# Curvature correction term as a constraint for the Skyrme interaction

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The curvature correction term to the surface tension is used as a criterion for the efficiency of the Skyrme interaction in describing surface properties. Based on the nuclear equation of state, the curvature correction term to the surface tension coefficient is calculated for 97 standard Skyrme interaction parameter sets in the vicinity of nuclear saturation density at zero temperature. The main idea is to find those parametrizations that give the Tolman  $\delta$  correction close to the available theoretical predictions from the statistical theory. Only 59 out of 97 models give satisfactory results. Comparison of the obtained results with the results of the implementation of different macroscopic and microscopic constraints to Skyrme parametrizations available in the literature allows us to select four models that satisfy all the constraints.

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# I. INTRODUCTION

The derivation of the equation of state (EOS) of nuclear matter is among the most important goals and long-standing unsolved problems in nuclear physics and astrophysics [1]. Various approaches to the description of infinite nuclear matter exist. Among them are purely microscopic ones based on the realistic description of the nucleon-nucleon (NN) interaction [2], in which case the result depends not only on the chosen interaction but also on the way many-body effects are treated. These may be handled either by a direct description of the Bruckner-Hartree-Fock method, the Dirac-Bruckner-Hartree-Fock formalism [3,4], self-consistent Green's functions [5], etc.

At the same time, in describing the experimental data and in computer simulation, the most widely used models are those based on effective density-dependent NN and NNN interactions rather than on realistic ones (e.g., the models introduced by Skyrme [6] and Gogny [7]). The main problem of such approaches is in the infinite number of possible sets of model parameters providing satisfactory description of the ground-state properties of stable nuclei.

A large number of various Skyrme-force parametrizations and theoretical models attempting to describe nuclear matter and finite nuclei in a wide range of external parameters exists. They all were constructed under specific assumptions that reduce their predictive power [8]. Combined with the indirect model-dependent experimental methods used to evaluate nuclear matter properties, it makes the selection of the realistic sets of parameters quite a difficult task.

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from the literature. They result from the analysis of various observables, leading to different predictions concerning the behavior of nuclear matter away from equilibrium. Recently, a number of interesting and important papers systematically checking the sets of parameters for nuclear matter constraints appeared [8,9]. Important work on the construction of new parametrizations with systematic variation of the parameters to improve the precision of the results for some crucial nuclear matter properties was done in Refs. [10–12]. Such investigations may result in the improvement of the equation of state of nuclear matter, applicable in a wide range of parameters. Thus, the search for model-independent constraints connected with the specific properties of the nuclear matter is timely and important.

Presently over 200 sets of Skyrme parameters are known

Such constraints may result from the properties of the interfaces. To start with, let us mention that the droplet model of nuclei [13] plays a special role among macroscopic models. It makes possible the description of average properties of a saturated system, such as a nucleus, consisting of two components (neutrons and protons), with account for the boundary effects and the presence of a diffuse layer. The surface energy and the properties of the surfaces in nuclear matter have been studied in a number of papers [14–16]. And even though the dependence of the surface tension (and surface energy) on the surface curvature as well as its impact on different physical properties were also studied by several groups of authors [17,18], still, for decades it has remained one of the most controversial issues in mesoscopic thermodynamics [19–21].

In studies of surface properties of nuclei with mass number *A*, the account for the curvature effects is important. Within the droplet model this requires the inclusion of additional terms proportional to  $A^{\frac{1}{3}}$  in any expansion concerning the nuclear properties in terms of the fundamental dimensionless ratio, given by  $\frac{r_0}{R} = A^{-\frac{1}{3}}$ , which is the ratio of the interparticle

