Collective Hamiltonian for wobbling modes

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The simple, longitudinal, and transverse wobblers are systematically studied within the framework of a collective Hamiltonian, where the collective potential and mass parameter included are obtained based on the tilted axis cranking approach. Solving the collective Hamiltonian by diagonalization, the energies and the wave functions of the wobbling states are obtained. The obtained results are compared with those by the harmonic approximation formula and particle rotor model. The wobbling energies calculated by the collective Hamiltonian are closer to the exact solutions by the particle rotor model than the harmonic approximation formula. It is confirmed that the wobbling frequency increases with the rotational frequency in simple and longitudinal wobbling motions while decreases in transverse wobbling motion. These variation trends are related to the stiffness of the collective potential in the collective Hamiltonian.

momentum *I*, are

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I. INTRODUCTION

Atomic nuclei possess a wide variety of shapes in both their ground and excited states. The shapes may range from spherical to deformed, from quadrupole to octupole, and even more exotic shapes, such as superdeformed and tetrahedral. For deformed nuclei, they in general possess an axially symmetric shape. The loss of axial symmetry would lead to a triaxial shape. The triaxiality has been invoked to describe many interesting phenomena including the *γ* band [\[1\]](#page-11-0), signature inversion [\[2\]](#page-11-0), anomalous signature splitting [\[3\]](#page-11-0), chiral symmetry breaking [\[4–6\]](#page-11-0), and the wobbling motion [\[1\]](#page-11-0). The wobbling motion and chirality are regarded as fingerprints of stable triaxial nuclei.

The wobbling motion within nuclear rotation was originally introduced by Bohr and Mottelson [\[1\]](#page-11-0) in the context of the triaxial rotor model (TRM). For a rotating triaxial even-even nuclei, the rotation motions about any of the axes are all possible and the corresponding TRM Hamiltonian reads

$$
\hat{H}_{\text{rot}} = \frac{\hat{I}_1^2}{2\mathcal{J}_1} + \frac{\hat{I}_2^2}{2\mathcal{J}_2} + \frac{\hat{I}_3^2}{2\mathcal{J}_3},\tag{1}
$$

with three distinct moments of inertia \mathcal{J}_k (usually defines \mathcal{J}_1) as maximal) associating with each of the principle axes. It is pointed out that although the triaxial nucleus energetically favors the rotation about the axis with the largest moment of inertia (i.e., one axis), contributions from rotations about the other two axes (two and three axes) would quantum mechanically disturb this rotation and force the angular momentum vector off the one axis. As a consequence, besides the uniform rotation about one axis, there is a wobbling motion [\[1\]](#page-11-0). The energies of wobbling states, characterized by the wobbling phonon number *n* together with total angular

$$
E(n,I) = \frac{I(I+1)}{2\mathcal{J}_1} + \left(n + \frac{1}{2}\right)\hbar\Omega_{\text{wob}}.\tag{2}
$$

The quantum number *n* describes the wobbling motion of the axes with respect to the direction of *I* . For small amplitudes, this motion has the character of a harmonic vibration with wobbling frequency given by

$$
\hbar\Omega_{\text{wob}} = 2I\sqrt{\left(\frac{\hbar^2}{2\mathcal{J}_2} - \frac{\hbar^2}{2\mathcal{J}_1}\right)\left(\frac{\hbar^2}{2\mathcal{J}_3} - \frac{\hbar^2}{2\mathcal{J}_1}\right)},\qquad(3)
$$

which is related to the moments of inertia of three axes and found to be proportional to the spin. Similar to Ref. [\[7\]](#page-11-0), such type of wobbling motion for a triaxial rotor is also denoted as *simple wobbler* at the present investigation.

The wobbling motion appears not only in the even-even nuclei but also in the odd-*A* nuclei. For rotating odd-*A* triaxial nuclei, there are two types of wobbling motions suggested by Frauendorf and Dönau [[7\]](#page-11-0) very recently according to the relation between the orientation of quasiparticle angular momentum vector with respect to the rotor axis with the largest moment of inertia. If the quasiparticle angular momentum vector is aligned with the axis with the largest moment of inertia, it is called *longitudinal wobbler*. If the quasiparticle angular momentum vector is perpendicular to the axis with the largest moment of inertia, it is called *transverse wobbler*. Assuming frozen alignment of the quasiparticle with one of the rotor axes and harmonic oscillations (HFA), a rather simple analytic expression for the wobbling frequency of these two types of wobbling motions is derived [\[7\]](#page-11-0). According to this analytic expression, the increasing trend of wobbling frequency for a longitudinal wobbling motion and decreasing trend for transverse wobbling can be expected.

On the experimental side, although the wobbling phenomenon has been predicted for a long time [\[1\]](#page-11-0), it was not observed until the beginning of this century when the first

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experimental evidence was reported in 163 Lu [\[8\]](#page-11-0). Subsequently, it has been extensively studied in the triaxial strongly deformed (TSD) region around $N = 94$, where the wobbling bands have been identified in 161,163,165,167 Lu $[9-14]$ and 167 Ta [\[15\]](#page-11-0). All wobbling bands in this mass region are based on the $\pi i_{13/2}$ configuration. Very recently, a new candidate wobbling band was proposed in 135 Pr [\[7\]](#page-11-0), which is built on $\pi h_{11/2}$ configuration, differing from the configuration of previous known examples. For even-even nuclei, however, the wobbling spectra are scarce since stable triaxial ground states are rare. The best example identified so far is 112 Ru [\[16\]](#page-11-0).

On the theoretical side, the wobbling motion was first investigated by TRM [\[1\]](#page-11-0). Following the discovery of the first wobbling structure in odd- A^{163} Lu [\[8\]](#page-11-0), the quantal particle rotor model (PRM) was used to describe the wobbling mode, see Refs. [\[7,17–20\]](#page-11-0). Based on the framework of mean field theory, there are many efforts to extend the cranking model to study the wobbling motion. Due to the mean-field approximation, the cranking model yields only the yrast sequence for a given configuration. Therefore, in order to describe the wobbling excitations, one has to go beyond the mean-filed approximation. At present, this has been done by incorporating the quantum correlations by means of random phase approximation (RPA) [\[21–28\]](#page-11-0) or by the generator coordinate method after angular momentum projection $(GCM + AMP)$ based on the cranking intrinsic states [\[29\]](#page-11-0).

Another promising method is to construct a collective Hamiltonian on the top of cranking mean field solutions. By taking into account the quantum fluctuation along the collective degree of freedom, the collective Hamiltonian goes beyond the mean-field approximation and restores the broken symmetry [\[30\]](#page-11-0). This has been implemented based on the framework of tilted axis cranking (TAC) single-*j* shell model to investigate the chiral vibration and rotation motions [\[31,32\]](#page-11-0). The chiral symmetry broken in the intrinsic reference frame is restored and chiral doublet bands are obtained in the laboratory reference frame. For the wobbling motion, the wobbling states are formed due to the quantum fluctuation of the total angular momentum deviating from the principle axes of the rotor. It is thus interesting to extend the collective Hamiltonian to describe the phenomenon of wobbling motion.

In this work, the collective Hamiltonian will be extended to study the simple, longitudinal, and transverse wobbling motions, in particular, to examine the trend of the wobbling frequency with respect to the rotational frequency. In the collective Hamiltonian, the collective potentials are calculated from the TAC model and the mass parameter is obtained with the assumption of harmonic approximation (HA) for simple wobbling motion or HFA approximation for longitudinal and transverse wobbling motions. The energy levels and wave function of wobbling states are obtained by diagonalizing the collective Hamiltonian. The corresponding energy spectra will be in comparison with the results obtained by HA (HFA) analytic expression as well as TRM (PRM) for simple (longitudinal and transverse) wobbling to evaluate the accuracy of the collective Hamiltonian.

