Structure and decays of nuclear three-body systems: The Gamow coupled-channel method in Jacobi coordinates

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Background: Weakly bound and unbound nuclear states appearing around particle thresholds are prototypical open quantum systems. Theories of such states must take into account configuration mixing effects in the presence of strong coupling to the particle continuum space.

Purpose: To describe structure and decays of three-body systems, we developed a Gamow coupled-channel (GCC) approach in Jacobi coordinates by employing the complex-momentum formalism. We benchmarked the complex-energy Gamow shell model (GSM) against the new framework.

Methods: The GCC formalism is expressed in Jacobi coordinates, so that the center-of-mass motion is automatically eliminated. To solve the coupled-channel equations, we use hyperspherical harmonics to describe the angular wave functions while the radial wave functions are expanded in the Berggren ensemble, which includes bound, scattering, and Gamow states.

Results: We show that the GCC method is both accurate and robust. Its results for energies, decay widths, and nucleon-nucleon angular correlations are in good agreement with the GSM results.

Conclusions: We have demonstrated that a three-body GSM formalism explicitly constructed in the cluster-orbital shell model coordinates provides results similar to those with a GCC framework expressed in Jacobi coordinates, provided that a large configuration space is employed. Our calculations for A = 6 systems and ²⁶O show that nucleon-nucleon angular correlations are sensitive to the valence-neutron interaction. The new GCC technique has many attractive features when applied to bound and unbound states of three-body systems: it is precise, is efficient, and can be extended by introducing a microscopic model of the core.

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I. INTRODUCTION

Properties of rare isotopes that inhabit remote regions of the nuclear landscape at and beyond the particle driplines are in the forefront of nuclear structure and reaction research [1–6]. The next generation of rare isotope beam facilities will provide unique data on dripline systems that will test theory, highlight shortcomings, and identify areas for improvement. The challenge for nuclear theory is to develop methodologies to reliably calculate and understand the properties and dynamics of new physical systems with different properties due to large neutron-to-proton asymmetries and low-lying reaction thresholds. Here, dripline systems are of particular interest, as they can exhibit exotic radioactive decay modes such as two-nucleon emission [7–13]. Theories of such nuclei must take into account their open quantum nature.

Theoretically, a powerful suite of *A*-body approaches based on internucleon interactions provides a quantitative description of light- and medium-mass nuclei and their reactions [14–16]. To unify nuclear bound states with resonances and scattering continuum within one consistent framework, advanced continuum shell-model approaches have been introduced [17–19]. Microscopic models of exotic nuclear states have been supplemented with a suite of powerful, albeit more phenomenological models, based on effective degrees of freedom such as cluster structures. While such models provide a "lower-resolution" picture of the nucleus, they can be extremely useful when interpreting experimental data, providing guidance for future measurements, and offer guidance for more microscopic approaches.

The objective of this work is to develop a new three-body method to describe both reaction and structural aspects of two-particle emission. A prototype system of interest is the two-neutron-unbound ground state (g.s.) of ²⁶O [13,20,21]. According to theory, ²⁶O exhibits dineutron-type correlations [21–25]. To describe such a system, the nuclear model should be based on a fine-tuned interaction capable of describing particle-emission thresholds, a sound many-body method, and the capability to treat bound and unbound states simultaneously.

If one considers bound three-body systems, few-body models are very useful [26], especially models based on the Lagrange-mesh technique [27] or cluster-orbital shell model (COSM) [28]. However, for the description of resonances, the outgoing wave function in the asymptotic region needs to be treated very carefully. For example, one can divide the coordinate space into internal and asymptotic regions, where the *R*-matrix theory [29,30], microscopic cluster model [31], and diagonalization of the Coulomb interaction [32] can be used. Other useful techniques include the Green function method [23] and complex scaling [33,34].

Our strategy is to construct a three-body framework for weakly bound and unbound systems similar to that of the Gamow shell model (GSM) [35], but with center-of-

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mass motion and asymptotic behavior treated precisely. The attractive feature of the GSM is that-by employing the Berggren ensemble [36]-it treats bound, scattering, and outgoing Gamow states on the same footing. Consequently, energies and decay widths are obtained simultaneously as the real and imaginary parts of the complex eigenenergies of the shell-model Hamiltonian [17]. In this study, we develop a three-body Gamow coupled-channel (GCC) approach in Jacobi coordinates with the Berggren basis. Since the Jacobi coordinates allow for the exact treatment of nuclear wave functions in both nuclear and asymptotic regions, and as the Berggren basis explicitly takes into account continuum effects, a comprehensive description of weakly bound three-body systems can be achieved. As the GSM is based on the COSM coordinates, a recoil term appears due to the center-of-mass motion. Hence, it is of interest to compare Jacobi- and COSM-based frameworks for the description of weakly bound and resonant nuclear states.

This article is organized as follows. Section II describes the models and approximations. In particular, it lays out the new GCC approach and GSM used for benchmarking and defines the configuration spaces used. The results for A = 6 systems and 26 O are reported in Sec. III. Finally, the summary and outlook are provided in Sec. IV.

II. THE MODEL

A. Gamow coupled-channel approach

In the three-body GCC model, the nucleus is described in terms of a core and two valence nucleons (or clusters). The GCC Hamiltonian can be written as

$$\hat{H} = \sum_{i=1}^{3} \frac{\hat{\boldsymbol{p}}_{i}^{2}}{2m_{i}} + \sum_{i>j=1}^{3} V_{ij}(\boldsymbol{r}_{ij}) - \hat{T}_{\text{c.m.}}, \qquad (1)$$

where V_{ij} is the interaction between cluster *i* and cluster *j*, including central, spin-orbit and Coulomb terms, and $\hat{T}_{c.m.}$ represents the kinetic energy of the center of mass.

The unwanted feature of three-body models is the appearance of Pauli-forbidden states arising from the lack of antisymmetrization between core and valence particles. In order to eliminate the Pauli-forbidden states, we implemented the orthogonal projection method [37–39] by adding to the GCC Hamiltonian the Pauli operator

$$\hat{Q} = \Lambda \sum_{c} |\varphi^{j_{c}m_{c}}\rangle \langle \varphi^{j_{c}m_{c}}|, \qquad (2)$$

where Λ is a constant and $|\varphi^{j_c m_c}\rangle$ is a two-body state involving forbidden single-particle (s.p.) states of core nucleons. At large values of Λ , Pauli-forbidden states appear at high energies, so that they are effectively suppressed.