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spacing  $r_0$  to the nuclear radius R [13]. In these studies, effects connected with the surface curvature are specified by the curvature correction coefficient  $a_3$  accompanying terms of the order  $A^{\frac{1}{3}}$ . In statistical mechanics and in Gibbs-Tolman (GT) thermodynamics of interfaces, it corresponds to the Tolman length, called also the  $\delta$  correction [22]. The basic parameter  $\delta$  was first introduced by Tolman [23]. It is equal to the distance between the equimolar surface  $R_{em}$  and the surface of tension R at the interphase boundary:

$$\delta = R_{em} - R. \tag{1}$$

According to the GT theory, the surface tension  $\sigma$  of the curved interface, in the leading order approximation, can be defined as

$$\sigma(R) = \sigma_{\infty} \left( 1 - \frac{2\delta}{R} + \cdots \right), \tag{2}$$

where *R* is the droplet radius (equal to the radius of the surface of tension [24,25]) and  $\sigma_{\infty}$  is the surface tension of the planar interface. Originally introduced for ordinary liquids, it can be defined for any system with curved interface of a non-negligible boundary layer [19], such as nuclei and nuclear systems with a finite diffuse layer [26].

The first theoretical estimates of the correction term were done by Tolman [27]. It appeared to be close to the average interparticle distance  $r_0$ . Namely,  $\delta \sim (0.3-0.6)r_0$ , that, for a nuclear systems, is  $r_0 \sim 0.7$  fm at normal density of  $\rho \sim$ 0.17 fm<sup>-3</sup>. Present calculations from statistical mechanics yield the Tolman length of the order of the interparticle distance  $\delta \sim r_0 = 1.14$  fm [16]. Thus, mathematically the term  $\frac{2\delta}{R}$  in Eq. (2) becomes important for the systems with  $R < 14 \text{ fm} \left[\frac{2\delta}{R} > 0.1 \text{ in Eq. (2)} \text{ and even more so for nuclear systems} with <math>R \approx 0.7A^{\frac{1}{3}} = 0.7(277)^{\frac{1}{3}} = <4.6 \text{ fm for heavy nuclei}$ ].

In view of the importance of the curvature correction for nuclear systems, checking different sets of standard Skyrme parameters as for their ability to reproduce the theoretically predicted values for  $\delta$  correction becomes an important task. Therefore, following the idea of Ref. [9] in this work we present an attempt to use the Tolman correction as a constraint for different sets of Skyrme parameters.

#### **II. THEORETICAL MODEL**

In the present paper the Tolman  $\delta$  corrections are calculated for 97 different sets of Skyrme parameters known from the literature (see Appendix A). In our analysis we included the most popular parametrizations. From large families (e.g., BSk and SkSC), we selected several representative members. The Tolman length was calculated for the whole families in which  $\delta$  coincides with or is close to that calculated statistically. The results are compared with the theoretical predictions. This analysis is interesting as a test of various models regarding their ability to describe interphase interfaces.

Various approaches for the evaluation of the curvature correction exist. A method to calculate the Tolman  $\delta$  correction from the EOS of nuclear mater was introduced earlier in Ref. [28].

In that model, one gets for  $\delta$  from the EOS of symmetric nuclear matter with isospin-independent effective mass (see Appendix B) in the case of T = 0, at normal density  $\rho_0$ ,

$$\delta = \frac{2}{3} \frac{1}{\rho_0^2} \frac{-33t_0 - 160W\rho_0^{-1/3} + t_3(1+\alpha)\rho_0^{\alpha}\frac{1}{12}(7(3\alpha+6) - 3(3\alpha+6)^2)}{\left(15t_0 + \frac{1}{12}t_3(1+\alpha)((3\alpha+6) - (3\alpha+6)^2)\right)^2} \sigma_{\infty},\tag{3}$$

where

$$W = \frac{h^2}{10m} \left(\frac{3}{8\pi g}\right)^{\frac{2}{3}} \left(\frac{5 - 3\frac{m*}{m}}{\frac{m*}{m}}\right).$$
 (4)

We use the above results to calculate corrections to 97 sets of Skyrme parameters, using them in testing various parametrizations describing surface effects properly. Let us recall that, following Tolman's first estimates [27], subsequent calculations based on statistical mechanics that use minimum information regarding the details of the interaction about the studied liquids but operating with dimensionless parameters such as the size of the diffuse layer divided by the interparticle distance (e.g., penetrable sphere model in the mean field approximation) yield  $|\delta|$  in a range from  $\frac{1}{3}r_0$  to  $\frac{5}{9}r_0$  (see Refs. [24,29–31] and references therein). Even though the sign of the curvature corrections reported in the literature varies in different calculations, the absolute value is approximately the same, varying within the distance where the density profile faces rapid changes. The density profiles in the diffuse layer in the dimensionless coordinates are very similar for ordinary liquids and nuclear matter. This allows us to extend the results

of the physics of ordinary liquids to nuclear matter. Thus, with account for the uncertainty of the theoretical values of  $\delta$ , the admissible range for the curvature corrections adopted in the work is

$$|\delta| = (0.3 \sim 0.6)r_0 = (0.34 \sim 0.68) \text{ fm.}$$
 (5)

