Supersymmetric solutions from N = 6 gauged supergravity

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We study supersymmetric solutions in four-dimensional N = 6 gauged supergravity with the SO(6) gauge group. There is a unique N = 6 supersymmetric AdS₄ vacuum with SO(6) symmetry dual to an N = 6 superconformal field theory (SCFT) in three dimensions. We find a number of domain walls interpolating between this AdS₄ vacuum and singular geometries in the IR with $SO(2) \times SO(4)$, U(3), SO(3), and $SO(2) \times SO(2) \times SO(2)$ symmetries. The SO(3) case admits N = 6 or N = 2 solutions depending on whether the pseudoscalars are present or not. On the other hand, all the remaining solutions preserve N = 6 supersymmetry. These solutions describe renormalization-group (RG) flows from the N = 6 SCFT to nonconformal field theories driven by mass deformations. In particular, the $SO(2) \times$ SO(4) solution is in agreement with the previously known mass deformations of the dual N = 6 SCFT. We also give a supersymmetric Janus solution with $SO(2) \times SO(4)$ symmetry, describing two-dimensional conformal defects in the N = 6 SCFT with unbroken N = (4, 2) supersymmetry. Finally, we find an N = 2supersymmetric AdS₂ × H^2 solution with $SO(2) \times SO(4)$ symmetry and the corresponding domain wall interpolating between this fixed point and the AdS_4 vacuum. The solution describes an AdS_4 black hole with a magnetic charge and is dual to a twisted compactification of the N = 6 SCFT on a hyperbolic space H^2 . We also give a domain wall interpolating between a locally supersymmetric AdS₄ and a curved domain wall with $SO(2) \times SO(2) \times SO(2)$ symmetry dual to an RG flow across dimensions from the N = 6 SCFT to a supersymmetric quantum mechanics.

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I. INTRODUCTION

Supersymmetric solutions of gauged supergravities in various space-time dimensions play an important role in string and M theory. In the AdS/CFT correspondence [1–3], these solutions provide holographic descriptions of strongly coupled systems such as (non)conformal field theories, conformal defects, AdS black holes, and condensed matter physics. In many cases, solutions of lower-dimensional gauged supergravities can be uplifted to ten or 11 dimensions via consistent truncations resulting in complete AdS/CFT dualities in the context of string and M theory.

In this paper, we are interested in supersymmetric solutions of four-dimensional N = 6 gauged supergravity with the SO(6) gauge group; see [4] for all timelike supersymmetric solutions in ungauged N = 6 supergravity.

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This has been constructed in Ref. [5] in the embedding tensor formalism obtained from truncating the maximal N = 8 gauged supergravity [6]; see also [7–9]. The N = 6 gauged supergravity has been shown to admit a unique N = 6 supersymmetric AdS₄ vacuum, with the full SO(6) symmetry unbroken, dual to an N = 6 superconformal field theory (SCFT) in three dimensions. The uniqueness of the N = 6 AdS₄ vacua has also been shown in a recent result on supersymmetric AdS vacua [10]. It has been pointed out in Ref. [5] that this AdS₄ fixed point describes a truncation of type IIA theory on CP^3 , so the AdS₄ vacuum can be uplifted to AdS₄ × CP^3 geometry in type IIA theory. This truncation has been studied long ago in Ref. [11], in which the full mass spectrum has also been given; for more recent studies, see, for example, [12–15].

The very first example of the dual N = 6 SCFT from type IIA theory has been given in Ref. [16]. In general, SCFTs in three dimensions take the form of Chern-Simonsmatter (CSM) theories, since the usual gauge theories with Yang-Mills gauge kinetic terms are not conformal. A number of these SCFTs with different numbers of supersymmetries have already been constructed; see [17–36] for an incomplete list. These SCFTs arise as world-volume theories of M2-branes on various transverse spaces and play an important role in understanding the dynamics of

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M2-branes. Supersymmetric solutions of four-dimensional gauged supergravities are expected to be very useful for their holographic descriptions, at least in the large-*N* limit.

Various types of supersymmetric solutions from gauged supergravities have been considered and given interpretations in terms of the corresponding dual field theories. We will study these solutions in N = 6 gauged supergravity beginning with supersymmetric domain walls interpolating between the N = 6 supersymmetric AdS₄ vacuum and singular geometries. The solutions describe holographic renormalization-group (RG) flows from the dual N = 6SCFT in the UV to nonconformal phases in the IR arising from mass deformations of the UV N = 6 SCFT. Similar solutions have been extensively studied in N = 8 and N = 2 gauged supergravities; see, for example, [37–45]. Solutions in gauged supergravities with N = 3, 4, 5supersymmetries have also been considered recently in Refs. [46–51]. This work could hopefully fill the existing gap by providing a number of supersymmetric solutions in N = 6 gauged supergravity.

We will also find Janus solutions in the form of AdS₃sliced domain walls interpolating between asymptotically AdS₄ spaces. These are holographically dual to twodimensional conformal defects within the N = 6 SCFT and break the superconformal symmetry in the threedimensional bulk to a smaller one on the two-dimensional surfaces. Solutions of this type in other four-dimensional gauged supergravities have previously been studied in Refs. [48,49,51–55]. Finally, we will look for $AdS_2 \times \Sigma^2$ geometries with Σ^2 being a Riemann surface together with solutions interpolating between these backgrounds and the supersymmetric AdS₄ vacuum. The solutions describe supersymmetric black holes in asymptotically AdS₄ space, and a number of these solutions have already been studied in other gauged supergravities in Refs. [51,56–66]. In the dual field theory, the solutions describe RG flows from the N = 6 SCFT to superconformal quantum mechanics in the IR which play a prominent role in microscopic computation of black hole entropy in asymptotically AdS₄ spaces; see, for example, [67–69]. In this context, the superconformal quantum mechanics, or one-dimensional SCFT, is obtained from the N = 6 SCFT via twisted compactifications on Σ^2 .

Four-dimensional N = 6 gauged supergravity has $SO^*(12)$ global symmetry with the maximal compact subgroup $U(6) \sim SU(6) \times U(1)$. There are 30 scalars encoded in the $SO^*(12)/U(6)$ coset manifold. The SO(6) gauging of this supergravity can be obtained from a consistent truncation of the maximal N = 8 gauged supergravity with the SO(8) gauge group. The latter is, in turn, a consistent truncation of 11-dimensional super-gravity on S^7 [70–75]. The N = 6 gauged supergravity with the SO(6) gauge group can accordingly be uplifted to 11 dimensions via a series of consistent truncations. On the other hand, the SO(6), N = 6 gauged supergravity is a

consistent truncation of type IIA theory on \mathbb{CP}^3 . Therefore, all the solutions given here have known higher-dimensional origins and can be embedded in ten- or 11-dimensional supergravities. The scalar potential of the N = 6 gauged supergravity has been analyzed for a long time in Ref. [76]. More recently, this gauged supergravity has been rewritten in a more general setting of the embedding tensor formalism in Ref. [5], in which the fermion-shift matrices and the scalar potential have been given by truncating the N = 8 theory. In this paper, we first complete the task by extending the truncation to all terms in the bosonic Lagrangian and fermionic supersymmetry transformations. Both of these are, of course, a relevant part in the present analysis.

The paper is organized as follows. In Sec. II, we review four-dimensional N = 6 gauged supergravity with SO(6)gauge group in the embedding tensor formalism. In Sec. III, we study supersymmetric domain wall solutions describing RG flows in the dual N = 6 SCFT to nonconformal phases in the IR. We then turn to supersymmetric Janus solutions in Sec. IV and finally look for possible supersymmetric AdS₂ × Σ^2 solutions together with flow solutions interpolating between the AdS₄ vacuum and these geometries in Sec. V. Conclusions and comments are given in Sec. VI.

II. N = 6 GAUGED SUPERGRAVITY WITH SO(6) GAUGE GROUP

We first give a review of N = 6 gauged supergravity in the embedding tensor formalism as described in Ref. [5]. We will follow most of the convention in Ref. [5] but with a mostly plus signature for the space-time metric. The only supermultiplet in N = 6 supersymmetry is the gravity multiplet with the field content

$$(e^{\hat{\mu}}_{\mu},\psi_{\mu A},A^{AB}_{\mu},A^{0}_{\mu},\chi_{ABC},\chi_{A},\phi_{AB}).$$
 (1)

The component fields correspond to the graviton e^{μ}_{μ} , six gravitini $\psi_{\mu A}$, 16 vectors $A^{AB}_{\mu} = -A^{BA}_{\mu}$ and A^{0}_{μ} , and 26 spin-¹/₂ fields $\chi_{ABC} = \chi_{[ABC]}$ and χ_A together with 15 complex scalars $\phi_{AB} = -\phi_{BA}$. Real and imaginary parts of ϕ_{AB} are usually called scalars and pseudoscalars, respectively.

In this work, space-time and tangent space indices are denoted by $\mu, \nu, ... = 0, 1, 2, 3$ and $\hat{\mu}, \hat{\nu}, ... = 0, 1, 2, 3$, respectively. Indices A, B, ... = 1, 2, ..., 6 correspond to the fundamental representation of SU(6) which is, in turn, a subgroup of the R-symmetry $U(6) \sim SU(6) \times U(1)$. The 30 real scalars within ϕ_{AB} are coordinates of the scalar manifold $SO^*(12)/U(6)$ and can be described by the coset representative in representation **32** of $SO^*(12)$ of the form

$$\mathcal{V}_M{}^{\underline{M}} = \mathcal{A}^{\dagger} e^Y \tag{2}$$

with the Cayley matrix

$$\mathcal{A} = \frac{1}{\sqrt{2}} \begin{pmatrix} \mathbb{I}_{16} & i\mathbb{I}_{16} \\ \mathbb{I}_{16} & -i\mathbb{I}_{16} \end{pmatrix}$$
(3)

and

$$Y = \begin{pmatrix} 0 & 0_{1\times15} & 0 & \phi_{CD} \\ 0_{15\times1} & 0_{15\times15} & \phi_{AB} & \frac{1}{2}\epsilon_{ABCDEF}\bar{\phi}^{EF}\phi_{CD} \\ 0 & \bar{\phi}^{CD} & 0 & 0_{1\times15} \\ \bar{\phi}^{AB} & \frac{1}{2}\epsilon^{ABCDEF}\phi_{EF} & 0_{15\times1} & 0_{15\times15} \end{pmatrix}.$$
(4)

We also note that $\bar{\phi}^{AB} = (\phi_{AB})^*$.

