## Dilepton emission in high-energy heavy-ion collisions with viscous hydrodynamics

Gojko Vujanovic,<sup>1</sup> Clint Young,<sup>2</sup> Björn Schenke,<sup>3</sup> Ralf Rapp,<sup>4</sup> Sangyong Jeon,<sup>1</sup> and Charles Gale<sup>1</sup>

<sup>1</sup>Department of Physics, McGill University, 3600 University Street, Montreal, Quebec H3A 2T8, Canada

<sup>2</sup>School of Physics and Astronomy, University of Minnesota, Minneapolis, Minnesota 55455, USA

<sup>3</sup>Physics Department, Building 510A, Brookhaven National Laboratory, Upton, New York 11973, USA

<sup>4</sup>Cyclotron Institute and Department of Physics and Astronomy, Texas A&M University, College Station, Texas 77843-3366, USA

(Received 20 December 2013; published 10 March 2014)

The invariant mass spectrum and the elliptic flow of lepton pairs produced in relativistic heavy-ion collisions at RHIC are studied with viscous hydrodynamics. The effects of viscous corrections on dilepton observables are explored. The lepton pairs originating from charm quarks evolving in the viscous background are seen to be a good probe of quark energy loss and gain, as quantified by the dilepton spectrum and by the dilepton elliptic flow.

DOI: 10.1103/PhysRevC.89.034904

PACS number(s): 25.75.-q, 24.10.Jv, 24.10.Nz, 47.75.+f

## I. INTRODUCTION

The collision of large nuclei at relativistic energies constitutes the only practical way to heat and compress nuclear matter in the laboratory. Therefore, a vibrant experimental program that aims to elucidate the bulk properties of hot and dense strongly interacting matter is being pursed at several accelerator facilities around the world, notably at the Relativistic Heavy Ion Collider (RHIC, at Brookhaven National Laboratory) and at the Large Hadron Collider (LHC, at CERN). This enterprise has several ambitious goals, but two of them are to map out of the phase diagram of quantum chromodynamics (QCD, the theory of the strong interaction) and to determine the transport coefficients of QCD, which regulate departures from equilibrium. From the volume of data acquired over the past 30 years, together with the theoretical approaches used to interpret them, a "standard picture" is emerging: The nuclear initial states interact strongly, leading to rapid apparent thermalization, followed by a period of quasi-ideal hydrodynamic evolution, which lasts until kinetic freeze-out. The details of this scenario may differ according to approaches-and such details will have empirical consequences-but it is fair to write that the global picture currently appears robust. The endurance of this conceptual framework owes much to the modeling success of relativistic hydrodynamics [1].

Much of the data measured at RHIC and at the LHC consist of hadronic particles which mostly reflect the final stages of the interacting system. Electromagnetic probes, on the other hand, have the potential to provide an unambiguous measurement of the interior dynamics in heavy-ion collisions, owing to the facts that their interaction is dictated by  $\alpha_{\rm EM}$ (the electromagnetic fine-structure constant) and that  $\alpha_{\rm EM} \ll$  $\alpha_s$  (where  $\alpha_s = g^2/4\pi$ , with g being the strong interaction coupling constant). The photons-real and virtual-produced in the hot and dense medium are thus penetrating probes essentially impervious to final-state interactions and as such can reveal details of the underlying particle distributions, including the degree of departure from thermal equilibrium. It is important, however, that precise measurements be accompanied by equally precise modeling. Indeed, the past few years have seen much progress in the simulation of heavy-ion collisions with relativistic hydrodynamics. More specifically, three- plus one-dimensional (3 + 1D) relativistic viscous hydrodynamics models are available and have been used to characterize the matter formed in the relativistic collisions of large nuclei. The current standard input to those simulation tools are hydrodynamic equations derived up to second order in flow velocity gradients for conformal and nonconformal fluids [2,3]. This degree of theoretical and modeling sophistication, together with progress in the quantitative description of the initial states, has brought closer one of the goals described earlier: the extraction of a shear viscosity coefficient for QCD [4].

Going back to electromagnetic radiation, the real photon spectrum has also shown sensitivity to a finite shear viscosity coefficient, as well as to the morphology of the initial states [5–7]. They are therefore observables capable of carrying information from the earliest moments of the collision. The drawback is that, as photons and leptons are emitted throughout the space-time history of the nuclear collision, precise knowledge of the emission rates and the spatial and temporal evolution is required to interpret the measured signal. Dileptons offer the same penetrating advantages as real photons, but their production is suppressed by an extra factor of  $\alpha_{\rm EM}$ . However, lepton pairs have an additional degree of freedom, as the pair's invariant mass and three-momentum are independent.