In order to describe three-body asymptotics and to eliminate the spurious center-of-mass motion exactly, we express the GCC model in the relative (Jacobi) coordinates [15,30,39,40],

$$\mathbf{x} = \sqrt{\mu_{ij}} (\mathbf{r}_i - \mathbf{r}_j),$$

$$\mathbf{y} = \sqrt{\mu_{(ij)k}} \left(\mathbf{r}_k - \frac{A_i \mathbf{r}_i + A_j \mathbf{r}_j}{A_i + A_j} \right),$$
(3)



FIG. 1. Jacobi coordinates in a three-body system.

where \mathbf{r}_i is the position vector of the *i*th cluster, A_i is the *i*th cluster mass number, and μ_{ij} and $\mu_{(ij)k}$ are the reduced masses associated with \mathbf{x} and \mathbf{y} , respectively:

$$\mu_{ij} = \frac{A_i A_j}{A_i + A_j},$$

$$\mu_{(ij)k} = \frac{(A_i + A_j)A_k}{A_i + A_j + A_k}.$$
 (4)

As one can see in Fig. 1, Jacobi coordinates can be expressed as T and Y types, each associated with a complete basis set. In practice, it is convenient to calculate the matrix elements of the two-body interaction individually in T- and Y-type coordinates and then transform them into one single Jacobi set. To describe the transformation between different types of Jacobi coordinates, it is convenient to introduce the basis of hyperspherical harmonics [41,42]. The hyperspherical coordinates are constructed from a five-dimensional hyperangular coordinates Ω_5 and a hyperradial coordinate $\rho = \sqrt{x^2 + y^2}$. The transformation between different sets of Jacobi coordinates is given by the Raynal-Revai coefficients [43].

Expressed in hyperspherical harmonics, the total wave function can be written as [39]

$$\Psi^{JM\pi}(\rho,\Omega_5) = \rho^{-5/2} \sum_{\gamma K} \psi^{J\pi}_{\gamma K}(\rho) \mathcal{Y}^{JM}_{\gamma K}(\Omega_5), \qquad (5)$$

where *K* is the hyperspherical quantum number and $\gamma = \{s_1, s_2, s_3, S_{12}, S, \ell_x, \ell_y, L\}$ is a set of quantum numbers other than *K*. The quantum numbers *s* and ℓ stand for spin and orbital

angular momentum, respectively, $\psi_{\gamma K}^{J\pi}(\rho)$ is the hyperradial wave function, and $\mathcal{Y}_{\gamma K}^{JM}(\Omega_5)$ is the hyperspherical harmonic.

The resulting Schrödinger equation for the hyperradial wave functions can be written as a set of coupled-channel equations,

$$\left[-\frac{\hbar^2}{2m} \left(\frac{d^2}{d\rho^2} - \frac{(K+3/2)(K+5/2)}{\rho^2} \right) - \tilde{E} \right] \psi_{\gamma K}^{J\pi}(\rho)$$

+ $\sum_{K'\gamma'} V_{K'\gamma',K\gamma}^{J\pi}(\rho) \psi_{\gamma'K'}^{J\pi}(\rho)$
+ $\sum_{K'\gamma'} \int_0^\infty W_{K'\gamma',K\gamma}(\rho,\rho') \psi_{\gamma'K'}^{J\pi}(\rho') d\rho' = 0,$ (6)

where

$$V_{K'\gamma',K\gamma}^{J\pi}(\rho) = \left\langle \mathcal{Y}_{\gamma'K'}^{JM} \right| \sum_{i>j=1}^{3} V_{ij}(\boldsymbol{r}_{ij}) \left| \mathcal{Y}_{\gamma K}^{JM} \right\rangle$$
(7)

and

$$W_{K'\gamma',K\gamma}(\rho,\rho') = \left\langle \mathcal{Y}_{\gamma'K'}^{JM} \middle| \hat{Q} \middle| \mathcal{Y}_{\gamma K}^{JM} \right\rangle \tag{8}$$

is the nonlocal potential generated by the Pauli projection operator, (2).

In order to treat the positive-energy continuum space precisely, we use the Berggren expansion technique for the hyperradial wave function,

$$\psi_{\gamma K}^{J\pi}(\rho) = \sum_{n} C_{\gamma n K}^{J\pi M} \mathcal{B}_{\gamma n}^{J\pi}(\rho), \qquad (9)$$

where $\mathcal{B}_{\gamma n}^{J\pi}(\rho)$ represents an s.p. state belonging to the Berggren ensemble [36]. The Berggren ensemble defines a basis in the complex-momentum plane, which includes bound, decaying, and scattering states. The completeness relation for the Berggren ensemble can be written as

$$\sum_{\mathbf{n}\in b,d} \mathcal{B}_{\mathbf{n}}(k_{\mathbf{n}},\rho)\mathcal{B}_{\mathbf{n}}(k_{\mathbf{n}},\rho') + \int_{L^{+}} \mathcal{B}(k,\rho)\mathcal{B}(k,\rho')dk$$
$$= \delta(\rho - \rho'), \tag{10}$$

where *b* are bound states and *d* are decaying resonant (or Gamow) states lying between the real-*k* momentum axis in the fourth quadrant of the complex-*k* plane and the L^+ contour representing the complex-*k* scattering continuum. For numerical purposes, L^+ has to be discretized, e.g., by adopting the Gauss-Legendre quadrature [44]. In principle, the contour L^+ can be chosen arbitrarily as long as it encompasses the resonances of interest. If the contour L^+ is chosen to lie along the real-*k* axis, the Berggren completeness relation reduces to the Newton completeness relation [45] involving bound and real-momentum scattering states.

