

Flat space holography and the complex Sachdev-Ye-Kitaev model

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We provide the first steps toward a flat space holographic correspondence in two bulk spacetime dimensions. The gravity side is described by a conformally transformed version of the matterless Callan-Giddings-Harvey-Strominger model. The field theory side follows from the complex Sachdev-Ye-Kitaev model in the limit of large specific heat and vanishing compressibility. We derive the boundary action analogous to the Schwarzian as the key link between gravity and field theory sides and show that it coincides with a geometric action discovered recently by one of us [H. R. Afshar, Warped Schwarzian theory, *J. High Energy Phys.* **02** (2020) 126].

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I. INTRODUCTION

The Sachdev-Ye-Kitaev (SYK) model [1–3] reinvigorated studies of Jackiw-Teitelboim (JT) gravity [4,5] since in a certain limit it is the gravity dual of the former [6]. This holographic relationship, dubbed $n\text{AdS}_2/n\text{CFT}_1$, inspired numerous research activities in the past few years in gravity and condensed-matter communities; see e.g., [7–46]. One crucial piece of the puzzle is the Schwarzian action [1,6] that arises in the large N and strong coupling limit on the quantum mechanics side and, upon imposing suitable boundary conditions, also on the gravity side.

Given the impressive evidence for AdS/CFT realizations of holography, it is justified to ask how general the holographic principle works, if it works beyond the AdS/CFT correspondence and in particular if and how it works in asymptotically flat spacetimes. See [47–70] and references therein for selected earlier results in flat space holography.

The main goal of our paper is to find a model analogous to JT that leads to a holographic relationship involving flat

space instead of AdS_2 , with some suitable replacement of the Schwarzian action.

The Callan-Giddings-Harvey-Strominger (CGHS) model [71] is a prime candidate for the gravity side of flat space holography since all solutions have asymptotically vanishing Ricci scalar. This opens up the prospect to construct a concrete holographic correspondence between flat space dilaton gravity in $1+1$ dimensions and some cleverly designed quantum system of (complex) fermions in $0+1$ dimensions.

The principal result of this paper is that the flat space analog of the Schwarzian action is given by

$$I^{\text{tw}}[h, g] = \kappa \int_0^\beta d\tau \left(\mathcal{T} h'^2 - g' \left(i\mathcal{P}h' + \frac{h''}{h'} \right) + g'' \right), \quad (1)$$

where κ is a coupling constant, β is inverse temperature, and prime denotes derivative with respect to τ , the time direction along the boundary. The time-reparametrization field $h(\tau + \beta) = h(\tau) + \beta$ is quasiperiodic and the phase field $g(\tau)$, in the absence of winding, is periodic. When the functions \mathcal{T} and \mathcal{P} are constant, we refer to them as mass and charge, respectively. While mass can be arbitrary, it will turn out that regularity demands a linear relationship between charge and temperature. The superscript tw stands for “twisted warped” and stems from the symmetries (7) that govern our action. On the gravity side κ is essentially the inverse Newton constant, as evident from our starting point (2). On the field theory side, κ is essentially the geometric mean of specific heat at constant charge and zero

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temperature compressibility, as evident from our final equation (32).

The remainder of our paper is organized as follows. We start by gaining some intuition about our gravity model in the metric formulation and then switch to a gauge theoretic formulation. The latter is employed to derive our main result (1). Finally, we recover the boundary action (1) from a scaling limit of complex SYK.

II. METRIC FORMULATION OF CGHS

Following Cangemi and Jackiw [72], we manipulate the CGHS action [71] in three ways: (i) for simplicity, we set all matter fields to zero, (ii) we perform a Weyl rescaling (depending on the dilaton X) of the metric $g_{\mu\nu}$, and (iii) we “integrate in” an Abelian gauge field A_μ and an auxiliary scalar field Y that is constant on shell. The action ($\varepsilon^{\mu\nu}$ is the ε -tensor)

$$I_{\widehat{\text{CGHS}}} = \frac{\kappa}{2} \int d^2x \sqrt{-g} (XR - 2Y + 2Y\varepsilon^{\mu\nu} \partial_\mu A_\nu) \quad (2)$$

provides a reformulation of the CGHS model referred to as $\widehat{\text{CGHS}}$. We solve now the $\widehat{\text{CGHS}}$ field equations,

$$R = 0, \quad (3)$$

$$\varepsilon^{\mu\nu} \partial_\mu A_\nu = 1, \quad (4)$$

$$\nabla_\mu \nabla_\nu X - g_{\mu\nu} \nabla^2 X = g_{\mu\nu} Y \quad Y = \Lambda = \text{const}, \quad (5)$$

with suitable boundary and gauge fixing conditions.