In subsequent analysis, it is useful to define 16×16 submatrices of \mathcal{V}_M^M by the following identification:

$$\mathcal{V}_{M}{}^{\underline{M}} = \begin{pmatrix} \bar{h}_{\Lambda}{}^{\underline{\Lambda}} & h_{\Lambda}{}_{\underline{\Lambda}} \\ \bar{f}^{\Lambda}{}^{\underline{\Lambda}} & f^{\Lambda}{}_{\underline{\Lambda}} \end{pmatrix}$$
(5)

in which \mathbf{f} , \mathbf{h} , $\mathbf{\bar{f}}$, and $\mathbf{\bar{h}}$ satisfy the relations

$$(\mathbf{f}\mathbf{f}^{\dagger})^{T} = \mathbf{f}\mathbf{f}^{\dagger}, \quad (\mathbf{h}\mathbf{h}^{\dagger})^{T} = \mathbf{h}\mathbf{h}^{\dagger}, \quad \mathbf{f}\mathbf{h}^{\dagger} - \bar{\mathbf{f}}\mathbf{h}^{T} = i\mathbb{I}_{16},$$

$$\mathbf{f}^{\dagger}\mathbf{h} - \mathbf{h}^{\dagger}\mathbf{f} = -i\mathbb{I}_{16}, \quad \mathbf{f}^{T}\mathbf{h} - \mathbf{h}^{T}\mathbf{f} = 0.$$
(6)

The inverse of $\mathcal{V}_M \stackrel{\underline{M}}{=}$ can accordingly be written in terms of **f** and **h** as

$$\mathcal{V}_{\underline{M}}{}^{\underline{M}} = \begin{pmatrix} -if^{\Lambda}{}_{\underline{\Lambda}} & ih_{\underline{\Lambda}\underline{\Lambda}} \\ i\bar{f}^{\underline{\Lambda}\underline{\Lambda}} & -i\bar{h}_{\underline{\Lambda}}\underline{\Lambda} \end{pmatrix}.$$
(7)

The 16 electric gauge fields A^{AB} and A^0 combine into $A^{\Lambda} = (A^0, A^{AB})$. Together with the magnetic dual A_{Λ} , the gauge fields transform as the **32** representation of $SO^*(12)$:

$$A^M = (A^\Lambda, A_\Lambda). \tag{8}$$

Gaugings are efficiently described by the embedding tensor formalism in which the corresponding gauge generators are defined as

$$X_M = \theta_M{}^n t_n \tag{9}$$

with t^n being the $SO^*(12)$ generators. $\theta_M{}^n$ is called the embedding tensor in terms of which the covariant derivative implementing the minimal coupling of various fields can be written as

$$D_{\mu} = \nabla_{\mu} - g A^M_{\mu} X_M. \tag{10}$$

 ∇_{μ} is the usual space-time covariant derivative including (possibly) the local U(6) composite connection. The parameter g is the gauge coupling constant which can be absorbed in the definition of $\theta_M{}^m$.

In **32** representation, with $SO^*(12)$ generators $(t_n)_M{}^N$, the embedding tensor can be described by the generalized structure constants

$$X_{MN}^{P} = \theta_M^{n} (t_n)_N^{P}.$$
(11)

To define a proper gauging and preserve the full supersymmetry of the ungauged theory, the embedding tensor needs to satisfy the so-called linear and quadratic constraints given, respectively, by

$$X_{(MN}{}^{L}\Omega_{P)L} = 0 \quad \text{and} \quad \theta_{M}{}^{m}\theta_{N}{}^{n}f_{mn}{}^{p} + X_{MN}{}^{P}\theta_{P}{}^{p} = 0$$
(12)

with $f_{mn}{}^{p}$ being the $SO^{*}(12)$ structure constants. The former implies that the embedding tensor $\theta_{M}{}^{m}$ is in the representation **351** of $SO^{*}(12)$, while the latter gives rise to

$$[X_M, X_N] = -X_{MN}{}^P X_P. (13)$$

The gauge generators then form a closed subalgebra for which X_{MN}^{P} act as the corresponding structure constants.

As usual in gauging a supergravity theory, supersymmetry requires some modifications to the ungauged Lagrangian and supersymmetry transformations. These modifications are of first and second order in the gauge coupling constant and can be written in terms of the socalled T tensor:

$$T_{\underline{M}\underline{N}}{}^{\underline{P}} = \mathcal{V}_{\underline{M}}{}^{\underline{M}}\mathcal{V}_{\underline{N}}{}^{\underline{N}}\mathcal{V}_{\underline{P}}{}^{\underline{P}}X_{\underline{M}\underline{N}}{}^{\underline{P}}.$$
 (14)

The bosonic Lagrangian of the N = 6 gauged supergravity can be written as

$$e^{-1}\mathcal{L} = \frac{1}{2}R - \frac{1}{24}P_{\mu ABCD}P^{\mu ABCD} - \frac{i}{4}(\mathcal{N}_{\Lambda\Sigma}F^{+\Lambda}_{\mu\nu}F^{+\Sigma\mu\nu} - \bar{\mathcal{N}}_{\Lambda\Sigma}F^{-\Lambda}_{\mu\nu}F^{-\Sigma\mu\nu}) - V. \quad (15)$$

The scalar kinetic term is given in terms of the vielbein $P_{\mu}^{ABCD} = (P_{\mu ABCD})^*$ on the $SO^*(12)/U(6)$ coset which is defined by

$$P^{ABCD}_{\mu} = \mathcal{V}^{ABM} D_{\mu} \mathcal{V}_{M}^{CD}$$

= $i(\bar{f}^{AAB} D_{\mu} \bar{h}_{\Lambda}^{CD} - \bar{h}_{\Lambda}^{AB} D_{\mu} \bar{f}^{\Lambda CD}).$ (16)

The scalar matrix appearing in the gauge kinetic terms is given by

$$\mathcal{N}_{\Lambda\Sigma} = -\bar{h}_{\Lambda}{}^{\underline{\Lambda}}(f^{-1})_{\underline{\Lambda}\Sigma} \tag{17}$$

with $\mathcal{N}_{\Lambda\Sigma}$ being its complex conjugate. The complex self-dual and anti-self-dual gauge field strengths are defined by

$$F_{\mu\nu}^{\pm\Lambda} = \frac{1}{2} \left(F_{\mu\nu}^{\Lambda} \pm \frac{i}{2} \epsilon_{\mu\nu\rho\sigma} F^{\Lambda\rho\sigma} \right)$$
(18)

with $F^{\Lambda}_{\mu\nu}$ given by

$$F^{\Lambda}_{\mu\nu} = \partial_{\mu}A^{\Lambda}_{\nu} - \partial_{\nu}A^{\Lambda}_{\mu} + X_{\Gamma\Sigma}{}^{\Lambda}A^{\Gamma}_{\mu}A^{\Sigma}_{\nu}.$$
 (19)

The scalar potential is obtained from the fermion-shift matrices as follows:

$$V = -2S_{AB}S^{AB} + \frac{1}{36}N_A{}^{BCD}N^A{}_{BCD} + \frac{1}{6}N_A{}^BN^A{}_B.$$
 (20)

We note that upper and lower SU(6) indices are related by complex conjugation. In terms of the various components of the *T* tensor with the splitting of indices $\underline{\Lambda}, \underline{\Sigma}, \ldots$ as (0, [AB]), we have

$$S_{AB} = \frac{\sqrt{2}}{5} T_{C(A,B)E}{}^{CE}, \qquad N_{AB} = -\frac{8\sqrt{2}}{3} T_{C[A,B]E}{}^{CE},$$
$$N^{A}{}_{B} = -2\sqrt{2}T_{0,BC}{}^{AC},$$
$$N^{A}{}_{BCD} = -2\sqrt{2}T_{[CD,B]E}{}^{AE} - \frac{1}{4}\delta^{A}_{[B}N_{CD]}.$$
(21)

The fermionic supersymmetry transformations, with all fermionic fields vanishing, are given by

$$\delta\psi_{\mu A} = D_{\mu}\epsilon_{A} - S_{AB}\gamma_{\mu}\epsilon^{B} - \frac{1}{4\sqrt{2}}\hat{F}^{+}_{\rho\sigma AB}\gamma^{\rho\sigma}\gamma_{\mu}\epsilon^{B}, \quad (22)$$

$$\delta\chi_A = -\frac{1}{4!} \epsilon_{ABCDEF} P^{BCDE}_{\mu} \gamma^{\mu} \epsilon^F + N^B_{\ A} \epsilon_B - \frac{1}{2\sqrt{2}} \hat{F}^+_{\mu\nu} \gamma^{\mu\nu} \epsilon_A,$$
(23)

$$\delta\chi_{ABC} = -P_{\mu ABCD}\gamma^{\mu}\epsilon^{D} + N^{D}{}_{ABC}\epsilon_{D} - \frac{3}{2\sqrt{2}}\hat{F}^{+}_{\mu\nu[AB}\epsilon_{C]}.$$
(24)

We note here the chiralities of the fermionic fields

$$\gamma_5 \psi_{\mu A} = -\psi_{\mu A}, \quad \gamma_5 \chi_{ABC} = -\chi_{ABC}, \quad \gamma_5 \chi_A = -\chi_A \tag{25}$$

with ψ^A_{μ} , χ^{ABC} , and χ^A having opposite chiralities. The tensors $\hat{F}^+_{\mu\nu AB} = (\hat{F}^{-AB}_{\mu\nu})^*$ can be obtained from

$$\hat{F}^{-AB}_{\mu\nu} = \mathcal{V}_M{}^{AB}G^{-M}_{\mu\nu} \tag{26}$$

with

$$G^{M}_{\mu\nu} = \begin{pmatrix} F^{\Lambda}_{\mu\nu} \\ G_{\Lambda\mu\nu} \end{pmatrix}$$
(27)

and $G_{\Lambda\mu\nu} = i\epsilon_{\mu\nu\rho\sigma} \frac{\partial \mathcal{L}}{\partial F^{\Lambda}_{\rho\sigma}}$. Similarly, we have $\hat{F}^{+}_{\mu\nu} = (\mathcal{V}_{M}{}^{0}G^{-M}_{\mu\nu})^{*}$.

The covariant derivative of ϵ_A is defined by

$$D_{\mu}\epsilon_{A} = \partial_{\mu}\epsilon_{A} + \frac{1}{4}\omega_{\mu}{}^{ab}\gamma_{ab}\epsilon_{A} + \frac{1}{2}Q_{\mu A}{}^{B}\epsilon_{B}.$$
 (28)

The connection $Q_{\mu A}{}^B$ is given by

$$Q_{\mu A}{}^{B} = \frac{2i}{3} (h_{\Lambda AC} \partial_{\mu} \bar{f}^{\Lambda AB} - f^{\Lambda}{}_{AC} \partial_{\mu} \bar{h}_{\Lambda}{}^{BC}) - g A^{M}_{\mu} Q_{MA}{}^{B}$$
(29)

with Q_{MA}^{B} obtained from

$$Q_{MAB}{}^{CD} = \mathcal{V}_{AB}{}^{P}X_{MP}{}^{N}\mathcal{V}_{N}{}^{CD}$$
(30)

by the relation $Q_{MAB}^{CD} = 4\delta^{[C}_{[A}Q_{MB]}^{D]}$.