To calculate radial matrix elements with the Berggren basis, we employ exterior complex scaling [46], where integrals are calculated along a complex radial path:

$$\langle \mathcal{B}_{n} | V(\rho) | \mathcal{B}_{m} \rangle = \int_{0}^{R} \mathcal{B}_{n}(\rho) V(\rho) \mathcal{B}_{m}(\rho) d\rho + \int_{0}^{+\infty} \mathcal{B}_{n}(R + \rho e^{i\theta}) V(R + \rho e^{i\theta}) \times \mathcal{B}_{m}(R + \rho e^{i\theta}) d\rho.$$
(11)

For potentials that decrease as $O(1/\rho^2)$ (centrifugal potential) or more rapidly (nuclear potential), R should be sufficiently large to bypass all singularities and the scaling angle θ is chosen so that the integral converges (see Ref. [47] for details). As the Coulomb potential is not square integrable, its matrix elements diverge when $k_n = k_m$. A practical solution is provided by the so-called "off-diagonal method" proposed in Ref. [48]. Basically, a small offset $\pm \delta k$ is added to the linear momenta k_n and k_m of involved scattering wave functions, so that the resulting diagonal Coulomb matrix element converges. By diagonalizing the complex symmetric Hamiltonian, energies and decay widths are obtained as the real and imaginary parts of the complex eigenenergies.

B. Gamow shell model

In the GSM, expressed in COSM coordinates, one deals with the center-of-mass motion by adding a recoil term ($\hat{p}_1 \cdot \hat{p}_2/m_n A_{core}$) [28,35]. The GSM Hamiltonian is diagonalized in the basis of Slater determinants built from the one-body Berggren ensemble. In this case, it is convenient to deal with the Pauli principle by eliminating spurious excitations at a level of the s.p. basis. In practice, one just needs to construct a valence s.p. space that does not contain the orbits occupied in the core. It is equivalent to the projection technique used in GCC wherein the Pauli operator, (2), expressed in Jacobi coordinates has a two-body character. The treatment of the interactions is the same in the GSM and GCC method. In both cases, we use the complex scaling method to calculate matrix elements [47] and the off-diagonal method to deal with the Coulomb potential [48].

The two-body recoil term is treated in the GSM by expanding it in a truncated basis of the harmonic oscillator (HO). The HO basis depends on the oscillator length b and the number of states used in the expansion. As demonstrated in Refs. [44] and [49], GSM eigenvalues and eigenfunctions converge for a sufficient number of HO states, and the dependence of the results on b is very weak.

Let us note in passing that one has to be careful when using arguments based on the variational principle when comparing the performance of the GSM vs the GCC technique. Indeed, the treatment of the Pauli-forbidden states is slightly different in the two approaches. Moreover, the recoil effect in the GSM is not removed exactly. (There is no recoil term in the GCC approach, as the center-of-mass motion is eliminated through the use of Jacobi coordinates.)

C. Two-nucleon correlations

In order to study the correlations between the two valence nucleons, we utilize the two-nucleon density [50–52] $\rho_{nn'}(r,r',\theta) = \langle \Psi | \delta(r_1 - r) \delta(r_2 - r') \delta(\theta_{12} - \theta) | \Psi \rangle$, where r_1 , r_2 , and θ_{12} are defined in Fig. 1(a). In the following, we apply the normalization convention of Ref. [52] in which the Jacobian $8\pi^2 r^2 r'^2 \sin \theta$ is incorporated into the definition of $\rho_{nn'}$, i.e., it does not appear explicitly. The angular density of the two valence nucleons is obtained by integrating $\rho_{nn'}(r,r',\theta)$

over radial coordinates:

$$\rho(\theta) = \int \rho_{nn'}(r, r', \theta) dr dr'.$$
(12)

The angular density is normalized to 1: $\int \rho(\theta) d\theta = 1$.

While it is straightforward to calculate $\rho_{nn'}$ with COSM coordinates, the angular density cannot be calculated directly with the Jacobi T-type coordinates used to diagonalize the GCC Hamiltonian. Consequently, either one can calculate the density distribution $\rho_{\rm T}(x, y, \varphi)$ in T-type coordinates and then transform it to $\rho(r_1, r_2, \theta_{12})$ in COSM coordinates by using the geometric relations of Fig. 1(a) or—as we do in this study—one can apply the T-type-to-COSM coordinate transformation. This transformation [43] provides an analytical relation between hyperspherical harmonics in COSM coordinates $\mathcal{Y}_{\gamma'K'}^{JM}(\mathbf{r}'_1, \mathbf{r}'_2)$ and the T-type Jacobi coordinates $\mathcal{Y}_{\gamma'K'}^{JM}(\mathbf{x}', \mathbf{y}')$, where \mathbf{r}'_1 , \mathbf{r}'_2 , \mathbf{x}' , and \mathbf{y}' are

$$\mathbf{r}_{1}' = \sqrt{A_{i}}\mathbf{r}_{1},$$

$$\mathbf{r}_{2}' = \sqrt{A_{j}}\mathbf{r}_{2},$$

$$\mathbf{x}' = \mathbf{x} = \sqrt{\mu_{ij}}(\mathbf{r}_{1} - \mathbf{r}_{2}),$$

$$\mathbf{y}' = \sqrt{\frac{A_{i} + A_{j}}{\mu_{(ij)k}}}\mathbf{y} = \frac{A_{i}\mathbf{r}_{1} + A_{j}\mathbf{r}_{2}}{\sqrt{A_{i} + A_{j}}}.$$
 (13)

D. Model space and parameters

In order to compare approaches formulated in Jacobi and COSM coordinates, we consider model spaces defined by the cutoff value ℓ_{max} , which is the maximum orbital angular momentum associated with (r_1, r_2) in the GSM and (x, y) in the GCC method. The remaining truncations come from the Berggren basis itself.

The nuclear two-body interaction between valence nucleons has been approximated by the finite-range Minnesota force with the original parameters in Ref. [53]. For the core-valence Hamiltonian, we took a Woods-Saxon (WS) potential with parameters fitted to the resonances of the core + n system. The Coulomb interaction has been considered when valence protons are present. For both the GSM and the GCC approach, the forbidden states of core nucleons, (2), are generated by the s.p. levels of the core-valence WS (+ Coulomb) potential.

