

Neutron skin in CsI and low-energy effective weak mixing angle from COHERENT data

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Both the neutron skin thickness ΔR_{np} of atomic nuclei and the low-energy neutrino-nucleon (νN) interactions are of fundamental importance in nuclear and particle physics, astrophysics as well as new physics beyond the standard model (SM) but largely uncertain currently, and the coherent elastic neutrino-nucleus scattering (CE ν NS) provides a clean way to extract their information. New physics beyond the SM may cause effectively a shift of the SM weak mixing angle θ_W in low-energy νN interactions, leading to an effective weak mixing angle θ_W^* . By analyzing the CE ν NS data of the COHERENT experiment, we find that while a one-parameter fit to the COHERENT data by varying ΔR_{np} produces $\Delta R_{np}^{\text{CsI}} \simeq 0.68_{-1.13}^{+0.91}$ fm for CsI with an unrealistically large central value by fixing $\sin^2\theta_W^*$ at the low-energy SM value of $\sin^2\theta_W^{\text{SM}} = 0.23857$, a two-dimensional fit by varying ΔR_{np} and $\sin^2\theta_W^*$ leads to a strong positive correlation between ΔR_{np} and $\sin^2\theta_W^*$ with significantly smaller central values of $\Delta R_{np}^{\text{CsI}} \simeq 0.24_{-2.03}^{+2.30}$ fm and $\sin^2\theta_W^* = 0.21_{-0.10}^{+0.13}$. Although the uncertainty is too large to claim a determination of $\Delta R_{np}^{\text{CsI}}$ and $\sin^2\theta_W^*$, the present study suggests that the multidimensional fit is important in future analyses of high-precision CE ν NS data. The implication of the possible deviation of $\sin^2\theta_W^*$ from $\sin^2\theta_W^{\text{SM}}$ on new physics beyond the SM is also discussed.

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I. INTRODUCTION

The neutron skin thickness of atomic nuclei, defined as $\Delta R_{np} = R_n - R_p$ where $R_{n(p)} = \langle r_{n(p)}^2 \rangle^{1/2}$ is the neutron (proton) rms radius of the nucleus, provides a good probe of the equation of state (EOS) for isospin asymmetric nuclear matter [1–10], which is critically important due to its multifaceted roles in nuclear physics and astrophysics [11–14] as well as some issues of new physics beyond the standard model (SM) [15–19]. While the R_p can be measured precisely from electromagnetic processes (see, e.g., Refs. [20–22]), the R_n is largely uncertain since it is usually determined from strong processes, which is generally model dependent due to the complicated nonperturbative effects. This provides a strong motivation for the lead radius experiment (PREX) being performed at the Jefferson Laboratory to determine the R_n of ^{208}Pb to about 1% accuracy by measuring the parity-violating electroweak asymmetry in the elastic scattering of polarized electrons from ^{208}Pb [23]. The PREX Collaboration reported the first result of the parity violating weak neutral interaction measurement of the ΔR_{np} for ^{208}Pb ,

i.e., $\Delta R_{np}^{208} = 0.33_{-0.18}^{+0.16}$ fm [24] (see, also, Ref. [25]). The central value of 0.33 fm means a surprisingly large neutron skin thickness in ^{208}Pb although there is no compelling reason to rule out a such large value [26].

Recently, the COHERENT Collaboration [27] reported the first observation of the coherent elastic neutrino-nucleus scattering (CE ν NS) [28,29]. In Ref. [30], a value of the averaged ΔR_{np} of ^{133}Cs and ^{127}I , i.e., $\Delta R_{np}^{\text{CsI}} \simeq 0.7_{-1.1}^{+0.9}$ fm, is extracted from analyzing the COHERENT data. The extracted central value of $\Delta R_{np}^{\text{CsI}} \simeq 0.7$ fm is unrealistically large. To the best of our knowledge, $\Delta R_{np}^{\text{CsI}} \simeq 0.7$ fm is actually much larger than all the predictions of current nuclear models. Moreover, since ^{208}Pb is much more neutron-rich than ^{133}Cs and ^{127}I , the $\Delta R_{np}^{\text{CsI}}$ is expected to be smaller than the ΔR_{np}^{208} according to the neutron skin systematics [31,32], and thus $\Delta R_{np}^{\text{CsI}} \simeq 0.7$ fm is inconsistent with the PREX result. Although the uncertainty is too large to claim a determination of $\Delta R_{np}^{\text{CsI}}$, the best-fit value $\Delta R_{np}^{\text{CsI}} \simeq 0.7$ fm indicates the possibility of a unrealistically large neutron skin thickness. The possible inconsistency could be a hint of new physics in neutrino physics and this provides the main motivation of the present work.

