# Robustness of the baryon-stopping signal for the onset of deconfinement in relativistic heavy-ion collisions

Yu. B. Ivanov<sup>1,2,\*</sup> and D. Blaschke<sup>3,4,†</sup>

<sup>1</sup>National Research Centre "Kurchatov Institute" (NRC "Kurchatov Institute"), 123182 Moscow, Russia <sup>2</sup>National Research Nuclear University "MEPhI" (Moscow Engineering Physics Institute), 115409 Moscow, Russia <sup>3</sup>Institute of Theoretical Physics, University of Wroclaw, 50-204 Wroclaw, Poland <sup>4</sup>Bogoliubov Laboratory of Theoretical Physics, Joint Institute for Nuclear Research Dubna, 141980 Dubna, Russia (Received 29 May 2015; published 24 August 2015)

The impact of the experimental acceptance, i.e. transverse-momentum ( $p_T$ ) cutoff and limited rapidity region, on the earlier predicted irregularity in the excitation function of the baryon stopping is studied. This irregularity is a consequence of the onset of deconfinement occurring in the compression stage of a nuclear collision and manifests itself as a wiggle in the excitation function of the reduced curvature ( $C_y$ ) of the net-proton rapidity distribution at midrapidity. It is demonstrated that the wiggle is a very robust signal of a first-order phase transition that survives even under conditions of a very limited acceptance. At the same time the  $C_y$  for pure hadronic and crossover transition scenarios become hardly distinguishable, if the acceptance cuts off too much of the low- $p_T$  proton spectrum and/or puts a rapidity window that is too narrow around midrapidity. It is found that the shape of the net-proton rapidity distribution near midrapidity depends on the  $p_T$  cutoff. This implies that the measurements should be taken at the same acceptance for all collision energies in order to reliably conclude the presence or absence of the irregularity.

DOI: 10.1103/PhysRevC.92.024916 PACS number(s): 25.75.Nq, 24.10.Nz

#### I. INTRODUCTION

The onset of deconfinement in relativistic heavy-ion collisions is now in the focus of theoretical and experimental studies of the equation of state (EoS) and the phase diagram of strongly interacting matter. This problem is one of the main motivations for the currently running beam-energy scan [1] at the Relativistic Heavy-Ion Collider (RHIC) at Brookhaven National Laboratory (BNL) and the low-energy-scan program [2] at the Super Proton Synchrotron (SPS) of the European Organization for Nuclear Research (CERN) as well as for constructing the Facility for Antiproton and Ion Research (FAIR) in Darmstadt [3] and the Nuclotron-Based Ion Collider Facility (NICA) in Dubna [4].

In Refs. [5–8], it was argued that the baryon stopping in nuclear collision can be a sensitive probe for the onset of deconfinement. Rapidity distributions of net protons were calculated [7,8] within a model of the three-fluid dynamics (3FD) [9] in scenarios with and without deconfinement transition. These calculations were performed employing three different types of EoS: a purely hadronic EoS [10] (hadr. EoS) and two versions of the EoS involving deconfinement [11]. The latter two versions are an EoS with a first-order phase transition (2-phase EoS) and one with a smooth crossover transition (crossover EoS).

It was found that 3FD predictions within the first-ordertransition scenario exhibit a peak-dip-peak-dip irregularity in the incident energy dependence of the form of the net-proton rapidity distributions in central collisions. At low energies, rapidity distributions have a peak at the midrapidity. With the incident energy rise it transforms into a dip, then again into a peak, and with further rising energy the midrapidity peak again changes into a dip, which already survives up to arbitrary high energies. The behavior of the type peak-dip-peak-dip in central collisions within the 2-phase-EoS scenario is very robust with respect to variations of the model parameters in a wide range. This behavior is in contrast with that for the hadronic-EoS scenario, where the form of distribution at midrapidity gradually evolves from one with a peak to one with a dip. The case of the crossover EoS is intermediate. Only a weak wiggle of the type of peak-dip-peak-dip takes place.

