# $\Lambda\Lambda$ pairing in multistrange hypernuclei

H. Güven and K. Bozkurt

Physics Department, Yildiz Technical University, 34220 Esenler, Istanbul, Turkey

E. Khan

Institut de Physique Nucléaire, Université Paris-Sud, IN2P3-CNRS, F-91406 Orsay Cedex, France

## J. Margueron

Institute for Nuclear Theory, University of Washington, Seattle, Washington 98195, USA and Institut de Physique Nucléaire de Lyon, CNRS/IN2P3, Université de Lyon, Université Claude Bernard Lyon 1, F-69622 Villeurbanne Cedex, France

(Received 16 March 2018; published 17 July 2018)

Multistrange Ca, Sn, and Pb hypernuclei with  $\Lambda\Lambda$  pairing interaction are investigated within the Hartree-Fock-Bogoliubov approach. The unknown  $\Lambda\Lambda$  pairing strength is calibrated to match with the maximal value for the prediction of the  $\Lambda$  pairing gap in uniform matter for densities and isospin asymmetries equivalent to those existing in multi- $\Lambda$  hypernuclei. In this way, we provide an upper bound for the prediction of the  $\Lambda$  pairing gap and its effects in hypernuclei. The condensation energy is predicted to be about 3 MeV as a maximum value, yielding small corrections on density distributions and shell structure. In addition, conditions on both Fermi energies and orbital angular momenta are expected to quench the nucleon- $\Lambda$  pairing for most of hypernuclei.

DOI: 10.1103/PhysRevC.98.014318

#### I. INTRODUCTION

Since the discovery of the first hyperfragment in an emulsion exposed to cosmic rays [1] in 1952, hypernuclei have often been considered as the best systems to investigate the nuclear interaction in the baryon octet [2,3]. Despite the numerous theoretical works about hypernuclei physics within various frameworks, such as relativistic mean field (RMF) [4-7], G matrix combined with Skyrme-Hartree-Fock (SHF) for finitenuclei [8–11], generalized liquid drop model [12], as well as more recently quantum Monte Carlo approach [13,14], there are still open questions concerning the understanding of multistrange nuclei and the hypernuclear equation of state possibly running the properties of the inner core of neutron stars [15,16]. A very well-known issue is indeed the so-called hyperonization puzzle in neutron stars [17]. It refers to the difficulty for many approaches to reach the highest neutron stars masses observed, of about  $1.9-2.0M_{\odot}$  [18,19] when considering the softening induced by the onset of hyperons in the nuclear matter [16, 17, 20]. The possible solutions may be that the hyperon interaction turns strongly repulsive in dense matter, counterbalancing the softening of the phase transition [13,14,20], or that quarks appear before hyperons. Our present experimental investigation could not yet resolve this question. In general the main difficulties for theoretical approaches is the very scarce amount of data. Not only there are few  $N\Lambda$  scattering data, but the number of hypernuclei produced in laboratories is also very small. Due to experimental limitations, most of the produced hypernuclei are single- $\Lambda$  ones, and there are only a few light double- $\Lambda$ and single- $\Xi$  hypernuclei which are known. Constraints on the hyperon interactions are therefore still weak. As an example, the  $NN\Lambda$  interaction is still subject of debate [13,14].

Most of the recent theoretical approaches predict binding energies and single particle energies of single- $\Lambda$  systems such as  ${}_{\Lambda}^{5}$ He,  ${}_{\Lambda}^{9}$ Be,  ${}_{\Lambda}^{13}$ C,  ${}_{\Lambda}^{209}$ Pb in good agreement with the experimental data [4,7]. In the present work, for instance, we consider density functional approaches where the nucleon sector is treated with Skyrme interaction and the *N*  $\Lambda$  channel is based on *G*-matrix calculations starting from various bare interactions such as NSC89, NSC97a–f (Nijmegen soft core potentials) or ESC08 (extended soft core potentials) [9,11]. The oldest DF-NSC89 functional can reproduce with a good accuracy the experimental single particle energies of  $\Lambda$  hyperon for light hypernuclei such as  ${}_{\Lambda}^{5}$ He or  ${}_{\Lambda}^{13}$ C, but for the heavier hypernuclei like  ${}_{\Lambda}^{41}$ Ca or  ${}_{\Lambda}^{209}$ Pb, DF-NSC97a–f and DF-ESC08 are better [9,11]. It should be noted that this problem can be removed with adequate fitting such as adding new terms in the energy functional for the single  $\Lambda$  hypernuclei [9].

There are still many open questions related to multistrange systems. The first one is concerning the strength of the  $\Lambda\Lambda$  force. In general, the experimental bond energy of multistrange systems, such as  ${}^6_{\Lambda\Lambda}$ He or  ${}^{10}_{\Lambda\Lambda}$ Be, are not reproduced by the usual density functionals [7]: the DF-NSC89 and DF-NSC97f forces produce bond energies which goes from -0.34 MeV ( ${}^{10}_{\Lambda\Lambda}$ Be) to -0.12 MeV ( ${}^{210}_{\Lambda\Lambda}$ Pb) while the DF-NSC97a, which has strong  $\Lambda\Lambda$  interaction, predicts bond energies from 0.37 MeV ( ${}^{10}_{\Lambda\Lambda}$ Be) to 0.01 MeV ( ${}^{210}_{\Lambda\Lambda}$ Pb) [11]. It should be noted that more recent density functionals, e.g., the one derived from the ESC08 potential, are not better at reproducing the bond energy (-0.57 MeV for ESC08 in  ${}^6_{\Lambda\Lambda}$ He [9]). In order to improve the present description of the  $\Lambda\Lambda$  force, an empirical correction for DF-NSC89 and DF-NSC97a–f in the  $\Lambda\Lambda$  channel has been proposed and fitted to the bond energies

issues exists with relativistic approaches for hypernuclei and it was recently proposed than the fit of the bond energy shall enter directly in the definition of the Lagrangian [25]. Therefore we use DF-NSC89, DF-NSC97a, and DF-NSC97f potentials, due to the compatible results of  $\Lambda\Lambda$  channel.

There is however a question which has not been addressed yet and may modify our understanding of the  $\Lambda\Lambda$  channel in finite hypernuclei: are  $\Lambda$  particles paired, and how much  $\Lambda$ pairing impacts the ground state properties (density distributions, binding energies, etc.)? It should be noted that although there is currently no microscopic calculation in hypernuclei including  $\Lambda$  pairing, the pairing gap in hypernuclear matter has been calculated within the BCS approximation [26–30]. The present work aims to provide a first investigation to the  $\Lambda$ pairing in finite hypernuclei by considering a rather optimistic scenario for the strength of  $\Lambda\Lambda$  pairing. We will show that even under this optimistic case,  $\Lambda\Lambda$  pairing is still rather weak.