We note that in Ref. [30], the $\Delta R_{np}^{\text{CsI}}$ is extracted from a one-parameter fit to the COHERENT data by varying $\Delta R_{np}^{\text{CsI}}$ with the low-energy weak mixing angle θ_W fixed at the SM value $\sin^2\theta_W^{\text{SM}} = 0.23865$ obtained in the modified minimal subtraction ($\overline{\text{MS}}$) renormalization scheme at near

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zero momentum transfer $Q = 0$ [33] (the newest value is $\sin^2\theta_W^{\text{SM}} = 0.23857(5)$ [34]). Experimentally, the precise determination of $\sin^2\theta_W$ at low Q^2 is an ongoing issue [35], and the atomic parity violation (APV) experiments offer the most precise results to date. For example, by measuring the $6s_{1/2} - 7s_{1/2}$ electric dipole transition in ^{133}Cs atom, a value of $\sin^2\theta_W = 0.2356(20)$ at $\langle Q \rangle \simeq 2.4$ MeV is obtained [36–38], which is smaller than $\sin^2\theta_W^{\text{SM}}$ by about 1.5σ . In the midenergy regime, the Qweak Collaboration reported the recent measurement on proton's weak charge and obtained $\sin^2\theta_W = 0.2383(11)$ at $Q = 0.158$ GeV [39], agreeing well with the SM prediction. On the other hand, the low-energy neutrino-nucleon (νN) interactions could involve new physics beyond the SM [35,40–47], which may cause effectively a shift of the SM weak mixing angle θ_W in the νN interactions, leading to a low-energy effective weak mixing angle θ_W^* . Any experimental constraints on θ_W^* would provide useful information on new physics beyond the SM.

In this work, we extract the values of $\Delta R_{np}^{\text{Csl}}$ and $\sin^2\theta_W^*$ using a two-dimensional (2D) fit to the COHERENT data by varying $\Delta R_{np}^{\text{Csl}}$ and $\sin^2\theta_W^*$. Compared to the results using one-parameter fit with $\sin^2\theta_W^*$ fixed at $\sin^2\theta_W^{\text{SM}}$, we find a strong positive correlation between ΔR_{np} and $\sin^2\theta_W^*$ with significantly smaller central values of $\Delta R_{np}^{\text{Csl}} \simeq 0.24_{-2.03}^{+2.30}$ fm and $\sin^2\theta_W^* = 0.21_{-0.10}^{+0.13}$ at $Q \simeq 0.05$ GeV (corresponding to the energy scale of COHERENT experiment), indicating that the $\sin^2\theta_W^*$ may play an important role in extracting neutron skin information from analyzing the CE ν NS data.

II. CE ν NS IN THE COHERENT EXPERIMENT

The differential cross section for coherent elastic neutrino-nucleus scattering has a straightforward SM prediction in the case with different proton and neutron distributions (form factors) in the nucleus. By neglecting the radiative corrections and axial contributions, the cross section can be expressed as [41,42,48–50]:

$$\frac{d\sigma}{dT}(E_\nu, T) = \frac{G_F^2 M}{2\pi} G_V^2 \left[1 - \frac{MT}{E_\nu^2} + \left(1 - \frac{T}{E_\nu} \right)^2 \right], \quad (1)$$

