Transport coefficients for multicomponent gas of hadrons using Chapman-Enskog method

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The transport coefficients of a multicomponent hadronic gas at zero and nonzero baryon chemical potential are calculated using the Chapman-Enskog method. The calculations are done within the framework of an S-matrix based interacting hadron resonance gas model. In this model, the phase shifts and cross sections are calculated using K-matrix formalism and where required, by parametrizing the experimental phase shifts. Using the energy dependence of the cross section, we find the temperature dependence of various transport coefficients such as shear viscosity, bulk viscosity, heat conductivity, and diffusion coefficient. We finally compare our results regarding various transport coefficients with previous results in the literature.

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I. INTRODUCTION

One of the important discoveries from experiments at the Relativistic Heavy Ion Collider (RHIC) and the Large Hadron Collider (LHC) in search of the quark-gluon plasma (QGP) is the fact that the deconfined quark-gluon matter behaves as an almost-perfect fluid [1–9]. The property that quantifies the "liquidness" of a liquid are its transport coefficients, for, e.g., the ratio of shear viscosity to its entropy density η_s/s or the ratio of bulk viscosity to its entropy density η_v/s . Experimentally measured elliptic flow, through the azimuthal correlation of produced particles with respect to the reaction plane is a sensitive probe for obtaining the transport coefficients. For example, the magnitude of elliptic flow depends quite sensitively on the shear viscosity of the QGP fluid, which is estimated to be around $\eta_s/s \approx 0.08 - 0.20$ during the hydrodynamic evolution [10-12]. These small values of η_s/s make the system (QGP), a near perfect fluid.

Theoretically, the value of the shear viscosity depends on the model under consideration. For example, at weak coupling the dimensionless ratio η_s/s is proportional to the ratio of mean free path to the mean spacing between the

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particles and weaker coupling means a larger value of this ratio [13–15]. For example, for a weakly coupled gas of gluons the $\eta_s/s \sim (\alpha_s^4 \log(1/\alpha_s))^{-1}$, where α_s is the QCD coupling constant. For $\alpha_s = 0.1$, the value of $\eta_s/s \sim 4 \times 10^3$. On the other hand, in a strongly coupled system there is efficient momentum transfer and the ratio η_s/s is significantly smaller [16]. For example, a lower bound of the ratio $\eta_s/s = 1/4\pi$ for strongly coupled field theories using the anti–de Sitter/conformal field theory (AdS/CFT) correspondence has been conjectured in Ref. [17]. In this work, we shall find that the ratio η_s/s values varies from weak coupling to strong coupling regime as one goes from lower to higher temperatures.

There are two other important reasons for studying the temperature dependence of transport coefficients. First, experimentally it has been observed [18,19] that η/s shows a minimum near the liquid-gas phase transition for different substances, this might help in studying QCD phase diagram. Such a minimum has also been observed in model calculation. For example in Ref. [20], it was shown for massless pions the ratio η_s/s diverges as temperature $T \rightarrow 0$, whereas in the quark gluon phase, at one loop order in α_s , the η_s/s is an increasing function of *T*. Second, it has been predicted that the ratio η_v/s should show a maximum near the phase transition [21–23]. For example in Ref. [24] the ratio η_v/s for massless pions goes to zero in the $T \rightarrow 0$ limit and also to zero in the quark gluon phase for asymptotically high *T* [25].

In this work, we calculate various transport coefficients of a hadronic gas consisting of baryon and meson octets namely $\pi, \eta, K, N, \Lambda, \Sigma, \Xi$. The corresponding resonances which appear as interactions among these hadronic

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constituents are handled using the K- matrix formalism [26,27]. The formalism that we use for calculating the equilibrium thermodynamic quantities like entropy density, enthalpy density, number density etc., is through the S-matrix based hadron resonance gas (HRG) model [27,28]. The cross sections that are used in the calculation of transport coefficients are calculated in the K-matrix formalism for all hadrons except for the nucleons, where we directly use the experimental phase shifts [29]. The calculations are done for a system with vanishing baryon chemical potential (μ_B) as well as for finite $\mu_B = 100$ MeV. The transport coefficients are obtained using Chapman-Enskog (CE) method developed in Refs. [30-32]. In this method the solution of the transport equation, i.e., the distribution function to be determined, is first written as an infinite series of Laguerre polynomials. With the help of this expansion, the transport equation could be transformed into an infinite set of linear, algebraic equations. From this infinite set of equations, a finite number of equations are taken and solved to get an approximate solution for the distribution function. This solution is used to compute the transport coefficients.

Pioneering work on transport coefficients in the CE method has been done in Ref. [33] for quark and gluon system and in Ref. [20], for various binary combinations in a system consisting of $\pi - K - N$ using experimental cross sections. Similarly, in Ref. [26] the calculations of η/s for a multicomponent system consisting of $\pi - K - N - \eta$ at vanishing μ_B in the K-matrix formalism, has been carried out but without including NN interaction. In Ref. [34], CE method was used for calculating η_s/s and η_v/s using UrQMD cross sections and in Ref. [35] using in-medium cross sections. The current work improves upon all these previous work to incorporate a larger spectrum of interacting hadronic states (7 stable hadrons +112 resonances) and also extends to finite chemical potential. By adding more stable hadrons into the mixture, one hopes that new channels of interaction (through resonance formation) could open up, which would relax the system to equilibrium quicker, than with fewer hadrons considered in earlier works. Also, the degeneracy of the system changes, which affects equilibrium quantities like entropy density and number density etc., and in turn also affects the dimensionless transport coefficient ratios. It is also interesting to compute transport coefficients at nonzero chemical potential, since finite baryon density, affects the concentration of various species interacting in the mixture and thus the overall weight coming from different channels, on the final value of transport coefficient. In regards to other formalism, e.g., in Refs. [22,36,37], which uses relaxation time approximation (RTA), the present formalism is better in the sense that small angle scattering is taken care of naturally, whereas RTA uses thermal averaged cross sections. Similarly, compared to models like ideal hadron resonance gas, excluded volume approach [38–42] which uses constant values of a cross section, the present formalism utilizes the energy dependence of cross sections to calculate the temperature dependence of transport coefficients. Calculations of shear viscosity has also been done using the Kubo formalism in transport models [43,44]. Our results on transport coefficients are in reasonable agreement with that from the transport models in the temperature range of T = 80-110 MeV.

The paper is organized as follows. In Sec. II, we describe the *K*-matrix formalism for calculating the scattering phase shifts and cross section. In Sec. III we have discussed the thermodynamics of interacting hadron gas using the formalism of Sec. II. Then in Sec. IV, we describe the CE method for calculating the transport coefficients for single, binary, and multicomponent system of hadrons at zero and finite μ_B . Finally, in Sec. V, we summarize our findings.

II. K-MATRIX FORMALISM

A theoretical way of calculating the attractive part of the phase shifts is to use the *K*-matrix formalism. In this section, we briefly discuss the *K*-matrix formalism.

For the process $ab \rightarrow cd$

$$S_{ab\to cd} = \langle cd|S|ab\rangle,\tag{1}$$

where *S* is the scattering matrix operator. The scattering amplitude for the process can be expressed in terms of the interaction matrix T^{l} as,

$$f(\sqrt{s},\theta) = \frac{1}{q_{ab}} \sum_{l} (2l+1)T^{l}P_{l}(\cos\theta), \qquad (2)$$

where $P_l(\cos \theta)$ are the Legendre polynomials for the angular momentum l and θ is the scattering angle in the center of mass frame. The cross section for the process can be given in terms of scattering amplitude,

$$\sigma(\sqrt{s},\theta) = |f(\sqrt{s},\theta)|^2.$$
(3)

The T matrix is related to the S-matrix by the following equation

$$S = I + 2iT, \tag{4}$$

where I is the unit matrix. Using the unitarity of S-matrix one can show

$$(T^{-1} + iI)^{\dagger} = T^{-1} + iI.$$
(5)

Therefore, one can define a Hermitian K matrix through

$$K^{-1} = T^{-1} + iI. (6)$$

The *K*-matrix formalism preserves the unitarity of *S*-matrix and neatly handles multiple resonances [45].

In addition to that, widths of the resonances are handled naturally in the above formalism. For overlapping resonances the *K*-matrix gives a good description of the phase shifts.