Experimental data also reveal a trend of the peak-dip-peak-dip irregularity in the energy range  $8A \text{ GeV} \leq$  $E_{\rm lab} \leq 40 A$  GeV, which is qualitatively similar to that in the first-order-transition scenario while quantitatively it differs. It is very likely that the quantitative discrepancy is due to the inadequacy of the model for the 2-phase EoS for which the onset of the phase transition lies at rather high densities, above 8 times saturation density at low temperatures. Recent models suggest an onset density of half that value [12-14], which would place the onset of the wiggle structure closer to the experimentally indicated position. However, the experimental trend is based on preliminary data at energies of 20A and 30A GeV. Therefore, updated experimental results at energies 20A and 30A GeV are badly needed to pin down the preferable EoS and to check the hint of wiggle behavior of the type peak-dip-peak-dip in the net-proton rapidity distributions. Moreover, it would be highly desirable if data in this energy range were taken within the same experimental setup and with the same experimental acceptance.

The calculations of Refs. [7,8] were performed assuming the acceptance for net protons to be wide enough to include almost all emitted particles. In practice, the extension of the acceptance beyond the range of  $0 < p_T < 2$  GeV/c does not practically change the net-proton rapidity distributions in the incident-energy range of interest. However, the actual

<sup>\*</sup>Y.Ivanov@gsi.de

<sup>†</sup>blaschke@ift.uni.wroc.pl

experimental acceptance can be narrower. For the NICA multipurpose-detector (MPD) experiment it is restricted by the proton identification capabilities in the time-of-flight (TOF) detector to  $0.4 \, {\rm GeV/c} < p_T < 1.0 \, {\rm GeV/c}$  in the central rapidity range |y| < 0.5 [17]. For the case of the STAR beam energy scan the analysis of the excitation function of the net-proton rapidity distribution is still under way. In this case the acceptance will be restricted to the range  $0.4 \, {\rm GeV/c} < p_T < 3.0 \, {\rm GeV/c}$  and |y| < 0.5 [18].

The purpose of the present paper is to investigate the question whether the peak-dip-peak-dip irregularity survives when experimental circumstances force a narrowing of the accessible acceptance region.

#### II. EQUATIONS OF STATE

Figure 1 illustrates the differences between the three considered EoS. The deconfinement transition makes an EoS softer at high temperatures and/or densities. The 2-phase EoS is based on the Gibbs construction, taking into account simultaneous conservation of baryon and strange charges. However, the displayed result looks very similar to the Maxwell construction, corresponding to the conservation of just the baryon charge, with the only difference that the plateau is slightly tilted, which is practically invisible.

As demonstrated in Refs. [7,8], the deconfinement transition in central Au+Au collisions starts at the top AGS energies in both cases. It gets practically completed at low SPS energies in the case of the 2-phase EoS. In the crossover scenario it lasts until very high incident energies. We would like to note that the density range for the 2-phase EoS in Fig. 1 with an onset of deconfinement above  $n \sim 8n_0$  is rather high. In recent models for the spinodal decomposition accompanying the phase transition [12–14] the onset of the phase coexistence is at about  $4n_0$ . This corresponds to a lower limit for the

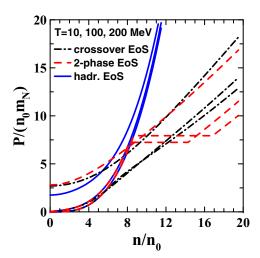


FIG. 1. (Color online) Pressure scaled by the product of normal nuclear density ( $n_0 = 0.15 \text{ fm}^{-3}$ ) and nucleon mass ( $m_N$ ) vs baryon density scaled by the normal nuclear density for three considered equations of state. Results are presented for three different temperatures T = 10, 100, and 200 MeV (from bottom upwards for corresponding curves).

onset of the deconfinement phase transition in cold, symmetric nuclear matter, which was obtained by a hybrid EoS model from constraints on the occurrence of quark matter phases in massive neutron stars [15]; see also Ref. [16].

#### III. IRREGULARITY AT CONSTRAINED ACCEPTANCE

The calculations at all collision energies were performed for  $\operatorname{Au+Au}(b=2\,\mathrm{fm})$  central collisions despite the fact that some experimental data were taken for central Pb+Pb collisions. This was done in order to avoid uncertainties associated with different colliding nuclei. However, in fact at the same incident energy the computed results for Pb+Pb collisions at  $b=2.4\,\mathrm{fm}$  are very close to those for  $\operatorname{Au+Au}$  at  $b=2\,\mathrm{fm}$ . The calculations were performed for four different acceptance ranges for the transverse momentum  $(p_T)$  and rapidity (y) of the measured proton:

- (i)  $0 < p_T < 2 \text{ GeV/c}$  and a very unrestrictive constraint to the rapidity range  $|y| < 0.7 \ y_{\text{beam}}$ , where  $y_{\text{beam}}$  is the beam rapidity in the collider mode, which is practically equivalent to the full acceptance;
- (ii)  $0.4 < p_T < 1$  GeV/c and |y| < 0.5, the expected MPD acceptance [17];
- (iii)  $1 < p_T < 2$  GeV/c and |y| < 0.5, an acceptance range where low-momentum particles witnessing collective behavior are largely eliminated;
- (iv)  $0.4 < p_T < 3 \text{ GeV/c}$  and |y| < 0.5, the range of the STAR acceptance [18].