Another question related to multistrange nuclei is concerning the presence of other strange baryons than  $\Lambda$ , such as  $\Xi$ ,  $\Sigma$ , or  $\Omega$ . The  $\Xi$  is the most crucial one because the  $\Lambda + \Lambda \rightarrow$  $N + \Xi$  decay channel—also called  $\Xi$  instability—can make it appear in the ground state of hypernuclei [21,31]. The attractive nature of the  $\Xi$  potential in nuclear matter ( $U_{\Xi} = -14 \text{ MeV}$ ) [32] stabilizes  $\Xi$  and both  $\Xi^0$  and  $\Xi^-$  hyperons are predicted in hypernuclei with a strangeness number  $-S \ge 20 - 30$  [21]. Besides,  $\Lambda$  hyperons can also decay to  $\Sigma^0$ ,  $\Sigma^+$ , and  $\Sigma^-$ . However, due to the high  $Q^{\text{free}}$  values of  $\Sigma$  hyperons— $Q_{\Sigma}^{\text{free}} =$ -80 MeV for  $\Sigma$  and  $Q_{\Xi}^{\text{free}} = -26$  MeV for  $\Xi$  [21,31]—the decay of  $\Lambda$  to  $\Sigma^{\pm,0}$  is not favored in finite hypernuclei. Since the presence of  $\Xi$  in the ground state is not expected to enhance the  $\Lambda$  pairing, and since the pairing in the  $\Xi$  channel is even less known that one in the  $\Lambda$  channel, we focus our present study on the pairing in  $\Lambda$  hypernuclei. We do not expect our conclusions to be strongly modified by the presence of hyperons other than  $\Lambda$  in finite nuclei.

In this work, ground state properties of single and multi- $\Lambda$ hypernuclei are investigated with Hartree-Fock-Bogoliubov (HFB) formalism considering  $\Lambda\Lambda$  pairing interactions. On this purpose we neglect the  $\Lambda$  spin-orbit interaction which is estimated to be very small [33–35]. The three-body interactions such as  $NN\Lambda$  [13,14] is effectively included from the functional approach. We have considered a zero-range pairing force in the  $\Lambda\Lambda$  channel, opening the possibility to calculate accurately open- $\Lambda$  shell nuclei. In addition, our calculation are performed in spherical symmetry since deformation is not expected to increase greatly pairing correlations. We have considered hypernuclei with proton and neutron closed shells, e.g.,  ${}^{40-S}_{-S\Lambda}$ Ca,  ${}^{132-S}_{-S\Lambda}$ Sn, and  ${}^{208-S}_{-S\Lambda}$ Pb, since semimagicity often guarantees that nuclei remain at, or close to, sphericity. A future study evaluating the effect of deformation on the pairing correlation for open shell  $\Lambda$  states would however be interesting to confirm our conclusions.

The HFB equations for multi-strange hypernuclei are presented in Sec. II. The general features of shell evolution for multistrange hypernuclei are discussed in Sec. III. The possibility of NA pairing is discussed in Sec. IV, and, in Sec. V results with and without pairing interaction are

PHYSICAL REVIEW C 98, 014318 (2018)

discussed. Conclusions and outlooks are given in the last Sec. VI.

## **II. THEORETICAL FRAMEWORK**

Considering a nonrelativistic system composed of interacting nucleons N = (p, n) and  $\Lambda$ 's, the total Hamiltonian reads

$$\widehat{H} = \widehat{T}_N + \widehat{T}_\Lambda + \widehat{H}_{NN} + \widehat{H}_{N\Lambda} + \widehat{H}_{\Lambda\Lambda}, \qquad (1)$$

where  $\widehat{T}_A$  are the kinetic energy operators and  $\widehat{H}_{AB}$  are the interaction operator terms acting between A and B species  $(A, B = N, \Lambda)$ .

## A. Mean-field approximation

In the mean field approximation the ground state of the system is the tensor product  $|\Psi_N\rangle \otimes |\Psi_\Lambda\rangle$ , where  $|\Psi_N\rangle (|\Psi_\Lambda\rangle)$  is a slater determinant of the nucleon ( $\Lambda$ ) states. The total Hamiltonian (1) can be turned into a density functional  $\epsilon(\rho_N, \rho_\Lambda)$ , function of the particle densities  $\rho_N$  and  $\rho_\Lambda$ , as  $\hat{H} = \int \epsilon(\rho_N, \rho_\Lambda) d^3r$ . The energy functional  $\epsilon$  is often expressed as [11,36]

$$\epsilon(\rho_N, \rho_\Lambda) = \frac{\hbar}{2m_N} \tau_N + \frac{\hbar}{2m_\Lambda} \tau_\Lambda + \epsilon_{NN}(\rho_N) + \epsilon_{N\Lambda}(\rho_N, \rho_\Lambda) + \epsilon_{\Lambda\Lambda}(\rho_\Lambda), \qquad (2)$$

where  $\tau_N$  ( $\tau_\Lambda$ ) is the nucleonic ( $\Lambda$ ) kinetic energy density and  $\epsilon_{ij}$  are the interaction terms of the energy density functional describing the *NN*, *N* $\Lambda$ , and  $\Lambda\Lambda$  channels. In the following, the nucleonic terms will be deduced from the well-known SLy5 Skyrme interaction [37] widely used for the description of the structure of finite nuclei, while the *N* $\Lambda$  channel is given by a density functional  $\epsilon_{N\Lambda}$  adjusted to BHF predictions in uniform matter [11,36],

$$\epsilon_{N\Lambda}(\rho_N,\rho_\Lambda) = -f_1(\rho_N)\rho_N\rho_\Lambda + f_2(\rho_N)\rho_N\rho_\Lambda^{5/3}.$$
 (3)

Since the spin-orbit doublets are experimentally undistinguishable [33,34], the spin-orbit interaction among  $\Lambda$  particles is neglected [38]. The following density functionals are considered for the  $N\Lambda$  channel: DF-NSC89 [36], DF-NSC97a [11], DF-NSC97f [11]. In the  $\Lambda\Lambda$  channel  $\epsilon_{\Lambda\Lambda}$  is adjusted to the experimental bond energy in  ${}^{6}_{\Lambda\Lambda}$ He [21] from the so-called Nagara event [24],

$$\epsilon_{\Lambda\Lambda}(\rho_{\Lambda}) = -f_3(\rho_{\Lambda})\rho_{\Lambda}^2. \tag{4}$$

The corresponding empirical approach EmpC [21] for the  $\Lambda\Lambda$  channel is considered in the present work.