In general, both electric and magnetic gauge fields can participate in the gaugings, leading to many possibilities of viable gauge groups. However, in this work, we will consider only the SO(6) gauge group embedded electrically in $U(6) \subset SO^*(12)$. This gauging involves only electric gauge fields A^{AB} . In this case, we have

$$X_{I_1J_1,I_2J_2}{}^{I_3J_3} = 4g\delta^{[I_3}_{[I_1}\delta_{I_2][J_2}\delta^{J_3]}_{J_2]}$$

and $X_{I_1J_1}{}^{I_3J_3}{}_{I_2J_2} = -X_{I_1J_1,I_2J_2}{}^{I_3J_3}$ (31)

with all remaining components vanishing. In particular, there are no $X^{\Lambda}{}_{M}{}^{N}$ components which couple to magnetic gauge fields.

With the splitting of indices $\Lambda, \Sigma, ... = (0, [IJ])$, we find from the definition (14) that

$$T_{AB}{}^{CD} = -\frac{1}{2} f^{I_1 J}{}_0 (f^{JJ_1}{}_{AB} \bar{h}_{I_1 J_1}{}^{CD} + h_{I_1 J_1, AB} \bar{f}^{JJ_1, CD}), \quad (32)$$

$$T_{EF, AB}{}^{CD} = -\frac{1}{2} f^{I_1 J}{}_{EF} (f^{JJ_1}{}_{AB} \bar{h}_{I_1 J_1}{}^{CD} + h_{I_1 J_1, AB} \bar{f}^{JJ_1, CD}). \quad (33)$$

From these, it is straightforward to obtain all the fermionshift matrices and the scalar potential.

In subsequent sections, we will look for various types of supersymmetric solutions to this N = 6 gauged supergravity with the SO(6) gauge group. It has been shown in Ref. [5] that this gauged supergravity admits a supersymmetric N = 6 AdS₄ vacuum with the cosmological constant $V_0 = -48g^2$ when all scalars vanish. According to the AdS/CFT correspondence, this is dual to an N = 6SCFT in three dimensions. We will find solutions that are asymptotic to this AdS₄ geometry and can be interpreted as different types of deformations of the dual N = 6 SCFT.

III. HOLOGRAPHIC RG FLOWS

We first consider holographic RG flow solutions in the form of domain walls interpolating between the supersymmetric AdS_4 vacuum and another AdS_4 vacuum (if it exists) or a singular geometry. These solutions correspond, respectively, to RG flows of the dual UV N = 6 SCFT to another conformal fixed point or to a nonconformal phase in the IR.

The metric ansatz is taken to be

$$ds^2 = e^{2A(r)} dx_{1,2}^2 + dr^2 \tag{34}$$

with $dx_{1,2}^2$ being the flat metric on three-dimensional Minkowski space. Scalar fields are allowed to depend only on the radial coordinate *r* with all the other fields set to zero.

A. Solutions with $SO(2) \times SO(4)$ symmetry

We first consider a simple case of solutions with $SO(2) \times SO(4)$ symmetry. The embedding of SO(6) implies that the scalars ϕ_{AB} transform as an adjoint representation of SO(6). There is one singlet scalar under $SO(2) \times SO(4) \subset SO(6)$ given explicitly by

$$\phi_{AB} = \phi(\delta_A^1 \delta_B^2 - \delta_B^1 \delta_A^2). \tag{35}$$

We will also write

$$\phi = \varphi e^{i\zeta} \tag{36}$$

for real scalars φ and ζ depending only on r.

By a straightforward computation, we find the tensor S_{AB} of the form

$$S_{AB} = 2g\cosh\varphi\delta_{AB} = \frac{1}{2}\mathcal{W}\delta_{AB}.$$
 (37)

We have introduced the "superpotential" W for convenience. In general, the function W is related to the eigenvalue of S_{AB} corresponding to the unbroken supersymmetry. In the present case, S_{AB} is proportional to the identity matrix, indicating that the solutions will preserve either N = 6 supersymmetry with all ϵ_A nonvanishing or no supersymmetry at all. Note also that W has a critical point at $\varphi = 0$ which is the supersymmetric N = 6AdS₄ vacuum mentioned above.

To solve all the Bogomol'nyi-Prasad-Sommerfield (BPS) conditions, we will, as in other previous works, impose the following projector:

$$\gamma_{\hat{r}}\epsilon_A = e^{i\Lambda}\epsilon^A \tag{38}$$

for a real function Λ . Throughout the paper, we will use Majorana representation for gamma matrices in which all γ^{μ} are real but γ_5 is purely imaginary. This implies that

 ϵ_A and ϵ^A are related by complex conjugation. Note also that the projector (38) relates the two chiralities of ϵ^A , so the full flow solutions will preserve only half of the original supersymmetry or 12 supercharges.

Considering the conditions $\delta \psi_{\mu A} = 0$ for $\mu = 0, 1, 2,$ we find

$$e^{i\Lambda}A' - \mathcal{W} = 0 \tag{39}$$

with the prime denoting r derivatives. This equation gives

$$A' = \pm |\mathcal{W}| \quad \text{and} \quad e^{i\Lambda} = \pm \frac{\mathcal{W}}{|\mathcal{W}|}.$$
 (40)

In what follows, we will write W = |W| for convenience. We will also choose the upper signs in order to make the supersymmetric AdS_4 critical point correspond to $r \to \infty$. Since, in this case, the superpotential is real, we simply have

$$A' = 4g \cosh \varphi \quad \text{and} \quad e^{i\Lambda} = 1.$$
 (41)

The condition $\delta \psi_{rA} = 0$ gives the standard Killing spinors of the domain walls

$$\epsilon_A = e^{A/2} \epsilon_{A(0)} \tag{42}$$

for spinors $\epsilon_{A(0)}$ satisfying Eq. (38).

Using the projection (38) in the variations $\delta \chi_{ABC}$ and $\delta \chi_A$ gives the following BPS equations:

$$\varphi' = -4g \sinh \varphi \quad \text{and} \quad \zeta' = 0.$$
 (43)

We have now obtained the BPS equations that solve all the supersymmetry conditions. It can also be readily verified that these equations imply the second-order field equations.

We can analytically solve the above BPS equations with the following solution:

$$4gr = \ln(1 + e^{\varphi}) - \ln(1 - e^{\varphi}), \tag{44}$$

$$A = \varphi - \ln(1 - e^{2\varphi}). \tag{45}$$

We have neglected the integration constants in these equations, since they can be removed by shifting the radial coordinate and scaling the $x^{0,1,2}$ coordinates, respectively. As $r \to \infty$, we find that

$$\varphi \sim e^{-4gr} \sim e^{-r/L}$$
 and $A \sim 4gr \sim \frac{r}{L}$ (46)

with *L* being the AdS_4 radius related to the cosmological constant by

$$L = \sqrt{-\frac{3}{V_0}} = \frac{1}{4g}.$$
 (47)

We have also taken g > 0 for convenience.

The behavior of φ implies that φ is dual to a relevant operator of dimensions $\Delta = 1$, 2 in the dual SCFT. In addition, the solution is singular at $r \rightarrow 0$ with

$$\varphi \sim \pm \ln(4gr)$$
 and $A \sim \ln(4gr)$. (48)

We then find that $\varphi \to \pm \infty$ near the singularity. From the explicit form of the scalar potential, we have

$$V \sim -8g^2 e^{\pm 2\varphi} \to -\infty. \tag{49}$$

By the criterion given in Ref. [77], we conclude that the singularity is physical. Therefore, the solution describes an RG flow from the UV N = 6 SCFT to a nonconformal phase in the IR. The flow is driven by an operator of dimensions $\Delta = 1$, 2 corresponding to scalar or fermion mass terms in three dimensions. The flow breaks superconformal symmetry but preserves the full N = 6 Poincaré supersymmetry. Moreover, the R symmetry SO(6) is broken to $SO(2) \times SO(4)$ subgroup. This is precisely in agreement with the field theory result given in Ref. [25]. We then expect the solution to describe mass deformations of the three-dimensional N = 6 SCFT.

B. Solutions with U(3) symmetry

We now consider another residual symmetry, namely, $U(3) \sim SU(3) \times U(1) \subset SO(6)$. The U(3) generators in the fundamental representation of SO(6) can be written as

$$X = \begin{pmatrix} A_{3\times3} & S_{3\times3} \\ -S_{3\times3} & A_{3\times3} \end{pmatrix}$$
(50)

in which $A_{3\times3}$ and $S_{3\times3}$ are antisymmetric and symmetric 3×3 matrices, respectively. The matrices $A_{3\times3}$ generate an $SO(3) \subset SU(3)$ which is a diagonal subgroup of $SO(3) \times SO(3) \subset SO(6)$. The U(1) factor corresponds to $S_{3\times3} = \mathbb{I}_3$. There is only one U(3) singlet scalar given by

$$\phi_{AB} = \begin{pmatrix} 0_{3\times3} & \phi \mathbb{I}_3 \\ -\phi \mathbb{I}_3 & 0_{3\times3} \end{pmatrix} = \phi J_{AB}.$$
 (51)

The matrix J_{AB} is identified with the Kähler form of CP^3 on which the ten-dimensional type IIA theory compactifies [5].

By writing $\phi = \varphi e^{i\zeta}$ and repeating the same analysis as in the previous case, we find the scalar potential

$$V = -24g^2 e^{-2\varphi} (1 + e^{4\varphi}), \tag{52}$$

which is exactly the same as that given in Ref. [5]. As in the $SO(2) \times SO(4)$ case, this potential admits an N = 6 AdS₄

critical point at $\varphi = 0$ dual to an N = 6 SCFT in three dimensions.

The matrix S_{AB} is proportional to the identity

$$S_{AB} = \frac{1}{2} \mathcal{W} \delta_{AB} \tag{53}$$

with a complex superpotential

$$\mathcal{W} = \frac{1}{2}e^{-3\varphi - i\zeta} [(e^{6\varphi} + 3e^{2\varphi})(1 + e^{i\zeta}) + (1 + e^{4\varphi})(e^{i\zeta} - 1)].$$
(54)

The variations $\delta \chi_A$ and $\delta \chi_{ABC}$ lead to

$$e^{-i\Lambda}(2\varphi' \pm i\sinh(2\varphi)\zeta') = -ge^{-3\varphi}(e^{4\varphi} - 1)[1 - e^{i\zeta} + e^{2\varphi}(1 + e^{i\zeta})], \quad (55)$$

which implies $\zeta' = 0$. It turns out that $\zeta' = 0$ is also required by the field equations. For constant $\zeta = \zeta_0$, we have verified that all the resulting BPS equations are compactible with the field equations. In the following analysis, we will set $\zeta_0 = 0$ and end up with the BPS equations

$$\varphi' = -e^{-\varphi}(e^{4\varphi} - 1)$$
 and $A' = ge^{-\varphi}(3 + e^{4\varphi}).$ (56)

The solution can be readily obtained:

$$A = 3\varphi - \ln(1 - e^{4\varphi}),$$
 (57)

$$4gr = 2\tan^{-1}e^{\varphi} - \ln(1 - e^{\varphi}) + \ln(1 + e^{\varphi}).$$
 (58)

As in the previous case, the solution is asymptotic to the supersymmetric AdS_4 with φ dual to an operator of dimensions $\Delta = 1, 2$, while at r = 0 the solution is singular with

$$\varphi \sim \ln(gr)$$
 and $A \sim 3\varphi \sim 3\ln(gr)$ (59)

and

$$\varphi \sim -\ln(gr)$$
 and $A \sim -\varphi \sim \ln(gr)$. (60)

Both of these give

$$V \sim -24g^2 e^{\pm 2\varphi} \to -\infty,\tag{61}$$

so the two singularities are physical. We can accordingly interpret the solution as a holographic dual of RG flows from the N = 6 SCFT to nonconformal phases in the IR. The flow preserves N = 6 Poincaré supersymmetry in three dimensions as in the $SO(2) \times SO(4)$ case but breaks the SO(6) R symmetry to U(3). It would be interesting to identify the corresponding mass deformations in the dual N = 6 SCFT similar to the $SO(2) \times SO(4)$ case.

