Decays of 21 Na, 23 Ne, 23 Mg, 24 Ne, and 31 S[†]

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The radioactive decays of ²¹Na, ²³Ne, ²³Mg, ²⁴Ne, and ³¹S have been studied using Ge (Li) and plastic scintillation detectors to measure γ and β rays. ²¹Na, ²³Ne, and ²⁴Ne were formed by bombarding a neon target with 3-MeV deuterons or tritons, followed by gas transfer to a remote counting station. ²³Mg and ³¹S were made in (p,n) reactions using 10-MeV protons from an MP tandem Van de Graaff together with a rabbit target-transfer facility. The following half-lives (in sec) were obtained: ²¹Na, 22.55±0.10; ²³Ne, 37.24±0.12; ²³Mg, 11.36±0.04; and ³¹S, 2.605±0.012. Relative β -ray intensities were derived from the measured γ -ray intensities. The various branching results are in reasonable agreement with previous work except for ²¹Na whose measured β^+ branch of 5.1±0.2% to the 350-keV state of ²¹Ne is a factor of 2 larger than earlier reports. However, the result is in excellent agreement with an estimate by Wilkinson.

 $\begin{bmatrix} \text{RADIOACTIVITY} & {}^{21}\text{Na}, & {}^{23}\text{Ne}, & {}^{23}\text{Mg}, & {}^{24}\text{Ne}, & {}^{31}\text{S}; \text{ measured } T_{1/2}, & I_{\gamma}; \text{ deduced } \beta \\ & \text{branches, } \log ft. \end{bmatrix}$

INTRODUCTION

During the past few years there has been a considerable amount of renewed interest in the theory of nuclear β decay. Various aspects of the β interaction have been examined by studying the systematic behavior of related cases such as the $0^+ \rightarrow 0^+$ superallowed Fermi β decays,¹ or the various examples of asymmetry in the ft values of Gamow-Teller β decay.² A common difficulty in all such studies based on systematics has been the frequent unreliability of experimentally determined halflives and β -ray branching ratios which has caused some glaring and perplexing discrepancies. Renewed efforts have been made recently to recheck the "literature" values given for these β -decay properties. Greater consistency among the various laboratories doing such work is now being obtained, and many of the discrepancies in the systematics are starting to be cleared up. This is probably the combined result of using more modern equipment and techniques, greater care in taking data, and a better understanding of the problems of data analysis.

The present work was begun partly because of an outstanding discrepancy between the measured β^+ branching ratio of ²¹Na to the 350-keV first excited state of ²¹Ne and the value predicted by Wilkinson³ from recent studies of the systematics of allowed Gamow-Teller β decay compared with strengths computed from shell-model wave functions.⁴ Two measurements of this ²¹Na branch have been reported, one of $2.2 \pm 0.3\%$ by Talbert and Stewart⁵ and the other of $2.3 \pm 0.2\%$ by Arnell and Wernbom.⁶ The mean of these results, 2.27 $\pm 0.20\%$, compares with a branch of $5 \pm 1\%$ predicted by Wilkinson³ on the basis of theoretical wave functions,⁷ where the error in the prediction comes from the width of the observed distribution of the ratio between theoretical and experimental *ft* values.⁴ Such a large discrepancy would be very difficult to accept.

As shown below a considerably larger β_{350}^+ branch has been found here for ²¹Na agreeing with the Wilkinson estimate. Since it was thought that contaminant positron activities might have affected the previous derivations of the β^+ branch it is possible that the earlier half-life values might also have been affected. It was therefore decided to recheck the half-life of ²¹Na.

²³Ne, which is produced along with ²¹Na in the deuteron bombardment of neon, was also studied. Aside from the need to recheck its half-life there were separate reasons for measuring the relative intensities of the ²³Ne γ rays. In some previous work8 on short-lived activities formed in heavyion reactions ²³Ne and ²⁰F were produced simultaneously. The analysis of the β -ray spectrum in coincidence with the 1633-keV 20 F γ rays depended on knowing the contribution due to the 1636-keV γ rays of ²³Ne. Since the 1636-keV peak due to ²³Ne was not resolved from the ²⁰F peak, it was necessary to use the 1636/440 γ -ray intensity ratio of ²³Ne from the literature, together with the observed intensity of the 440-keV peak and the detector efficiency function, to determine the ²³Ne contribution to the 1633-1636-keV doublet. It was therefore felt that a recheck of the ²³Ne γ -

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ray intensities, particularly those of 440 and 1636 keV, would be desirable.

