## **B(E2;**  $5/2^-$  →  $1/2^-$ ) in <sup>17</sup>N and <sup>17</sup>Ne

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A simple model of E2 strengths previously applied to several neutron-excess light nuclei is used to investigate the  $1/2^- \rightarrow 5/2^-$  transition strength in <sup>17</sup>Ne, with the aid of mirror symmetry. The calculation is found to be in good agreement with results of a recent measurement and emphasizes a likely problem with the same transition in  $17<sub>N</sub>$ .

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## **I. INTRODUCTION**

I recently performed an analysis [\[1\]](#page-1-0) of E2 transition strengths [\[2–7\]](#page-1-0) in several neutron-excess light nuclei. Transitions were chosen so that the  $J^{\pi}$  values excluded the possibility of M1 competition. The included nuclei were those with one or two sd-shell neutrons outside a p-shell core, so that a reasonable assumption was that the transitions involved only neutrons. Transitions were of the type  $1d_{5/2} \leftrightarrow 2s_{1/2}$  or  $(sd)^2$ <sup>2</sup>  $\leftrightarrow$   $(sd)^2$ <sub>0</sub>. Given information from <sup>17,18</sup>O, a combination of weak coupling and the shell model allowed parameterfree predictions for E2 transitions in  $^{16,17}$ N and  $^{15,16}$ C. Results for those four nuclei are summarized in Table I and plotted in Fig. 1. Agreement between experimental and predicted strengths is reasonable for all four nuclei, but agreement is worst for  ${}^{17}N$ .

A realistic shell-model calculation  $[8]$  for <sup>17</sup>N also predicted a B(E2) significantly larger than the current experimental value [ $exp/\text{shell-model} = 0.68(13)$ ], prompting me to suggest [\[1\]](#page-1-0) "a remeasurement of the  $17N$  gamma width might be warranted."

## **II. CALCULATIONS AND RESULTS**

Because of mirror symmetry, information from <sup>17</sup>Ne should help in assessing the  $17N$  problem. However, in  $17Ne$ , the 5/2<sup>−</sup> state is unbound and decays by *2p* emission, making a

TABLE I. E2 transition strengths in relevant nuclei.<sup>a</sup>

				$B(E2) (e^2 fm^4)$	
<b>Nucleus</b>	$J_i^{\mu}$	$J_f{}^n$	Calculated		Measured
			$e_n = 0.5e$	$e_n = (Z/A)e$	
$^{17}N$	$5/2^{-}$	$1/2^{-}$	4.02	3.52	$2.25(44)$ <sup>b</sup>
16 <sub>N</sub>	$0-$	$2^{-}$	4.25	3.66	$4.25(5)^{b}$
${}^{16}C$	$2^+$	$0+$	4.05	3.00	$3.50(30)$ <sup>c</sup>
${}^{15}C$	$5/2^{+}$	$1/2^+$	1.40	1.03	0.98(2) <sup>d</sup>

a Nuclei that have one or two sd-shell neutrons outside a p-shell core [\[1\]](#page-1-0).

<sup>b</sup>Reference [\[2\]](#page-1-0).

<sup>c</sup>Simple average of four most recent values [\[4–7\]](#page-1-0).

 ${}^{\text{d}}$ Reference [\[3\]](#page-1-0).

measurement of gamma width difficult. But, the inverse B(E2) transition strength can be deduced from Coulomb excitation measurements. Such experiments have been done. The relevant equations are

$$
2 B(E2; 1/2^- \to 5/2^-) = 6 B(E2; 5/2^- \to 1/2^-);
$$
  
 
$$
B(E2 \text{ in}^{17} \text{Ne}) = (e_p/e_n)^2 B(E2 \text{ in}^{17} \text{N}).
$$

By consideration of B(E2)'s and quadrupole moments in  $17O/I^{7}F$ , Lawson *et al.* [\[9\]](#page-1-0) determined effective charges of  $e_p = 1.5e$ ,  $e_n = 0.5e$  in the sd space. I use those values here. Thus, the expected B(E2;  $1/2^- \rightarrow 5/2^-$ ) in <sup>17</sup>Ne is 109  $e^2$ fm<sup>4</sup> (Table [II\)](#page-1-0).