We separately study effects of the  $p_T$  and y constraints in order to reveal their relative importance.

A direct measure of the baryon stopping is the net-baryon (i.e., baryons minus antibaryons) rapidity distribution. However, since experimental information on neutrons is unavailable, we have to rely on net-proton (i.e., proton minus antiproton) data. Presently there exist experimental data on proton (or net-proton) rapidity spectra at AGS [19–22] and SPS [23–27] energies. At AGS energies, the yield of antiprotons is negligible, and therefore the proton rapidity spectra serve as a good probe of the baryon stopping.

In order to quantify the previously discussed peak-dippeak-dip irregularity, it is useful to make use of the method proposed in Ref. [5]. For this purpose, the data on the net-proton rapidity distributions are fitted by a simple formula

$$\frac{dN}{dy} = a(\exp\{-(1/w_s)\cosh(y - y_s)\} + \exp\{-(1/w_s)\cosh(y + y_s)\}),$$
(1)

where a,  $y_s$ , and  $w_s$  are parameters of the fit. The form (1) is a sum of two thermal sources shifted by  $\pm y_s$  from the midrapidity, which is put to be  $y_{\rm mid} = 0$  as it is in the collider mode. The width  $w_s$  of the sources can be interpreted as  $w_s = ({\rm temperature})/({\rm transverse\ mass})$ , if we assume that collective velocities in the sources have no spread with respect to the source rapidities  $\pm y_s$ . The parameters of the two sources are identical (up to the sign of  $y_s$ ) because only collisions of identical nuclei are considered.

Here we apply the two-source fit, though results of simulations sometime indicate a need of a three-source fit

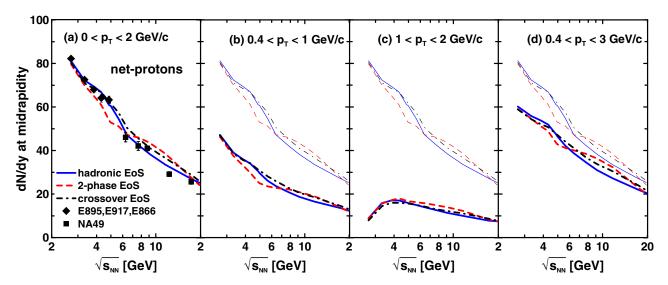


FIG. 2. (Color online) Midrapidity value of the net-proton rapidity distribution as a function of the collision energy in central (b = 2 fm) Au+Au collisions in different windows of the transverse momentum  $p_T$ : (a)  $0 < p_T < 2$  GeV/c, (b)  $0.4 < p_T < 1$  GeV/c, (c)  $1 < p_T < 2$  GeV/c, and (d)  $0.4 < p_T < 3$  GeV/c. Results for the wide  $p_T$  window ( $0 < p_T < 2$  GeV/c) are also presented in panels (b)–(d) by the corresponding thin lines for the sake of comparison. Results of simulations with different EoS's are presented. Experimental data are from the collaborations E895 [19], E877 [20], E917 [21], E866 [22], NA49 [23–27], and STAR [28].

(see below). At the same time all available experimental data manifest either one or two peaks in the net-proton rapidity distributions and thus are well approximated by the two-source fit. Therefore, we inclined to consider the indications of a three-source fit as an artifact of a simplified treatment of the complicated nonequilibrium stage of the collision within the three-fluid approximation.