The other activities studied in the present work included ²⁴Ne, ²³Mg, and ³¹S. The ²⁴Ne investigation was made because of the availability of the gas transfer system used for the work on ²¹Na and ²³Ne, and the desire to recheck the γ -ray intensity ratio in ²⁴Ne decay. On the other hand, relative γ -ray intensities in the decays of both ²³Mg and ³¹S had been reported by Talbert and Stewart⁵ in the same paper giving their result on the ²¹Na γ -ray intensities.

EXPERIMENTAL METHODS

For some earlier work⁹ on the decay of ²⁵Ne at the Brookhaven National Laboratory tandem Van de Graaff facility a gas target and transfer system was developed consisting of a bombardment cell, a transfer line including a liquid nitrogen trap, a remote counting cell, and appropriate valves operated by an automatic timer. This system is ideally suited to the study of a radioactive noble gas such as ²³Ne and was therefore adapted for use at the 3.5-MeV Van de Graaff. The cell was periodically filled with ~0.4 atm of neon and bombarded through a 0.0003-cm-thick Ni entrance window for 10 sec with a $0.3-\mu A$ beam of 3.0-MeVdeuterons in order to make the activity via the ²²Ne $(d, p)^{23}$ Ne reaction. After cutting off the beam the gas was expanded through the trap to the counting cell located in the control room and data were recorded for a specified period.

For the work on ²³Ne tests were made to find the most effective means of trapping out the principal contaminant activity ²¹Na which is formed in the ²⁰Ne $(d, n)^{21}$ Na reaction. This was done by measuring in a Ge(Li) detector the intensity of the 511-keV annihilation radiation line, due to the β^+ associated with ²¹Na, relative to the 1636keV γ ray in ²³Ne decay. The most effective trap consisted of a stainless-steel u tube filled with activated charcoal, dampened slightly before assembly by blowing through it, and then immersed in liquid nitrogen. Under these conditions the ratio of activities reaching the counting cell was found, after making all of the necessary corrections to the ratio of $511/1636 \gamma$ -ray peak intensities, to be 21 Na $/{}^{23}$ Ne = 2.4 × 10⁻⁴. Since the isotopic abundance ratio of ²⁰Ne to ²²Ne is ~10 this indicates a discrimination of $\sim 10^5$ against the passage of ²¹Na through the trap, assuming that the cross sections for the (d, p) and (d, n)reactions are comparable. Such a relatively small amount of ²¹Na reaching the counting cell has negligible effects on the γ -ray spectrum or on the measurement of the half-life for β -ray emission

in the decay of ²³Ne.

For the measurements of the γ -ray spectra of ²³Ne and ²⁴Ne, where the latter was formed in the ²²Ne(t, p)²⁴Ne reaction, the counting cell had internal dimensions of 7.5 cm diam by 2 cm high and the top and bottom were made of 1.5-mm-thick Al. This same cell was also used for the ²³Ne half-life measurements by placing a 5-cm-diam by 2.5-cm-high NE102 plastic scintillator next to the cell for detecting the emerging β rays. Following standard procedures¹⁰ the counts were multiscaled after the valves had been closed so as to isolate the ²³Ne activity in the counting cell.

The effectiveness of the charcoal trap used in the ²³Ne and ²⁴Ne work suggested a way of making a source for studies of ²¹Na. This consisted of a small flat cylindrical cell 3 cm in diam by 3 mm high filled with granular activated charcoal. The top and bottom were of 0.5-mm-thick Al and inlet and outlet tubes were attached at opposite edges of the cylinder. After bombardment of the neon target the gas was expanded through the small cell, but in this case the outlet was pumped on continuously for 15 sec in order to remove all of the ²³Ne activity from the system leaving the ²¹Na sticking to the charcoal in the small cell. Of course for these experiments the large charcoal trap between the bombardment and counting cells used in the ²³Ne and ²⁴Ne was removed. Counting was started at the end of the 15-sec pumping period after closing the pump-out valve.

The above arrangement was used both for measuring the half-life of ²¹Na by observing the positrons in the NE102 detector and for determining the relative γ -ray intensities in a Ge(Li) detector.