C. Solutions with $SO(2) \times SO(2) \times SO(2)$ symmetry

To obtain more interesting and more complicated solutions, we consider solutions with a smaller symmetry, namely, $SO(2) \times SO(2) \times SO(2) \subset SO(6)$ symmetry. There are three complex scalars which are singlets under this $SO(2) \times SO(2) \times SO(2)$. The explicit parametrization of these singlets can be written as

$$\phi_{AB} = \begin{pmatrix} \phi_1 i \sigma_2 & 0_{2\times 2} & 0_{2\times 2} \\ 0_{2\times 2} & \phi_2 i \sigma_2 & 0_{2\times 2} \\ 0_{2\times 2} & 0_{2\times 2} & \phi_3 i \sigma_2 \end{pmatrix}.$$
 (62)

By setting

$$\phi_{\alpha} = \varphi_{\alpha} e^{i\zeta_{\alpha}}, \qquad \alpha = 1, 2, 3, \tag{63}$$

we find the scalar potential

$$V = -16g^{2}[\cosh(2\varphi_{1}) + \cosh(2\varphi_{2}) + \cosh(2\varphi_{3})].$$
(64)

It is clearly seen that this potential admits only one critical point at $\varphi_1 = \varphi_2 = \varphi_3 = 0$ which is the aforementioned $N = 6 \text{ AdS}_4$ vacuum.

The matrix S_{AB} is given by

$$S_{AB} = \frac{1}{2} \begin{pmatrix} \mathcal{W}_1 \mathbb{I}_2 & 0_{2\times 2} & 0_{2\times 2} \\ 0_{2\times 2} & \mathcal{W}_2 \mathbb{I}_2 & 0_{2\times 2} \\ 0_{2\times 2} & 0_{2\times 2} & \mathcal{W}_3 \mathbb{I}_2 \end{pmatrix}$$
(65)

with

$$\mathcal{W}_{1} = \frac{1}{2}ge^{-\varphi_{1}-\varphi_{2}-\varphi_{3}} \times [e^{i(\zeta_{1}-\zeta_{2}-\zeta_{3})}(e^{2\varphi_{1}}-1)(e^{2\varphi_{2}}-1)(e^{2\varphi_{3}}-1) - (1+e^{2\varphi_{1}})(1+e^{2\varphi_{2}})(1+e^{2\varphi_{3}})].$$
(66)

 W_2 and W_3 take a similar form with the phase $e^{i(\zeta_1-\zeta_2-\zeta_3)}$ replaced by $e^{i(\zeta_2-\zeta_1-\zeta_3)}$ and $e^{i(\zeta_3-\zeta_1-\zeta_2)}$, respectively.

It turns out that none of these W_{α} gives rise to the superpotential in term of which the scalar potential (64) can be written unless $\zeta_1 = \zeta_2 = \zeta_3 = 0$. This is also implied by the consistency between the resulting BPS equations and the field equations. We now set $\zeta_1 = \zeta_2 = \zeta_3 = 0$, resulting in $e^{i\Lambda} = \pm 1$, and obtain the following BPS equations:

$$\varphi_1' = -ge^{-\varphi_1 - \varphi_2 - \varphi_3} [e^{2(\varphi_1 + \varphi_2)} + e^{2(\varphi_1 + \varphi_3)} - e^{2(\varphi_2 + \varphi_3)} - 1],$$
(67)

$$\varphi_2' = -ge^{-\varphi_1 - \varphi_2 - \varphi_3} [e^{2(\varphi_1 + \varphi_2)} + e^{2(\varphi_2 + \varphi_3)} - e^{2(\varphi_1 + \varphi_3)} - 1],$$
(68)

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$$\varphi_{3}' = -ge^{-\varphi_{1}-\varphi_{2}-\varphi_{3}}[e^{2(\varphi_{1}+\varphi_{3})} + e^{2(\varphi_{2}+\varphi_{3})} - e^{2(\varphi_{1}+\varphi_{2})} - 1],$$
(69)

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$$A' = ge^{-\varphi_1 - \varphi_2 - \varphi_3} [e^{2(\varphi_1 + \varphi_2)} + e^{2(\varphi_1 + \varphi_3)} + e^{2(\varphi_3 + \varphi_3)} + 1].$$
(70)

To find the solution to these equations, we first take a linear combination:

$$\varphi_1' + \varphi_2' = -2ge^{-\varphi_1 - \varphi_2 - \varphi_3}(e^{2(\varphi_1 + \varphi_2)} - 1).$$
(71)

After changing to a new radial coordinate ρ defined by

$$\frac{d\rho}{dr} = e^{-\varphi_1 - \varphi_2 - \varphi_3},\tag{72}$$

we find

$$\varphi_2 = 2g\rho - \varphi_1 - \frac{1}{2}\ln(e^{4g\rho} + C_2) \tag{73}$$

for a constant C_2 .

Similarly, taking a linear combination $\varphi'_1 + \varphi'_3$ gives rise to

$$\varphi_3 = 2g\rho - \varphi_1 - \frac{1}{2}\ln(e^{4g\rho} + C_3).$$
 (74)

Using these results in Eq. (67), we find

$$\varphi_1 = \frac{1}{4} \ln \left[\frac{e^{4g\rho} (e^{4g\rho} + C_1)}{(e^{4g\rho} + C_2)(e^{4g\rho} + C_3)} \right].$$
(75)

Finally, with all these results, the solution for A is given by

$$A = g\rho + \frac{1}{4}\ln(e^{4g\rho} + C_1) + \frac{1}{4}\ln(e^{4g\rho} + C_2) + \frac{1}{4}\ln(e^{4g\rho} + C_3).$$
(76)

We now look at the behavior of the solution as $\varphi_{\alpha} \sim 0$ which gives $\rho \sim r$ and

$$\varphi_1 \sim \frac{1}{4} (C_1 - C_2 - C_3) e^{-4g\rho}, \quad \varphi_{2,3} \sim -\frac{1}{4} (C_1 - C_{3,2}) e^{-4g\rho},$$

$$A \sim 4g\rho. \tag{77}$$

This is the expected behavior of the solution asymptotic to the supersymmetric AdS_4 vacuum. As in the previous cases, the solution is singular as $4g\rho \rightarrow \ln(-C_{\alpha})$. For $C_1 \neq C_2 \neq C_3$, there are three possibilities. when $4g\rho \rightarrow \ln(-C_1)$ with

$$\varphi_1 \sim \frac{1}{4} \ln(4g\rho - \tilde{C}_1), \qquad \tilde{C}_1 = \ln(-C_1),$$

 $\varphi_{2,3} \sim -\varphi_1, \qquad A \sim \varphi_1.$
(78)

(ii) $C_2 < C_{1,3}$ or $C_3 < C_{1,2}$.—In this case, we find

$$\varphi_1 \sim -\frac{1}{4} \ln(4g\rho - \tilde{C}_{2,3}), \qquad \tilde{C}_{2,3} = \ln(-C_{2,3}),$$

 $\varphi_{2,3} \sim \varphi_1, \qquad A \sim -\varphi_1.$
(79)

In the first case, we have $\varphi_1 \to -\infty$ and $\varphi_{2,3} \to \infty$, while in the second case, the solution gives $\varphi_{1,2,3} \to \infty$. It can be easily verified that all of these behaviors lead to $V \to -\infty$, so the singularities are physically acceptable. The solution then describes different types of mass deformations within the dual N = 6 SCFT to nonconformal phases with $SO(2) \times SO(2) \times SO(2)$ symmetry. The solution also preserves N = 6 Poincaré supersymmetry as in the previous cases.

D. Solutions with SO(3) symmetry

We further reduce the residual symmetry to $SO(3) \subset$ $SO(3) \times SO(3) \subset SO(6)$ generated by the antisymmetric matrices $A_{3\times 3}$ in the upper-left block of Eq. (50). There are three singlet scalars parametrized by

$$\phi_{AB} = \begin{pmatrix} 0_{3\times3} & 0_{3\times3} \\ 0_{3\times3} & \hat{A}_{3\times3} \end{pmatrix}$$
(80)

with

$$\hat{A} = \begin{pmatrix} 0 & \tilde{\phi}_1 & \tilde{\phi}_2 \\ -\tilde{\phi}_1 & 0 & \tilde{\phi}_3 \\ -\tilde{\phi}_2 & -\tilde{\phi}_3 & 0 \end{pmatrix}.$$
 (81)

By writing $\tilde{\phi}_{\alpha} = \varphi_{\alpha} e^{i\zeta_{\alpha}}$ with

$$\varphi_1 = \Phi \cos \theta, \qquad \varphi_2 = \Phi \sin \theta \cos \vartheta, \qquad \varphi_3 = \Phi \sin \theta \sin \vartheta$$
(82)

and

$$\zeta_1 = \zeta, \qquad \zeta_2 = \zeta + \eta, \qquad \zeta_3 = \zeta + \xi, \qquad (83)$$

we find the scalar potential

$$V = -g^{2}[16\cos^{4}\theta(2 + \cosh 2\Phi) + 16\cosh^{4}\theta\sin^{4}\theta(2 + \cosh 2\Phi) + 16\sin^{4}\theta\sin^{4}\theta(2 + \cosh 2\Phi) - \cos^{2}\theta\sin^{2}\theta\cos^{2}\theta \times (\cosh 4\Phi - 8\cos 2\eta\sinh^{4}\Phi - 36\cosh 2\Phi - 61) + \sin^{2}\theta\sin^{2}\theta \times [8\sinh^{4}\Phi(\cos^{2}\theta\cos 2\xi + \cos^{2}\theta\sin^{2}\theta\cos[2(\eta - \xi)]) + (61 + 36\cosh 2\Phi - \cosh 4\Phi)(\cos^{2}\theta + \cos^{2}\theta\sin^{2}\theta)]].$$

$$(84)$$

In this case, the scalar potential depends on phases of the complex scalars, and the analysis is more complicated. To make the analysis more traceable, we will further truncate to two scalars by setting $\vartheta = 0$ and $\xi = -\zeta$. This is equivalent to setting $\tilde{\phi}_3 = 0$. We now begin with the eigenvalues of the matrix S_{AB} . These are of the form, after diagonalization,

$$S_{AB}^{\text{diag}} = \text{diag}\left(-2g\cosh\Phi_{\times 4}, \frac{1}{2}\mathcal{W}_{+}, \frac{1}{2}\mathcal{W}_{-}\right)$$
(85)

with \mathcal{W}_{\pm} given by

$$\mathcal{W}_{\pm} = 2g(\cos 2\eta + 2\sin \eta)\sinh^4\frac{\Phi}{2}(\cos 4\theta\sin\eta \pm i\sin 2\theta) - \frac{1}{4}g(3 + 12\cosh\Phi + \cosh 2\Phi). \tag{86}$$