The above fit has been prepared using the least-squares method and applied to both available data and results of calculations. The fit was performed in the rapidity range  $|y| < 0.7 y_{\text{beam}}$ , where  $y_{\text{beam}}$  is the beam rapidity in the collider mode. The choice of this range is dictated by the data. As a rule, the data are available in this rapidity range, although sometimes the data range is even more narrow (80A GeV and new data at 158A GeV [27]). We apply the above restriction in order to deal with different data in approximately the same rapidity range. Another reason for this cut is that the rapidity range should not be too wide in order to exclude contributions of cold spectators. The fit in the rapidity range  $|y| < 0.5 y_{\text{beam}}$ has been also done in order to estimate uncertainty of the fit parameters associated with the choice of fit range. An additional set of fits has been done under the constraint |y| < 0.5. This fit was applied only to the 3FD simulation results. Due to low number of experimental points within the |y| < 0.5 range and their insufficient accuracy such a fit gives too large error bars for the deduced parameters of the fit (1) at  $\sqrt{s_{NN}} > 5$  GeV, whereas it is practically identical to the  $|y| < 0.7 y_{\text{beam}}$  fit at  $\sqrt{s_{NN}} < 5$  GeV.

A useful quantity, which characterizes the shape of the rapidity distribution, is a reduced curvature of the spectrum at midrapidity, defined as follows:

$$C_{y} = \left(y_{\text{beam}}^{3} \frac{d^{3} N}{dy^{3}}\right)_{y=0} / \left(y_{\text{beam}} \frac{d N}{dy}\right)_{y=0}$$
$$= (y_{\text{beam}} / w_{s})^{2} (\sinh^{2} y_{s} - w_{s} \cosh y_{s}). \tag{2}$$

The factor  $1/(y_{\text{beam}}dN/dy)_{y=0}$  is introduced in order to cancel out the overall normalization of the spectrum. The second part of Eq. (2) presents this curvature in terms of the parameters of the fit (1). The reduced curvature,  $C_y$ , and the midrapidity value,  $(dN/dy)_{y=0}$ , are two independent quantities quantifying the the spectrum in the midrapidity range. Excitation functions of these quantities deduced both from experimental data and from results of the 3FD calculations with different EoS's are displayed in Figs. 2 and 3.

In Fig. 2 the midrapidity values of the rapidity spectra were taken directly from experimental data and calculated results. Therefore, only experimental error bars are displayed there. All presently available experimental data shown in the leftmost panel of Fig. 2 approximately correspond to acceptance range (i). As seen from Fig. 2, approximately 60% (for different collision energies and EoS's) of the protons produced at midrapidity will be covered by the MPD acceptance window, while the corresponding coverage of the STAR acceptance is ~80%.

The reduced curvature  $C_{\nu}$ , displayed in Fig. 3, was deduced from fit (1) of both the experimental data and results of simulations. The fit of the experimental data was done in the rapidity range  $|y| < 0.7 y_{\text{beam}}$ , where it is possible, and in a narrow range, if the data are unavailable in the above range. The fit of the available experimental data in the rapidity range |y| < 0.5 is not presented because of huge error bars resulting from very restricted number of experimental points near midrapidity and their statistical errors. Therefore, all presently available experimental data shown in the top panels of Fig. 3 approximately correspond to acceptance range (i). To evaluate errors of  $C_v$  values deduced from data, errors produced by the least-squares method, as well as performed fits in different the rapidity ranges, where it is appropriate, were estimated. The errors of the least-squares method, in particular, result from error bars of the experimental points. The error bars of the experimental points in Fig. 3 present

