B(E2) values in neutron-excess nuclei near A = 16

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A simple model is used to compute B(E2)'s in several nuclei that have one or two *sd*-shell neutrons and no *sd*-shell protons. The model works well for all six nuclei if I use later experimental values for ¹⁶C for which the measured B(E2) is about four to eight times earlier values.

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I. INTRODUCTION

In ¹⁶C, the B(E2) from the first-excited 2⁺ state to the ground state (g.s.) has had an interesting history, both experimentally and theoretically. The first experiments reported values that were remarkably small. A value of 0.63(11)(16 syst) e^2 fm⁴ was inferred from the 2⁺ lifetime of 77(14)(19 syst) ps measured by a recoil shadow method [1]. In another work from about the same time, inelastic scattering of ¹⁶C from ²⁰⁸Pb was used (via Coulomb-nuclear interference) to obtain the ratio of mass to charge deformation lengths, which were, in turn, converted into neutron and proton *E*2 matrix elements from which the *B*(*E*2) was computed as *B*(*E*2) = M_p^2 [2]. This result is stated as *B*(*E*2) = 0.28(6) Weiskopf units (W. u.) to be compared with 0.26(5)(7 syst) W. u. from the lifetime mentioned above. In terms of W. u. this *B*(*E*2) is the smallest ever measured for a 2_1^+ to g.s. transition in any nucleus.

The ground and first-excited states of ¹⁶C are very well described [3] in terms of excitations of two neutrons in the sd shell coupled to two proton holes in the 1p shell. Within this space, the E2 involves only neutron transitions and hence will depend on the neutron effective charge. It is a certainty, then, that the B(E2)—no matter how small—can be fitted by some value of the neutron effective charge as has been demonstrated [4]. Suzuki *et al.* [4], in a ${}^{14}C + n + n$ model, found they could fit the experimental value if they used a neutron effective charge of $e_n = 0.10 e$, compared to their value of $e_n = 0.16 e$ in ¹⁵C. Imai *et al.* [1] had suggested that different E2 amplitudes interfered destructively. Heyde et al. [5] explained that such destructive interference was extremely unlikely for the E2 between the lowest 0^+ and 2^+ states. They proposed the existence of an undiscovered 0^+ state about 800 keV below the 2^+ state that had the majority of the E2 strength. This was also an extremely unlikely scenario.

Horiuchi and Suzuki [6], also in a ${}^{14}C + n + n$ model, found that their calculated B(E2) was about twice the experimental value if they used the same neutron effective charge in ${}^{16}C$ and ${}^{15}C$. They pointed out that the inclusion of S = 1 components in the 0⁺ and 2⁺ wave functions caused a slight reduction in the calculated B(E2). Hagino and Sagawa [7] used the same e_n in ${}^{15}C$ and ${}^{16}C$ and were able to compute values in the range of 0.94–1.07 e^2 fm⁴. They pointed out that the dominance of the $(d_{5/2})^2$ configuration in the g.s. plays a crucial role in the agreement and that such $(d_{5/2})^2$ dominance was in disagreement with other works [4,6]. It also disagrees with our wave functions [3]. In a no-core shell-model calculation, Fujii *et al.* [8] found that they needed to use "dressed" single-particle energies in order to get the correct level ordering in ¹⁶C. Their dressed B(E2) was 0.84 e^2 fm⁴ (undressed was 1.30). Other groups [9–11] performed calculations that came close to reproducing the extremely small E2 strength in ¹⁶C. Thus, many different sets of workers were able to produce a small value for ¹⁶C.

On the other hand, systematics of other nearby nuclei would have led one to expect a much larger B(E2). A calculation by Sagawa et al. [12] provided values from 3.80 to $6.85 e^2$ fm⁴. Fortunately, another set of experiments soon followed, and they drastically changed the situation. Ong et al. [13] measured the 2^+ mean life to be 18.3(1.4)(4.8 syst)ps, about 25% of the previous value, and thus B(E2) = 2.6(0.2)(0.7 syst), about four times larger than earlier. Elekes et al. [14] reexamined their analysis and found a similar result of $B(E2) = 3.04 e^2 \text{ fm}^4$. Then, Wiedeking *et al.* [15] used the recoil distance method following the fusion reaction ${}^{9}\text{Be}({}^{9}\text{Be}, 2p)$ and measured $\tau = 11.7(20)$ ps, giving B(E2) = $4.15(73)e^2$ fm⁴. This value is larger even than those of Ong et al. and the reexamination of Elekes et al. The simple average of the three is about $3.26(43)e^2$ fm⁴, approximately 5.2 times the first values. These experimental values and the average of the last four are listed in Table I.

