## Photoelectric disintegration of <sup>16</sup>O

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The photoelectric cross section of  ${}^{16}O(\gamma, \alpha){}^{12}C$  is estimated to be larger than the radiative capture cross section of  ${}^{12}C(\alpha, \gamma){}^{16}O$ . The predicted cross section and the angular distribution of the  $\alpha$  particle are illustrated for the future experiment. The cross section just above the  $\alpha$ -particle threshold is found to be dominated by the *E*2 excitation.

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The low-energy  ${}^{12}C(\alpha,\gamma){}^{16}O$  reaction plays a crucial role in the nucleosynthesis of elements in a star. However, the cross section is very small at energies corresponding to helium burning temperatures,  $E_{c.m.} \approx 0.3$  MeV, because of the Coulomb barrier, and it is far out of reach by present laboratory technologies [1]. ( $E_{c.m.}$  is the center-of-mass energy of the  $\alpha + {}^{12}C$  system.) To cope with these difficulties, experimental efforts have been made (e.g., [2–9]), as well as theoretical predictions [10,11]. Current experimental projects with high intensity lasers [2,12,13] are also trying to precisely measure the tiny cross section.

The probability of the photodisintegration of <sup>16</sup>O just above the  $\alpha$ -particle threshold restricts the reaction rates of the  ${}^{12}C(\alpha,\gamma){}^{16}O$  reaction. In particular, the angular distribution of the emitted  $\alpha$  particle is expected to reduce the large uncertainties of the cross section. The subthreshold  $1_1^-$  state at the excitation energy  $E_x = 7.12$  MeV has been believed to couple strongly with the  $1_2^-$  state at  $E_x = 9.585$  MeV. The strong interference in two  $1^{-}$  states leads to the large enhancement of the low-energy cross section, and it has been presumed to play an important role in the derived reaction rates at helium burning temperatures. In contrast, Refs. [10,11,14] have predicted that the E1 cross section is not enhanced by the high energy tail of the subthreshold  $1_1^-$  state because the  $\alpha + {}^{12}C$  system can be described by the weak coupling. In addition, the low-energy cross section has been reported to be dominated by the E2 transition. This is endorsed by the  $\gamma$ -ray angular distribution at  $E_{c.m.} = 1.254-1.34$  MeV [8,9] and the transparency of the  $\alpha$  + <sup>12</sup>C system at low energies [11,14–16].

In the present Brief Report, the photoelectric disintegration of <sup>16</sup>O is studied with a potential model. The expected photoelectric cross section of the <sup>16</sup>O( $\gamma, \alpha$ )<sup>12</sup>C reaction is illustrated with the reciprocity theorem. The purpose of this report is to show the prediction of the photoelectric cross section from the potential model [10].

The cross section of the  ${}^{16}O(\gamma,\alpha){}^{12}C$  reaction is given by the inverse reaction in the following expression:

$$\sigma_{\gamma\alpha}(E_{\gamma}) = \frac{k_{\rm c}^2}{2k_{\gamma}^2} \sigma_{\alpha\gamma}(E_{\rm c.m.}),\tag{1}$$

where  $k_c$  is the wave number of the relative motion between the  $\alpha$  particle and the <sup>12</sup>C nuclei;  $k_{\gamma}$  is the wave number of the photon  $k_{\gamma} = E_{\gamma}/(\hbar c)$ ,  $E_{\gamma} = E_{c.m.} + 7.162$  MeV;  $\sigma_{\alpha\gamma}$  is the capture cross section for the  ${}^{12}C(\alpha,\gamma){}^{16}O$  reaction. To obtain the predicted values, the same potential model as in [10] is adopted. The  $\alpha$  particle does not emerge easily because of the Coulomb barrier in the  $\alpha + {}^{12}C$  channel. Consequently,  $\sigma_{\gamma\alpha}$  is very small. To compensate for the rapid energy variation below the barrier, it is multiplied by the Gamow factor,  $\exp(2\pi\eta)$ , in the present report.  $\eta$  is the Sommerfeld parameter,  $\eta = Z_1 Z_2 e^2/(\hbar v)$ .  $Z_1$  and  $Z_2$  are the charges of the interacting nuclei. v is the velocity of relative motion between the  $\alpha$  particle and  ${}^{12}C$  nuclei. The calculation displayed here is the cross section obtained with the unpolarized  $\gamma$  ray, because the  ${}^{12}C(\alpha,\gamma){}^{16}O$  cross section is made by the incoherent sum over two independent polarizations of the  $\gamma$  ray [10,17].