The functions  $f_{1-3}$  in Eqs. (3),(4) are expressed as

$$f_1(\rho_N) = \alpha_1 - \alpha_2 \rho_N + \alpha_3 \rho_N^2, \tag{5}$$

$$f_2(\rho_N) = \alpha_4 - \alpha_5 \rho_N + \alpha_6 \rho_N^2, \tag{6}$$

$$f_3(\rho_\Lambda) = \alpha_7 - \alpha_8 \rho_\Lambda + \alpha_9 \rho_\Lambda^2, \tag{7}$$

where  $\alpha_{1-7}$  are constants given in Table I. Note that  $\alpha_8 = \alpha_9 = 0$  [21].

TABLE I. Parameters of the functionals DF-NSC89, DF-NSC97a, and DF-NSC97f considering EmpC prescription for  $\alpha_7$  [21].

Functional	$\alpha_1$	$\alpha_2$	α <sub>3</sub>	$\alpha_4$	$\alpha_5$	$\alpha_6$	$\alpha_7$
DF-NSC89 +EmpC	327	1159	1163	335	1102	1660	22.81
DF-NSC97a +EmpC	423	1899	3795	577	4017	11061	21.12
DF-NSC97f +EmpC	384	1473	1933	635	1829	4100	33.25

In uniform nuclear matter the single particle energies read

$$\epsilon_N(k) = \frac{\hbar^2 k^2}{2m_N^*} + v_{NN}^{\text{matt.}} \quad \text{and} \ \epsilon_\Lambda(k) = \frac{\hbar^2 k^2}{2m_\Lambda^*} + v_\Lambda^{\text{matt.}}, \quad (8)$$

where the  $v_{\Lambda}^{\text{matt.}}$  decomposes as

$$v_{\Lambda}^{\text{matt.}} = v_{N\Lambda}^{\text{matt.}} + v_{\Lambda\Lambda}^{\text{matt.}}.$$
 (9)

The potentials  $v_{NN}^{\text{matt.}}$ ,  $v_{N\Lambda}^{\text{matt.}}$ , and  $v_{\Lambda\Lambda}^{\text{matt.}}$  derive from the energy functionals. Namely,

$$v_{NN}^{\text{matt.}}(\rho_N, \rho_\Lambda) = v_N^{\text{Skyrme}} + \frac{\partial \epsilon_{N\Lambda}}{\partial \rho_N},$$
 (10)

$$v_{N\Lambda}^{\text{matt.}}(\rho_{\Lambda}) = \frac{\partial \epsilon_{N\Lambda}}{\partial \rho_{\Lambda}},\tag{11}$$

$$v_{\Lambda\Lambda}^{\text{matt.}}(\rho_{\Lambda}) = \frac{\partial \epsilon_{\Lambda\Lambda}}{\partial \rho_{\Lambda}}.$$
 (12)

The nucleon effective mass is given from Skyrme interaction [39] and the  $\Lambda$  effective mass is expressed as a polynomial in the nucleonic density  $\rho_N$  as [36]

$$\frac{m_{\Lambda}^{*}(\rho_{N})}{m_{\Lambda}} = \mu_{1} - \mu_{2}\rho_{N} + \mu_{3}\rho_{N}^{2} - \mu_{4}\rho_{N}^{3}.$$
 (13)

The values for the parameters  $\mu_{1-4}$  are given in Table II.

In hypernuclei, the potentials  $v_N$  and  $v_\Lambda$  are corrected by the effective mass term as (see Ref. [21] and therein)

$$v_{NN}^{\text{nucl.}} = v_{NN}^{\text{matt.}} - \frac{3\hbar^2}{10m_N} \rho_N^{5/3} \left(\frac{6\pi^2}{g_N}\right)^{2/3} \left[\frac{m_N}{m_N^*} - 1\right], \quad (14)$$

$$v_{N\Lambda}^{\text{nucl.}} = v_{N\Lambda}^{\text{matt.}} - \frac{3\hbar^2}{10m_{\Lambda}} \rho_{\Lambda}^{5/3} \left(\frac{6\pi^2}{g_{\Lambda}}\right)^{2/3} \left[\frac{m_{\Lambda}}{m_{\Lambda}^*} - 1\right], \quad (15)$$

$$v_{\Lambda\Lambda}^{\text{nucl.}} = v_{\Lambda\Lambda}^{\text{matt.}}.$$
(16)

The present functional (SLy5 in the NN channel, DF-NSC in the N $\Lambda$  channel and EmpC in the  $\Lambda\Lambda$  channel) therefore yields an optimal set to perform HF calculations in hypernuclei (see [21] for details).

TABLE II. The parameters of the  $\Lambda$ -effective mass.

Force	$\mu_1$	$\mu_2$	$\mu_3$	$\mu_4$
DF-NSC89	1.00	1.83	5.33	6.07
DF-NSC97a	0.98	1.72	3.18	0
DF-NSC97f	0.93	2.19	3.89	0

### B. Mean field + pairing approximation

The HFB framework is well designed for the treatment of pairing both for strongly and weakly bound systems. In this work, we study hypernuclei which are magic in neutron and proton and open shell in  $\Lambda$ . We thus consider the Hartree-Fock-Bogoliubov (HFB) framework in the  $\Lambda\Lambda$  channel, and the *NN* channel is treated within Hartree-Fock (HF). In addition, because of their magic properties in the nucleon sector, which still contains the majority of particles, we consider spherical symmetry. In the HFB approach the mean field matrix that characterizes the system is obtained from the particle and pairing energy densities [40]. Particle and pairing densities can be expressed as

$$\rho(\mathbf{r}\sigma q, \mathbf{r}'\sigma' q') = \langle \psi | a^+_{\mathbf{r}'\sigma' q'} a_{\mathbf{r}\sigma q} | \psi \rangle, \tag{17}$$

$$\tilde{\rho}(\mathbf{r}\sigma q, \mathbf{r}'\sigma' q') = -2\sigma' \langle \psi | a^+_{\mathbf{r}' - \sigma' q'} a_{\mathbf{r}\sigma q} | \psi \rangle, \qquad (18)$$

where  $a_{r'\sigma'q'}^+$  and  $a_{r\sigma q}$  are creation and annihilation operators which affect nucleon with  $\sigma = \pm 1/2$  spin and q isospin from nucleon to hyperon at point r. The mean field matrix elements are obtained by variation of the expectation value of the energy with respect to the particle and pairing densities:

$$h(\mathbf{r}\sigma, \mathbf{r}'\sigma') = \frac{\delta E}{\delta\rho(\mathbf{r}\sigma, \mathbf{r}'\sigma')},$$
(19)

$$\tilde{h}(\mathbf{r}\sigma,\mathbf{r}'\sigma') = \frac{\delta E}{\delta\tilde{\rho}(\mathbf{r}\sigma,\mathbf{r}'\sigma')},$$
(20)

$$\int d^{3}r' \begin{pmatrix} h(\mathbf{r}\sigma, \mathbf{r}'\sigma') & \tilde{h}(\mathbf{r}\sigma, \mathbf{r}'\sigma') \\ \tilde{h}(\mathbf{r}\sigma, \mathbf{r}'\sigma') & -h(\mathbf{r}\sigma, \mathbf{r}'\sigma') \end{pmatrix} \begin{pmatrix} \psi_{1}(E, \mathbf{r}'\sigma') \\ \psi_{2}(E, \mathbf{r}'\sigma') \end{pmatrix}$$
$$= \begin{pmatrix} E + \epsilon & 0 \\ 0 & E - \epsilon \end{pmatrix} \begin{pmatrix} \psi_{1}(E, \mathbf{r}'\sigma') \\ \psi_{2}(E, \mathbf{r}'\sigma') \end{pmatrix}.$$
(21)

In Eq. (19), the diagonal elements of the matrix in the integral correspond to the particle-hole mean field while the nondiagonal elements of the matrix correspond to contributions of the pairing to the mean field of the hypernucleus. To calculate particle-hole channel, we use total energy functional by summing up each interaction channel [Eqs. (14), (15), and (16)] and generated particle energy densities with Eq. (19).

For the particle-particle channel, due to scarce available information, especially for the  $\Lambda$  pairing channel, it is convenient to consider a volume type zero range pairing interaction in the  $\Lambda\Lambda$  channel as

$$V_{\Lambda_{\text{pair}}} = V_{\Lambda_0} \delta(\mathbf{r}_1 - \mathbf{r}_2), \qquad (22)$$

where  $V_{\Lambda_0}$  is the  $\Lambda$  pairing strength. We use a 60 MeV cutoff energy and 15 $\hbar$  cutoff total angular momentum for quasiparticles, allowing for a large configuration space for all hypernuclei under study.

We now discuss the strength  $V_{\Lambda_0}$  of the  $\Lambda$  pairing interaction. At variance with the NN pairing interaction, there are not enough experimental data to set the  $\Lambda\Lambda$  pairing interaction. We therefore choose to calibrate the  $\Lambda\Lambda$  pairing interaction to calculations of  $\Lambda$  pairing gaps in uniform matter, see for instance the very recent work in Ref. [41]. There are several predictions for the  $\Lambda$  pairing gap in uniform matter which have been employed in cooling models for neutron stars. These predictions are substantially different for several reasons: they were calibrated on either the old [26,27] or the more recent [28,30] value for the Nagara event [24]; they were considering nonrelativistic [26,27] or relativistic mean field [28,30] approaches; as a consequence, they incorporate different density dependencies of the nucleon and  $\Lambda$  effective masses; they are based on various  $\Lambda$  interactions which are weakly constrained. As a result, qualitatively different predictions have been performed in nuclear matter: the influence of the nucleon density on the  $\Lambda$  pairing gap has been found opposite between nonrelativistic approaches [26,27] and relativistic ones [28]. Despite these differences, the predictions of the  $\Lambda$  pairing gap at saturation density and for  $k_{F_{\Lambda}} \approx 0.8 \text{ fm}^{-1}$ (corresponding to the average  $\Lambda$  density  $\rho_{sat}/5$  in hypernuclei) are rather consistent across the different predictions and reach a maximum at about 0.5 MeV-0.8 MeV. For instance, under these conditions the  $\Lambda$  pairing gap is predicted to be about 0.5 MeV for  $\rho_N = \rho_{\text{sat}}$  with HS-m2 parameters [28], and 0.5 (0.75) MeV for NL3 (TM1) parameters with ESC00  $\Lambda$ force sets [30]. These values are also consistent with the extrapolations of earlier calculations [26,27] in hypernuclear matter. Some interactions predict however lower values. In the following, we will therefore calibrate our  $\Lambda\Lambda$  pairing interaction on hypernuclear matter predictions of Ref. [28], which represents an average prediction for the maximum possible  $\Lambda$  pairing gap.

In addition to the  $\Lambda\Lambda$  pairing, let us mention the existence of a prediction suggesting a strong  $N\Lambda$  pairing interaction in nuclear matter [29]. In finite nuclei, large  $N\Lambda$  pairing gaps may however be quenched by shell effects, due to large single particle energy differences between the N and  $\Lambda$  states, or mismatch of the associated single particle wave functions. This will be discussed in more details in Sec. III B.

## **III. RESULTS**

We present in this section the prediction for  $\Lambda$  pairing gap and its effect in multi- $\Lambda$  hypernuclei. We first discuss the relative gaps between N and  $\Lambda$  single particle energies predicted by HF calculation without pairing to assert our calculation without  $N\Lambda$  pairing. Then, we employ HFB framework with pairing in the  $\Lambda\Lambda$  channel to study binding energies and density profiles in multi- $\Lambda$  hypernuclei.

## A. Shell structure of hypernuclei without ΛΛ pairing interaction

Let us first discuss the hypernuclei of interest in this work, without  $\Lambda\Lambda$  pairing interaction. On this purpose, we investigate closed shell hypernuclei such as  $^{60}_{20\Lambda}$ Ca,  $^{172}_{40\Lambda}$ Sn,  $^{278}_{70\Lambda}$ Pb shown in Fig. 1. These nuclei are triply magic. Due to the absence of the spin-orbit term, the shell structure of hyperon is also expected to be similar to that of the spherical harmonic oscillator, with magic numbers 2, 8, 20, 40, 70, etc., and the energy gaps are larger than in ordinary nuclei. While the central potential is also expected to modify the details of the single particle spectrum, the gross increase of the single particle gap is mostly due to the absence of spin-orbit interaction. The new magic number in the  $\Lambda$  channel is clearly caused by absence of spin-orbit interaction, increasing the degeneracy of the states as



DF-NSC89+EmpC DF-NSC97a+EmpC DF-NSC97f+EmpC

FIG. 1. The  $\Lambda$  single particle spectrum  $^{60}_{20\Lambda}$  Ca (a),  $^{172}_{40\Lambda}$  Sn (b), and  $^{278}_{70\Lambda}$  Pb (c) hypernuclei, calculated with the HF approach.

well as the average energy gaps between single particle states. In order to check this statement, we have calculated the single particle spectrum for other Skyrme interactions, e.g., SGII, SAMI, and SIII, and we found the same gross conclusions.