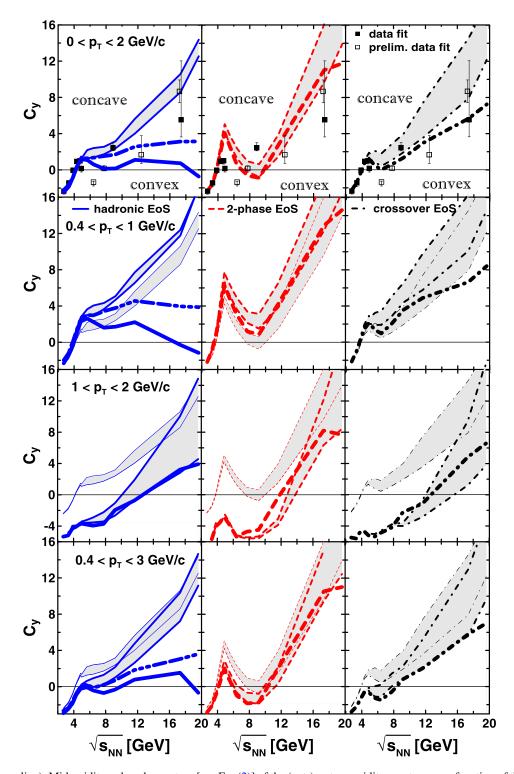


FIG. 3. (Color online) Midrapidity reduced curvature [see Eq. (2)] of the (net-)proton rapidity spectrum as a function of the center-of-mass energy of colliding nuclei as deduced from experimental data and predicted by 3FD calculations with different EoS's: the hadronic EoS (hadr. EoS) [10] (left column of panels), the EoS involving a first-order phase transition (2-ph. EoS, middle column of panels), and the EoS with a crossover transition (crossover EoS, right column of panels) into the quark-gluon phase [11]. Upper bounds of the shaded areas correspond to fits confined in the region of |y| < 0.7  $y_{\text{beam}}$ , and lower bounds correspond to |y| < 0.5  $y_{\text{beam}}$ . Results are presented for four different windows of the transverse momentum  $p_T$ :  $0 < p_T < 2$  GeV/c (top row of panels),  $0.4 < p_T < 1$  GeV/c (second row of panels),  $1 < p_T < 2$  GeV/c (third row of panels), and  $0.4 < p_T < 3$  GeV/c (bottom row of panels). Results for the  $p_T$  window of the top panel with experimental data ( $0 < p_T < 2$  GeV) are also presented in the lower panels with different restricted  $p_T$  windows by shaded areas bounded by the corresponding thin lines for the sake of comparison. Results of fits within the range of |y| < 0.5 are displayed by corresponding bold lines. In those cases, when (|y| < 0.5) fit substantially differs from that in the |y| < 0.9 range, the (|y| < 0.9) results are also displayed by bold double-dash-triple-dotted lines.

largest uncertainties among mentioned above. The uncertainty associated with the choice of the rapidity range turned out to be the dominant one for the  $C_y$  quantities deduced from simulation results. Therefore, in Fig. 3 results for the curvature  $C_y$  in the wide rapidity range are presented by shaded areas with borders corresponding to the fit ranges |y| < 0.7  $y_{\text{beam}}$  and |y| < 0.5  $y_{\text{beam}}$ . The  $C_y$  results in the narrow rapidity range |y| < 0.5, corresponding to the MPD and STAR acceptance, are also displayed by bold lines. In order to control the error induced by the narrow range |y| < 0.5, the  $C_y$  calculations were also performed in the range |y| < 0.9. In most cases the (|y| < 0.5) and (|y| < 0.9) results turned out to be very close to each other. In those few cases, when (|y| < 0.5) and (|y| < 0.9) results substantially differ, the (|y| < 0.9) results are also displayed in Fig. 3.

The irregularity in the data is distinctly seen as a strong wiggle in the excitation function of  $C_y$ ; see Fig. 3. Various data used to deduce  $C_y$  approximately correspond to the wide acceptance window (i). Of course, this is only a hint of an irregularity since this wiggle is formed only in preliminary data of the NA49 Collaboration. In the wide acceptance window (i) (shaded bands in the upper row of panels in Fig. 3) the  $C_y$  excitation function in the first-order-transition scenario manifests qualitatively (though not quantitatively) the same wiggle irregularity (middle-column upper-row panel in Fig. 3) like that in the data fit, while the hadronic scenario produces purely monotonous behavior. The crossover EoS represents a very smooth transition. Therefore, it is not surprising that it produces only a weak wiggle in  $C_y$ .

The application of various  $p_T$  cuts without confining the rapidity range (shaded bands in Fig. 3) does not qualitatively change the picture. As seen from Fig. 3, the wiggle in the energy dependence of  $C_y$  is a very robust signal of the first-order phase transition. It survives even at very limited  $p_T$  acceptance. The important difference between different  $p_T$  acceptances is that the wiggle is completely located in the range of positive curvatures  $C_y$  (concave shapes of the rapidity distribution near midrapidity) at low- $p_T$  acceptance, like the MPD one (ii). While for the high- $p_T$  acceptance (iii), the wiggle entirely lies in the range of negative  $C_{\nu}$  (convex shape). Therefore, the name of peak-dip-peak-dip for this irregularity is not quite correct. The amplitude of the wiggle becomes somewhat weaker when the  $p_T$  cutoff from below is too strong, e.g., when  $1 < p_T < 2$  GeV/c. It is expected because the wiggle is an effect of collective behavior of the system, in which predominantly low-momentum particles participate. If these low-momentum particles are cut off by the acceptance, the collective effects, in particular, the wiggle, become less manifested.