$$G_V = Zg_V^p F_p(q^2) + Ng_V^n F_n(q^2), \quad (2)$$

where G_F is the Fermi coupling constant, M is the nucleus mass, E_ν and T are neutrino energy and nuclear recoil kinetic energy, respectively. For a given E_ν , the corresponding T varies from 0 to $T^{\text{max}} = 2E_\nu^2/(M + 2E_\nu)$. The proton and neutron neutral current vector couplings are defined, respectively, as $g_V^p = \frac{1}{2} - 2\sin^2\theta_W$ and $g_V^n = -\frac{1}{2}$. The form factor $F_{n(p)}(q^2)$ encapsulate the neutron (proton) number density distribution in nuclei, where the momentum transfer q is

given by $q^2 = 2E_\nu^2 TM/(E_\nu^2 - E_\nu T) \simeq 2MT$ under the condition of $E_\nu \gg T$.

In the case of the COHERENT experiment, the measurement is performed using a CsI detector which is dominantly composed of ^{133}Cs and ^{127}I . The mass of a nucleus with $N(Z)$ neutrons (protons) is determined by its corresponding total binding energy (E_B) from $M = N \times m_n + Z \times m_p - E_B$ where $m_{n(p)}$ is the rest mass of neutrons (protons). The binding energies per nucleon are 8.40998 MeV and 8.44549 MeV [51] for isotopes ^{133}Cs and ^{127}I , respectively. As for their density distributions, in order to test the model dependence, two analytic nuclear form factors are adopted, namely, the symmetrized Fermi (SF) form factor and the Helm form factor, which are two very successful and well-tested forms of nuclear form factors for medium to heavy nuclei [30,52–55]. Both form factors are characterized by two parameters related to the nuclear radius and the surface thickness (diffuseness), respectively.

The SF form factor has the form (See, e.g., Ref. [55])

$$F_{\text{SF}}(q^2) = \frac{3}{qc[(qc)^2 + (\pi qa)^2]} \left[\frac{\pi qa}{\sinh(\pi qa)} \right] \times \left[\frac{\pi qa}{\tanh(\pi qa)} \sin(qc) - qc \cos(qc) \right], \quad (3)$$

and the corresponding rms radius is expressed as

$$R_{\text{SF}}^2 \equiv \langle r^2 \rangle = \frac{3}{5} c^2 + \frac{7}{5} (\pi a)^2. \quad (4)$$

where c is the half-density radius and a quantifies the surface thickness $t = 4a \ln 3$. Experimentally, the proton distribution has been determined precisely, and we take the same parameters for proton distribution as in Ref. [30], which are obtained by fitting the proton structure data of ^{133}Cs and ^{127}I measured in muonic atom spectroscopy, namely, $t_p = 2.30$ fm, $c_{p,\text{Cs}} = 5.6710 \pm 0.0001$ fm and $c_{p,\text{I}} = 5.5931 \pm 0.0001$ fm. The corresponding proton rms radii for ^{133}Cs and ^{127}I are $R_p^{\text{Cs}} = 4.804$ fm and $R_p^{\text{I}} = 4.749$ fm, respectively.

The Helm form factor is expressed as [52]

$$F_{\text{Helm}}(q^2) = 3 \frac{j_1(qR_0)}{qR_0} e^{-q^2 s^2/2}, \quad (5)$$

where $j_1(x)$ is the spherical Bessel function of order one, i.e., $j_1(x) = \sin(x)/x^2 - \cos(x)/x$. The rms radius is simply given by

$$R_{\text{Helm}}^2 \equiv \langle r^2 \rangle = \frac{3}{5} R_0^2 + 3s^2, \quad (6)$$

where R_0 is the box radius and s quantifies the surface thickness. Again, for the proton distributions in ^{133}Cs and ^{127}I , we use $s_p = 0.9$ fm following Ref. [30], which

was determined for the proton form factor of similar nuclei [56], and the $R_{0,p}$ is determined by the corresponding R_p .

For the parameters of the neutron distributions in ^{133}Cs and ^{127}I , they are essentially unknown. In these neutron-rich nuclei, in principle, the neutron distributions should be different from the proton distributions because of the charge difference, which means that the neutron distributions could have different radius parameters (c_n and $R_{0,n}$) and diffuseness (surface thickness) parameters (t_n and s_n) compared to the proton distributions. We will examine these effects in the following.