TABLE III. Energy difference of each shell between DF-NSC97a+EmpC and DF-NSC89+EmpC force sets. The detailed spectra are shown in Fig. 1.

Shell	E	nergy difference (Me	V)
	<sup>60</sup> 20л <b>Са</b>	$^{172}_{40\Lambda}{ m Sn}$	$^{278}_{70\Lambda}{ m Pb}$
1 <i>s</i>	6.00	7.50	8.50
1 <i>p</i>	6.00	6.87	7.10
1d	3.32	5.80	6.36
2 <i>s</i>	2.59	5.57	6.42
1 <i>f</i>		4.81	6.30
2p		4.20	5.20

The average single particle gaps between two neighboring orbitals can be estimated from Fig. 1, where the  $\Lambda$  spectrum is shown for  ${}^{60}_{20\Lambda}$ Ca,  ${}^{172}_{40\Lambda}$ Sn, and  ${}^{278}_{70\Lambda}$ Pb hypernuclei and for three different density functional in the  $\Lambda$  channel (the Skyrme interaction SLy5 is fixed in the nucleon channel): the average single-particle gap is found to be generally larger than 4 MeV, except for the gap between the 2s-1d and 3s-2d states, where it is between 1 and 3 MeV. These smaller energy gaps may be related to the pseudospin symmetry [42,43], since the 2s-1d and 3s-2d states are pseudospin partners. The small energy gap between these states makes them good candidates for pairing correlations: these states could largely mix against pairing correlations when they are close to the Fermi level, represented in dashed lines in Fig. 1. For the selected nuclei in Fig. 1, the Fermi level is indeed close to either the 2s-1dor the 3s-2d states in the cases of Ca and Pb hypernuclei, respectively.

The energy spectra predicted by DF-NSC89+EmpC and DF-NSC97f+EmpC are mainly identical, while the single particle states predicted by DF-NSC97a+EmpC are systematically more bound, since the  $N\Lambda$  potential which is deeper for DF-NSC97a+EmpC compared to the two others functionals [21,38]. We give more quantitative estimation of the single particle energy differences between the predictions of DF-NSC97a+EmpC and DF-NSC89+EmpC in Table III. As expected, the larger the number of hyperons, the larger the differences as well, since the difference is larger for the deep single particle states than for the weakly bound ones.

#### B. The NA and AA pairing channels

We now discuss the  $N\Lambda$  and  $\Lambda\Lambda$  pairing channels. These two pairing channels are expected to compete: a  $\Lambda$  can be paired either to a nucleon or to another  $\Lambda$ . Drawing an analogy with T = 0 and T = 1 pairing channels in ordinary nuclei, the pairing interaction between two different particles (T = 0) can occur under the condition of a good matching between their wave functions. It also requires a good matching between their single particle energies. This is the main reason why T = 0pairing is expected to appear mainly at (or close to) N = Znuclei [44,45].

Let us first focus on the  $N\Lambda$  pairing. A necessary condition for this pairing to occur is that the neutron or proton Fermi energy is close to the  $\Lambda$  one. The neutron, proton, and  $\Lambda$  Fermi energies are displayed on Fig. 2 as function of the strangeness number -S for the three representative nuclei:  ${}^{40-S}_{-S\Lambda}$ Ca,  ${}^{132-S}_{-S\Lambda}$ Sn, and  ${}^{208-S}_{-S\Lambda}$ Pb. The intersections of nucleons and  $\Lambda$  Fermi energies occur at -S = 4 (neutrons) and 8 (protons) for  ${}^{40-S}_{-S\Lambda}$ Ca, -S = 10-16 (proton) and 20–32 (neutrons) for  ${}^{132-S}_{-S\Lambda}$ Sn and for -S = 34-40 for both neutrons and protons  ${}^{208-S}_{-S\Lambda}$ Pb hypernuclei.

Let us now take typical examples of the nuclei which are located at these crossings.  ${}^{44}_{4\Lambda}$ Ca and  ${}^{48}_{8\Lambda}$ Ca single-particle levels are shown in Fig. 3 and the ones of  ${}^{244}_{36\Lambda}$ Pb is shown in Fig. 4. The  $\Lambda$  Fermi level is mainly the 1*p* state in  ${}^{44}_{4\Lambda}$ Ca and  ${}^{48}_{8\Lambda}$ Ca, and there are no *p* states in the neutron and proton spectrum around the Fermi energy. The conditions for  $N\Lambda$  pairing are therefore not satisfied for Ca isotopes.

A similar analysis can be made for the Sn isotopes. We also calculated  ${}^{142}_{10\Lambda}$  Sn,  ${}^{52}_{20\Lambda}$  Sn, and  ${}^{56}_{24\Lambda}$  Sn hypernuclei for which the proton or neutron and the  $\Lambda$  levels cross. The last occupied  $\Lambda$  states is 1*d* for  ${}^{142}_{10\Lambda}$  Sn (respectively 2*s* for  ${}^{152}_{20\Lambda}$  Sn and 1*f* for  ${}^{156}_{24\Lambda}$  Sn) while the corresponding proton (neutron) state is 1 $g_{9/2}$  (1 $h_{11/2}$ ). Since the quantum orbital numbers does not coincide between the nucleons and the  $\Lambda$  states in the cases where their respective Fermi energies cross, the  $N\Lambda$  pairing is not favored for these Ca and Sn nuclei.

The case of  $\frac{208-S}{-S\Lambda}$ Pb hypernuclei is different. Figure 4 displays the single particle spectrum for  $\frac{244}{36\Lambda}$ Pb hypernucleus, since the crossing of the nucleon (neutrons and protons) and  $\Lambda$  Fermi levels occurs at around S = -36 (Fig. 3). Figure 4 shows that the last filled orbits are  $3s_{1/2}$  for proton,  $3p_{1/2}$  for neutron and 2p for  $\Lambda$ . Since Pb is magic in the proton, only neutrons and  $\Lambda$  may be paired. We therefore predict that  $n\Lambda$  pairing may occur for  $\frac{208-S}{-S\Lambda}$ Pb hypernuclei and for  $\Lambda$  between S = -34 and S = -40. For lower or higher values of *S*, the mismatching of the single particles orbitals does not favor  $n\Lambda$  pairing. Since the level density increases with increasing masses, it is expected the general trend that  $N\Lambda$  pairing may occur more frequently as *A* increases.

