}\right)+\left(\partial_{\nu} \rho_{\rho}^{-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{+}\right)\right] \\
& +2 \cos \beta_{w_{3}}\left[\left(\partial_{\nu} \omega_{\rho}\right) \cos \beta_{v}-\left(\partial_{\nu} \phi_{\rho}\right) \sin \beta_{v}\right] \times\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right) \cos \beta_{p}-\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right) \sin \beta_{p}\right] \\
& \left.+2 \sqrt{2} \sin \beta_{w_{3}}\left[\left(\partial_{\nu} \omega_{\rho}\right) \sin \beta_{v}+\left(\partial_{\nu} \phi_{\rho}\right) \cos \beta_{v}\right] \times\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right) \sin \beta_{p}+\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right) \cos \beta_{p}\right]\right\} \\
& +\phi_{3, \mu \alpha \beta}\left\{\left(-\sin \beta_{w_{3}}+\sqrt{2} \cos \beta_{w_{3}}\right) \times\left[\left(\partial_{\nu} \bar{K}_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{0}\right)+\left(\partial_{\nu} K_{\rho}^{* 0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \bar{K}^{0}\right)\right.\right. \\
& \left.+\left(\partial_{\nu} K_{\rho}^{*+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{-}\right)+\left(\partial_{\nu} K_{\rho}^{*-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{+}\right)\right]+ \\
& -2 \sin \beta_{w_{3}}\left[\left(\partial_{\nu} \rho_{\rho}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+\left(\partial_{\nu} \rho_{\rho}^{+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{-}\right)+\left(\partial_{\nu} \rho_{\rho}^{-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{+}\right)\right]+ \\
& -2 \sin \beta_{w_{3}}\left[\left(\partial_{\nu} \omega_{\rho}\right) \cos \beta_{v}-\left(\partial_{\nu} \phi_{\rho}\right) \sin \beta_{v}\right] \times\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right) \cos \beta_{p}-\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right) \sin \beta_{p}\right] \\
& \left.\left.+2 \sqrt{2} \cos \beta_{w_{3}}\left[\left(\partial_{\nu} \omega_{\rho}\right) \sin \beta_{v}+\left(\partial_{\nu} \phi_{\rho}\right) \cos \beta_{\nu}\right] \times\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right) \sin \beta_{p}+\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right) \cos \beta_{p}\right]\right\}\right) . \tag{B2}
\end{align*}
$$

$$
\begin{align*}
\mathcal{L}_{w_{3} \gamma p}= & g_{w_{3} v_{1} p} \frac{e}{g_{\rho}} \varepsilon^{\mu \nu \rho \sigma}\left(\partial_{\nu} a_{\rho}\right) \operatorname{tr}\left[W_{3, \mu \alpha \beta}\left\{Q,\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right] \\
= & \frac{g_{w_{3} v_{1} p}}{6} \frac{e}{g_{\rho}} \varepsilon^{\mu \nu \rho \sigma}\left(\partial_{\nu} a_{\rho}\right)\left(+\rho_{3, \mu \alpha \beta}^{0}\left\{\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+3\left[\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right) \cos \left(\beta_{p}\right)-\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right) \sin \left(\beta_{p}\right)\right]\right\}\right. \\
& +\rho_{3, \mu \alpha \beta}^{+}\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{-}\right)+\rho_{3, \mu \alpha \beta}^{-}\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{+}\right)-2 \bar{K}_{3, \mu \alpha \beta}^{* 0}\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{0}\right)-2 K_{3, \mu \alpha \beta}^{* 0}\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \bar{K}^{0}\right)+\bar{K}_{3, \mu \alpha \beta}^{*+}\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{-}\right) \\
& +K_{3, \mu \alpha \beta}^{*-}\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{+}\right)+\omega_{3, \mu \alpha \beta}\left\{3 \cos \left(\beta_{w_{3}}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+\left[\cos \left(\beta_{w_{3}}\right) \cos \left(\beta_{p}\right)-2 \sin \left(\beta_{w_{3}}\right) \sin \left(\beta_{p}\right)\right] \times\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right)+\right. \\
& \left.-\left[\cos \left(\beta_{w_{3}}\right) \sin \left(\beta_{p}\right)+2 \sin \left(\beta_{w_{3}}\right) \cos \left(\beta_{p}\right)\right] \times\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right)\right\}-\phi_{3, \mu \alpha \beta}\left\{3 \sin \left(\beta_{w_{3}}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)\right. \\
& +\left[\sin \left(\beta_{w_{3}}\right) \cos \left(\beta_{p}\right)+2 \cos \left(\beta_{w_{3}}\right) \sin \left(\beta_{p}\right)\right] \times\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right)+-\left[\sin \left(\beta_{w_{3}}\right) \sin \left(\beta_{p}\right)-2 \cos \left(\beta_{w_{3}}\right) \cos \left(\beta_{p}\right)\right] \\
& \left.\times\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right)\right\} . \tag{B3}
\end{align*}
$$

$$
\begin{align*}
& \mathcal{L}_{w_{3} a_{2} p}=g_{w_{3} a_{2} p} \varepsilon_{\mu \nu \rho \sigma} \operatorname{tr}\left(W_{3, \mu \alpha \beta}\left[\left(\partial^{\nu} A_{2}^{\rho \alpha}\right),\left(\partial^{\sigma} \partial^{\beta} P\right)\right]_{-}\right) \\
& =\frac{g_{w_{3} a_{2} p}}{4} \varepsilon_{\mu \nu \rho \sigma}\left(+\rho_{3 \alpha \beta}^{0, \mu}\left\{+\left(\partial^{\nu} \bar{K}_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{0}\right)-\left(\partial^{\nu} K_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \bar{K}^{0}\right)+\left(\partial^{\nu} K_{2}^{*+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{-}\right)\right.\right. \\
& \left.-\left(\partial^{\nu} K_{2}^{*-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{+}\right)+2\left[\left(\partial^{\nu} a_{2}^{+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{-}\right)-\left(\partial^{\nu} a_{2}^{-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{+}\right)\right]\right\} \\
& +\rho_{3 \alpha \beta}^{+, \mu}\left\{\sqrt{2}\left[-\left(\partial^{\nu} K_{2}^{*-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{0}\right)+\left(\partial^{\nu} K_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{-}\right)\right]+2\left[\left(\partial^{\nu} a_{2}^{-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{0}\right)-\left(\partial^{\nu} a_{2}^{0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{-}\right)\right]\right\} \\
& +\rho_{3 \alpha \beta}^{-, \mu}\left\{\sqrt{2}\left[\left(\partial^{\nu} K_{2}^{*+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \bar{K}^{0}\right)-\left(\partial^{\nu} \bar{K}_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{+}\right)\right]+2\left[-\left(\partial^{\nu} a_{2}^{+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{0}\right)+\left(\partial^{\nu} a_{2}^{0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{+}\right)\right]\right\} \\
& +\bar{K}_{3 \alpha \beta}^{* 0, \mu}\left\{+\left(\partial^{\nu} K_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{0}\right)-\sqrt{2}\left(\partial^{\nu} K_{2}^{*+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{-}\right)-\left(\partial^{\nu} a_{2}^{0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{0}\right)+\sqrt{2}\left(\partial^{\nu} a_{2}^{-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{+}\right)\right. \\
& +\left(\partial^{\nu} K_{2}^{* 0, \rho \alpha}\right)\left[\left(\partial^{\sigma} \partial^{\beta} \eta\right)\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial^{\sigma} \partial^{\beta} \eta^{\prime}\right)\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\left(\partial^{\mu} f_{2}^{\rho \alpha}\right)\left(\cos \beta_{a_{2}}-\sqrt{2} \sin \beta_{a_{2}}\right)+\left(\partial^{\mu} f_{2}^{\rho \alpha \alpha}\right)\left(-\sin \beta_{a_{2}}-\sqrt{2} \cos \beta_{a_{2}}\right)\right]\left(\partial^{\sigma} \partial^{\beta} K^{0}\right)\right\} \\
& +K_{3 \alpha \beta}^{* 0, \mu}\left\{-\left(\partial^{\nu} \bar{K}_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{0}\right)+\sqrt{2}\left(\partial^{\nu} K_{2}^{*-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{+}\right)+\left(\partial^{\nu} a_{2}^{0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \bar{K}^{0}\right)-\sqrt{2}\left(\partial^{\nu} a_{2}^{+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{-}\right)\right. \\
& +\left(\partial^{\nu} \bar{K}_{2}^{* 0, \rho \alpha}\right)\left[\left(\partial^{\sigma} \partial^{\beta} \eta\right)\left(\cos \beta_{p}-\sqrt{2} \sin \beta_{p}\right)+\left(\partial^{\sigma} \partial^{\beta} \eta^{\prime}\right)\left(-\sin \beta_{p}-\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\left(\partial^{\mu} f_{2}^{\rho \alpha}\right)\left(-\cos \beta_{a_{2}}+\sqrt{2} \sin \beta_{a_{2}}\right)+\left(\partial^{\mu} f_{2}^{\prime \rho \alpha}\right)\left(\sin \beta_{a_{2}}+\sqrt{2} \cos \beta_{a_{2}}\right)\right]\left(\partial^{\sigma} \partial^{\beta} \bar{K}^{0}\right)\right\} \\
& +K_{3 \alpha \beta}^{*+, \mu}\left\{+\left(\partial^{\nu} K_{2}^{*-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{0}\right)+\sqrt{2}\left(\partial^{\nu} \bar{K}_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{-}\right)-\left(\partial^{\nu} a_{2}^{0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{-}\right)-\sqrt{2}\left(\partial^{\nu} a_{2}^{-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \bar{K}^{0}\right)\right. \\
& +\left(\partial^{\nu} K_{2}^{*-, \rho \alpha}\right)\left[\left(\partial^{\sigma} \partial^{\beta} \eta\right)\left(\cos \beta_{p}-\sqrt{2} \sin \beta_{p}\right)+\left(\partial^{\sigma} \partial^{\beta} \eta^{\prime}\right)\left(-\sin \beta_{p}-\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\left(\partial^{\mu} f_{2}^{\rho \alpha}\right)\left(-\cos \beta_{a_{2}}+\sqrt{2} \sin \beta_{a_{2}}\right)+\left(\partial^{\mu} f_{2}^{\prime \rho \alpha}\right)\left(\sin \beta_{a_{2}}+\sqrt{2} \cos \beta_{a_{2}}\right)\right]\left(\partial^{\sigma} \partial^{\beta} K^{-}\right)\right\} \\
& +K_{3 \alpha \beta}^{*-, \mu}\left\{-\left(\partial^{\nu} K_{2}^{*+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{0}\right)-\sqrt{2}\left(\partial^{\nu} K_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \pi^{+}\right)+\left(\partial^{\nu} a_{2}^{0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{+}\right)+\sqrt{2}\left(\partial^{\nu} a_{2}^{+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{0}\right)\right. \\
& +\left(\partial^{\nu} K_{2}^{*+, \rho \alpha}\right)\left[\left(\partial^{\sigma} \partial^{\beta} \eta\right)\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial^{\sigma} \partial^{\beta} \eta^{\prime}\right)\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[\left(\partial^{\mu} f_{2}^{\rho \alpha}\right)\left(\cos \beta_{a_{2}}-\sqrt{2} \sin \beta_{a_{2}}\right)+\left(\partial^{\mu} f_{2}^{\prime \rho \alpha}\right)\left(-\sin \beta_{a_{2}}-\sqrt{2} \cos \beta_{a_{2}}\right)\right]\left(\partial^{\sigma} \partial^{\beta} K^{-}\right)\right\} \\
& +\omega_{3, \alpha \beta}^{\mu}\left\{\left(-\cos \beta_{w_{3}}+\sqrt{2} \sin \beta_{w_{3}}\right) \times\left[\left(\partial^{\nu} \bar{K}_{2}^{* 0 . \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{0}\right)-\left(\partial^{\nu} K_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \bar{K}^{0}\right)+\right.\right. \\
& \left.\left.-\left(\partial^{\nu} K_{2}^{*+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{-}\right)+\left(\partial^{\nu} K_{2}^{*-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{+}\right)\right]\right\}+\phi_{3, \alpha \beta}^{\mu}\left\{\left(\sin \beta_{w_{3}}+\sqrt{2} \cos \beta_{w_{3}}\right)\right. \\
& \times\left[\left(\partial^{\nu} \bar{K}_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{0}\right)-\left(\partial^{\nu} K_{2}^{* 0, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} \bar{K}^{0}\right)+\right. \\
& \left.\left.\left.-\left(\partial^{\nu} K_{2}^{*+, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{-}\right)+\left(\partial^{\nu} K_{2}^{*-, \rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} K^{+}\right)\right]\right\}\right) .  \tag{B4}\\
& \mathcal{L}_{w_{3} b_{1} p}=g_{w_{3} b_{1} p} \operatorname{tr}\left[W_{3}^{\mu \nu \rho}\left\{B_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right\}_{+}\right] \\
& =\frac{g_{w_{3} b_{1} p}}{4}\left(+\rho_{3}^{0, \mu \nu \rho}\left\{-\bar{K}_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)-K_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)+K_{1, B, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)+K_{1, B, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right.\right. \\
& \left.+2\left[h_{1, \mu} \cos \beta_{b_{1}}-h_{1, \mu}^{\prime} \sin \beta_{b_{1}}\right]\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+2 b_{1, \mu}^{0}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right) \cos \beta_{p}-\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right) \sin \beta_{p}\right]\right\} \\
& +\rho_{3}^{+, \mu \nu \rho}\left\{\sqrt{2}\left[K_{1, B, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)+K_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)\right]+2\left[h_{1, \mu} \cos \beta_{b_{1}}-h_{1, \mu}^{\prime} \sin \beta_{b_{1}}\right]\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)\right. \\
& \left.+2 b_{1, \mu}^{-}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right) \cos \beta_{p}-\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right) \sin \beta_{p}\right]\right\}+\rho_{3}^{-, \mu \nu \rho}\left\{\sqrt{2}\left[K_{1, B, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)+\bar{K}_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right]\right. \\
& \left.+2\left[h_{1, \mu} \cos \beta_{b_{1}}-h_{1, \mu}^{\prime} \sin \beta_{b_{1}}\right]\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)+2 b_{1, \mu}^{+}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right) \cos \beta_{p}-\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right) \sin \beta_{p}\right]\right\} \\
& +\bar{K}_{3}^{* 0, \mu \nu \rho}\left\{-K_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+\sqrt{2} K_{1, B, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)-b_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)+\sqrt{2} b_{1, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right. \\
& +K_{1, B, \mu}^{0}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right)\left(\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(-\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[h_{1, \mu}\left(\cos \beta_{b_{1}}+\sqrt{2} \sin \beta_{b_{1}}\right)+h_{1, \mu}^{\prime}\left(-\sin \beta_{b_{1}}+\sqrt{2} \cos \beta_{b_{1}}\right)\right]\left(\partial_{\nu} \partial_{\rho} K^{0}\right)\right\} \\
& +K_{3}^{* 0, \mu \nu \rho}\left\{-\bar{K}_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+\sqrt{2} K_{1, B, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)-b_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)+\sqrt{2} b_{1, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)\right. \\
& +\bar{K}_{1, B, \mu}^{0}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right)\left(\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(-\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[h_{1, \mu}\left(\cos \beta_{b_{1}}+\sqrt{2} \sin \beta_{b_{1}}\right)+h_{1, \mu}^{\prime}\left(-\sin \beta_{b_{1}}+\sqrt{2} \cos \beta_{b_{1}}\right)\right]\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)\right\}
\end{align*}
$$