The cross section of the photodisintegration is expected to be larger than that of the <sup>12</sup>C( $\alpha, \gamma$ )<sup>16</sup>O reaction. From Eq. (1),  $\sigma_{\gamma\alpha} \approx 50 \times \sigma_{\alpha\gamma}$  is found at  $E_{\gamma} \approx 8.41$  MeV, corresponding to  $E_{\rm c.m.} = 1.25$  MeV close to the lowest energy of the angular distribution measurement of <sup>12</sup>C( $\alpha, \gamma$ )<sup>16</sup>O [9]. This means that the study of <sup>16</sup>O( $\gamma, \alpha$ )<sup>12</sup>C gives the capture cross section and the *E*2/*E*1 ratio more accurately.

The calculated photoelectric cross section of the  ${}^{16}O(\gamma,\alpha){}^{12}C$  reaction is shown in Fig. 1. The solid curve is the result obtained from the potential model. The dashed and dotted curves are the E1 and E2 components, respectively. The vertical thin line at  $E_{\gamma} = 7.162$  MeV indicates the energy position of the  $\alpha$ -particle threshold in <sup>16</sup>O. The  $\alpha$  particle is emitted above this energy. The peak at the  $1_2^-$  state can be seen at  $E_{\gamma} \approx 9.5$  MeV. The E1 component is approximately constant below  $E_{\gamma} = 8$  MeV. The photoelectric disintegration is found to be dominated by the E2 excitation below  $E_{\gamma} \approx$ 8 MeV. This is due to the high energy tail of the subthreshold  $2_1^+$ state at  $E_x = 6.92$  MeV, which has the well-developed  $\alpha + {}^{12}C$ cluster structure [14]. The photonuclear reaction may remind you of the discussion about the dipole excitation. However, the photomagnetic dipole excitation (s wave) is forbidden. The electric dipole transition (p wave) is hindered by the isospin selection rule. Thus, the E2 excitation (d wave) could be the dominant mode of the transition in the vicinity of the  $\alpha$ -particle threshold. The E3 and E4 transitions (f and g waves) are negligible.

The  $\sigma_{\gamma\alpha} \exp(2\pi \eta)$  in Fig. 1 resembles the astrophysical *S* factor in the <sup>12</sup>C( $\alpha,\gamma$ )<sup>16</sup>O reaction [10]. The energy variation, in fact, corresponds to  $\sigma_{\gamma\alpha} \exp(2\pi \eta) \propto S/E_{\gamma}^2$ .

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FIG. 1. Photoelectric cross section for the  ${}^{16}O(\gamma, \alpha){}^{12}C$  reaction as a function of the  $\gamma$ -ray energy. The solid curve is the result obtained from the potential model. The dashed and dotted curves are the *E*1 and *E*2 components, respectively. The cross section is multiplied by  $\exp(2\pi\eta)$  to compensate for the rapid energy variation in the vicinity of the  $\alpha$ -particle threshold. The vertical thin line at  $E_{\gamma} = 7.162$  MeV indicates the energy position of the threshold.

The angular distributions of the  $\alpha$  particle at  $E_{\gamma} = 8.0, 8.5, 9.0,$  and 9.5 MeV are shown in Fig. 2. The solid curves are the results obtained from the potential model.  $\theta$  is the angle between  $\mathbf{k}_c$  and  $\mathbf{k}_{\gamma}$ . At  $E_{\gamma} = 9.5$  MeV, the angular distribution appears to be the single peak due to the  $1_2^-$  state at  $E_x = 9.585$  MeV. The angular distribution becomes the double peaks at  $E_{\gamma} = 8.0$  MeV. The dashed and dotted curves are the pure E1 and E2 components. The E2 contribution caused by the subthreshold  $2_1^+$  state interferes in the dipole component.