The paper is organized as follows. In Sec. II, a brief introduction to the collective Hamiltonian is given. The corresponding numerical details adopted in the calculations are presented in Sec. [III.](#page-3-0) In Sec. [IV,](#page-4-0) the obtained potential energy and the mass parameter are respectively shown for the three types of wobbling motions and the corresponding energy levels and wave functions obtained by the collective Hamiltonian are discussed in details. A brief summary is given in Sec. [V.](#page-10-0)

II. THEORETICAL FRAMEWORK

The collective Hamiltonian, in terms of a few numbers of collective coordinates and momenta, is an effective method for describing various collective processes which involve small velocities. The well-known Bohr Hamiltonian describes the collective rotational and vibrational degrees of freedom with the five collective intrinsic variables β , γ , and Euler angles Ω [\[1\]](#page-11-0). In Ref. [\[30\]](#page-11-0), to describe the chiral motions in triaxial rotational nuclei, a collective Hamiltonian based on the TAC solutions was constructed. Therein, the orientation of nucleus in rotating mean field, described by polar angle *θ* and azimuth angle φ in the spherical coordinate as illustrated in Fig. 1, is considered as collective variable. As the motion along the φ direction is much easier than the *θ* direction, the collective Hamiltonian has been restricted to one-dimensional motion along φ direction [\[30\]](#page-11-0).

For the wobblers caused by the quantum fluctuation of the total angular momentum orientation, the azimuth angle φ can also be taken as collective coordinate to the wobbling motions and the wobbling excitation is restricted to the onedimensional motion along the φ direction. It is necessary to mention that in a semiclassical model for wobbling motion, the azimuth angle defined as $cos^{-1}(\frac{\langle \hat{I}_1 \rangle}{I_2})$ has been interpreted as the wobbling angle of the total angular momentum vector [\[33\]](#page-11-0). The two azimuth angles are consistent with each other when $\theta = 90^\circ$.

The detailed theoretical framework of collective Hamiltonian based on the TAC solutions has been formulated in Ref. [\[30\]](#page-11-0). The formalism can be analogized to describe the wobbling motion. Here for completeness, a brief introduction to the formalism is presented.

FIG. 1. (Color online) Orientation of the rotational frequency *ω* with respect to the principal axes.

A. Collective Hamiltonian

Taking φ as the collective variable, the classical form of the collective Hamiltonian is written as the sum of kinetic and potential terms

$$
H_{\text{coll}} = T_{\text{kin}}(\varphi) + V(\varphi) = \frac{1}{2}B(\varphi)\dot{\varphi}^{2} + V(\varphi), \qquad (4)
$$

where $V(\varphi)$ is collective potential and $B(\varphi)$ mass parameter. The quantized form of the collective Hamiltonian is obtained according to the Pauli prescription [\[34\]](#page-11-0)

$$
\hat{H}_{\text{coll}} = -\frac{\hbar^2}{2\sqrt{B(\varphi)}}\frac{\partial}{\partial\varphi}\frac{1}{\sqrt{B(\varphi)}}\frac{\partial}{\partial\varphi} + V(\varphi). \tag{5}
$$

By solving this Hamiltonian on the basis states with appropriate boundary condition on φ , e.g., box boundary condition [\[30\]](#page-11-0), the wobbling levels and corresponding wave functions can be obtained.

B. Collective potential $V(\varphi)$

Both the collective potential $V(\varphi)$ and the mass parameter $B(\varphi)$ in the collective Hamiltonian (4) can be determined based on the TAC model.

Let us first discuss $V(\varphi)$ for the cases of longitudinal and transverse wobblers. For schematic discussions, we consider a system of a high-*j* particle coupled to a triaxial rotor. The cases for more than one particle coupled to triaxial rotor can be easily extended as well. The cranking Hamiltonian reads

$$
\hat{h}' = \hat{h}_{def} - \omega \cdot \hat{j}, \n\omega = (\omega \sin \theta \cos \varphi, \omega \sin \theta \sin \varphi, \omega \cos \theta),
$$
\n(6)

where \hat{j} is the single particle angular momentum and deformed single particle Hamiltonian \hat{h}_{def} is taken as the single-*j* shell Hamiltonian

$$
\hat{h}_{\text{def}} = \frac{1}{2} C \left\{ \left(\hat{j}_3^2 - \frac{j(j+1)}{3} \right) \cos \gamma + \frac{1}{2\sqrt{3}} (\hat{j}_+^2 + \hat{j}_-^2) \sin \gamma \right\}.
$$
\n(7)

Diagonalizing the cranking Hamiltonian, one obtains the total Routhian

$$
E'(\theta, \varphi) = \langle h' \rangle - \frac{1}{2} \sum_{k=1}^{3} \mathcal{J}_k \omega_k^2, \quad \mathcal{J}_k : \text{moments of inertia.}
$$
\n(8)

Minimizing the total Routhian with respect to θ for given φ , the collective potential $V(\varphi)$ is finally obtained.

For simple wobbler, i.e., a simple triaxial rotor without coupling any particles, the total Routhian (8) is degenerated to

$$
E'(\theta, \varphi) = -\frac{1}{2} \sum_{k=1}^{3} \mathcal{J}_k \omega_k^2, \tag{9}
$$

and similarly the collective potential $V(\varphi)$ is obtained by minimizing the total Routhian with respect to θ for given φ .

C. Mass parameter *B*

In Ref. [\[30\]](#page-11-0), the mass parameter is calculated by the cranking method. However, for the simple wobbler with a frozen rotor assumption, the full single particle levels are not available in the present TAC model. Therefore, in the following, we apply another approximated method to determine the mass parameter.

Before discussing how to calculate the mass parameter, it is worth noting once more that as pointed out by Bohr and Mottelson [\[1\]](#page-11-0), the wobbling motion as a small amplitude vibration has the character of a harmonic oscillation with frequency Ω_{wob} . As is well known, the oscillation frequency Ω for a harmonic oscillator system is related to the mass parameter *B* of the oscillator and the stiffness parameter *C* of the harmonic oscillator potential by

$$
\Omega = \sqrt{\frac{C}{B}}.\tag{10}
$$

Therefore, once the stiffness parameter *C* and oscillation frequency Ω are determined, the mass parameter *B* can be obtained.

To extract the stiffness parameter *C* of the collective potential $V(\varphi)$, one can expand the collective potential $V(\varphi)$ by Taylor series at $\varphi = 0^{\circ}$ up to $\sim \varphi^2$ terms, i.e., the harmonic approximation (HA) is adopted. For the total Routhian surface calculation (9) of a simple wobbler, one can find that its minimum along the θ direction is always at $\theta = 90^\circ$ for any value of φ . Therefore, the collective potential becomes

$$
V(\varphi) = -\frac{1}{2}\omega^2 (\mathcal{J}_1 \cos^2 \varphi + \mathcal{J}_2 \sin^2 \varphi)
$$
 (11)

$$
\approx -\frac{1}{2}\mathcal{J}_1\omega^2 + \frac{1}{2}\omega^2(\mathcal{J}_1 - \mathcal{J}_2)\varphi^2, \quad \text{for } \varphi \to 0^\circ. \tag{12}
$$

Equation (12) suggests that the collective potential can be regarded as the sum of a rotational energy term along the one axis with frequency*ω* and a harmonic oscillation potential term along φ direction with stiffness parameter $C = \omega^2(\mathcal{J}_1 - \mathcal{J}_2)$. Thus the wobbling frequency Ω_{wob} and the mass parameter *B* are related each other by

$$
\hbar\Omega_{\text{wob}} = \hbar\sqrt{\frac{C}{B}} = \hbar\omega\sqrt{\frac{\mathcal{J}_1 - \mathcal{J}_2}{B}}.\tag{13}
$$

To determine the mass parameter B in Eq. (13), we further recall the wobbling frequency [\(3\)](#page-0-0) given by Bohr and Mottelson [\[1\]](#page-11-0)

$$
\hbar\Omega_{\text{wob}} = 2I \sqrt{\left(\frac{\hbar^2}{2\mathcal{J}_2} - \frac{\hbar^2}{2\mathcal{J}_1}\right) \left(\frac{\hbar^2}{2\mathcal{J}_3} - \frac{\hbar^2}{2\mathcal{J}_1}\right)}
$$

$$
= \frac{\hbar^2 I}{\mathcal{J}_1} \sqrt{\frac{(\mathcal{J}_1 - \mathcal{J}_2)(\mathcal{J}_1 - \mathcal{J}_3)}{\mathcal{J}_3 \mathcal{J}_2}}
$$

$$
= \hbar\omega \sqrt{\frac{(\mathcal{J}_1 - \mathcal{J}_2)(\mathcal{J}_1 - \mathcal{J}_3)}{\mathcal{J}_3 \mathcal{J}_2}}.
$$
(14)

Combining Eqs. (13) and (14) , the mass parameter is obtained for simple wobbler

$$
B = \frac{\mathcal{J}_2 \mathcal{J}_3}{\mathcal{J}_1 - \mathcal{J}_3}.\tag{15}
$$

It is determined only by the moments of inertia of three principal axes and independent of rotational frequency.