One can write real and imaginary part of the T-matrix in terms of K matrix as

$$\frac{\mathrm{Im}T}{\mathrm{Re}T} = K.$$
 (7)

In Ref. [26] the *K*-matrix formalism was used to study the shear viscosity of an interacting gas of hadrons. Recently in Refs. [27,28], *K*-matrix formalism is used to study the equation of state and susceptibilities pertaining to conserved charges.

For the process $ab \rightarrow R \rightarrow cd$, resonances appear as sum of poles in the *K*-matrix as $K_{ab\rightarrow cd}^{I,l}$

$$K_{ab\to cd}^{I,l} = \sum_{R} \frac{g_{R\to ab}(\sqrt{s})g_{R\to cd}(\sqrt{s})}{m_{R}^{2} - s},$$
(8)

where *a*, *b* and *c*, *d* are hadrons, *R* is the resonance with mass m_R . The sum over *R* is restricted to the addition of resonances which have the same spin *l* and isospin *I*. The residue functions are given by

$$g_{R \to ab}^2(\sqrt{s}) = m_R \Gamma_{R \to ab}(\sqrt{s}), \qquad (9)$$

where \sqrt{s} is the energy in the center of mass frame and $\Gamma_{R \to ab}(\sqrt{s})$ is the energy dependent partial decay widths. For the process $R \to ab$, decay width can be written as [45]

$$\Gamma_{R \to ab}(\sqrt{s}) = \Gamma_{R \to ab}^{0} \frac{m_{R}}{\sqrt{s}} \frac{q_{ab}}{q_{ab0}} (B^{l}(q_{ab}, q_{ab0}))^{2}, \quad (10)$$

where $\Gamma^0_{R \to ab}$ is the partial width of the pole of the resonance at half maximum for the channel $R \to ab$, $B^l(q_{ab}, q_{ab0})$ are the Blatt-Weisskopf barrier factors [45] which can be expressed in terms of daughter momentum q_{ab} and resonance momentum q_{ab0} for the orbital angular momentum *l*. The momentum q_{ab} in the last expression is given as

$$q_{ab}(\sqrt{s}) = \frac{1}{2\sqrt{s}}\sqrt{(s - (m_a + m_b)^2)(s - (m_a - m_b)^2)},$$
(11)

where m_a and m_b are the mass of hadrons *a* and *b* decaying from resonance *R*. In Eq. (10), $q_{ab0} = q_{ab}(m_R)$ is the resonance momentum at $\sqrt{s} = m_R$.

Therefore, to compute the *K*-matrix one needs the relevant masses and widths of resonances. Further, in partial wave decomposition, the *K*-matrix can be written in terms of phase shift δ_I^I as [45],

$$K = \tan \delta_1^I. \tag{12}$$

Since the *K*-matrix formalism is applicable only for attractive interaction, for the repulsive as well as NN interaction we have used the experimental data of phase shift [29]. In the present work we do not incorporate baryon-hyperon and hyperon-hyperon interactions.

III. THERMODYNAMICS OF INTERACTING HADRONS

The most natural way to incorporate interaction among a gas of hadrons is to use relativistic virial expansion [46,47] where the logarithm of the partition function can be separated into two parts, noninteracting (Z_0) and interacting (Z_{int}) parts, i.e.,

$$\ln Z = \ln Z_0 + \ln Z_{\rm int}.\tag{13}$$

The noninteracting part of the partition function can be written as

$$\ln Z_0 = \sum_h \frac{V g_h m_h^2 T}{2\pi^2} \sum_{j=1}^{\infty} (\pm 1)^{j-1} (z_h^j / j^2) K_2(j m_h \beta), \quad (14)$$

where *h* denotes the index of stable hadron, *V* is the volume of the system, g_h is the degeneracy, m_h is the mass of the hadron, $z_h = \exp(\mu_h \beta)$ is the fugacity, β is the inverse of the temperature (T), and K_2 is the modified Bessel function of the second kind. For the conserved quantities like baryon number, strangeness, and electric charge, μ_h can be written as $\mu_h = B_h \mu_B + S_h \mu_S + Q_h \mu_Q$. Here B_h , S_h , Q_h are respectively the baryon number, strangeness and electric charge of the hadron and the μ 's are the chemical potentials of the corresponding conserved charges. In Eq. (14), +(-) sign refer to bosons (fermions) and j = 1 term corresponds to the classical ideal gas.

The interacting part of Eq. (13) can be written as

$$\ln Z_{\rm int} = \sum_{i_1, i_2} z_1^{i_1} z_2^{i_2} b(i_1, i_2), \qquad (15)$$

where $b(i_1, i_2)$ is the virial coefficient defined as,

$$b(i_1, i_2) = \frac{V}{4\pi i} \int \frac{d^3 p}{(2\pi)^3} \int d\varepsilon \exp\left(-\beta (p^2 + \varepsilon^2)^{1/2}\right) \\ \times \left[A \left\{ S^{-1} \frac{\partial S}{\partial \varepsilon} - \frac{\partial S^{-1}}{\partial \varepsilon} S \right\} \right]_c.$$
(16)

In the above equation, the labels i_1 and i_2 refer to channel of the *S*-matrix which has initial state containing $i_1 + i_2$ particles. The symbol *A* denotes the symmetrization (antisymmetrization) operator for a system of bosons (fermions), the subscript *c* refers to trace over all the linked diagrams.



FIG. 1. Temperature dependence of normalized number density calculated in S-matrix formalism. Left panel shows the n/T^3 for the mesons and the right panel shows the same for the baryons. Closed symbols correspond to result at $\mu_B = 0$ MeV and open symbols at $\mu_B = 100$ MeV.

The lowest virial coefficient i.e., the second virial coefficient, $b_2 = b(i_1, i_2)/V$ as $V \to \infty$, corresponds to interaction between two hadrons $(i_1 = i_2 = 1)$. The higher order virial coefficients give interaction among many hadrons. In the present, work we will consider only up to the second virial coefficient.

The S-matrix can be expressed in terms of phase shifts δ_l^I as [46]

$$S(\varepsilon) = \sum_{l,l} (2l+1)(2l+1) \exp(2i\delta_l^l), \qquad (17)$$

where l and I denote angular momentum and isospin, respectively. Integrating Eq. (16) over the total momentum we get

$$b_{2} = \frac{1}{2\pi^{3}\beta} \int_{M}^{\infty} d\varepsilon \varepsilon^{2} K_{2}(\beta \varepsilon) \sum_{l,l} g_{h} \frac{\partial \delta_{l}^{l}(\varepsilon)}{\partial \varepsilon}.$$
 (18)

The factor $g_h = (2I + 1)(2l + 1)$ is the degeneracy factor, *M* is the invariant mass of the interacting pair at zero center of mass momentum. The prime over the summation sign denotes that for given *l* the sum over *I* is restricted to values consistent with statistics.

The Eq. (18) shows that the contribution arising from interaction to thermodynamic variable depends on the derivative of the phase shift. The positive values of derivative of phase shifts (attractive interactions) give positive contributions to the thermodynamical variables and negative value of derivative of phase shifts (repulsive interactions) give negative contributions.

Once we know the partition function [Eq. (13)] we can calculate various thermodynamic quantities like pressure, energy density, entropy density, number density etc. In Fig. 1, we show the scaled number density (n/T^3) as a function of temperature, using the interacting model described above, for mesons (π, K, η) and baryons $(N, \Lambda, \Sigma, \Xi)$ at $\mu_B = 0$ MeV and at $\mu_B = 100$ MeV.