The differential cross section at  $\theta = 90^{\circ}$  is made only from the *E*1 component, because the *E*2 component vanishes at this angle. Using the 90° cross section, the integrated *E*1 cross section is expressed as

$$\sigma_{\gamma\alpha}^{E1}(E_{\gamma}) = \frac{8\pi}{3} \frac{d\sigma_{\gamma\alpha}}{d\Omega} \bigg|_{\theta = 90^{\circ}}.$$
 (2)

The dashed curve in Fig. 1 can be confirmed from this relation. The numerical values are listed in Table I.

The 90° cross section basically determines the E1 contribution of the photodisintegration of <sup>16</sup>O. As shown in Fig. 2, the angular distribution has the minimum at  $\theta \approx 90^{\circ}$  at low energies, because the E2 contribution becomes large. In addition, the total cross section is very small below the barrier. In this circumstance, the E1 component may be susceptible to the background noise. The absolute value of the E1 cross section could not be determined without precise measurement of the angular distribution. Hopefully, the E2/E1 ratio will be provided more accurately by the future experiment.

The calculated  $\sigma_{E2}/\sigma_{E1}$  value is shown in Fig. 3. The solid curve is the result obtained from the potential model. As a comparison, the calculated value of [13] is also shown by the dot-dashed curve. The difference between the two curves at low energies stems from the assumed reaction mechanism. Reference [13] uses the compound nuclear model with the strong coupling. The potential model describes the direct reaction process, in which only a few degrees of freedom of motion are activated.



FIG. 2. Differential cross sections for the  ${}^{16}O(\gamma,\alpha){}^{12}C$  reaction at  $E_{\gamma} = 8.0, 8.5, 9.0$ , and 9.5 MeV. The solid curves are the results obtained from the potential model. The dashed and dotted curves are the pure E1 and E2 components.

Finally, let me describe a supplementary explanation about the angular distribution with a linearly polarized  $\gamma$  ray. The angular distribution of  ${}^{16}O(\vec{\gamma},\alpha){}^{12}C$  is given from [17,18] as follows:

$$\frac{d\sigma_{\bar{\gamma}\alpha}}{d\Omega}(\theta,\phi) = \frac{d\sigma_{\gamma\alpha}}{d\Omega}(\theta)[1+f\cos(2\phi)],\tag{3}$$

where  $d\sigma_{\gamma\alpha}/d\Omega$  is the angular distribution by the unpolarized  $\gamma$  ray shown in Fig. 2; f is the linear polarization of the  $\gamma$ -ray beam,  $f \approx 1$ ; the  $\phi$  is the angle between the direction of polarization and the x axis. The Cartesian coordinate is defined by the right-handed system: the z axis is along with  $k_{\gamma}$ ; the y axis is along with the vector  $k_{\gamma} \times k_c$ . The reaction occurs on the x - z plane. The angular distribution varies with a factor of 0–2 by the direction of the polarization of  $\gamma$ -ray

TABLE I. Photoelectric dipole cross section and 90° cross section. The value of  $\exp(2\pi \eta)$  is also listed.

$E_{\gamma}$ (MeV)	$d\sigma_{\gamma\alpha}/d\Omega (90^\circ) (b/sr)$	$\sigma_{\gamma\alpha}^{E1}$ (b)	$\exp(2\pi\eta)$
8.0	$3.11 \times 10^{-12}$	$2.60 \times 10^{-11}$	$5.72 \times 10^{9}$
8.5	$4.22 \times 10^{-10}$	$3.53 \times 10^{-9}$	$5.27 \times 10^{7}$
9.0	$1.19 \times 10^{-8}$	$9.98  imes 10^{-8}$	$3.88 \times 10^{6}$
9.5	$3.47 \times 10^{-7}$	$2.91 \times 10^{-6}$	$6.94 \times 10^{5}$



FIG. 3. Ratio of *E*2 component to *E*1 in  ${}^{16}O(\gamma, \alpha){}^{12}C$ . The solid curve is the result obtained from the potential model. The dot-dashed curve is from the resonance theory [13], of which the value is taken from Fig. 2 of [13].