In the case of the GSM, we use the Berggren basis for the spd partial waves and an HO basis for channels with higher orbital angular momenta. For ⁶He, ⁶Li, and ⁶Be we assume a ⁴He core. For ⁶He and ⁶Be, in the GSM we took a complex-momentum contour defined by the segments $k = 0 \rightarrow 0.17 - 0.17i \rightarrow 0.34 \rightarrow 3$ (all in fm⁻¹) for the $p_{3/2}$ partial wave and $0 \rightarrow 0.5 \rightarrow 1 \rightarrow 3 \text{ fm}^{-1}$ for the remaining *spd* partial waves. For ⁶Li, we took the contours $0 \rightarrow 0.18 0.17i \rightarrow 0.5 \rightarrow 3 \text{ fm}^{-1}$ for the $p_{1/2}, 0 \rightarrow 0.15 - 0.14i \rightarrow$ $0.5 \rightarrow 3 \text{ fm}^{-1}$ for the $p_{3/2}$, and $0 \rightarrow 0.25 \rightarrow 0.5 \rightarrow 3 \text{ fm}^{-1}$ for the sd partial waves. Each segment was discretized with 10 points. This is sufficient for the energies and most the other physical quantities, but one may need more points to describe wave functions precisely, especially for the unbound resonant states that are affected by Coulomb interaction. Hence, we choose 15 points for each segment to calculate the two-proton angular correlation of the unbound ⁶Be. The HO basis was defined through the oscillator length b = 2 fm and the maximum radial quantum number $n_{\text{max}} = 10$. The WS parameters for the A = 6 nuclei are the depth of the central term, $V_0 = 47$ MeV; the spin-orbit strength, $V_{\text{s.o.}} = 30$ MeV; the diffuseness, a = 0.65 fm; and the WS (and charge) radius, R = 2 fm. With these parameters we predict the $3/2^-$ g.s. of ⁵He at E = 0.732 MeV ($\Gamma = 0.622$ MeV) and its first excited $1/2^-$ state at E = 2.126 MeV ($\Gamma = 5.838$ MeV). For ²⁶O, we consider the ²⁴O core [23,54,55]. In the

For ²⁶O, we consider the ²⁴O core [23,54,55]. In the GSM variant, we used the contour $0 \rightarrow 0.2 - 0.15i \rightarrow 0.4 \rightarrow 3 \text{ fm}^{-1}$ for $d_{3/2}$ and $0 \rightarrow 0.5 \rightarrow 1 \rightarrow 3 \text{ fm}^{-1}$ for the remaining *spd* partial waves. For the HO basis we took b = 1.75 fm and $n_{\text{max}} = 10$. The WS potential for ²⁶O was fitted in Ref. [23] to the resonances of ²⁵O. Its parameters are $V_0 = 44.1$ MeV, $V_{\text{s.o.}} = 45.87$ MeV, a = 0.73 fm, and R = 3.6 fm.

GCC calculations have been carried out with the maximal hyperspherical quantum number $K_{\text{max}} = 40$, which is sufficient for all the physical quantities we study. We checked that the calculated energies differ by as little as 2 keV when varying K_{max} from 30 to 40. Similarly to the GSM, in the GCC technique we used the Berggren basis for $K \leq 6$ channels and the HO basis for higher-angular-momentum channels. The complex-momentum contour of the Berggren basis is defined as $k = 0 \rightarrow 0.3 - 0.2i \rightarrow 0.5 \rightarrow 0.8 \rightarrow 1.2 \rightarrow 4$ (all in fm⁻¹), with each segment discretized with 10 points. We took the HO basis with b = 2 fm and $n_{\text{max}} = 20$. As $k_{\rho}^2 = k_x^2 + k_y^2$, the energy range covered by the GCC basis is roughly doubled compared to that of the GSM. For the constant Λ in the Pauli operator, (2), we use 10⁷ MeV. We have checked that our results do not depend on Λ in the range $\Lambda = 10^5$ to 10^9 MeV.

For the core-valence Coulomb potential, we use the dilatation-analytic form [49,56,57]

$$U_{\rm c}^{(Z)}(r) = e^2 Z_{\rm c} \frac{{\rm erf}(r/\nu_{\rm c})}{r},$$
 (14)

where $v_c = 4R_0/(3\sqrt{\pi})$ fm, R_0 is the radius of the WS potential, and Z_c is the number of core protons.

We emphasize that the large continuum space, containing states of both parities, is essential for the formation of the dineutron structure in nuclei such as ⁶He and ²⁶O [24,25,52,58–60]. In the following, we study the effect of including positive- and negative-parity continuum shells on the stability of threshold configurations.

III. RESULTS

A. Structure of A = 6 systems

We begin with the GCC-GSM benchmarking for A = 6 systems. Figure 2 shows the convergence rate for the g.s. energies of ⁶He, ⁶Li, and ⁶Be with respect to ℓ_{max} . (See Ref. [61] for a similar comparison between GSM and complex scaling results.) While the g.s. energies of ⁶He and ⁶Be are in reasonable agreement with experiment, that of ⁶Li is overbound. This is because the Minnesota interaction does not explicitly separate the T = 0 and T = 1 channels. The structure of ⁶He and ⁶Be is given by the T = 1 force, while



FIG. 2. Comparison between GSM and GCC results for the twonucleon separation energies of ⁶Be, ⁶Li, and ⁶He obtained in different model spaces defined by ℓ_{max} . Vertical bars in (a) represent decay widths.

the T = 0 channel, which is crucial for ⁶Li, has not been optimized. This is of minor importance for this study, as our goal is to benchmark the GCC method and GSM, not to provide quantitative predictions. As we use different coordinates in the GCC basis and the GSM, their model spaces are manifestly different. Still for $\ell_{max} = 10$ both approaches provide very similar results, which is most encouraging.

One can see in Fig. 2 that calculations done with Jacobi coordinates converge more rapidly than those done with COSM coordinates. This derives from the attractive character of the nucleon-nucleon interaction, which results in the presence of a dinucleon structure (see discussion below). Consequently, as T-type Jacobi coordinates well describe the dinucleon cluster, they are able to capture correlations in a more efficient way than COSM coordinates. This is in agreement with the findings in Ref. [34] based on the complex scaling method with COSM coordinates, which obtained a g.s. energy for ⁶He that was slightly less bound compared to the results in Ref. [39] using Jacobi coordinates. In any case, our calculations have demonstrated that one obtains very similar results in the GCC technique and GSM when sufficiently large model spaces are considered. As reported in Table I, the energy difference between GCC and GSM predictions for A = 6 systems is very small, around 20 keV for the majority of states. The maximum deviation of \sim 70 keV is obtained for the

TABLE I. Comparison between energies (in MeV) and widths (in keV) predicted for ⁶He, ⁶Li, and ⁶Be in the GSM and GCC approach in the $\ell_{max} = 10$ model space.