IV. TRANSPORT COEFFICIENTS

The relativistic Boltzmann equation, describing the spacetime evolution of the phase space density f = f(x, p), where x is position and p is momentum, is given by [48],

$$p^{\mu}\partial_{\mu}f_1 = C[f, f]. \tag{19}$$

The collision term C[f, f], in the Boltzmann approximation, is given by,

$$C[f,f] = \frac{1}{2} \int \frac{d^3 p_2}{p_2^0} \frac{d^3 p_3}{p_3^0} \frac{d^3 p_4}{p_4^0} [f_3 f_4 (1+\theta f_1)(1+\theta f_2) - f_1 f_2 (1+\theta f_3)(1+\theta f_4)] W(p_3, p_4|p_1, p_2),$$
(20)

where p_1 , p_2 are momenta of incoming and p_3 , p_4 are momenta of outgoing particles, respectively. $W(p_3, p_4|p_1, p_2)$ is the transition rate in the collision process $p_1 + p_2 \Leftrightarrow p_3 + p_4$. The constant $\theta = \pm 1$ for bosons or fermions and 0 for classical Maxwellian particles. We shall employ the Chapman-Enskog method as discussed in Refs. [30–32] to linearize and solve the kinetic equation Eq. (20). We split the derivative operator ∂^{μ} into a timelike and spacelike part

$$\partial^{\mu} \to U^{\mu}D + \nabla^{\mu},$$
 (21)

where $D = U^{\nu}\partial_{\nu}$ and $\nabla^{\mu} = \Delta^{\mu\nu}\partial_{\nu}$ and $\Delta^{\mu\nu} = g^{\mu\nu} - U^{\mu}U^{\nu}$ is the projection operator. Here, U^{μ} is the hydrodynamic four velocity, as discussed in Ref. [30]. Taking $\theta = 0$, i.e., assuming the particles to be classical, we expand the distribution function f into an equilibrium part $f^{(0)}$ and a deviation $\epsilon f^{(1)}$, i.e.,

$$f = f^{(0)} + \epsilon f^{(1)}.$$
 (22)

To order ϵ , substituting Eq. (22) into the transport equation Eq. (19) gives

$$p^{\mu}U_{\mu}Df_{1}^{(0)} + p^{\mu}\nabla_{\mu}f_{1}^{(0)} = -f_{1}^{(0)}\mathcal{L}[\phi], \qquad (23)$$

where $\mathcal{L}[\phi]$ is the linearized collision operator [found from Eq. (20), using Eq. (22) and invoking the principle of detailed balance given as, $f_1^{(0)}f_2^{(0)} = f_3^{(0)}f_4^{(0)}$].

Hence,

$$\mathcal{L}[\phi] = \frac{1}{2} \int \frac{d^3 p_2}{p_2^0} \frac{d^3 p_3}{p_3^0} \frac{d^3 p_4}{p_4^0} f_2^{(0)}(\phi_1 + \phi_2 - \phi_3 - \phi_4) \\ \times W(p_3, p_4|p_1, p_2).$$
(24)

The ϕ_i is the ratio $f_i^{(1)}/f_1^{(0)}$. The equilibrium distribution functions $f_i^{(0)}$ are assumed to be Maxwell-Boltzmann type

$$f_i^{(0)} = \exp\left(\frac{\mu_i(x) - p_i^{\nu} U_{\nu}(x)}{T(x)}\right).$$
 (25)

To identify the functions $\mu(x)$, $U^{\mu}(x)$, and T(x) with the usual definitions of chemical potential, hydrodynamic velocity and temperature of the system, we demand that the particle density *n* and energy density *en* be determined solely by the local equilibrium distribution function in Eq. (25) as,

$$n = \int \frac{d^3 p}{(2\pi)^3 p^0} (p^{\mu} U_{\mu}) f^{(0)}, \qquad (26)$$

$$en = \int \frac{d^3p}{(2\pi)^3 p^0} (p^{\mu} U_{\mu})^2 f^{(0)}.$$
 (27)

The choice of distribution function given in Eq. (25) along with condition given in Eqs. (26) and (27) determines the set of independent variables T, μ, U^{ν} . The derivative of the distribution function $f^{(0)}$, then depends only on the above set of independent variables. Then one can express $Df^{(0)}$ as,

$$Df^{(0)} = \frac{\partial f^{(0)}}{\partial n} Dn + \frac{\partial f^{(0)}}{\partial T} DT + \frac{\partial f^{(0)}}{\partial U^{\mu}} DU^{\mu}$$
$$= \left[\frac{\partial \mu}{\partial n} Dn + \left(T^{2} \frac{\partial}{\partial T} \left(\frac{\mu}{T} \right) + p^{\mu} U_{\mu} \right) \right.$$
$$\times D(\log T) - p_{\mu} DU^{\mu} \left] \frac{f^{(0)}}{T}, \qquad (28)$$

and $\nabla^{\alpha} f^{(0)}$ as,

$$\nabla^{\alpha} f^{(0)} = \left[T \nabla^{\alpha} \left(\frac{\mu}{T} \right) + p^{\mu} U_{\mu} \nabla^{\alpha} \log T - p^{\mu} \nabla^{\alpha} U_{\mu} \right] \frac{f^{(0)}}{T},$$
(29)

expressed in terms of temperature *T*, density *n*, hydrodynamic four-velocity U^{μ} and the chemical potential μ . Multiplying Eq. (23) with $\int d^3p p^{\mu}/p^0$ and contracting with U^{μ} , gives [48],

$$DT = -(\gamma - 1)T\nabla_{\mu}U^{\mu} \tag{30}$$

where $\gamma = c_P/c_V$ is the ratio of heat capacities at constant pressure c_P and constant volume c_V . Similarly, on multiplying Eq. (23) with $\int d^3p p^{\mu}/p^0$ and contracting with projection operator $\Delta^{\mu\nu}$, gives the equation of motion

$$DU^{\mu} = \frac{1}{wn} \nabla^{\mu} P, \qquad (31)$$

where, wn is the enthalpy density, wn = en + P and P is the pressure [30]. Also, the continuity equation, e.g., given in Ref. [48]

$$Dn = -n\nabla_{\mu}U^{\mu}.$$
 (32)

can be used to express the time derivative of number density in terms of gradients of hydrodynamic velocity. Equations (30)–(32) are used in Eqs. (28) and (29), to express time derivative of T, n, and U^{μ} in terms of gradients of U^{μ} and P, respectively.

The expressions of $Df^{(0)}$ and $\nabla^{\mu} f^{(0)}$ given in Eqs. (28) and (29) can be substituted in the linearized transport equation Eq. (23). Thus, one can express the transport equation in terms of thermodynamics forces, whose components include scalar force, vectorial force and tensorial force, respectively. The scalar force can be expressed as the divergence of hydrodynamic velocity

$$X = -\nabla_{\mu} U^{\mu}, \qquad (33)$$

the vectorial force, due to temperature gradient and pressure gradient is given as

$$Y^{\mu} = \nabla^{\mu} \log T - \frac{1}{wn} \nabla^{\mu} P, \qquad (34)$$

and tensorial forces (traceless indicated by " $\langle \rangle$ "), due to gradient of hydrodynamic velocity is given as

$$\langle Z^{\mu\nu} \rangle = \frac{1}{2} \nabla^{\mu} U^{\nu} + \frac{1}{2} \nabla^{\nu} U^{\mu} - \frac{1}{3} \Delta^{\mu\nu} \nabla_{\alpha} U^{\alpha}.$$
(35)

In terms of these forces, the transport equation is then given as

$$QX - p_{\nu}(p^{\mu}U_{\mu} - w)Y^{\nu} + p_{\mu}p_{\nu}\langle Z^{\mu\nu}\rangle = T\mathcal{L}[\phi].$$
(36)

The quantity Q is defined as

$$Q = \left(\frac{4}{3} - \gamma\right) (p_{\mu}U^{\mu})^{2} + ((\gamma - 1)w - \gamma T)p^{\mu}U_{\mu} - \frac{m^{2}}{3},$$
(37)

where the relativistic version of Gibbs-Duhem relation [48] $T\nabla^{\mu}(\mu/T) = -w(\nabla^{\mu}T/T - \nabla^{\mu}P/wn)$ was used for the derivation of Eq. (36).

An equation similar to Eq. (36) can also be derived for a two component mixture with components labeled by subscripts 1 and 2. Here, we indicate the few differences pertaining to extension of derivation of Eq. (36) for binary mixtures. Interested readers may refer [31] for the complete derivation.

The analogous linearized transport equation for mixtures can be written as,

$$p_1^{\mu} U_{\mu} D f_1^{(0)} + p_1^{\mu} \nabla_{\mu} f_1^{(0)} = -f_1^{(0)} \sum_{k=1}^2 \mathcal{L}_{1k}[\phi]. \quad (38)$$

An equation similar to the Eq. (38) also holds for component 2. The right-hand side takes collisions of the form $1(2) + 1(2) \rightarrow 1(2) + 1(2)$ and $1 + 2 \rightarrow 1 + 2$ into account. The linearized operator is given by

$$\mathcal{L}_{1k}[\phi] = \left(1 - \frac{\delta_{1k}}{2}\right) \int \frac{d^3 p_2}{p_2^0} \frac{d^3 p_3}{p_3^0} \frac{d^3 p_4}{p_4^0} f_k^{(0)} \\ \times (\phi_1 + \phi_2 - \phi_3 - \phi_4) W_{1k}(p_3, p_4|p_1, p_2). \tag{39}$$

The factor $(1 - \delta_{1k}/2)$ takes into account the correct weighting for same or different species which interact in the scattering process.