In all calculations the surface tension of the semi-infinite matter at T = 0 is calculated for the symmetric case with  $\rho_n = \rho_p = \frac{1}{2}\rho$  (without Coulomb interaction) within the restricted extended Thomas-Fermi approach [26]. Terms up to fourth order are considered. All the calculations include effective mass and spin-orbit contribution. The function  $\rho(z)$  used to minimize the surface energy is the one-parameter Fermi function:

$$\rho(z) = \frac{\rho_{\infty}}{1 + exp(\alpha z)}.$$
(6)

Our calculations yield values for  $\sigma_{\infty}$  in the range 0.93–1.21 MeV fm<sup>-2</sup>. The results of our calculation are presented in Fig. 1.

In Table I the parametrization yielding values of  $\delta$  within the admissible range (5) and, at the same time, satisfying

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FIG. 1. (Color online) Tolman's  $\delta$  length. Solid lines show the boundaries defined by Eq. (5); dashed line corresponds to the extended range up to the interparticle distance.

the constraints of Ref. [9] are shown (marked with the label "+" in the rightmost column), appended by the macroscopic properties of the corresponding parametrizations. The label "part" indicates that the parametrization passed the test of Ref. [9] for all but one of the applied constraints, and the failure for that one constraint is less than 5%.

#### **III. RESULTS AND DISCUSSION**

Only 59 parametrizations of those analyzed satisfy the constraint on the  $\delta$  corrections imposed in Eq. (5). Among them Ska25s20 and SV-min0 satisfy also all the criteria of Ref. [9], while Ska35s20, SkT1, SkT2, SkT3, and Skxs20 satisfy all but one of the constraints of Ref. [9]. If one increases the admissible range of  $\delta$  up to the interparticle distance, then two more parametrizations satisfying all the constraints of Ref. [9] and three more satisfying all but one will pass the  $\delta$ -correction test (see Table I). At the same time, as seen in Fig. 1, many parametrizations yield values of  $\delta$  correction close to the range allowed by Eq. (5). An interesting but not surprising observation is that, while some of the parametrizations of the family SkT pass the  $\delta$ -constraint test, others do not, although the parametrizations are based on the same inputs and use the same method. This observation may suggest that some of the parametrizations pass the test just by chance.

The main features of the parametrizations that passed the test are shown in Table II (see Appendix C). As seen from the table, the families with acceptable values of the Tolman correction use finite nuclei properties connected with the nuclear surface (e.g., surface properties of selected magic and semimagic nuclei, surface thickness, and neutron rms radius) as input data. At the same time, families that did not pass the test use different input data (e.g., v070 or skz2). It should be also noted that parametrizations elaborated for neutron matter (e.g., the SLy family) fail to produce an acceptable curvature correction to the surface tension in symmetric nuclear matter.

The obtained values appear negative for all chosen parametrizations except for SV (see Fig. 1), which means that the surface of tension is located closer to the liquid phase with respect to the equimolar surface.

An important observation is that all parametrizations constructed by systematic variations of the parameters (e.g., SV and SkS families) do pass the  $\delta$ -correction test.