II. CALCULATIONS AND RESULTS

Weak-coupling considerations have worked extremely well in accounting for B(E2)'s in a particle-hole nucleus in terms of those in the particle and hole nuclei separately. For example, in the $1/2^-$ band of ¹⁹F, the B(E2)'s [17] are in excellent agreement [18] with those expected for a 4p-1h band in which the 4p's are the 0⁺ band of ²⁰Ne [19]. Many other examples exist.

A proton inelastic-scattering experiment [20] on ¹⁶C has confirmed that the g.s.-2⁺₁ transition is predominantly a neutron excitation. The deformation parameter $\beta_{pp'} = 0.47(5)$ determined in their experiment is consistent with the trend in other nearby nuclei. Because the ¹⁶C transition primarily involves *sd*-shell neutron transitions, I have examined lowlying *B*(*E*2)'s in other nearby nuclei for which the "action" is mostly in the *sd*-shell neutron space. I restrict attention to pure *E*2's connecting the g.s. to a low-lying state whose *J* value does not allow the presence of a competing *M*1. These *B*(*E*2)'s [17,21,22] are listed in Table II along with values of *M*(*E*2) defined as *B*(*E*2 : $i \rightarrow f$) = $M^2/(2J_i + 1)$.

TABLE I. Experimental $B(E2: 2^+ \rightarrow 0^+)$ (e^2 fm⁴) in ¹⁶C.

Reference	Value
Imai <i>et al.</i> [1]	0.63[11(15)]
Elekes <i>et al.</i> [2]	0.67(14)
Ong <i>et al.</i> [13]	2.6[2(7)]
Elekes <i>et al.</i> reexamination [14]	3.04
Wiedeking et al. [15]	4.15(73)
Petri et al. [16]	$4.21^{+0.34}_{-0.26}(\text{stat})^{+0.28}_{-0.24}$ syst
Simple average of the last four	3.50(30)

The cases of ¹⁸O and (to a lesser extent) ¹⁷O require special mention. The lowest 0^+ and 2^+ states of ${}^{18}O$ have long been known [23,24] to contain significant core-excited (mostly 4p-2h) components. Even though these collective components are only about 10% of the wave functions of the first two states, the strong 4p E2 causes a much larger contribution to the B(E2). In a fit [24] to many of the properties of the low-lying states of ¹⁸O, the fitted value of the $(sd)^2$ part of M(E2) was 3.91 e fm² Re-doing the least-squares fit, fit with the updated experimental value from Table II, provides M(E2) $\approx 4.03 \, e \, \text{fm}^2$ for the $(sd)^2$ part. As this value arises from only 90% of the wave functions (91% for the g.s. and 88% for 2_1^+), the "complete" M(E2) for two sd-shell neutrons would thus be $4.5e \,\mathrm{fm^2}$. Even before this renormalization, I would have expected $M({}^{16}\text{C})/M({}^{15}\text{C}) \ge M[{}^{18}\text{O}(sd)^2]/M({}^{17}\text{O})$ for several reasons:

- (1) The non- $(sd)^2$ part of ¹⁶C likely involves $(sd)^4 \times {}^{12}$ C, where $(sd)^4$ represents ²⁰O, and ²⁰O(2⁺ $\rightarrow 0^+$) is reasonably large [$M(E2) = 5.43(11)e \, \text{fm}^2$].
- (2) The E2 in 17 O is likely more collective (from core excitation) than is 15 C.
- (3) The ¹⁵C value may be slightly suppressed because the $5/2^+$ state is not pure single particle. Its spectroscopic factor [25] is only 0.69, and some other amplitudes

	TABLE III.	. Experimental	and	weak-coupling	values	of	M(E2)
(i	$n e fm^2$) and	their ratios.					