The goal of this article is to explore the sensitivity of the virtual photon spectrum—through its conversion to lepton pairs—to a nonzero value of the shear viscosity to entropy density ratio,  $\eta/s$ . As is the case for photons, a good knowledge of the different sources of dileptons is required in order to extract meaningful quantitative information. We concentrate on dileptons produced in nuclear collisions at full RHIC energy,  $\sqrt{s} = 200$  A GeV. There, dileptons from high-temperature QCD processes, dileptons from the hot hadronic medium, and dileptons from charm decays are each expected to dominate in different ranges of invariant mass. We include these sources and study the effect of viscosity on the final lepton pair spectrum, which includes the coefficient of elliptic flow,  $v_2$ .

In Sec. II, we first discuss a formulation of relativistic hydrodynamics that incorporates a nonzero coefficient of

shear viscosity, and we highlight the differences between the inviscid and viscous evolutions. Then, the derivation of ideal rates and their viscous corrections for both the quark-gluon plasma phase and the hadronic phase are summarized. The contribution to the dilepton yields in the intermediate mass range from charm decays and its modification in a heavy-ion collision are discussed in Sec. III. The dilepton production in the low- and intermediate-mass regions is discussed in detail in Sec. IV, with emphasis on how viscous corrections change observables. After a complete hydrodynamic simulation of the Au-Au collisions, we compare our yields against recent experimental data from the STAR Collaboration at RHIC. Finally, we conclude in Sec. V.

#### II. IDEAL THERMAL RATES AND THEIR VISCOUS CORRECTIONS

#### A. Viscous relativistic fluid dynamics

Before deriving the viscous correction to the emission rates of electromagnetic radiation, it is appropriate to summarize the effect of shear viscosity on the bulk dynamics. In this paper, we consider no other transport coefficients. As reviews can be found in the recent literature [1,8,9], this summary may be brief. In the conformal Israel-Stewart formalism [10], the stress-energy tensor is usually expressed as

$$T^{\mu\nu} = T^{\mu\nu}_{\text{ideal}} + \pi^{\mu\nu}, \qquad (1)$$

where  $T_{\text{ideal}}^{\mu\nu} = (\epsilon + \mathcal{P})u^{\mu}u^{\nu} - \mathcal{P}g^{\mu\nu}$  is the part of the stressenergy tensor that is unaffected by viscous corrections. The metric tensor is  $g_{\mu\nu} = g^{\mu\nu} = \text{diag}(1, -1, -1, -1)$  and the flow velocity is  $u^{\mu} = (\gamma, \gamma \mathbf{v})$ , which reduces to  $u^{\mu} = (1, 0, 0, 0)$ in the fluid rest frame. The quantities  $\epsilon$  and  $\mathcal{P}$  are the energy density and the pressure, respectively. To second order in flow velocity gradients, the equations that dictate the hydrodynamic evolution are

$$\partial_{\mu}T^{\mu\nu} = 0,$$
  
$$\Delta^{\mu}_{\alpha}\Delta^{\nu}_{\beta}u^{\sigma}\partial_{\sigma}\pi^{\alpha\beta} = -\frac{1}{\tau_{\pi}}(\pi^{\mu\nu} - S^{\mu\nu}) - \frac{4}{3}\pi^{\mu\nu}(\partial_{\alpha}u^{\alpha}).$$
(2)

The viscous part of the stress-energy tensor to first order in flow velocity gradients (the Navier-Stokes limit) is  $S^{\mu\nu} = \eta(\nabla^{\mu}u^{\nu} + \nabla^{\nu}u^{\mu} - \frac{2}{3}\Delta^{\mu\nu}\nabla_{\alpha}u^{\alpha})$  and  $\tau_{\pi}$  is the shear relaxation time. The local three-metric and spatial derivative are  $\Delta^{\mu\nu} = g^{\mu\nu} - u^{\mu}u^{\nu}$  and  $\nabla^{\mu} = \Delta^{\mu\nu}\partial_{\nu}$ , respectively, and  $\eta$  is the coefficient of shear viscosity.