 $^{208-S}_{-S\Lambda}$ Pb is a typical case representing heavy hypernuclei. Since the  $\Xi$  instability is expected to occur around -S = 70 [21], we can infer that pairing may occur for about 10% of  $^{208-S}_{-S\Lambda}$ Pb isotopes. This number may be considered as the maximum percentage of heavy hypernuclei where  $N\Lambda$  pairing may occur. The amount of hypernuclei where  $N\Lambda$  pairing is possible is therefore predicted to be small. In the following, we will avoid the cases where it may occur.

#### C. $\Lambda\Lambda$ pairing and binding energies

We now focus on the  $\Lambda\Lambda$  pairing and consider the cases of semimagical hypernuclei, such as  ${}^{40-S}_{-S\Lambda}$ Ca,  ${}^{132-S}_{-S\Lambda}$ Sn, and  ${}^{208-S}_{-S\Lambda}$ Pb. It should be noted that these nuclei are magic in both proton and neutron numbers, which helps most of these hypernuclei to resist against deformation, as in the case of normal hypernuclei. They have however an open shell in the  $\Lambda$  channels.

The  $\Lambda$  pairing strengths, mean gaps, and averaged mean gaps of isotopic chains are displayed in Table IV. The fitting procedure for the  $\Lambda\Lambda$  pairing is the following: we first remind that the  $\Lambda\Lambda$  mean-field interaction is calibrated to the  $\Lambda\Lambda$ 



FIG. 2. Evolution of proton, neutron, and  $\Lambda$  Fermi energies function of strangeness number -S for  $^{40-S}_{-S\Lambda}Ca$  (a),  $^{132-S}_{-S\Lambda}Sn$  (b), and  $^{208-S}_{-S\Lambda}Pb$  (c) hypernuclei with the HF approach.

bond energy in <sup>6</sup>He (Nagara event). Then we consider openshell nuclei and calibrate the average  $\Lambda$ -pairing gap to its expectation from uniform matter calculations. Densities are averaged from r = 0.2 fm to 10 fm for each hypernucleus and each force set using HF results. Fermi momenta corresponding to these densities are calculated as  $k_{F_{\Lambda}} = (\frac{3\pi^2}{2}\rho_{\Lambda})^{1/3}$ . The density profile of hypernuclear matter calculations [28] which has corresponding Fermi momentum and density fraction allows to extract  $\Lambda\Lambda$  pairing gap for each hypernuclei for each force sets. For finding adequate  $\Lambda$  pairing strength [V\_{\Lambda\_{\rm pair}} in Eq. (22)], starting from  $-50 \text{ MeV fm}^3$  to  $-300 \text{ MeV fm}^3$ , we iterated and determined mean gap values for each hypernuclei chain in HFB calculation. On each iteration, the mean gap values are averaged over all isotopic chain until similar values of pairing gaps of hypernuclear matter calculation are obtained. Namely for the  $\frac{40-S}{-S\Lambda}$ Ca isotopic chain, the average mean gap was calculated by summing each mean gap of hypernucleus starting from -S = 6 to -S = 20 and dividing by the total isotope number. Similarly for the  $^{132-S}_{-S\Lambda}$ Sn ( $^{208-S}_{-S\Lambda}$ Pb) isotopes, the average man gap determined between -S = 18(-S = 58)to -S = 40(-S = 70) region. The average mean gaps for each isotopes with each force set is given in Table IV. A typical

0.5 MeV gap is obtained in all nuclei, leading to a pairing effect independent of the number of  $\Lambda$  involved.

The effect of  $\Lambda\Lambda$  pairing on the binding energy can be estimated from the condensation energy, defined as  $E_{cond} = E_{HF} - E_{HFB}$ . The condensation energy measures the impact of the pairing effect on the binding energy. Figure 5 displays the condensation energy for a set of  $\frac{40-S}{-S\Lambda}$ Ca and  $\frac{132-S}{-S\Lambda}$ Sn semimagical hypernuclei. As in the case of normal nuclei, the condensation energy evolves as arches, with zero values at closed shells and maximum values for middle-open shells. The condensation energy can reach about 3 MeV in midopen shell hypernuclei for Ca and Sn isotopes. Since the  $\Lambda\Lambda$  interaction considered here is calibrated on the maximum prediction for the  $\Lambda$  gap in uniform matter with respect to  $\Lambda$  force sets, the condensation energy represents the estimation of the maximum value for the condensation energy generated by  $\Lambda\Lambda$ interaction.

The  $\Lambda$  numbers at which condensation energy is zero signs the occurrence of shell closure. It is therefore not surprising to recover the magical numbers 8, 20, 40, as we previously discussed. Strong subshell closure occurs for  $\Lambda = 34$  in Sn isotopes corresponding the filling of the 1 *f* state.



FIG. 3. A complete single particle spectrum of  ${}^{44}_{4\Lambda}$  Ca (a) and  ${}^{48}_{8\Lambda}$  Ca (b) hypernuclei, calculated with the HF approach.

Investigating the effect of  $\Lambda\Lambda$  pairing on the single particle energies, it turns out to be weak: states around the Fermi level are shifted by about 100–200 keV at maximum. The impact of



FIG. 4. A complete single particle spectrum of  $^{244}_{36\Lambda}$  Pb hypernucleus with the HF approach.

 $\Lambda$  pairing on single particle energies remains therefore rather small.

## D. $\Lambda\Lambda$ pairing and densities

We now discuss the effect of  $\Lambda\Lambda$  pairing on both normal and pairing densities. Figure 6 shows normal density profiles for  ${}^{40-S}_{-S\Lambda}$ Ca,  ${}^{132-S}_{-S\Lambda}$ Sn, and  ${}^{208-S}_{-S\Lambda}$ Pb series of hypernuclei. For the  ${}^{40-S}_{-S\Lambda}$ Ca series we consider cases where the  $N\Lambda$  pairing is not expected to occur. As mentioned above, the 1*d* and 2*s* states are almost degenerate, and can largely mix due to pairing correlations. Namely, before the 1*d* orbital is completely filled,  $\Lambda$  hyperons start to fill the 2*s* state due to the pairing interaction, resulting in a small increase at the center of the hypernucleus which corresponds the effect of the *s* state. Similar results can be seen on the density profile of  ${}^{208-S}_{-S\Lambda}$ Pb hypernucleus: before the 2*d* state is completely filled,  $\Lambda$  hyperons start to fill the 3*s* state due to the pairing interaction resulting from the almost degeneracy of the 2*d* and 3*s*  $\Lambda$  states. In the case of  ${}^{132-S}_{-S}$ Sn, there is no major difference on density profiles: because of the large gap between 1 *f* and 2*p* states, the  $\Lambda$  pairing changes only the total energy of the  ${}^{132-S}_{-S}$ Sn isotopic chain in -S = 24

TABLE IV. Pairing strength,  $\Lambda$  density, and calculated, averaged mean gap, and hypernuclear pairing gap for each force sets.