Nucleus	J^{π}	GSM	GCC
⁶ He	0^{+}	-0.933	-0.934
	2^{+}	0.800(98)	0.817(42)
⁶ Li	1^{+}	-5.680	-5.698
	3+	-2.097	-2.167
	0^+	-0.041	-0.048
⁶ Be	0^+	1.314(25)	1.275(54)

 3^+ state of ⁶Li. However, because of the attractive character of the T = 0 interaction, the GSM calculation for this state has not fully converged at $\ell_{\text{max}} = 10$.

Motivated by the discussion in Ref. [39], we have also studied the effect of the ℓ -dependent core-nucleus potential. To this end, we changed the WS strength V_0 from 47 to 49 MeV for $\ell = 1$ partial waves while keeping the standard strength for the remaining ℓ values. As shown in Fig. 3, the convergence behavior obtained with Jacobi and COSM coordinates is fairly similar to that shown in Fig. 2, where the WS strength V_0 is the same for all partial waves. For $\ell_{max} = 12$, the difference between the GSM and the GCC energies of ⁶He becomes very small. This result is consistent with the findings in Ref. [62] that the recoil effect can indeed be successfully eliminated using COSM coordinates at the expense of reduced convergence.

In order to see whether the difference between the model space of the GCC approach and that of the GSM can be compensated by renormalizing the effective Hamiltonian, we slightly readjusted the depth of the WS potential in GSM calculations to reproduce the g.s. GCC energy of ⁶He at the model space of $\ell_{\text{max}} = 7$. As a result, the strength V_0 changed from 47 to 47.05 MeV. Except for the 2⁺ state of ⁶He, the GSM and GCC energies for A = 6 systems became significantly closer as a result of this renormalization. This indicates that the differences between COSM and Jacobi coordinates can be partly accounted for by refitting the interaction parameters,



FIG. 3. Similar to Fig. 2, but for the two-neutron separation energy of ⁶He obtained with the angular-momentum dependent Hamiltonian; see text for details.

even though the model spaces and asymptotic behavior are different.

The GCC method is also in rough agreement with the GSM when comparing decay widths, considering that they are very sensitive to the asymptotic behavior of the wave function, which is treated differently with Jacobi and COSM coordinates. Also, the presence of the recoil term in the GSM, which is dealt with by means of the HO expansion, is expected to impact the GSM results for decay widths.

In order to check the precision of decay widths calculated with the GCC technique we adopted the current expression [63]

$$\Gamma = i \frac{\int (\Psi^{\dagger} \hat{H} \Psi - \Psi \hat{H} \Psi^{\dagger}) d\mathbf{x} d\mathbf{y}}{\int |\Psi|^2 d\mathbf{x} d\mathbf{y}},$$
(15)

which can be expressed in hyperspherical coordinates as [64,65]

$$\Gamma = i \frac{\hbar^2}{m} \frac{\int d\Omega_5 \text{Im} \left[\psi \frac{\partial}{\partial \rho} \psi^{\dagger} \right] \Big|_{\rho = \rho_{\text{max}}}}{\int_0^{\rho_{\text{max}}} |\psi|^2 d\rho d\Omega_5},$$
(16)

where ρ_{max} is larger than the nuclear radius (in general, the decay width should not depend on the choice of ρ_{max}). By using the current expression, we obtain $\Gamma = 42$ keV for the 2^+ state of ⁶He and $\Gamma = 54$ keV for the 0^+ state of ⁶Be, which are practically the same as the GCC values in Table I obtained from the direct diagonalization.

We now discuss the angular correlation of the two valence neutrons in the g.s. of ⁶He. Figure 4 shows GSM and GCC results for model spaces defined by different values of ℓ_{max} . The distribution $\rho(\theta)$ shows two maxima [24,34,51,52,62,66,67]. The higher peak, at a small opening angle, can be associated with a dineutron configuration. The second maximum, found in the region of large angles, represents a cigarlike configuration. The GCC results for $\ell_{max} = 2$ and 10 are already very close. This is not the case for the GSM, which shows sensitivity to the cutoff value of ℓ . This is because a large continuum space, including states of positive and negative parity, is needed in the COSM picture to describe dineutron correlations [25,52,58–60]. Indeed, as ℓ_{max} increases, the angular correlations obtained in the GSM and GCC method



FIG. 4. Comparison between GSM and GCC results for the twoneutron angular correlation in ⁶He for different model spaces defined by ℓ_{max} .

are very similar. This indicates that the Jacobi and COSM descriptions of $\rho(\theta)$ are essentially equivalent provided that the model space is sufficiently large.

In order to benchmark the GCC and GSM calculations for the valence-proton case, in Fig. 5 we compare two-nucleon angular correlations for the A = 6 nuclei ⁶He, ⁶Li, and ⁶Be. Similarly to Refs. [51] and [52], we find that the T = 1configurations have a dominant S = 0 component, in which the two neutrons in ⁶He or two protons in ⁶Be are in the spin singlet state. The amplitude of the S = 1 density component is low. For all nuclei, the GCC and GSM angular correlations are close.

Similarly to ⁶He, the two peaks in ⁶Be indicate diproton and cigarlike configurations [68] (see also Refs. [69–73]). It is noteworthy that the dineutron peak in ⁶He is slightly higher than the diproton maximum in ⁶Be. This is due to the repulsive character of the Coulomb interaction between valence protons. The large maximum at small opening angles seen in ⁶Li corresponds to a deuteronlike structure. As discussed in Ref. [66], this peak is higher than the dineutron peak in ⁶He. Indeed, the valence proton-neutron pair in ⁶Li is very strongly correlated because the T = 0 interaction is much stronger



FIG. 5. Two-nucleon angular densities (total and in the S = 1 channel) in the g.s. configurations of (a) ⁶He, (b) ⁶Li, and (c) ⁶Be obtained in the GSM and GCC approach with $\ell_{max} = 10$.



FIG. 6. Two-neutron separation energy of the g.s. of 26 O computed with the GSM and GCC method for different values of ℓ_{max} .

than the T = 1 interaction. The different features in the two-nucleon angular correlations in the three A = 6 systems shown in Fig. 5 demonstrate that the angular correlations contain useful information on the effective interaction between valence nucleons.

B. Structure of unbound ²⁶O

After benchmarking the GSM and GCC approach for A = 6 systems, we apply both models to ²⁶O, which is believed to be a threshold dineutron structure [13,20–25]. It is a theoretical challenge to reproduce the resonances in ²⁶O as both continuum and high partial waves must be considered. As ²⁴O can be associated with a subshell closure in which the $0d_{5/2}$ and $1s_{1/2}$ neutron shells are occupied [74], it can be used as the core in our three-body model.