In Eddington-Finkelstein gauge, the most general solution to the Ricci-flatness condition (3) is given by Rindler-type black hole metrics of the form

$$ds^2 = -2dudr + 2(\mathcal{P}(u)r + \mathcal{T}(u))du^2. \quad (6)$$

The loosest set of boundary conditions compatible with the gauge fixing (6) allows fluctuations of both free functions, $\delta\mathcal{P} \neq 0 \neq \delta\mathcal{T}$. These boundary- and gauge-fixing conditions are preserved by asymptotic Killing vectors $\xi(\varepsilon, \eta) = \varepsilon(u)\partial_u - (\varepsilon'(u)r + \eta(u))\partial_r$ since $\mathcal{L}_\xi g_{\mu\nu} = \mathcal{O}(\delta g_{\mu\nu})$, namely, $\mathcal{L}_\xi g_{rr} = 0 = \mathcal{L}_\xi g_{ur}$ and $\mathcal{L}_\xi g_{uu} = \delta_\xi \mathcal{P}r + \delta_\xi \mathcal{T}$, with $\delta_\xi \mathcal{P} = \varepsilon\mathcal{P}' + \varepsilon'\mathcal{P} + \varepsilon''$ and $\delta_\xi \mathcal{T} = \varepsilon\mathcal{T}' + 2\varepsilon'\mathcal{T} + \eta' - \eta\mathcal{P}$. Prime is the derivative along retarded time u and \mathcal{L}_ξ the Lie-derivative along ξ .

The Lie-bracket algebra of the asymptotic Killing vectors $[\xi(\varepsilon_1, \eta_1), \xi(\varepsilon_2, \eta_2)]_{\text{Lie}} = \xi(\varepsilon_1\varepsilon_2' - \varepsilon_2\varepsilon_1', (\varepsilon_1\eta_2 - \varepsilon_2\eta_1)')$ in terms of Laurent modes, $L_n = \xi(\varepsilon = -u^{n+1}, \eta = 0)$ and $M_n = \xi(\varepsilon = 0, \eta = u^{n-1})$, yields $[L_n, L_m]_{\text{Lie}} = (n-m)L_{n+m}$, $[L_n, M_m]_{\text{Lie}} = -(n+m)M_{n+m}$, and $[M_n, M_m]_{\text{Lie}} = 0$. This algebra consists of a Witt subalgebra generated by L_n and spin-0 supertranslations generated by M_n . We refer to it as BMS_2 [73].

In axial gauge for the $U(1)$ connection, the field equation (4) is solved by the two-dimensional Coulomb connection $A = rdu$. Its preservation under combined diffeomorphisms and gauge transformations, $\delta_{\xi, \sigma} A_\nu = \xi^\mu \partial_\mu A_\nu + A_\mu \partial_\nu \xi^\mu + \partial_\nu \sigma$, relates the functions η and σ , $\eta = \sigma'$, which can be interpreted in two different ways. Either one concludes that σ has to contain a $\ln u$ term, since M_0 is allowed to be nonzero, or one concludes that M_0 is forbidden, since σ is assumed to have a Laurent series around $u = 0$. The first option leads to the BMS_2 symmetries discussed above. The second option leads to a slight modification of the transformation properties,

$$\delta_\xi \mathcal{P} = \varepsilon\mathcal{P}' + \varepsilon'\mathcal{P} + \varepsilon'', \quad (7a)$$

$$\delta_\xi \mathcal{T} = \varepsilon\mathcal{T}' + 2\varepsilon'\mathcal{T} + \sigma'' - \sigma'\mathcal{P}. \quad (7b)$$

In the present paper, we focus on the second option, since it guarantees that the Wilson loop in the complex u plane encircling the origin $u = 0$ is gauge invariant, $\delta_\sigma \oint A = 0$; in other words, there are no winding modes.