TABLE I. γ -ray intensities in the decay of ²³Ne normalized to a value of 33.0 for the 440-keV γ ray.

	Intensity						
E_{γ}		Previous					
(keV)	Present	I ^a	пь	III c			
440	33.0	33.0	33.0	33,0			
1636	1.00 ± 0.04	0.93	0.90	1.0			
2076	0.101 ± 0.006	0.10	0.10	0.10			
2542	0.027 ± 0.002	0.026	0.027	0.025			
2982	$\textbf{0.038} \pm \textbf{0.002}$	0.031	0.033	0.045			

^a H. Lancman, A. Jasinski, J. Kownacki, and

J. Ludziejewski, Nucl. Phys. <u>69</u>, 384 (1965). No errors were quoted.

^b Deduced from decay scheme given by R. Moreh, J. Balderman, and Y. Gozez, Nucl. Phys. <u>A107</u>, 236 (1968). Over-all errors on relative γ -ray intensities were not quoted.

^c Deduced from decay scheme given by J. E. Christiansson, J. Dubois, and L. Jarneborn, Phys. Scripta <u>5</u>, 159 (1972). No errors were quoted.

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In the latter case identical absorbers of either 1.3-cm-thick Lucite or 0.63-cm-thick Al were clamped on either side of the cell so as to completely absorb the β^+ particles in a symmetrical geometry that could be corrected for by simple calculations.

²³Mg and ³¹S were both studied at the tandem Van de Graaff facility by making use of the "rabbit" target-transfer system described previously.¹¹ For the ²³Mg experiments the target consisted of a thick layer of Na₂CO₃ and the activity was made in the ²³Na(p, n)²³Mg reaction using a 5-nA beam of 10-MeV protons. ³¹S was also made via the (p, n) reaction by bombarding a thick target of red phosphorus, deposited on a Ta backing, with a 1-nA beam of 10-MeV protons. In both cases the half-lives for β emission were measured with the NE102 scintillator and the γ -ray spectra were studied with a Ge(Li) detector.

RESULTS AND DISCUSSION

²¹Na

Runs were made on the decay of ²¹Na at β detector biases from 0.25 to 1.5 MeV. Following the procedures described previously¹⁰ the adopted half-life of ²¹Na based on the analysis of six runs is 22.55 ± 0.10 sec. This is in agreement with the compilation value¹² of 22.8 ± 0.2 sec. There was no evidence in the γ -ray spectrum for the 1633keV line of ²⁰F which could be produced in the ²²Ne(d, α)²⁰F reaction. Furthermore, the fitting of data over various regions of the decay curves at different biases did not suggest the presence of a ²⁰F component ($T_{1/2} \cong 11$ sec) of sufficient intensity to alter the results on ²¹Na.

The γ -ray spectrum from ²¹Na was recorded with a $70-cm^3$ Ge(Li) detector for which the efficiency function had been accurately established with calibrated sources. This was placed 10 cm from the cell. Net areas under the 350- and 511keV peaks were measured. In addition to correcting for γ -ray efficiency versus energy one must allow for the fact that the source of 511-keV γ rays is extended due to the penetration of the β^+ particles into the absorbers clamped on either side of the cell. Based on the mean range of the β^{*} particles in the absorber the correction to the counting rate of the 511-keV line was calculated to be 0.5% for the source-detector distance used. Corrections were also made for the differences in the attenuation of 350- and 511-keV γ rays in passing through the β^+ absorbers. Summing corrections were shown to be negligible. When all factors were applied (and allowing for the two 511-keV γ rays per decay) the β^+ branching of ²¹Na to the 350-keV state of ²¹Ne was found from

the average results of two runs using Lucite and Al absorbers to be $5.1 \pm 0.2\%$. This is more than a factor of 2 greater than the mean of the two previous measurements, ^{5, 6} 2.27 $\pm 0.20\%$.

Log ft values may be derived for the ²¹Na branches using the results presented above together with the total decay energy based on the known masses of ²¹Na and ²¹Ne. For all such calculations in this paper masses were taken from the 1971 tables of Wapstra and Gove.¹³ The resulting log ft values are 3.61 ± 0.01 for the decay of ²¹Na to the ground state of ²¹Ne and 4.60 ± 0.02 for decay to the 350-keV state.