$$
\begin{align*}
& +K_{3}^{*+, \mu \nu \rho}\left\{+K_{1, B, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+\sqrt{2} \bar{K}_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)+b_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)+\sqrt{2} b_{1, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)\right. \\
& +K_{1, B, \mu}^{-}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right)\left(\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(-\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[h_{1, \mu}\left(\cos \beta_{b_{1}}+\sqrt{2} \sin \beta_{b_{1}}\right)+h_{1, \mu}^{\prime}\left(-\sin \beta_{b_{1}}+\sqrt{2} \cos \beta_{b_{1}}\right)\right]\left(\partial_{\nu} \partial_{\rho} K^{-}\right)\right\} \\
& +K_{3}^{*-, \mu \nu \rho}\left\{+K_{1, B, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+\sqrt{2} K_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)+b_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)+\sqrt{2} b_{1, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)\right. \\
& +K_{1, B, \mu}^{+}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right)\left(\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(-\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right] \\
& \left.+\left[h_{1, \mu}\left(\cos \beta_{b_{1}}+\sqrt{2} \sin \beta_{b_{1}}\right)+h_{1, \mu}^{\prime}\left(-\sin \beta_{b_{1}}+\sqrt{2} \cos \beta_{b_{1}}\right)\right]\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right\} \\
& +\omega_{3}^{\mu \nu \rho}\left\{\left(\cos \beta_{w_{3}}+\sqrt{2} \sin \beta_{w_{3}}\right) \times\left[\bar{K}_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)+K_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)+K_{1, B, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)+K_{1, B, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right]\right. \\
& +2 \cos \beta_{w_{3}}\left[b_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+b_{1, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)+b_{1, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)\right] \\
& +2 \cos \beta_{w_{3}}\left[h_{1, \mu} \cos \beta_{b_{1}}-h_{1, \mu}^{\prime} \sin \beta_{b_{1}}\right] \times\left[\left(\partial_{\nu} \partial_{\rho} \eta\right) \cos \beta_{p}-\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right) \sin \beta_{p}\right] \\
& \left.+2 \sqrt{2} \sin \beta_{w_{3}}\left[h_{1, \mu} \sin \beta_{b_{1}}+h_{1, \mu}^{\prime} \cos \beta_{b_{1}}\right] \times\left[\left(\partial_{\nu} \partial_{\rho} \eta\right) \sin \beta_{p}+\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right) \cos \beta_{p}\right]\right\} \\
& +\phi_{3}^{\mu \nu \rho}\left\{\left(-\sin \beta_{w_{3}}+\sqrt{2} \cos \beta_{w_{3}}\right) \times\left[\bar{K}_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)+K_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)+K_{1, B, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)+K_{1, B, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right]+\right. \\
& -2 \sin \beta_{w_{3}}\left[b_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+b_{1, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)+b_{1, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)\right]+ \\
& -2 \sin \beta_{w_{3}}\left[h_{1, \mu} \cos \beta_{b_{1}}-h_{1, \mu}^{\prime} \sin \beta_{b_{1}}\right] \times\left[\left(\partial_{\nu} \partial_{\rho} \eta\right) \cos \beta_{p}-\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right) \sin \beta_{p}\right] \\
& \left.\left.+2 \sqrt{2} \cos \beta_{w_{3}}\left[h_{1, \mu} \sin \beta_{b_{1}}+h_{1, \mu}^{\prime} \cos \beta_{b_{1}}\right] \times\left[\left(\partial_{\nu} \partial_{\rho} \eta\right) \sin \beta_{p}+\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right) \cos \beta_{p}\right]\right\}\right) . \tag{B5}
\end{align*}
$$

$$
\mathcal{L}_{w_{3} a_{1} p}=g_{w_{3} a_{1} p} \operatorname{tr}\left(W_{3}^{\mu \nu \rho}\left[A_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right]_{-}\right)
$$

$$
=\frac{g_{w_{3} a_{1} p}}{4}\left(+\rho_{3}^{0, \mu \nu \rho}\left\{+\bar{K}_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)-K_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)\right.\right.
$$

$$
\left.+K_{1, A, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)-K_{1, A, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)+2\left[a_{1, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)-a_{1, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)\right]\right\}
$$

$$
+\rho_{3}^{+, \mu \nu \rho}\left\{\sqrt{2}\left[-K_{1, A, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)+K_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)\right]+2\left[a_{1, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)-a_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)\right]\right\}
$$

$$
+\rho_{3}^{-, \mu \nu \rho}\left\{\sqrt{2}\left[K_{1, A, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)-\bar{K}_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right]+2\left[-a_{1, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+a_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)\right]\right\}
$$

$$
+\bar{K}_{3}^{* 0, \mu \nu \rho}\left\{+K_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)-\sqrt{2} K_{1, A, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)-a_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)+\sqrt{2} a_{1, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right.
$$

$$
+K_{1, A, \mu}^{0}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right)\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right]
$$

$$
\left.+\left[f_{1, \mu}\left(\cos \beta_{a_{1}}-\sqrt{2} \sin \beta_{a_{1}}\right)+f_{1, \mu}^{\prime}\left(-\sin \beta_{a_{1}}-\sqrt{2} \cos \beta_{a_{1}}\right)\right]\left(\partial_{\nu} \partial_{\rho} K^{0}\right)\right\}
$$

$$
+K_{3}^{* 0, \mu \nu \rho}\left\{-\bar{K}_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+\sqrt{2} K_{1, A, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)+a_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)-\sqrt{2} a_{1, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)+\right.
$$

$$
+\bar{K}_{1, A, \mu}^{0}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right)\left(\cos \beta_{p}-\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(-\sin \beta_{p}-\sqrt{2} \cos \beta_{p}\right)\right]
$$

$$
\left.+\left[f_{1, \mu}\left(-\cos \beta_{a_{1}}+\sqrt{2} \sin \beta_{a_{1}}\right)+f_{1, \mu}^{\prime}\left(\sin \beta_{a_{1}}+\sqrt{2} \cos \beta_{a_{1}}\right)\right]\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)\right\}
$$

$$
+K_{3}^{*+, \mu \nu \rho}\left\{+K_{1, A, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+\sqrt{2} \bar{K}_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)-a_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)-\sqrt{2} a_{1, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)\right.
$$

$$
+K_{1, A, \mu}^{-}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right)\left(\cos \beta_{p}-\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\iota} \partial_{\rho} \eta^{\prime}\right)\left(-\sin \beta_{p}-\sqrt{2} \cos \beta_{p}\right)\right]
$$

$$
\left.+\left[f_{1, \mu}\left(-\cos \beta_{a_{1}}+\sqrt{2} \sin \beta_{a_{1}}\right)+f_{1, \mu}^{\prime}\left(\sin \beta_{a_{1}}+\sqrt{2} \cos \beta_{a_{1}}\right)\right]\left(\partial_{\nu} \partial_{\rho} K^{-}\right)\right\}
$$

$$
+K_{3}^{*-\mu \nu \rho}\left\{-K_{1, A, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)-\sqrt{2} K_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)+a_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)+\sqrt{2} a_{1, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)\right.
$$

$$
+K_{1, A, \mu}^{+}\left[\left(\partial_{\nu} \partial_{\rho} \eta\right)\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)+\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right]
$$

$$
\left.+\left[f_{1, \mu}\left(\cos \beta_{a_{1}}-\sqrt{2} \sin \beta_{a_{1}}\right)+f_{1, \mu}^{\prime}\left(-\sin \beta_{a_{1}}-\sqrt{2} \cos \beta_{a_{1}}\right)\right]\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right\}
$$

$$
+\omega_{3}^{\mu \nu \rho}\left\{\left(-\cos \beta_{w_{3}}+\sqrt{2} \sin \beta_{w_{3}}\right) \times\left[\bar{K}_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)-K_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)-K_{1, A, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)+K_{1, A, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right]\right\}
$$

$$
\begin{equation*}
\left.+\phi_{3}^{\mu \nu \rho}\left\{\left(\sin \beta_{w_{3}}+\sqrt{2} \cos \beta_{w_{3}}\right) \times\left[\bar{K}_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)-K_{1, A, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)-K_{1, A, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)+K_{1, A, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right]\right\}\right) . \tag{B6}
\end{equation*}
$$

$$
\begin{align*}
& \mathcal{L}_{w_{3} v_{1} v_{1}}=g_{w_{3} v_{1} v_{1}} \operatorname{tr}\left(W_{3}^{\mu \nu \rho}\left[\left(\partial_{\mu} V_{1, \nu}\right), V_{1, \rho}\right]_{-}\right) \\
& =\frac{g_{w_{3}} v_{1} v_{1}}{4}\left(+\rho_{3}^{0, \mu \nu \rho}\left\{+\left(\partial_{\mu} \bar{K}_{\nu}^{* 0}\right) K_{\rho}^{* 0}-\left(\partial_{\mu} K_{\nu}^{* 0}\right) \bar{K}_{\rho}^{* 0}+\left(\partial_{\mu} K_{\nu}^{*+}\right) K_{\rho}^{*-}-\left(\partial_{\mu} K_{\nu}^{*-}\right) K_{\rho}^{*+}+2\left[\left(\partial_{\mu} \rho_{\nu}^{+}\right) \rho_{\rho}^{-}-\left(\partial_{\mu} \rho_{\nu}^{-}\right) \rho_{\rho}^{+}\right]\right\}\right. \\
& +\rho_{3}^{+, \mu \nu \rho}\left\{\sqrt{2}\left[-\left(\partial_{\mu} K_{\nu}^{*-}\right) K_{\rho}^{* 0}+\left(\partial_{\mu} K_{\nu}^{* 0}\right) K_{\rho}^{*-}\right]+2\left[\left(\partial_{\mu} \rho_{\nu}^{-}\right) \rho_{\rho}^{0}-\left(\partial_{\mu} \rho_{\nu}^{0}\right) \rho_{\rho}^{-}\right]\right\} \\
& +\rho_{3}^{-, \mu \nu \rho}\left\{\sqrt{2}\left[\left(\partial_{\mu} K_{\nu}^{*+}\right) \bar{K}_{\rho}^{* 0}-\left(\partial_{\mu} \bar{K}_{\nu}^{* 0}\right) K_{\rho}^{*+}\right]+2\left[-\left(\partial_{\mu} \rho_{\nu}^{+}\right) \rho_{\rho}^{0}+\left(\partial_{\mu} \rho_{\nu}^{0}\right) \rho_{\rho}^{+}\right]\right\} \\
& +\bar{K}_{3}^{* 0, \mu \nu \rho}\left\{+\left(\partial_{\mu} K_{\nu}^{* 0}\right) \rho_{\rho}^{0}-\left(\partial_{\mu} \rho_{\nu}^{0}\right) K_{\rho}^{* 0}+\sqrt{2}\left[-\left(\partial_{\mu} K_{\nu}^{*+}\right) \rho_{\rho}^{-}+\left(\partial_{\mu} \rho_{\nu}^{-}\right) K_{\rho}^{*+}\right]\right. \\
& +\left(\partial_{\mu} K_{\nu}^{* 0}\right)\left[\omega_{\rho}\left(-\cos \beta_{v_{1}}+\sqrt{2} \sin \beta_{v_{1}}\right)+\phi_{\rho}\left(\sin \beta_{v_{1}}+\sqrt{2} \cos \beta_{v_{1}}\right)\right] \\
& \left.-\left[\left(\partial_{\mu} \omega_{\nu}\right)\left(-\cos \beta_{v_{1}}+\sqrt{2} \sin \beta_{v_{1}}\right)+\left(\partial_{\mu} \phi_{\nu}\right)\left(\sin \beta_{v_{1}}+\sqrt{2} \cos \beta_{v_{1}}\right)\right] K_{\rho}^{* 0}\right\} \\
& +K_{3}^{* 0, \mu \nu \rho}\left\{-\left(\partial_{\mu} \bar{K}_{\nu}^{* 0}\right) \rho_{\rho}^{0}+\left(\partial_{\mu} \rho_{\nu}^{0}\right) \bar{K}_{\rho}^{* 0}+\sqrt{2}\left[\left(\partial_{\mu} K_{\nu}^{*-}\right) \rho_{\rho}^{+}-\left(\partial_{\mu} \rho_{\nu}^{+}\right) K_{\rho}^{*-}\right]\right. \\
& -\left(\partial_{\mu} \bar{K}_{\nu}^{* 0}\right)\left[\omega_{\rho}\left(-\cos \beta_{v_{1}}+\sqrt{2} \sin \beta_{v_{1}}\right)+\phi_{\rho}\left(\sin \beta_{v_{1}}+\sqrt{2} \cos \beta_{v_{1}}\right)\right] \\
& \left.+\left[\left(\partial_{\mu} \omega_{\nu}\right)\left(-\cos \beta_{v_{1}}+\sqrt{2} \sin \beta_{v_{1}}\right)+\left(\partial_{\mu} \phi_{\nu}\right)\left(\sin \beta_{v_{1}}+\sqrt{2} \cos \beta_{v_{1}}\right)\right] \bar{K}_{\rho}^{* 0}\right\} \\
& +K_{3}^{*+, \mu \nu \rho}\left\{+\left(\partial_{\mu} K_{\nu}^{*-}\right) \rho_{\rho}^{0}-\left(\partial_{\mu} \rho_{\nu}^{0}\right) K_{\rho}^{*-}+\sqrt{2}\left[\left(\partial_{\mu} \bar{K}_{\nu}^{* 0}\right) \rho_{\rho}^{-}-\left(\partial_{\mu} \rho_{\nu}^{-}\right) \bar{K}_{\rho}^{* 0}\right]\right. \\
& -\left(\partial_{\mu} K_{\nu}^{*-}\right)\left[\omega_{\rho}\left(-\cos \beta_{v_{1}}+\sqrt{2} \sin \beta_{v_{1}}\right)+\phi_{\rho}\left(\sin \beta_{v_{1}}+\sqrt{2} \cos \beta_{v_{1}}\right)\right] \\
& \left.+\left[\left(\partial_{\mu} \omega_{\nu}\right)\left(-\cos \beta_{v_{1}}+\sqrt{2} \sin \beta_{v_{1}}\right)+\left(\partial_{\mu} \phi_{\nu}\right)\left(\sin \beta_{v_{1}}+\sqrt{2} \cos \beta_{v_{1}}\right)\right] K_{\rho}^{*-}\right\} \\
& +K_{3}^{*-, \mu \nu \rho}\left\{-\left(\partial_{\mu} K_{\nu}^{*+}\right) \rho_{\rho}^{0}+\left(\partial_{\mu} \rho_{\nu}^{0}\right) K_{\rho}^{*+}+\sqrt{2}\left[-\left(\partial_{\mu} K_{\nu}^{* 0}\right) \rho_{\rho}^{+}+\left(\partial_{\mu} \rho_{\nu}^{+}\right) K_{\rho}^{* 0}\right]\right. \\
& +\left(\partial_{\mu} K_{\nu}^{*+}\right)\left[\omega_{\rho}\left(-\cos \beta_{v_{1}}+\sqrt{2} \sin \beta_{v_{1}}\right)+\phi_{\rho}\left(\sin \beta_{v_{1}}+\sqrt{2} \cos \beta_{v_{1}}\right)\right] \\
& \left.-\left[\left(\partial_{\mu} \omega_{\nu}\right)\left(-\cos \beta_{v_{1}}+\sqrt{2} \sin \beta_{v_{1}}\right)+\left(\partial_{\mu} \phi_{\nu}\right)\left(\sin \beta_{v_{1}}+\sqrt{2} \cos \beta_{v_{1}}\right)\right] K_{\rho}^{*+}\right\} \\
& +\omega_{3}^{\mu \nu \rho}\left\{\left(-\cos \beta_{w_{3}}+\sqrt{2} \sin \beta_{w_{3}}\right) \times\left[\left(\partial_{\mu} \bar{K}_{\nu}^{* 0}\right) K_{\rho}^{* 0}-\left(\partial_{\mu} K_{\nu}^{* 0}\right) \bar{K}_{\rho}^{* 0}-\left(\partial_{\mu} K_{\nu}^{*+}\right) K_{\rho}^{*-}+\left(\partial_{\mu} K_{\nu}^{*-}\right) K_{\rho}^{*+}\right]\right\} \\
& \left.+\phi_{3}^{\mu \nu \rho}\left\{\left(\sin \beta_{w_{3}}+\sqrt{2} \cos \beta_{w_{3}}\right) \times\left[\left(\partial_{\mu} \bar{K}_{\nu}^{* 0}\right) K_{\rho}^{* 0}-\left(\partial_{\mu} K_{\nu}^{* 0}\right) \bar{K}_{\rho}^{* 0}-\left(\partial_{\mu} K_{\nu}^{*+}\right) K_{\rho}^{*-}+\left(\partial_{\mu} K_{\nu}^{*-}\right) K_{\rho}^{*+}\right]\right\}\right) .  \tag{B7}\\
& \mathcal{L}_{g_{3} v_{1} p}=c_{g_{3} v_{1} P} G_{3, \mu \alpha \beta} \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[\left\{\left(\partial_{\nu} V_{1, \rho}\right),\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right] \\
& =c_{g_{3} v_{1} p} G_{3, \mu \alpha \beta} \varepsilon^{\mu \nu \rho \sigma}\left\{+\left(\partial_{\nu} \bar{K}_{1, \mu}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{0}\right)+\left(\partial_{\nu} K_{1, \mu}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \bar{K}^{0}\right)+\left(\partial_{\nu} K_{1, \mu}^{+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{-}\right)\right. \\
& +\left(\partial_{\nu} K_{1, \mu}^{-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} K^{+}\right)+\left(\partial_{\nu} \rho_{1, \mu}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{0}\right)+\left(\partial_{\nu} \rho_{1, \mu}^{+}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{-}\right)+\left(\partial_{\nu} \rho_{1, \mu}^{-}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \pi^{+}\right) \\
& +\left[\left(\partial_{\nu} \omega_{1, \mu}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right)+\left(\partial_{\nu} \phi_{1, \mu}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right)\right] \times \cos \left(\beta_{p}-\beta_{v_{1}}\right) \\
& \left.+\left[\left(\partial_{\nu} \phi_{1, \mu}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta\right)-\left(\partial_{\nu} \omega_{1, \mu}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} \eta^{\prime}\right)\right] \times \sin \left(\beta_{p}-\beta_{v_{1}}\right)\right\} .  \tag{B8}\\
& \mathcal{L}_{g_{3} \gamma p}=c_{g_{3} v_{1} p} \frac{e}{g_{\rho}} G_{3, \mu \alpha \beta}\left(\partial_{\nu} a_{\rho}\right) \varepsilon^{\mu \nu \rho \sigma} \operatorname{tr}\left[\left\{Q,\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right)\right\}_{+}\right] \\
& =\frac{c_{g_{3} v_{1} p}}{3} \frac{e}{g_{\rho}} G_{3, \mu \alpha \beta}\left(\partial_{\nu} a_{\rho}\right) \varepsilon^{\mu \nu \rho \sigma}\left\{+3 \pi^{0}+\left[\cos \left(\beta_{p}\right)-\sqrt{2} \sin \left(\beta_{p}\right)\right] \eta-\left[\sin \left(\beta_{p}\right)+\sqrt{2} \cos \left(\beta_{p}\right)\right] \eta^{\prime}\right\} .  \tag{B9}\\
& \mathcal{L}_{g_{3} v_{1} a_{1}}=c_{g_{3} v_{1} a_{1}} G_{3, \mu \alpha \beta} \beta^{\mu \nu \rho \sigma} \operatorname{tr}\left[\left\{\left(\partial_{\nu} V_{1, \rho}\right),\left(\partial^{\alpha} \partial^{\beta} A_{1, \sigma}\right)\right\}_{+}\right] \\
& =c_{g_{3} v_{1} a_{1}} G_{3, \mu \alpha \beta} \varepsilon^{\mu \nu \rho \sigma}\left\{+\left(\partial_{\nu} \bar{K}_{1, \mu}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} K_{1, A, \sigma}^{0}\right)+\left(\partial_{\nu} K_{1, \mu}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} \bar{K}_{1, A, \sigma}^{0}\right)\right. \\
& +\left(\partial_{\nu} K_{1, \mu}^{+}\right)\left(\partial^{\alpha} \partial^{\beta} K_{1, A, \sigma}^{-}\right)+\left(\partial_{\nu} K_{1, \mu}^{-}\right)\left(\partial^{\alpha} \partial^{\beta} K_{1, A, \sigma}^{+}\right)+\left(\partial_{\nu} \rho_{1, \mu}^{0}\right)\left(\partial^{\alpha} \partial^{\beta} a_{1, \sigma}^{0}\right)+\left(\partial_{\nu} \rho_{1, \mu}^{+}\right)\left(\partial^{\alpha} \partial^{\beta} a_{1, \sigma}^{-}\right) \\
& +\left(\partial_{\nu} \rho_{1, \mu}^{-}\right)\left(\partial^{\alpha} \partial^{\beta} a_{1, \sigma}^{+}\right)+\left[\left(\partial_{\nu} \omega_{1, \mu}\right)\left(\partial^{\alpha} \partial^{\beta} f_{1, \sigma}\right)+\left(\partial_{\nu} \phi_{1, \mu}\right)\left(\partial^{\alpha} \partial^{\beta} f_{1, \sigma}^{\prime}\right)\right] \times \cos \left(\beta_{a_{1}}-\beta_{v_{1}}\right) \\
& \left.+\left[\left(\partial_{\nu} \phi_{1, \mu}\right)\left(\partial^{\alpha} \partial^{\beta} f_{1, \sigma}\right)-\left(\partial_{\nu} \omega_{1, \mu}\right)\left(\partial^{\alpha} \partial^{\beta} f_{1, \sigma}^{\prime}\right)\right] \times \sin \left(\beta_{a_{1}}-\beta_{v_{1}}\right)\right\} . \tag{B10}
\end{align*}
$$

