Spectroscopy of ⁷⁶Se: Prolate-to-oblate shape transition

C. Xu, ¹ X. Q. Li, ¹ J. Meng, ¹ S. Q. Zhang, ^{1,*} H. Hua, ^{1,†} S. Y. Wang, ² B. Qi, ² C. Liu, ² Z. G. Xiao, ^{3,4} H. J. Li, ³ L. H. Zhu, ⁵ Z. Shi, ⁵ Z. H. Li, ¹ Y. L. Ye, ¹ D. X. Jiang, ¹ J. J. Sun, ¹ Z. H. Zhang, ¹ Y. Shi, ¹ P. W. Zhao, ¹ Q. B. Chen, ¹ W. Y. Liang, ¹ R. Han, ¹ C. Y. Niu, ¹ C. G. Li, ¹ C. G. Wang, ¹ Z. H. Li, ¹ S. M. Wyngaardt, ⁶ R. A. Bark, ⁷ P. Papka, ⁶ T. D. Bucher, ^{6,7} A. Kamblawe, ^{6,7} E. Khaleel, ^{6,7} N. Khumalo, ^{7,8,9} E. A. Lawrie, ⁷ J. J. Lawrie, ⁷ P. Jones, ⁷ S. M. Mullins, ⁷ S. Murray, ⁷ M. Wiedeking, ⁷ J. F. Sharpey-Schafer, ^{7,8} S. N. T. Majola, ^{7,10} J. Ndayishimye, ^{6,7} D. Negi, ⁷ S. P. Noncolela, ^{7,8} S. S. Ntshangase, ⁹ O. Shirinda, ⁷ P. Sithole, ^{7,8} M. A. Stankiewicz, ^{7,10} J. N. Orce, ⁸ T. Dinoko, ^{7,8} J. Easton, ^{7,8} B. M. Nyakó, ¹¹ and K. Juhász¹² ¹School of Physics and State Key Laboratory of Nuclear Physics and Technology, Peking University, Beijing 100871, China ²Shandong Provincial Key Laboratory of Optical Astronomy and Solar-Terrestrial Environment, School of Space Science and Physics, Shandong University, Weihai 264209, China ³Department of Physics, Tsinghua University, Beijing 100084, China ⁴Collaborative Innovation Center of Quantum Matter, Beijing 100084, China ⁶Department of Physics, University of Stellenbosch, Matieland 7602, South Africa ⁷iThemba LABS, 7129 Somerset West, South Africa ⁸Department of Physics, University of Tuluand, Private Bag X1001, KwaDlangezwa 3886, South Africa ¹⁰Department of Physics, University of Zululand, Private Bag X1001, KwaDlangezwa 3886, South Africa

¹¹Institute of Nuclear Research of the Hungarian Academy of Sciences (ATOMKI), H-4001 Debrecen, P.O. Box: 51, Hungary

¹²Department of Information Technology, University of Debrecen, Egyetem tér 1, Debrecen, Hungary

(Received 5 May 2015; revised manuscript received 26 May 2015; published 22 June 2015)

The spectroscopy of ⁷⁶Se has been studied using the ⁷⁰Zn(12 C, $\alpha 2n$)⁷⁶Se fusion evaporation reaction. The yrast band of ⁷⁶Se has been extended to substantially higher spin, allowing observation of the second band crossing. The much-delayed $g_{9/2}$ proton-pair alignment is discussed in terms of the cranked shell model and most likely is caused by a shape transition from prolate to oblate along the yrast line occurring in ⁷⁶Se. Based on the systematic investigation of the band crossings associated with the $g_{9/2}$ quasiparticle alignments and their relationships with the shape evolutions in the even-A Se and Kr isotopes, a comprehensive picture of shape evolution along with spin and isospin in these nuclei is obtained.

DOI: 10.1103/PhysRevC.91.061303

PACS number(s): 21.10.Re, 23.20.Lv, 27.50.+e

The nuclear shape that reflects the spatial distribution of the nucleons is a fundamental characteristic of the nucleus. For the neutron-deficient Kr and Se nuclei in the $A \sim 70$ mass region, due to the large subshell gaps around Fermi surfaces in the single-particle spectra at prolate and/or oblate deformation for proton and neutron numbers 34, 36, and 38, they exhibit rich and varied shape-related phenomena, such as rapid shape transition, shape coexistence, and triaxiality, and have been the focus of intense theoretical and experimental investigations in recent years.