Force set	Pairing strength (MeV fm <sup>3</sup> )	$ ho_{\rm av}$ (fm <sup>-3</sup> )	Hypernucleus	Calculated mean gap (MeV)	Average mean gap (MeV)	Pairing gap in uniform matter [28] (MeV)
DF-NSC89+EmpC	-139	0.0264	${}^{46}_{6\Delta}$ Ca	0.82	0.59	0.42
DF-NSC97a+EmpC	-148	0.0349	$^{46}_{6\Delta}$ Ca	1.04	0.50	0.44
DF-NSC97f+EmpC	-180	0.0241	$^{46}_{6\Lambda}$ Ca	0.98	0.49	0.43
DF-NSC89+EmpC	-158	0.0421	$\frac{160}{28\Lambda}$ Sn	0.84	0.46	0.43
DF-NSC97a+EmpC	-145	0.0499	$\frac{160}{28\Lambda}$ Sn	0.82	0.45	0.50
DF-NSC97f+EmpC	-180	0.0382	$^{160}_{28\Lambda}$ Sn	0.82	0.46	0.45
DF-NSC89+EmpC	-184	0.0400	$^{272}_{64\Lambda}$ Pb	0.69	0.44	0.47
DF-NSC97a+EmpC	-180	0.0320	$^{272}_{64\Lambda}$ Pb	0.76	0.46	0.45
DF-NSC97f+EmpC	-220	0.0270	$^{272}_{64\Lambda} \mathrm{Pb}$	0.71	0.44	0.40



FIG. 5. Difference of binding energies between HF and HFB for  ${}^{40-S}_{-S\Lambda}$ Ca (a) and  ${}^{132-S}_{-S\Lambda}$ Sn (b) hypernuclei with DF-NSC89+EmpC, DF-NSC97a+EmpC, and DF-NSC97f+EmpC force sets.

to -S = 30 zone but does not impact the occupation numbers of 1f and 2p orbitals.

Figure 7 displays the  $\Lambda$  pairing density for  ${}^{40-S}_{-S\Lambda}$ Ca,  ${}^{132-S}_{-S\Lambda}$ Sn, and  $\frac{208-S}{-S\Lambda}$  Pb. As mentioned above, pairing interaction effects result in the partial occupation of  $\Lambda$  hyperons in the *s* and *d* states. The pairing density of <sup>54</sup><sub>14</sub> Ca hypernucleus peaks at 3 fm due to half-filled 1d orbital. As strangeness number increases, hyperons start to fill the 2s state and the contribution of the 1dstate decreases. For -S = 18,  $\Lambda$  hyperons starts to largely fill the 2s state before the 1d state is completely full, resulting in a pairing density having non-negligible contributions of both s and d states. Similar result can be seen for the pairing density of  $\frac{208-s}{-s\Lambda}$  Pb hypernuclei which has 2*d*-3*s* coupling. At -S = 64, pairing densities are mainly built from the 2d state but as the strangeness number increases, the pairing of 2dorbital decreases while pairing density at 3s state increases. However for  $\frac{132-s}{-s\Lambda}$ Sn hypernuclei, the situation is different. Due to the large energy gap between 2s and 1f states, the pairing interaction does not change the occupation of these states. For this reason, the pairing density is only built from the



FIG. 6. Normal density profiles with on/off pairing for  ${}^{40-S}_{-S\Lambda}$ Ca (a),  ${}^{132-S}_{-S\Lambda}$ Sn (b), and  ${}^{208-S}_{-S\Lambda}$ Pb (c) hypernuclei, calculated with the HFB approach.

1 f orbital and its magnitude increases when the occupation of the 1 f orbital increases until it is half-filled. When the 1 f state is more than half-filled, the magnitude of the pairing density starts to decrease. The spatial distribution of pairing effect in



FIG. 7. A pairing densities for  ${}^{40-S}_{-SA}$ Ca (a),  ${}^{132-S}_{-SA}$ Sn (b), and  ${}^{208-S}_{-SA}$ Pb (c) hypernuclei, calculated with the HFB approach.

PHYSICAL REVIEW C 98, 014318 (2018)

hypernuclei is therefore expected to exhibit strong variations from one hypernucleus to another, because of the weak spinorbit effect, giving rise to well separated sets of states.

### **IV. CONCLUSIONS**

In this work we have investigated the effect of  $\Lambda$  pairing on the ground state properties of hypernuclei within the Hartree-Fock-Bogoliubov formalism. The SLy5 Skyrme functional is used in the NN channel, while for  $N\Lambda$  channel we employ three functionals fitted from microscopic Brueckner-Hartree-Fock calculations: DF-NSC89, DF-NSC97a, and DF-NSC97f. These functionals reproduce the sequence of single- $\Lambda$  experimental binding energies from light to heavy hypernuclei. For the  $\Lambda\Lambda$  channel, we used the empirical prescription EmpC which is calibrated to 1 MeV which is the experimental bond energy in  ${}^{6}\text{He}_{\Lambda\Lambda}$ . Based on these density-functional approach, several nuclei have been studied with nucleon closed-shells and  $\Lambda$  open shells. A  $\Lambda\Lambda$  pairing interaction is therefore introduced, which magnitude is calibrated to be consistent with the maximum BCS predictions for the  $\Lambda$  pairing gap in hypernuclear matter.

Since the energy difference between the N and  $\Lambda$  Fermi levels is usually large (more than 5 MeV) in the considered nuclei, the  $N\Lambda$  pairing is quenched in most of the cases. The impact of  $\Lambda \Lambda$  pairing on the binding energies, density profiles, and single particle energies have been analyzed for  $\frac{40-S}{-S\Lambda}$ Ca,  $^{132-S}_{-S\Lambda}$ Sn, and  $^{208-S}_{-S\Lambda}$ Pb chains. We have shown that the effects of the  $\Lambda\Lambda$  pairing depends on hypernuclei. At maximum, the condensation energy in these chains is about 3 MeV. Density profiles reflect the occurrence of almost degenerate states in the  $\Lambda$  single particle spectrum, such as for instance the almost degeneracy between the 1*d* and 2*s* states in  $\frac{40-s}{-s\Lambda}$ Ca hypernuclei and 2d and 3s almost degeneracy in  $\frac{276}{68\Lambda}$  Pb. The effects of the A pairing also depend on the  $N\Lambda$  and  $\Lambda\Lambda$  force sets, but we found only a small overall impact. Generally, we found that  $\Lambda\Lambda$  pairing could be active if the energy gap between orbitals is smaller than 3 MeV. Under this condition, A pairing could impact densities and binding energies. Since only a weak spin-orbit interaction is expected in the  $\Lambda$  channel,  $\Lambda$  states are highly degenerated and usually distant by more than 3 MeV in energy. In conclusion, the present microscopic approach shows that the  $\Lambda$ -related pairing effect can usually be neglected in most of hypernuclei, except for hypernuclei which have a single particle gap lower than 3 MeV around the Fermi level.