Attempting to find an apparent pattern of the forces' performance in describing properties of the nuclear surfaces, we analyzed the dependence of  $\delta$  on various force parameters and the macroscopic properties characteristic for the tested forces. We also calculated the coupling constants found in recent years and aimed to improve the relevance of the Skyrme forces to different physical properties [9] and given as linear combinations of individual parameters:

$$C_{0}^{\rho} = \frac{3}{8}t_{0} + \frac{3}{48}t_{3}\rho_{0}^{\alpha},$$

$$C_{1}^{\rho} = -\frac{1}{4}t_{0}(\frac{1}{2} + x_{0}) - \frac{1}{24}t_{3}(\frac{1}{2} + x_{3})\rho_{0}^{\alpha}.$$
(7)

Other combinations of the individual parameters are those providing the most compact formulation of the energy functional and the residual interaction [32]:

$$b_0 = t_0 \left( 1 + \frac{1}{2} x_0 \right),$$
  

$$b_3 = \frac{1}{4} t_3 \left( 1 + \frac{1}{2} x_3 \right).$$
(8)

Unfortunately, no particular dependence of the value of the  $\delta$  correction in almost all of the force parameters, coupling

Skyrme force	$\delta$ (fm)	$ ho_0 (\mathrm{fm}^{-3})$	K (MeV)	$E_0$	$C_0^{\rho}$ (MeV fm <sup>3</sup> )	Ref. [9] results
Ska25s20	-0.74	0.1746	210.78	- 15.32	- 250.23	+
SV-min	-0.74	0.1746	210.78	-15.32	-250.23	+
Ska35s20	-0.78	0.16	230	- 16	-272.70	part
SkT1	-0.58	0.162	201.95	-15.81	-240.38	part
SkT2	-0.58	0.162	201.95	-15.81	-240.38	part
SkT3	-0.58	0.162	201.95	-15.81	-240.38	part
Skxs20	-0.58	0.1595	234	- 15.94	-253.67	part
Extended						
LNS	-0.74	0.1746	210.78	-15.32	-250.23	+
SQMC700	-0.74	0.1746	210.78	-15.32	-250.23	+
MSL0	-0.58	0.1595	234	- 15.94	-253.67	part
SKRA	-0.58	0.1595	234	- 15.94	-253.67	part
KDE0v1	-0.58	0.1595	234	- 15.94	- 253.67	part

TABLE I. Tolman  $\delta$  correction and macroscopic properties for various Skyrme parametrizations. Forces below "Extended" pass the  $\delta$ -correction test with the extended range (5) up to interparticle distance.

constants, or macroscopic properties was found. The only observable correlations are in the slight increase of the absolute value of  $\delta$  with increasing  $t_1$  (Fig. 2), decreasing effective mass (Fig. 3) and increasing absolute value of the coefficient  $C_0^{\rho}$  (Fig. 4).

From standard statistical analysis of the forces that pass the curvature correction test one may suggest that with probability 0.95 the most probable values of the above parameters should be

$$C_0^{\rho} = -245.1 \pm 3.6,$$
  

$$m* = 0.96 \pm 0.02,$$
 (9)  

$$t_1 = 304.1 \pm 10.2.$$

The standard deviations are  $SD_{C_0^{\rho}} = 13.8$ ,  $SD_{m*} = 0.07$  and  $SD_{t_1} = 36.7$ . One can see from the Figs. 2–4 that there exists a dependence of  $\delta$  on  $t_1$ , m\*, and  $C_0^{\rho}$ . Even though this may be the necessary condition for the force ability to describe



FIG. 3. Dependence of the absolute value of  $\delta$  on effective mass.



FIG. 2. Dependence of the absolute value of  $\delta$  on the  $t_1$  coefficient of the Skyrme force.



FIG. 4. Dependence of the absolute value of  $\delta$  on  $C_0^{\rho}$ .

nuclear surfaces, certainly it is not sufficient, since some of the parametrizations, including those with  $\delta$  outside our range [Eq. (5)], produce  $t_1$ , m\*, and  $C_0^{\rho}$  within the ranges imposed by Eq. (9). This may be an indication of some basic problem in the Skyrme-type parametrizations, maybe connected with the large freedom in choosing the force parameters.

#### **IV. CONCLUSIONS**

In this paper we calculated the curvature correction term of the surface tension from the nuclear equation of state for 97 different standard Skyrme parametrizations available in the literature. The obtained results show strong dependence of the curvature correction term on the EOS.

To summarize, our study shows that not all the existing parametrizations are capable of describing adequately the interphase interfaces in nuclear matter. In spite of the considerable uncertainty regarding the absolute value of  $\delta$ , the suggested constraint, even with a wide admissible range for  $\delta$  allows one to test various Skyrme forces regarding their capacity to describe curvature effects. It should be mentioned that better agreement is observed with the use of EOS that account for the surface effects with respect to those that do not (or were designed for neutron matter).