In the COHERENT experiment, the photoelectrons are counted to monitor the scattering events and extract the nuclear recoil energy, with approximately 1.17 photoelectrons expected per keV of nuclear recoil energy, denoted as $\zeta = 1.17 \text{ keV}^{-1}$ [27]. The number of event counts in a nuclear recoil energy bin $[T^i, T^{i+1}]$ can be obtained as

$$N_i^{\text{th}} = N_{\text{CsI}} \sum_{\nu_l} \sum_{\mathcal{N}=\text{Cs,I}} \int_{T^i}^{T^{i+1}} dT \mathcal{A}(\zeta T) \times \int_{E_\nu^{\min}}^{E_\nu^{\max}} dE_\nu \frac{dN_{\nu_l}}{dE_\nu} \frac{d\sigma_{\nu-\mathcal{N}}}{dT}, \quad (7)$$

where N_{CsI} is the number of CsI in the detector and is given by $N_A m_{\text{det}} / M_{\text{CsI}}$ with N_A being the Avogadro constant, $m_{\text{det}} = 14.57 \text{ kg}$ the detector mass and $M_{\text{CsI}} = 259.8 \text{ g/mol}$ the molar mass of CsI. The acceptance efficiency function $\mathcal{A}(x)$ is described by [57]

$$\mathcal{A}(x) = \frac{a}{1 + e^{-k(x-x_0)}} \Theta(x-5), \quad (8)$$

where the parameter values are taken as $a = 0.6655^{+0.0212}_{-0.0384}$, $k = 0.4942^{+0.0335}_{-0.0131}$ and $x_0 = 10.8507^{+0.1838}_{-0.3995}$, and the $\Theta(x)$ is a modified Heaviside step function defined as

$$\Theta(x-5) = \begin{cases} 0 & x < 5, \\ 0.5 & 5 \leq x < 6, \\ 1 & x \geq 6. \end{cases} \quad (9)$$

The value of E_ν^{\min} depends on T , and the E_ν^{\max} is related to the neutrino source. At the Spallation Neutron Source, the neutrino flux is generated from the stopped pion decays $\pi^+ \rightarrow \mu^+ + \nu_\mu$ as well as the subsequent muon decays $\mu^+ \rightarrow e^+ + \bar{\nu}_\mu + \nu_e$. The neutrino population has the following energy distributions [30,43]

$$\begin{aligned} \frac{dN_{\nu_\mu}}{dE_\nu} &= \eta \delta\left(E_\nu - \frac{m_\pi^2 - m_\mu^2}{2m_\pi}\right), \\ \frac{dN_{\bar{\nu}_\mu}}{dE_\nu} &= \eta \frac{64E_\nu^2}{m_\mu^3} \left(\frac{3}{4} - \frac{E_\nu}{m_\mu}\right), \\ \frac{dN_{\nu_e}}{dE_\nu} &= \eta \frac{192E_\nu^2}{m_\mu^3} \left(\frac{1}{2} - \frac{E_\nu}{m_\mu}\right), \end{aligned} \quad (10)$$

with $E_\nu^{\max} \leq m_\mu/2$. The normalization factor η is defined as $\eta = r N_{\text{POT}} / (4\pi L^2)$, where $r = 0.08$ is the averaged production rate of the decay-at-rest (DAR) neutrinos for each flavor per proton on target, $N_{\text{POT}} = 1.76 \times 10^{23}$ is the total number of protons delivered to the target and $L = 19.3 \text{ m}$ is the distance between the neutrino source and the CsI detector [27].