For longitudinal and transverse wobblers, we introduce the harmonic frozen alignment (HFA) approximation as in Ref. [\[7\]](#page-11-0), i.e., the angular momentum of the odd particle is assumed to be firmly aligned with the short axis (one axis) and can be considered as a number. Then for a given rotational frequency *ω*, the moment of inertia of the one axis is treated as a *ω*-dependent effective moment of inertia

$$
\mathcal{J}_1^*(\omega) = \frac{\mathcal{J}_1 \omega + j}{\omega} = \mathcal{J}_1 + \frac{j}{\omega}.\tag{16}
$$

The odd-particle contributes a *ω*-dependent term to the effective moment of inertia, which will decrease with the increasing rotational frequency.

Similar to a simple wobbler, it can be also found that for the longitudinal and transverse wobblers the collective potential obtained from the total Routhian surface calculation $E'(\theta, \varphi)$ [\(8\)](#page-2-0) is minimized at $\theta = 90^\circ$ for any given φ . Therefore, the collective potential is written as

$$
V(\varphi) = \langle \hat{h}_{\text{def}} \rangle - \omega j \cos \varphi
$$

\n
$$
- \frac{1}{2} \omega^2 (\mathcal{J}_1 \cos^2 \varphi + \mathcal{J}_2 \sin^2 \varphi)
$$
(17)
\n
$$
\approx \langle \hat{h}_{\text{def}} \rangle - \omega j \left(1 - \frac{\varphi^2}{2} \right) - \frac{1}{2} \mathcal{J}_1 \omega^2
$$

\n
$$
+ \frac{1}{2} \omega^2 (\mathcal{J}_1 - \mathcal{J}_2) \varphi^2, \text{ for } \varphi \to 0
$$

\n
$$
= \langle \hat{h}_{\text{def}} \rangle - \frac{1}{2} \omega j - \frac{1}{2} \left(\mathcal{J}_1 + \frac{j}{\omega} \right) \omega^2
$$

\n
$$
+ \frac{1}{2} \omega^2 \left[\left(\mathcal{J}_1 + \frac{j}{\omega} \right) - \mathcal{J}_2 \right] \varphi^2
$$

\n
$$
= \langle \hat{h}_{\text{def}} \rangle - \frac{1}{2} \omega j - \frac{1}{2} \mathcal{J}_1^* \omega^2 + \frac{1}{2} \omega^2 [\mathcal{J}_1^* (\omega) - \mathcal{J}_2] \varphi^2.
$$
(18)

This formula is similar to Eq. (12) except that the moment of inertia of the one axis \mathcal{J}_1 has been replaced by the effect moment of inertia $\mathcal{J}_1^*(\omega)$, thereby one expects that the mass parameter for longitudinal and transverse wobblers has the similar form as a simple wobbler

$$
B(\omega) = \frac{\mathcal{J}_2 \mathcal{J}_3}{\mathcal{J}_1^*(\omega) - \mathcal{J}_3} = \frac{\mathcal{J}_2 \mathcal{J}_3}{(\mathcal{J}_1 - \mathcal{J}_3) + \frac{j}{\omega}}.
$$
(19)

Differing from the mass parameter (15) for a simple wobbler, it is determined not only by the moments of inertia of three principal axes, but also by the angular momentum of the odd particle and the rotational frequency. As the rotational frequency increases, the mass parameter for longitudinal and transverse wobblers will increase as well.

The wobbling frequency for the longitudinal and transverse wobbling motions can be then obtained from Eq. (13)

$$
\hbar\Omega_{\text{wob}} = \sqrt{\frac{\mathcal{J}_1^*(\omega) - \mathcal{J}_2}{B(\omega)}} \hbar\omega
$$

$$
= \hbar \sqrt{\frac{[(\mathcal{J}_1 - \mathcal{J}_3)\omega + j][(\mathcal{J}_1 - \mathcal{J}_2)\omega + j]}{\mathcal{J}_2\mathcal{J}_3}}.
$$
(20)

This formula is nothing but the HFA formula in Ref. [\[7\]](#page-11-0) by replacing the spin with $\mathcal{J}_1\omega + j$. For longitudinal wobbling motion, since $\mathcal{J}_1 > \mathcal{J}_2, \mathcal{J}_3$, the wobbling frequency increases with the rotational frequency. While for transverse wobbling motion, since $\mathcal{J}_2 > \mathcal{J}_1$, the wobbling frequency decreases with the rotational frequency, and will reach zero at a critical rotational frequency $\hbar \omega_c = j/(\mathcal{J}_2 - \mathcal{J}_1)$.

III. NUMERICAL DETAILS

In the following calculations, a triaxial rotor with the deformation parameters $\beta = 0.25$ and $\gamma = -30^\circ$ is considered to investigate the simple wobbling motion. Following the notation in Ref. [\[35\]](#page-11-0), for such deformation, three principal axes, the one, tow, and three axis respectively correspond to short (*s*), intermediate (*i*), and long (*l*) axis. For the investigation of the longitudinal and transverse wobbling motions, the triaxial rotor is assumed to be further coupled with a $h_{11/2}$ proton particle. Thus the proton aligns its angular momentum along short axis (namely, one axis). The longitudinal (transverse) wobblers are achieved by choosing the one axis to be (perpendicular to) the axis with largest moments of inertia.

With regard to the moments of inertia, both the rigid body type

$$
\mathcal{J}_k^{\text{rig}} = \frac{2}{5} m A R_0^2 \left[1 - \sqrt{\frac{5}{4\pi}} \beta \cos \left(\gamma - \frac{2\pi}{3} k \right) \right]
$$

$$
= \mathcal{J}_0^{\text{rig}} \left[1 - \sqrt{\frac{5}{4\pi}} \beta \cos \left(\gamma - \frac{2\pi}{3} k \right) \right], \quad k = 1, 2, 3 \tag{21}
$$

and the irrotational flow type

$$
\mathcal{J}_k^{\text{irr}} = \frac{3}{2\pi} m A R_0^2 \beta^2 \sin^2 \left(\gamma - \frac{2\pi}{3} k\right)
$$

$$
= \mathcal{J}_0^{\text{irr}} \sin^2 \left(\gamma - \frac{2\pi}{3} k\right), \quad k = 1, 2, 3 \qquad (22)
$$

are often assumed [\[35\]](#page-11-0). $\mathcal{J}_k^{\text{rig}}$ shows less dependence on the deformation β than $\mathcal{J}_k^{\text{irr}}$ (∼ β^2). In the γ dependence, \mathcal{J}^{irr} vanishes about the symmetry axes while \mathcal{T}^{rig} does not and the largest moment of inertia axes of them are different. For the present deformation parameters $\beta = 0.25$ and $\gamma =$ −30◦, the largest moment of inertia axis is the one axis (*s*-axis) for rigid body type while the two axis (*i*-axis) for irrotational flow type.