Transition	Nucleus	Mexp	$M_{ m wc}$	$M_{\rm exp}/M_{\rm wc}$
$\overline{s \leftrightarrow d}$	$^{17}\mathrm{O}(s \leftrightarrow d)$	2.90 ^a		≡ 1.00
	¹⁶ N	2.06(1)	2.05	1.00
			1.91 ^c	1.08 ^c
	¹⁵ C	2.42(4)	2.90	0.834(14)
			2.49 ^c	0.972(16) ^c
$(sd)^2 \leftrightarrow (sd)^2$	$^{18}\mathrm{O}(sd)^{2\mathrm{a,b}}$	4.50		
	^{17}N	3.67(37)	4.90	0.75(8)
			4.60 ^c	0.80(8) ^c
	¹⁶ C	4.18(18)	4.50	0.93(4)
			3.86 ^c	1.08(5) ^c

^aAfter removing a collective contribution.

^bRenormalized to 100% of the wave function.

^cValues on the second line use $e_n = (Z/A)e$, where Z, A refer to the core.

are present [26]. But they are not expected to have an appreciable impact on the M(E2). The $(sd)^3 \times {}^{12}C$ component will involve an E2 in ${}^{19}O$ that is also reduced by a $5/2^+$ spectroscopic factor [27] considerably less than unity.

However, from the initial values in Table I (with $M_{18} = 4.5e \text{ fm}^2$) values are $M_{^{16}\text{C}}/M_{^{15}\text{C}} = 0.73$ and $M_{^{18}}(sd)^2/M_{^{17}}=1.27$, clearly demonstrating that the earlier value of ^{16}C B(E2) was less than one-third of the expected value. With the average of the more recent values, the $^{16}\text{C}/^{15}\text{C}$ ratio is 1.67, clearly satisfying the expected inequality.

For a more quantitative computation, I have removed the collective part of the ¹⁷O E2 to get a "local" value for $M(s \leftrightarrow d)$, viz. 2.9*e* fm². With these two values now of $M[(sd)^2 \leftrightarrow (sd)^2]$ and $M(s \leftrightarrow d)$, I can compute M(E2) for the other

Nucleus	J_i^π	J_f^π	$E\gamma$ (MeV)	$\Gamma\gamma$ (eV)	$B(E2) (e^2 \text{ fm}^4)$	M(E2) (e fm ²)
¹⁸ O ^b	2^{+}	0^+	1.982	$2.35(6) \times 10^{-4}$	9.3(3)	6.8(1)
¹⁷ O ^c	$1/2^{+}$	$5/2^{+}$	0.8707	$2.55(3) \times 10^{-6}$	6.3(1)	3.55(2)
$^{17}N^{c}$	$5/2^{-}-$	$1/2^{-}-$	1.907	$4.6(9) \times 10^{-5}$	2.25(44)	3.67(37)
¹⁶ N ^c	$0^{-}-$	$2^{-}-$	0.1204	$8.7(1) \times 10^{-11}$	4.25(5)	2.06(1)
¹⁶ C	2+	0^{+}	1.766	$\tau(ps) = 77(14)(19 \text{ syst})^{d}$ $18.3(1.4)(4.8 \text{ syst})^{e}$ $11.7(20)^{f}$ $11.4_{-0.9}^{-0.9}(7 \text{ syst})$	3.50(30) ^g	4.18(18)
${}^{15}C^{h}$	$5/2^{+}$	$1/2^{+}$	0.7400	$1.75(5) \times 10^{-7}$	0.98(2)	2.42(4)

TABLE II. Experimental B(E2)'s in relevant nuclei.^a

^aNuclei that have one or two *sd*-shell neutrons outside a *p*-shell core.

^bReference [17].

- ^eReference [13].
- ^fReference [15].
- ^gAverage from Table I.

^hReference [21].

^cReference [19].

^dReference [1].

TABLE IV. Wave functions of relevant states in ¹⁸O, ¹⁷N, and ¹⁶C.