For the neutron-deficient Kr isotopes, a clear and coherent picture of shape transitions and shape coexistence has been obtained. With the increase of isospin, both the experimental results [1–6] and theoretical calculations [7–14] have demonstrated that the shape transition of the ground state from oblate to prolate occurs at ⁷⁴Kr. Meanwhile, the coexistence of prolate and oblate shapes has also been discovered in these transitional Kr isotopes. For example, a low-lying prolate deformed excited 0^+ state, which coexists with the oblate 0^+ ground state, has been found in ⁷²Kr [3,15], while low-lying oblate deformed excited 0^+ states, which coexist with the prolate 0^+ ground

states, have been found in ^{74,76}Kr [2,5,16–19]. These two coexistent shapes, prolate and oblate, interchange their roles between ⁷²Kr and ⁷⁴Kr. In addition, with the increase of angular momentum, a rapid shape transition from oblate to prolate along the yrast line has been suggested to occur at very low spins in ⁷²Kr with the recent lifetime measurement [6]. In contrast, the *g*-factor [20] and lifetime [21] measurements, as well as the excited Vampir calculations [9], indicated that ⁷⁸Kr undergoes a prolate-to-oblate shape transition along the yrast line. For ^{74,76}Kr, along the yrast line, the prolate shapes of their ground states were found to be relatively stable and persist to higher spins [19,22–26].

Compared to the neutron-deficient Kr isotopes, the situation for the neutron-deficient Se isotopes is not that clear. The early Coulomb excitation experiments in the late 1970s have shown that all the ground states of ^{74,76,78,80,82}Se have prolate shapes [27,28], while the ⁶⁸Se was found to have an oblate ground state [29,30]. For ^{70,72}Se, total-Routhian-surface [31] and excited Vampir [32] calculations predicted an oblate ground state. In contrast to the predictions, a Coulomb excitation experiment of ⁷⁰Se suggested its ground state has a prolate shape like the heavier Se isotopes [33]. Later, new lifetime measurements for ^{70,72}Se [34] revised the conclusions drawn from the Coulomb excitation of ⁷⁰Se. The experimental results in Ref. [34] seem to favor a positive value of the

^{*}sqzhang@pku.edu.cn

[†]hhua@pku.edu.cn

spectroscopic quadrupole moment, i.e., oblate shape, for the ⁷⁰Se. However, due to the large uncertainty of the Coulomb excitation cross section and its relatively weak dependence on the quadrupole moment, in Ref. [34], a precise determination of sign of the quadrupole moment for ⁷⁰Se was not achieved. More accurate experimental confirmation of exact location where shape transitions of the ground state from oblate to prolate occur in Se isotopes is still needed. Meanwhile, similar to 72 Kr, both 70 Se and 72 Se were found to have a prolate shape at low spins as indicated by the lifetime measurements [34] and the high-spin studies [31]. For ⁷⁴Se, although direct quadrupole moments measurements have not been made for excited states above the 2^+ level, a high-spin study has revealed that a prolate shape is necessary to reproduce the character of the observed band crossings [35]. As a neighboring nucleus of ⁷⁴Se, ⁷⁶Se is also suggested to have a prolate ground state [27]. So far the high-spin structure of ⁷⁶Se has not been well established and the yrast band was observed only up to spin 12^+ [36]. Due to little experimental information on the high-spin structure, no definite conclusion of shape evolution along the yrast line in ⁷⁶Se has been drawn. An interesting question therefore arose whether the prolate shape of the ground state in ⁷⁶Se persists to high spins as in the neighboring isotope ⁷⁴Se or undergoes a prolate-to-oblate shape transition along the yrast line as in the neighboring isotone ⁷⁸Kr. To answer this question and to get a complete picture of shape evolution with spin and isospin in the neutron-deficient Se isotopes, extending the high-spin spectroscopic study to ⁷⁶Se is very meaningful.