- [1] C. E. Rolfs and W. S. Rodney, *Cauldrons in the Cosmos* (University of Chicago, Chicago, 1988).
- [2] M. Gai, Nucl. Phys. A 928, 313 (2014).
- [3] M. Gai, Phys. Rev. C 88, 062801 (2013).
- [4] C. R. Brune and D. B. Sayre, J. Phys.: Conf. Ser. 420, 012140 (2013).
- [5] D. Schürmann, L. Gialanella, R. Kunz, and F. Strieder, Phys. Lett. B 711, 35 (2012).
- [6] R. Plag, R. Reifarth, M. Heil, F. Käppeler, G. Rupp, F. Voss, and K. Wisshak, Phys. Rev. C 86, 015805 (2012).
- [7] H. Makii, Y. Nagai, T. Shima, M. Segawa, K. Mishima, H. Ueda, M. Igashira, and T. Ohsaki, Phys. Rev. C 80, 065802 (2009).
- [8] M. Assunção, M. Fey, A. Lefebvre-Schuhl, J. Kiener, V. Tatischeff, J. W. Hammer, C. Beck, C. Boukari-Pelissie, A. Coc, J. J. Correia *et al.*, Phys. Rev. C 73, 055801 (2006).
- [9] R. Kunz, M. Jaeger, A. Mayer, J. W. Hammer, G. Staudt, S. Harissopulos, and T. Paradellis, Phys. Rev. Lett. 86, 3244 (2001).
- [10] M. Katsuma, Phys. Rev. C 78, 034606 (2008); 81, 029804 (2010); Astrophys. J. 745, 192 (2012).

beam [13]. From Eq. (3), the  $d\sigma_{\gamma\alpha}/d\Omega$  of Fig. 2 is found to give the basic quantities of the photoelectric disintegration of <sup>16</sup>O, even when a linearly polarized  $\gamma$  ray is used.

In summary, the  ${}^{16}O(\gamma, \alpha){}^{12}C$  reaction has been studied with the potential model. The calculated photoelectric cross section and the angular distribution of the  $\alpha$  particle have been shown in figures. The cross section just above the  $\alpha$ -particle threshold is dominated by the *E*2 excitation. The cross section of the photoelectric disintegration is found to be larger than that of the  ${}^{12}C(\alpha,\gamma){}^{16}O$  reaction. Therefore, the forthcoming experimental projects [2,12,13] are expected to determine the  ${}^{12}C(\alpha,\gamma){}^{16}O$  cross section more accurately.

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- [11] M. Katsuma, arXiv:1404.3966.
- [12] M. Ahmed, A. Champagne, B. Holstein, C. Howell, W. Snow, R. Springer, and Y. Wu, arXiv:1307.8178.
- [13] Y. Xu, W. Xu, Y. G. Ma, W. Guo, J. G. Chen, X. Z. Cai, H. W. Wang, C. B. Wang, G. C. Lu, and W. Q. Shen, Nucl. Instrum. Methods A 581, 866 (2007).
- [14] M. Katsuma, Phys. Rev. C 81, 067603 (2010); J. Phys. G 40, 025107 (2013); EPJ Web Conf. 66, 03041 (2014).
- [15] P. Tischhauser, A. Couture, R. Detwiler, J. Görres, C. Ugalde, E. Stech, M. Wiescher, M. Heil, F. Käppeler, R. E. Azuma *et al.*, Phys. Rev. C **79**, 055803 (2009).
- [16] R. Plaga, H. W. Becker, A. Redder, C. Rolfs, H. P. Trautvetter, and K. Langanke, Nucl. Phys. A 465, 291 (1987).
- [17] H. J. Rose and D. M. Brink, Rev. Mod. Phys. 39, 306 (1967).
- [18] H. R. Weller, J. Langenbrunner, R. M. Chasteler, E. L. Tomusiak, J. Asai, R. G. Seyler, and D. R. Lehman, At. Data Nucl. Data Tables 50, 29 (1992).