### ACKNOWLEDGMENTS

This work is supported by the Yildiz Technical University Project (BAPK Project No. FBI-2018-3325). J.M. was partially supported by the IN2P3 Master Project MAC, New-CompStar COST Action MP1304, PHAROS COST Action MP16214.

<sup>[1]</sup> M. Danysz and J. Pniewski, Philos. Mag. Ser. 44, 348 (1953).

<sup>[2]</sup> A. Feliciello and T. Nagae, Rep. Prog. Phys. 78, 096301 (2015).

- [3] A. Gal, E. V. Hungerford, and D. J. Millener, Rev. Mod. Phys. 88, 035004 (2016).
- [4] S. H. Ren, T. T. Sun, and W. Zhang, Phys. Rev. C 95, 054318 (2017).
- [5] E. N. E. van Dalen, G. Colucci, and A. Sedrakian, Phys. Lett. B 734, 383 (2014).
- [6] H. F. Lü, Chinese Phys. Lett. 25, 3613 (2008).
- [7] Shen Yao-song and Ren Zhong-zhou, Acta Phys. Sin. (Overseas Edn.) 7, 258 (1998).
- [8] N. Guleria, S. K. Dhiman, and R. Shyam, Nucl. Phys. A 886, 71 (2012).
- [9] H.-J. Schulze and T. Rijken, Phys. Rev. C 88, 024322 (2013).
- [10] X. R. Zhou, E. Hiyama, and H. Sagawa, Phys. Rev. C 94, 024331 (2016).
- [11] I. Vidaña, A. Polls, A. Ramos, and H.-J. Schulze, Phys. Rev. C 64, 044301 (2001).
- [12] C. Samanta, P. R. Chowdhury, and D. N. Basu, J. Phys. G: Nucl. Part. Phys. 37, 075104 (2010).
- [13] D. Lonardoni, S. Gandolfi, and F. Pederiva, Phys. Rev. C 87, 041303(R) (2013).
- [14] D. Lonardoni, A. Lovato, S. Gandolfi, and F. Pederiva, Phys. Rev. Lett. **114**, 092301 (2015).
- [15] M. Baldo, G. F. Burgio, and H.-J. Schulze, Phys. Rev. C 58, 3688 (1998); 61, 055801 (2000); I. Vidaña, A. Polls, A. Ramos, L. Engvik, and M. Hjorth-Jensen, *ibid.* 62, 035801 (2000).
- [16] J. Schaffner-Bielich, Nucl. Phys. A 804, 309 (2008); 835, 279 (2010).
- [17] G. Colucci and A. Sedrakian, Phys. Rev. C 87, 055806 (2013).
- [18] J. Antoniadis *et al.*, Science **340**, 1233232 (2013).
- [19] E. Fonseca et al., Astrophys. J. 832, 167 (2016).
- [20] I. Bombaci, JPS Conf. Proc. 17, 101002 (2017).
- [21] J. Margueron, E. Khan, and F. Gulminelli, Phys. Rev. C 96, 054317 (2017).
- [22] J. K. Ahn et al., Phys. Rev. C 88, 014003 (2013).
- [23] S. Aoki, S. Y. Bahk, S. H. Chung *et al.* (KEK E176 Collaboration), Nucl. Phys. A 828, 191 (2009).

- [24] H. Takahashi et al., Phys. Rev. Lett. 87, 212502 (2001).
- [25] M. Fortin, S. S. Avancini, C. Providência, and I. Vidaña, Phys. Rev. C 95, 065803 (2017).
- [26] S. Balberg and N. Barnea, Phys. Rev. C 57, 409 (1998).
- [27] T. Takatsuka and R. Tamagaki, Nucl. Phys. A 670, 222 (2000).
- [28] T. Tanigawa, M. Matsuzaki, and S. Chiba, Phys. Rev. C 68, 015801 (2003).
- [29] X. R. Zhou, H.-J. Schulze, F. Pan, and J. P. Draayer, Phys. Rev. Lett. 95, 051101 (2005).
- [30] Y. N. Wang and H. Shen, Phys. Rev. C 81, 025801 (2010).
- [31] J. Schaffner-Bielich and A. Gal, Phys. Rev. C 62, 034311 (2000).
- [32] P. Khaustov *et al.* (The AGS E885 Collaboration), Phys. Rev. C 61, 054603 (2000).
- [33] O. Hashimoto and H. Tamura, Prog. Part. Nucl. Phys. 57, 564 (2006).
- [34] T. Motoba, D. E. Lanskoy, D. J. Millener, and Y. Yamamoto, Nucl. Phys. A 804, 99 (2008).
- [35] P. Finelli, N. Kaiser, D. Vretenar, and W. Weise, Phys. Lett. B 658, 90 (2007); Nucl. Phys. A 831, 163 (2009).
- [36] J. Cugnon, A. Lejeune, and H.-J. Schulze, Phys. Rev. C 62, 064308 (2000).
- [37] M. Bender, P. H. Heenen, and P.-G. Reinhard, Rev. Mod. Phys. 75, 121 (2003).
- [38] E. Khan, J. Margueron, F. Gulminelli, and Ad. R. Raduta, Phys. Rev. C 92, 044313 (2015).
- [39] J. Bartel, K. Bencheikh, and J. Meyer, Phys. Rev. C 77, 024311 (2008).
- [40] K. Bennaceur and J. Dobaczewski, Comp. Phys. Commun. 168, 96 (2005).
- [41] A. R. Raduta, A. Sedrakian, and F. Weber, Mon. Not. R. Astron. Soc. 475, 4347 (2018).
- [42] J. N. Ginocchio, Phys. Rep. 414, 165 (2005).
- [43] H. Liang, J. Meng, and S.-G. Zhou, Phys. Rep. 570, 1 (2015).
- [44] H. H. Wolter et al., Nucl. Phys. A 167, 108 (1971).
- [45] A. L. Goodman, Phys. Rev. C 60, 014311 (1999).