Here, we report an experimental investigation on the high-spin properties of ⁷⁶Se via the ⁷⁰Zn(¹²C, α 2n)⁷⁶Se fusion-evaporation reaction. The collective structure of ⁷⁶Se is

PHYSICAL REVIEW C 91, 061303(R) (2015)

expanded significantly, allowing the observation of the second band crossing. The characters of band crossings in ⁷⁶Se are discussed in terms of the cranked shell model (CSM) and suggest that a possible shape transition from prolate to oblate along the yrast line occurs in ⁷⁶Se.

The present experiment was performed at iThemba LABS in South Africa. The high-spin states of ⁷⁶Se were populated via the ⁷⁰Zn(¹²C, α 2n)⁷⁶Se fusion-evaporation reaction at beam energies of 60 and 65 MeV. The target was selfsupporting ⁷⁰Zn with a thickness of 0.85 mg/cm². The in-beam γ -rays were detected by the AFRODITE array [37], which consists of eight Compton-suppressed clover detectors. The clover detectors have been arranged in two rings at 90° (four clovers) and 135° (four clovers) with respect to the beam direction. To select specific reaction channels, the DIAMANT array [38,39], which consisted of 64 CsI(Tl) scintillators in the present experiment, was also used with the AFRODITE array.

A total of $1.24 \times 10^8 \alpha - \gamma - \gamma$ coincident events were collected, from which a symmetric matrix was built. The level scheme analysis was performed using the RADWARE program [40]. The γ -ray spectra gated on the known γ -ray transitions in ⁷⁶Se are shown in Fig. 1. In order to obtain Directional Correlations of γ rays deexciting Oriented states (DCO) intensity ratios to determine the multipolarities of γ -ray transitions, the detectors around 90° with respect to the beam direction were sorted against the detectors around 135° to produce a two-dimensional angular correlation matrix. To get clean DCO values for transitions in ⁷⁶Se, gates were set on uncontaminated stretched E2 transitions. In general, stretched quadrupole transitions were adopted if DCO ratios were larger



FIG. 1. Coincidence γ -ray spectra generated from (a) the sum of gates on 1007.1-, 1029.6-, and 1131.2-keV transitions, (b) the sum of gates on 942.8- and 962.0-keV transitions. The peaks marked with stars are known contaminants.

SPECTROSCOPY OF ⁷⁶Se: PROLATE-TO-OBLATE ...



FIG. 2. (Color online) Partial level scheme of ⁷⁶Se. Energies are in keV. New observed transitions are indicated by red lines.

than 1.0, and stretched dipole transitions were assumed if DCO ratios were less than 0.8.

The partial level scheme of ⁷⁶Se deduced from the present work is shown in Fig. 2. It was constructed from $\gamma - \gamma$ coincidence relationships, intensity balances, and DCO analyses. As shown in Fig. 2, the previously reported positive-parity bands [41,42] have been considerably extended. The yrast band (band 1) of ⁷⁶Se is extended from spin 12⁺ at 5429.4 keV to spin 22⁺ at 13678.9 keV and the γ -vibrational band (band 2) is extended from spin 10⁺ at 4685.3 keV to spin 19⁺ at 11144.9 keV.

The high spin states of ⁷⁶Se, observed in the present work, allow band crossing phenomena and shape evolution to be studied. In neutron-deficient nuclei in the $A \sim 70$ region, both protons and neutrons occupy the same high- $j g_{9/2}$ intruder subshell and give rise to strong competition between the $g_{9/2}$ neutron and proton alignments, which together with the variety of shapes in these nuclei, means it is usually difficult to discern which alignment will be favored for the band crossing. To identify the origin of the observed band crossings in this mass region, systematics and theoretical studies of the characters of



FIG. 3. (Color online) The kinematic $J^{(1)}$ and dynamic $J^{(2)}$ moments of inertia as functions of rotational frequency for the yrast bands in ⁷⁴Se, ⁷⁶Se, and ⁷⁸Kr. New data of ⁷⁶Se are indicated by solid red triangles.

the band crossings are very helpful. In Fig. 3, the kinematic $J^{(1)}$ and dynamic $J^{(2)}$ moments of inertia as functions of rotational frequency for the yrast band in ⁷⁶Se, in comparison with those for the yrast bands in neighboring isotope ⁷⁴Se and isotone ⁷⁸Kr, are plotted.