Defining instead of M_n new Fourier modes $J_n := \xi(0, \sigma = u^n)$ yield the asymptotic symmetry algebra in terms of asymptotic Killing vector modes $[L_n, L_m]_{\text{Lie}} = (n-m)L_{n+m}$, $[L_n, J_m]_{\text{Lie}} = -mJ_{n+m}$, and $[J_n, J_m]_{\text{Lie}} = 0$, which is known as “warped Witt algebra,” the centerless version of either the warped conformal algebra [76] or the twisted warped conformal algebra [77].

Finally, the rr component of the field equation (5) is solved by dilaton fields linear in the radial coordinate

$$X = x_1(u)r + x_0(u). \quad (8)$$

The remaining components of the field equation (5), which involve the functions $x_i(u)$, will be determined in the gauge theoretic formulation of $\widehat{\text{CGHS}}$.

III. GAUGE THEORY FORMULATION OF CGHS

For the gauge theoretic formulation as non-Abelian BF theory [72,78], we use conventions analogous to [79]. The first order form of the CGHS bulk action (2) is given by

$$I^{\text{BF}}[B, \mathcal{A}] = \kappa \int \langle B, F \rangle, \quad (9)$$

where κ is the coupling constant, B is a scalar, and $F = d\mathcal{A} + \mathcal{A} \wedge \mathcal{A}$ the non-Abelian field strength. The connection

$$\mathcal{A} = \omega J + e^a P_a + AZ \quad (10)$$

contains dualized spin-connection ω , zweibein e^a , and $U(1)$ connection A . The generators obey the Maxwell algebra [80] whose nonzero commutators read $[P_a, P_b] = \varepsilon_{ab} Z$ and $[P_a, J] = \varepsilon_a^b P_b$. The scalar field

$$B = XZ + X^a \epsilon_a{}^b P_b + YJ \quad (11)$$

comprises the dilaton X , Lagrange multipliers X^a for torsion constraints, and the auxiliary field Y . Finally, \langle, \rangle denotes the bilinear form with nonvanishing entries,

$$\langle J, Z \rangle = -1 \langle P_a, P_b \rangle = \eta_{ab}. \quad (12)$$

We use light cone gauge for the Minkowski metric, $\eta_{+-} = 1$, in terms of which the Levi-Civita symbol is $\epsilon_{\pm}^{\pm} = \pm 1$, and the gauge algebra reads $[P_+, P_-] = Z$ and $[P_{\pm}, J] = \pm P_{\pm}$. Integrating out the Lagrange multipliers X^a and solving the torsion constraints, the action (9) with (10)–(12) can be shown to be equivalent to (2).

Boundary conditions compatible with the ones in the metric formulation are given by

$$\mathcal{A} = b^{-1}(d + a)b \quad B = b^{-1}xb, \quad (13)$$

with $b = \exp(-rP_+)$ and

$$a = (T(u)P_+ + P_- + \mathcal{P}(u)J)du, \quad (14)$$

$$x = x^+(u)P_+ + x_1(u)P_- + YJ + x_0(u)Z, \quad (15)$$

where both functions in the connection are allowed to vary, $\delta T \neq 0 \neq \delta \mathcal{P}$.

The equations of motion reduce to

$$da + a \wedge a = 0 = dx + [a, x]. \quad (16)$$

The first one is obeyed automatically by our ansatz (14). The second one, which states that x is the stabilizer of a , holds provided $Y = \Lambda = \text{const.}$, and the following differential equations are fulfilled: $(x^+)' = \mathcal{P}x^+ - TY$, $x_1' = -\mathcal{P}x_1 + Y$, and $x_0' = x^+ - \mathcal{T}x_1$. Using them, x is conveniently parametrized by two functions x_1 and x_0 ,

$$x = (x_0' + \mathcal{T}x_1)P_+ + x_1P_- + (x_1' + \mathcal{P}x_1)J + x_0Z. \quad (17)$$

IV. ASYMPTOTIC SYMMETRIES

The boundary condition preserving gauge transformations $\delta_\lambda B = [B, \lambda]$ and

$$\delta_\lambda \mathcal{A} = d\lambda + [\mathcal{A}, \lambda] \stackrel{\dagger}{=} \mathcal{O}(\delta \mathcal{A}) = (\mathcal{O}(r)P_+ + \mathcal{O}(1)J)du \quad (18)$$

are generated by gauge parameters $\lambda = b^{-1}\epsilon b$ with

$$\epsilon = \epsilon^+(u)P_+ + \epsilon(u)P_- + \epsilon^J(u)J + \sigma(u)Z. \quad (19)$$