The dynamics of relativistic heavy-ion collisions make it especially advantageous to work in so-called hyperbolic coordinates, such that the coordinate transformation is  $x^{\mu} = (t,x,y,z) \rightarrow (\tau,x,y,\eta_s)$ , with  $\tau = \sqrt{t^2 - z^2}$  and  $\eta_s = (1/2) \ln[(t+z)/(t-z)]$ : the space-time rapidity. In addition, it is straightforward to show that  $t = \tau \cosh \eta_s$ ,  $z = \tau \sinh \eta_s$ , and  $g_{\mu\nu} = \text{diag}(1, -1, -1, -\tau^2)$ . The equations of motion [Eq. (2)] are integrated forward in time with the help of the equation of state,  $\mathcal{P}(\epsilon)$ , which is obtained from lattice QCD analyses.<sup>1</sup> The code which realizes this is MUSIC [11], a 3 + 1D numerical hydrodynamics simulation which relies on the Kurganov-Tadmor algorithm [12]. In this work, the hydrodynamic evolution is done as in Ref. [11], with the hydroparameters specified in the last line of Table I there. The viscous evolution here uses a value of  $\eta/s = 1/4\pi$  and requires the initial energy density,  $\epsilon_0$ , to be 90% of the value in the inviscid case in order to account for entropy buildup by the dissipative dynamics. The hydrodynamic initial state here is free of fluctuations: A quantitative study of these effects will be done in an upcoming work.

#### B. Dilepton rates from perturbative QCD at high temperature

The production rate of dileptons depends on the local temperature. For massless quarks and antiquarks annihilating into lepton pairs, the (four-momentum-integrated) rate will go as  $R \sim T^4$  [13]. Most hydrodynamical simulation models used at RHIC energies require initial and kinematical freeze-out temperatures such that the net dilepton signal should originate from temperatures both above and below the range where lattice simulations predict a transition from hadronic degrees of freedom to a phase known as the quark-gluon plasma (QGP). In the limit of zero net baryon density, lattice calculations do not show a first- or even second-order phase transition but instead a smooth crossover centered at about 175 MeV (for a recent review, see, e.g., Ref. [14]). The dilepton contribution from the QGP phase will be the most visible for invariant masses above 1 GeV, where the mass scale is large compared to both  $\Lambda_{OCD}$  and T. For this channel, quark-antiquark annihilation at leading order (the Born approximation, for which the cross section actually does not depend on  $\alpha_s$ ) is often used as an approximation to the production rate at high temperature. At lower invariant mass, processes other than the Born term will contribute. Those may be parametrically of higher order in  $\alpha_s$  but typically contain collinear divergences that, when correctly resummed, produce a final result complete to leading order in the strong coupling. Those hard thermal loop (HTL) augmented rates rise over the Born result as the invariant mass is lowered but are still only available either in the zero momentum ( $\mathbf{q} = 0$  [15]) or large energy ( $q^0 \gtrsim T$  [16]) limits. Recently, the vector current correlation function has also been evaluated on the lattice, enabling the extraction of a thermal dilepton production rate at a given temperature and vanishing pair three-momentum [17]. The Born rate is used in this work, in part because of the kinematical restrictions still associated with the newer dilepton production rates but mostly because the rates from the confined hadronic sector of QCD will produce a dominant contribution at the invariant masses considered here (and for conditions prevalent at RHIC) [18]. In addition, this formulation is readily amenable to a viscous correction, as is discussed in the next section.

We start with high temperatures and use perturbative QCD. The cross section for  $q\bar{q} \rightarrow \ell^+ \ell^-$  is, neglecting quark and lepton masses,

$$\sigma = \frac{16\pi \alpha_{\rm EM}^2 \left(\sum_{q'} e_{q'}^2\right) N_c}{3q^2},$$
(3)

<sup>&</sup>lt;sup>1</sup>See the discussion in Ref. [1] and references therein.

where the index q' runs over quark flavors and q is the four-momentum of the virtual photon. The rate of dilepton production in the Born approximation is related to the cross section through

$$\frac{d^4R}{d^4q} = \int \frac{d^3p_1 d^3p_2}{(2\pi)^6 p_1^0 p_2^0} n_F(p_1) n_F(p_2) \frac{q^2}{2} \sigma \delta^4(q-p_1-p_2),$$
(4)

where  $p_1$  and  $p_2$  label the momenta of the incoming quarks and  $n_F$  is the Fermi-Dirac distribution. Integrating gives

$$\frac{d^4 R}{d^4 q} = \frac{\alpha_{\rm EM}^2}{6\pi^4} \frac{1}{\exp(\beta q^0) - 1} \left\{ 1 - \frac{2}{\beta |\mathbf{q}|} \ln\left[\frac{n_-}{n_+}\right] \right\}, \quad (5)$$

where  $n_{-} = 1 + \exp[\frac{-\beta(q^0 - |\mathbf{q}|)}{2}]$ ,  $n_{+} = 1 + \exp[\frac{-\beta(q^0 + |\mathbf{q}|)}{2}]$ ,  $\beta = 1/T$ ,  $|\mathbf{q}|$  is the norm of the three-momentum of the virtual photon,  $q^0$  is its energy, and we used q' = u, d, s.