However, an extra thermodynamics force called the diffusion force [31], given by

$$Y_1^{\mu} = (\nabla^{\mu} \mu_1)_{P,T} - (\nabla^{\mu} \mu_2)_{P,T} - \frac{w_1 - w_2}{wn} \nabla^{\mu} P, \qquad (40)$$

needs to be introduced when dealing with mixtures. Here, n is the particle density and w_i is the enthalpy per particle of component *i*. Further derivation of the transport equation in terms of thermodynamic forces proceed along lines similar to single component system and can be found in Ref. [31]. Here we state the final result analogous to Eq. (36) for component 1 as

$$Q_{1}X - p_{1}^{\nu}(p_{1}^{\mu}U_{\mu} - w_{1})Y_{\nu} - x_{2}p_{1}^{\mu}Y_{1\mu} + p_{1}^{\mu}p_{1}^{\nu}\langle Z_{\mu\nu}\rangle$$

= $T\sum_{k=1}^{2}\mathcal{L}_{1k}[\phi],$ (41)

where, $x_i = n_i/(n_1 + n_2)$ being the particle number density fraction. An equation similar to above, also holds for component 2. The linear equations given in, Eqs. (36)

and (41) are used in the later sections to derive explicit expressions for the transport coefficients.

A. Single component system

In the present section we derive the transport coefficients for a single component system as described by the transport equation given in Eq. (36).

The observation that thermodynamic forces X, Y^{μ} and $\langle Z^{\mu\nu} \rangle$ appear as linearly independent quantities in Eq. (36), enables us to write the function ϕ of Eq. (24) as

$$\phi = AX - B_{\mu}Y^{\mu} + C_{\mu\nu}\langle Z^{\mu\nu}\rangle, \qquad (42)$$

where the unknown coefficients A, B_{μ} , and $C_{\mu\nu}$ are still to be determined. The sign of B_{μ} is chosen in accordance with the sign of the vector force in Eq. (36). Inserting Eq. (42) into Eq. (36), the transport equation can be separated into three independent equations, given as

$$QX = T\mathcal{L}[AX] \tag{43}$$

$$-(p^{\mu}U_{\mu} - w)p_{\mu}Y^{\mu} = T\mathcal{L}[-B_{\mu}Y^{\mu}]$$
(44)

$$p_{\mu}p_{\nu}\langle Z^{\mu\nu}\rangle = T\mathcal{L}[C_{\mu\nu}Z^{\mu\nu}], \qquad (45)$$

where $\mathcal{L}[\phi]$ is the linearized collision operator, as defined in Eq. (24).

We next define the macroscopic dissipative quantities, such as the viscous pressure and the heat flow which are functions of ϕ . The viscous pressure is defined as [30]

$$\Pi = -\frac{1}{3} \int \frac{d^3 p}{p^0} \Delta_{\mu\nu} p^{\mu} p^{\nu} f^{(0)} \phi, \qquad (46)$$

the heat flow is defined as

$$I_{q}^{\mu} = \int \frac{d^{3}p}{p^{0}} \Delta^{\mu\alpha} p_{\alpha}(p^{\mu}U_{\mu} - w) f^{(0)}\phi, \qquad (47)$$

and the traceless viscous pressure is defined as

$$\langle \Pi^{\mu\nu} \rangle = \int \frac{d^3p}{p^0} \left(\Delta^{\mu}_{\alpha} \Delta^{\nu}_{\beta} - \frac{1}{3} \Delta_{\alpha\beta} \Delta^{\mu\nu} \right) p^{\alpha} p^{\beta} f^{(0)} \phi.$$
 (48)

The dissipative quantities can be written in a more transparent way using the following dimensionless inner product bracket notation

$$(F,G) = \frac{T}{n} \int \frac{d^3 p}{p^0} F(p) G(p) f^{(0)}$$

= $\frac{1}{4\pi^2 z^2 K_2(z) T^2} \int \frac{d^3 p}{p^0} F(p) G(p) e^{-\tau},$ (49)

where the quantities z = m/T and $\tau = p^{\mu}U_{\mu}/T$ have been used. Inserting the expression for function ϕ , given in Eq. (42) into the definitions of dissipative quantities defined in Eqs. (46)–(48), expresses these dissipative quantities in terms of bracket notation. Hence, the bulk viscous pressure is given as,

$$\Pi = -\frac{1}{3}nT(\pi^{\mu}\pi_{\mu}, A)X = \eta_{v}X,$$
(50)

such that $\pi^{\mu} = \Delta^{\mu\nu} p_{\nu}/T$. The heat flow is given as

$$I_q^{\mu} = -nT\left(\pi^{\mu}\left(\tau - \frac{w}{T}\right), B_{\nu}\right)Y^{\nu} = T\lambda_{\nu}^{\mu}Y^{\nu}, \quad (51)$$

and the shear viscous flow as

$$\langle \Pi^{\mu\nu} \rangle = nT(\langle \pi^{\mu}\pi^{\nu} \rangle, C_{\alpha\beta}) \langle Z^{\alpha\beta} \rangle = 2\eta_s \langle Z^{\mu\nu} \rangle.$$
 (52)

The quantities η_v , $\lambda = \Delta_{\mu\nu}\lambda^{\mu\nu}/3$ and η_s stand for the bulk (volume) viscosity, heat conductivity, and shear viscosity coefficients that appear as a constant of proportionality between thermodynamic forces and the dissipative quantities.

The technical details needed to compute the still unknown quantities A, B^{μ} , and $C^{\mu\nu}$ into a tractable form, using collision integrals is given in Ref. [30]. Here, we simply write the expressions that can be used for computational purposes. The bulk viscosity is given by

$$\eta_v = T \frac{\alpha_2^2}{a_{22}},\tag{53}$$

the heat conductivity is given by

$$\lambda = \frac{T}{3m} \frac{\beta_1^2}{b_{11}},\tag{54}$$

and the shear viscosity is given by

$$\eta_s = \frac{T}{10} \frac{\gamma_0^2}{c_{00}}.$$
 (55)

The definitions of symbols α_2 , β_1 , and γ_0 and the expression for the quantities a_{22} , b_{11} , and c_{00} are given in the Appendix.

In Fig. 2 we use the relations as given in Eqs. (53)–(55) to calculate various transport coefficients for single component gas of baryons and mesons. The differential cross sections that go into the expression of a_{22} , b_{11} , and c_{00} are calculated using *K*-matrix formalism described in Sec. II for π , *K* and η while for nucleons (*N*) differential cross section, we use the experimental phase-shift data from Ref. [29]. One must note that the temperature dependence of transport coefficients are highly dependent on the energy dependence of differential cross sections. This is because the transport coefficients, through the quantities



FIG. 2. Variation of bulk viscosity, shear viscosity and heat conductivity of the single component gas with temperature. The lower triangle correspond to the results of transport coefficient computed using current algebra cross sections [49].



FIG. 3. Energy dependence of the cross section for single component gas using *K*-matrix formalism (solid, dot and dashed lines) and (dot dashed) line for pions using current algebra (CA) cross sections [49].

 a_{22} , b_{11} , and c_{00} (as given in the Appendix) are inversely proportional to the interaction cross sections.

Figure 3 shows, the role of cross sections on the temperature dependence of transport coefficients. It can be seen from the comparision of the current algebra (CA) cross sections of massive pions [49] which increase with the center of mass energy to that of K-matrix cross section which shows various peaks corresponding to various resonances that occur in $\pi\pi$ interaction throughout the energy spectrum. This makes the transport coefficients as shown in Figs. 2(a)-(c) to decrease with T for current algebra and increase with T for K-matrix. Similarly, for nn interaction which has only a few resonances, the temperature dependence of transport coefficients show a dip at some given range of temperature, which can be alluded to the sharp rise in the cross sections at corresponding energies (shown in Fig. 3). Comparing the transport coefficients among various mesons, we find that transport coefficients of a gas of $\eta > K > \pi$. This is because the total cross section of $\pi > K > \eta$. For nucleons, the elastic cross section decreases with the center of mass energy, the same is reflected in the transport coefficients of nucleons at low T, where it drops even lower than for π s, but with increasing T, increases faster than for π s.