The  $C_y$  excitation functions for hadronic and crossover scenarios are very similar to each other already in the wide acceptance window (i) (the upper raw of panels in Fig. 3). The crossover scenario results in a very weak wiggle in  $C_y$ . Under the MPD (ii) and STAR (iv)  $p_T$  cuts the basic features of the  $C_y$  excitation functions remain the same, only the weak crossover wiggle turns out to be shifted to higher (the MPD case) or lower (the STAR case)  $C_y$ . At the high- $p_T$  cut (iii) the qualitative difference between predictions of the hadronic and crossover scenarios is practically washed out. This a consequence of the

above-mentioned lack of collectivity in the behavior of the high-momentum particles.

The actual MPD (ii) and STAR (iv) acceptance conditions also include the restriction |y| < 0.5. Applying this restriction in addition to the  $p_T$  cuts leads to the corresponding bold lines in Fig. 3. In order to study how restrictive the condition |y| < 0.5 is, calculations were also done for the rapidity window |y| < 0.9. In most cases the  $C_y$  results for the range |y| < 0.5 practically coincide with those in the |y| < 0.9 window and are quite close to those in the wide rapidity window |y| < 0.5  $y_{\text{beam}}$ . In these cases the |y| < 0.9 results are not displayed in Fig. 3. In a few cases the |y| < 0.5 and |y| < 0.9 results considerably differ. These cases correspond to the hadronic EoS when low- $p_T$  protons are included in the analysis. For these cases the |y| < 0.9 results are displayed by bold double-dash-dash-triple-dotted lines in Fig. 3 (in contrast to solid bold lines for the |y| < 0.5 results).

As seen from Fig. 3, under the constraints of the MPD (ii) and STAR (iv) acceptance conditions, the wiggle in the energy dependence of  $C_y$  is very robust for the first-order phase transition and the crossover one. However, this is not the case for the hadronic EoS. The hadronic-EoS  $C_v$  excitation functions now exhibit a local weak wiggle in the region 8  $<\sqrt{s_{NN}}<$  12 GeV (in all cases except for that with the high- $p_T$  range  $1 < p_T < 2 \text{ GeV/c}$ ) qualitatively similar to that for the crossover transition. Moreover, the  $C_y$  curvature becomes negative at  $<\sqrt{s_{NN}}=19.6$  GeV (again in all cases except for that with the high- $p_T$  range) that contradicts natural expectations that the net-proton distribution at midrapidity proceeds from convex shape (strong stopping) to a concave one (an increasing transparancy) with the beam energy rise. These peculiarities are a consequence of a fine structure of the rapidity distribution near midrapidity that becomes dominant in the narrow rapidity window |y| < 0.5. This situation is illustrated in Fig. 4. It is seen that at  $\sqrt{s_{NN}} = 7.7$  and 19.6 GeV there are tiny maxima at the midrapidity in the hadronic-EoS rapidity distributions, which result in the weak wiggle and negative curvature, respectively, observed in Fig. 3. The analysis of the net-proton distributions in the narrow rapidity window reveals only a fine structure of the rapidity distribution near midrapidity, which could be just an artifact of a simplified treatment of the complicated nonequilibrium stage of the collision based on the three-fluid approximation. In order to conclude on the baryon stopping we need to analyze the global shape of the rapidity distributions. The global shape of the hadronic-EoS rapidity distributions does not exhibit such peculiarities, as it is clear from  $C_{\nu}$  calculations in a wider rapidity range. For a comparison, see also Fig. 4 of Ref. [29].

Therefore, we can conclude that under the conditions of the MPD and STAR acceptance it is possible to distinguish the first-order phase transition, the onset of which is signalled by a strong wiggle in the excitation function of  $C_y$ . However, the difference between the purely hadronic case and that with the smooth crossover transition becomes ambiguous.