In the present investigation, the wobbling angle φ in a collective Hamiltonian is restricted to $-\pi/2 \le \varphi \le \pi/2$, or in other words the wobbling motion happens around the one axis. For a simple wobbler, the rigid body type of moment of inertia (21) is adopted for the one axis being the axis with the largest moment of inertia. Similarly, the rigid body type of moment of inertia is also applied to longitudinal wobbler so that the orientation of the proton angular momentum (one axis) is parallel to the axis with largest moments of inertia. For the transverse wobbler, while the orientation of proton angular momentum (one axis) is required to be perpendicular to the axis with largest moments of inertia, the irrotational flow type of moment of inertia [\(22\)](#page-3-0) is adopted. In the calculations, the constants $\mathcal{J}_{0 \text{r} \text{io}}^{\text{rig}}$ and $\mathcal{J}_0^{\text{irr}}$ in Eqs. [\(21\)](#page-3-0) and [\(22\)](#page-3-0) are respectively taken as $J_0^{\text{rig}} = 256\pi/15 \hbar^2/\text{MeV}$ and $\mathcal{J}_0^{\text{irr}} = 40 \hbar^2/\text{MeV}$. In details, for a rigid type of moment of inertia, $\mathcal{J}_1^{\text{rig}} = 60.9 \hbar^2/\text{MeV}$, $\mathcal{J}_2^{\text{rig}} = 53.6 \hbar^2/\text{MeV}$, and $\mathcal{J}_3^{\text{rig}} = 46.3 \hbar^2/\text{MeV}$; for an irrotational flow type of moment of inertia, $\mathcal{J}_2^{\text{irr}} = 40 \hbar^2/\text{MeV}$ and $\mathcal{J}_1^{\text{irr}} = \mathcal{J}_3^{\text{irr}} = 10 \hbar^2/\text{MeV}$.

IV. RESULTS AND DISCUSSION

A. Simple wobbler

We first present the results of our calculations for the simple wobbling motion by means of the TAC model and collective Hamiltonian. As described in Sec. [II,](#page-1-0) the collective potential and the mass parameter included in the collective Hamiltonian are respectively calculated by the TAC model and Eq. [\(15\)](#page-2-0). The obtained wobbling energies will be compared with the HA formula and the exact TRM.

1. Collective potential

In the contour plots of Figs. $2(a)-2(d)$, the total Routhian surface calculation $E'(\theta, \varphi)$ [\(9\)](#page-2-0) in the (θ, φ) plane at the rotational frequencies $\hbar \omega = 0.1, 0.2, 0.3,$ and 0.4 MeV are shown. All the potential energy surfaces are symmetrical with respect to $\varphi = 0^\circ$ line. With the increasing rotational frequency, the minima in the potential energy surfaces always locate at $(\theta = 90^\circ, \varphi = 0^\circ)$, which corresponds to uniform rotation about the axis with the largest moment of inertia.

Minimizing the total Routhian $E'(\theta, \varphi)$ with θ for a given φ , we find that the minimum along the θ direction is always at $\theta = 90^\circ$ for any value of φ at each rotational frequency. The corresponding extracted collective potentials $V(\varphi)$ are shown in the upper panels of Figs. $2(a)-2(d)$, respectively, for $\hbar \omega = 0.1, 0.2, 0.3,$ and 0.4 MeV. Again, the potential energy

FIG. 2. (Color online) Lower panels: Contour plots of the total Routhian surface calculation $E'(\theta, \varphi)$ for a triaxial rigid body rotor with $\gamma = -30^\circ$ at the frequencies $\hbar \omega = 0.1, 0.2, 0.3$, and 0.4 MeV. All energies at each rotational frequency are normalized with respect to the absolute minimum. Upper panels: The collective potential $V(\varphi)$ as a function of φ extracted from the corresponding total Routhian surface calculation.

is symmetrical about $\varphi = 0^\circ$ in correspondence with the results displayed in the lower panels of Figs. $2(a)-2(d)$. For all cases, the potential $V(\varphi)$ is a harmonic oscillator type that has only one minimum at $\varphi = 0^\circ$, corresponding to the rotation about the one axis. The stiffness of the collective potential becomes larger as the rotational frequency increases. This is directly reflected by the increase of the energy difference between $\varphi = \pm 90^\circ$ and $\varphi = 0^\circ$. For example, the value is only ~30 keV at $\hbar \omega = 0.1$ MeV while reaches ~600 keV at $\hbar \omega = 0.4$ MeV.

2. Collective levels and wave functions

The collective potential obtained above and the mass parameter obtained using Eq. (15) are combined to construct the collective Hamiltonian for investigating the simple wobbling motion. Diagonalizing the collective Hamiltonian, the collective energy levels and wave functions at each cranking frequency are yielded. Taking $\hbar \omega = 0.1 \text{ MeV}$ and 0.4 MeV for example, the obtained ten lowest wobbling energy levels and corresponding wave functions are presented in Fig. [3.](#page-5-0) It is obviously seen that the wave functions are symmetric for even-*n* levels and antisymmetric for odd-*n* levels with respect to the $\varphi \rightarrow -\varphi$ transformation. Thus the broken signature symmetry in the TAC model is restored in the collective Hamiltonian by the quantization of wobbling angle φ and the consideration of quantum fluctuation along the *ϕ* motion. In addition, it is also shown that the wave function of the most favored wobbling energy levels are symmetric.

The wobbling frequency $\hbar\Omega_{\rm wob}$ defined as the energy difference between the lowest two levels for a certain rotational frequency in the collective Hamiltonian is shown in Fig. [4](#page-5-0) in comparison with those from the HA formula (14) . It is seen that both collective Hamiltonian and HA give the linear increasing trend of wobbling frequency with respect to rotational frequency. For the HA results, this is just expected since the coefficient $\sqrt{\frac{(\mathcal{J}_1 - \mathcal{J}_2)(\mathcal{J}_1 - \mathcal{J}_3)}{\mathcal{J}_3\mathcal{J}_2}}$ in the HA formula [\(14\)](#page-2-0) is a positive constant value. For the collective Hamiltonian results, this can be also readily understood according to the stiffness of the collective potential, as shown in the upper panels of Figs. $2(a)-2(d)$, which becomes larger with increasing rotational frequency. The wobbling frequency given by the HA formula is a bit larger than that by the collective Hamiltonian results from the fact that the simple harmonic approximation for the collective potential would overestimate the stiffness of the potential, as shown in Fig. [5.](#page-5-0)

3. Comparison with TRM solutions

The simple wobbler solutions discussed here can be exactly obtained by TRM. To study the accuracy of collective Hamiltonian scheme, in Fig. [6,](#page-6-0) the energies of the four lowest wobbling bands $n = 1, 2, 3, 4$ relative to the $n = 0$ yrast sequence obtained by the collective Hamiltonian are displayed in comparison with those from TRM and HA. In the TRM, the states possesses D_2 symmetry so that the spectrum is restricted to the states with $(-1)^n = (-1)^l$ [\[1\]](#page-11-0), i.e., only even spins for even-*n* wobbling bands while only odd appear for odd-*n* wobbling bands. Hence, the wobbling excitation energies are calculated in different ways for even-*n* and odd-*n* wobbling

FIG. 3. (Color online) The collective levels and wave functions obtained from the collective Hamiltonian. Upper panel: The ten lowest energy levels labeled as $n = 0-9$ (left) and the corresponding wave functions for even-*n* (middle) and odd-*n* (right) states at the frequency $\hbar \omega = 0.1$ MeV. Lower panel: The ten lowest energy levels labeled as $n = 0-9$ (left) and the corresponding wave functions for even-*n* (middle) and odd-*n* (right) states at the frequency $\hbar \omega = 0.4 \text{ MeV}$.

bands. For even-*n* wobbling bands, the wobbling energies are directly calculated as the energy difference with respect

FIG. 4. (Color online) The simple wobbling frequency $\hbar\Omega_{\text{wob}}$ obtained by the collective Hamiltonian in comparison with those of the HA formula [\(14\)](#page-2-0).

to $n = 0$ wobbling bands $E_{\text{wob}}^n = E_n(I) - E_0(I)$, while for odd-*n* wobbling bands calculated as the energy difference with respect to the interpolated energies by $n = 0$ wobbling band $E_{\text{wob}}^n(I) = E_n(I) - [E_0(I+1) + E_0(I-1)]/2$. In the TRM, the spin is a good quantum number, while in the collective Hamiltonian not but an expectation value of angular

FIG. 5. (Color online) The collective potentials obtained by HA [\(12\)](#page-2-0) in comparison with those by TAC [\(11\)](#page-2-0) at frequencies $\hbar \omega = 0.1$ and 0.4 MeV.