Finally, we clarify that our calculations of transport coefficients is limited in scattering energy within the region where resonance dominates. This constraint on scattering energy can be translated to the limitation on the temperature in the following way. Considering each collision takes place between particles in thermal equilibrium, the scattering energy squared *s* fluctuates around the average $\langle s \rangle$ with a standard deviation σ . The average and standard deviation can be defined as [50]



FIG. 4. Temperature dependence of maximum scattering energy $\sqrt{s_{\text{max}}}$ for single component gas of pions, kaons, and nucleons.

$$\langle s \rangle = \frac{\int d^3 p_1 d^3 p_2 s(p_1, p_2) f^{(0)}(p_1) f^{(0)}(p_2)}{\int d^3 p_1 d^3 p_2 f^{(0)}(p_1) f^{(0)}(p_2)}$$
(56)

$$\sigma = \sqrt{\langle s^2 \rangle - \langle s \rangle^2}.$$
 (57)

Thus, we can define the maximum of scattering energy squared as $s_{\text{max}}(T) = \langle s \rangle + \sigma$. The temperature dependence of $\sqrt{s_{\text{max}}}$ for a single component gas of pions, kaons and nucleons is shown in Fig. 4. The $\sqrt{s_{\text{max}}}$ for $\pi\pi$, *KK* and *NN*, determined from their resonance mass cutoff, are 2.01 GeV, 2.29 GeV, and 2.34 GeV respectively. From Fig. 4, the corresponding temperature T_{max} are found to be 318 MeV, 297 MeV, and 150 MeV respectively.

B. Binary component system

The equation needed to obtain the transport coefficients for a mixture of two component gas is given in Eq. (41). The trial function is a linear combination of thermodynamic forces, i.e.,

$$\phi_{k} = \left(A_{k}X - B_{k\mu}Y_{q}^{\mu} - \frac{1}{T}B_{1k}^{\mu}Y_{1\mu} + C_{k}^{\mu\nu}\langle Z_{\mu\nu}\rangle\right).$$
(58)

The only differences between the trial function for single component system Eq. (42) and ϕ_k of binary-component system is the diffusion force Y_1^{μ} . Substituting function ϕ_k in Eq. (41) gives us

$$Q_1 = T \sum_{k=1}^{2} \mathcal{L}_{1k}[A_1],$$
 (59)

$$-(p_1^{\mu}U_{\mu} - w_1)p_1^{\nu} = T\sum_{k=1}^2 \mathcal{L}_{1k}[-B_1^{\nu}], \qquad (60)$$

$$-x_2 p_1^{\nu} = T \sum_{k=1}^2 \mathcal{L}_{1k} \left[-\frac{1}{T} B_{1k}^{\mu} \right], \tag{61}$$

$$p_1^{\mu} p_1^{\nu} = T \sum_{k=1}^2 \mathcal{L}_{1k}[C_1^{\mu\nu}], \qquad (62)$$

where the factors A_1 , B_1^{μ} , B_{1k}^{μ} , and $C_1^{\mu\nu}$ are unknown functions that are determined later. The law relating the traceless viscous pressure tensor to the hydrodynamic velocity and the law relating the viscous pressure to the divergence of hydrodynamic velocity as in Eqs. (50) and (55) do not change for mixtures. However, the law relating the heat flow to the temperature and pressure gradient, as in Eq. (51) needs to modified as,

$$\bar{I}_{q}^{\mu} = l_{qq} X_{q}^{\mu} + l_{q1} X_{1}^{\mu}, \tag{63}$$

where X_q^{μ} is the generalized driving force of heat flow and X_1^{μ} is the diffusion driving force, which accounts for the flow due to gradients of different constituents of the system. The transport coefficients are defined as

$$l_{qq} = \lambda T = -\frac{T}{3} \sum_{k=1}^{2} x_k \left(\pi_k^{\mu} \left(\tau_k - \frac{w_k}{T} \right), B_k \pi_{\mu k} \right), \quad (64)$$

for the thermal conductivity and

$$l_{q1} = -\frac{1}{3} \sum_{k=1}^{2} x_k \left(\pi_k^{\mu} \left(\tau_k - \frac{w_k}{T} \right), B_{1k} \pi_{\mu k} \right), \quad (65)$$

for the Dufour coefficients. This coefficients accounts for the heat flow in the presence of density gradients in a mixture. The other new coefficient for a mixture is the diffusion flow given by

$$I_q^{\mu} = l_{11} X_q^{\mu} + l_{1q} X_1^{\mu}, \tag{66}$$

where the coefficient l_{1q} is equal to the Dufour coefficient l_{q1} . The second coefficient l_{11} is related to diffusion coefficient through the relation [31], $D_d = \frac{l_{11}T}{nx_1x_2}$. This is given as

$$l_{11} = -\frac{1}{3T} \sum_{k=1}^{2} \left(\delta_{1k} - x_1 \right) x_k (\pi_k^{\mu}, B_{1k} \pi_{\mu k}).$$
(67)

As in a single component system, the transport coefficients can be written in collision bracket form, the details of which can be found in Ref. [31] and in the Appendix. Here we write the expression which can be used for computational purposes. The bulk viscosity is given as

$$\eta_v = T \frac{\alpha_2^2}{a_{22}},\tag{68}$$

the shear viscosity is given as

$$\eta_s = \frac{T}{10\Delta_c} ((x_1\gamma_1)^2 c_{22} - 2x_1 x_2 \gamma_1 \gamma_2 c_{12} + (x_2\gamma_2)^2 c_{11}),$$
(69)

and the diffusion coefficient is given as

$$D_d = \frac{\rho T}{3n^2 m_1 m_2 c_1 c_2} \frac{\delta_2^2}{b_{22}}.$$
 (70)

The symbols and their relations to collision brackets are explained in the Appendix.

One should note that the expressions given in Eqs. (53)-(55) for single component system and Eqs. (68)-(70) for binary component system corresponds to the first nonvanishing approximation of the transport coefficients (by approximation, we mean that the unknown coefficients B^{μ} , $C^{\mu\nu}$, etc., are expanded using a infinite series of Laguerre polynomials truncated at some order). Except for bulk viscosity, the first approximation corresponds to first nonvanishing value. For bulk viscosity, the nonvanishing value happens to be the third order approximation for single component system and second order approximation for binary component system. Thus, bulk viscosity for binary mixtures in the second order approximation calculated in this work depends only on the interaction among dissimilar species. The coefficient of shear viscosity, on the other hand, depends on c_{12} , c_{11} and c_{22} where, c_{12} describes the interaction between dissimilar species and c_{11}, c_{22} , describe the interaction among similar species [see Eqs. (A13)–(A15)].

The resulting transport coefficients for various binary mixtures are shown in Fig. 5. We have found both shear and bulk viscosities of the mixtures of two species lie in between the transport coefficients of the individual species. The dip seen in the shear viscosities of πN and KN can be attributed to resonances that appear in πN and KNinteraction at the relevant energies which leads to an increase in the cross section and thus lowering the value of shear viscosity. Similarly, we show the diffusion coefficient of various binary components in Fig. 5(c) which depends on the density gradients in a mixture. We find that KN system has largest diffusion coefficient at smaller temperatures and πN system the lowest, but with increasing temperature, the coefficient for KN system, shows a sharp decrease in its value and the πK system shows a minimum. The open symbols in Fig. 5 correspond to transport coefficients at $\mu_B = 100$ MeV. In the CE approximation μ_{B} enters implicitly in the expressions of transport coefficients via concentration or number densities of various reacting mixtures. The number densities were calculated



FIG. 5. Temperature dependence of shear viscosity, bulk viscosity, and the diffusion coefficient of the binary gas mixture. Close and open symbols correspond to the results at $\mu_B = 0$ and $\mu_B = 100$ MeV, respectively.

using virial expansion that was described in Sec. II and are themselves function of temperature. We find that values of bulk viscosities are larger for large μ_B but gradually asymptotes towards $\mu_B = 0$ MeV value, while shear viscosities values are smaller for large μ_B and gradually asymptotes toward the $\mu_B = 0$ MeV values. The diffusion coefficient are mostly unaffected by the value of μ_B considered in the work.