As shown in Figs. 3(a) and 3(c), at low rotational frequency <0.55 MeV, the kinematic moments of inertia for ⁷⁴Se, ⁷⁶Se, and ⁷⁸Kr increase rapidly. Above a rotational frequency of 0.55 MeV, the kinematic moments of inertia of 76 Se and 78 Kr become flat until a rotational frequency of 0.80 MeV, while the kinematic moment of inertia of ⁷⁴Se continues to increase until a rotational frequency of 0.70 MeV and then becomes relatively flat. Above the rotational frequency of 0.80 MeV, both ⁷⁶Se and ⁷⁸Kr display a similar onset of other upbending. As shown in Fig. 3(b), corresponding to the long upbending in the kinematic moments of inertia, there are two close peaks around frequencies of 0.50 and 0.65 MeV, in the dynamic moment of inertia of the yrast band in ⁷⁴Se. In Ref. [35], the long upbending observed in ⁷⁴Se was ascribed to the successive alignments of a pair of $g_{9/2}$ protons around $\hbar \omega = 0.50$ MeV and a pair of $g_{9/2}$ neutrons around $\hbar \omega = 0.65$ MeV, meanwhile its prolate shape was assumed to persist up to the second band crossing region. For 78 Kr, the *g*-factor [20] and lifetime [21]



FIG. 4. (Color online) Experimental and calculated moments of inertia for the yrast band of 76 Se.

measurements have indicated that the two well-separated band crossings occurring around $\hbar\omega = 0.55$ and 0.90 MeV can be well ascribed to the alignments of a pair of $g_{9/2}$ neutrons and a pair of $g_{9/2}$ protons, respectively. Meanwhile, a shape transition from prolate to oblate around the first band crossing region for ⁷⁸Kr is suggested. The striking similarity in the moments of inertia between ⁷⁶Se and ⁷⁸Kr indicates that ⁷⁶Se has a similar alignment behavior to that of isotone ⁷⁸Kr, which is different from that of isotope ⁷⁴Se. In addition, it is well known that the nuclear moment of inertia is very sensitive to the nuclear shape. Considering that previous experimental investigations [27,28,43,44] have indicated that all the ground states of ⁷⁴Se, ⁷⁶Se, and ⁷⁸Kr have similar prolate shapes, the similar moment of inertia behavior between ⁷⁶Se and ⁷⁸Kr implies that ⁷⁶Se may undergo a similar prolate-to-oblate shape evolution along the yrast line as ⁷⁸Kr.

To further understand the details of band-crossing phenomena and shape evolution in ⁷⁶Se, the CSM calculations [7,45–47] using a Woods-Saxon potential with the deformation parameters $(\beta_2, \beta_4, \gamma) = (0.267, 0.014, 0^\circ)$ and $(\beta_2, \beta_4, \gamma) =$ $(0.267, 0.014, -60^{\circ})$ for both quasineutrons and quasiprotons in ⁷⁶Se are performed. The quadrupole and hexadecapole deformation parameters $\beta_2 = 0.267$ and $\beta_4 = 0.014$ are deduced from the coupled-channel analyses of the polarized-protons inelastic scattering measurements [48]. With the prolate shape, the $g_{9/2}$ proton-pair alignment and $g_{9/2}$ neutron-pair alignment occur around the similar rotational frequency, while with the oblate shape, the band crossing shifts to a much higher rotational frequency for the $g_{9/2}$ proton-pair alignment but remains nearly unchanged for the $g_{9/2}$ neutron-pair alignment. In Fig. 4, we compare the experimental moments of inertia with the theoretical results using the oblate deformation parameters for the yrast band of ⁷⁶Se. The overall agreements between the experimental data and theoretical results are reasonable. Therefore, the pronounced delay of the $g_{9/2}$ proton-pair alignment can be seen as a sign of oblate shape, and a prolate-to-oblate shape transition along the yrast line appears to well explain the underlying reason for the observation of



PHYSICAL REVIEW C 91, 061303(R) (2015)

FIG. 5. (Color online) Calculated crossing frequency, $\hbar\omega_c$, for the yrast band as a function of the triaxiality parameter γ for the fixed quadrupole and hexadecapole deformation parameters $\beta_2 =$ 0.267 and $\beta_4 = 0.014$.

the delayed band crossing for the $g_{9/2}$ proton-pair alignment in ⁷⁶Se.