Recently, Marganiec *et al.* [\[10\]](#page-1-0) studied dissociation of relativistic <sup>17</sup>Ne projectiles incident on targets of lead, carbon, and polyethylene, paying special attention to the excitation and decay of narrow resonant states in <sup>17</sup>Ne. Comparison of data from the heavy and light targets yielded cross sections and transition probabilities for the Coulomb excitations of these narrow states. In particular, the subsequent analysis produced a B(E2) value for  $1/2^- \rightarrow 5/2^-$  of 90(18) $e^2$ fm<sup>4</sup>. Those authors stated that their B(E2) suggested that the  $s^2$  intensity in <sup>17</sup>Ne ground state (g.s.) was either  $23^{+9}$ –6 or  $53^{+5}$ –9%, where I have read the uncertainties from their graph (their Fig. 7). The



FIG. 1. Ratios of experimental to calculated E2 transition strengths  $[1]$  in  $^{15,16}$ C and  $^{16,17}$ N are plotted vs. A. Points labeled Set 1 were obtained with a neutron effective charge of  $e_n = 0.5e$ ; Set 2 used  $e_n = (Z/A)e$ , where Z and A refer to the core.

<span id="page-1-0"></span>



smaller of their two values is consistent with our shell-model value of 0.284 [11]. Several other works [12–14] had suggested dominance of  $s^2$  in this state. Ozawa *et al.* [12] concluded that the experimental value of the  $17$ Ne interaction cross section required that the g.s. has protons mostly in the  $s_{1/2}$ orbit. Timofeyuk *et al.* [13] used a three-cluster  $({}^{15}O + p + p)$ generator coordinate model and concluded that the last two neutrons in  $17N$  and the two external protons in  $17Ne$  occupy  $s_{1/2}$  states rather than  $d_{5/2}$ . Nakamura *et al.* [14] computed Coulomb energies for the  $A = 17$ ,  $T = 3/2$  isobaric quartet. They concluded that the "last two protons in  $^{17}$ Ne (g.s.) occupy the  $s_{1/2}$  orbit." However, Millener [15], in addressing the asymmetry in  $\beta \pm$  decays of <sup>17</sup>N and <sup>17</sup>Ne, found dominance of  $d^2$  over  $s^2$  configurations. His  $s^2$  occupancies were 0.15 in  $17N$  and 0.22 in  $17Ne$ , whereas the other workers assumed the same configuration amplitudes for the mirror nuclei. Sherr and I computed the  $17N/\sqrt{17}$ Ne mirror energy difference as a function of the  $s^2/d^2$  ratio and concluded a value of 0.22 for the  $s<sup>2</sup>$  occupancy [16]. These g.s. results are summarized in Table III.

An earlier study of Coulomb excitation [17] yielded an E2 strength of  $124(18) e^{2} fm^{4}$  for this transition in <sup>17</sup>Ne. Reference [10] stated that this value should be corrected to 179(26)  $e^2fm^4$ because of an error of a factor of  $e^2$ . The various experimental strengths are listed in Table II. Reference [10] suggested that the incorrectly large strength in Ref. [17] might have been

TABLE III. Wave function of  $^{17}N$  and  $^{17}Ne$  (g.s.) from various sources.

Source	$s2$ occupancy	Reference
Shell model calculation <sup>a</sup>	0.28 s <sup>2</sup>	[11]
Interaction cross section	Mostly $s^2$	[12]
Three-cluster calculation	Predominantly $s^2$	$\lceil 13 \rceil$
Coulomb energies	Mostly $s^2$	[14]
$17N/17$ Ne mirror energy	$\beta \pm$ decays of <sup>17</sup> N and <sup>17</sup> Ne 0.15 s <sup>2</sup> in <sup>17</sup> N, 0.22 in <sup>17</sup> Ne	[15]
difference $B(E2)$ in <sup>17</sup> Ne	$0.22 s2$ in both <sup>17</sup> N and <sup>17</sup> Ne $0.23^{+0.09}$ <sub>-0.06</sub> or $0.53^{+0.05}$ <sub>-0.09</sub>	[16] [10]

<sup>a</sup>Gave good agreement with <sup>15</sup>N(t, p).

caused by the assumption of pure Coulomb excitation, when in fact nuclear excitation also contributes.

We note that the most recent  $B(E2)$  in <sup>17</sup>Ne is in good agreement with my prediction, whereas the earlier value (if the proposed correction is valid) disagrees. This agreement in  $17$ Ne increases the likelihood that the experimental B(E2) in  $17N$  is too small and strengthens the argument that this gamma width should be remeasured. It appears that the old  $17$ N strength depends solely on a branching ratio of 0.78(3) and a single measurement of the mean life of 11(2) ps, using the recoil-distance technique [18].

## **III. SUMMARY**

In summary, a recent measurement of B(E2;  $1/2^- \rightarrow 5/2^-$ ) in  $\frac{17}{10}$  is in good agreement with the calculation in a simple model that combines weak coupling, the shell model, and mirror symmetry. The  $s^2$  occupancy of <sup>17</sup>Ne (g.s.) that is deduced from the measured value is also in agreement with our earlier wave function. Results indicate that the corresponding strength in  $17N$  should be remeasured.

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