C. Multicomponent system

The derivation of transport coefficients for the multicomponent system follows the same line of reasoning as in the case of the single and binary component system. The transport coefficient can be expressed transparently using the bracket notation which can be found in Refs. [31,32]. Here, we only give the final expressions which can used for computational purposes. The bulk viscosity of a N component gas can be written as

$$\eta_v = n^2 T \sum_{k=1}^N \sum_{l=1}^N a_k a_l a_{kl},$$
(71)

while the coefficients a_k satisfy the linear equations

$$\sum_{l=1}^{N} a_{kl} a_l = \frac{\alpha_k}{n},\tag{72}$$

and the shear viscosity can be written as

$$\eta_s = \frac{T^3 \rho^2}{10} \sum_{k=1}^N \sum_{l=1}^N c_k c_l c_{kl}, \tag{73}$$

and the coefficients c_1 are solutions of

$$\sum_{l=1}^{N} c_l c_{kl} = \frac{\gamma_k}{\rho T} = \gamma_k^*.$$
(74)

In this work the Eq. (74) for the multicomponent system can be written as

$$\begin{pmatrix} c_{\pi} \\ c_{K} \\ c_{N} \\ c_{N} \\ c_{\eta} \\ c_{\Lambda} \\ c_{\Sigma} \\ c_{\Xi} \end{pmatrix} \begin{pmatrix} c_{\pi\pi} & c_{\pi K} & c_{\pi N} & c_{\pi\eta} & c_{\pi\Lambda} & c_{\pi\Sigma} & c_{\pi\Xi} \\ c_{\pi K} & c_{KK} & c_{KN} & 0 & 0 & c_{K\Sigma} & 0 \\ c_{\pi N} & c_{KN} & c_{NN} & c_{\eta N} & 0 & 0 & 0 \\ c_{\pi \eta} & 0 & c_{\eta N} & c_{\eta \eta} & c_{\eta \Lambda} & c_{\eta\Sigma} & 0 \\ c_{\pi\Lambda} & 0 & 0 & c_{\eta\Lambda} & c_{\Lambda\Lambda} & 0 & 0 \\ c_{\pi\Sigma} & c_{K\Sigma} & 0 & c_{\eta\Sigma} & 0 & c_{\Sigma\Sigma} & 0 \\ c_{\pi\Xi} & 0 & 0 & 0 & 0 & 0 & c_{\Xi\Xi} \end{pmatrix} = \begin{pmatrix} \gamma_{\pi}^{*} \\ \gamma_{K}^{*} \\ \gamma_{N}^{*} \\ \gamma_{N}^{*} \\ \gamma_{\Lambda}^{*} \\ \gamma_{\Sigma}^{*} \\ \gamma_{\Xi}^{*} \end{pmatrix},$$
(75)

and similarly for Eq. (72). The coefficients c_{kl} and a_{kl} depend on the scattering cross section of the given channel k and l and the expressions in terms of collision integrals are given in the Appendix [see Eqs. (A18)–(A20)]. The zeros in c_{kl} occur, when we do not have a resonance decaying in a channel kl.

The result of transport coefficients (η_v, η_s) , for various multi-channel processes is shown in Figs. 6(a), 6(c) at $\mu_B = 0$ MeV and Fig. 6(b), 6(d) at $\mu_B = 100$ MeV.

We find that bulk viscosity turns out to be additive for a mixture of hadrons, in contrast to the shear viscosity, which decreases with the increase in number of components. This also explains why in RTA, for shear viscosities one should not add the relaxation time but the inverse of relaxation time for a multicomponent system. The decrease in shear viscosities due to the increase in the components of the reacting mixture is evident, since it opens additional channels for reactions to occur and thus the overall cross section of the system. Comparing the result of η_v at $\mu_B = 0$ MeV with that at $\mu_B = 100$ MeV, we find that the values of η_v are larger at large μ_B . Similarly, we notice that at low T the shear viscosities at finite μ_B is slightly lower than at zero μ_B . However, with increasing temperature, the value of shear viscosity at finite μ_B overshoots that at zero μ_B . This can be understood, since at large T contributions from heavier baryonic states which have smaller cross section increases and thus increases the viscosity. At lower μ_B their concentration is smaller, hence their effect is not noticeable but increasing μ_B increases their concentration (the cross section remains the same) and hence their effect on viscosity also increases.



FIG. 6. Temperature dependence of bulk viscosity, shear viscosity at $\mu_B = 0$ MeV and $\mu_B = 100$ MeV for multicomponent gas of hadrons.



FIG. 7. Variation of normalized bulk and shear viscosity for the multicomponent hadronic gas. The black solid line is the value of η_v/s and η_s/s at $\mu_B = 0$ MeV and the black dotted line at $\mu_B = 100$ MeV. The red dashed line is AdS/CFT bound for η_s/s [55] and from a nonconformal model [54] for η_v/s . Other symbols are the results of transport coefficients, at $\mu_B = 0$ MeV, previously reported in the literature [26,34,38,40,51].

The variation of η_v/s and η_s/s as a function of temperature is shown in Figs. 7(a)–(b). Our results of η_v/s is an increasing function of T for T < 150 MeV and decreasing for T > 150 MeV. At $\mu_B = 100$ MeV, we find the magnitude of the peak, seen in η_v/s is larger than at $\mu_B = 0$ MeV. Similarly, we find that η_s/s decreases with temperature consistent with previous results in this regard [26,34,38,40,51]. However, we find that the result of η_s/s at $\mu_B = 0$ violates the AdS/CFT bound around a temperature of T = 160 MeV, while the result of $\mu_B = 100$ MeV remains above the bound and asymptotically approaches it at higher temperatures. Of course, the temperature where the violation of the AdS/CFT bound occurs, is in between the deconfinement temperature which is around $T \approx$ 155–165 MeV [52,53], where our model should break down. It is also interesting to note that peak in the ratio η_v/s happens to be around the same temperature where the ratio η_s/s violates the AdS/CFT bound. It is also important to note that recently the temperature dependence of bulk viscosity has also been calculated in nonconformal field theories in the context of heavy ion collisions [54]. A comparison of η_v/s with one such model from Ref. [54] is shown in Fig. 7(a), with the minimum of the potential ϕ_M set at $\phi_M = 2$. It is interesting to see that the nonconformal model compares well with our results.

Let us now discuss the comparison of our result with the calculations that has been previously reported in the literature at $\mu_B = 0$ MeV. In EVHRG (excluded volume HRG model), η_v/s monotonically decreases as a function of temperature *T* in contrast to our results which shows a peak structure and further one can note that magnitude of η_v/s in the EVHRG model is a factor of ten more than our results. The first reason for this is that, the calculation of η_v/s is carried out using RTA [40], in the EVHRG model using momentum independent relaxation times which is

quantitatively different from that of CE method used in the current work. The difference in temperature variation can be attributed to use of constant cross section in the EVHRG model calculations compared to energy dependent cross sections used in our work. Moroz [34] uses cross sections from the UrQMD model, including elastic plus resonance processes calculated in the CE approximation. The η_v/s result from Moroz calculation is qualitatively and quantitatively similar to our calculations. Some discrepancies are still present because of the use of some constant cross sections to describe nonresonant interaction in Moroz's calculation.

The η_s/s calculation in EVHRG model [38] is done assuming all hadrons have the same hard-core radius r = 0.5 fm. Apart from the fact that the value of r used is model dependent, one must note that, they also assume that the shear viscosity is additive for a mixture of hadrons, contrary to our results. Although η_s/s decreases with temperature, but the slope is less steeper than our calculation. This is because in Ref. [38] both η_s and s increase, as degeneracies increase. However, in our case η decreases and s increases as degeneracies increase. Both this feature make the slope of η_s/s steeper than Ref. [38]. Wiranata *et al.* [26] used K-matrix formalism for calculating η_s/s in a hadronic gas consisting of $\pi - K - N - \eta$. Their result is around six times larger than ours at low T and about two times larger in high T. The discrepancies between the two results are first, due to the fact that we have used a larger spectrum of interacting hadrons and resonances. Second, and an important difference is that Ref. [26] did not include the NN mutual interaction, since their cross section were solely using K- matrix formalism, where as we parametrize NN experimental phase shifts to calculate the differential cross section. Owing, to the fact that NN cross section are larger at small \sqrt{s} as has been previously discussed, their contribution to transport coefficients is quite different and dramatic than other resonant interaction present in Kmatrix formalism. SMASH (simulating many accelerated strongly interacting hadrons) [51], which is a transport code, uses Green-Kubo formalism to calculate η_s/s for hadronic gas mixture. One of the common features between our model and SMASH is the treatment of interactions through resonances, which have a non-zero lifetime. Our result of η/s is in good agreement with SMASH within temperature range of 80–110 MeV. But after $T \sim 120$ MeV, we find that the SMASH result saturates and forms a plateau at higher temperature. The same trend is also seen in other transport codes, e.g., UrQMD [43]. The crucial difference between our approach and SMASH is that, SMASH utilizes a feedback loop between the relaxation time and resonance lifetimes whereas our approach does not [44].