Comparison of the obtained results with the available data on different Skyrme forces with nuclear matter constraints suggests two parametrizations, namely Ska25s20 and SV-min, that satisfy all constraints. For an extended admissible range for  $\delta$ , that number increases and two more parametrizations, namely LNS and SCMC700, that come into play. It can be seen from our study that the systematic variation of the parameters is quite efficient, giving promising results in studies of the basic properties of nuclear matter. At the same time, further progress with Skyrme-type forces requires a better understanding of the physical meaning of different parameters rather than the introduction of new parameters.

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## APPENDIX A: PARAMETRIZATIONS USED

The Skyrme parametrizations analyzed in the work are BSk1 [33], BSk10 [34], E [10], E  $\sigma$  [10], FPLyon [35], G $\sigma$  [10], KDE [36], KDE0v [36], KDE0v1 [36], LNS [37], MSk1 [38], MSk2 [38], MSk3 [38], MSk4 [38], MSk5 [38], MSk5\* [39], MSk6 [38], MSkA [40], MSL0 [41], NRAPR [42], PRC45 [43], RATP [44], R  $\sigma$  [10], SGII [45], SGO2 [46], SII [47], SIII [47], SIV [47], SK255 [48], SK272 [48], Ska [49], Ska25s20 [9], Ska35s20 [9], SkI4 [32], SkM\* [50],

SkMP [46], SkO [32], SkP [50], SKRA [51], SkS1 [11], SkS2 [11], SkS3 [11], SkS4 [11], SkSC1 [52], SkSC10 [53], SkSC2 [52], SkSC3 [52], SkSC4 [53], SkSC5 [53], SkSC6 [53], SkSP1 [39], SkT1 [54], SkT1\* [54], SkT2 [54], SkT3 [54], SkT3\* [54], SkT4 [54], SkT5 [54], SkT6 [54], SkT7 [54], SkT8 [54], SkT9 [54], SKX [55], SKXce [55], SKXm [55], Skxs15 [56], Skxs20 [56], Skxs25 [56], Skz2 [57], SLy4 [58], SLy5 [58], SLy6 [58], SLy7 [58], SQMC650 [9], SQMC700 [9], SV [47], SV-bas [12], SVI [47], SV-K226 [12], SV-kap02 [12], SV-mas08 [12], SV-min [12], SV-sym32 [12], SV-tls [12], T [10], v070 [59], Z [10], ZR1a [43], ZR1b [43], ZR1c [43], ZR2a [43], ZR2b [43], ZR2c [43], ZR3a [43], ZR3b [43], ZR3c [43], and Z  $\sigma$  [10].

## APPENDIX B: EQUATION OF STATE OF NUCLEAR MATTER

To calculate the  $\delta$  correction, an EOS of nuclear matter at low temperatures and in the high-density limit was used, where  $\lambda^3 \rho \gg 1$  (i.e., when the average de Broglie thermal wavelength  $\lambda$  is larger than the average interparticle separation  $\rho^{-\frac{1}{3}}$ ). In this case, the EOS takes the form [60]

$$P(\rho_q, T) = \sum_q \left[ \frac{5}{3} \varepsilon^*_{kq}(\rho_q, T) - \varepsilon_{kq}(\rho_q, T) \right] + \frac{t_0}{2} \left( 1 + \frac{x_0}{2} \right) \rho^2 + \frac{t_3}{12} \left( 1 + \frac{x_3}{2} \right) (\alpha + 1) \rho^{\alpha + 2} - \frac{t_0}{2} \left( x_0 + \frac{1}{2} \right) \sum_q \rho_q^2 - \frac{t_3}{12} \left( \frac{1}{2} + x_3 \right) (\alpha + 1) \rho^{\alpha} \sum_q \rho_q^2,$$
(B1)

with

$$\varepsilon_{kq} = \frac{m_q^*}{m} \frac{1}{\beta} \frac{2g}{\sqrt{\pi}} \lambda_q^{-3} F_{\frac{3}{2}}(\eta_q),$$
  