FIG. 6. (Color online) Energy spectra of four simple wobbling bands $n = 1,2,3,4$ relative to the $n = 0$ yrast sequence obtained by collective Hamiltonian in comparison with TRM and HA. In the TRM, the wobbling energies for even-*n* wobbling bands are calculated as $E_{\text{wob}}^{n} = E_n(I) - E_0(I)$, while for odd-*n* wobbling bands $E_{\text{wob}}^n(I) = E_n(I) - [E_0(I + 1) + E_0(I - 1)]/2.$

momentum operator on the rotational state with given rotational frequency.

As shown in Fig. 6, for each wobbling band, the wobbling energy is found to increase with spin. With an increase in *n*, the HA results gradually deviate from TRM, which indicates that the wobbling motion gradually deviates from the harmonic oscillation character. The collective Hamiltonian excellently reproduces the TRM results even for the large-*n* wobbling bands. The collective Hamiltonian based on TAC approach, however, provides a new perspective to interpret the variation trend of wobbling frequency with spin by exploring the variation trend of stiffness of the collective potential.

B. Longitudinal wobbler

Now we discuss the longitudinal wobbler, where a $h_{11/2}$ proton particle is assumed to couple to a triaxial rotor and its angular momentum is parallel to the axis with the largest moment of inertia. The rigid body type of moment of inertia (21) is used here too.

1. Collective potential

In the contour plots of Figs. $7(a)$ – $7(d)$, the total Routhian surface calculations for longitudinal wobbling motions obtained by TAC are shown at the rotational frequencies $\hbar \omega = 0.1, 0.2, 0.3,$ and 0.4 MeV. Similar to the case of simple wobbling, the total Routhian is also symmetrical with respect to the $\varphi = 0^\circ$ line and the minima always locate at $(\theta = 90^\circ, \varphi = 0^\circ)$ regardless of how fast the nucleus rotates. This is very clear since both the proton particle and triaxial rotor angular momenta in the longitudinal wobbling system are oriented along the short axis, the axis with the largest moment of inertia.

With the total Routhian, the extracted collective potentials $V(\varphi)$ are presented at $\hbar \omega = 0.1, 0.2, 0.3,$ and 0.4 MeV in the upper panels of Figs. $7(a)$ – $7(d)$. It is clearly seen that the collective potentials presented here are very similar to those presented in the upper panels of Figs. $2(a)-2(d)$ for simple

FIG. 7. (Color online) Same as Fig. [2](#page-4-0) but for longitudinal wobbling motion, where a proton $h_{11/2}$ particle coupled to a triaxial rigid body rotor with $\gamma = -30^\circ$.

wobbling motions, while the only difference is that the stiffness here becomes larger. Therefore, similar discussions for simple wobbling motion still hold true here. It is worth stressing that the deeper potentials here are attributed to the proton particle and its contribution would become larger at larger rotational frequency. For example, at $\hbar \omega = 0.1 \text{ MeV}$, the energy difference between $\varphi = \pm 90^\circ$ and $\varphi = 0^\circ$ is ~540 keV and reaches \sim 2080 keV at $\hbar \omega = 0.4$ MeV. Comparing with the simple wobbling motions, one obtains the contribution from proton increases from ~510 keV at $\hbar \omega = 0.1 \text{ MeV}$ to 1480 keV at 0*.*4 MeV.

2. Mass parameter

The mass parameter for longitudinal wobbling motion is calculated by Eq. [\(19\)](#page-3-0) and shown in Fig. 8. As discussed in Sec. [II,](#page-1-0) since the effective moments of inertia for the one axis decreases with the rotational frequency, the mass parameter increases with increasing rotational frequency. This increasing characteristic is different from the simple wobbler, where the mass parameter is constant at any rotational frequency.

FIG. 8. (Color online) The calculated mass parameter as a function of rotational frequency $\hbar \omega$ for longitudinal wobbling motion.

3. Collective levels and wave functions

The obtained collective energy levels and corresponding wave functions are illustrated in Fig. 9 for $\hbar \omega = 0.1$ and 0*.*4 MeV. Again, the wave functions presented here are similar as those presented in Fig. [3](#page-5-0) for simple wobbling motions.

In Fig. 10, the obtained wobbling frequency calculated by the collective Hamiltonian, is in comparison with the results obtained by the HFA approximation (20) . It is found that both the collective Hamiltonian and HFA give the increased wobbling frequency as function of rotational frequency. However, the HFA results are larger than the collective Hamiltonian ones over the whole range of rotational frequency. To understand the origin of the differences between HFA and collective Hamiltonian, the collective potential obtained by the HFA approximation in comparison with the results obtained by TAC at $\hbar \omega = 0.1 \text{ MeV}$ and $\hbar \omega = 0.4 \text{ MeV}$ are shown in Fig. [11.](#page-8-0) It is seen that the stiffness of the collective potential calculated by HFA are larger than the collective Hamiltonian at both $\hbar \omega = 0.1$ and 0.4 MeV. Since the mass parameter in the collective Hamiltonian is the same as that in the HFA, the wobbling frequency of HFA is larger than the collective Hamiltonian. Aside from the harmonic approximation as a

FIG. 10. (Color online) The longitudinal wobbling frequency $\hbar\Omega_{\rm wob}$ obtained by collective Hamiltonian in comparison with those of the HFA approximation [\(20\)](#page-3-0).

simple wobbling motion, the HFA further introduces that the proton particle rigidly aligns its angular momentum along the short axis. Hence it deviates larger from the collective Hamiltonian for the longitudinal wobbling motion (∼20 keV,

FIG. 9. (Color online) Same as Fig. [3](#page-5-0) but for longitudinal wobbling motion.

FIG. 11. (Color online) The collective potentials obtained by HFA (18) in comparison with those by TAC (17) at frequencies $\hbar \omega = 0.1$ and 0.4 MeV.

see Fig. [10\)](#page-7-0) than HA for the simple wobbling motion (∼5 keV, see Fig. 4).

4. Comparison with PRM solutions

The exact solutions for longitudinal wobbling motion can be obtained by PRM. In order to investigate the quality of the collective Hamiltonian, the energies of the two lowest wobbling bands $n = 1,2$ relative to the $n = 0$ yrast sequence obtained by the collective Hamiltonian are shown in Fig. 12 in comparison with those from PRM. In PRM, for even-*n*, the wobbling energies are calculated as $E_{\text{wob}}^n = E_n(I) - E_0(I)$, while for odd-*n* wobbling bands are calculated as $E_{\text{wob}}^n(I) =$ $E_n(I) - [E_0(I + 1) + E_0(I - 1)]/2$. It is found that the collective Hamiltonian can reproduce the PRM very well. With increasing spin, the wobbling energy increases. The results calculated by HFA are also shown in Fig. 12. The wobbling energies given by HFA are larger than those obtained by both PRM and collective Hamiltonian.