The absence of the P_- and Z components on the right-hand side of (18) yields consistency relations between the functions in the gauge parameter (19).

$$\epsilon^J = \epsilon' + \mathcal{P}\epsilon \quad \epsilon^+ = \sigma' + \mathcal{T}\epsilon. \quad (20)$$

The boundary condition preserving gauge transformations (18)–(20) imply precisely the transformation laws (7), which is the twisted warped conformal transformation behavior introduced in [77]. Therefore, the analog of the conformal symmetries that govern the Schwarzian action are twisted warped conformal symmetries, which govern our boundary action (1).

V. BOUNDARY ACTIONS

Our derivation of the boundary action for $\widehat{\text{CGHS}}$ follows closely the derivation of the Schwarzian action in Sec. 3 of [79].

From now on, we work in Euclidean signature with periodic boundary time, $t \sim t + \beta$, where β is inverse temperature. Mapping Lorentzian to Euclidean results requires the following replacements: $u \rightarrow iu = t$, $\eta_{ab} \rightarrow \delta_{ab}$, $P_{\pm} = P_{\pm}^E \pm iP_0^E$, $Z = iZ^E$, $J = iJ^E$, and $x_{0,1} = -ix_{0,1}^E$ (with real $x_{0,1}^E$). The fields are given by (10) and (11), with all quantities replaced by their Euclidean counterparts. We use the definition $\oint dt = \int_0^\beta dt$ and dot means $\frac{d}{dt}$.

The variation of the BF action (9) apart from the bulk equations of motion yields a boundary term [79],

$$\delta I^{\text{tw}} = -\kappa \oint dt \langle x, \delta a_t \rangle. \quad (21)$$

Our aim is to cancel this boundary term by variation of a boundary action. To this end, we use a convenient representation of the connection, $a_t = f_t x + G^{-1} \partial_t G$. A consistent choice is $G = \exp(x_0^E (iP_1^E - P_0^E)) \exp(-i \ln(-ix_1^E) J^E) \exp(-\int^t x_0^E/x_1^E Z^E)$ and $f_t = 1/x_1^E$. We impose as integrability condition that the function f_t has a fixed zero mode (which we set to unity with no loss of generality). As shown below, this guarantees that the first variation of the full action vanishes for all variations preserving our boundary conditions.

The variation of the boundary action (21) expands as

$$\begin{aligned} \delta I^{\text{tw}} = & -\kappa \oint dt [\delta(f_t C) + C \delta f_t \\ & - \langle (\partial_t x + [G^{-1} \partial_t G, x]), G^{-1} \delta G \rangle + \partial_t \langle x, G^{-1} \delta G \rangle], \end{aligned} \quad (22)$$

with the bilinear Casimir [83]

$$C = \frac{1}{2} \langle B, B \rangle = \frac{1}{2} \langle x, x \rangle. \quad (23)$$

The second term in (22) vanishes on shell since C is constant and f_t has a fixed zero mode, while the terms in

the second line vanish on shell or because they integrate to zero; thus, only the first term remains, which is integrable in field space, leading to the boundary action

$$I^{\text{tw}} = -\kappa \oint dt f_i C. \quad (24)$$

For future purposes, we define $f_i = \dot{f}$, where $f = t + \dots$ is a quasiperiodic function that has arbitrary Fourier modes but a fixed linear term. The full action

$$\Gamma^{\text{BF}}[B, \mathcal{A}] = I^{\text{BF}}[B, \mathcal{A}] - \kappa \oint dt \dot{f} C \quad (25)$$

has a well-defined variational principle, i.e., its first variation vanishes for all variations that preserve our boundary and on shell conditions (13)–(16).

Plugging the Euclidean version of the expression (17) for x into the bilinear Casimir (23) yields

$$C = -\frac{1}{(f)^2} (\mathcal{T} - \dot{f} x_0^E \mathcal{P} + i \dot{f} \dot{x}_0^E + i \dot{f} \ddot{x}_0^E), \quad (26)$$

where we used the relation $1/x_1^E = \dot{f}$. Defining additionally $\dot{g} := i x_0^E \dot{f}$, the boundary action (24) with the expression for the Casimir (26) is nearly our final result.

$$I^{\text{tw}} = \kappa \oint \frac{dt}{f} (\mathcal{T} + i \dot{g} \mathcal{P} + \ddot{g}). \quad (27)$$

The boundary action (27) depends functionally on f and g , both of which are boundary scalars, as evident from their transformation behavior under asymptotic symmetries, $\delta_\lambda f = \varepsilon \dot{f}$ and $\delta_\lambda g = \sigma + \varepsilon \dot{g}$. The latter also shows that g is a phase under $U(1)$ gauge transformations.