When gradients of the flow or the temperature exist in a viscous fluid, departures from thermal equilibrium must occur, and those will modify the distribution functions. These departures should be reflected in the traces taken over the density matrix to determine what were previously thermal averages:  $n_{eq}(p) \rightarrow n_{eq}(p) + \delta n(p)$ . Various ansatzes for the form of  $\delta n(p)$  were examined in Ref. [19]. In this work we use

$$\delta n(p) = \frac{C}{2} n(p) [1 \pm n(p)] \frac{p^{\alpha} p^{\beta}}{T^2} \frac{\pi_{\alpha\beta}}{\epsilon + P}, \qquad (6)$$

where *C* is at this point an undetermined proportionality constant. Calculating  $T^{\mu\nu}$  with this expression and matching this to  $T^{\mu\nu}_{\text{ideal}} + \pi^{\mu\nu}$  (defined in the previous section) gives  $C_q = \frac{7\pi^4}{675\zeta(5)} \approx 0.97$  for the Fermi-Dirac distribution of a single-component massless quark fluid.

We modify the thermal average with this change in the Fermi-Dirac distribution where it appears, as was done in Ref. [20]. The details of this derivation are presented in Appendix A. This yields a viscous correction to dilepton production dependent on  $\pi^{\mu\nu}$ :

$$\frac{d^{4}R}{d^{4}q} = \frac{d^{4}R_{0}}{d^{4}q} + \frac{d^{4}\delta R}{d^{4}q},$$

$$\frac{d^{4}R_{0}}{d^{4}q} = \frac{q^{2}}{2}\frac{\sigma}{(2\pi)^{5}}\frac{1}{\exp(\beta q^{0}) - 1}\left\{1 - \frac{2}{\beta|\mathbf{q}|}\ln\left[\frac{n_{-}}{n_{+}}\right]\right\},$$

$$\frac{d^{4}\delta R}{d^{4}q} = \frac{q^{2}}{2}\frac{\sigma}{(2\pi)^{5}}C_{q}\frac{q^{\alpha}q^{\beta}}{T^{2}}\frac{\pi_{\alpha\beta}}{\epsilon + P}\frac{1}{2|\mathbf{q}|^{5}}$$

$$\times \int dE_{1}n(E_{1})n(q^{0} - E_{1})[1 - n(E_{1})]D,$$

$$D = \left[\left(3q_{0}^{2} - |\mathbf{q}|^{2}\right)E_{1}^{2} - 3q^{0}E_{1}q^{2} + \frac{3}{4}q^{4}\right].$$
(7)

The rates now depend on accurate calculation of the nonideal corrections to the energy-momentum tensor and therefore require viscous hydrodynamical simulation at RHIC and the LHC, such as MUSIC. In Sec. IV, the implications of the viscous correction on observables are explored.

#### C. Rates from a hadronic medium

Most of the thermal production of dileptons from a hadronic medium (HM) are produced by low-mass vector mesons V ( $V = \rho, \omega, \phi$ ). In fact the majority of the the emission below M < 1.2 GeV is of hadronic origin. A description for this production requires an accurate effective coupling between the electromagnetic field and the hadrons, which are composite particles and have, in general, complicated and mostly unknown electromagnetic form factors. The vector meson dominance model (VMD), first proposed by Sakurai [21], successfully describes dilepton production [22,23]. The effective coupling in this model is given by

$$\mathcal{L} = \mathcal{L}_{QED} - \sum_{V=\rho,\omega,\phi} \left[ \frac{\sqrt{4\pi \alpha_{\rm EM}}}{g_V} m_V^2 V^\mu A_\mu + \frac{1}{4} F_V^{\mu\nu} F_{\mu\nu}^V \right],$$
(8)

where  $F_V^{\mu\nu} = (\partial^{\mu}V^{\nu} - \partial^{\nu}V^{\mu})$  and  $\mathcal{L}_{QED} = \bar{\psi}_{\ell}(i\partial - m_{\ell})\psi_{\ell} - \sqrt{4\pi\alpha_{\rm EM}}\,\bar{\psi}_{\ell}\gamma^{\mu}\psi_{\ell}A_{\mu} - \frac{1}{4}F^{\mu\nu}F_{\mu\nu}$ . The coupling constants  $g_V$  are determined by measuring the vacuum decay rate of vector mesons to dileptons. The thermal rate of dilepton production for each low-mass vector meson *V* is then

$$\frac{d^4 R_V}{d^4 q} = -\frac{\alpha_{\rm EM}^2}{\pi^3} \frac{L(M)}{M^2} \frac{m_V^4}{g_V^2} \left[ \frac{{\rm Im} \, D_V^{\rm R}}{e^{\beta q^0} - 1} \right],\tag{9}$$

where  $L(M) = (1 + \frac{2m_{\ell}^2}{M^2})\sqrt{1 - \frac{4m_{\ell}^2}{M^2}}$ , and the imaginary part of the retarded vector propagator,  $\text{Im } D_V^R = \frac{1}{3}\text{Im} D_{\mu}^{\mu R}$ , was obtained using the Kubo-Martin-Schwinger (KMS) relation and VMD [22]. From here on, we set the lepton mass  $m_{\ell}$  to zero.