$$
\begin{align*}
\mathcal{L}_{g_{3} b_{1} p}= & c_{g_{3} b_{1} p} G_{3}^{\mu \nu \rho} \operatorname{tr}\left[\left\{B_{1, \mu},\left(\partial_{\nu} \partial_{\rho} P\right)\right\}_{+}\right] \\
= & c_{g_{3} b_{1} p} G_{3}^{\mu \nu \rho}\left\{+\bar{K}_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} K^{0}\right)+K_{1, B, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \bar{K}^{0}\right)+K_{1, B, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} K^{-}\right)+K_{1, B, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} K^{+}\right)\right. \\
& +b_{1, \mu}^{0}\left(\partial_{\nu} \partial_{\rho} \pi^{0}\right)+b_{1, \mu}^{+}\left(\partial_{\nu} \partial_{\rho} \pi^{-}\right)+b_{1, \mu}^{-}\left(\partial_{\nu} \partial_{\rho} \pi^{+}\right)+\left[h_{1, \mu}\left(\partial_{\nu} \partial_{\rho} \eta\right)+h_{1, \mu}^{\prime}\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\right] \cos \left(\beta_{p}-\beta_{b_{1}}\right) \\
& \left.+\left[h_{1, \mu}^{\prime}\left(\partial_{\nu} \partial_{\rho} \eta\right)-h_{1, \mu}\left(\partial_{\nu} \partial_{\rho} \eta^{\prime}\right)\right] \sin \left(\beta_{p}-\beta_{b_{1}}\right)\right\} .  \tag{B11}\\
\mathcal{L}_{g_{3} b_{1} a_{1}}= & c_{g_{3} b_{1} a_{1}} G_{3}^{\mu \nu \rho} \operatorname{tr}\left[\left\{B_{1, \mu},\left(\partial_{\nu} A_{1, \rho}\right)\right\}_{+}\right] \\
= & c_{g_{3} b_{1} a_{1}} G_{3}^{\mu \nu \rho}\left\{+\bar{K}_{1, B, \mu}^{0}\left(\partial_{\nu} K_{1, A, \rho}^{0}\right)+K_{1, B, \mu}^{0}\left(\partial_{\nu} \bar{K}_{1, A, \rho}^{0}\right)+K_{1, B, \mu}^{+}\left(\partial_{\nu} K_{1, A, \rho}^{-}\right)+K_{1, B, \mu}^{-}\left(\partial_{\nu} K_{1, A, \rho}^{+}\right)\right. \\
& \quad+b_{1, \mu}^{0}\left(\partial_{\nu} a_{1, \rho}^{0}\right)+b_{1, \mu}^{+}\left(\partial_{\nu} a_{1, \rho}^{-}\right)+b_{1, \mu}^{-}\left(\partial_{\nu} a_{1, \rho}^{+}\right)+\left[h_{1, \mu}\left(\partial_{\nu} f_{1, \rho}\right)+h_{1, \mu}^{\prime}\left(\partial_{\nu} f_{1, \rho}^{\prime}\right)\right] \cos \left(\beta_{a_{1}}-\beta_{b_{1}}\right) \\
& \left.\quad+\left[h_{1, \mu}^{\prime}\left(\partial_{\nu} f_{1, \rho}\right)-h_{1, \mu}\left(\partial_{\nu} f_{1, \rho}^{\prime}\right)\right] \sin \left(\beta_{a_{1}}-\beta_{b_{1}}\right)\right\} . \tag{B12}
\end{align*}
$$

## APPENDIX C: COEFFICIENTS FOR THE DECAY CHANNELS

In this Appendix we present the explicit forms of the coefficients $\kappa_{i}$ and $\kappa_{i}^{\gamma}$ in Eq. (23) based on the extended form of Lagrangians presented in the previous section in Tables XIX-XXX.

TABLE XIX. Coefficients $\kappa_{i}$ for the decay channels of $W_{3} \rightarrow$ $P+P$ decays that are not kinematically forbidden.

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $\rho_{3}(1690) \rightarrow \pi \pi$ | 1 |
| $\rho_{3}(1690) \rightarrow \bar{K} K$ | $2\left(\frac{1}{2}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \pi \bar{K}$ | $\left(\frac{1}{2}\right)^{2}+\left(\frac{\sqrt{2}}{2}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K} \eta$ | $\left[\frac{1}{2}\left(-\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)\right]^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K} \eta^{\prime}(958)$ | $\left[\frac{1}{2}\left(\sqrt{2} \cos \beta_{p}+\sin \beta_{p}\right)\right]^{2}$ |
| $\omega_{3}(1670) \rightarrow \bar{K} K$ | $2\left[\frac{1}{2}\left(-\cos \beta_{w_{3}}+\sqrt{2} \sin \beta_{w_{3}}\right)\right]^{2}$ |
| $\phi_{3}(1850) \rightarrow \bar{K} K$ | $2\left[\frac{1}{2}\left(\sqrt{2} \cos \beta_{w_{3}}+\sin \beta_{w_{3}}\right)\right]^{2}$ |

TABLE XX. Coefficients $\kappa_{i}^{\gamma}$ for the decay channels of radioactive $W_{3} \rightarrow \gamma P$ decays.

| Decay process | $\kappa_{i}^{\gamma}$ |
| :--- | :---: |
| $\rho_{3}^{ \pm / 0}(1690) \rightarrow \gamma \pi^{ \pm / 0}$ | $\left(\frac{1}{6}\right)^{2}$ |
| $\rho_{3}^{0}(1690) \rightarrow \gamma \eta$ | $\left(\frac{1}{2} \cos \beta_{p}\right)^{2}$ |
| $\rho_{3}^{0}(1690) \rightarrow \gamma \eta^{\prime}(958)$ | $\left(\frac{1}{2} \sin \beta_{p}\right)^{2}$ |
| $K_{3}^{ \pm}(1780) \rightarrow \gamma K^{ \pm}$ | $\left(\frac{1}{6}\right)^{2}$ |
| $K_{3}^{0}(1780) \rightarrow \gamma K^{0}$ | $\left(\frac{1}{3}\right)^{2}$ |
| $\omega_{3}(1670) \rightarrow \gamma \pi^{0}$ | $\left(\frac{1}{2} \cos \beta_{w_{3}}\right)^{2}$ |
| $\omega_{3}(1670) \rightarrow \gamma \eta$ | $\left[\frac{1}{6}\left(\cos \beta_{w_{3}} \cos \beta_{p}-2 \sin \beta_{w_{3}} \sin \beta_{p}\right)\right]^{2}$ |
| $\omega_{3}(1670) \rightarrow \gamma \eta^{\prime}(958)$ | $\left[\frac{1}{6}\left(\cos \beta_{w_{3}} \sin \beta_{p}+2 \sin \beta_{w_{3}} \cos \beta_{p}\right)\right]^{2}$ |
| $\phi_{3}(1850) \rightarrow \gamma \pi^{0}$ | $\left(\frac{1}{2} \sin \beta_{w_{3}}\right)^{2}$ |
| $\phi_{3}(1850) \rightarrow \gamma \eta$ | $\left[\frac{1}{6}\left(\sin \beta_{w_{3}} \cos \beta_{p}+2 \cos \beta_{w_{3}} \sin \beta_{p}\right)\right]^{2}$ |
| $\phi_{3}(1850) \rightarrow \gamma \eta^{\prime}(958)$ | $\left[\frac{1}{6}\left(\sin \beta_{w_{3}} \sin \beta_{p}-2 \cos \beta_{w_{3}} \cos \beta_{p}\right)\right]^{2}$ |

TABLE XXI. Coefficients $\kappa_{i}$ for the decay channels of $W_{3} \rightarrow$ $A_{2}+P$ decays that are not kinematically forbidden.

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $\rho_{3}(1690) \rightarrow a_{2}(1320) \pi$ | $2\left(\frac{1}{2}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}_{2}^{*}(1430) \pi$ | $\left(\frac{1}{4}\right)^{2}+\left(\frac{\sqrt{2}}{4}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow f_{2}(1270) \bar{K}$ | $\left[\frac{1}{4}\left(\cos \beta_{a_{2}}-\sqrt{2} \sin \beta_{a_{2}}\right)\right]^{2}$ |

TABLE XXII. Coefficients $\kappa_{i}$ for the decay channels of $W_{3} \rightarrow$ $B_{1}+P$ decays that are not kinematically forbidden.

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $\rho_{3}(1690) \rightarrow h_{1}(1170) \pi$ | $\left(\frac{1}{2} \cos \beta_{b_{1}}\right)^{2}$ |
| $\rho_{3}(1690) \rightarrow h_{1}(1415) \pi$ | $\left(\frac{1}{2} \sin \beta_{b_{1}}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow b_{1}(1235) K$ | $\left(\frac{1}{4}\right)^{2}+\left(\frac{\sqrt{2}}{4}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}_{1, B} \pi$ | $\left(\frac{1}{4}\right)^{2}+\left(\frac{\sqrt{2}}{4}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow h_{1}(1170) \bar{K}$ | $\left[\frac{1}{4}\left(\cos \beta_{b_{1}}+\sqrt{2} \sin \beta_{b_{1}}\right)\right]^{2}$ |
| $\omega_{3}(1670) \rightarrow b_{1}(1235) \pi$ | $3\left(\frac{1}{2} \cos \beta_{w_{3}}\right)^{2}$ |
| $\phi_{3}(1850) \rightarrow b_{1}(1235) \pi$ | $3\left(\frac{1}{2} \sin \beta_{w_{3}}\right)^{2}$ |
| $\phi_{3}(1850) \rightarrow \bar{K}_{1, B} K$ | $4\left[\frac{1}{4}\left(\sqrt{2} \cos \beta_{w_{3}}-\sin \beta_{w_{3}}\right)\right]^{2}$ |
| $\phi_{3}(1850) \rightarrow h_{1}(1170) \eta$ | $\left[\frac{1}{2}\left(-\sin \beta_{w_{3}} \cos \beta_{b_{1}} \cos \beta_{p}+\right.\right.$ |
|  | $\left.\left.+\sqrt{2} \cos \beta_{w_{3}} \sin \beta_{b_{1}} \sin \beta_{p}\right)\right]^{2}$ |

TABLE XXIII. Coefficients $\kappa_{i}$ for the decay channels of $W_{3} \rightarrow$ $A_{1}+P$ decays that are not kinematically forbidden.

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $\rho_{3}(1690) \rightarrow a_{1}(1260) \pi$ | $2\left(\frac{1}{2}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow K_{1, A} \pi$ | $\left(\frac{1}{4}\right)^{2}+\left(\frac{\sqrt{2}}{4}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow a_{1}(1260) K$ | $\left(\frac{1}{4}\right)^{2}+\left(\frac{\sqrt{2}}{4}\right)^{2}$ |

TABLE XXIV. Coefficients $\kappa_{i}$ for the decay channels of $W_{3} \rightarrow V_{1}+V_{1}$ decays that are not kinematically forbidden.