V. SUMMARY

In summary, we have calculated the transport coefficients for a multicomponent hadronic gas. The thermodynamic quantities are calculated using the *S*-matrix based hadron resonance gas model. The phase shifts required for the calculation of *S*-matrix was calculated using the *K*matrix formalism for all hadrons except for nucleons, for which we directly parametrize the experimental phase shifts. The transport coefficients were calculated using the Chapman-Enskog method. Such a method utilizes the energy dependence of cross sections to calculate the temperature dependence of transport coefficients.

We start with various single component gas systems and gradually add different species of hadrons to finally form a multicomponent gas mixture. We found that adding new species into the mixture, opens up new channels for interaction to occur, which leads to an increase in cross section and thus reducing the shear viscosity. Similarly, we calculate the transport coefficients at zero and nonzero μ_B . We found that increasing μ_B increase the contribution of higher mass baryons, which have a smaller cross section, in the transport coefficients. This leads to the increase in the value of η/s at higher temperatures. Interestingly, we found that at the temperature around $T \approx 160$ MeV, the ratio η_v/s shows a maximum and around the same temperature, the η_s/s starts violating the AdS/CFT bound. A maximum in η_v/s is a signature of crossover transition, that has been seen previously in molecular gases [56]. Similarly, the violation of AdS/CFT bound may signal the breakdown of a simple model like the HRG and that the nonperturbative nature of physics in this regime. However, increasing μ_B , evades such a violation of the bound to larger temperature. Finally, we compute and compare the ratio of η_s/s and η_v/s , with other models in the literature. Our calculation shows qualitatively similar features, with models that use energy dependent cross section in the relevant temperature ranges. It is also interesting to see that a model which assumes the hadrons to be gases with interaction governed by S-matrix elements, which are basically resonances, capture the essential physics of transport coefficients in the $T - \mu_B$ plane.

A few future directions for this work. The crucial assumption that has been done in this work, is the use of Maxwell-Boltzmann (MB) distribution function, which may not be valid for large chemical potentials. Then, one needs to solve the full quantum Boltzmann equation in the CE method. In that case, the polynomials (the Laguerre polynomials in the case of MB) satisfying them are not known, and we have to find them order by order, as has been done in Ref. [50]. Another important direction would be to include a feedback mechanism between the relaxation time and resonance lifetimes as is done in transport codes [51]. For example, in this work we have considered resonances like ρ , Δ , etc., as unstable particles, which although contribute to the cross section, by themselves are not a part of the mixture. This is only valid, if the lifetime of resonance is shorter than mean free time of the system. However, if the resonance lifetime is comparable or larger to the mean free time of the system, interaction can only occur until the resonance decays. Thus, the relaxation time in such cases is limited by the lifetime of the resonance and not by the mean free time.

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APPENDIX: DEFINITIONS OF VARIOUS SYMBOLS USED IN THE TEXT

In the following Appendix, we define the various symbols and expressions that were used in the main text.

For single component system the symbols α_2 , β_1 , and γ_0 are defined as

$$\alpha_2 = \frac{5w}{T} - 3\gamma \left(1 + \frac{w}{T}\right),\tag{A1}$$

$$\beta_1 = \frac{3\gamma}{\gamma - 1},\tag{A2}$$

$$\gamma_0 = \frac{10w}{T},\tag{A3}$$

where $\gamma = c_p/c_v$. The quantities a_{22} , b_{11} , and c_{00} are defined in terms of relativistic omega integrals, $\omega_i^{(j)}$

$$a_{22} = 2\omega_0^{(2)},\tag{A4}$$

$$b_{11} = 8(\omega_1^{(2)} + z^{-1}\omega_0^{(2)}), \tag{A5}$$

$$c_{00} = 16 \left(\omega_2^{(2)} - z^{-1} \omega_1^{(2)} + \frac{1}{3} z^{-2} \omega_0^{(2)} \right), \quad (A6)$$

where the definitions of relativistic omega integrals are given in [30] and can be written as

$$\omega_i^{(s)}(z) = \left(\frac{2\pi z^3}{K_2^2(z)}\right) \int_0^\infty d\psi \sinh^7 \psi \cosh^i \psi K_j(2z \cosh \psi)$$
$$\times \int_0^\pi d\theta \sin \theta \sigma(\psi, \theta) (1 - \cos^s \theta)$$
$$j = \frac{5}{2} + \frac{1}{2} (-1)^i, \quad i = 0, \pm 1, \pm 2, \dots, \quad s = 2, 4, 6, \dots$$
(A7)

where $\sigma(\psi, \theta)$ is the differential cross section for interaction between two identical particles, expressed through the quantities ψ and angle θ between the initial and final hadrons defined as

$$\sinh\psi = \frac{\sqrt{(p_1 - p_2)^2}}{2m}, \quad \cosh\psi = \frac{\sqrt{(p_1 + p_2)^2}}{2m}$$
 (A8)

where p_1 and p_2 are the initial four-momenta of the two colliding hadrons.

For binary component system the symbols α_i , δ_i , and γ_i , where i = 1, 2 are defined as

$$\alpha_i = x_i \frac{\gamma_{(i)} - \gamma}{\gamma_{(i)} - 1},\tag{A9}$$

$$\delta_i = (-1)^i 3c_1 c_2, \tag{A10}$$

$$\gamma_i = -10c_i h_i, \tag{A11}$$

where $h_i = K_3(z_i)/K_2(z_i)$ is the specific enthalpy of species *i*, $c_i = \rho_i/\rho$ is the mass fraction of species *i*. ρ_i is the mass density, which is mass times the number density of species *i* and ρ is the total mass density. Similarly $x_i = n_i/n$ is the number density fractions of species *i*, where n_i is particle number density of species *i* and *n* is the total number density. The quantity $\gamma_{(i)} = c_{p,i}/c_{v,i}$ is the ratio of specific heats of species *i*. The quantities a_{ii} , c_{ij} , b_{ii} , and Δ_c are defined in terms of relativistic omega integrals, $\omega_{iikl}^{(m)}(\sigma_{uv})$

$$a_{22} = \frac{16\rho_1\rho_2}{M^2n^2}\omega_{1200}^{(1)}(\sigma_{12}),\tag{A12}$$

$$c_{12} = \frac{32\rho^2 x_1^2 x_2^2}{3M^2 n^2 x_1 x_2} (-10z_1 z_2 \zeta^{-1} Z^{-1} \omega_{1211}^{(1)}(\sigma_{12}) - 10z_1 z_2 \zeta^{-1} Z^{-2} \omega_{1311}^{(1)}(\sigma_{12}) + 3\omega_{2100}^{(2)}(\sigma_{12}) - 3Z^{-1} \omega_{2200}^{(2)}(\sigma_{12}) + Z^{-2} \omega_{2300}^{(2)}(\sigma_{12})),$$
(A13)

$$c_{11} = c_{00}(z_1) + \frac{32\rho^2 x_1^2 x_2^2}{3M^2 n^2 x_1 x_2} (10z_1^2 \zeta^{-1} Z^{-1} \omega_{1220}^{(1)}(\sigma_{12}) + 10z_1^2 \zeta^{-1} Z^{-2} \omega_{1320}^{(1)}(\sigma_{12}) + 3\omega_{2100}^{(2)}(\sigma_{12}) - 3Z^{-1} \omega_{2200}^{(2)}(\sigma_{12}) + Z^{-2} \omega_{2300}^{(2)}(\sigma_{12})),$$
(A14)

$$c_{22} = c_{00}(z_2) + \frac{32\rho^2 x_1^2 x_2^2}{3M^2 n^2 x_1 x_2} (10z_2^2 \zeta^{-1} Z^{-1} \omega_{1202}^{(1)}(\sigma_{12}) + 10z_2^2 \zeta^{-1} Z^{-2} \omega_{1302}^{(1)}(\sigma_{12}) + 3\omega_{2100}^{(2)}(\sigma_{12}) - 3Z^{-1} \omega_{2200}^{(2)}(\sigma_{12}) + Z^{-2} \omega_{2300}^{(2)}(\sigma_{12})),$$
(A15)