To evaluate the fitting quality on the COHERENT data in Fig. 3A of Ref. [27], following Ref. [30], we apply the following least-squares function with only the 12 energy bins from $i = 4$ to $i = 15$, i.e.,

$$\begin{aligned} \chi^2 &= \sum_{i=4}^{15} \left(\frac{N_i^{\text{exp}} - (1 + \alpha)N_i^{\text{th}} - (1 + \beta)B_i}{\sigma_i} \right)^2 \\ &+ \left(\frac{\alpha}{\sigma_\alpha} \right)^2 + \left(\frac{\beta}{\sigma_\beta} \right)^2. \end{aligned} \quad (11)$$

Here for each energy bin, the experimental number of events, denoted as N_i^{exp} , is generated from the C-AC differences, and B_i is the estimated beam-on background with only prompt neutrons included [27]. The $\sigma_i = \sqrt{N_i^{\text{exp}} + 2B_i^{\text{ss}} + B_i}$ is the statistical uncertainty where B_i^{ss} is the estimated steady-state background determined with AC data [27]. The α and β are the systematic parameters corresponding to the uncertainties on the signal rate and the beam-on background rate, respectively. The fractional uncertainties corresponding to 1- σ variation are $\sigma_\alpha = 0.28$ and $\sigma_\beta = 0.25$ [27]. All the experimental data are taken from the COHERENT release [57].

III. RESULTS AND DISCUSSIONS

In the present work for CE ν NS calculations, we replace the θ_W in Eq. (1) by θ_W^* to effectively consider the possible effects of new physics in νN interactions. We first assume that the neutron and proton distributions have the same diffuseness parameters (i.e., $t_n = t_p$ and $s_n = s_p$) and the value of $\sin^2 \theta_W^*$ is fixed at the SM value of $\sin^2 \theta_W^{\text{SM}} = 0.23857$, and then perform a one-parameter fit to the COHERENT data by varying R_n to extract the neutron rms radius R_n^{CsI} of CsI (^{133}Cs and ^{127}I are assumed to have equal R_n). Our calculations lead to $R_n = 5.46^{+0.91}_{-1.13} \text{ fm}$ with the Helm form factor and $R_n = 5.47^{+0.91}_{-1.13} \text{ fm}$ with the SF form factor. Our results thus

nicely confirm the value of $R_n = 5.5^{+0.9}_{-1.1}$ fm extracted in Ref. [30] with the same assumptions.

In addition, we explore the effects of the neutron diffuseness parameters. To this end, we perform a one-parameter fit to the COHERENT data by varying R_n^{CsI} with various fixed values of the diffuseness parameter while the effective weak mixing angle is fixed at $\sin^2\theta_W^* = \sin^2\theta_W^{\text{SM}}$. The results indicate that a variation of ± 0.02 fm for $\Delta R_{np}^{\text{CsI}}$ arises when s_n changes from 0.63 to 1.17 fm (corresponding to a variation of $\pm 30\%$ for $s_n = 0.9$ fm) in the Helm form factor. The same conclusion is obtained when the SF form factor is used. Therefore, compared to the obtained neutron skin thickness of $\Delta R_{np}^{\text{CsI}} \simeq 0.68^{+0.91}_{-1.13}$ fm, the effects of the neutron diffuseness parameters are indeed quite small, consistent with the statement in Ref. [30].

Now we turn to examining the effects of the low-energy effective weak mixing angle. The possible nonstandard running of $\sin^2\theta_W^*$ in low-energy regime is expected to influence the extraction of the neutron distribution from the low-energy CE ν NS experiments. The simultaneous precise determination of the neutron distribution and the low-energy $\sin^2\theta_W^*$ through CE ν NS experiments can (in)validate our knowledge of nuclear physics and neutrino physics. Hence, we perform a 2D fit to the COHERENT data by varying R_n and $\sin^2\theta_W^*$ using the Helm form factor with $s_n = s_p$. The resulting number of CE ν NS event counts as a function of the number of photoelectrons is shown in Fig. 1 while the corresponding χ^2 contours are displayed in Fig. 2.