Both HFA and collective Hamiltonian are approximate solutions with respect to PRM. In the HFA approximation, the harmonic oscillator potential and the frozen alignment of proton particle are assumed. In the collective Hamiltonian, however, only the mass parameter is calculated with the HFA approximation, while the collective potential is calculated by

FIG. 12. (Color online) Energy spectra of two longitudinal wobbling bands $n = 1,2$ relative to the $n = 0$ yrast sequence obtained by collective Hamiltonian in comparison with PRM and HFA. In the PRM, the wobbling energies for even-*n*, the wobbling energies are calculated as $E_{\text{wob}}^n = E_n(I) - E_0(I)$, while for odd-*n* $E_{\text{wob}}^n(I) =$ $E_n(I) - [E_0(I+1) + E_0(I-1)]/2.$

FIG. 13. (Color online) Same as Fig. [2](#page-4-0) but for transverse wobbling, where a proton $h_{11/2}$ particle coupled to a triaxial irrotational flow rotor with $\gamma = -30^\circ$.

the TAC model without prior assuming the frozen alignment with respect to any axis for proton particle. The PRM exactly diagonalizes the particle rotor coupling Hamiltonian and thus gives the exact solutions. From this point of view, the collective Hamiltonian has improved the descriptions for the collective potential and provides a more accurate solution than HFA.

C. Transverse wobbler

For transverse wobbling motions, the proton particle angular momentum is supposed to be perpendicular to the axis with the largest moment of inertia. In the present investigation, the irrotational flow type of moment of inertia [\(22\)](#page-3-0) is employed to satisfy this requirement.

1. Collective potential

The total Routhian calculated by TAC for a $h_{11/2}$ proton particle coupled to a triaxial irrotational flow rotor with $\gamma = -30^\circ$ in the (θ, φ) plane are displayed at the rotational frequencies $\hbar \omega = 0.1, 0.2, 0.3, 0.4 \,\text{MeV}$ in contour plots of Figs. $13(a)$ – $13(d)$. The potential energy surfaces are also

FIG. 14. (Color online) Same as Fig. [8](#page-6-0) but for transverse wobbling motion.

FIG. 15. (Color online) Same as Fig. [3](#page-5-0) but for transverse wobbling motion.

symmetric with the $\varphi = 0^\circ$ line. In contrast to the simple and longitudinal wobbling motions, the minima in the potential energy surfaces change from $\varphi = 0^\circ$ to $\varphi \neq 0^\circ$ with the increasing frequency. As discussed in Ref. [\[7\]](#page-11-0), this implies the axis of uniform rotation is tilted from *s* axis into the *s*-*i* plane.

The extracted collective potentials $V(\varphi)$ for transverse wobbling motion are shown in the upper panels of Fig. $13(a)$ – [13\(d\).](#page-8-0) For $\hbar \omega = 0.1 \text{ MeV}$, the potential $V(\varphi)$ is a harmonic oscillator type which has only one minimum at $\varphi = 0^\circ$, which corresponds to the uniform rotation around the one axis. For $\hbar \omega \geqslant 0.20 \,\text{MeV}$, the potential $V(\varphi)$ has two symmetrical minima, which correspond to the tilted rotation. Due to the appearance of the potential barrier, the tilted solutions are achieved in the body-fixed frame. The heights of barrier defined as $\Delta V = V(0) - V_{\text{min}}$ (in MeV) with V_{min} being the value of the potential at the minimum presented also in the figure. It is found that the potential barrier increases with the rotational frequency, e.g., from 0.046 MeV at $\hbar \omega =$ 0.20 MeV to 1.145 MeV at $\hbar \omega = 0.40 \text{ MeV}$.

2. Mass parameter

The obtained mass parameter calculated by Eq. [\(19\)](#page-3-0) as a function of rotational frequency is shown in Fig. [14.](#page-8-0) Since here the irrotational flow type of moment of inertia [\(22\)](#page-3-0) with γ = -30° assumed, i.e., $\mathcal{J}_1 = \mathcal{J}_3$, the deduced mass parameter is linearly dependent on rotational frequency. The mass parameter [\(19\)](#page-3-0) is derived based on the assumption of the harmonic frozen alignment approximation, therefore, it is strictly speaking valid only at the wobbling motion region and will becomes invalid in the tilted rotation region. Nevertheless, as a rough approximation, the mass parameter formula (19) is used for the calculations over the whole range of rotational frequency.

3. Collective levels and wave functions

The obtained collective levels and wave functions are shown in Fig. 15 for $\hbar \omega = 0.1$, 0.4 MeV. Similar to the simple and longitudinal wobbling motions, the wave functions are symmetric for even-*n* levels and antisymmetric for odd-*n* levels. For $n = 0$, the peak of the wave function locates

FIG. 16. (Color online) Same as Fig. [10](#page-7-0) but for transverse wobbling motion.

around $\varphi = 0^\circ$ at $\hbar \omega = 0.1 \text{ MeV}$, while moves towards to $\varphi = 90^\circ$ at $\hbar \omega = 0.4 \,\text{MeV}$. In addition, when the rotational frequency increases, the probability distributions determined by the absolute square of wave functions tend to show a similar pattern for $n = 0$ and $n = 1$ levels. This is consistent with their energy differences, as shown in the left panel of Fig. [15,](#page-9-0) tending to zero.

The calculated wobbling frequencies are shown in Fig. 16. It can be seen from Fig. 16 that the wobbling frequency decreases with the rotational frequency. This decreasing is attributed to the increase of the potential barrier, as shown in the upper panels of Fig. [13,](#page-8-0) which will suppress the tunneling probability between the two symmetrical TAC solutions. At $\hbar \omega \geq 0.3$ MeV, the wobbling frequency tends to zero, which implies the transverse wobbling motion is terminated. For comparison, the wobbling frequencies calculated by HFA are also shown in Fig. 16. The decreasing trend is clearly observed. As discussed above, at the critical rotational frequency $\hbar \omega_c = j_\pi / (\mathcal{J}_2 - \mathcal{J}_1) \approx 0.183 \text{ MeV}$ the wobbling frequency becomes zero and above it the HFA formula becomes invalid. Comparing with the collective Hamiltonian, the HFA gives about 50 keV smaller values of wobbling frequency. It is also worthy to mention that the HFA gives a more rapid decreasing trend than the collective Hamiltonian since the quantum fluctuations are not taken into account in the HFA beyond the region of transverse wobbling motion.

FIG. 17. (Color online) Same as Fig. [12](#page-8-0) but for transverse wobbling motion.

4. Comparison with PRM solutions

In Fig. 17, the energies of the two lowest wobbling bands $n = 1,2$ relative to the $n = 0$ yrast sequence obtained by collective Hamiltonian are shown in comparison with the PRM solutions and HFA results. It is found that the collective Hamiltonian can reproduce the PRM results well at the region of wobbling motions. For $I \geq 16.5\hbar$, the wobbling energies of $n = 1$ increase in the PRM, which indicates the onset of transitions from the transverse to longitudinal wobbling motions as discussed in Ref. [\[7\]](#page-11-0). In the present collective Hamiltonian, however, the longitudinal wobbling motion could not be given since the adopted box boundary condition prohibits any of the quantal fluctuation around $\varphi = 90^\circ$. Therefore, the transition is not reproduced. If one wants to reproduce the longitudinal wobbling and to describe the transition from the transverse to the longitudinal wobbler, one should use the periodic boundary condition to replace the box boundary condition or alternatively choose a new collective coordinate, which is defined as the angle between the angular momentum and the two axis. Further investigation on this topic will be done in the future.

V. SUMMARY

In summary, three types of wobbling modes for the nucleus have been studied in the framework of the collective Hamiltonian. The simple wobbler is a pure triaxial rotor assumed with rigid body type of moments of inertia. With an odd $h_{11/2}$ proton of particle character coupling to the triaxial rotor, the longitudinal wobbler is achieved by arranging the moments of inertia as rigid body type, while the transverse wobbler achieved as irrotational body type. The collective potentials in the collective Hamiltonian are calculated based on the TAC approach. The mass parameters are obtained by HA for simple wobbling motion, while by the HFA approximation for longitudinal and transverse wobbling motions.

Diagonalizing the collective Hamiltonian, the energies and the wave functions of the wobbling states are yielded. The obtained wobbling energies of the simple wobbler are compared with the results calculated by HA and TRM, while those of longitudinal and transverse wobblers energies are compared with HFA and PRM. It is found that the results of the collective Hamiltonian are in good agreement with those exact solutions by TRM or PRM.