State		Wave-function amplitudes ^a						
	d^2	ds	Coll	dd'	d's			
$\overline{{}^{18}O(2^+_1)}$	0.774	0.485	-0.347	0.056	-0.204	24		
$^{17}N(5/2_1^-)$	0.787	0.572		0.054	-0.226	30		
${}^{16}\mathrm{C}(2^+_1)$	0.606	0.755		0.054	-0.244	3		
	d^2	s^2	Coll					
¹⁸ O(g.s.)	0.848	0.438	-0.297			24		
¹⁷ N(g.s.)	0.845	0.536				30		
¹⁶ C(g.s.)	0.682	0.731				3		

^aThe $1d_{5/2}$, $2s_{1/2}$, and $1d_{3/2}$ are denoted by *d*, *s*, and *d'*, respectively.

nuclei in Table II within the weak-coupling approximation and compare with the experimental values. The relevant formula is [28] as follows:

$$B[E2: (jJ_c)I_i \to (jJ_c')I_f] = (2I_f + 1)(2J_c + 1)W^2(jJ_cI_f2; I_iJ_c')B(E2: J_c \to J_c'),$$

where W is a Racah coefficient. Results are listed in Table III. We see that agreement with weak coupling is reasonably good for all cases—even for 16 C with the newer experimental values. The agreement for ${}^{17}N$ and ${}^{15}C$ is made better if I use local neutron effective charges to renormalize those cases relative to the others, viz. if I use $e_n = (Z/A)e$, where Z and A refer to the core. The M's from this renormalization are on the second line for each nucleus in Table III. The values for ^{15,16}C and ^{16,17}N are all within expectations for such a simple model. For ¹⁷N, our ratio of experimental to calculated M(E2) is 0.75(8) or 0.80(8). A realistic shell-model calculation [29] gave 0.82(8) for this ratio. So, a remeasurement of the ${}^{17}N\gamma$ width might be warranted. In ¹⁶C the ratio of d^2 to s^2 in the g.s. and the ratio of d^2 to ds in the 2⁺ [3] (Table IV) are somewhat different from those in ¹⁸O [24], but computing M(E2) with the different wave functions makes only a small difference. [The calculated ${}^{16}C M(E2)$ becomes larger by 3.7%.] Table IV also lists the relevant wave functions in ¹⁷N [30].

TABLE V. Dimensionles M(E2) ratios in ^{15,16}C and ^{16,17}N.

Transition	Quantity	Calculated ratio		Expt, ratio	
$\overline{s \leftrightarrow d} \\ (sd)^2 \leftrightarrow (sd)^2$	$M(^{15}C)/M(^{16}N)$	1.41 ^a	1.30 ^b	1.17(2)	
	$M(^{16}C)/M(^{17}N)$	0.913 ^a	0.838 ^b	1.14(10)	

^aUsing the constant effective charge and oscillator parameter. ^bUsing the local effective charge $e_n = (Z/A)e$.

It is possible to estimate the value of M(E2) in ¹⁶C without reference to ¹⁸O and ¹⁷O, both of which have collective admixtures of a type that is much less likely in the other nuclei in Table II. And, one can make this estimate by forming dimensionless ratios. The E2's in both ¹⁶N and ¹⁵C involve the transition $s \leftrightarrow d$. Simply from coupling coefficients, I expect to have $M(^{15}C)/M(^{16}N) = \sqrt{2}$. If I use the same *n* effective charge and oscillator parameter in the two nuclei, this ratio does not depend on either quantity. If I use the local effective charge $e_n = (Z/A)e$, this ratio is reduced by a factor of (6/14)/(7/15) = 0.918, resulting in an expected M(E2)ratio of 1.30 to be compared with the experimental ratio of 2.42/2.06 = 1.17.

In a similar fashion, the *E*2's in both ¹⁷N and ¹⁶C involve the transitions $(sd)^2 \leftrightarrow (sd)^2$. If I temporarily ignore the small differences in the $(sd)^2$ amplitudes in these two nuclei (Table IV), I expect to find $M({}^{16}C)/M({}^{17}N) = \sqrt{(5/6)}$ with a constant effective charge and oscillator parameter. With local effective charges the expectation is 0.838 for the M(E2)ratio. The experimental ratio is 1.14(10). Accounting for the differences in $(sd)^2$ amplitudes in Table IV causes an increase of 4% in ${}^{16}C$ relative to ${}^{17}N$. These expectations are listed in Table V.

III. SUMMARY

In conclusion, I have used a simple model to compute the B(E2)'s from the ground state of several nuclei in which the transition is dominated by transitions within the neutron *sd*-shell space. The model is satisfactory for ^{17,18}O, ^{16,17}N, and ^{15,16}C. For ¹⁶C, the most recent experimental B(E2) values agree with the expectation from weak coupling, whereas the earlier values were about a factor of 4–6 too small.