Viscous effects will modify the retarded self-energy and any averages that were originally thermal. To estimate these changes, we need to determine the viscous corrections to the thermal emission rate, Eq. (9). In a kinetic theory formulation of the dilepton production, the microscopic equilibrium distribution functions can be replaced with their nonequilibrium counterparts, and the net rates can then be recalculated in the different hadronic channels, as was done for real photons [5], and for the Born contribution of the previous section. Note that an empirically successful modeling of the electromagnetic current-current correlator for an interacting ensemble of baryons and mesons is achieved through hadronic many-body theory [24] and builds on Eq. (9). At this point, the viscous corrections to dilepton rates resulting from this approach have vet to be derived, and the nonperturbative extension of the KMS relation to the nonequilibrium realm is still a topic in development. Therefore, in this work a formalism based on experimental data is adopted and the consequences of shear viscous corrections to the self-energy are explored as detailed below

The total vector meson self-energy is given by [25]

$$\Pi_{V}^{\text{tot}}(M,|\mathbf{p}|,T) = \Pi_{V}^{\text{vac}}(M) + \Pi_{V}^{\text{T}}(M,|\mathbf{p}|,T) + \delta \Pi_{V}^{\text{T}}(M,|\mathbf{p}|,T).$$
(10)

The vacuum self-energy depends only on invariant mass  $M = \sqrt{E^2 - |\mathbf{p}|^2}$ , while the finite temperature contributions depend on *E* and  $|\mathbf{p}|$  (or *M* and  $|\mathbf{p}|$ ) as do the viscous corrections to the self-energy. Here, as in Ref. [25],  $\Pi_V^T$  is evaluated on the mass shell of *V*.

The calculations of  $\Pi_V^{\text{vac}}$  are present in Refs. [25–27]; there, terms in the Lagrangian describe all interactions contributing to  $\Pi_V^{\text{vac}}$ . We make Eq. (10) more explicit by writing  $\Pi_{Va}^{\text{T}}$  and relating it to the vacuum forward scattering amplitude  $f_{Va}(s)$  as in [25,28]:

$$\Pi_{Va}^{\rm T} = -4\pi \int \frac{d^3k}{(2\pi)^3} n_a(u \cdot k) \frac{\sqrt{s}}{\omega} f_{Va}(s), \qquad (11)$$

where *V* represents the vector meson and *a* is the particle with which the vector meson interacts. Also,  $n_a$  is a Fermi-Dirac or Bose-Einstein distribution function of the particle of type *a*,  $k^{\mu}$  is its four-momentum, while  $u^{\mu}$  is the velocity of the fluid cell. The work by Eletsky *et al.* [25] describes how to obtain the thermal correction to the vacuum self-energy of the vector meson *V* using  $f_{Va}(s)$  and Eq. (11). For a more in-depth derivation of the viscous correction to the self-energy, the reader is referred to Appendix B.

Using the relation between the self-energy and  $f_{Va}$ , and modifying the thermal average with viscous corrections, gives an expression for  $\delta \Pi_{Va}^{T}$  that is dependent on the nonideal correction to  $T^{\mu\nu}$ . Quoting from Appendix B, the result is

$$\Pi_{V}^{\text{tot}}(M, |\mathbf{p}|) = \Pi_{V}^{\text{vac}}(M) - \sum_{a=N,\bar{N},\pi} \frac{m_{V}m_{a}T}{\pi |\mathbf{p}|} \times \int_{m_{a}}^{\infty} d\omega' \ln \left[\frac{1 \pm \exp(-\omega_{+}/T)}{1 \pm \exp(-\omega_{-}/T)}\right] f_{Va}^{a's \text{ rest}}\left(\frac{m_{V}}{m_{a}}\omega'\right) + \sum_{a=N,\bar{N},\pi} C_{a}B_{2,Va}\frac{p_{V}^{\alpha}p_{V}^{\beta}}{T^{2}}\frac{\pi_{\alpha\beta}}{\epsilon + P},$$
(12)

where Eq. (B8) gives an expression for  $B_{2,Va}$ . Unlike the case of the QGP, where considering a single-component fluid was sufficient to compute  $C_q$ , the hadronic medium is a mixture of many particle species. Simplifying assumptions were made:  $C_a$  is particle independent [29] and  $\forall a : C_a = 1$ . Also, this work is done in the limit where the self-energies arising from interactions with antinucleons and nucleons are the same.