As is well known, triaxiality plays an important role in this transitional nuclear region [11,49–51]. To investigate the dependence of the $g_{9/2}$ quasiparticle crossing frequencies on the triaxiality, in Fig. 5, the calculated $g_{9/2}$ quasineutron and $g_{9/2}$ quasiproton crossing frequencies as functions of the triaxial deformed parameter γ are plotted. Here the Lund convention for the γ value is used in the CSM [52]. It can be seen that with a prolate-like deformation ($-20^{\circ} < \gamma < 30^{\circ}$), the $g_{9/2}$ quasiproton crossing frequencies are close to each other, while with an oblate-like deformation ($-60^{\circ} < \gamma < -40^{\circ}$), the $g_{9/2}$ quasiproton crossing frequency is much higher than that of $g_{9/2}$ quasiparton. The calculations further illustrate the role of oblate shape for the delayed second band crossing in ⁷⁶Se.

With the new results in ⁷⁶Se, to achieve an overall picture of the $g_{9/2}$ quasiparticle alignments and their relationships with the shape evolutions in $A \sim 70$ mass region, the systematics of the $g_{9/2}$ neutron and proton crossing frequencies are plotted as functions of neutron number for even-A Se and Kr isotopes, in Fig. 6. Several interesting systematical features can be clearly seen. (1) For the lighter ^{70,72}Se and ^{72,74,76}Kr isotopes, they have similar alignment schemes: simultaneous $g_{9/2}$ proton-pair and neutron-pair alignments. The ⁷²Kr has been found to have an oblate ground state and undergo a rapid shape evolution from oblate to prolate at low spins. For ^{70,72}Se, previous studies have shown that these two isotopes also have prolate shapes at low spins, although the shapes of their ground states have not been definitely determined. Meanwhile, the prolate shapes of ground states of ^{74,76}Kr were found to be preserved to higher spins. Thus the simultaneous occurrence of $g_{9/2}$ proton-pair and neutron-pair alignments in these light Se and Kr isotopes can be interpreted in terms of their prolate shapes at the band crossing region. (2) For ⁷⁴ Se, the alignment behavior is somewhat different from that of ^{70,72}Se and 72,74,76 Kr. The frequency of $g_{9/2}$ quasiproton alignment



FIG. 6. (Color online) Systematics of experimental crossing frequencies for the $g_{9/2}$ proton-pair and $g_{9/2}$ neutron-pair alignments in the (a) Se and (b) Kr isotopes. The arrow indicates that only the lower limit can be deduced from the present data. Experimental data are taken from Refs. [24,26,31,35,53–59].

is 0.15 MeV lower than that of $g_{9/2}$ quasineutron. Based on the measured B(E2) value, an average prolate quadrupole deformation $\beta_2 = 0.30$ at the band crossing was deduced for ⁷⁴Se [35]. The CSM calculations in Ref. [35] have shown that if the prolate quadrupole deformation β_2 of ⁷⁴Se is larger than 0.25, the $g_{9/2}$ quasiproton alignment will occur before the $g_{9/2}$ quasineutron alignment. (3) For the N = 42 isotones ⁷⁶Se and 78 Kr, which have prolate ground states, the $g_{9/2}$ quasiproton crossing frequency is far above the $g_{9/2}$ quasineutron crossing frequency. As discussed above, such a delay of $g_{9/2}$ proton-pair alignment can be well attributed to the shape transition from prolate to oblate along the yrast line in these two isotones. (4) For the heavier nucleus 80 Kr, the nearly simultaneous $g_{9/2}$ proton and neutron pair alignments occur again. In Ref. [54], the almost simultaneous alignments of $g_{9/2}$ neutron and proton pairs in ⁸⁰Kr can be explained by assuming an oblate-to-prolate shape change. In contrast to ⁸⁰Kr, which has an oblate ground state, the N = 44 isotone ⁷⁸Se was suggested to have a prolate ground state with the Coulomb excitation experiments [27]. So far the yrast band of ⁷⁸Se was only established up to spin 12⁺ and the first band crossing was ascribed to the $g_{9/2}$ neutron-pair alignment. The $g_{9/2}$ proton-pair alignment, which is sensitive to the nuclear shape as discussed, has not been observed in this nucleus. Thus, it would be very interesting to extend the high-spin spectroscopic study to ⁷⁸Se.