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $\rho_{3}(1690) \rightarrow \rho(770) \rho(770)$ | 1 |
| $K_{3}^{*}(1780) \rightarrow \rho(770) \bar{K}^{*}(892)$ | $\left(\frac{1}{2}\right)^{2}+\left(\frac{\sqrt{2}}{2}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}^{*}(892) \omega(782)$ | $\left[\frac{1}{2}\left(\cos \beta_{v_{1}}-\sqrt{2} \sin \beta_{v_{1}}\right)\right]^{2}$ |
| $\phi_{3}(1850) \rightarrow \bar{K}^{*}(892) K^{*}(892)$ | $2\left[\frac{1}{2}\left(\sqrt{2} \cos \beta_{w_{3}}+\sin \beta_{w_{3}}\right)\right]^{2}$ |

TABLE XXV. Coefficients $\kappa_{i}$ for the decay channels of $G_{3}(4200) \rightarrow V_{1}+P$ decays that are not kinematically forbidden.

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $G_{3}(4200) \rightarrow \rho(770) \pi$ | 3 |
| $G_{3}(4200) \rightarrow \bar{K}^{*}(892) K$ | 4 |
| $G_{3}(4200) \rightarrow \omega(782) \eta$ | $\left[\cos \left(\beta_{p}-\beta_{v_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow \omega(782) \eta^{\prime}(958)$ | $\left[\sin \left(\beta_{p}-\beta_{v_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow \phi(1020) \eta$ | $\left[\sin \left(\beta_{p}-\beta_{v_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow \phi(1020) \eta^{\prime}(958)$ | $\left[\cos \left(\beta_{p}-\beta_{v_{1}}\right)\right]^{2}$ |

TABLE XXVI. Coefficients $\kappa_{i}^{\gamma}$ for the decay channels of radioactive $G_{3}(4200) \rightarrow \gamma+P$ decays.

| Decay process | $\kappa_{i}^{\gamma}$ |
| :--- | :---: |
| $G_{3}(4200) \rightarrow \gamma \pi^{0}$ | 1 |
| $G_{3}(4200) \rightarrow \gamma \eta$ | $\left[\frac{1}{3}\left(\cos \beta_{p}-\sqrt{2} \sin \beta_{p}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow \gamma \eta^{\prime}(958)$ | $\left[\frac{1}{3}\left(\sin \beta_{p}+\sqrt{2} \cos \beta_{p}\right)\right]^{2}$ |

TABLE XXVII. Coefficients $\kappa_{i}$ for the decay channels of $G_{3}(4200) \rightarrow V_{1}+A_{1}$ decays that are not kinematically forbidden.

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $G_{3}(4200) \rightarrow \rho(770) a_{1}(1260)$ | 3 |
| $G_{3}(4200) \rightarrow \bar{K}^{*}(892) K_{1, A}$ | 4 |
| $G_{3}(4200) \rightarrow \omega(782) f_{1}(1285)$ | $\left[\cos \left(\beta_{a_{1}}-\beta_{v_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow \omega(782) f_{1}^{\prime}(1420)$ | $\left[\sin \left(\beta_{a_{1}}-\beta_{v_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow \phi(1020) f_{1}(1285)$ | $\left[\sin \left(\beta_{a_{1}}-\beta_{v_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow \phi(1020) f_{1}^{\prime}(1420)$ | $\left[\cos \left(\beta_{a_{1}}-\beta_{v_{1}}\right)\right]^{2}$ |

TABLE XXVIII. Coefficients $\kappa_{i}$ for the decay channels of $G_{3}(4200) \rightarrow B_{1}+P$ decays that are not kinematically forbidden.

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $G_{3}(4200) \rightarrow b_{1}(1235) \pi$ | 3 |
| $G_{3}(4200) \rightarrow K_{1, B} K$ | 4 |
| $G_{3}(4200) \rightarrow h_{1}(1170) \eta$ | $\left[\cos \left(\beta_{p}-\beta_{b_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow h_{1}(1170) \eta^{\prime}(958)$ | $\left[\sin \left(\beta_{p}-\beta_{b_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow h_{1}(1415) \eta$ | $\left[\sin \left(\beta_{p}-\beta_{b_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow h_{1}(1415) \eta^{\prime}(958)$ | $\left[\cos \left(\beta_{p}-\beta_{b_{1}}\right)\right]^{2}$ |

TABLE XXIX. Coefficients $\kappa_{i}$ for the decay channels of $G_{3}(4200) \rightarrow B_{1}+A_{1}$ decays that are not kinematically forbidden.

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $G_{3}(4200) \rightarrow b_{1}(1235) a_{1}(1260)$ | 3 |
| $G_{3}(4200) \rightarrow K_{1, B} K_{1, A}$ | 4 |
| $G_{3}(4200) \rightarrow h_{1}(1170) f_{1}(1285)$ | $\left[\cos \left(\beta_{a_{1}}-\beta_{b_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow h_{1}(1170) f_{1}^{\prime}(1420)$ | $\left[\sin \left(\beta_{a_{1}}-\beta_{b_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow h_{1}(1415) f_{1}(1285)$ | $\left[\sin \left(\beta_{a_{1}}-\beta_{b_{1}}\right)\right]^{2}$ |
| $G_{3}(4200) \rightarrow h_{1}(1415) f_{1}^{\prime}(1420)$ | $\left[\cos \left(\beta_{a_{1}}-\beta_{b_{1}}\right)\right]^{2}$ |

## APPENDIX D: UNPOLARIZED INVARIANT DECAY AMPLITUDES

This Appendix is devoted to the calculation of the unpolarized invariant decay amplitudes. These form the normal decay amplitudes by multiplication with factors $\kappa_{i}$ depending on the specific decay channels. We apply the following Feynman rules for each interaction vertex in order to derive the expressions for the decay amplitudes,

$$
\begin{gather*}
\partial^{\mu} \mapsto \mathrm{i} k_{\rho}^{\mu},  \tag{D1}\\
V_{1}^{\mu} \mapsto \epsilon^{\mu}\left(\lambda_{v_{1}}, \vec{k}_{v_{1}}\right),  \tag{D2}\\
B_{1}^{\mu} \mapsto \epsilon^{\mu}\left(\lambda_{b_{1}}, \vec{k}_{b_{1}}\right),  \tag{D3}\\
A_{1}^{\mu} \mapsto \epsilon^{\mu}\left(\lambda_{a_{1}}, \vec{k}_{a_{1}}\right),  \tag{D4}\\
A_{2}^{\mu \nu} \mapsto \epsilon^{\mu \nu}\left(\lambda_{a_{2}}, \vec{k}_{a_{2}}\right),  \tag{D5}\\
W_{3}^{\mu \nu \rho} \mapsto \epsilon^{\mu \nu \rho}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right),  \tag{D6}\\
G_{3}^{\mu \nu \rho} \mapsto \epsilon^{\mu \nu \rho}\left(\lambda_{g_{3}}, \vec{k}_{g_{3}}\right),  \tag{D7}\\
a^{\mu} \mapsto \epsilon^{\mu}\left(\lambda_{\gamma}, \vec{k}_{\gamma}\right), \tag{D8}
\end{gather*}
$$

where $\epsilon^{\mu(\nu)(\rho)}$ are the polarization vectors (tensors), $\lambda$ is the specific polarization, ${ }^{6} k^{\mu}$ is the four momentum, and $\vec{k}$ is the three momentum.
(1) For the interaction Lagrangian

$$
\begin{equation*}
\mathcal{L}_{w_{3} p p}=g_{w_{3} p p} W_{3, \mu \nu \rho} P^{(1)}\left(\partial^{\mu} \partial^{\nu} \partial^{\rho} P^{(2)}\right) \tag{D9}
\end{equation*}
$$

[^5]$\underline{\underline{\text { TABLE XXX. Coefficients } \kappa_{i} \text { for the decay channels of } W_{3} \rightarrow V_{1}+P \text { decays that are not kinematically forbidden. }}}$

| Decay process | $\kappa_{i}$ |
| :--- | :---: |
| $\rho_{3}(1690) \rightarrow \rho(770) \eta$ | $\left(\frac{1}{2} \cos \beta_{p}\right)^{2}$ |
| $\rho_{3}(1690) \rightarrow \bar{K}^{*}(892) K$ | $4\left(\frac{1}{4}\right)^{2}$ |
| $\rho_{3}(1690) \rightarrow \omega(782) \pi$ | $\left(\frac{1}{2} \cos \beta_{v_{1}}\right)^{2}$ |
| $\rho_{3}(1690) \rightarrow \phi(1020) \pi$ | $\left(\frac{1}{2} \sin \beta_{v_{1}}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \rho(770) K$ | $\left(\frac{1}{4}\right)^{2}+\left(\frac{\sqrt{2}}{4}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}^{*}(892) \pi$ | $\left(\frac{1}{4}\right)^{2}+\left(\frac{\sqrt{2}}{4}\right)^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}^{*}(892) \eta$ | $\left[\frac{1}{4}\left(\cos \beta_{p}+\sqrt{2} \sin \beta_{p}\right)\right]^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \bar{K}^{*}(892) \eta^{\prime}(958)$ | $\left[\frac{1}{4}\left(\sqrt{2} \cos \beta_{p}-\sin \beta_{p}\right)\right]^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \omega(782) \bar{K}$ | $\left[\frac{1}{4}\left(\cos \beta_{v_{1}}+\sqrt{2} \sin \beta_{v_{1}}\right)\right]^{2}$ |
| $K_{3}^{*}(1780) \rightarrow \phi(1020) \bar{K}$ | $\left[\frac{1}{4}\left(\sqrt{2} \cos \beta_{v_{1}}-\sin \beta_{v_{1}}\right)\right]^{2}$ |
| $\omega_{3}(1670) \rightarrow \rho(770) \pi$ | $3\left(\frac{1}{2} \cos \beta_{w_{3}}\right)^{2}$ |
| $\omega_{3}(1670) \rightarrow \bar{K}^{*}(892) K$ | $4\left[\frac{1}{4}\left(\cos \beta_{w_{3}}+\sqrt{2} \sin \beta_{w_{3}}\right)\right]^{2}$ |
| $\omega_{3}(1670) \rightarrow \omega(782) \eta$ | $\left[\frac{1}{2}\left(\cos \beta_{w_{3}} \cos \beta_{v_{1}} \cos \beta_{p}+\sqrt{2} \sin \beta_{w_{3}} \sin \beta_{v_{1}} \sin \beta_{p}\right)\right]^{2}$ |
| $\omega_{3}(1670) \rightarrow \phi(1020) \eta$ | $\left[\frac{1}{2}\left(-\cos \beta_{w_{3}} \sin \beta_{v_{1}} \cos \beta_{p}+\sqrt{2} \sin \beta_{w_{3}} \cos \beta_{v_{1}} \sin \beta_{p}\right)\right]^{2}$ |
| $\phi_{3}(1850) \rightarrow \rho(770) \pi$ | $3\left(\frac{1}{2} \sin \beta_{w_{3}}\right)^{2}$ |
| $\phi_{3}(1850) \rightarrow \bar{K}^{*}(892) K$ | $4\left[\frac{1}{4}\left(\sqrt{2} \cos \beta_{w_{3}}-\sin \beta_{w_{3}}\right)\right]^{2}$ |
| $\phi_{3}(1850) \rightarrow \omega(782) \eta$ | $\left[\frac{1}{2}\left(-\sin \beta_{w_{3}} \cos \beta_{v_{1}} \cos \beta_{p}+\sqrt{2} \cos \beta_{w_{3}} \sin \beta_{v_{1}} \sin \beta_{p}\right)\right]^{2}$ |
| $\phi_{3}(1850) \rightarrow \omega(782) \eta^{\prime}(958)$ | $\left[\frac{1}{2}\left(\sin \beta_{w_{3}} \cos \beta_{v_{1}} \sin \beta_{p}+\sqrt{2} \cos \beta_{w_{3}} \sin \beta_{v_{1}} \cos \beta_{p}\right)\right]^{2}$ |
| $\phi_{3}(1850) \rightarrow \phi(1020) \eta$ | $\left[\frac{1}{2}\left(\sin \beta_{w_{3}} \sin \beta_{v_{1}} \cos \beta_{p}+\sqrt{2} \cos \beta_{w_{3}} \cos \beta_{v_{1}} \sin \beta_{p}\right)\right]^{2}$ |

which describes decays of massive spin-3 fields into two massive pseudoscalar fields, the unpolarized invariant decay amplitude has the following form

$$
\begin{equation*}
\mathrm{i} \mathcal{M}_{w_{3} p p}=\mathrm{i} g_{w_{3} p p} \epsilon_{\mu \nu \rho}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) k_{p^{(2)}}^{\mu} k_{p^{(2)}}^{\nu} k_{p^{(2)}}^{\rho} \tag{D10}
\end{equation*}
$$

The square of the amplitude reads

$$
\begin{equation*}
\frac{1}{7}\left|\mathcal{M}_{w_{3} p p}\right|^{2}=g_{w_{3} p p}^{2} \frac{1}{7} \sum_{\lambda_{w_{3}}=-3}^{+3} \epsilon_{\mu \nu \rho}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \epsilon_{\bar{\mu} \bar{\nu} \bar{\rho}}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) k_{p^{(2)}}^{\mu} k_{p^{(2)}}^{\nu} k_{p^{(2)}}^{\rho} k_{p^{(2)}}^{\bar{\mu}} k_{p^{(2)}}^{\bar{\nu}} k_{p^{(2)}}^{\bar{\rho}}=g_{w_{3} p p}^{2} \frac{2}{35}\left|\vec{k}_{p^{(1)}, p^{(2)}}\right|^{6} \tag{D11}
\end{equation*}
$$

(2) The unpolarized invariant decay amplitude for the interaction of massive spin-3 fields, massive vector fields, and massive pseudoscalar fields

$$
\begin{equation*}
\mathcal{L}_{w_{3} v_{1} p}=g_{w_{3} v_{1} p} \varepsilon^{\mu \nu \rho \sigma} W_{3, \mu \alpha \beta}\left(\partial_{\nu} V_{1, \rho}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right) \tag{D12}
\end{equation*}
$$

reads

$$
\begin{equation*}
\mathrm{i} \mathcal{M}_{w_{3} v_{1} p}=g_{w_{3} v_{1} p} \varepsilon^{\mu \nu \rho \sigma} \epsilon_{\mu \alpha \beta}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) k_{v_{1} \nu} \epsilon_{\rho}\left(\lambda_{v_{1}}, \vec{k}_{v_{1}}\right) k_{p}^{\alpha} k_{p}^{\beta} k_{p \sigma} \tag{D13}
\end{equation*}
$$

The square of the amplitude is

$$
\begin{align*}
\frac{1}{7}\left|\mathcal{M}_{w_{3} v_{1} p}\right|^{2}= & g_{w_{3} v_{1} p}^{2} \frac{1}{7} \varepsilon^{\mu \nu \rho \sigma} \varepsilon^{\bar{\mu} \bar{\nu} \bar{\rho} \bar{\sigma}} \sum_{\lambda_{w_{3}}=-3}^{+3} \epsilon_{\mu \alpha \beta}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \epsilon_{\bar{\mu} \bar{\alpha} \bar{\beta}}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \\
& \times k_{v_{1} \nu} k_{v_{1} \bar{\nu}} \sum_{\lambda_{v_{1}}=-1}^{+1} \epsilon_{\rho}\left(\lambda_{v_{1}}, \vec{k}_{v_{1}}\right) \epsilon_{\bar{\rho}}\left(\lambda_{v_{1}}, \vec{k}_{v_{1}}\right) k_{p}^{\alpha} k_{p}^{\beta} k_{p \sigma} k_{p}^{\bar{\alpha}} k_{p}^{\bar{\beta}} k_{p \bar{\sigma}} \\
= & g_{w_{3} v_{1} p}^{2} \frac{8}{105}\left|\vec{k}_{v_{1}, p}\right|{ }^{6} m_{w_{3}}^{2} \tag{D14}
\end{align*}
$$