As pointed out in the Introduction, the present result on the ²¹Na branch to the 350-keV state of ²¹Ne is in excellent agreement with the $5 \pm 1\%$ branch predicted by Wilkinson.³ Recent nuclearstructure calculations on ²¹Ne and ²³Na have been carried out by Gunye¹⁴ in the framework of Hartree-Fock projection formalism. Among his results are the calculated $\log ft$ values for the β decays of ²¹Na and ²³Ne to states of ²¹Ne and ²³Na, respectively. The present experimental $\log ft$ of 4.60 ± 0.02 for the ²¹Na β^+ branch to the first excited state of ²¹Ne is in better agreement with the theoretical value of 4.71 than was the previous experimental $\log ft$ of 5.0. However, it is pointed out by Gunye that his calculations of $\log ft$ values give only semiquantitative agreement with experiment, generally differing from experimental values by 0.2 to 0.3.

²³Ne

Half-life data were taken on ²³Ne decay at β -ray biases from 0.5 to 2 MeV. Based on the analysis of six runs the adopted half-life is 37.24 ± 0.12 sec. This differs from the mean of previous results (37.55 ± 0.10 sec) by somewhat more than the sum of the errors, the earlier measurements being 37.6 ± 0.1 , ¹⁵ 37.5 ± 0.1 , ¹⁶ and 38.0 ± 0.3 sec.¹⁷

 γ -ray intensity data for ²³Ne are presented in Table I where the results are compared with previous measurements. The observed γ -ray peak intensities were corrected for attenuation in a 0.63-mm-thick brass β -ray absorber, for the detector efficiency function and for summing effects. The agreement of the present and previous results is good, although errors were not quoted for the earlier measurements. The decay scheme of ²³Ne therefore remains the same, and comparisons of the experimental and theoretical¹⁴ log ft values are unaffected by the present work.

²³Mg

Based on the analysis of four runs at β detector biases from 1.0 to 2.0 MeV the half-life adopted for ²³Mg is 11.36±0.04 sec. Data were not taken below 1.0-MeV bias because of the presence of a weak long-lived background of low-energy positrons presumably due to ¹³N produced in the ¹³C- $(p, n)^{13}$ N reaction. The present result is substantially different from the compilation value¹² of 12.04±0.09 sec representing the weighted mean of six prior measurements. But in the meantime Goss *et al.*¹⁸ have reported a value of 11.41±0.05 sec for ²³Mg with which the present result is in excellent agreement.

The β^{+} branching ratio of ²³Mg to the 440-keV state of ²³Na was found, as in the case of ²¹Na, by measuring the intensity ratio of the 440- and 511-keV peaks with a calibrated Ge(Li) detector. However, allowance had to be made here for the long-lived 511-keV component. This was done by recording the γ -ray spectrum for 15 sec in one section of the analyzer and again in another section after a delay of 75 sec, a procedure that was repeated often enough to acquire sufficient statistical accuracy. After making a small correction for the decay of ¹³N ($T_{1/2}$ = 9.96 min) this component of the 511-keV peak in the first counting interval was found to be about 10% of the total. Further corrections were made as in the ²¹Na analysis, where in this case β^+ absorbers were clamped around the rabbit line so as to make the absorption as uniform as possible.

The β^+ branching of ²³Mg was found to be 9.1 $\pm 0.4\%$ to the 440-keV ²³Na state. This is in good

agreement both with the $9.1 \pm 0.5\%$ result of Talbert and Stewart⁵ and with a value of $8.6 \pm 0.3\%$ obtained by Gorodetzky *et al.*¹⁹ From all measurements a mean value of $8.9 \pm 0.3\%$ is adopted and may be used to calculate log*ft* values of 3.68 ± 0.01 for decay to the ground state of ²³Na and 4.39 ± 0.02 for decay to the 440-keV first excited state.