For comparison, we also include in Fig. 1 the corresponding results from the COHERENT data, the similar 2D fit by using the SF form factor with $t_n = t_p$, and the

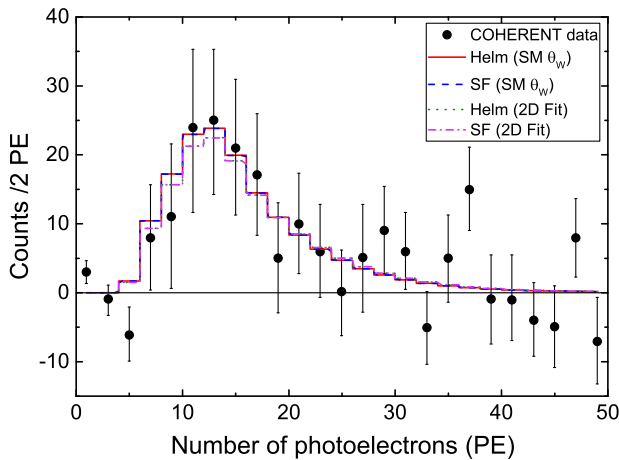


FIG. 1. The CE ν NS event counts as a function of the number of photoelectrons in the COHERENT experiment. The solid (dashed) line shows the results with best-fit neutron rms radius using the Helm (SF) form factor in the one-parameter fit when the $\sin^2\theta_W^*$ is fixed at the SM prediction. The dotted (dash-dotted) line shows the results with best-fit parameters in the 2D fit using the Helm (SF) form factor. Data are taken from Ref. [27].

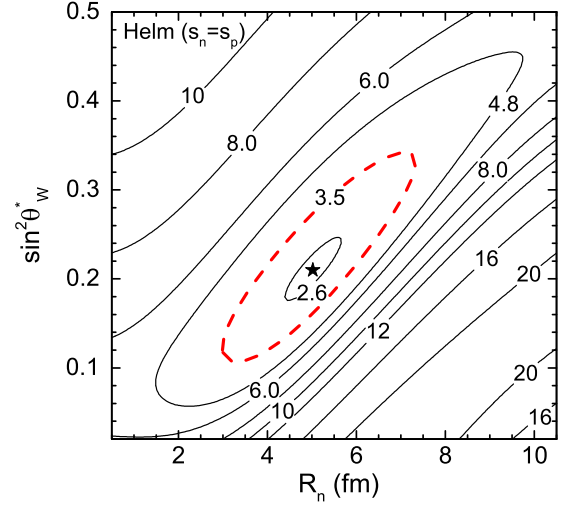


FIG. 2. The χ^2 contours in the plane of R_n vs $\sin^2\theta_W^*$ obtained from a 2D fit to the COHERENT data using the Helm form factor with $s_n = s_p$. The star marks the center values of $R_n = 5.02$ fm and $\sin^2\theta_W^* = 0.21$ at $\chi^2_{\text{min}} = 2.498$. The dashed curve corresponds to the contour at $\chi^2 = \chi^2_{\text{min}} + 1$.

one-parameter fit by varying R_n with fixed $\sin^2\theta_W^* = \sin^2\theta_W^{\text{SM}}$ using both the Helm and SF form factors. It is seen from Fig. 1 that for both one-parameter and 2D fits, the SF and Helm form factors produce almost identical results, indicating the independence of our results on the form of nuclear form factors. Furthermore, Fig. 1 indicates that compared to the one-parameter fit, the 2D fit predicts a fewer event counts in the region of 7–15 for the photoelectron number, leading to a decrease by $\sim 3.2\%$ for the number of total event counts.

From Fig. 2, one sees clearly that there exhibits a strong positive correlation between R_n and $\sin^2\theta_W^*$. Particularly interesting is that there exists favored center values for R_n and $\sin^2\theta_W^*$, i.e.,

$$R_n^{\text{Helm}} = 5.02^{+2.30}_{-2.03} \text{ fm}, \quad \sin^2\theta_W^* = 0.21^{+0.13}_{-0.10}. \quad (12)$$

We note that very similar results are obtained when the SF form factor is used. With the averaged rms radii of protons and neutrons in ^{133}Cs and ^{127}I , we then obtain the averaged neutron skin thickness of CsI as

$$\Delta R_{np}^{\text{CsI}} \simeq 0.24^{+2.30}_{-2.03} \text{ fm}. \quad (13)$$

The favored central value $\Delta R_{np}^{\text{CsI}} \simeq 0.24$ fm is significantly smaller than $\Delta R_{np}^{\text{CsI}} \simeq 0.68$ fm extracted from the one-parameter fit to the COHERENT data with fixed $\sin^2\theta_W^* = \sin^2\theta_W^{\text{SM}}$, indicating the importance of the $\sin^2\theta_W^*$ in the extraction of $\Delta R_{np}^{\text{CsI}}$ from CE ν NS.