(3) For the decays of massive spin-3 fields into massive pseudoscalar fields and photons, the generic interaction Lagrangian is of the general type

$$
\begin{equation*}
\mathcal{L}_{w_{3} \gamma p}=g_{w_{3} v_{1} p} \frac{e}{g_{\rho}} \varepsilon^{\mu \nu \rho \sigma} W_{3, \mu \alpha \beta}\left(\partial_{\nu} a_{\rho}\right)\left(\partial^{\alpha} \partial^{\beta} \partial_{\sigma} P\right) \tag{D15}
\end{equation*}
$$

The corresponding decay amplitude reads

$$
\begin{equation*}
\text { i. } \mathcal{M}_{w_{3} \gamma p}=g_{w_{3} v_{1} p} \frac{e}{g_{\rho}} \varepsilon^{\mu \nu \rho \sigma} \epsilon_{\mu \alpha \beta}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) k_{\gamma, \nu} \epsilon_{\rho}\left(\lambda_{\gamma}, \vec{k}_{\gamma}\right) k_{p}^{\alpha} k_{p}^{\beta} k_{p, \sigma} . \tag{D16}
\end{equation*}
$$

The averaged square of the amplitude is

$$
\begin{align*}
\frac{1}{7}\left|\mathcal{M}_{w_{3} \gamma p}\right|^{2}= & g_{w_{3} v_{1} p}^{2}\left(\frac{e}{g_{\rho}}\right)^{2} \varepsilon^{\mu \nu \rho \sigma} \varepsilon^{\bar{\mu} \bar{\nu} \bar{\rho} \bar{\sigma}} k_{p}^{\alpha} k_{p}^{\beta} k_{p, \sigma} k_{p}^{\bar{\alpha}} k_{p}^{\bar{\beta}} k_{p, \bar{\sigma}} \\
& \times \sum_{\lambda_{w_{3}}=-3}^{+3} \epsilon_{\mu \alpha \beta}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \epsilon_{\bar{\mu} \bar{\alpha} \bar{\beta}}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) k_{\gamma, \nu} k_{\gamma, \bar{\nu}} \sum_{\lambda_{\gamma}=1}^{2} \epsilon_{\rho}\left(\lambda_{\gamma}, \vec{k}_{\gamma}\right) \epsilon_{\bar{\rho}}\left(\lambda_{\gamma}, \vec{k}_{\gamma}\right) \\
= & g_{w_{3} v_{1} p}^{2}\left(\frac{e}{g_{\rho}}\right)^{2} \frac{8}{105}\left|\vec{k}_{\gamma, p}\right|^{6} m_{w_{3}}^{2} . \tag{D17}
\end{align*}
$$

(4) Decays of massive spin-3 fields into massive spin-2 tensor and massive pseudoscalar fields is described by the following interaction

$$
\begin{equation*}
\mathcal{L}_{w_{3} a_{2} p}=g_{w_{3} a_{2} p} \varepsilon_{\mu \nu \rho \sigma} W_{3,{ }_{\alpha \beta}}{ }_{\alpha \beta}\left(\partial^{\nu} A_{2}^{\rho \alpha}\right)\left(\partial^{\sigma} \partial^{\beta} P\right) \tag{D18}
\end{equation*}
$$

which leads to the amplitude

$$
\begin{equation*}
\mathrm{i} \mathcal{M}_{w_{3} a_{2} p}=-\mathrm{i} g_{w_{3} a_{2} p} \varepsilon_{\mu \nu \rho \sigma} \eta_{\alpha \gamma} \epsilon^{\mu \alpha \beta}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) k_{a_{2}}^{\nu} \epsilon^{\rho \gamma}\left(\lambda_{a_{2}}, \vec{k}_{a_{2}}\right) k_{p}^{\sigma} k_{p \beta} \tag{D19}
\end{equation*}
$$

Squaring the amplitude, we find

$$
\begin{align*}
\frac{1}{7}\left|\mathcal{M}_{w_{3} a_{2} p}\right|^{2}= & g_{w_{3} a_{2} p}^{2} \frac{1}{7} \varepsilon_{\mu \nu \rho \sigma} \varepsilon_{\bar{\mu} \bar{\nu} \bar{\sigma}} \eta_{\alpha \gamma} \eta_{\bar{\alpha} \bar{\gamma}} \sum_{\lambda_{w_{3}}=-3}^{+3} \epsilon^{\mu \alpha \beta}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \epsilon^{\bar{\mu} \bar{\alpha} \bar{\beta}}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \\
& \times \sum_{\lambda_{a_{2}}=-2}^{2} \epsilon^{\rho \gamma}\left(\lambda_{a_{2}}, \vec{k}_{a_{2}}\right) \epsilon^{\bar{\rho} \bar{\gamma}}\left(\lambda_{a_{2}}, \vec{k}_{a_{2}}\right) k_{a_{2}}^{\nu} k_{a_{2}}^{\bar{\nu}} k_{p}^{\sigma} k_{p \beta} k_{p}^{\bar{\sigma}} k_{p \bar{\beta}} \\
= & g_{w_{3} a_{2} p}^{2} \frac{2}{105}\left|\vec{k}_{a_{2}, p}\right|^{4} \frac{m_{w_{3}}^{2}}{m_{a_{2}}^{2}}\left(2\left|\vec{k}_{a_{2}, p}\right|^{2}+7 m_{a_{2}}^{2}\right) . \tag{D20}
\end{align*}
$$

(5) The amplitudes for the decays of massive spin-3 fields into massive pseudoscalar and massive axial-/pseudovector fields, which are described via the interaction Lagrangians

$$
\begin{equation*}
\mathcal{L}_{w_{3} b_{1} p}=g_{w_{3} b_{1} p} W_{3}^{\mu \nu \rho} B_{1, \mu}\left(\partial_{\nu} \partial_{\rho} P\right) \tag{D21}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathcal{L}_{w_{3} a_{1} p}=g_{w_{3} a_{1} p} W_{3}^{\mu \nu \rho} A_{1, \mu}\left(\partial_{\nu} \partial_{\rho} P\right) \tag{D22}
\end{equation*}
$$

have the same shape,

$$
\begin{align*}
& \mathrm{i} \mathcal{M}_{w_{3} b_{1} p}=-g_{w_{3} b_{1} p} \epsilon^{\mu \nu \rho}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \epsilon_{\mu}\left(\lambda_{b_{1}}, \vec{k}_{b_{1}}\right) k_{p \nu} k_{p \rho}  \tag{D23}\\
& \text { i } \mathcal{M}_{w_{3} a_{1} p}=-g_{w_{3} a_{1} p} \epsilon^{\mu \nu \rho}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \epsilon_{\mu}\left(\lambda_{a_{1}}, \vec{k}_{a_{1}}\right) k_{p \nu} k_{p \rho} . \tag{D24}
\end{align*}
$$

Their squares have the following form:

$$
\begin{align*}
\frac{1}{7}\left|\mathcal{M}_{w_{3} b_{1} p}\right|^{2} & =g_{w_{3} b_{1} p}^{2} \frac{1}{7} k_{p \nu} k_{p \rho} k_{p \bar{\nu}} k_{p \bar{\rho}} \sum_{\lambda_{w_{3}}=-3}^{+3} \epsilon^{\mu \nu \rho}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) e^{\bar{\mu} \bar{\nu} \bar{\rho}}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \sum_{\lambda_{b_{1}}=-1}^{+1} \epsilon_{\mu}\left(\lambda_{b_{1}}, \vec{k}_{b_{1}}\right) \epsilon_{\bar{\mu}}\left(\lambda_{b_{1}}, \vec{k}_{b_{1}}\right) \\
& =g_{w_{3} b_{1} p}^{2} \frac{2}{105}\left|\vec{k}_{b_{1}, p}\right|^{4}\left(7+3 \frac{\mid \vec{k}_{b_{1}, p} p^{2}}{m_{b_{1}}^{2}}\right),  \tag{D25}\\
\frac{1}{7}\left|\mathcal{M}_{w_{3} a_{1} p}\right|^{2} & =g_{w_{3} a_{1} p}^{2} \frac{1}{7} k_{p \nu} k_{p \rho} k_{p \bar{\nu}} k_{p \bar{\rho}} \sum_{\lambda_{w_{3}}=-3}^{+3} \epsilon^{\mu \nu \rho}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) e^{\bar{\mu} \bar{\nu} \bar{\rho}}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \sum_{\lambda_{a_{1}}=-1}^{+1} \epsilon_{\mu}\left(\lambda_{a_{1}}, \vec{k}_{a_{1}}\right) \epsilon_{\bar{\mu}}\left(\lambda_{a_{1}}, \vec{k}_{a_{1}}\right) \\
& =g_{w_{3} a_{1} p}^{2} \frac{2}{105}\left|\vec{k}_{a_{1}, p}\right|^{4}\left(7+3 \frac{\left|\vec{k}_{a_{1}, p}\right|^{2}}{m_{a_{1}}^{2}}\right) . \tag{D26}
\end{align*}
$$

(6) The interaction Lagrangian, which describes the decay of massive spin-3 fields into two massive vector fields

$$
\begin{equation*}
\mathcal{L}_{w_{3} v_{1} v_{1}}=g_{w_{3} v_{1} v_{1}} W_{3}^{\mu \nu \rho}\left(\partial_{\mu} V_{1, \nu}^{(1)}\right) V_{1, \rho}^{(2)}, \tag{D27}
\end{equation*}
$$

leads to the following amplitude

$$
\begin{equation*}
\mathrm{i} \mathcal{M}_{w_{3} v_{1} v_{1}}=\mathrm{i} g_{w_{3} v_{1} v_{1}} \epsilon^{\mu \nu \rho}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) k_{v_{1} \mu}^{(1)} \epsilon_{\nu}\left(\lambda_{v_{1}}^{(1)}, \vec{k}_{v_{1}}^{(1)}\right) \epsilon_{\rho}\left(\lambda_{v_{1}}^{(2)}, k_{v_{2}}^{(2)}\right) \tag{D28}
\end{equation*}
$$

The square of this amplitude is given by

$$
\begin{align*}
\frac{1}{7}\left|\mathcal{M}_{w_{3} v_{1} v_{1}}\right|^{2} & =g_{w_{3} v_{1} v_{1}}^{2} \frac{1}{7} \sum_{\lambda_{w_{3}}=-3}^{+3} \epsilon^{\mu \nu \rho}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \epsilon^{\bar{\mu} \bar{\nu} \bar{\rho}}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \\
& =\times k_{v_{1} \mu}^{(1)} k_{v_{1} \bar{\mu}}^{(1)} \sum_{\lambda_{v_{1}}^{(1)}=-1}^{+1} \epsilon_{\nu}\left(\lambda_{v_{1}}^{(1)}, \vec{k}_{v_{1}}^{(1)}\right) \epsilon_{\bar{\nu}}\left(\lambda_{v_{1}}^{(1)}, \vec{k}_{v_{1}}^{(1)}\right) \sum_{\lambda_{v_{1}}^{(2)}=-1}^{+1} \epsilon_{\rho}\left(\lambda_{v_{1}}^{(2)}, \vec{k}_{v_{1}}^{(2)}\right) \epsilon_{\bar{\rho}}\left(\lambda_{v_{1}}^{(2)}, \vec{k}_{v_{1}}^{(2)}\right) \\
& =g_{w_{3} v_{1} v_{1}}^{2} \frac{1}{105}\left(m_{v_{1}^{(1)}}^{2} m_{v_{1}^{(2)}}^{2}\right)^{-1}\left|\vec{k}_{v_{1}^{(1)}, v_{1}^{(2)}}\right|^{2}\left[6\left|\vec{k}_{v_{1}^{(1)}, v_{1}^{(2)}}\right|^{4}+35 m_{v_{1}^{(1)}}^{2} m_{v_{1}^{(2)}}^{2}+14\left|\vec{k}_{v_{1}^{(1)}, v_{1}^{(2)}}\right|^{2}\left(m_{v_{1}^{(1)}}^{2}+m_{v_{1}^{(2)}}^{2}\right)\right] . \tag{D29}
\end{align*}
$$

(7) For the decays of massive spin-3 fields into massive vector and axial-vector fields, we use the following Lagrangian

$$
\begin{equation*}
\mathcal{L}_{w_{3} v_{1} a_{1}}=g_{w_{3} v_{1} a_{1}} \varepsilon^{\mu \nu \rho \sigma} W_{3, \mu \alpha \beta}\left(\partial_{\nu} V_{1, \rho}\right)\left(\partial^{\alpha} \partial^{\beta} A_{1, \sigma}\right) \tag{D30}
\end{equation*}
$$

to obtain the amplitude

$$
\begin{equation*}
\mathrm{i} \mathcal{M}_{w_{3} v_{1} a_{1}}=-\mathrm{i} g_{w_{3} v_{1} a_{1}} \varepsilon^{\mu \nu \rho \sigma} \epsilon_{\mu \alpha \beta}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) k_{v_{1} \nu} \epsilon_{\rho}\left(\lambda_{v_{1}}, \vec{k}_{v_{1}}\right) k_{a_{1}}^{\alpha} k_{a_{1}}^{\beta} \epsilon_{\sigma}\left(\lambda_{a_{1}}, \vec{k}_{a_{1}}\right) \tag{D31}
\end{equation*}
$$

The square of the amplitude is

$$
\begin{align*}
\frac{1}{7}\left|\mathcal{M}_{w_{3} v_{1} a_{1}}\right|^{2}= & g_{w_{3} v_{1} a_{1}}^{2} \frac{1}{7} \varepsilon^{\mu \nu \rho \sigma} \varepsilon^{\bar{\mu} \bar{\nu} \bar{\rho} \bar{\sigma}} \sum_{\lambda_{w_{3}}=-3}^{+3} \epsilon_{\mu \alpha \beta}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) \epsilon_{\bar{\mu} \bar{\alpha} \bar{\beta}}\left(\lambda_{w_{3}}, \vec{k}_{w_{3}}\right) k_{v_{1} \nu} k_{v_{1} \bar{\nu}} \sum_{\lambda_{v_{1}}=-1}^{+1} \epsilon_{\rho}\left(\lambda_{v_{1}}, \vec{k}_{v_{1}}\right) \epsilon_{\bar{\rho}}\left(\lambda_{v_{1}}, \vec{k}_{v_{1}}\right) \\
& \times k_{a_{1}}^{\alpha} k_{a_{1}}^{\beta} k_{a_{1}}^{\bar{\alpha}} k_{a_{1}}^{\bar{\beta}} \sum_{\lambda_{a_{1}}=-1}^{+1} \epsilon_{\sigma}\left(\lambda_{a_{1}}, \vec{k}_{a_{1}}\right) \epsilon_{\bar{\sigma}}\left(\lambda_{a_{1}}, \vec{k}_{a_{1}}\right) \\
= & g_{w_{3} v_{1} a_{1}}^{2} \frac{4}{105}\left|\vec{k}_{v_{1}, a_{1}}\right|^{[ }\left[\left|\vec{k}_{v_{1}, a_{1}}\right|^{2}\left(3+2 \frac{m_{w_{2}}^{2}}{m_{a_{1}}^{2}}\right)+7 m_{v_{1}}^{2}\right] \tag{D32}
\end{align*}
$$

We remark that the form of the amplitudes $\mathcal{M}$ of most of the decay channels was already derived in Ref. [16]. In contrast to Ref. [16] we neglect higher order derivative couplings and purely rely on the lowest order contribution in derivatives of the fields. Furthermore, we do not include form factors.