As in the JT case [79], we reparametrize the time coordinate along the boundary by a diffeomorphism $\tau := f(t)$, where τ is our new (Euclidean) time coordinate with period β , and introduce the inverse of f as a new field $h(\tau) := -f^{-1}(\tau)$. The other field, g , now also depends on τ and prime from now on means derivative with respect to τ . Implementing this diffeomorphism in the boundary action (27) establishes the boundary action (1) announced in the introduction, where prime means $\frac{d}{d\tau}$.

The action (1) is our main result and constitutes the analog of the Schwarzian. Since it has a geometric interpretation as group action for twisted warped coadjoint orbits, governed by the symmetries (7), we refer to it as “twisted warped action.” This is analogous to the interpretation of the Schwarzian action as group action for Virasoro coadjoint orbits [86,87]. We refer to [88] and references therein for more on these mathematical aspects.

VI. SOLUTIONS TO TWISTED WARPED THEORY

We study now classical solutions of the action (1) for constant representatives, $\mathcal{T} = \mathcal{T}_0$ and $\mathcal{P} = \mathcal{P}_0$. The Hamiltonian formulation involves three canonical pairs ($i = 1, 2, 3$),

$$I^{\text{tw}}[q_i, p_i] = -\kappa \int_0^\beta d\tau (p_i q_i' - p_1 p_2 - e^{q_1} p_3). \quad (28)$$

The relation to the original variables is $q_3 = \exp(i\mathcal{P}_0 h)$, $q_2 = g + ih\mathcal{T}_0/\mathcal{P}_0$ while all other canonical variables are of auxiliary nature to get rid of higher derivatives. The interaction term with the exponential in q_1 also appears in the Schwarzian theory; see Eq. (2.1) in [39]. The key difference is the kinetic term, p_1^2 for the Schwarzian and $p_1 p_2$ for the twisted warped Hamiltonian.

Solving the Hamiltonian equations of motion yields $q_3 = h_0 + h_1 e^{i\tau/\tau_0}$ and $q_2 = g_0 - i g_1 \tau + g_2 e^{i\tau/\tau_0}$. These solutions depend on six integration constants, $g_0, g_1, g_2, h_0, h_1, \tau_0$, the latter playing the role of the periodicity, $\tau_0 = \frac{\beta}{2\pi}$. The integration constants h_0 and g_0 are constant shifts, while h_1 and g_2 are amplitudes in front of oscillating terms. The remaining constant, g_1 , captures the nonperiodicity of q_2 and is responsible for the on shell action being nonzero, $I^{\text{tw}}[q_i, p_i]|_{\text{EOM}} = -2\pi\kappa g_1$.

VII. THERMODYNAMICS

Assuming g_1 is independent from temperature allows to deduce the entropy $S = -I^{\text{tw}}[q_i, p_i]|_{\text{EOM}} = 2\pi\kappa g_1$ from the on shell action. Inserting all our definitions we recover the well-known fact that entropy is given by the dilaton at the horizon [89],

$$S = 2\pi\kappa X|_{\text{horizon}}. \quad (29)$$

The result for entropy (29) can be derived along the lines of [7,43]. One aspect of this derivation is worth highlighting: the holonomy of a along the thermal cycle must belong to the center of our gauge group for regularity (in order to have contractible thermal cycles). Assuming a single cover, we find that this regularity condition relates to temperature $T = \beta^{-1}$ and charge

$$\mathcal{P}_0 = 2\pi T \quad (30)$$

while mass \mathcal{T}_0 remains arbitrary. The label “charge” is justified for \mathcal{P}_0 since the equations of motion imply $\mathcal{P}_0 = Y$ and Y is the $U(1)$ charge. The label “mass” is justified for \mathcal{T}_0 as it is the subleading term in the metric (6) and since the associated function \mathcal{T} transforms like a stress tensor (7b) in a twisted warped field theory [77].