Figure 6 illustrates the convergence of the g.s. of ²⁶O with respect to ℓ_{max} in GSM and GCC calculations. It is shown that in the GCC approach the energy converges nearly exponentially and that the stable result is practically reached at $\ell_{max} = 7$. While slightly higher in energy, the GSM results are quite satisfactory, as they differ only by about 30 keV from the GCC benchmark. Still, it is clear that $\ell_{max} = 12$ is not sufficient to reach full convergence in the GSM.

The calculated energies and widths of the g.s. and 2⁺ state of ²⁶O are listed in Table II; they are both consistent with the most recent experimental values [21]. The amplitudes of dominant configurations listed in Table II illustrate the importance of

TABLE II. Energies and widths (all in keV) predicted for ²⁶O in GSM and GCC in the $\ell_{max} = 12$ model space. Also listed are the dominant GSM (ℓ_1 , ℓ_2) and GCC (ℓ_x , ℓ_y) configurations.

J^{π}	G	SM	GCC	
0+	101	81% (d,d) 11% (f,f) 7% (p,p)	69	46% (p,p) 44% (s,s) 3% (d,d)
2+	1137(33)	77% (d,d) 7% (p,p) 7% (d,s)	1150(14)	28% (f,p) 27% (p,f) 10% (d,d)



FIG. 7. GCC wave function of the g.s. of ${}^{26}O$ in the Jacobi coordinates *nn* and ${}^{24}O-2n$.

considering partial waves of different parities in the GSM description of a dineutron g.s. configuration in ²⁶O [25].

The g.s. wave function of ²⁶O computed in the GCC approach is shown in Fig. 7 in Jacobi coordinates. The corresponding angular distribution is displayed in Fig. 8. Three pronounced peaks, associated with the dineutron, triangular, and cigarlike configurations [23,75], can be identified. In the GCC approach, the $(\ell_x, \ell_y) = (s, s)$, (p, p) components dominate the g.s. wave function of ²⁶O; this is consistent with a sizable clusterization of the two neutrons. In COSM coordinates, it is the $(\ell_1, \ell_2) = (d, d)$ configuration that dominates, but the negative-parity (f, f) and (p, p) channels contribute



FIG. 8. Two-neutron angular correlation for (a) the 0⁺ g.s. and (b) the 2⁺₁-state configurations of ²⁶O computed with the GCC approach (solid line) and GSM (dashed line) with $\ell_{max} = 10$. The dash-dotted curve labeled GCC' in (a) shows GCC results obtained with the strength of the neutron-neutron interaction reduced by 50%.

~20%. Again, it is encouraging to see that at $\ell_{max} = 10$ both approaches predict very similar two-nucleon densities.

In Table II we also list the predicted structure of the excited 2^+ state of 26 O. The predicted energy is close to experiment [21] and other theoretical studies (see, e.g., [22], [23], and [76–78]). We obtain a small width for this state, which is consistent with the GSM + density-matrix renormalizationgroup calculations in Ref. [25]. The GCC occupations in Table II indicate that the wave function of the 2^+ state is spread out in space, as the main three configurations, of the cluster type, contribute only 65% to the wave function. Considering the GSM wave function, the (d,d) configuration dominates. The corresponding two-neutron angular correlation shown in Fig. 8(b) exhibits a broad distribution with a maximum around 90°. This situation is fairly similar to what has been predicted for the 2^+ state of ⁶He [34,52].

Finally, it is interesting to study how the neutron-neutron interaction impacts the angular correlation. To this end, Fig. 8(a) shows $\rho(\theta)$ values obtained with the Minnesota neutron-neutron interaction, whose strength has been reduced by 50%. While there are still three peaks present, the distribution becomes more uniform and the dineutron component no longer dominates. We can thus conclude that the *nn* angular correlation can be used as an indicator of the interaction between valence nucleons.

IV. CONCLUSIONS

We have developed a Gamow coupled-channel (GCC) approach in Jacobi coordinates with the Berggren basis to describe structure and decays of three-body systems. We have benchmarked the performance of the Gamow shell model (GSM) against the new approach. Both methods are capable of considering large continuum spaces but differ in their treatment of the three-body asymptotics, center-of-mass motion, and Pauli operator. We have demonstrated that the Jacobi-coordinate-based framework (GCC) is accurate and robust for three-body systems. It is also encouraging to see that, in spite of these differences, the COSM-coordinate-based framework (GSM) yields fairly similar results, provided that the continuum space is sufficiently large.

For benchmarking and illustrative examples we choose ⁶He, ⁶Li, ⁶Be, and ²⁶O—all viewed as a core-plus-twonucleon systems. We discuss the spectra, decay widths, and nucleon-nucleon angular correlations in these nuclei. The Jacobi coordinates capture cluster correlations (such as dineutron and deuteron type) more efficiently; hence, the convergence rate of the GCC method is faster than that of the GSM.

For 26 O, we have demonstrated the sensitivity of the *nn* angular correlation to the valence-neutron interaction. It will be interesting to investigate this aspect further to provide guidance for future experimental investigations of dinucleon correlations in bound and unbound states of dripline nuclei.

In summary, we have developed an efficient approach to study of structure and decays of three-cluster systems. The GCC method is based on a Hamiltonian involving an interaction between valence nucleons and a core-nucleon potential. The advantage of the model is its ability to correctly describe the three-body asymptotic behavior and the efficient treatment of the continuum space, which is of particular importance for the treatment of threshold states and narrow resonances. The model can be easily extended along the lines of the resonating-group method by introducing a microscopic picture of the core [15,79]. Meanwhile, it can be used to elucidate experimental findings on dripline systems and to provide fine-tuned predictions to guide *A*-body approaches.

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- J. Dobaczewski and W. Nazarewicz, Philos. Trans. R. Soc. London A 356, 2007 (1998).
- [2] Scientific Opportunities with a Rare-Isotope Facility in the United States. Report of the NAS/NRC Rare Isotope Science Assessment Committee (National Academies Press, Washington, DC, 2007).
- [3] J. Dobaczewski, N. Michel, W. Nazarewicz, M. Płoszajczak, and J. Rotureau, Prog. Part. Nucl. Phys. 59, 432 (2007).
- [4] C. Forssén, G. Hagen, M. Hjorth-Jensen, W. Nazarewicz, and J. Rotureau, Phys. Scripta 2013, 014022 (2013).
- [5] A. B. Balantekin et al., Mod. Phys. Lett. A 29, 1430010 (2014).
- [6] The 2015 Long Range Plan in Nuclear Science: Reaching for the Horizon. NSAC Long Range Plan Report (Nuclear Science Advisory Committee, Washington, DC, 2015).