In order to obtain the results of this appendix one has to average over all incoming spin polarizations, sum up all possible outgoing polarizations, and consider the following
completeness relations for massive particles, which are discussed in the subsequent appendix,

$$
\begin{gather*}
\sum_{\lambda=-1}^{+1} \epsilon_{\mu}(\lambda, \vec{k}) \epsilon_{\nu}(\lambda, \vec{k})=-G_{\mu \nu}  \tag{D33}\\
\sum_{\lambda=-2}^{+2} \epsilon_{\mu \nu}(\lambda, \vec{k}) \epsilon_{\alpha \beta}(\lambda, \vec{k})=-\frac{1}{3} G_{\mu \nu} G_{\alpha \beta}+\frac{1}{2}\left(G_{\mu \alpha} G_{\nu \beta}+G_{\mu \beta} G_{\nu \alpha}\right), \tag{D34}
\end{gather*}
$$

$$
\begin{align*}
\sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k})= & \frac{1}{15}\left[G_{\mu \nu}\left(G_{\rho \alpha} G_{\beta \gamma}+G_{\rho \beta} G_{\alpha \gamma}+G_{\rho \gamma} G_{\alpha \beta}\right)+G_{\mu \rho}\left(G_{\nu \alpha} G_{\beta \gamma}+G_{\nu \beta} G_{\alpha \gamma}+G_{\nu \gamma} G_{\alpha \beta}\right)\right. \\
& \left.+G_{\nu \rho}\left(G_{\mu \alpha} G_{\beta \gamma}+G_{\mu \beta} G_{\alpha \gamma}+G_{\mu \gamma} G_{\alpha \beta}\right)\right] \\
& -\frac{1}{6}\left[G_{\mu \alpha}\left(G_{\nu \beta} G_{\rho \gamma}+G_{\nu \gamma} G_{\rho \beta}\right)+G_{\mu \beta}\left(G_{\nu \alpha} G_{\rho \gamma}+G_{\nu \gamma} G_{\rho \alpha}\right)+G_{\mu \gamma}\left(G_{\nu \alpha} G_{\rho \beta}+G_{\nu \beta} G_{\rho \alpha}\right)\right] \tag{D35}
\end{align*}
$$

where

$$
\begin{equation*}
G_{\mu \nu} \equiv \eta_{\mu \nu}-\frac{k_{\mu} k_{\nu}}{k^{2}}, \quad k^{2}=m^{2} \tag{D36}
\end{equation*}
$$

and $\left(\eta_{\mu \nu}\right)=\operatorname{diag}(+1,-1,-1,-1)$.
For the square of the amplitudes of the radioactive decays in Eq. (D17) the following completeness relation for photons is used [130]:

$$
\begin{align*}
& \sum_{\lambda=1}^{2} \epsilon_{\mu}(\lambda, \vec{k}) \epsilon_{\nu}(\lambda, \vec{k}) \\
& \quad=-\eta_{\mu \nu}-\frac{k_{\mu} k_{\nu}}{(k \cdot n)^{2}}+\frac{k_{\mu} n_{\nu}+k_{\nu} n_{\mu}}{(k \cdot n)} . \tag{D37}
\end{align*}
$$

Here, $n=(1,0,0,0)^{T}$ is the unit vector in temporal direction and $k$ is the four momentum of the outgoing photon.

The summations over Minkwoski spacetime indices in the squares of the amplitudes were performed on a symbolic level with Wolfram Mathematica 12.1 [131] using SUM and PARALLELSUM. We work in the center of mass frame of the decaying particle. In order to reduce the number of addends in the sums drastically, w.l.o.g. we choose the three momenta of the outgoing particles in the $z$ direction.

## APPENDIX E: COMPLETENESS RELATIONS

In this Appendix, we present how the completeness relations for the polarization vectors and tensors of massive higher-spin fields, like Eqs. (D33)-(D35), are constructed. These completeness relations are used for the calculation of the unpolarized decay amplitudes in the previous Appendix D.

We begin by deriving the completeness relation for massive spin-1 fields explicitly. This derivation is based on the discussion in Ref. [130] and presented here for reasons of consistency. Starting from this result for spin-1 fields, we generalize the procedure and show, how the completeness relations for massive higher-spin fields can be constructed. We then provide an explicit derivation of the completeness relations for massive spin-3 tensor fields. We close the discussion by elaborating on the degrees of freedom of massive fields of arbitrary integer spin- $J$ and their description via rank- $J$ tensor fields.

None of the results in this appendix is original to our work. However, we think that the discussion might still be useful for a nonexpert reader.

## 1. Spin- 1 fields

The construction of the completeness relation for the polarization vectors $\epsilon^{\mu}(\lambda, \vec{k})$ of the massive (real) vector fields $V^{\mu}(x)$ is based on the field equations that describe free massive vector particles. ${ }^{7}$ These field equations are given by the Proca equations, which are the Klein-Gordon (KG) equation and a transversality condition (which is explicitly not a gauge),

$$
\begin{equation*}
0=\left(\square+m^{2}\right) V^{\mu}(x), \quad 0=\partial_{\mu} V^{\mu}(x) \tag{E1}
\end{equation*}
$$

For neutral massive vector mesons, the solutions to the Klein-Gordon equation are plane waves,

$$
\begin{align*}
V^{\mu}(x)= & \int \frac{\mathrm{d}^{3} k}{(2 \pi)^{3}} \frac{1}{2 E(\vec{k})} \sum_{\lambda=-1}^{+1} \epsilon^{\mu}(\lambda, \vec{k})\left\{a(\lambda, \vec{k}) \mathrm{e}^{+\mathrm{i}[E(\vec{k}) t-\vec{k} \cdot \vec{x}]}\right. \\
& \left.+a^{*}(\lambda, \vec{k}) \mathrm{e}^{-\mathrm{i}[E(\vec{k}) t+\vec{k} \cdot \vec{x}]}\right\} \tag{E2}
\end{align*}
$$

[^6]where $a(\lambda, \vec{k})=a^{*}(\lambda,-\vec{k})$ are the Fourier amplitudes and $E^{2}(\vec{k})=\vec{k}^{2}+m^{2}$ is the square of the energy. The transversality condition further implies that the polarization vectors $\epsilon(\lambda, \vec{k})$ have to be transversal to the four momentum $k=(E(\vec{k}), \vec{k})^{T}$ of the fields,
\[

$$
\begin{equation*}
0=k_{\mu} \epsilon^{\mu}(\lambda, \vec{k}) \tag{E3}
\end{equation*}
$$

\]

From this equation, we can construct the completeness relation for the polarization vectors.

We start with the most general ansatz for a dimensionless rank-two tensor that only depends on the four momentum $k=(E(\vec{k}), \vec{k})^{T}$,

$$
\begin{equation*}
\sum_{\lambda=-1}^{+1} \epsilon_{\mu}(\lambda, \vec{k}) \epsilon_{\nu}(\lambda, \vec{k})=a \eta_{\mu \nu}+b \frac{k_{\mu} k_{\nu}}{k^{2}} \tag{E4}
\end{equation*}
$$

where $a$ and $b$ are dimensionless constants. The transversality condition (E3) implies

$$
\begin{equation*}
0 \stackrel{!}{=} k^{\mu} \sum_{\lambda=-1}^{+1} \epsilon_{\mu}(\lambda, \vec{k}) \epsilon_{\nu}(\lambda, \vec{k})=k_{\nu}(a+b) \tag{E5}
\end{equation*}
$$

which simplifies our ansatz to

$$
\begin{equation*}
\sum_{\lambda=-1}^{+1} \epsilon_{\mu}(\lambda, \vec{k}) \epsilon_{\nu}(\lambda, \vec{k})=a\left(\eta_{\mu \nu}-\frac{k_{\mu} k_{\nu}}{k^{2}}\right) \tag{E6}
\end{equation*}
$$

The coefficient $a$ can be fixed by choosing an appropriate normalization condition. This is done by demanding that in the rest frame $k=(m, \overrightarrow{0})^{T}$ for the spatial components the completeness relation should reduce to the Kronecker delta, to have a three-dimensional orthonormal basis of polarization vectors:

$$
\begin{equation*}
\delta_{i j} \stackrel{!}{=} \sum_{\lambda=-1}^{+1} \epsilon_{i}(\lambda, \overrightarrow{0}) \epsilon_{j}(\lambda, \overrightarrow{0})=-a \delta_{i j} \tag{E7}
\end{equation*}
$$

We obtain the completeness relation

$$
\begin{equation*}
\sum_{\lambda=-1}^{+1} \epsilon_{\mu}(\lambda, \vec{k}) \epsilon_{\nu}(\lambda, \vec{k})=-\left(\eta_{\mu \nu}-\frac{k_{\mu} k_{\nu}}{k^{2}}\right) \equiv-G_{\mu \nu} \tag{E8}
\end{equation*}
$$

and define the projector $G_{\mu \nu}$, which projects on the subspace of the Minkowski space that is transversal to the fourmomentum vector $k=(E(\vec{k}), \vec{k})^{T}$ of the massive spin-1 field.

## 2. Higher-spin fields

For all kind of higher-spin tensor fields, we proceed analogously. The starting point are the so called Fierz-Pauli equations for a massive field with integer spin $J$. The field is described by a rank- $J$ tensor field that fulfills the following equations (see also Refs. [130,132-137] for more details on wave equations and degrees of freedom of higher-spin fields),

$$
\begin{equation*}
0=\left(\square+m^{2}\right) \varphi^{\alpha_{1} \cdots \alpha_{J}}(x), \tag{E9}
\end{equation*}
$$

and $\forall 1 \leq m \neq n \leq J$,

$$
\begin{gather*}
0=\varphi^{\alpha_{1} \cdots \alpha_{m} \cdots \alpha_{n} \cdots \alpha_{J}}(x)-\varphi^{\alpha_{1} \cdots \alpha_{n} \cdots \alpha_{m} \cdots \alpha_{J}}(x),  \tag{E10}\\
0=\eta_{\alpha_{m} \alpha_{n}} \varphi^{\alpha_{1} \cdots \alpha_{m} \cdots \alpha_{n} \cdots \alpha_{J}}(x), \tag{E11}
\end{gather*}
$$

and $\forall 1 \leq m \leq J$,

$$
\begin{equation*}
0=\partial_{\alpha_{m}} \varphi^{\alpha_{1} \cdots \alpha_{m} \cdots \alpha_{J}}(x) \tag{E12}
\end{equation*}
$$

These equations reduce the $4^{J}$ degrees of freedom of a general rank- $J$ tensor field in four-dimensional Minkowski space to the $(2 J+1)$ degrees of freedom that are necessary to describe a massive spin- $J$ field. (We will come back to this point at the end of this Appendix.) Note that the Proca equations (E1) for spin-1 fields are merely a special case of the Fierz-Pauli equations (E9)-(E12).

Completely analogous to the vector field, we can thus find a solution to the KG equation and derive the following constraints for the corresponding polarization tensors, $\forall 1 \leq m \neq n \leq J:$

$$
\begin{gather*}
0=\epsilon^{\alpha_{1} \cdots \alpha_{m} \cdots \alpha_{n} \cdots \alpha_{J}}(\lambda, \vec{k})-\epsilon^{\alpha_{1} \cdots \alpha_{n} \cdots \alpha_{m} \cdots \alpha_{J}}(\lambda, \vec{k}),  \tag{E13}\\
0=\eta_{\alpha_{m} \alpha_{n}} \epsilon^{\alpha_{1} \cdots \alpha_{m} \cdots \alpha_{n} \cdots \alpha_{J}}(\lambda, \vec{k}), \tag{E14}
\end{gather*}
$$

$\forall 1 \leq m \leq J$,

$$
\begin{equation*}
0=k_{\alpha_{m}} \epsilon^{\alpha_{1} \cdots \alpha_{m} \cdots \alpha_{J}}(\lambda, \vec{k}) \tag{E15}
\end{equation*}
$$

Next, to find the corresponding completeness relation, one has to start with an appropriate ansatz for

$$
\begin{equation*}
\sum_{\lambda=-J}^{+J} \epsilon_{\alpha_{1} \cdots \alpha_{J}}(\lambda, \vec{k}) \epsilon_{\beta_{1} \cdots \beta_{J}}(\lambda, \vec{k})=\ldots \tag{E16}
\end{equation*}
$$

Consequently, one should write down the most general rank- $2 J$ tensor, which only depends on the four momentum $k=(E(\vec{k}), \vec{k})^{T}$ and the metric tensor. This means that all possible combinations of $\eta_{\mu \nu}$ and $\frac{k_{\mu} k_{\nu}}{k^{2}}$, with all possible permutations of the indices $\alpha_{1}, \ldots, \alpha_{J}, \beta_{1}, \ldots, \beta_{J}$ of the polarization tensors have to be included:

$$
\begin{align*}
\sum_{\lambda=-J}^{J} & \epsilon_{\alpha_{1} \cdots \alpha_{J}}(\lambda, \vec{k}) \epsilon_{\beta_{1} \cdots \beta_{J}}(\lambda, \vec{k}) \\
= & +a_{1} \eta_{\alpha_{1} \alpha_{2}} \cdots \eta_{\beta_{J-1} \beta_{J}}+\text { all permutations } \\
& +b_{1} \eta_{\alpha_{1} \alpha_{2}} \cdots \frac{k_{\beta_{J-1}} k_{\beta_{J}}}{k^{2}}+\text { all permutations } \\
& +\ldots \\
& +z_{1} \frac{k_{\alpha_{1}} k_{\alpha_{2}}}{k^{2}} \cdots \frac{k_{\beta_{J-1}} k_{\beta_{J}}}{k^{2}}+\text { all permutations. } \tag{E17}
\end{align*}
$$

Here, $a_{i}, b_{i}, \ldots, z_{i}$ stand for the various dimensionless coefficients of the ansatz. An alternative approach is to write down the ansatz for the rank- $2 J$ tensor on the rhs of the above equation in terms of projection operators

$$
\begin{equation*}
G_{\mu \nu}=\eta_{\mu \nu}-\frac{k_{\mu} k_{\nu}}{k^{2}}, \quad P_{\mu \nu}=\frac{k_{\mu} k_{\nu}}{k^{2}} \tag{E18}
\end{equation*}
$$

which also form a complete basis.
What follows is identical for both approaches. To determine the various coefficients, one has to use Eqs. (E13)-(E15) (or combinations of these constraints) successively until one reaches the point, where only one overall coefficient/factor is left, see for example Ref. [132] for massive spin-2 fields. The leftover constant can only be fixed by appropriate normalization. For most problems, a normalization, which produces an orthonormal basis of polarization tensors in the rest frame, like for the vectors in Eq. (E7), is suitable.

Following these procedures, one eventually arrives at the completeness relations (D34) and (D35) for the polarization tensors of spin-2 and spin-3 fields.

We remark, that an expression for arbitrary spin fields was already presented in Refs. [19,138]. The expressions for spin-2 and spin-3 were already presented among others in Refs. [16-18,132,139].

## 3. Spin-3 fields revisited

Here, we explicitly present the derivation of the completeness relation for massive spin-3 fields.