Physics relevant to interacting hadrons now enters this expression through  $f_{Va}$ , which receives contributions at both low and high energies;  $f_{Va}$  at low energy has both resonance and pomeron contributions. In the center-of-mass (c.m.) frame [25]

$$f_{Va}^{\text{c.m.}}(s) = \frac{1}{2q_{\text{c.m.}}} \sum_{R} W_{Va}^{R} \frac{\Gamma_{R \to Va}}{M_{R} - \sqrt{s} - \frac{1}{2}i\Gamma_{R}} - \frac{q_{\text{cm}}}{4\pi s} \frac{1 + \exp(-i\pi\alpha_{P})}{\sin(\pi\alpha_{P})} r_{Va}^{P} s^{\alpha_{P}}.$$
 (13)

The center-of-mass momentum is  $q_{c.m.}$ , which can be expressed in terms of the Mandelstam variable *s*. Here the sum ranges over resonances *R* that decay into the vector meson *V* and the particle *a*, which is either a nucleon or a pion. The spin and isospin are averaged, leading to the factor  $W_{Va}^R = \frac{(2s_R+1)}{(2s_V+1)(2s_a+1)} \frac{(2t_R+1)}{(2t_V+1)(2t_a+1)}$ , with  $s_i$  being the spin of particle *i*, and  $t_i$ , its isospin.  $\Gamma_{R \to Va}$  is an effective width of

the decay of the resonance *R* into *Va*. Its internal structure, and the types of resonances contributing to  $f_{Va}$ , are all discussed in detail in [25–27]. The values of the Regge residues  $r_{Va}$ , intercept  $\alpha_P$ , and the resonances included in  $f_{Va}$  are all given there. The transformation of the distribution function from the rest frame of *a* [as used in Eq. (12)] to the c.m. frame of particles *V* and *a* [Eq. (13)] is straightforward [22].

The high-energy limit of  $f_{Va}$  is described by a Regge parametrization [25–27]:

$$f_{Va}^{\text{c.m.}}(s) = -\frac{q_{\text{c.m.}}}{4\pi s} \sum_{i} \left[ \frac{1 + \exp(-i\pi\alpha_i)}{\sin(\pi\alpha_i)} \right] r_{Va}^i s^{\alpha_i}.$$
 (14)

The low-energy and the high-energy pieces are then matched onto one another at  $E_V - m_V \sim 4$  GeV for pions and  $E_V - m_V \sim 1$  GeV for nucleons, where  $E_V$  is evaluated in the rest frame of pions and nucleons respectively. To verify that the matching does not introduce violations of the Kramers-Kronig relations, a dispersion integral formula relating the real part of  $f_{Va}$  to a principal value integral over its imaginary part [25,27] is used:

$$\operatorname{Re}[f_{Va}(E_V)] = \operatorname{Re}[f_{Va}(0)] + \frac{2E_V^2}{\pi} \operatorname{P} \int_{m_V}^{\infty} \frac{\operatorname{Im}[f_{Va}(E')]dE'}{E'(E' + E_V)(E' - E_V)}$$
(15)

Indeed, as was shown in Refs. [25–27], the effect of the matching procedure on the shape of the forward-scattering amplitude is not significant. The dilepton rates derived in Secs. II B and II C have been used previously in an interpretation of NA60 data [30], taken at the CERN SPS.

#### **III. LEPTON PAIRS FROM CHARM DECAYS**

Dileptons originate not only from electromagnetic transitions but also from weak decays: A charm quark decays semileptonically into an electron with a branching fraction of approximately 10%. In a proton-proton collisions with center-of-mass energies of 200 GeV, dileptons produced from pairs of charm quarks dominate the yield in the intermediate mass range (from 1.2 to 2.5 GeV) [31]. Therefore, the analysis of the dilepton spectrum provides a measurement of the charm cross section. In this work, however, the emphasis is placed on the interaction (energy loss and gain, angular deflection) of heavy quarks with the hot and dense viscous matter, and how this will reflect itself in the dilepton spectrum. The production of heavy quarks in relativistic nuclear collisions—and their interaction with the hot and dense medium—is a topic that has received much attention over recent years [32–34].