Furthermore, we examine the effects of neutron diffuseness parameters using the 2D fit to the COHERENT data by varying R_n and $\sin^2\theta_W^*$ with s_n and t_n fixed at various

values. Our results indicate that the central value of $\Delta R_{np}^{\text{CSI}}$ varies by ± 0.03 fm (the corresponding R_n varies from 4.99 fm to 5.05 fm) when the value of s_n in the Helm form factor changes from 0.63 fm to 1.17 fm (corresponding to a variation of $\pm 30\%$ for $s_n = 0.9$ fm). Similarly, we find the central value of $\Delta R_{np}^{\text{CSI}}$ varies by ± 0.04 fm (the corresponding R_n varies from 4.99 fm to 5.07 fm) when the value of t_n in the SF form factor changes from 1.61 fm to 2.99 fm (corresponding to a variation of $\pm 30\%$ for $t_n = 2.3$ fm). Meanwhile, we note the central value variation of $\sin^2\theta_W^*$ is tiny, namely, from 0.209 to 0.211 when s_n (t_n) changes from 0.63 (1.61) fm to 1.17 (2.99) fm. The variation of $\pm(0.03-0.04)$ fm is appreciable compared to the central value $\Delta R_{np}^{\text{CSI}} \simeq 0.24$ fm, implying that one may extract useful information on the neutron diffuseness parameters in atomic nuclei from analyzing the future high-precise data of CE ν NS via a three-dimensional fit by varying $\sin^2\theta_W^*$, $\Delta R_{np}^{\text{CSI}}$ and the diffuseness parameters (s_n and t_n). This can help to address the interesting question about whether the neutron skin structure is really from the bulk radius difference or the surface diffuseness difference between the neutron and proton distributions in atomic nuclei [31,58,59]. Therefore, our results suggest that a multidimensional fit is important to extract the value of $\sin^2\theta_W^*$ and the neutron skin information including its size (i.e., $\Delta R_{np}^{\text{CSI}}$) and shape (e.g., s_n and t_n) in future analyses of high-precision CE ν NS data. Nevertheless, the extracted central value of $\Delta R_{np}^{\text{CSI}} \simeq 0.24$ fm with an uncertainty of $\pm(0.03-0.04)$ fm obtained in the present work is consistent with some carefully calibrated nuclear models (see, e.g., Refs. [26,30]).

On the other hand, a possible substantial deviation of $\sin^2\theta_W^*$ from $\sin^2\theta_W^{\text{SM}}$, i.e., $\Delta\sin^2\theta_W^* = -0.02857$, is obtained with the best-fit value of $\sin^2\theta_W^* = 0.21$. This deviation could be a hint of new physics beyond SM in neutrino physics. For example, one new physics scenario is to introduce the nonstandard interactions (NSIs) in the SM interactions, which has been widely discussed [40–44]. To make a rough estimate on the parameters in NSIs, we introduce an *ad hoc* nonstandard charge G_V^{NSI} to replace the G_V in Eq. (2), i.e.,

$$G_V^{\text{NSI}} = Zg_V^p F_p(q^2) + Ng_V^n F_n(q^2) + 3\delta_{\text{NSI}}[ZF_p(q^2) + NF_n(q^2)], \quad (14)$$

where $\delta_{\text{NSI}} = \epsilon_{\alpha\alpha}^{\mu V} = \epsilon_{\alpha\alpha}^{dV}$ ($\alpha = e, \mu, \tau$ represents the neutrino flavor) denotes the NSI parameters. Equation (14) can be obtained from the more general NSIs (see, e.g., Refs. [40–42,44]) by neglecting the flavor-changing couplings $\epsilon_{\alpha\beta}^{qV}$ ($\alpha \neq \beta$) and assuming that the new flavor-preserving couplings ($\epsilon_{\alpha\alpha}^{qV}$) are flavor symmetric for neutrinos and the first-generation quarks ($q = u, d$). Then one can estimate the value of δ_{NSI} as