$$\varepsilon_{kq}^* = \frac{1}{\beta} \frac{2g}{\sqrt{\pi}} \lambda_q^{-3} F_{\frac{3}{2}}(\eta_q),$$
(B2)

where *m* and *m*<sup>\*</sup> are the mass and effective mass, respectively; *T* and  $\rho$  are temperature and density; *q* is the particle type (*q* = proton, neutron); *F* is the Fermi integral;  $\lambda = \sqrt{\frac{2\pi\hbar^2}{m^*T}}$  is the average de Broglie thermal wavelength; *g* = 2 is the spin degeneracy factor; *t*<sub>0</sub>, *t*<sub>3</sub>, *x*<sub>0</sub>, *x*<sub>3</sub>, and  $\alpha$  are the Skyrme force parameters; and  $\beta = \frac{1}{T}$ .

## APPENDIX C: METHODS AND INPUT DATA USED IN CONSTRUCTION OF DIFFERENT SKYRME PARAMETRIZATIONS

In Table II several representatives of Skyrme parametrizations that satisfy the constraint on the  $\delta$  correction are shown. The input data and the methods used in their construction are compared.

Skyrme forces	Method	Input data
KDE0v1	Simulated annealing method	Ground-state properties of normal and exotic nuclei: binding energy, charge radii and spin-orbit splitting, radii for $1d_{5/2}$ and $1f_{7/2}$ neutron orbits in <sup>17</sup> O and <sup>41</sup> Ca breathing mode energies, critical density $\rho_{cr}$ , positive slope of the symmetry energy up to 3 $\rho_0$ , enhancement factor associated with GDR L andau parameter $G'$
LNS	Bruckner-Hartree-Fock with two- and three-body forces (homogeneous matter) Hartree-Fock (finite matter)	Nucleon effective mass in symmetric nuclear matter (SNM) and asymmetric nuclear matter (ANM), energy per particle in SNM and ANM as function of density and proton neutron asymmetry; constraint on the Landau parameter $G_0$ , surface properties of selected magic and semimagic nuclei, are arbitrary in $P_0$ and $P_0$
SV family	Hartree-Fock	Properties of finite nuclei (energies, radii, and surface thickness), energies of giant resonances,
MSkA	Density-dependent Hartree-Fock	Empirical binding energies and charge radii of the closed-shell nuclei <sup>16</sup> O, <sup>40</sup> Ca, <sup>90</sup> Zr, and <sup>208</sup> Pb; isotopes <sup>116,124</sup> Sn and <sup>214</sup> Pb; doubly closed <sup>132</sup> Sn
SK255, SK272	Hartree-Fock-based random phase approximation (RPA) approach	Nuclear binding energies, charge radii, and neutron radii; rms charge radius of <sup>208</sup> Pb; incompressibility coefficient from relativistic and nonrelativistic mean-field-based RPA calculations
SkS family	Hartree-Fock seniority	Symmetric nuclear matter ground-state properties; binding energies of <sup>16</sup> O, <sup>40,48</sup> Ca, <sup>90</sup> Zr, and <sup>208</sup> Pb; spin-orbit splitting in <sup>16</sup> O; surface parameter and symmetry properties; fission barrier of <sup>240</sup> Pu; restrictions on L andau parameter for SkS4
SkT family	Hartree-Fock; extended Thomas-Fermi	Nuclear radii and binding energies; droplet model parameters; experimental masses, charge radii, charge distribution, neutron skin thickness, semiclassical fission barriers, and Landau parameters
Skxs family	Skyrme Hartree-Fock; relativistic mean-field models	Binding energies, rms charge radii, and single-particle energies; binding energy difference <sup>48</sup> Ni- <sup>48</sup> Ca; charge density of <sup>208</sup> Pb; constraint of $\alpha = 1/6$ for the density-dependent potential
E, Z, T, E $_{\sigma}$ , G $_{\sigma}$ , R $_{\sigma}$ , Z $_{\sigma}$	Hartree-Fock-Bogoliubov	Binding energy, diffraction radius, surface width of <sup>16</sup> O, <sup>40,48</sup> Ca, <sup>58</sup> Ni, <sup>90</sup> Zr, <sup>116,124</sup> Sn, and <sup>208</sup> Pb; <i>l-s</i> splitting of the 1 <i>p</i> level in <sup>16</sup> O

TABLE II. Comparison of the methods and input data used in construction of different Skyrme parametrizations.