The mass of a charm quark pair is much greater than the temperature reached in any model of the heavy-ion collisions at RHIC or the LHC; thermal production is negligible in comparison with the partonic annihilation in the initial collision. The mass of a charm quark pair is also significantly larger than  $\Lambda_{QCD}$ , and the production can be treated perturbatively. For proton-proton collisions, fixed-order next-to-leading-log (FONLL) calculations [35] fit the available experimental data well by including both next-to-leading order results at



FIG. 1. (Color online) Left panel (a): Dilepton yield from the hadronic medium (HM) only, in the 0–10% centrality class and fixed invariant mass  $M = m_{\rho}$ . The contributions from (i) ideal hydroevolution (dashed line), (ii) viscous hydroevolution alone (solid line), and (iii) viscous hydroevolution including viscous corrections to ideal dilepton rates are shown (square dots). Right panel (b): Dilepton yield from the QGP only, in the same centrality class, and for M = 1.5 GeV.

low momenta and terms proportional to  $\alpha_s \log(p/m)$  and  $\alpha_s^2 [\log(p/m)]^2$ , and by treating the heavy quarks as effectively massless at large  $p_T$ . In heavy-ion collisions, the initial production of charm (and anticharm) is affected by changes in the parton distribution functions: There can be-depending on the energy scale-shadowing and antishadowing of the parton distribution functions as well as isospin dependence of the heavy quark cross sections. The measured nuclear parton distribution functions can be evolved to different values of Q with the Dokshitzer-Gribov-Lipatov-Altarelli-Parisi (DGLAP) equations. Then, one needs to calculate the effect of the in-medium evolution of heavy quarks in heavy-ion collisions. The transport coefficients for heavy quarks have proven to be difficult to estimate reliably with hard-thermalloop effective theory [36]; however, for heavy quark momenta both less than and on the order of the heavy quark mass, the evolution of heavy quarks can be approximated to be diffusive and relativistic Langevin equations describe their dynamics [34], allowing the heavy quark diffusion coefficient to be estimated phenomenologically.

We use PYTHIA8 to generate events with heavy quarks. We also use EKS98 to determine the initial parton distribution functions in the nuclei. Then, using the same hydrodynamical description as was used to determine the thermal dilepton production, the heavy quarks are evolved using relativistic Langevin dynamics and the heavy quark spatial diffusion coefficient  $D_c = 3/(2\pi T)$ . The heavy quarks then hadronize according to Peterson fragmentation [37] into D,  $\overline{D}$ ,  $D^*$ , and  $\Lambda_c$  particles that then decay semileptonically. The quantitative results of our modeling are reported in Sec. IV C.

#### **IV. RESULTS**

# A. Thermal dilepton yield: the transverse momentum and invariant mass dependence

The yield of lepton pairs is obtained in our approach by integrating the production rates over the space-time history of the collision, using relativistic hydrodynamics to simulate the time- and space-dependent background fields. It is instructive to compare the transverse momentum spectra associated with different values of the dilepton invariant mass. In order to highlight in turn the hadronic and QGP thermal contributions, two values chosen can be associated with the "low-mass region" ( $M = m_{\rho}$ ), and the "intermediate-mass region" (M =1.5 GeV), respectively. We first consider the effect of viscous corrections only on the dileptons originating from the hadronic matter phase. In Fig. 1 (left panel), the dilepton yields as functions of  $p_T$  for the 0–10% centrality at a fixed invariant mass  $M = m_{\rho}$  are plotted, considering in turn three cases: that of inviscid hydrodynamics, then allowing for viscous corrections to the bulk evolution but not to the rates, and then finally correcting both the rates and the bulk evolution. The viscous effects on the bulk evolution in the hadronic phase raise the yield slightly ( $\sim$ 60%) at momenta from 3 to 4 GeV, as the viscous evolution slows down the temperature drop in the high-T portion of the hadronic phase [5,38]. We also notice, on the scale of the plot in the left panel of Fig. 1, that viscous corrections to the hadronic emission rates have basically no effect over that of the viscous evolution. The physical reason explaining the irrelevance of  $\delta n$  corrections on the yield arises from the fact that dileptons from the hadronic phase are mostly emitted late ( $\tau \gtrsim 4 \text{ fm/c}$ ) at which time  $\pi^{\mu\nu}$  is small (see the left panel of Fig. 2). Note that this explanation is somewhat qualitative as many cells with different temperatures contribute to the net dilepton yield. The statement is verified by a direct calculation, and viscous photon yields exhibit the same behavior [5]. Turning to dileptons from the QGP phase only, this receives further support from the dilepton transverse momentum spectrum for M = 1.5 GeV, shown on the right panel of Fig. 1. Correcting the bulk evolution only leads to a slight decrease of the yield at transverse momentum values of  $p_T \sim 2-4$  GeV. This is because the initial temperature in the viscous case is lower than that in the inviscid case, owing to entropy generation:<sup>2</sup> Recall that the entropy in the final state is

<sup>&</sup>lt;sup>2</sup>See, for example, Fig. 1 in Ref. [5].