The corresponding eigenvectors are

$$\hat{\epsilon}_{\pm} = -\frac{1}{2}\sec 2\theta \Big(2\cos\eta\sin 2\theta \mp \sqrt{3 + \cos 2\eta + 2\cos 4\theta\sin^2\eta} \Big) \epsilon_5 + \epsilon_6.$$
(87)

The scalar kinetic term is given by

$$\mathcal{L}_{\rm kin} = -\frac{1}{2}G_{\alpha\beta}\phi^{\alpha\prime}\phi^{\beta\prime}$$
$$= -\Phi^{\prime 2} - \sinh^2\Phi\theta^{\prime 2} - \frac{1}{4}\sinh^22\Phi\zeta^{\prime 2} - \frac{1}{2}\sin^2\theta\sinh^22\Phi\zeta^{\prime}\eta^{\prime} - \frac{1}{4}\sin^2\theta\sinh^2\Phi(3 + \cosh 2\Phi - 2\cos 2\theta\sinh^2\Phi)\eta^{\prime 2}$$
(88)

with $\phi^{\alpha} = (\Phi, \theta, \zeta, \eta)$. It is useful to give an explicit form of the inverse of $G_{\alpha\beta}$ here:

$$G^{\alpha\beta} = -\frac{1}{2} \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \operatorname{csch}^2 \Phi & 0 & 0 \\ 0 & 0 & -\operatorname{sech}^2 \Phi + \operatorname{csch}^2 \Phi \operatorname{sec}^2 \theta & -\operatorname{csch}^2 \Phi \operatorname{sec}^2 \theta \\ 0 & 0 & -\operatorname{csch}^2 \Phi \operatorname{sec}^2 \theta & 4\operatorname{csc}^2 2\theta \operatorname{csch} \Phi \end{pmatrix}.$$
 (89)

The scalar potential can be written in term of the real superpotential $W = |W_+| = |W_-|$ as

$$V = -2G^{\alpha\beta} \frac{\partial W}{\partial \phi^{\alpha}} \frac{\partial W}{\partial \phi^{\beta}} - 3W^{2}$$

= $g^{2} [\cos^{2}\theta \sin^{2}\theta (\cosh 4\Phi - 8\cos 2\eta \sinh^{4}\Phi - 36\cosh 2\Phi - 61) - 4(3 + \cos 4\theta)(2 + \cosh 2\Phi)].$ (90)

After setting $\epsilon^{1,2,3,4} = 0$ and imposing the projection conditions

$$\gamma_{\hat{r}}\epsilon_{\pm} = e^{\pm i\Lambda}\epsilon^{\pm} \quad \text{with} \quad e^{\pm i\Lambda} = \frac{\mathcal{W}_{\pm}}{W},$$
(91)

we find the following BPS equations:

$$\Phi' = \frac{1}{16W} g^2 [8\sinh^3\Phi\cosh\Phi(\cos 2\eta + 2\cos 4\theta\sin^3\eta) - 30\sin 2\Phi - \sinh 4\Phi],$$
(92)

$$\theta' = -\frac{1}{W}g^2 \sin^2 \eta \sin 4\theta \sinh^2 \Phi, \tag{93}$$

$$\zeta' = \frac{2}{W}g^2 \sin 2\eta \sin^2 \theta \sinh^2 \Phi, \qquad (94)$$

$$\eta' = -\frac{2}{W}g^2 \sin 2\eta \sinh^2 \Phi, \qquad (95)$$

$$A' = W. \tag{96}$$

The flow equations for the scalars can be written in a compact form as

$$\phi^{\alpha\prime} = 2G^{\alpha\beta} \frac{\partial W}{\partial \phi^{\beta}}.$$
(97)

It can straightforwardly be verified that all these equations satisfy the second-order field equations. We also see that, from these equations, there is only one supersymmetric critical point, with $\Phi' = \theta' = \zeta' = \eta' = 0$, at $\phi^{\alpha} = 0$.

We note that, although the superpotential and the scalar potential do not depend on ζ , ζ' is still nonvanishing due to the mixed terms between ζ and η in the matrix $G_{\alpha\beta}$. Moreover, further truncations such as $\eta = 0$ or $\theta = 0$ will lead to the BPS equations in the case of N = 6 supersymmetry with all six eigenvalues of S_{AB} leading to

$$\mathcal{W} = 4g\cosh\Phi. \tag{98}$$

This is very similar to the N = 5 gauged supergravity studied in Ref. [51] in which the differences in the phases of the scalars are crucial for breaking the original supersymmetry to a lower amount.

We now look at the solution to the above equations. Combining η' and θ' equations gives

$$\frac{d\theta}{d\eta} = \frac{1}{4}\sin 4\theta \tan \eta \tag{99}$$

with the solution given by

$$\cot 2\theta = C_1 \cos \eta. \tag{100}$$

Similarly, taking the combination between ζ' and η' equations leads to

$$\frac{d\zeta}{d\eta} = -\sin^2\theta. \tag{101}$$

After using the above solution for θ , we find the solution

$$\zeta = \zeta_0 - \frac{\eta}{2} + \frac{\sqrt{C_1^2 + \sec^2 \eta} \cos \eta \tan^1 \frac{\sqrt{2}C_1 \sin \eta}{\sqrt{2 + C_1^2 (1 + \cos 2\eta)}}}{\sqrt{4 + 2C_1^2 (1 + \cos 2\eta)}}$$
(102)

for constant ζ_0 .

Combining Φ' and η' equations by taking into account all the previous results together with a redefinition

$$\tilde{\Phi} = \sinh \Phi, \tag{103}$$

we find

$$\frac{d\tilde{\Phi}}{d\eta} = \csc 2\eta (1 + \tilde{\Phi}^2) \left(\frac{\tilde{\Phi} \tan^2 \eta}{C_1^2 + \sec^2 \eta} + \frac{2}{\tilde{\Phi}} \right),\tag{104}$$

whose solution is given by

$$\frac{\tilde{\Phi}^2}{4} = -\frac{1 + C_1^2 \cos^2 \eta - C_2 \sqrt{(1 + \cos 2\eta)(2 + C_1^2(1 + \cos 2\eta))}}{3 + 4C_1^2 \cos^2 \eta + \cos 2\eta - 4C_2 \sqrt{(1 + \cos 2\eta)(2 + C_1^2(1 + \cos 2\eta))}}.$$
(105)

Using all these results in the η' equation, we find the solution for $\eta(r)$ implicitly from

$$8gr = \sinh^{-1}\left[2C_2\sqrt{\frac{\Xi-1}{(1+C_1^2)[2+(C_1^2-4C_2^2)(1+\Xi)]}}\right] - \tanh^{-1}\sqrt{\frac{(1+C_1^2-4C_2^2)(1+\Xi)}{\Xi-1}},$$
(106)

in which we have defined

$$\Xi = \cos 2\eta. \tag{107}$$

Finally, we can find the solution for $A(\Xi)$ as

$$A = \frac{1}{4} (\tanh^{-1} \alpha_{+} - \tanh^{-1} \alpha_{-}) - \frac{1}{2} \tanh^{-1} \left[2C_{2} \sqrt{\frac{\Xi + 1}{2 + C_{1}^{2}(1 + \Xi)}} \right] - \frac{1}{8} \ln \left[4C_{1}^{4}(1 + \Xi)^{2} + (3 + \Xi)^{2} - 4(1 + \Xi) \left[8C_{2}^{2} + C_{1}^{2}(4C_{2}^{2}(1 + \Xi) - 3 - \Xi) \right] \right] + \frac{1}{4} \ln \left[2 + (1 + \Xi)(C_{1}^{2} - 4C_{2}^{2}) \right]$$
(108)

with α_{\pm} defined by

$$\alpha_{\pm} = \sqrt{-\frac{2 + C_1^2 (1 + \Xi)}{(1 + \Xi)[1 + C_1^2 - 4C_2(2C_2 + \pm\sqrt{4C_2^2 - C_1^2 - 1})]}}.$$
(109)

The solution preserves N = 2 supersymmetry and breaks the SO(6) R symmetry to SO(3). The solution is singular when

$$\cos^2 \eta = -\frac{1}{1 + 2C_1^2 - 8C_2^2 \pm 4C_2\sqrt{4C_2^2 - C_1^2 - 1}}.$$
 (110)

This gives $\tilde{\Phi} \to \pm \infty$ or $\Phi \to \pm \infty$, which, in turn, leads to

$$V \to g^2 e^{4|\Phi|} \cos^2 \theta \sin^2 \eta \sin^2 \theta. \tag{111}$$

We can see that the scalar potential is unbounded from above, $V \rightarrow +\infty$, unless $\theta = 0$ or $\eta = 0$, both of which give the N = 6 solution as previously mentioned. Therefore, the

IR singularities of the N = 2 solutions are unphysical by the criterion of Ref. [77].

IV. SUPERSYMMETRIC JANUS SOLUTIONS

In this section, we consider supersymmetric Janus solutions in the form of curved domain walls. The solutions can be obtained from an AdS₃-sliced domain wall ansatz

$$ds^{2} = e^{2A} (e^{2\xi/\ell} dx_{1,1}^{2} + d\xi^{2}) + dr^{2}.$$
 (112)

Since the analysis closely follows that given in Ref. [52] (see also [53]), we will not repeat all the detail here but mainly review relevant results for deriving the corresponding BPS equations.