²⁴Ne

For the ²⁴Ne measurements the neon target was bombarded with a $0.2-\mu A$ beam of 3.2-MeV tritons. Sources of ⁵⁴Mn and ⁸⁸Y, emitting γ rays of 834.81 and 898.02 keV, respectively, were superposed in order to determine the energy of 874-keV γ ray of ²⁴Ne while at the same time measuring its intensity relative to the 472-keV peak. By applying the efficiency function for the Ge(Li) detector the intensity ratio was found to be $I_{874}/I_{472} = (7.9 \pm 0.2)/$ 100. This differs from a previous result²⁰ of $(8.9 \pm 0.5)/100$ by somewhat more than the sum of the errors. The energy of the 874-keV γ ray was found to be 874.35 ± 0.14 keV, in agreement with the previously reported²⁰ energy of 874.3 ± 0.3 keV.

Based on the present results and on the halflife²¹ of 3.38 ± 0.02 min the log *ft* values are 4.37 ± 0.01 for the decay of ²⁴Ne to the 472-keV first excited state of ²⁴Na and 4.41 ± 0.02 for decay to the 1347-keV level.

	$T_{1/2}$ (in sec, unless otherwise noted)		State of daughter nucleus	β branch (%)					
Nuclide	Present	Previous	(keV)	Present	Previous	Logft			
²¹ Na	22.55 ± 0.10	22.8 $\pm 0.2^{a}$	Ground	94.9 ± 0.2	97.7 ± 0.2^{b}	3.61 ± 0.01 ^c			
²³ Ne	37.24 ± 0.12	37.55 ± 0.10 ^d	350 e	5.1 ±0.2 e	2.27±0.20° e	4.60 ± 0.02 °			
^{23}Mg	11.36 ± 0.04	12.04 ± 0.09^{a}	Ground	90.9 ± 0.4	91.2 $\pm 0.3^{f}$	3.68 ± 0.01 g			
²⁴ Ne	•••	11.41 ± 0.05 ^h 3.38 ± 0.02 ⁱ	440 472	9.1 ± 0.4 92.1 ± 0.2	8.8 $\pm 0.3^{\text{f}}$ 91.1 $\pm 0.5^{\text{j}}$	4.39 ± 0.02 g 4.37 ± 0.01 k			
³¹ S	2.605 ± 0.012	min 2.61 ± 0.03 ^a	1347 Ground 1266	$\begin{array}{r} 7.9 \pm 0.2 \\ 98.75 \pm 0.06 \\ 1.25 \pm 0.06 \end{array}$	$\begin{array}{r} 8.9 \pm 0.5^{\text{J}} \\ 98.9 \pm 0.1^{\text{a}} \\ 1.1 \pm 0.1^{\text{a}} \end{array}$	4.41 ± 0.02 k 3.71 ± 0.01 c 4.94 ± 0.02 c			

TABLE II. Summary of results.

^a Reference 12.

^b Mean value based on Refs. 5 and 6.

^c Based on present measurements only.

^d Mean value based on Refs. 15, 16, and 17.

^e See Table I for relative γ -ray intensities.

^f Mean value based on Refs. 5 and 19.

^g Based on present half-life and on mean of present and previous results on β branching.

^h Reference 18.

ⁱ Reference 21.

^j Reference 20.

^k Based on present β branches and previous half-life.

Runs on the decay of ³¹S were made at β detector biases from 0.5 to 1.7 MeV. Based on the analysis of six runs the adopted half-life is 2.605 \pm 0.012 sec. This compares with the compilation value¹² of 2.61 \pm 0.03 sec.

The γ -ray spectrum of ³¹S was recorded with the efficiency-calibrated Ge(Li) detector and the intensities of the 511- and 1266-keV peaks were measured. After making the necessary corrections for γ -ray efficiency and absorption the β^+ branching of ³¹S to the 1266-keV state of ³¹P was found to be $1.25 \pm 0.06\%$. This is in agreement with the branch of $1.1 \pm 0.1\%$ given previously.¹² In separate runs an energy calibration source of ²²Na ($E_{\gamma} = 1274.5$ keV) was superposed and the energy of the ³¹S γ ray was found to be 1266.2±0.2 keV, in excellent agreement with a value of 1266.13 ±0.12 keV due to Wolff, Meyer, and Endt.²²

Log ft values for the 31 S branches based on the present work are 3.71 ± 0.01 for decay to the ground state of 31 P and 4.94 ± 0.02 for decay to the 1266-keV first excited state.

Additional very weak γ rays of 2240, 3134, and 3506 keV have been reported²³ as occurring in the decay of ³¹S. No attempt was made to confirm these lines.

A summary of the various results described in this paper is presented in Table II.

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