We start by adopting the constraints (E13), (E14), and (E15) to the polarization tensors of rank-3 tensor fields. Thus, the polarization tensors of spin-3 fields have to fulfill the following conditions,

$$
\begin{align*}
& 0=\epsilon^{(\mu \nu \rho)}(\lambda, \vec{k})  \tag{E19}\\
& 0=\epsilon_{\mu}{ }^{\mu \nu}(\lambda, \vec{k})  \tag{E20}\\
& 0=k_{\mu} \epsilon^{\mu \nu \rho}(\lambda, \vec{k}) \tag{E21}
\end{align*}
$$

Next we construct the completeness relation for these polarization tensors. We use the definition

$$
\begin{equation*}
G_{\mu \nu} \equiv \eta_{\mu \nu}-\frac{k_{\mu} k_{\nu}}{k^{2}} \tag{E22}
\end{equation*}
$$

and write down the most general ansatz for the completeness relation, which contains all possible combinations of the metric and the four momentum or the projection operators (E18), respectively,

$$
\begin{align*}
& \sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k})=+a_{1} G_{\mu \nu} G_{\rho \alpha} G_{\beta \gamma}+a_{2} G_{\mu \nu} G_{\rho \beta} G_{\alpha \gamma}+a_{3} G_{\mu \nu} G_{\rho \gamma} G_{\alpha \beta}+a_{4} G_{\mu \rho} G_{\nu \alpha} G_{\beta \gamma}+a_{5} G_{\mu \rho} G_{\nu \beta} G_{\alpha \gamma} \\
& +a_{6} G_{\mu \rho} G_{\nu \gamma} G_{\alpha \beta}+a_{7} G_{\mu \alpha} G_{\nu \rho} G_{\beta \gamma}+a_{8} G_{\mu \alpha} G_{\nu \beta} G_{\rho \gamma}+a_{9} G_{\mu \alpha} G_{\nu \gamma} G_{\rho \beta}+a_{10} G_{\mu \beta} G_{\nu \rho} G_{\alpha \gamma} \\
& +a_{11} G_{\mu \beta} G_{\nu \alpha} G_{\rho \gamma}+a_{12} G_{\mu \beta} G_{\nu \gamma} G_{\rho \alpha}+a_{13} G_{\mu \gamma} G_{\nu \rho} G_{\alpha \beta}+a_{14} G_{\mu \gamma} G_{\nu \alpha} G_{\rho \beta}+a_{15} G_{\mu \gamma} G_{\nu \beta} G_{\rho \alpha} \\
& +b_{1} G_{\mu \nu} G_{\rho \alpha} k_{\beta} k_{\gamma}+b_{2} G_{\mu \nu} G_{\rho \beta} k_{\alpha} k_{\gamma}+b_{3} G_{\mu \nu} G_{\rho \gamma} k_{\alpha} k_{\beta}+b_{4} G_{\mu \nu} G_{\alpha \beta} k_{\rho} k_{\gamma}+b_{5} G_{\mu \nu} G_{\alpha \gamma} k_{\rho} k_{\beta} \\
& +b_{6} G_{\mu \nu} G_{\beta \gamma} k_{\rho} k_{\alpha}+b_{7} G_{\mu \rho} G_{\nu \alpha} k_{\beta} k_{\gamma}+b_{8} G_{\mu \rho} G_{\nu \beta} k_{\alpha} k_{\gamma}+b_{9} G_{\mu \rho} G_{\nu \gamma} k_{\alpha} k_{\beta}+b_{10} G_{\mu \rho} G_{\alpha \beta} k_{\nu} k_{\gamma} \\
& +b_{11} G_{\mu \rho} G_{\alpha \gamma} k_{\nu} k_{\beta}+b_{12} G_{\mu \rho} G_{\beta \gamma} k_{\nu} k_{\alpha}+b_{13} G_{\mu \alpha} G_{\nu \rho} k_{\beta} k_{\gamma}+b_{14} G_{\mu \alpha} G_{\nu \beta} k_{\rho} k_{\gamma}+b_{15} G_{\mu \alpha} G_{\nu \gamma} k_{\rho} k_{\beta} \\
& +b_{16} G_{\mu \alpha} G_{\rho \beta} k_{\nu} k_{\gamma}+b_{17} G_{\mu \alpha} G_{\rho \gamma} k_{\nu} k_{\beta}+b_{18} G_{\mu \alpha} G_{\beta \gamma} k_{\nu} k_{\rho}+b_{19} G_{\mu \beta} G_{\nu \rho} k_{\alpha} k_{\gamma}+b_{20} G_{\mu \beta} G_{\nu \alpha} k_{\rho} k_{\gamma} \\
& +b_{21} G_{\mu \beta} G_{\nu \gamma} k_{\rho} k_{\alpha}+b_{22} G_{\mu \beta} G_{\rho \alpha} k_{\nu} k_{\gamma}+b_{23} G_{\mu \beta} G_{\rho \gamma} k_{\nu} k_{\alpha}+b_{24} G_{\mu \beta} G_{\alpha \gamma} k_{\nu} k_{\rho}+b_{25} G_{\mu \gamma} G_{\nu \rho} k_{\alpha} k_{\beta} \\
& +b_{26} G_{\mu \gamma} G_{\nu \alpha} k_{\rho} k_{\beta}+b_{27} G_{\mu \gamma} G_{\nu \beta} k_{\rho} k_{\alpha}+b_{28} G_{\mu \gamma} G_{\rho \alpha} k_{\nu} k_{\beta}+b_{29} G_{\mu \gamma} G_{\rho \beta} k_{\nu} k_{\alpha}+b_{30} G_{\mu \gamma} G_{\alpha \beta} k_{\nu} k_{\rho} \\
& +b_{31} k_{\mu} k_{\nu} G_{\rho \alpha} G_{\beta \gamma}+b_{32} k_{\mu} k_{\nu} G_{\rho \beta} G_{\alpha \gamma}+b_{33} k_{\mu} k_{\nu} G_{\rho \gamma} G_{\alpha \beta}+b_{34} k_{\mu} k_{\rho} G_{\nu \alpha} G_{\beta \gamma}+b_{35} k_{\mu} k_{\rho} G_{\nu \beta} G_{\alpha \gamma} \\
& +b_{36} k_{\mu} k_{\rho} G_{\nu \gamma} G_{\alpha \beta}+b_{37} k_{\mu} k_{\alpha} G_{\nu \rho} G_{\beta \gamma}+b_{38} k_{\mu} k_{\alpha} G_{\nu \beta} G_{\rho \gamma}+b_{39} k_{\mu} k_{\alpha} G_{\nu \gamma} G_{\rho \beta}+b_{40} k_{\mu} k_{\beta} G_{\nu \rho} G_{\alpha \gamma} \\
& +b_{41} k_{\mu} k_{\beta} G_{\nu \alpha} G_{\rho \gamma}+b_{42} k_{\mu} k_{\beta} G_{\nu \gamma} G_{\rho \alpha}+b_{43} k_{\mu} k_{\gamma} G_{\nu \rho} G_{\alpha \beta}+b_{44} k_{\mu} k_{\gamma} G_{\nu \alpha} G_{\rho \beta}+b_{45} k_{\mu} k_{\gamma} G_{\nu \beta} G_{\rho \alpha} \\
& +c_{1} G_{\mu \nu} k_{\rho} k_{\alpha} k_{\beta} k_{\gamma}+c_{2} G_{\mu \rho} k_{\nu} k_{\alpha} k_{\beta} k_{\gamma}+c_{3} G_{\mu \alpha} k_{\nu} k_{\rho} k_{\beta} k_{\gamma}+c_{4} G_{\mu \beta} k_{\nu} k_{\rho} k_{\alpha} k_{\gamma}+c_{5} G_{\mu \gamma} k_{\nu} k_{\rho} k_{\alpha} k_{\beta} \\
& +c_{6} G_{\nu \rho} k_{\mu} k_{\alpha} k_{\beta} k_{\gamma}+c_{7} G_{\nu \alpha} k_{\mu} k_{\rho} k_{\beta} k_{\gamma}+c_{8} G_{\nu \beta} k_{\mu} k_{\rho} k_{\alpha} k_{\gamma}+c_{9} G_{\nu \gamma} k_{\mu} k_{\rho} k_{\alpha} k_{\beta}+c_{10} G_{\rho \alpha} k_{\mu} k_{\nu} k_{\beta} k_{\gamma} \\
& +c_{11} G_{\rho \beta} k_{\mu} k_{\nu} k_{\alpha} k_{\gamma}+c_{12} G_{\rho \gamma} k_{\mu} k_{\nu} k_{\alpha} k_{\beta}+c_{13} G_{\alpha \beta} k_{\mu} k_{\nu} k_{\rho} k_{\gamma}+c_{14} G_{\alpha \gamma} k_{\mu} k_{\nu} k_{\rho} k_{\beta}+c_{15} G_{\beta \gamma} k_{\mu} k_{\nu} k_{\rho} k_{\alpha} \\
& +d k_{\mu} k_{\nu} k_{\rho} k_{\alpha} k_{\beta} k_{\gamma} \text {. } \tag{E23}
\end{align*}
$$

Next, we have to determine the coefficients $a_{i}, b_{i}, c_{i}$, and $d$ by applying the constraints (E19), (E20), and (E21). Note, that for convenience only, the coefficients $a_{i}$ are dimensionless, while $b_{i}, c_{i}$, and $d$ contain factors of $\frac{1}{k^{2}}$, in order to have the rhs of the equation dimensionless.

First, we use Eq. (E19). The symmetry under the exchange of $\mu \leftrightarrow \nu$ yields

$$
\begin{array}{rlrl}
a_{4} & =a_{7}, & & a_{5}=a_{10}, \\
a_{8} & =a_{11}, & & a_{6}=a_{13}, \\
b_{7} & =b_{13}, & & a_{12}=a_{15}, \\
b_{10} & =b_{43}, & b_{8}=b_{19}, & \\
b_{11}=b_{40}, & & b_{12}=b_{25}, \\
b_{14} & =b_{20}, \\
b_{17} & =b_{41}, & & b_{15}=b_{26}, \\
b_{18}=b_{34}, & & b_{16}=b_{44}, \\
b_{22} & =b_{25}, & & b_{23}=b_{27}, \\
b_{28} & =b_{42}, & & b_{24}=b_{35},  \tag{E24}\\
c_{2} & =c_{6}, & c_{3}=b_{39}, & b_{30}=b_{36}, \\
c_{5} & =c_{9}, & &
\end{array}
$$

whereas for the symmetry of the completeness relation under the exchange $\nu \leftrightarrow \rho$ we obtain

$$
\begin{array}{rlrl}
a_{1} & =a_{4}, & & a_{2}=a_{5}, \\
a_{8} & =a_{9}, & & a_{11}=a_{12}, \\
b_{1} & =b_{7}, & a_{14}=a_{15}, \\
b_{4} & =b_{10}, & b_{2}=b_{8}, & b_{3}=b_{9}, \\
b_{14} & =b_{16}, & & b_{6}=b_{12}, \\
b_{21} & =b_{23}, & b_{26}=b_{17}, & b_{20}=b_{28}, \\
b_{31} & =b_{34}, & b_{27}=b_{29}, \\
b_{38} & =b_{39}, & b_{41}=b_{35}, & b_{33}=b_{36}, \\
c_{1} & =c_{2}, & c_{7}=b_{10}, & c_{8}=b_{45}, \\
c_{9} & =c_{12} . & & \tag{E25}
\end{array}
$$

The last symmetry, which includes all other possible index interchanges, is $\mu \nu \rho \leftrightarrow \alpha \beta \gamma$, from which follows that

$$
\begin{array}{rlrl}
a_{1} & =a_{13}, & & a_{2}=a_{6}, \\
a_{12} & =a_{14}, & & a_{4}=a_{10}, \\
b_{3} & =b_{33}, & & b_{2}=b_{36}, \\
b_{7} & =b_{24}, & & b_{5}=b_{10}, \\
b_{8}=b_{35}, & & b_{6}=b_{43}, \\
b_{12} & =b_{40}, \\
b_{19} & =b_{32}, \\
b_{23} & & b_{13}=b_{18}, & \\
b_{41}, & & b_{21}=b_{15}=b_{16}, \\
b_{25} & & b_{22}=b_{26},  \tag{E26}\\
b_{29} & =b_{42}, & & b_{27}=b_{45}, \\
c_{4} & =c_{7}, & & c_{5}=c_{10}, \\
c_{9} & =c_{11}, & & c_{6}=c_{14}, \\
& & &
\end{array}
$$

Therefore, we define new coefficients

$$
\begin{aligned}
& a \equiv a_{1}=a_{2}=a_{3}=a_{4}=a_{5}=a_{6}=a_{7}=a_{10}=a_{13} \\
& a^{\prime} \equiv a_{8}=a_{9}=a_{11}=a_{12}=a_{14}=a_{15} \\
& b \equiv b_{1}=b_{2}=b_{3}=b_{7}=b_{8}=b_{9}=b_{13}=b_{18}=b_{19} \\
&=b_{24}=b_{25}=b_{30}=b_{31}=b_{32}=b_{33}=b_{34}=b_{35} \\
&=b_{36} \\
& b^{\prime} \equiv b_{14}=b_{15}=b_{16}=b_{17}=b_{20}=b_{21}=b_{22}=b_{23} \\
&=b_{26}=b_{27}=b_{28}=b_{29}=b_{38}=b_{39}=b_{41}=b_{42} \\
&=b_{44}=b_{45} \\
& \tilde{b} \equiv b_{4}=b_{5}=b_{6}=b_{10}=b_{11}=b_{12}=b_{37}=b_{40}=b_{43} \\
& c \equiv c_{1}=c_{2}=c_{6}=c_{13}=c_{14}=c_{15} \\
& c^{\prime} \equiv c_{3}=c_{4}=c_{5}=c_{7}=c_{8}=c_{9}=c_{10}=c_{11}=c_{12}
\end{aligned}
$$

and replace the coefficients of our ansatz:

$$
\begin{aligned}
\sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k})= & +a G_{\mu \nu} G_{\rho \alpha} G_{\beta \gamma}+a G_{\mu \nu} G_{\rho \beta} G_{\alpha \gamma}+a G_{\mu \nu} G_{\rho \gamma} G_{\alpha \beta}+a G_{\mu \rho} G_{\nu \alpha} G_{\beta \gamma}+a G_{\mu \rho} G_{\nu \beta} G_{\alpha \gamma}+a G_{\mu \rho} G_{\nu \gamma} G_{\alpha \beta} \\
& +a G_{\mu \alpha} G_{\nu \rho} G_{\beta \gamma}+a^{\prime} G_{\mu \alpha} G_{\nu \beta} G_{\rho \gamma}+a^{\prime} G_{\mu \alpha} G_{\nu \gamma} G_{\rho \beta}+a G_{\mu \beta} G_{\nu \rho} G_{\alpha \gamma}+a^{\prime} G_{\mu \beta} G_{\nu \alpha} G_{\rho \gamma}+a^{\prime} G_{\mu \beta} G_{\nu \gamma} G_{\rho \alpha} \\
& +a G_{\mu \gamma} G_{\nu \rho} G_{\alpha \beta}+a^{\prime} G_{\mu \gamma} G_{\nu \alpha} G_{\rho \beta}+a^{\prime} G_{\mu \gamma} G_{\nu \beta} G_{\rho \alpha}+b G_{\mu \nu} G_{\rho \alpha} k_{\beta} k_{\gamma}+b G_{\mu \nu} G_{\rho \beta} k_{\alpha} k_{\gamma} \\
& +b G_{\mu \nu} G_{\rho \gamma} k_{\alpha} k_{\beta}+\tilde{b} G_{\mu \nu} G_{\alpha \beta} k_{\rho} k_{\gamma}+\tilde{b} G_{\mu \nu} G_{\alpha \gamma} k_{\rho} k_{\beta}+\tilde{b} G_{\mu \nu} G_{\beta \gamma} k_{\rho} k_{\alpha}+b G_{\mu \rho} G_{\nu \alpha} k_{\beta} k_{\gamma} \\
& +b G_{\mu \rho} G_{\nu \beta} k_{\alpha} k_{\gamma}+b G_{\mu \rho} G_{\nu \gamma} k_{\alpha} k_{\beta}+\tilde{b} G_{\mu \rho} G_{\alpha \beta} k_{\nu} k_{\gamma}+\tilde{b} G_{\mu \rho} G_{\alpha \gamma} k_{\nu} k_{\beta}+\tilde{b} G_{\mu \rho} G_{\beta \gamma} k_{\nu} k_{\alpha} \\
& +b G_{\mu \alpha} G_{\nu \rho} k_{\beta} k_{\gamma}+b^{\prime} G_{\mu \alpha} G_{\nu \beta} k_{\rho} k_{\gamma}+b^{\prime} G_{\mu \alpha} G_{\nu \gamma} k_{\rho} k_{\beta}+b^{\prime} G_{\mu \alpha} G_{\rho \beta} k_{\nu} k_{\gamma}+b^{\prime} G_{\mu \alpha} G_{\rho \gamma} k_{\nu} k_{\beta} \\
& +b G_{\mu \alpha} G_{\beta \gamma} k_{\nu} k_{\rho}+b G_{\mu \beta} G_{\nu \rho} k_{\alpha} k_{\gamma}+b^{\prime} G_{\mu \beta} G_{\nu \alpha} k_{\rho} k_{\gamma}+b^{\prime} G_{\mu \beta} G_{\nu \gamma} k_{\rho} k_{\alpha}+b^{\prime} G_{\mu \beta} G_{\rho \alpha} k_{\nu} k_{\gamma} \\
& +b^{\prime} G_{\mu \beta} G_{\rho \gamma} k_{\nu} k_{\alpha}+b G_{\mu \beta} G_{\alpha \gamma} k_{\nu} k_{\rho}+b G_{\mu \gamma} G_{\nu \rho} k_{\alpha} k_{\beta}+b^{\prime} G_{\mu \gamma} G_{\nu \alpha} k_{\rho} k_{\beta}+b^{\prime} G_{\mu \gamma} G_{\nu \beta} k_{\rho} k_{\alpha} \\
& +b^{\prime} G_{\mu \gamma} G_{\rho \alpha} k_{\nu} k_{\beta}+b^{\prime} G_{\mu \gamma} G_{\rho \beta} k_{\nu} k_{\alpha}+b G_{\mu \gamma} G_{\alpha \beta} k_{\nu} k_{\rho}+b k_{\mu} k_{\nu} G_{\rho \alpha} G_{\beta \gamma}+b k_{\mu} k_{\nu} G_{\rho \beta} G_{\alpha \gamma}
\end{aligned}
$$