### IV. CONCLUSION

An irregularity in the baryon stopping is a natural consequence of deconfinement occurring in the compression stage

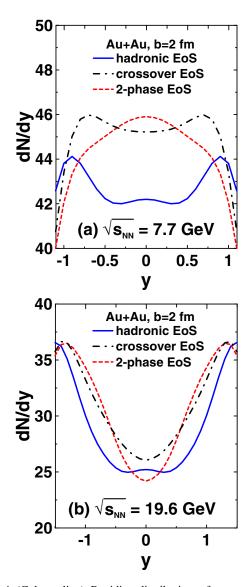


FIG. 4. (Color online) Rapidity distribution of net protons in central (b=2 fm) Au+Au collisions at collision energies  $\sqrt{s_{NN}}=7.7$  GeV (a) and 19.6 GeV (b) for the case of the  $p_T$  range  $0 < p_T < 2$  GeV/c.

of a nuclear collision. It is a combined effect of the softest point [30,31] of an EoS and a change in the nonequilibrium regime from hadronic to partonic one. As was demonstrated in Refs. [7,8], this irregularity manifests itself as a wiggle in the excitation function of a reduced curvature ( $C_y$ ) of the netproton rapidity distribution at midrapidity. These calculations were performed in the full acceptance range for protons. In the present paper we studied the effect of a restricted acceptance, in particular, the one expected for the NICA MPD experiment and the one of the STAR beam energy scan program, on the earlier predicted  $C_y$  wiggle.

It was found that the wiggle in the excitation function of  $C_y$  is a quite robust signal of the onset of deconfinement as a first-order phase transition. It survives even at very limited acceptance. Therefore, the MPD experiment will be suitable for the experimental investigation of this effect. It was also found that the shape of the net-proton rapidity

distribution near the midrapidity depends on the experimental acceptance. Hence, only if the measurements are taken at the same acceptance for all collision energies we can reliably conclude on the presence or absence of the wiggle. Notice that this is not the case for the data analyzed so far. These data were measured in four different experiments. Therefore, the question if the presently available data really indicate a wiggle in the  $C_y$  excitation function is still open. It is the more so because some of the data in the region of the expected wiggle still have a preliminary status. Data in preparation by the STAR experiment will meet the requirement that the excitation function of the curvature  $C_v$  shall be measured by one experiment and under same acceptance conditions since they stem from a collider experiment with an acceptance range widely independent of the beam energy. These data, however, come from the energy range above  $\sqrt{s_{NN}} = 7.7$  GeV, while the wiggle is expected in the range between 4 and 8 GeV. Therefore, systematic measurements with the NICA MPD collider experiment in the whole energy range of interest would be highly desirable to clarify this problem.

It is important to emphasize that the discussed irregularity is a signal from the hot and dense stage of the nuclear collision. It is formed at the (nonequilibrium) compression stage of the collision, as it was argued in Refs. [7,8]. Therefore, a theoretical approach should be able to treat differences in the models for the EoS (in particular, absence or presence of a first-order deconfinement transition) already at this early nonequilibrium compression stage of the collision. The 3FD model does it by means of the three-fluid interactions. This, at least, makes the baryon stopping EoS dependent. At the same time, a hybrid fluid model does not distinguish different EoS models at the early collision stage because the initial state for the hydrodynamical evolution is prepared by means of the same hadronic kinetic model for any EoS. Hence, the hybrid fluid model mainly reveals the baryon stopping inherent in the hadronic kinetic model and cannot produce an irregularity of this stopping relevant to the hydrodynamical EoS. Here the situation is similar to the case of the directed flow that is also predominantly formed at the early nonequilibrium stage of the collision (see discussion in Ref. [32]). The kinetic transport approach of the parton-hadron string dynamics (PHSD) [33] takes into account the possibility of the deconfinement transition at the early nonequilibrium stage of the collision. Therefore, in principle, the PHSD could reveal the discussed irregularity. However, the deconfinement transition in the PHSD model is of the crossover type. This implies that the wiggle irregularity is very weak, if present at all, in accordance with results obtained in the present paper.

## ACKNOWLEDGMENTS

We are grateful to A. S. Khvorostukhin, V. V. Skokov, and V. D. Toneev for providing us with the tabulated 2-phase and crossover EoS's. Our thanks go to O. Rogachevsky and V. Voronyuk for information about the acceptance range of the TOF detector in the NICA MPD experiment and to Nu Xu for the case of the STAR experiment. D.B. acknowledges discussions with M. Bleicher, Y. Karpenko, H. Petersen, and J. Steinheimer during his stay at FIAS Frankfurt. The

calculations were performed at the computer cluster of GSI (Darmstadt). Y.B.I. received partial support from the Russian Ministry of Science and Education Grant No. NS-932.2014.2;

the work of D.B. was supported in part by the Polish NCN under Grant No. UMO-2011/02/A/ST2/00306 and by the Hessian LOEWE initiative through HIC for FAIR.