$$\delta_{\text{NSI}} \simeq -\frac{2Z}{3A} \Delta\sin^2\theta_W^* = 0.008, \quad (15)$$

by assuming $F_p(q^2) \simeq F_n(q^2)$. These results indicate that the NSI contribution into the proton and neutron neutral current vector couplings is $3\delta_{\text{NSI}} = 0.024$, which is even larger than the SM proton coupling $g_V^p = \frac{1}{2} - 2\sin^2\theta_W^{\text{SM}} = 0.02286$.

Moreover, we would like to point out that the deviation of $\sin^2\theta_W^*$ from $\sin^2\theta_W^{\text{SM}}$ in neutrino physics can also potentially arise from the neutrino electromagnetic properties, e.g., the neutrino charge radius $\langle r_\nu^2 \rangle$ [44–46]. Furthermore, the deviation could be as well from the dark parity violation [35,47]. All these scenarios beyond the SM can effectively shift the low-energy weak mixing angle in νN interactions and worthy of further investigation with forthcoming more precise CE ν NS data in future. It will be also very interesting to check the similar effects in other weak neutral interaction measurements, e.g., APV and PREX.

Finally, it should be pointed out that the uncertainty of the extracted values for both $\Delta R_{np}^{\text{CSI}}$ and $\sin^2\theta_W^*$ is very large due to the poor statistics of the current COHERENT data, and this hinders us from claiming a determination of the $\Delta R_{np}^{\text{CSI}}$ and $\sin^2\theta_W^*$. Nevertheless, our results indicate that the $\Delta R_{np}^{\text{CSI}}$ is positively correlated with $\sin^2\theta_W^*$ and the best-fit values lead to the possibility of significantly smaller values of $\Delta R_{np}^{\text{CSI}}$ and $\sin^2\theta_W^*$ compared to the one-parameter fit to the COHERENT data with $\sin^2\theta_W^* = \sin^2\theta_W^{\text{SM}}$. The present work thus suggests that the $\sin^2\theta_W^*$ may play an important role in extracting neutron skin information from analyzing the CE ν NS data and a multidimensional fit is important in future analyses of high-precision CE ν NS data.

IV. SUMMARY AND OUTLOOK

We have demonstrated that the low-energy effective weak mixing angle θ_W^* plays an important role in the extraction of neutron skin information of atomic nuclei from the CE ν NS experiments. By analyzing the CE ν NS data of the COHERENT experiment, we have found that while a one-parameter fit to the COHERENT data produces $\Delta R_{np}^{\text{CSI}} \simeq 0.68_{-1.13}^{+0.91}$ fm with $\sin^2\theta_W^* = \sin^2\theta_W^{\text{SM}} = 0.23857$, a two-dimensional fit by varying ΔR_{np} and $\sin^2\theta_W^*$ leads to a strong positive correlation between ΔR_{np} and $\sin^2\theta_W^*$ with significantly smaller central values of $\Delta R_{np}^{\text{CSI}} \simeq 0.24_{-2.03}^{+2.30}$ fm and $\sin^2\theta_W^* = 0.21_{-0.10}^{+0.13}$. While the best-fit value $\Delta R_{np}^{\text{CSI}} \simeq 0.24$ fm seems to be reasonable, the substantial deviation of the best-fit value $\sin^2\theta_W^* = 0.21$ from $\sin^2\theta_W^{\text{SM}}$ could give a hint on new physics in ν -nucleon interactions.

Although the current large uncertainty does not allow us to claim a determination of the $\Delta R_{np}^{\text{CSI}}$ and $\sin^2\theta_W^*$, our present work suggests that a multidimensional fit is important to extract useful information on neutron skin information (including its size and shape) and the low-energy effective

$\sin^2 \theta_W^*$ from analyzing the high-precision data of future CE ν NS measurements. It will be also extremely interesting to explore the similar effects in other experiments of weak neutral interaction measurements.

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