In this paper we reported a spectroscopic study of ⁷⁶Se. The yrast sequence of ⁷⁶Se is first extended up to the second band crossing region. Based on the systematic comparison with the band crossings observed in the neighboring nuclei and CSM calculations, the much-delayed alignment of $g_{9/2}$ proton pair most likely is caused by a shape transition from prolate to oblate along the yrast line occurring in ⁷⁶Se. Systematic study of the underlying reason for the band crossings observed in the shape evolutions along with spin and isospin in this mass region. To get further insight into the shape evolutions in this mass region, more experimental spectroscopic and theoretical studies are needed; especially in determining the ground-state triaxiality in this region by using, for instance, the technique of rotational invariants [60,61].

ACKNOWLEDGMENTS

This work is support by the Natural Science Foundation of China under Grants No. 11175003, No. 11175108, No. 11235001, No. 11335002, No. 11320101004, No. 11375015, No. 11375017, No. 11079025, No. 11461141001, No. 11461141002, and No. J1103206, the Chinese Major State Basic Research Development Program under Grant No. 2013CB834400, and the SA/CHINA research collaboration in science and technology under Grant No. CS05-L06. The authors thank Dr. Q. W. Fan for making the target and the iThemba LABS technical staff and accelerator group for their support and providing the beam.

- [1] G. de Angelis et al., Phys. Lett. B 415, 217 (1997).
- [2] F. Becker *et al.*, Eur. Phys. J. A **4**, 103 (1999).
- [3] E. Bouchez et al., Phys. Rev. Lett. 90, 082502 (2003).
- [4] A. Gade *et al.*, Phys. Rev. Lett. **95**, 022502 (2005); **96**, 189901(E) (2006).
- [5] E. Clément et al., Phys. Rev. C 75, 054313 (2007).
- [6] H. Iwasaki et al., Phys. Rev. Lett. 112, 142502 (2014).
- [7] W. Nazarewicz et al., Nucl. Phys. A 435, 397 (1985).
- [8] A. Petrovici, K. W. Schmid, and A. Faessler, Nucl. Phys. A 665, 333 (2000).
- [9] A. Petrovici et al., J. Phys. G: Nucl. Part. Phys. 32, 583 (2006).
- [10] M. Bender, P. Bonche, and P.-H. Heenen, Phys. Rev. C 74, 024312 (2006).
- [11] M. Girod et al., Phys. Lett. B 676, 39 (2009).
- [12] P. Möller, A. J. Sierk, R. Bengtsson, H. Sagawa, and T. Ichikawa, Phys. Rev. Lett. **103**, 212501 (2009).
- [13] K. Sato and N. Hinohara, Nucl. Phys. A 849, 53 (2011).