Compared to the RG flow case, the BPS equations will be modified by the curvature of the three-dimensional slices. In addition, as pointed out in Ref. [52], the existence of Janus solutions requires nonvanishing pseudoscalars, resulting in a more complicated set of BPS equations in contrast to the simple flat domain wall or RG flow case. We begin with the conditions $\delta \psi_{\hat{\mu}}^{i} = 0$ for $\hat{\mu} = 0, 1$, which give

$$A'\gamma_{\hat{r}}\epsilon_A + \frac{1}{\ell'}e^{-A}\gamma_{\hat{\xi}}\epsilon_A - \mathcal{W}\epsilon^A = 0.$$
(113)

This leads to the following BPS equation:

$$A^{\prime 2} = W^2 - \frac{1}{\ell^2} e^{-2A} \tag{114}$$

for W = |W|. We still use the $\gamma_{\hat{r}}$ projection given in Eq. (38). Imposing the $\gamma_{\hat{\xi}}$ projection of the form

$$\gamma_{\hat{\varepsilon}}\epsilon_A = i\kappa e^{i\Lambda}\epsilon^A \tag{115}$$

with $\kappa^2 = 1$, we can solve the condition (113) for the $\gamma_{\hat{r}}$ projector, leading to the phase factor

$$e^{i\Lambda} = \frac{A'}{W} + \frac{i\kappa}{\ell} \frac{e^{-A}}{W}$$
(116)

for real \mathcal{W} and

$$e^{i\Lambda} = \frac{\mathcal{W}}{A' + \frac{i\kappa}{\ell}e^{-A}} \tag{117}$$

for complex W. We also note that the constant $\kappa = \pm 1$ corresponds to the chiralities of the Killing spinors on the two-dimensional conformal defects described by the AdS₃ slices. Finally, the conditions $\delta \psi_{\hat{\xi}}^A = 0$ and $\delta \psi_{\hat{r}}^A = 0$ can be solved to obtain the explicit form of the Killing spinors (see details in [52]):

$$\epsilon_A = e^{(A/2) + (\xi/2\ell) + i(\Lambda/2)} \epsilon_A^{(0)}, \qquad (118)$$

in which the constant spinors $\varepsilon_A^{(0)}$ could possibly have an *r*-dependent phase and satisfy

$$\gamma_{\hat{r}}\varepsilon_A^{(0)} = \varepsilon^{(0)A} \quad \text{and} \quad \gamma_{\hat{\xi}}\varepsilon_A^{(0)} = i\kappa\varepsilon^{(0)A}.$$
 (119)

It turns out that among the previously considered cases only $SO(2) \times SO(4)$ and SO(3) symmetric scalars can possibly possess supersymmetric Janus solutions. This is mainly a consequence of the consistency in turning on nonvanishing pseudoscalars. For the SO(3) case, the analysis is highly complicated as already seen in the case of RG flows considered in the previous section. Therefore, we will give only the Janus solution with $SO(2) \times SO(4)$ symmetry. This case is more traceable, and it turns out that the solution can be analytically obtained.

A. Janus solutions with $SO(2) \times SO(4)$ symmetry

In this case, the superpotential is real, so we will use the phase $e^{i\Lambda}$ from Eq. (116). In general, since $e^{1.2}$ and e^A with A = 3, 4, 5, 6 transform differently under $SO(2) \times SO(4)$, namely, as (2, 1) + (1, 4), the two sets of Killing spinors can satisfy different projectors. We find that, in order to obtain a consistent set of BPS equations, we need to choose opposite signs of κ for $e^{1.2}$ and $e^{3,\dots,6}$. Therefore, the surface defect will preserve N = (2, 4) or N = (4, 2) superconformal symmetry.

With the superpotential

$$\mathcal{W} = 4g\cosh\varphi,\tag{120}$$

we find the following BPS equations:

$$\varphi' = -\frac{8g^2 \ell'^2 A' e^{2A} \sinh(2\varphi)}{1 + \ell'^2 A'^2 e^{2A}},$$
(121)

$$\zeta' = -\frac{16g^2 \kappa \ell e^A}{1 + \ell^2 A'^2 e^{2A}},$$
(122)

$$A'^{2} + \frac{e^{-2A}}{\ell^{2}} = 16g^{2}\ell^{2}\cosh^{2}\varphi.$$
 (123)

It should be noted that, for $\ell \to \infty$, these equations reduce to those of the RG flow studied in the previous section. Furthermore, these equations take a very similar form to the SO(4) symmetric Janus solution in N = 5 gauged supergravity studied in Ref. [51].

By taking φ as an independent variable, we can solve for $A(\varphi)$ and $\zeta(\varphi)$. The complete solution is given by

$$A = C - \ln \sinh \varphi, \tag{124}$$

$$\cosh(2\varphi) = \frac{32g^2\ell^2 \tanh^2[4g(r-r_0)]}{16g^2\ell^2 - 1}, \quad (125)$$

$$\kappa \tan \zeta = -\sqrt{1 - 16g^2\ell^2} \sinh[4g(r - r_0)] \quad (126)$$

for constants *C* and r_0 . This solution takes the same form as the solution given in Refs. [52,51,53], respectively, in N = 8, 5, and 3 gauged supergravities. We also note the unbroken supersymmetries on the conformal defects in these cases as follows: N = (4, 4), N = (4, 1), and N = (2, 1). All of these solutions should be related by truncations of N = 8 gauged supergravity to N = 3 and N = 5, 6 theories. This indicates that the N = (4, 4) Janus solution of Ref. [52] survives in the truncation to N = 3, 5, 6 gauged supergravities. According to the AdS/CFT correspondence, we then expect the dual N = 3, 5, 6 SCFTs to possess the same two-dimensional conformal defect as in the N = 8 SCFT.

We end this section by giving a brief comment on the possible SO(3) symmetric Janus solution. A partial analysis shows that there appears to be no obstruction in obtaining the BPS equations for Janus solutions in this case as in a similar analysis of N = 5 theory in Ref. [51]. Therefore, we expect a supersymmetric N = 2 Janus solution with SO(3) symmetry to exist in N = 6 gauged supergravity as well. Since the analysis is far more complicated than the $SO(2) \times SO(4)$ case, we refrain from giving any definite result here.

V. SUPERSYMMETRIC AdS₄ BLACK HOLES

In this section, we look for supersymmetric AdS_4 black holes in the form of curved-domain wall solutions interpolating between a locally asymptotically AdS_4 and $AdS_2 \times \Sigma^2$ geometries. The $AdS_2 \times \Sigma^2$ space describes the near-horizon geometry of the black holes. In the following analysis, we will consider only the cases of Σ^2 being a twosphere (S^2) and a hyperbolic space (H^2).

The metric ansatz is taken to be

$$ds^{2} = -e^{2f(r)}dt^{2} + dr^{2} + e^{2h(r)}(d\theta^{2} + F^{2}(\theta)d\phi^{2}) \quad (127)$$

with the function $F(\theta)$ defined by

$$F(\theta) = \begin{cases} \sin \theta, & \Sigma_2 = S^2, \\ \sinh \theta, & \Sigma_2 = H^2. \end{cases}$$
(128)

It is straightforward to derive nonvanishing components of the spin connection

$$\omega^{\hat{t}\,\hat{r}} = f'e^{\hat{t}}, \qquad \omega^{\hat{\theta}\,\hat{r}} = h'e^{\hat{\theta}},$$
$$\omega^{\hat{\phi}\,\hat{r}} = h'e^{\hat{\phi}}, \qquad \omega^{\hat{\theta}\,\hat{\phi}} = \frac{F'}{F}e^{-h}e^{\hat{\phi}} \qquad (129)$$

with $F'(\theta) = \frac{dF}{d\theta}$.

In general, the curvature of the Σ^2 part on the world volume of the domain wall will completely break supersymmetry. However, it is well known that some amount of supersymmetry can be preserved by performing a topological twist. This can be achieved by turning on some gauge fields along Σ^2 in such a way that the corresponding spin connection, $\omega^{\hat{\theta}\hat{\phi}}$ in the above metric, is canceled. This turns the covariant derivative of ϵ_A along Σ^2 into a partial one. The resulting Killing spinors are accordingly given by spinors that are independent of the Σ^2 coordinates.

A. $SO(2) \times SO(2) \times SO(2)$ symmetric solutions

We first consider the $SO(2) \times SO(2) \times SO(2)$ twist with the following ansatz for the gauge fields:

$$A^{12} = A_1(r)dt - p_1 F'(\theta)d\phi,$$

$$A^{34} = A_2(r)dt - p_2 F'(\theta)d\phi,$$

$$A^{56} = A_3(r)dt - p_3 F'(\theta)d\phi.$$
 (130)

The constants p_i , i = 1, 2, 3, are identified with magnetic charges. The corresponding field strengths are given by

$$F^{12} = A'_{1}dr \wedge dt + \kappa p_{1}F(\theta)d\theta \wedge d\phi,$$

$$F^{34} = A'_{2}dr \wedge dt + \kappa p_{2}F(\theta)d\theta \wedge d\phi,$$

$$F^{56} = A'_{3}dr \wedge dt + \kappa p_{3}F(\theta)d\theta \wedge d\phi.$$
 (131)

We have introduced a parameter κ with $\kappa = 1$, -1 for Σ^2 being S^2 or H^2 , respectively. We also note that $F''(\theta) = -\kappa F(\theta)$.

With the $SO(2) \times SO(2) \times SO(2)$ singlet scalars given by Eq. (62), we find nonvanishing components of the composite connection:

$$Q_A{}^B = 2gi\sigma_2 \otimes \begin{pmatrix} A^{12} & & \\ & A^{34} & \\ & & A^{56} \end{pmatrix}.$$
(132)

With the component $Q_{\hat{\phi}A}{}^B$, the spin connection $\omega^{\hat{\theta}\hat{\phi}}$ can be canceled by imposing the following projector:

$$\gamma_{\hat{\theta}\hat{\phi}}\epsilon_A = (i\sigma_2 \otimes \mathbb{I}_3)_A{}^B\epsilon_B \tag{133}$$

and the twist conditions

$$2gp_1 = 2gp_2 = 2gp_3 = 1. \tag{134}$$

These conditions imply that $p_1 = p_2 = p_3$. The twist is then obtained from the diagonal subgroup $SO(2)_{\text{diag}} \subset$ $SO(2) \times SO(2) \times SO(2)$ as in the pure N = 4 and N = 5gauged supergravities studied in Refs. [78,51], respectively. For consistency, we will also set $A_3 = A_2 = A_1$.

For all p_i nonvanishing, the twists allow all the supersymmetries corresponding to ϵ^A , A = 1, 2, ..., 6, to be unbroken subject to the projector (133). We also note some useful relations for deriving the full set of BPS equations. Using $\epsilon_{\hat{0}\hat{r}\hat{\theta}\hat{\phi}} = 1$ and $\gamma_5\epsilon_A = -\epsilon_A$, we find

$$\gamma^{\hat{\theta}\hat{\tau}}\epsilon_A = -i\gamma^{\hat{\theta}\hat{\phi}}\epsilon_A = (\sigma_2 \otimes \mathbb{I}_3)_A{}^B\epsilon_B.$$
(135)

It turns out that we need to turn on the $SO(2) \sim U(1)$ gauge field of $U(6) \sim SU(6) \times U(1)$ in order to find a consistent set of BPS equations. We similarly take the ansatz for this U(1) gauge field to be

$$A^{0} = A_{0}(r)dt - p_{0}F'(\theta)d\phi \quad \text{and}$$

$$F^{0} = A_{0}'(r)dr \wedge dt + \kappa p_{0}d\theta \wedge d\phi. \quad (136)$$

We also note that both A^0 and A^{IJ} can appear in the BPS equations due to the off-diagonal element of the scalar coset representative. In particular, we have the relations

$$\hat{F}_{AB}^{+} = h_{\Lambda AB}F^{+\Lambda} = h_{0AB}F^{+0} + \frac{1}{2}h_{IJ,AB}F^{+IJ}, \quad (137)$$

$$\hat{F}^{+} = h_{\Lambda 0}F^{+\Lambda} = h_{00}F^{+0} + \frac{1}{2}h_{IJ0}F^{+IJ}.$$
 (138)

However, A^0 does not participate in the twist, since ϵ_A are not charged under the U(1) factor outside SU(6).