In accordance with those obtained by the HA or HFA formula [\[7\]](#page-11-0), and RPA model [\[36\]](#page-11-0), it is observed that the wobbling frequency increases with the rotational frequency for the simple and longitudinal wobbling motions, while decreases for the transverse wobbling motion. It is presented here that these variation trends of the wobbling frequency are in association with the stiffness of the collective potentials. It should be mentioned that the present work has provided a new way to understand the wobbling phenomena, which in particular may further contribute to the investigation of nuclear wobbling based on a realistic TAC theory such as tilted axis cranking density functional theory [\[37\]](#page-11-0). It should also be mentioned that only the φ is considered as the collective coordinate in the present collective Hamiltonian. To further quantitatively consider the fluctuations in both the θ and φ directions is a very interesting work. We are going to do this work in the near future, in which a twodimensional collective Hamiltonian including the dynamic motions of both φ and θ and their couplings would be constructed.

- [1] A. Bohr and B. R. Mottelson, *Nuclear Structure*, Vol. II (Benjamin, New York, 1975).
- [2] R. Bengtsson, H. Frisk, F. May, and J. Pinston, [Nucl. Phys. A](http://dx.doi.org/10.1016/0375-9474(84)90620-1) **[415](http://dx.doi.org/10.1016/0375-9474(84)90620-1)**, [189](http://dx.doi.org/10.1016/0375-9474(84)90620-1) [\(1984\)](http://dx.doi.org/10.1016/0375-9474(84)90620-1).
- [3] I. Hamamoto and H. Sagawa, [Phys. Lett. B](http://dx.doi.org/10.1016/0370-2693(88)90593-X) **[201](http://dx.doi.org/10.1016/0370-2693(88)90593-X)**, [415](http://dx.doi.org/10.1016/0370-2693(88)90593-X) [\(1988\)](http://dx.doi.org/10.1016/0370-2693(88)90593-X).
- [4] S. Frauendorf and J. Meng, [Nucl. Phys. A](http://dx.doi.org/10.1016/S0375-9474(97)00004-3) **[617](http://dx.doi.org/10.1016/S0375-9474(97)00004-3)**, [131](http://dx.doi.org/10.1016/S0375-9474(97)00004-3) [\(1997\)](http://dx.doi.org/10.1016/S0375-9474(97)00004-3).
- [5] S. Frauendorf, [Rev. Mod. Phys.](http://dx.doi.org/10.1103/RevModPhys.73.463) **[73](http://dx.doi.org/10.1103/RevModPhys.73.463)**, [463](http://dx.doi.org/10.1103/RevModPhys.73.463) [\(2001\)](http://dx.doi.org/10.1103/RevModPhys.73.463).
- [6] J. Meng and S. Q. Zhang, [J. Phys. G: Nucl. Part. Phys.](http://dx.doi.org/10.1088/0954-3899/37/6/064025) **[37](http://dx.doi.org/10.1088/0954-3899/37/6/064025)**, [064025](http://dx.doi.org/10.1088/0954-3899/37/6/064025) [\(2010\)](http://dx.doi.org/10.1088/0954-3899/37/6/064025).
- [7] S. Frauendorf and F. Dönau, *[Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.89.014322)* [89](http://dx.doi.org/10.1103/PhysRevC.89.014322), [014322](http://dx.doi.org/10.1103/PhysRevC.89.014322) [\(2014\)](http://dx.doi.org/10.1103/PhysRevC.89.014322).
- [8] S. W. Ødegård, G. B. Hagemann, D. R. Jensen, M. Bergström, B. Herskind, G. Sletten, S. Törmänen, J. N. Wilson, P. O. Tjøm, I. Hamamoto *et al.*, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.86.5866) **[86](http://dx.doi.org/10.1103/PhysRevLett.86.5866)**, [5866](http://dx.doi.org/10.1103/PhysRevLett.86.5866) [\(2001\)](http://dx.doi.org/10.1103/PhysRevLett.86.5866).
- [9] D. R. Jensen, G. B. Hagemann, I. Hamamoto, S. W. Ødegård, B. Herskind, G. Sletten, J. N. Wilson, K. Spohr, H. Hübel, P. Bringel *et al.*, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.89.142503) **[89](http://dx.doi.org/10.1103/PhysRevLett.89.142503)**, [142503](http://dx.doi.org/10.1103/PhysRevLett.89.142503) [\(2002\)](http://dx.doi.org/10.1103/PhysRevLett.89.142503).
- [10] D. R. Jensen, G. B. Hagemann, I. Hamamoto, S. W. Ødegård, M. Bergstrom, B. Herskind, G. Sletten, S. Tormanen, J. N. Wilson, P. O. Tjom *et al.*, [Nucl. Phys. A](http://dx.doi.org/10.1016/S0375-9474(01)01459-2) **[703](http://dx.doi.org/10.1016/S0375-9474(01)01459-2)**, [3](http://dx.doi.org/10.1016/S0375-9474(01)01459-2) [\(2002\)](http://dx.doi.org/10.1016/S0375-9474(01)01459-2).
- [11] G. Schönwaßer, H. Hübel, G. B. Hagemann, P. Bednarczyk, G. Benzoni, A. Bracco, P. Bringel, R. Chapman, D. Curien, J. Domscheit *et al.*, [Phys. Lett. B](http://dx.doi.org/10.1016/S0370-2693(02)03095-2) **[552](http://dx.doi.org/10.1016/S0370-2693(02)03095-2)**, [9](http://dx.doi.org/10.1016/S0370-2693(02)03095-2) [\(2003\)](http://dx.doi.org/10.1016/S0370-2693(02)03095-2).
- [12] P. Bringel, G. Hagemann, H. Hübel, A. Al-khatib, P. Bednarczyk, A. Bürger, D. Curien, G. Gangopadhyay, B. Herskind, D. Jensen *et al.*, [Eur. Phys. J. A](http://dx.doi.org/10.1140/epja/i2005-10005-7) **[24](http://dx.doi.org/10.1140/epja/i2005-10005-7)**, [167](http://dx.doi.org/10.1140/epja/i2005-10005-7) [\(2005\)](http://dx.doi.org/10.1140/epja/i2005-10005-7).
- [13] H. Amro, W. C. Ma, G. B. Hagemann, R. M. Diamond, J. Domscheit, P. Fallon, A. Gorgen, B. Herskind, H. Hubel, D. R. Jensen *et al.*, [Phys. Lett. B](http://dx.doi.org/10.1016/S0370-2693(02)03199-4) **[553](http://dx.doi.org/10.1016/S0370-2693(02)03199-4)**, [197](http://dx.doi.org/10.1016/S0370-2693(02)03199-4) [\(2003\)](http://dx.doi.org/10.1016/S0370-2693(02)03199-4).
- [14] G. B. Hagemann, [Eur. Phys. J. A](http://dx.doi.org/10.1140/epja/i2002-10349-4) **[20](http://dx.doi.org/10.1140/epja/i2002-10349-4)**, [183](http://dx.doi.org/10.1140/epja/i2002-10349-4) [\(2004\)](http://dx.doi.org/10.1140/epja/i2002-10349-4).
- [15] D. J. Hartley, R. V. F. Janssens, L. L. Riedinger, M. A. Riley, A. Aguilar, M. P. Carpenter, C. J. Chiara, P. Chowdhury, I. G. Darby, U. Garg *et al.*, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.80.041304) **[80](http://dx.doi.org/10.1103/PhysRevC.80.041304)**, [041304\(R\)](http://dx.doi.org/10.1103/PhysRevC.80.041304) [\(2009\)](http://dx.doi.org/10.1103/PhysRevC.80.041304).
- [16] S. J. Zhu, Y. X. Luo, J. H. Hamilton, A. V. Ramayya, X. L. Che, Z. Jiang, J. K. Hwang, J. L. Wood, M. A. Stoyer, R. Donangelo *et al.*, [Int. J. Mod. Phys. E](http://dx.doi.org/10.1142/S0218301309013798) **[18](http://dx.doi.org/10.1142/S0218301309013798)**, [1717](http://dx.doi.org/10.1142/S0218301309013798) [\(2009\)](http://dx.doi.org/10.1142/S0218301309013798).
- [17] I. Hamamoto, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.65.044305) **[65](http://dx.doi.org/10.1103/PhysRevC.65.044305)**, [044305](http://dx.doi.org/10.1103/PhysRevC.65.044305) [\(2002\)](http://dx.doi.org/10.1103/PhysRevC.65.044305).