- M. Baldo and C. Maieron, J. Phys. G: Nucl. Part. Phys. 34, R243 (2007).
- [2] R. Machleidt, Phys. Rev. C 63, 024001 (2001).
- [3] B. ter Haar and R. Malfiet, Phys. Rep. 149, 207 (1987).
- [4] I. Bombaci and U. Lombardo, Phys. Rev. C 44, 1892 (1991).
- [5] Y. Dewulf, W. H. Dickhoff, D. Van Neck, E. R. Stoddard, and M. Waroquier, Phys. Rev. Lett. 90, 152501 (2003).
- [6] T. H. R. Skyrme, Philos. Mag. 1, 1043 (1956).
- [7] J. Decharge and D. Gogny, Phys. Rev. C 21, 1568 (1980).
- [8] P. Stevenson and P. Goddard, in *Proceedings of the 31st International Workshop on Nuclear Theory*, edited by A. Georgieva and N. Minkov (Heron Press, Sofia, 2012), Vol. 31.
- [9] M. Dutra, O. Lourenco, J. S. Sá Martins, A. Delfino, J. R. Stone, and P. D. Stevenson, Phys. Rev. C 85, 035201 (2012).
- [10] J. Friedrich and P.-G. Reinhard, Phys. Rev. C 33, 335 (1986).

- [11] J. Gomez, C. Prieto, and J. Navarro, Nucl. Phys. A 549, 125 (1992).
- [12] P. Klupfel, P.-G. Reinhard, T. J. Burvenich, and J. A. Maruhn, Phys. Rev. C 79, 034310 (2009).
- [13] W. D. Myers and W. Swiatecki, Ann. Phys. 55, 395 (1969).
- [14] D. Ravenhall, C. Pethick, and J. Lattimer, Nucl. Phys. A 407, 571 (1983).
- [15] V. G. Boyko, L. L. Jenkovszky, and V. M. Sysoev, Zeitsch. Phys. C 45, 607 (1990).
- [16] L. L. Jenkovszky, B. Kampfer, and V. M. Sysoev, Sov. J. Nucl. Phys. 57, 1507 (1994).
- [17] L. G. Moretto, J. B. Elliott, P. T. Lake, and L. Phair, AIP Conf. Proc. 1491, 75 (2012).
- [18] K. Pomorski and J. Dudek, Phys. Rev. C 67, 044316 (2003).
- [19] M. A. Anisimov, Phys. Rev. Lett. 98, 035702 (2007).

- [20] E. M. Blokhius and J. Kuipers, J. Chem. Phys. 124, 074701 (2006).
- [21] V. M. Kolomietz, S. V. Lukyanov, and A. I. Sanzhur, Phys. Rev. C 86, 024304 (2012).
- [22] S. Ono and S. Kondo, *Molecular Theory of Surface Tension in Liquids* (Springer-Verlag, Berlin, 1960).
- [23] R. C. Tolman, J. Chem. Phys. 17, 333 (1949).
- [24] J. S. Rowlinson and B. Widom, *Molecular Theory of Capillarity* (Clarendon, Oxford, UK, 1982).
- [25] J. S. Rowlinson, J. Phys.: Condens. Matter 6, A1 (1994).
- [26] M. Brack, C. Guet, and H.-B. Hakansson, Phys. Rep. 123, 275 (1985).
- [27] R. Tolman, J. Chem. Phys. 17, 118 (1949).
- [28] K. V. Cherevko, L. A. Bulavin, L. L. Jenkovszky, V. M. Sysoev, and F.-S. Zhang, Phys. Rev. C 90, 017303 (2014).
- [29] A. Harasima, Advances in Chemical Physics, edited by I. Prigogine, Vol. 1 (Interscience Publishers Inc., New York, 1958), p. 203).
- [30] N. I. Gorski, Z. Phys. Chem. (Leipzig) 270, 817 (1989).
- [31] J. H. Irving and J. G. Kirkwood, J. Chem. Phys. 18, 817 (1950).
- [32] P.-G. Reinhard, D. J. Dean, W. Nazarewicz, J. Dobaczewski, J. A. Maruhn, and M. R. Strayer, Phys. Rev. C 60, 014316 (1999).
- [33] M. Samyn, S. Goriely, P.-H. Heenen, J. Pearson, and F. Tondeur, Nucl. Phys. A 700, 142 (2002).
- [34] S. Goriely, M. Samyn, and J. Pearson, Nucl. Phys. A 773, 279 (2006).
- [35] J. Meyer, in Les Noyaux en 1993 Une Nouvelle Facon D'exister, Ecole Internationale Joliot-Curie de Physique Nucleaire (L'Institut National de Physique Nucleaire et de Physique des Particules du CNRS en collaboration avec La Direction des Sciences de la Matiere du CEA et le Fonds National de la Recherche Scientifique de Belgique, 1993), p. 1.
- [36] B. K. Agrawal, S. Shlomo, and V. K. Au, Phys. Rev. C 72, 014310 (2005).
- [37] L. G. Cao, U. Lombardo, C. W. Shen, and N. V. Giai, Phys. Rev. C 73, 014313 (2006).
- [38] F. Tondeur, S. Goriely, J. M. Pearson, and M. Onsi, Phys. Rev. C 62, 024308 (2000).