- [7] M. Pfützner, M. Karny, L. V. Grigorenko, and K. Riisager, Rev. Mod. Phys. 84, 567 (2012).
- [8] M. Pfützner, Phys. Scripta 2013, 014014 (2013).
- [9] M. Thoennessen, Rep. Prog. Phys. 67, 1187 (2004).
- [10] B. Blank and M. Płoszajczak, Rep. Prog. Phys. 71, 046301 (2008).
- [11] L. V. Grigorenko, I. G. Mukha, C. Scheidenberger, and M. V. Zhukov, Phys. Rev. C 84, 021303 (2011).
- [12] E. Olsen, M. Pfützner, N. Birge, M. Brown, W. Nazarewicz, and A. Perhac, Phys. Rev. Lett. 111, 139903(E) (2013).
- [13] Z. Kohley, T. Baumann, D. Bazin, G. Christian, P. A. DeYoung, J. E. Finck, N. Frank, M. Jones, E. Lunderberg, B. Luther, S. Mosby, T. Nagi, J. K. Smith, J. Snyder, A. Spyrou, and M. Thoennessen, Phys. Rev. Lett. **110**, 152501 (2013).

- [14] S. Elhatisari, D. Lee, G. Rupak, E. Epelbaum, H. Krebs, T. A. Lähde, T. Luu, and U.-G. Meißner, Nature 528, 111 (2015).
- [15] P. Navrátil, S. Quaglioni, G. Hupin, C. Romero-Redondo, and A. Calci, Phys. Scripta 91, 053002 (2016).
- [16] A. Kumar, R. Kanungo, A. Calci, P. Navratil, A. Sanetullaev, M. Alcorta, V. Bildstein, G. Christian, B. Davids, J. Dohet-Eraly, J. Fallis, A. T. Gallant, G. Hackman, B. Hadinia, G. Hupin, S. Ishimoto, R. Krucken, A. T. Laffoley, J. Lighthall, D. Miller, S. Quaglioni, J. S. Randhawa, E. T. Rand, A. Rojas, R. Roth, A. Shotter, J. Tanaka, I. Tanihata, and C. Unsworth, Phys. Rev. Lett. 118, 262502 (2017).
- [17] N. Michel, W. Nazarewicz, M. Płoszajczak, and T. Vertse, J. Phys. G 36, 013101 (2009).
- [18] G. Hagen, M. Hjorth-Jensen, G. R. Jansen, R. Machleidt, and T. Papenbrock, Phys. Rev. Lett. 109, 032502 (2012).
- [19] G. Papadimitriou, J. Rotureau, N. Michel, M. Płoszajczak, and B. R. Barrett, Phys. Rev. C 88, 044318 (2013).
- [20] E. Lunderberg, P. A. DeYoung, Z. Kohley, H. Attanayake, T. Baumann, D. Bazin, G. Christian, D. Divaratne, S. M. Grimes, A. Haagsma, J. E. Finck, N. Frank, B. Luther, S. Mosby, T. Nagi, G. F. Peaslee, A. Schiller, J. Snyder, A. Spyrou, M. J. Strongman, and M. Thoennessen, Phys. Rev. Lett. 108, 142503 (2012).
- [21] Y. Kondo, T. Nakamura, R. Tanaka, R. Minakata, S Ogoshi, N. A. Orr, N. L. Achouri, T. Aumann, H. Baba, F. Delaunay, P. Doornenbal, N. Fukuda, J. Gibelin, J. W. Hwang, N. Inabe, T. Isobe, D. Kameda, D. Kanno, S. Kim, N. Kobayashi, T. Kobayashi, T. Kubo, S. Leblond, J. Lee, F. M. Marques, T. Motobayashi, D. Murai, T. Murakami, K. Muto, T. Nakashima, N. Nakatsuka, A. Navin, S. Nishi, H. Otsu, H. Sato, Y. Satou, Y. Shimizu, H. Suzuki, K. Takahashi, H. Takeda, S. Takeuchi, Y. Togano, A. G. Tuff, M. Vandebrouck, and K. Yoneda, Phys. Rev. Lett. **116**, 102503 (2016).
- [22] L. V. Grigorenko and M. V. Zhukov, Phys. Rev. C 91, 064617 (2015).
- [23] K. Hagino and H. Sagawa, Phys. Rev. C 93, 034330 (2016).
- [24] K. Hagino and S. Sagawa, Few-Body Syst. 57, 185 (2016).
- [25] K. Fossez, J. Rotureau, N. Michel, and W. Nazarewicz, Phys. Rev. C 96, 024308 (2017).
- [26] E. Braaten and H. W. Hammer, Phys. Rep. 428, 259 (2006).
- [27] D. Baye, P. Descouvemont, and N. K. Timofeyuk, Nucl. Phys. A 577, 624 (1994).
- [28] Y. Suzuki and K. Ikeda, Phys. Rev. C 38, 410 (1988).
- [29] P. Descouvemont, E. Tursunov, and D. Baye, Nucl. Phys. A 765, 370 (2006).
- [30] A. E. Lovell, F. M. Nunes, and I. J. Thompson, Phys. Rev. C 95, 034605 (2017).
- [31] A. Damman and P. Descouvemont, Phys. Rev. C 80, 044310 (2009).
- [32] L. V. Grigorenko, T. D. Wiser, K. Mercurio, R. J. Charity, R. Shane, L. G. Sobotka, J. M. Elson, A. H. Wuosmaa, A. Banu, M. McCleskey, L. Trache, R. E. Tribble, and M. V. Zhukov, Phys. Rev. C 80, 034602 (2009).
- [33] S. Aoyama, T. Myo, K. Katō, and K. Ikeda, Prog. Theor. Phys. 116, 1 (2006).
- [34] A. T. Kruppa, G. Papadimitriou, W. Nazarewicz, and N. Michel, Phys. Rev. C 89, 014330 (2014).
- [35] N. Michel, W. Nazarewicz, M. Płoszajczak, and K. Bennaceur, Phys. Rev. Lett. 89, 042502 (2002).
- [36] T. Berggren, Nucl. Phys. A 109, 265 (1968).
- [37] S. Saito, Prog. Theor. Phys. 41, 705 (1969).