A peculiar aspect of CGHS black hole thermodynamics is that the inverse specific heat (at fixed charge) vanishes,

$C^{-1} = \frac{1}{T} \frac{dT}{dS} \Big|_{\delta \mathcal{P}_0=0} = 0$, since the Hawking-Unruh temperature T trivially does not vary if the charge \mathcal{P}_0 is kept fixed due to the relation (30). This property is well known [90], but will be crucial for the scaling limit from complex SYK.

VIII. SCALING LIMIT FROM COMPLEX SYK

We turn now to the field theory side, starting with the complex SYK model [6,27,91–93]. The effective action governing the dynamics of the collective low temperature modes of complex SYK is given by (see [27] and Eq. (1.12) in [93])

$$I^{\text{cSYK}} = \frac{NK}{2} \int_0^\beta d\tau \left(g' + \frac{2\pi i \mathcal{E}}{\beta} h' \right)^2 - \frac{N\gamma}{4\pi^2} \int_0^\beta d\tau \left\{ \tan\left(\frac{\pi}{\beta} h\right); \tau \right\}, \quad (31)$$

where $\{f; \tau\} := f'''/f' - \frac{3}{2}(f''/f')^2$ is the Schwarzian derivative, N is the (large) number of complex fermions, NK is the zero temperature compressibility, $N\gamma$ is the specific heat at fixed charge, and \mathcal{E} is a spectral asymmetry parameter. The time-reparametrization field $h(\tau + \beta) = h(\tau) + \beta$ is quasiperiodic and the phase field $g(\tau)$, in the absence of winding, is periodic.

According to the thermodynamical discussion above, we are interested in the limit $N\gamma \rightarrow \infty$ in order to obtain our action (1) as limit from the complex SYK effective action (31). This is indeed possible by combining the actions (1) and (31) to the geometric action associated with the twisted warped Virasoro group, known as “warped Schwarzian” [88], $I^{\text{wSch}} = I^{\text{cSYK}} + I^{\text{tw}}$.

Starting from the effective action (31) and shifting g by [88] $g \rightarrow g - \frac{\kappa}{NK} (\log h' + \frac{2\pi i}{\beta} h)$ yields the action I^{wSch} with nonvanishing κ and shifted specific heat parameter $\hat{\gamma} = \gamma + \frac{36\pi^2 \kappa^2}{N^2 K}$. Thus, our boundary action (1) emerges by sending both $\hat{\gamma}$ and K to zero, while keeping fixed κ . At large N , this is achieved by the family of scaling limits,

$$\gamma = \gamma_0 N^a \quad K = -K_0 N^{-b} \quad \kappa = \frac{N^{1+\frac{a-b}{2}}}{6\pi} \sqrt{\gamma_0 K_0}. \quad (32)$$

The constants γ_0 and K_0 are independent from N and their product must be positive. The exponents $a > -1$, $b > 1$ lead to infinite specific heat and vanishing zero temperature

compressibility, respectively, in the large N limit. Two simple choices are $a = b = 2$, leading to $\kappa = \frac{N}{6\pi} \sqrt{\gamma_0 K_0}$, and $a = 0$, $b = 2$, leading to $\kappa = \frac{1}{6\pi} \sqrt{\gamma_0 K_0}$.

IX. CONCLUSIONS

We derived on the gravity side the boundary action (1) as a first step toward a two-dimensional model for flat space holography. We showed that the field theory side of our proposal for flat space holography emerges as a triple scaling limit of complex SYK: large N , large coupling (or small temperature) and large specific heat, while keeping fixed (with an adjustable scaling in N) the geometric mean of specific heat and zero temperature compressibility. As evident from (32), this geometric mean (up to a factor $N^{1+\frac{a-b}{2}}/(6\pi)$) is the coupling constant κ in (1).

Starting from our flat space holographic description, numerous further research avenues can now be pursued, inspired by corresponding SYK-related results or by generic aspects of two-dimensional dilaton gravity (see [94] for a review and [95] for a list of models). Not intending to do justice to the vast literature on these subjects we highlight just one intriguing aspect, namely, the role of chaos in flat space holography. By analogy to the AdS_2 case [11,96], we expect saturation of the chaos bound, i.e., a Lyapunov exponent given by $\lambda_L = 2\pi T$. It should be rewarding to verify this through explicit calculations.

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