- [1] N. Imai et al., Phys. Rev. Lett. 92, 062501 (2004).
- [2] Z. Elekes et al., Phys. Lett. B 586, 34 (2004).
- [3] H. T. Fortune, M. E. Cobern, S. Mordechai, G. E. Moore, S. Lafrance, and R. Middleton, Phys. Rev. Lett. 40, 1236 (1978).
- [4] Y. Suzuki, H. Matsumura, and B. Abu-Ibrahim, Phys. Rev. C 70, 051302(R) (2004).
- [5] K. Heyde, L. Fortunato, and J. L. Wood, Phys. Rev. Lett. 94, 199201 (2005).
- [6] W. Horiuchi and Y. Suzuki, Phys. Rev. C 73, 037304 (2006); 74, 019901(E) (2006).
- [7] K. Hagino and H. Sagawa, Phys. Rev. C 75, 021301(R) (2007).
- [8] S. Fujii, T. Mizusaki, T. Otsuka, T. Sebe, and A. Arima, Phys. Lett. B 650, 9 (2007).

- [9] M. Takashina, Y. Kanada-En'yo, and Y. Sakuragi, Phys. Rev. C 71, 054602 (2005).
- [10] H.-L. Ma, B.-G. Dong, and Y.-L. Yan, Phys. Lett. B 688, 150 (2010).
- [11] L. Coragio, A. Covello, A. Gargano, and N. Itaco, Phys. Rev. C 81, 064303 (2010).
- [12] H. Sagawa, X. R. Zhou, X. Z. Zhang, and T. Suzuki, Phys. Rev. C 70, 054316 (2004).
- [13] H. J. Ong *et al.*, Phys. Rev. C 78, 014308 (2008); Eur. Phys. J. A 42, 393 (2009).
- [14] Z. Elekes, N. Aoi, Zs. Dombrádi, Zs. Fülöp, T. Motobayashi, and H. Sakurai, Phys. Rev. C 78, 027301 (2008).
- [15] M. Wiedeking et al., Phys. Rev. Lett. 100, 152501 (2008).

- [16] M. Petri *et al.*, Phys. Rev. C **86**, 044329 (2012).
- [17] D. R. Tilley, H. R. Weller, C. M. Cheves, and R. M. Chasteler, Nucl. Phys. A 595, 1 (1995).
- [18] M. Harvey, Nucl. Phys. 52, 542 (1964).
- [19] D. R. Tilley, C. M. Cheves, J. H. Kelley, S. Raman, and H. R. Weller, Nucl. Phys. A 636, 249 (1998).
- [20] H. J. Ong et al., Phys. Rev. C 73, 024610 (2006).
- [21] D. R. Tilley, H. R. Weller, and C. M. Cheves, Nucl. Phys. A 564, 1 (1993).
- [22] F. Ajzenberg-Selove, Nucl. Phys. A 523, 1 (1991).
- [23] T. Engeland, Nucl. Phys. 72, 68 (1965); P. J. Ellis and T. Engeland, Nucl. Phys. A 144, 161 (1970); T. Engeland and P. J. Ellis, *ibid.* 181, 368 (1972).

- [24] R. D. Lawson, F. J. D. Serduke, and H. T. Fortune, Phys. Rev. C 14, 1245 (1974).
- [25] J. D. Goss, P. L. Jolivette, C. P. Browne, S. E. Darden, H. R. Weller, and R. A. Blue, Phys. Rev. C 12, 1730 (1975).
- [26] H. T. Fortune and R. Sherr, Phys. Rev. C 72, 024319 (2005).
- [27] S. Sen, S. E. Darden, H. R. Hiddleston, and W. A. Yoh, Nucl. Phys. A 219, 429 (1974).
- [28] R. D. Lawson, *Theory of the Nuclear Shell Model* (Clarendon, Oxford, 1980), p 327.
- [29] E. K. Warburton and D. J. Millener, Phys. Rev. C 39, 1120 (1989).
- [30] H. T. Fortune, G. E. Moore, L. Bland, M. E. Cobern, S. Mordechai, R. Middleton, and R. D. Lawson, Phys. Rev. C 20, 1228 (1979).