It is also useful to define the "central charge" matrix

$$\mathcal{Z}_{AB} = -\frac{1}{\sqrt{2}} (\hat{F}^+_{\hat{\theta}\hat{\phi}AC} - i\hat{F}^+_{\hat{0}\hat{r}AC}) (i\sigma_2 \otimes \mathbb{I}_3)^C{}_B.$$
(139)

In the present case, it turns out that Z_{AB} is proportional to the identity matrix: $Z_{AB} = Z \delta_{AB}$.

With all these and the projector (38), we find the following BPS equation, from $\delta \psi_{\hat{\theta}A}$ and $\delta \psi_{\hat{\theta}A}$:

$$h'e^{i\Lambda} - \mathcal{W} - \mathcal{Z} = 0, \tag{140}$$

which leads to

$$h' = \pm |\mathcal{W} + \mathcal{Z}|$$
 and $e^{i\Lambda} = \pm \frac{\mathcal{W} + \mathcal{Z}}{|\mathcal{W} + \mathcal{Z}|}.$ (141)

a|w + z|

With the projectors (38) and (135), the condition $\delta \psi_{\hat{0}A} = 0$ gives

$$(f' + 2igA_1)e^{i\Lambda} - \mathcal{W} + \mathcal{Z} = 0, \qquad (142)$$

which implies

$$f' = \operatorname{Re}[e^{-i\Lambda}(\mathcal{W} - \mathcal{Z})]$$
 and $2gA_1 = \operatorname{Im}[e^{-i\Lambda}(\mathcal{W} - \mathcal{Z})].$

(143)

The latter fixes the time component of the gauge fields. Finally, as in the case of domain walls and Janus solutions, the condition $\delta \psi_{\hat{r}A} = 0$ determines the *r* dependence of the Killing spinors giving rise to $\epsilon_A = e^{f/2} \epsilon_{A(0)}$.

Similar to the RG flow case, it turns out that we need to set $\zeta_1 = \zeta_2 = \zeta_3 = 0$ for consistency. This gives real Wand Z resulting in $e^{i\Lambda} = \pm 1$ and $A_1 = 0$. We will also set $A_0(r) = 0$ for simplicity, although it is not constrained by the previously obtained conditions. In addition, the compatibility between the BPS equations coming from $\delta \chi_A$ and $\delta \chi_{ABC}$ requires

$$p_0 = \kappa p_1. \tag{144}$$

We also note that, in this case, the choice $p_0 = 0$ breaks all supersymmetry. This implies that the $SO(2) \times SO(2) \times SO(2)$ twist needs to be accompanied by the U(1) gauge field A_u^0 .

With all these, we find a consistent set of BPS equations given by

$$\varphi_1' = -\frac{\partial |\nu \nu + 2|}{\partial \varphi_1}$$

= $\frac{1}{2} e^{-\varphi_1 - \varphi_2 - \varphi_3} [2g(1 + e^{2(\varphi_2 + \varphi_3)} - e^{2(\varphi_1 + \varphi_2)} - e^{2(\varphi_1 + \varphi_3)}) - p_1 \kappa e^{-2h + 2\varphi_1 + 2\varphi_2 + \varphi_3}],$ (145)

$$\varphi_{2}^{\prime} = -\frac{\partial |\mathcal{W} + \mathcal{Z}|}{\partial \varphi_{2}}$$

= $\frac{1}{2} e^{-\varphi_{1} - \varphi_{2} - \varphi_{3}} [2g(1 - e^{2(\varphi_{2} + \varphi_{3})} - e^{2(\varphi_{1} + \varphi_{2})} + e^{2(\varphi_{1} + \varphi_{3})}) - p_{1}\kappa e^{-2h + 2\varphi_{1} + 2\varphi_{2} + \varphi_{3}}],$ (146)

$$\varphi_{3}' = -\frac{\partial |\mathcal{W} + \mathcal{Z}|}{\partial \varphi_{3}}$$

= $\frac{1}{2} e^{-\varphi_{1} - \varphi_{2} - \varphi_{3}} [2g(1 - e^{2(\varphi_{2} + \varphi_{3})} + e^{2(\varphi_{1} + \varphi_{2})} - e^{2(\varphi_{1} + \varphi_{3})}) - p_{1}\kappa e^{-2h + 2\varphi_{1} + 2\varphi_{2} + \varphi_{3}}],$ (147)

$$h' = |\mathcal{W} + \mathcal{Z}|$$

= $\frac{1}{2}e^{-\varphi_1 - \varphi_2 - \varphi_3} [2g(1 + e^{2(\varphi_2 + \varphi_3)} + e^{2(\varphi_1 + \varphi_2)} + e^{2(\varphi_1 + \varphi_3)}) + p_1 \kappa e^{-2h + 2\varphi_1 + 2\varphi_2 + \varphi_3}],$ (148)

$$f' = |\mathcal{W} - \mathcal{Z}|$$

= $\frac{1}{2}e^{-\varphi_1 - \varphi_2 - \varphi_3}[2g(1 + e^{2(\varphi_2 + \varphi_3)} + e^{2(\varphi_1 + \varphi_2)} + e^{2(\varphi_1 + \varphi_3)}) - p_1\kappa e^{-2h + 2\varphi_1 + 2\varphi_2 + \varphi_3}].$ (149)

For an AdS₂ × Σ^2 fixed point to exist, we require that $\varphi'_1 = \varphi'_2 = \varphi'_3 = h' = 0$ and $f' \sim \frac{1}{L_{AdS_2}}$. It can be easily verified that the above equations do not admit any AdS₂ × Σ^2 fixed points.

Although there is no supersymmetric $AdS_2 \times \Sigma^2$ fixed point, we are able to analytically obtain the complete solution to these BPS equations. Since it might be useful for some holographic studies, we will present the solution here. By changing to a new radial coordinate ρ using the relation $\frac{d\rho}{dr} = e^{\varphi_3}$, we can form the following linear combinations:

$$\frac{d}{d\rho}(\varphi_1 - \varphi_2) = 2g(e^{\varphi_2 - \varphi_1} - e^{\varphi_1 - \varphi_2})$$
(150)

and

$$\frac{d}{d\rho}(\varphi_2 - \varphi_3) = 2g(e^{\varphi_1 - \varphi_2} - e^{\varphi_1 + \varphi_2 - 2\varphi_3}).$$
(151)

The first equation can be solved by

$$\varphi_1 = \ln \left[\frac{e^{\varphi_2} (e^{4g\rho} + e^{4g\rho_0})}{e^{4g\rho} - e^{4g\rho_0}} \right]$$
(152)

with an integration constant ρ_0 . Using this result in the second equation, we find the solution

$$\varphi_2 = \ln \left[\frac{e^{\varphi_3 - 2g\rho} (e^{4g\rho} - e^{4g\rho_0})}{\sqrt{e^{4g\rho} + e^{8g\rho_0} + 8gC}} \right]$$
(153)

with another integration constant C.

By treating f and h as functions of φ_3 , we find the following solutions for f and h:

$$f = -\frac{1}{2} \ln \left[e^{4g\rho_0} \left(256g^3 C \tilde{C} - 16g \tilde{C} e^{8g\rho_0} + \kappa p_1 \ln \left[\frac{1 + e^{8g(\rho - \rho_0)} + 8gC e^{4g\rho - 8g\rho_0}}{e^{8g(\rho - \rho_0)} - 1} \right] \right) - 8gC \kappa p_1 \tanh^{-1} e^{4g(\rho - \rho_0)} \right] + h,$$
(154)

$$h = \frac{1}{2} \ln \left[e^{4g\rho_0} \left(16g\tilde{C}(e^{8g\rho_0} - 16g^2C^2) - \kappa p_1 \ln \left[\frac{1 + e^{8g(\rho - \rho_0)} + 8gCe^{4g\rho - 8g\rho_0}}{e^{8g(\rho - \rho_0)} - 1} \right] \right) + 8gC\kappa p_1 \tanh^{-1} e^{4g(\rho - \rho_0)} \right] + \frac{1}{2} \ln \left[\frac{e^{12g\rho_0} (1 - e^{8g(\rho - \rho_0)})^2}{8g(e^{8g\rho_0} - 16g^2C^2)} \right] + \varphi_3 - 4g\rho.$$
(155)

Finally, the solution $\varphi_3(\rho)$ can be given implicitly in the following equation:

$$4C_0 e^{4g\rho} (e^{8g\rho_0} + e^{8g\rho} + 8gCe^{4g\rho}) = \beta_0 + \beta_1 \ln \left[\frac{e^{4g(\rho_0 - \rho)} + 1}{e^{4g(\rho_0 - \rho)} - 1} \right] + \beta_2 \ln \left[\frac{e^{8g(\rho - \rho_0)} - 1}{1 + e^{8g(\rho - \rho_0)} + 8gCe^{4g\rho}} \right]$$
(156)

in which C_0 is a constant and the coefficients β_0 , β_1 , and β_2 are defined, respectively, by

$$\beta_0 = -16g\tilde{C}e^{4\rho_0}(16Cg^2 - e^{8g\rho_0})[2e^{4\varphi_3 + 8g(\rho + \rho_0)} + 8gCe^{4g\rho}(e^{8g\rho} + e^{8g\rho_0}) + (e^{16g\rho_0} + e^{16g\rho})(1 - e^{4\varphi_3}) + 2e^{8g(\rho + \rho_0)}], \quad (157)$$

$$\beta_{1} = \left[\frac{\kappa p_{1}}{2(e^{4g\rho} + e^{4g(3\rho - 2\rho_{0})} + 8Cge^{8g(\rho - \rho_{0})})}\right] [e^{12g\rho} + e^{4g(\rho + 2\rho_{0})} + 4Cge^{8g\rho}(3 + e^{4\varphi_{3}}) + 16g^{2}C^{2}e^{4g\rho}(1 + e^{8g(\rho - \rho_{0})}) - 2Cg(e^{4\varphi_{3}} - 1)(e^{8g\rho_{0}} + e^{8g(2\rho - \rho_{0})})],$$
(158)

$$\beta_2 = \frac{\kappa p_1 [(e^{8g\rho_0} + e^{8g\rho} + 8Cge^{4g\rho})^2 - e^{4\varphi_3} (e^{8g\rho_0} - e^{8g\rho})^2]}{4e^{4g(\rho - \rho_0)} (e^{8g\rho_0} + e^{8g\rho} + 8Cge^{4g\rho})}.$$
(159)

Since there is no $AdS_2 \times \Sigma^2$ fixed point in the IR, the solution describes a flow from the locally supersymmetric AdS_4 vacuum to a curved domain wall with world volume $\mathbb{R} \times \Sigma^2$. According to the AdS/CFT correspondence, the solution is expected to describe an RG flow from the N = 6SCFT in three dimensions to a supersymmetric quantum mechanics in the IR. The latter arises from the former by a twisted compactification on Σ^2 .