- [39] M. Farine, J. Pearson, and F. Tondeur, Nucl. Phys. A 696, 396 (2001).
- [40] M. M. Sharma, G. Lalazissis, J. Konig, and P. Ring, Phys. Rev. Lett. 74, 3744 (1995).
- [41] L.-W. Chen, C. M. Ko, B.-A. Li, and J. Xu, Phys. Rev. C 82, 024321 (2010).
- [42] A. Steiner, M. Prakash, J. Lattimer, and P. Ellis, Phys. Rep. 411, 325 (2005).
- [43] S. J. Lee and A. Z. Mekjian, Phys. Rev. C 63, 044605 (2001).
- [44] M. Rayet, M. Arnould, F. Tondeur, and G. Paulus, Astron. Astrophys. 116, 183 (1982).
- [45] N. V. Giai and H. Sagawa, Phys. Lett. B 106, 379 (1981).
- [46] Q. B. Shen, Y. L. Han, and H. R. Guo, Phys. Rev. C 80, 024604 (2009).
- [47] M. Beiner, H. Flocard, N. V. Giai, and P. Quentin, Nucl. Phys. A 238, 29 (1975).
- [48] B. K. Agrawal, S. Shlomo, and V. Kim Au, Phys. Rev. C 68, 031304 (2003).
- [49] H. S. Kohler, Nucl. Phys. A 258, 301 (1976).
- [50] E. Chabanat, E. Bonche, E. Haensel, J. Meyer, and R. Schaeffer, Nucl. Phys. A 627, 710 (1997).
- [51] M. Rashdan, Mod. Phys. Lett. 15, 1287 (2000).
- [52] J. Pearson, Y. Aboussir, A. Dutta, R. Nayak, and M. Farine, Nucl. Phys. A 528, 1 (1991).
- [53] M. Onsi, H. Przysiezniak, and J. M. Pearson, Phys. Rev. C 50, 460 (1994).
- [54] F. Tondeur, M. Brack, M. Farine, and J. M. Pearson, Nucl. Phys. A 420, 297 (1984).
- [55] B. A. Brown, Phys. Rev. C 58, 220 (1998).
- [56] B. A. Brown, G. Shen, G. C. Hillhouse, J. Meng, and A. Trzcinska, Phys. Rev. C 76, 034305 (2007).
- [57] J. Margueron, J. Navarro, and N. V. Giai, Phys. Rev. C 66, 014303 (2002).
- [58] E. Chabanat, P. Bonche, R. Haensel, J. Meyer, and R. Schaeffer, Nucl. Phys. A 635, 231 (1998).
- [59] J. M. Pearson and S. Goriely, Phys. Rev. C 64, 027301 (2001).
- [60] S. J. Lee and A. Z. Mekjian, Phys. Rev. C 77, 054612 (2008).