- [14] Y. Fu et al., Phys. Rev. C 87, 054305 (2013).
- [15] B. J. Varley et al., Phys. Lett. B 194, 463 (1987).
- [16] R. B. Piercey et al., Phys. Rev. Lett. 47, 1514 (1981).
- [17] C. Chandler et al., Phys. Rev. C 56, R2924 (1997).
- [18] C. Chandler *et al.*, Phys. Rev. C **61**, 044309 (2000).
- [19] A. Görgen et al., Eur. Phys. J. A 26, 153 (2005).
- [20] J. Billowes et al., Phys. Rev. C 47, R917 (1993).
- [21] P. K. Joshi et al., Nucl. Phys. A 700, 59 (2002).
- [22] C. J. Gross et al., Nucl. Phys. A 501, 367 (1989).
- [23] S. L. Tabor *et al.*, Phys. Rev. C **41**, 2658 (1990).
- [24] J. Heese et al., Phys. Rev. C 43, R921 (1991).
- [25] D. Rudolph et al., Phys. Rev. C 56, 98 (1997).
- [26] J. J. Valiente-Dobón et al., Phys. Rev. C 71, 034311 (2005).
- [27] R. Lecomte et al., Nucl. Phys. A 284, 123 (1977).
- [28] R. Lecomte, S. Landsberger, P. Paradis, and S. Monaro, Phys. Rev. C 18, 2801 (1978).
- [29] S. M. Fischer et al., Phys. Rev. Lett. 84, 4064 (2000).

- [30] A. Obertelli *et al.*, Phys. Rev. C **80**, 031304(R) (2009).
- [31] T. Mylaeus et al., J. Phys. G: Nucl. Part. Phys. 15, L135 (1989).
- [32] A. Petrovici, K. W. Schmid, and Amand Faessler, Nucl. Phys. A 728, 396 (2003).
- [33] A. M. Hurst et al., Phys. Rev. Lett. 98, 072501 (2007).
- [34] J. Ljungvall et al., Phys. Rev. Lett. 100, 102502 (2008).
- [35] J. Döring et al., Phys. Rev. C 57, 2912 (1998).
- [36] Evaluated nuclear structure data file, http://www.nndc.bnl.gov/ ensdf/.
- [37] J. F. Sharpey-Schafer, Nucl. Phys. News 14, 5 (2004).
- [38] J. N. Scheurer *et al.*, Nucl. Instrum. Methods Phys. Res. A 385, 501 (1997).
- [39] J. Gál et al., Nucl. Instrum. Methods Phys. Res. A 516, 502 (2004).
- [40] D. C. Radford, Nucl. Instrum. Methods Phys. Res. A 361, 297 (1995).
- [41] J. C. Wells et al., Phys. Rev. C 22, 1126 (1980).
- [42] T. Matsuzaki and H. Taketani, Nucl. Phys. A 390, 413 (1982).
- [43] A. E. Kavka et al., Nucl. Phys. A 593, 177 (1995).
- [44] F. Becker et al., Nucl. Phys. A 770, 107 (2006).

- [45] S. Cwiok et al., Comput. Phys. Commun. 46, 379 (1987).
- [46] R. Wyss et al., Phys. Lett. B 215, 211 (1988).
- [47] F. R. Xu, W. Satuła, and R. Wyss, Nucl. Phys. A 669, 119 (2000).

- [48] S. Matsuki et al., Phys. Rev. Lett. 51, 1741 (1983).
- [49] W. Andrejtscheff and R. Petkov, Phys. Lett. B 329, 1 (1994).
- [50] S. Y. Wang et al., Phys. Lett. B 703, 40 (2011).
- [51] S. F. Shen, S. J. Zheng, F. R. Xu, and R. Wyss, Phys. Rev. C 84, 044315 (2011).
- [52] G. Andersson et al., Nucl. Phys. A 268, 205 (1976).
- [53] H. Sun et al., Phys. Rev. C 59, 655 (1999).
- [54] G. Mukherjee et al., Phys. Rev. C 64, 034316 (2001).
- [55] N. Yoshinaga, K. Higashiyama, and P. H. Regan, Phys. Rev. C 78, 044320 (2008).
- [56] G. Rainovski et al., J. Phys. G: Nucl. Part. Phys. 28, 2617 (2002).
- [57] R. Schwengner et al., Z. Phys. A 326, 287 (1987).
- [58] S. M. Fischer, C. J. Lister, and D. P. Balamuth, Phys. Rev. C 67, 064318 (2003).
- [59] P. Kemnitz et al., Nucl. Phys. A 425, 493 (1984).
- [60] K. Kumar, Phys. Rev. Lett. 28, 249 (1972).
- [61] D. Cline, Annu. Rev. Nucl. Part. Sci. 36, 683 (1986).