$$\Delta_c = c_{11}c_{22} - c_{12}^2, \tag{A16}$$

$$b_{22} = \frac{8\rho c_1 c_2}{Mn} \left(2\omega_{1100}^{(1)}(\sigma_{12}) - 3Z^{-1}\omega_{1200}^{(2)}(\sigma_{12}) \right), \quad (A17)$$

where σ_{uv} is the cross section between particles u and v. The coefficients $c_{00}(z_k)$ accounts for contribution from interaction between identical species of type k as given in Eq. (A5). The reduced mass μ is given as $\mu = m_1 m_2/(m_1 + m_2)$. The abbreviations Z and ζ are given as Z = M/T and $\zeta = 2\mu/T$, where $M = m_1 + m_2$ is the total mass. The definitions of relativistic omega integrals are given in Refs. [31,32] and we do not write them here. For multicomponent system, the coefficients a_{kl} , c_{kl} are given as

$$a_{kk} = -a_{kl} = \sum_{l=1}^{N} a_{22}(kl) \quad (l \neq k)$$
 (A18)

$$c_{kk} = c_{00}(z_k) + \sum_{l=1}^{N} c_{22}(kl) \quad (l \neq k)$$
 (A19)

$$c_{kl} = c_{12}(kl) \quad (l \neq k), \tag{A20}$$

where $a_{22}(kl)$, $c_{22}(kl)$, and $c_{12}(kl)$ are the expressions given in Eqs. (A12)–(A15), with subscripts 1 and 2 replaced by k and l.

- J. Adams *et al.* (STAR Collaboration), Nucl. Phys. A757, 102 (2005).
- [2] M. Gyulassy and L. McLerran, Nucl. Phys. A750, 30 (2005).
- [3] T. Hirano and M. Gyulassy, Nucl. Phys. A769, 71 (2006).
- [4] M. Luzum and P. Romatschke, Phys. Rev. C 78, 034915 (2008); 79, 039903(E) (2009).
- [5] H. Song, S. A. Bass, U. Heinz, T. Hirano, and C. Shen, Phys. Rev. Lett. **106**, 192301 (2011); **109**, 139904(E) (2012).
- [6] B. Schenke, S. Jeon, and C. Gale, Phys. Rev. Lett. 106, 042301 (2011).
- [7] B. Jacak and P. Steinberg, Phys. Today 63, 5, 39 (2010).
- [8] V. Roy and A. K. Chaudhuri, Phys. Lett. B 703, 313 (2011).
- [9] U. Heinz and R. Snellings, Annu. Rev. Nucl. Part. Sci. 63, 123 (2013).
- [10] H.-J. Drescher, A. Dumitru, C. Gombeaud, and J.-Y. Ollitrault, Phys. Rev. C 76, 024905 (2007).
- [11] A. K. Chaudhuri, Phys. Lett. B 681, 418 (2009).
- [12] V. Roy, B. Mohanty, and A. K. Chaudhuri, J. Phys. G 40, 065103 (2013).
- [13] P. B. Arnold, G. D. Moore, and L. G. Yaffe, J. High Energy Phys. 11 (2000) 001.
- [14] P. B. Arnold, G. D. Moore, and L. G. Yaffe, J. High Energy Phys. 05 (2003) 051.
- [15] M. A. York and G. D. Moore, Phys. Rev. D 79, 054011 (2009).
- [16] W. Busza, K. Rajagopal, and W. van der Schee, Annu. Rev. Nucl. Part. Sci. 68, 339 (2018).
- [17] G. Policastro, D. T. Son, and A. O. Starinets, Phys. Rev. Lett. 87, 081601 (2001).
- [18] R. A. Lacey, N. N. Ajitanand, J. M. Alexander, P. Chung, W. G. Holzmann, M. Issah, A. Taranenko, P. Danielewicz, and H. Stoecker, Phys. Rev. Lett. 98, 092301 (2007).
- [19] L. P. Csernai, J. Kapusta, and L. D. McLerran, Phys. Rev. Lett. 97, 152303 (2006).
- [20] M. Prakash, M. Prakash, R. Venugopalan, and G. Welke, Phys. Rep. 227, 321 (1993).
- [21] D. Kharzeev and K. Tuchin, J. High Energy Phys. 09 (2008) 093.
- [22] P. Chakraborty and J. I. Kapusta, Phys. Rev. C 83, 014906 (2011).
- [23] A. Dobado and J. M. Torres-Rincon, Phys. Rev. D 86, 074021 (2012).
- [24] J.-W. Chen and J. Wang, Phys. Rev. C 79, 044913 (2009).
- [25] P. B. Arnold, C. Dogan, and G. D. Moore, Phys. Rev. D 74, 085021 (2006).
- [26] A. Wiranata, V. Koch, M. Prakash, and X. N. Wang, Phys. Rev. C 88, 044917 (2013).
- [27] A. Dash, S. Samanta, and B. Mohanty, Phys. Rev. C 97, 055208 (2018).
- [28] A. Dash, S. Samanta, and B. Mohanty, Phys. Rev. C 99, 044919 (2019).

- [29] R. L. Workman, W. J. Briscoe, and I. I. Strakovsky, Phys. Rev. C 94, 065203 (2016).
- [30] W. van Leeuwen, P. Polak, and S. de Groot, Physica (Utrecht) 63, 65 (1973).
- [31] W. van Leeuwen, A. Kox, and S. de Groot, Phys. Lett. 47A, 31 (1974).
- [32] W. van Leeuwen, Physica (Amsterdam) 81A, 249 (1975).
- [33] D. W. von Oertzen, Phys. Lett. B 280, 103 (1992).
- [34] O. N. Moroz, arXiv:1301.6670.
- [35] U. Gangopadhyaya, S. Ghosh, S. Sarkar, and S. Mitra, Phys. Rev. C 94, 044914 (2016).
- [36] S. Gavin, Nucl. Phys. A435, 826 (1985).
- [37] G. Kadam, S. Pawar, and H. Mishra, J. Phys. G 46, 015102 (2019).
- [38] M. I. Gorenstein, M. Hauer, and O. N. Moroz, Phys. Rev. C 77, 024911 (2008).
- [39] G. S. Denicol, C. Gale, S. Jeon, and J. Noronha, Phys. Rev. C 88, 064901 (2013).
- [40] G. P. Kadam and H. Mishra, Phys. Rev. C **92**, 035203 (2015).
- [41] S. Ghosh, T. C. Peixoto, V. Roy, F. E. Serna, and G. Krein, Phys. Rev. C 93, 045205 (2016).
- [42] A. Dobado and F.J. Llanes-Estrada, Phys. Rev. D 69, 116004 (2004).
- [43] N. Demir and S. A. Bass, Phys. Rev. Lett. 102, 172302 (2009).
- [44] J. B. Rose, J. M. Torres-Rincon, A. Schfer, D. R. Oliinychenko, and H. Petersen, Phys. Rev. C 97, 055204 (2018).
- [45] S. U. Chung, J. Brose, R. Hackmann, E. Klempt, S. Spanier, and C. Strassburger, Ann. Phys. (Berlin) 507, 404 (1995).
- [46] R. Dashen, S.-K. Ma, and H. J. Bernstein, Phys. Rev. 187, 345 (1969).
- [47] R. Venugopalan and M. Prakash, Nucl. Phys. A546, 718 (1992).
- [48] S. R. De Groot, *Relativistic Kinetic Theory. Principles and Applications*, edited by W. A. Van Leeuwen and C. G. Van Weert (North Holland Publishing Company, Amsterdam, Netherlands, 1980).
- [49] S. Weinberg, Phys. Rev. Lett. 17, 616 (1966).
- [50] K. Itakura, O. Morimatsu, and H. Otomo, Phys. Rev. D 77, 014014 (2008).
- [51] J. Weil et al., Phys. Rev. C 94, 054905 (2016).
- [52] Y. Aoki, S. Borsanyi, S. Durr, Z. Fodor, S. D. Katz, S. Krieg, and K. K. Szabo, J. High Energy Phys. 06 (2009) 088.
- [53] A. Bazavov et al., Phys. Rev. D 85, 054503 (2012).
- [54] M. Attems, J. Casalderrey-Solana, D. Mateos, D. Santos-Olivn, C. F. Sopuerta, M. Triana, and M. Zilho, J. High Energy Phys. 06 (2017) 154.
- [55] P. Kovtun, D. T. Son, and A. O. Starinets, Phys. Rev. Lett. 94, 111601 (2005).
- [56] K. Meier, A. Laesecke, and S. Kabelac, J. Chem. Phys. 122, 014513 (2005).