$$
\begin{align*}
& +b k_{\mu} k_{\nu} G_{\rho \gamma} G_{\alpha \beta}+b k_{\mu} k_{\rho} G_{\nu \alpha} G_{\beta \gamma}+b k_{\mu} k_{\rho} G_{\nu \beta} G_{\alpha \gamma}+b k_{\mu} k_{\rho} G_{\nu \gamma} G_{\alpha \beta}+\tilde{b} k_{\mu} k_{\alpha} G_{\nu \rho} G_{\beta \gamma}+b^{\prime} k_{\mu} k_{\alpha} G_{\nu \beta} G_{\rho \gamma} \\
& +b^{\prime} k_{\mu} k_{\alpha} G_{\nu \gamma} G_{\rho \beta}+\tilde{b} k_{\mu} k_{\beta} G_{\nu \rho} G_{\alpha \gamma}+b^{\prime} k_{\mu} k_{\beta} G_{\nu \alpha} G_{\rho \gamma}+b^{\prime} k_{\mu} k_{\beta} G_{\nu \gamma} G_{\rho \alpha}+\tilde{b} k_{\mu} k_{\gamma} G_{\nu \rho} G_{\alpha \beta}+b^{\prime} k_{\mu} k_{\gamma} G_{\nu \alpha} G_{\rho \beta} \\
& +b^{\prime} k_{\mu} k_{\gamma} G_{\nu \beta} G_{\rho \alpha}+c G_{\mu \nu} k_{\rho} k_{\alpha} k_{\beta} k_{\gamma}+c G_{\mu \rho} k_{\nu} k_{\alpha} k_{\beta} k_{\gamma}+c^{\prime} G_{\mu \alpha} k_{\nu} k_{\rho} k_{\beta} k_{\gamma}+c^{\prime} G_{\mu \beta} k_{\nu} k_{\rho} k_{\alpha} k_{\gamma}+c^{\prime} G_{\mu \gamma} k_{\nu} k_{\rho} k_{\alpha} k_{\beta} \\
& +c G_{\nu \rho} k_{\mu} k_{\alpha} k_{\beta} k_{\gamma}+c^{\prime} G_{\nu \alpha} k_{\mu} k_{\rho} k_{\beta} c_{\gamma} G_{\nu \beta} k_{\mu} k_{\rho} k_{\alpha}+c^{\prime} G_{\nu \gamma} k_{\mu} k_{\rho} k_{\alpha}+c^{\prime} G_{\rho \alpha} k_{\mu} k_{\nu} k_{\gamma}+c^{\prime} G_{\rho \beta} k_{\mu} k_{\nu} k_{\alpha} k_{\gamma} \\
& +c^{\prime} G_{\rho \gamma} k_{\mu} k_{\nu} k_{\alpha} k_{\beta}+c G_{\alpha \beta} k_{\mu} k_{\nu} k_{\rho} k_{\gamma}+c G_{\alpha \gamma} k_{\mu} k_{\nu} k_{\rho} k_{\beta}+c G_{\beta \gamma} k_{\mu} k_{\nu} k_{\rho} k_{\alpha}+d k_{\mu} k_{\nu} k_{\rho} k_{\alpha} k_{\beta} k_{\gamma} . \tag{E27}
\end{align*}
$$

To determine the remaining eight constants, we use Eqs. (E20) and (E21). We note that

$$
\begin{equation*}
k^{\mu} G_{\mu \nu}=0 \tag{E28}
\end{equation*}
$$

and contract the last expression with $k^{\mu} k^{\nu} k^{\rho} k^{\alpha}$,

$$
\begin{align*}
0 & =k^{\mu} k^{\nu} k^{\rho} k^{\alpha} \sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k}) \\
& =\left(k^{2}\right)^{4}\left(c G_{\beta \gamma}+d k_{\beta} k_{\gamma}\right) \tag{E29}
\end{align*}
$$

The only solution to this equation is $c=d=0$, which eliminates two of the remaining constants. In complete analogy the contraction with $k^{\mu} k^{\nu} k^{\alpha} k^{\beta}$ leads to

$$
\begin{align*}
0 & =k^{\mu} k^{\nu} k^{\alpha} k^{\beta} \sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k}) \\
& =\left(k^{2}\right)^{4} c^{\prime} G_{\rho \gamma} \tag{E30}
\end{align*}
$$

which sets $c^{\prime}=0$. Next, we contract with $k^{\mu} k^{\nu} \eta^{\rho \alpha} \eta^{\beta \gamma}$ and find

$$
\begin{align*}
0 & =k^{\mu} k^{\nu} \eta^{\rho \alpha} \eta^{\beta \gamma} \sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k}) \\
& =\left(k^{2}\right)^{2} b\left(G_{\mu}{ }^{\mu} G_{\nu}{ }^{\nu}+2 G_{\mu \nu} G^{\mu \nu}\right) \\
& =15 b\left(k^{2}\right)^{2} \tag{E31}
\end{align*}
$$

which states that also $b=0$. To eliminate $b^{\prime}$ and $\tilde{b}$, we have to contract the whole expression with $k^{\mu} k^{\alpha} \eta^{\nu \beta} \eta^{\rho \gamma}$,

$$
\begin{align*}
0 & =k^{\mu} k^{\alpha} \eta^{\nu \beta} \eta^{\rho \gamma} \sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k}) \\
& =\left(k^{2}\right)^{2}\left(3 \tilde{b}+12 b^{\prime}\right) \tag{E32}
\end{align*}
$$

and with $k^{\mu} k^{\alpha} \eta^{\nu \rho} \eta^{\beta \gamma}$,

$$
\begin{align*}
0 & =k^{\mu} k^{\alpha} \eta^{\nu \rho} \eta^{\beta \gamma} \sum_{\lambda=1}^{7} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k}) \\
& =\left(k^{2}\right)^{2}\left(9 \tilde{b}+6 b^{\prime}\right) \tag{E33}
\end{align*}
$$

Thus, we conclude $b^{\prime}=\tilde{b}=0$. Before we calculate the relation between the coefficients $a$ and $a^{\prime}$, we provide the remainder expression for the completeness relation.

$$
\begin{align*}
\sum_{\lambda=1}^{7} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k})= & +a G_{\mu \nu} G_{\rho \alpha} G_{\beta \gamma}+a G_{\mu \nu} G_{\rho \beta} G_{\alpha \gamma}+a G_{\mu \nu} G_{\rho \gamma} G_{\alpha \beta}+a G_{\mu \rho} G_{\nu \alpha} G_{\beta \gamma}+a G_{\mu \rho} G_{\nu \beta} G_{\alpha \gamma}+a G_{\mu \rho} G_{\nu \gamma} G_{\alpha \beta} \\
& +a G_{\mu \alpha} G_{\nu \rho} G_{\beta \gamma}+a^{\prime} G_{\mu \alpha} G_{\nu \beta} G_{\rho \gamma}+a^{\prime} G_{\mu \alpha} G_{\nu \gamma} G_{\rho \beta}+a G_{\mu \beta} G_{\nu \rho} G_{\alpha \gamma}+a^{\prime} G_{\mu \beta} G_{\nu \alpha} G_{\rho \gamma}+a^{\prime} G_{\mu \beta} G_{\nu \gamma} G_{\rho \alpha} \\
& +a G_{\mu \gamma} G_{\nu \rho} G_{\alpha \beta}+a^{\prime} G_{\mu \gamma} G_{\nu \alpha} G_{\rho \beta}+a^{\prime} G_{\mu \gamma} G_{\nu \beta} G_{\rho \alpha} . \tag{E34}
\end{align*}
$$

The easiest way to eliminate $a^{\prime}$ in favor of $a$ is, to contract this expression with $\eta^{\mu \nu} \eta^{\rho \alpha} \eta^{\beta \gamma}$,

$$
\begin{equation*}
0=\eta^{\mu \nu} \eta^{\rho \alpha} \eta^{\beta \gamma} \sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k})=75 a+30 a^{\prime} \tag{E35}
\end{equation*}
$$

Thus, $a^{\prime}=-\frac{5}{2} a$. Reordering the completeness relation gives

$$
\begin{align*}
\sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k})= & a\left\{G_{\mu \nu}\left(G_{\rho \alpha} G_{\beta \gamma}+G_{\rho \beta} G_{\alpha \gamma}+G_{\rho \gamma} G_{\alpha \beta}\right)\right. \\
& +G_{\mu \rho}\left(G_{\nu \alpha} G_{\beta \gamma}+G_{\nu \beta} G_{\alpha \gamma}+G_{\nu \gamma} G_{\alpha \beta}\right)+G_{\nu \rho}\left(G_{\mu \alpha} G_{\beta \gamma}+G_{\mu \beta} G_{\alpha \gamma}+G_{\mu \gamma} G_{\alpha \beta}\right) \\
& \left.-\frac{5}{2}\left[G_{\mu \alpha}\left(G_{\nu \beta} G_{\rho \gamma}+G_{\nu \gamma} G_{\rho \beta}\right)+G_{\mu \beta}\left(G_{\nu \alpha} G_{\rho \gamma}+G_{\nu \gamma} G_{\rho \alpha}\right)+G_{\mu \gamma}\left(G_{\nu \alpha} G_{\rho \beta}+G_{\nu \beta} G_{\rho \alpha}\right)\right]\right\} \tag{E36}
\end{align*}
$$

The last constant $a$ can be chosen arbitrarily. Nevertheless it might be convenient to orthonormalize the polarization tensors,

$$
\begin{equation*}
\epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon^{\mu \nu \rho}\left(\lambda^{\prime}, \vec{k}\right)=-\delta_{\lambda \lambda^{\prime}} \tag{E37}
\end{equation*}
$$

which is the higher-spin equivalent to the orthonormality condition Eq. (E7). Then it directly follows that

$$
\begin{equation*}
-7=-\sum_{\lambda-3}^{+3} \delta_{\lambda \lambda}=\sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon^{\mu \nu \rho}(\lambda, \vec{k})=-105 a \tag{E38}
\end{equation*}
$$

The final result is

$$
\begin{align*}
\sum_{\lambda=-3}^{+3} \epsilon_{\mu \nu \rho}(\lambda, \vec{k}) \epsilon_{\alpha \beta \gamma}(\lambda, \vec{k})= & \frac{1}{15}\left[G_{\mu \nu}\left(G_{\rho \alpha} G_{\beta \gamma}+G_{\rho \beta} G_{\alpha \gamma}+G_{\rho \gamma} G_{\alpha \beta}\right)\right. \\
& \left.+G_{\mu \rho}\left(G_{\nu \alpha} G_{\beta \gamma}+G_{\nu \beta} G_{\alpha \gamma}+G_{\nu \gamma} G_{\alpha \beta}\right)+G_{\nu \rho}\left(G_{\mu \alpha} G_{\beta \gamma}+G_{\mu \beta} G_{\alpha \gamma}+G_{\mu \gamma} G_{\alpha \beta}\right)\right] \\
& -\frac{1}{6}\left[G_{\mu \alpha}\left(G_{\nu \beta} G_{\rho \gamma}+G_{\nu \gamma} G_{\rho \beta}\right)+G_{\mu \beta}\left(G_{\nu \alpha} G_{\rho \gamma}+G_{\nu \gamma} G_{\rho \alpha}\right)+G_{\mu \gamma}\left(G_{\nu \alpha} G_{\rho \beta}+G_{\nu \beta} G_{\rho \alpha}\right)\right] \tag{E39}
\end{align*}
$$

## 4. Spin- $J$ tensor degrees of freedom

As a last step, we recapitulate for the sake of completeness, how to determine the number of degrees of freedom for a massive bosonic field (see Refs. [133,134] for the original discussion). A Lorentz tensor $\phi^{\cdots \mu \nu \ldots}$ of arbitrary rank $J$ in four-dimensional Minkowski space has $4^{J}$ degrees of freedom. The Klein-Gordon equation (E9) does not reduce the degrees of freedom, whereas the symmetry condition (E10) does. A combinatorial analysis is necessary and we use for simplicity an urn model: for each index of the rank- $J$ tensor components there are four possible entries, which is equivalent to four different balls in the urn. The tensor field is symmetric in all indices, which means that the order of drawings does not matter. Indices can also be "drawn" several times. This leads to

$$
\begin{equation*}
\frac{(4-1+J)!}{J!(4-1)!}=\binom{4-1+J}{4-1} \tag{E40}
\end{equation*}
$$

possibilities, or better degrees of freedom. The next reduction is done using the traceless condition (E11), which leads to

$$
\begin{equation*}
\binom{4-1+(J-2)}{4-1} \tag{E41}
\end{equation*}
$$

additional constraints [this is exactly the number of degrees of freedom of a symmetric tensor of rank $(J-2)$ ]. The last constraint, also refereed to as Lorenz constraint or transversality constraint (E12), leads to restrictions:

$$
\begin{equation*}
0=\partial_{\mu} \phi^{\cdots \mu \nu \cdots} \tag{E42}
\end{equation*}
$$

Now it is easy to conclude, that the number of these constraints are exactly the degrees of freedom of a rank-$(J-1)$ symmetric traceless tensor, which are

$$
\begin{equation*}
\binom{4-1+(J-1)}{4-1}-\binom{4-1+(J-3)}{4-1} \tag{E43}
\end{equation*}
$$

In total, the Fierz-Pauli equations lead to

$$
\begin{align*}
& \binom{4-1+J}{4-1}-\binom{4-1+(J-2)}{4-1} \\
& \quad-\left[\binom{4-1+(J-1)}{4-1}-\binom{4-1+(J-3)}{4-1}\right] \\
& \quad=2 J+1 . \tag{E44}
\end{align*}
$$

degrees of freedom for a rank- $J$ massive tensor field.

On the other hand, we know that a massive particle of total spin- $J$ has also $2 J+1$ degrees of freedom, which correspond to the $2 J+1$ possible orientations of the particles spin in a magnetic field. Furthermore, one could
study the behavior of particles of spin- $J$ under rotations or Lorentz transformations, which would also emphasize the use massive rank- $J$ tensor fields to describe those spin- $J$ particles.
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[^0]:    ${ }^{1}$ Other mesonic nonets constructed in the same way can be found for instance in Refs. [34,37,53,79].

[^1]:    ${ }^{2}$ We refer to Appendix A for the description of the mass terms of the corresponding Lagrangian(s) and for the derivation of the PDG-mixing formula, which was applied in the PDG [1] in the case of ground-state vector and pseudoscalar mesons as well as for the ground-state tensor mesons with $J^{P C}=2^{++}$and $J^{P C}=3^{--}$. We shall also present the link between the singletoctet basis used in the PDG and the strange-nonstrange one employed in this work.

[^2]:    ${ }^{3}$ Note, Ref. [16] included higher derivative couplings and form factors in their approach to strong decays of spin-3 mesons involving charm quarks.

[^3]:    ${ }^{4}$ Since the decaying particles are spin- 3 mesons we multiply by the factor $\frac{1}{7}$, averaging over the different spin states.

[^4]:    ${ }^{5} \mathrm{We}$ are aware of the fact that some of the data might be correlated because it is taken from similar measurements at the same experiments. Within the level of accuracy of our approach, there is no need to take care of these effects.

[^5]:    ${ }^{6}$ For massive spin-3 tensor fields $\lambda=-3,-2,-1,0,+1$, $+2,+3$, for massive spin-2 tensor fields $\lambda=-2,-1,0$, $+1,+2$, whereas for massive vector fields $\lambda=-1,0,+1$, representing the spin orientations in a magnetic field, or respective degrees of freedom of the massive fields. For massless vector fields (photons) there are only two polarizations $\lambda=1,2$.

[^6]:    ${ }^{7}$ For free massive axial- and pseudovector fields, the construction is identical.