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- [18] I. Hamamoto and B. R. Mottelson, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.68.034312) **[68](http://dx.doi.org/10.1103/PhysRevC.68.034312)**, [034312](http://dx.doi.org/10.1103/PhysRevC.68.034312) [\(2003\)](http://dx.doi.org/10.1103/PhysRevC.68.034312).
- [19] K. Tanabe and K. Sugawara-Tanabe, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.73.034305) **[73](http://dx.doi.org/10.1103/PhysRevC.73.034305)**, [034305](http://dx.doi.org/10.1103/PhysRevC.73.034305) [\(2006\)](http://dx.doi.org/10.1103/PhysRevC.73.034305).
- [20] K. Tanabe and K. Sugawara-Tanabe, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.77.064318) **[77](http://dx.doi.org/10.1103/PhysRevC.77.064318)**, [064318](http://dx.doi.org/10.1103/PhysRevC.77.064318) [\(2008\)](http://dx.doi.org/10.1103/PhysRevC.77.064318).
- [21] Y. R. Shimizu and M. Matsuzaki, [Nucl. Phys. A](http://dx.doi.org/10.1016/0375-9474(94)00823-6) **[588](http://dx.doi.org/10.1016/0375-9474(94)00823-6)**, [559](http://dx.doi.org/10.1016/0375-9474(94)00823-6) [\(1995\)](http://dx.doi.org/10.1016/0375-9474(94)00823-6).
- [22] [M. Matsuzaki, Y. R. Shimizu, and K. Matsuyanagi,](http://dx.doi.org/10.1103/PhysRevC.65.041303) Phys. Rev. C **[65](http://dx.doi.org/10.1103/PhysRevC.65.041303)**, [041303\(R\)](http://dx.doi.org/10.1103/PhysRevC.65.041303) [\(2002\)](http://dx.doi.org/10.1103/PhysRevC.65.041303).
- [23] [M. Matsuzaki, Y. R. Shimizu, and K. Matsuyanagi,](http://dx.doi.org/10.1140/epja/i2002-10350-y) Eur. Phys. J. A **[20](http://dx.doi.org/10.1140/epja/i2002-10350-y)**, [189](http://dx.doi.org/10.1140/epja/i2002-10350-y) [\(2003\)](http://dx.doi.org/10.1140/epja/i2002-10350-y).
- [24] M. Matsuzaki and S.-I. Ohtsubo, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.69.064317) **[69](http://dx.doi.org/10.1103/PhysRevC.69.064317)**, [064317](http://dx.doi.org/10.1103/PhysRevC.69.064317) [\(2004\)](http://dx.doi.org/10.1103/PhysRevC.69.064317).
- [25] [M. Matsuzaki, Y. R. Shimizu, and K. Matsuyanagi,](http://dx.doi.org/10.1103/PhysRevC.69.034325) *Phys. Rev.* C **[69](http://dx.doi.org/10.1103/PhysRevC.69.034325)**, [034325](http://dx.doi.org/10.1103/PhysRevC.69.034325) [\(2004\)](http://dx.doi.org/10.1103/PhysRevC.69.034325).
- [26] [Y. R. Shimizu, M. Matsuzaki, and K. Matsuyanagi,](http://dx.doi.org/10.1103/PhysRevC.72.014306) Phys. Rev. C **[72](http://dx.doi.org/10.1103/PhysRevC.72.014306)**, [014306](http://dx.doi.org/10.1103/PhysRevC.72.014306) [\(2005\)](http://dx.doi.org/10.1103/PhysRevC.72.014306).
- [27] Y. R. Shimizu, T. Shoji, and M. Matsuzaki, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.77.024319) **[77](http://dx.doi.org/10.1103/PhysRevC.77.024319)**, [024319](http://dx.doi.org/10.1103/PhysRevC.77.024319) [\(2008\)](http://dx.doi.org/10.1103/PhysRevC.77.024319).
- [28] T. Shoji and Y. R. Shimizu, [Progr. Theor. Phys.](http://dx.doi.org/10.1143/PTP.121.319) **[121](http://dx.doi.org/10.1143/PTP.121.319)**, [319](http://dx.doi.org/10.1143/PTP.121.319) [\(2009\)](http://dx.doi.org/10.1143/PTP.121.319).
- [29] M. Oi, A. Ansari, T. Horibata, and N. Onishi, [Phys. Lett. B](http://dx.doi.org/10.1016/S0370-2693(00)00364-6) **[480](http://dx.doi.org/10.1016/S0370-2693(00)00364-6)**, [53](http://dx.doi.org/10.1016/S0370-2693(00)00364-6) [\(2000\)](http://dx.doi.org/10.1016/S0370-2693(00)00364-6).
- [30] Q. B. Chen, S. Q. Zhang, P. W. Zhao, R. V. Jolos, and J. Meng, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.87.024314) **[87](http://dx.doi.org/10.1103/PhysRevC.87.024314)**, [024314](http://dx.doi.org/10.1103/PhysRevC.87.024314) [\(2013\)](http://dx.doi.org/10.1103/PhysRevC.87.024314).
- [31] S. Mukhopadhyay, D. Almehed, U. Garg, S. Frauendorf, T. Li, P. V. M. Rao, X. Wang, S. S. Ghugre, M. P. Carpenter, S. Gros *et al.*, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.99.172501) **[99](http://dx.doi.org/10.1103/PhysRevLett.99.172501)**, [172501](http://dx.doi.org/10.1103/PhysRevLett.99.172501) [\(2007\)](http://dx.doi.org/10.1103/PhysRevLett.99.172501).
- [32] B. Qi, S. Q. Zhang, J. Meng, S. Y. Wang, and S. Frauendorf, [Phys. Lett. B](http://dx.doi.org/10.1016/j.physletb.2009.02.061) **[675](http://dx.doi.org/10.1016/j.physletb.2009.02.061)**, [175](http://dx.doi.org/10.1016/j.physletb.2009.02.061) [\(2009\)](http://dx.doi.org/10.1016/j.physletb.2009.02.061).
- [33] M. Oi, [Phys. Lett. B](http://dx.doi.org/10.1016/j.physletb.2005.12.061) **[634](http://dx.doi.org/10.1016/j.physletb.2005.12.061)**, [30](http://dx.doi.org/10.1016/j.physletb.2005.12.061) [\(2006\)](http://dx.doi.org/10.1016/j.physletb.2005.12.061).
- [34] W. Pauli, *Handbuch der Physik* (Springer Verlag, Berlin, 1933), Vol. XXIV, p. 120.
- [35] P. Ring and P. Schuck, *The Nuclear Many Body Problem* (Springer Verlag, Berlin, 1980).
- [36] Y. R. Shimizu, M. M. Matsuzaki, and K. Matsuyanagi, [arXiv:nucl-th/0404063v1.](http://arxiv.org/abs/arXiv:nucl-th/0404063v1)
- [37] J. Meng, J. Peng, S. Q. Zhang, and P. W. Zhao, [Front. Phys.](http://dx.doi.org/10.1007/s11467-013-0287-y) **[8](http://dx.doi.org/10.1007/s11467-013-0287-y)**, [55](http://dx.doi.org/10.1007/s11467-013-0287-y) [\(2013\)](http://dx.doi.org/10.1007/s11467-013-0287-y).