B. $SO(2) \times SO(4)$ symmetric solutions

We now look at a truncation of the previous result by setting $p_2 = p_3 = 0$ and $\varphi_2 = \varphi_3 = 0$. The resulting solutions will preserve $SO(2) \times SO(4)$ symmetry with the twist performed along the SO(2) factor. In this case, the supersymmetry corresponding to $e^{3,4,5,6}$ will be broken, since we cannot perform a twist along these directions. With $e^{3,4,5,6} = 0$, we find the BPS equations

$$\varphi' = \frac{1}{4}e^{-2h-\varphi}[8ge^{2h} - p_0 + \kappa p_1 - e^{2\varphi}(8ge^{2h} + p_0 + \kappa p_1)],$$
(160)

$$h' = \frac{1}{4}e^{-2h-\varphi}[8ge^{2h} - p_0 + \kappa p_1 + e^{2\varphi}(8ge^{2h} + p_0 + \kappa p_1)],$$
(161)

$$f' = \frac{1}{4}e^{-2h-\varphi}[8ge^{2h} + p_0 - \kappa p_1 + e^{2\varphi}(8ge^{2h} - p_0 - \kappa p_1)],$$
(162)

in which we have set $\varphi_1 = \varphi$. We also note that, with only the SO(2) twist, it is not necessary to set $p_0 = \kappa p_1$. However, the existence of an $AdS_2 \times \Sigma^2$ fixed point requires vanishing p_0 . For $p_0 = 0$, we find a fixed point

$$\varphi = \varphi_0, \qquad h = \frac{1}{2} \ln \left[-\frac{\kappa p_1}{8g} \right], \qquad L_{\text{AdS}_2} = \frac{1}{8g \cosh 2\varphi_0}.$$
(163)

for constant φ_0 . This is an AdS₂ × H^2 fixed point, since the reality of *h* implies $\kappa = -1$.

The complete flow solution can be obtained by using the same procedure as in the previous sections. The resulting solution is given by

$$h = \varphi - \ln(1 - e^{2\varphi}) + C,$$
 (164)

$$f = h - 2\varphi + \ln[\kappa p_1(1 + e^{4\varphi}) + 2e^{2\varphi}(4g - \kappa p_1)], \quad (165)$$

$$8g(\rho - \rho_0) = 2\sqrt{\frac{2g}{\kappa p_1 - 2g}} \tan^{-1} \left[\frac{4g + \kappa p_1(e^{2\varphi} - 1)}{2\sqrt{2g(\kappa p_1 - 2g)}}\right] + \ln \left[\frac{\kappa p_1(1 + e^{4\varphi}) + 2e^{2\varphi}(4g - \kappa p_1)}{(1 - e^{2\varphi})^2}\right],$$
(166)

in which we have defined the new radial coordinate ρ by $\frac{d\rho}{dr} = e^{\varphi}$. We have neglected the integration constant of f by absorbing it in the rescaling of the time coordinate t.

Near $r \sim \rho \rightarrow \infty$, we find

$$\varphi \sim e^{-4gr}, \qquad h \sim f \sim 4gr$$
 (167)

which gives an asymptotically locally AdS_4 critical point. On the other hand, by choosing $\varphi_0 = \frac{1}{2} \ln \left(1 - 2\sqrt{-\frac{2g}{\kappa p_1}} \right)$ and $C = -\varphi_0$, we find that as $\varphi \to \varphi_0$

$$h \sim \frac{1}{2} \ln \left[-\frac{\kappa p_1}{8g} \right] \quad \text{and} \quad f \sim 8gr \frac{1 - \sqrt{-\frac{2g}{\kappa p_1}}}{\sqrt{1 - 2\sqrt{-\frac{2g}{\kappa p_1}}}}, \quad (168)$$

which is the $AdS_2 \times H^2$ fixed point identified above.

It should be noted that, in this case, the solution can be regarded as a solution of a truncated N = 2 gauged supergravity. In particular, the solution with vanishing scalar corresponds to a universal RG flow across dimension of which the uplifts to M theory and massive type IIA theory have been extensively studied in Refs. [67].

C. U(3) symmetric solutions

As a final case, we consider U(3) symmetric solutions with a twist performed along the $SO(2) \sim U(1)$ factor. The corresponding gauge generator of this U(1) factor is given by $X_{14} + X_{25} + X_{36}$. We then turn on the following gauge fields:

$$\mathcal{A} = A^{14} = A^{25} = A^{36} = A(r)dt - \kappa p F'(\theta)d\phi.$$
(169)

With the U(3) singlet scalar given in Eq. (51), we find the composite connection

$$Q_A{}^B = 2gi\mathcal{A}(\mathbb{I}_3 \otimes \sigma_2)_A{}^B.$$
(170)

The twist is implemented by imposing

$$\gamma_{\hat{\theta}\hat{\phi}}\epsilon_A = (\mathbb{I}_3 \otimes i\sigma_2)_A{}^B\epsilon_B \quad \text{and} \quad 2gp = 1.$$
(171)

We also note that, similar to the $SO(2) \times SO(2) \times SO(2)$ twist, all ϵ_A can be nonvanishing. In addition, we also need nonvanishing A^0 , which we will again use the ansatz (136). In this case, consistency requires $p_0 = -\kappa p$. As in the RG flow case, we need to set $\zeta = 0$ for consistency between the BPS equations and the field equations. This again results in A(r) = 0. Repeating the same analysis as in the previous cases, we find the following BPS equations:

$$\varphi' = g(e^{-\varphi} - e^{3\varphi}) + \frac{1}{2}\kappa p e^{-2h-3\varphi},$$
 (172)

$$h' = g(3e^{-\varphi} + e^{3\varphi}) + \frac{1}{2}\kappa p e^{-2h-3\varphi}, \qquad (173)$$

$$\varphi' = g(3e^{-\varphi} + e^{3\varphi}) - \frac{1}{2}\kappa p e^{-2h-3\varphi}.$$
 (174)

It is easily verified that there is no $AdS_2 \times \Sigma^2$ fixed point in these equations. In this case, we are not able to obtain the analytic flow solution.

VI. CONCLUSIONS AND DISCUSSIONS

In this paper, we have studied N = 6 gauged supergravity in four dimensions with the SO(6) gauge group. The gauged supergravity can be obtained from a truncation of the maximal N = 8 theory with the SO(8) gauge group. There is a unique N = 6 supersymmetric AdS₄ vacuum preserving the full SO(6) gauge symmetry. This can be identified with $AdS_4 \times CP^3$ geometry in type IIA theory dual to an N = 6 SCFT in three dimensions. We have found a number of RG flow solutions with various symmetries from this N = 6 SCFT to possible nonconformal phases in the IR. In particular, there is one solution, breaking the SO(6) R symmetry to $SO(2) \times SO(4)$, with unbroken N = 6 Poincaré supersymmetry. This is precisely in agreement with the field theory result on mass deformations of N = 6 SCFTs given in Ref. [25]. Other solutions preserve U(3), SO(3), and $SO(2) \times SO(2) \times SO(2)$ symmetries. While most of the solutions preserve N = 6 supersymmetry, in the case of SO(3) symmetry, it is possible to find N = 2 supersymmetric solutions. We have analytically given all of these solutions and also checked that, except for the N = 2 solution, the resulting IR singularities are physical by the criterion given in Ref. [77].

We have also considered more complicated solutions by generalizing the flat domain walls to the curved ones. In the case of AdS₃-sliced domain walls, we have found a supersymmetric Janus solution, describing a two-dimensional conformal defect within the N = 6 SCFT, with $SO(2) \times SO(4)$ symmetry and N = (2, 4) supersymmetry on the defect. The resulting solution takes the same form as those given in N = 8, N = 5, and N = 3 gauged supergravities studied in Refs. [52,51,53], respectively. We therefore argue that these solutions are related to the N = 8 solution by truncations. In order for Janus solutions to exist, it is necessary that pseudoscalars are nonvanishing as pointed out in Ref. [52]. It turns out that, among the remaining cases studied in this work, only the SO(3) invariant sector could possibly admit supersymmetric Janus solutions.

Furthermore, we have studied supersymmetric solutions of the form $AdS_2 \times \Sigma^2$ and the interpolating solutions between these geometries and the $N = 6 AdS_4$ vacuum. We have found one $AdS_2 \times H^2$ fixed point with $SO(2) \times$ SO(4) symmetry from SO(2) twist. The solution interpolating between this fixed point and the AdS_4 vacuum preserves two supercharges, while the IR fixed point $AdS_2 \times H^2$ has four supercharges. Holographically, this solution corresponds to an RG flow from the N = 6SCFT to superconformal quantum mechanics which is useful in computing black hole entropy along the lines of Refs. [67–69].

For $SO(2) \times SO(2) \times SO(2)$ twist, the BPS equations are more complicated but admit no $AdS_2 \times \Sigma^2$ fixed point. However, in this case, we are able to obtain a complete flow solution between the AdS_4 critical point to a curved domain wall with world volume $\mathbb{R} \times \Sigma^2$ in the IR. The solution preserves N = 6 supersymmetry in three dimensions, or 12 supercharges, and $SO(2) \times SO(2) \times SO(2)$ symmetry. This should be dual to a twisted compactification on Σ^2 of the UV N = 6 SCFT to a supersymmetric quantum mechanics in the IR. We have also looked for $AdS_2 \times \Sigma^2$ geometries from an $SO(2) \sim U(1)$ twist in the case of U(3)symmetric solutions, but there do not exist any $AdS_2 \times \Sigma^2$ fixed points.

Since all the solutions presented here are fully analytic, we hope they could be useful in the study of gauge and gravity holography and other related aspects. We also note that most of the structures of the solutions are very similar to those of the N = 5 gauged supergravity studied in Ref. [51]. In particular, the N = 6 Poincaré supersymmetry in three dimensions is unbroken on the domain wall solutions if there are no nonvanishing pseudoscalars. Unlike in the N = 5 theory, we are not able to find a definite conclusion on whether this is true, in general, due to a more complicated scalar coset manifold. However, many similarities in the structures of various types of supersymmetric solutions suggest that this should be the case.

There are a number of directions to extend the present work, which is clearly only the first step in classifying supersymmetric solutions of N = 6 gauged supergravity. First of all, it would be interesting to uplift the RG flow solutions to M theory via the embedding in N = 8 gauged supergravity, which, in turn, can be obtained from a truncation of M theory on S^7 . The time component g_{00} of the 11-dimensional metric can be used to determine whether the four-dimensional singularities, in particular, the N = 2 case, are physically acceptable in M theory using the criterion given in Ref. [79]. This would lead to a complete holographic description of mass deformations of N = 6 CSM theory and possible related M-brane configurations. We have considered only gauged supergravity with the SO(6) gauge group electrically embedded in the global $SO^*(12)$ symmetry. It would be interesting to study magnetic and dyonic gaugings involving also magnetic gauge fields. In particular, performing a similar study in the case of N = 6 gauged supergravity with the electric-magnetic phase ω (see [80–82]) could be of particular interest, since, in the omega deformed N = 8 theory, the structure of vacua and domain walls are much richer than the electric counterpart; see [43,83–85] for more detail. In addition, the study of genuine N = 6 gaugings which cannot be embedded in the N = 8 theory is worth considering. In this case, the gaugings

do not satisfy extra quadratic constraints coming from the truncation of the N = 8 theory (see the discussion in Ref. [86]), so the corresponding solutions cannot be embedded in the maximal theory.

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