- [1] G. S. F. Stephans, J. Phys. G 32, S447 (2006).
- [2] NA49 Collaboration, P. Seyboth *et al.*, Addendum 1 to the NA49 Proposal, CERNSPSC- 97-26; M. Gazdzicki, arXiv:nuclth/9701050; NA61/SHINE Collaboration, M. Gazdzicki *et al.*, PoS CPOD2006, 016 (2006).
- [3] B. Friman, C. Hohne, J. Knoll, S. Leupold, J. Randrup, R. Rapp, and P. Senger, Lect. Notes Phys. 814, 1 (2011).
- [4] A. N. Sissakian, A. S. Sorin, and V. D. Toneev, Conf. Proc. C060726, 421 (2006).
- [5] Yu. B. Ivanov, Phys. Lett. B 690, 358 (2010).
- [6] Yu. B. Ivanov, Phys. At. Nucl. 75, 621 (2012).
- [7] Yu. B. Ivanov, Phys. Lett. B 721, 123 (2013).
- [8] Yu. B. Ivanov, Phys. Rev. C 87, 064904 (2013).
- [9] Yu. B. Ivanov, V. N. Russkikh, and V. D. Toneev, Phys. Rev. C 73, 044904 (2006).
- [10] V. M. Galitsky and I. N. Mishustin, Sov. J. Nucl. Phys. 29, 181 (1979).
- [11] A. S. Khvorostukhin, V. V. Skokov, K. Redlich, and V. D. Toneev, Eur. Phys. J. C 48, 531 (2006).
- [12] J. Steinheimer and J. Randrup, Phys. Rev. Lett. 109, 212301 (2012).
- [13] J. Steinheimer and J. Randrup, Phys. Rev. C 87, 054903 (2013).
- [14] J. Steinheimer, J. Randrup, and V. Koch, Phys. Rev. C 89, 034901 (2014).
- [15] T. Klähn, R. Łastowiecki, and D. B. Blaschke, Phys. Rev. D 88, 085001 (2013).
- [16] T. Klähn, D. Blaschke, and F. Weber, Phys. Part. Nucl. Lett. 9, 484 (2012).
- [17] S. P. Merz, S. V. Rasin, and O. V. Rogachevsky, http://mpd. jinr.ru/data/presentations/ras/merts.pdf.

- [18] N. Xu (private communication).
- [19] E-0895 Collaboration, J. L. Klay et al., Phys. Rev. C 68, 054905 (2003).
- [20] E877 Collaboration, J. Barrette et al., Phys. Rev. C 62, 024901 (2000).
- [21] E917 Collaboration, B. B. Back et al., Phys. Rev. Lett. 86, 1970 (2001).
- [22] J. Stachel, Nucl. Phys. A 654, 119c (1999).
- [23] NA49 Collaboration, H. Appelshäuser *et al.*, Phys. Rev. Lett. 82, 2471 (1999).
- [24] NA49 Collaboration, T. Anticic *et al.*, Phys. Rev. C **69**, 024902 (2004).
- [25] NA49 Collaboration, C. Alt et al., Phys. Rev. C 73, 044910 (2006).
- [26] NA49 Collaboration, C. Blume et al., J. Phys. G 34, S951 (2007).
- [27] NA49 Collaboration, T. Anticic *et al.*, Phys. Rev. C 83, 014901 (2011).
- [28] STAR Collaboration, B. I. Abelev *et al.*, Phys. Rev. C **79**, 034909 (2009).
- [29] M. Mitrovski, T. Schuster, G. Graf, H. Petersen, and M. Bleicher, Phys. Rev. C 79, 044901 (2009).
- [30] C. M. Hung and E. V. Shuryak, Phys. Rev. Lett. 75, 4003 (1995).
- [31] J. Steinheimer and M. Bleicher, Eur. Phys. J. A **48**, 100 (2012).
- [32] Y. B. Ivanov and A. A. Soldatov, Phys. Rev. C. 91, 024915 (2015).
- [33] W. Cassing and E. L. Bratkovskaya, Nucl. Phys. A 831, 215 (2009); Phys. Rev. C 78, 034919 (2008); W. Cassing, Nucl. Phys. A 791, 365 (2007).