- PHYSICAL REVIEW C 96, 044307 (2017)
- [38] V. Kukulin and V. Pomerantsev, Ann. Phys. (NY) 111, 330 (1978).
- [39] P. Descouvemont, C. Daniel, and D. Baye, Phys. Rev. C 67, 044309 (2003).
- [40] P. Navrátil, G. P. Kamuntavičius, and B. R. Barrett, Phys. Rev. C 61, 044001 (2000).
- [41] M. Fabre de la Ripelle, Ann. Phys. 147, 281 (1983).
- [42] A. Kievsky, S. Rosati, M. Viviani, L. E. Marcucci, and L. Girlanda, J. Phys. G 35, 063101 (2008).
- [43] J. Raynal and J. Revai, Nuovo Cimento A 68, 612 (1970).
- [44] G. Hagen, M. Hjorth-Jensen, and N. Michel, Phys. Rev. C 73, 064307 (2006).
- [45] R. Newton, Scattering Theory of Waves and Particles (Springer-Verlag, Berlin, 1982).
- [46] B. Gyarmati and T. Vertse, Nucl. Phys. A 160, 523 (1971).
- [47] N. Michel, W. Nazarewicz, M. Płoszajczak, and J. Okołowicz, Phys. Rev. C 67, 054311 (2003).
- [48] N. Michel, Phys. Rev. C 83, 034325 (2011).
- [49] N. Michel, W. Nazarewicz, and M. Płoszajczak, Phys. Rev. C 82, 044315 (2010).
- [50] G. Bertsch and H. Esbensen, Ann. Phys. (NY) 209, 327 (1991).
- [51] K. Hagino and H. Sagawa, Phys. Rev. C 72, 044321 (2005).
- [52] G. Papadimitriou, A. T. Kruppa, N. Michel, W. Nazarewicz, M. Płoszajczak, and J. Rotureau, Phys. Rev. C 84, 051304 (2011).
- [53] D. Thompson, M. Lemere, and Y. Tang, Nucl. Phys. A 286, 53 (1977).
- [54] R. Kanungo, C. Nociforo, A. Prochazka, T. Aumann, D. Boutin, D. Cortina-Gil, B. Davids, M. Diakaki, F. Farinon, H. Geissel *et al.*, Phys. Rev. Lett. **102**, 152501 (2009).
- [55] C. Hoffman et al., Phys. Lett. B 672, 17 (2009).
- [56] S. Saito, Suppl. Prog. Theor. Phys. 62, 11 (1977).
- [57] R. I. Betan, A. T. Kruppa, and T. Vertse, Phys. Rev. C 78, 044308 (2008).
- [58] F. Catara, A. Insolia, E. Maglione, and A. Vitturi, Phys. Rev. C 29, 1091 (1984).
- [59] N. Pillet, N. Sandulescu, and P. Schuck, Phys. Rev. C 76, 024310 (2007).
- [60] K. Hagino and H. Sagawa, Phys. Rev. C 90, 027303 (2014).
- [61] H. Masui, K. Katō, N. Michel, and M. Płoszajczak, Phys. Rev. C 89, 044317 (2014).
- [62] M. V. Zhukov, B. V. Danilin, D. V. Fedorov, J. M. Bang, I. J. Thompson, and J. S. Vaagen, Phys. Rep. 231, 151 (1993).
- [63] J. Humblet and L. Rosenfeld, Nucl. Phys. 26, 529 (1961).
- [64] L. V. Grigorenko, R. C. Johnson, I. G. Mukha, I. J. Thompson, and M. V. Zhukov, Phys. Rev. Lett. 85, 22 (2000).
- [65] L. V. Grigorenko and M. V. Zhukov, Phys. Rev. C 76, 014008 (2007).
- [66] W. Horiuchi and Y. Suzuki, Phys. Rev. C 76, 024311 (2007).
- [67] Y. Kikuchi, K. Katō, T. Myo, M. Takashina, and K. Ikeda, Phys. Rev. C 81, 044308 (2010).
- [68] T. Oishi, K. Hagino, and H. Sagawa, Phys. Rev. C 90, 034303 (2014).
- [69] E. Garrido, D. Fedorov, H. Fynbo, and A. Jensen, Nucl. Phys. A 781, 387 (2007).
- [70] L. V. Grigorenko et al., Phys. Lett. B 677, 30 (2009).
- [71] L. V. Grigorenko, I. A. Egorova, R. J. Charity, and M. V. Zhukov, Phys. Rev. C 86, 061602 (2012).
- [72] I. A. Egorova, R. J. Charity, L. V. Grigorenko, Z. Chajecki, D. Coupland, J. M. Elson, T. K. Ghosh, M. E. Howard, H. Iwasaki, M. Kilburn *et al.*, Phys. Rev. Lett. **109**, 202502 (2012).

- [73] R. Álvarez-Rodríguez, A. S. Jensen, E. Garrido, and D. V. Fedorov, Phys. Scripta 2012, 014002 (2012).
- [74] K. Tshoo, Y. Satou, H. Bhang, S. Choi, T. Nakamura, Y. Kondo, S. Deguchi, Y. Kawada, N. Kobayashi, Y. Nakayama *et al.*, Phys. Rev. Lett. **109**, 022501 (2012).
- [75] D. Hove, E. Garrido, P. Sarriguren, D. V. Fedorov, H. O. U. Fynbo, A. S. Jensen, and N. T. Zinner, Phys. Rev. C 95, 061301 (2017).
- [76] A. Volya and V. Zelevinsky, Phys. Rev. C 74, 064314 (2006).
- [77] K. Tsukiyama, T. Otsuka, and R. Fujimoto, Prog. Theor. Exp. Phys. 2015, 093D01 (2015).
- [78] S. K. Bogner, H. Hergert, J. D. Holt, A. Schwenk, S. Binder, A. Calci, J. Langhammer, and R. Roth, Phys. Rev. Lett. 113, 142501 (2014).
- [79] Y. Jaganathen, N. Michel, and M. Płoszajczak, Phys. Rev. C 89, 034624 (2014).