FIG. 2. (Color online) Left panel (a): Shear stress tensor in the local rest frame of the cell located at x = y = 8/3 fm z = 0 fm in the 0–10% centrality class. Right panel (b): Total thermal dilepton yield (HM + QGP) as a function of  $p_T$  and at two different invariant masses:  $M = m_\rho$  and M = 1.5 GeV.

directly related to the observed particle multiplicity. Unlike the case of the hadron medium, the  $\delta n$  correction does influence the net dilepton yield as the emission occurs at early times when the temperature is high, which coincides with the proper time interval where the magnitude of the shear pressure tensor is maximal.

The right panel of Fig. 2 displays the net thermal dilepton yield (including both HM and QGP contributions) as a function of transverse momentum in the 0-10% centrality class, for two values of invariant mass. For invariant masses in the lowmass region, the higher momentum yield's sensitivity to the shear viscosity coefficient manifests itself almost exclusively through that of the bulk evolution. On the other hand, the thermal yield at higher invariant masses shows that the initial conditions (here, mainly  $T_i$ , the initial hydrotemperature), the hydroevolution, and the viscous corrections to the distribution functions all have an effect. While the different ingredients invoked here leave a quantitative imprint on the dilepton transverse momentum spectrum that is still quantitatively modest, these findings do confirm the power and the potential of lepton pairs as both a precise thermometer and viscometer. We leave the search for the specific conditions (e.g., centrality classes, different initial state conditions, beam energy scans, etc.) that will accentuate and perhaps even maximize those differences to an upcoming study.

The effect of viscous corrections to the dilepton invariant mass distribution is now investigated. It is straightforward to show that, owing to defining symmetry properties of the shear pressure tensor  $(u_{\mu}\pi^{\mu\nu} = u_{\nu}\pi^{\mu\nu} = 0)$  in the fluid rest frame and  $\pi^{\mu}_{\mu} = 0$ ), the viscous corrections to the QGP and HM dilepton rates as a function of the invariant mass M vanish:  $d \,\delta R/dM = 0$ . Hence, the differences between the invariant mass profiles in the inviscid and viscous cases entirely stem from the different time evolutions. For the conditions in this study, those appropriate for RHIC, the viscous evolution has an effect on the thermal dilepton spectrum that is essentially indistinguishable for that of the ideal hydrodynamic evolution: Only the viscous case is plotted in Fig. 3. The dilepton yield itself is therefore a poor viscometer. The spectrum asymmetry—as quantified by the elliptic flow—is now investigated.

#### B. Thermal dilepton elliptic flow

Regarding the shear viscosity and its experimental signature in relativistic heavy-ion collisions, flow coefficients, for example, that of elliptic flow, are more sensitive to the presence of viscosity than any particle spectra. Penetrating probes such as photons and dileptons are ideal to study viscosity, as they are influenced by the entire evolution of the medium [39–41]. Hadrons, on the other hand, will reflect properties that prevailed at the point of their last scattering.

The elliptic flow of thermal lepton pairs is quantified through  $v_2$ , a Fourier coefficient of the azimuthal angle expansion of the yield spectrum with respect to the reaction plane

$$\frac{dN}{dMp_T dp_T d\phi dy} = \frac{1}{2\pi} \frac{dN}{dMp_T dp_T dy} \times \left\{ 1 + \sum_{n=1}^{\infty} 2v_n \cos[n(\phi - \psi_r)] \right\}.$$
 (16)



FIG. 3. (Color online) Dilepton yield from hadronic medium and QGP as a function of invariant mass, in the 0-10% centrality class.



FIG. 4. (Color online) Dilepton  $v_2$  from the hadronic medium and QGP as a function of  $p_T$  for two invariant masses. The panel on the left (a) is for  $M = m_\rho$ , whereas the one on the right (b) is for M = 1.5 GeV (note the scaling applied to the HM  $v_2$ ). The calculations shown here are for the 0–10% centrality class.

With the averaged initial conditions used in this study,  $\psi_r$  is set to zero.

Shear viscosity introduces friction between adjacent fluid layers, thus coupling faster moving fluid layers to slower moving ones, which ultimately isotropizes the angular velocity distribution of the medium and slows down its expansion. As is the case for hadrons, the elliptic flow  $(v_2)$  of dileptons as a function of invariant mass is modified by the presence of shear viscosity. Following a sequence similar to that of the previous section, we start by presenting our  $v_2$  results as a function of  $p_T$  at fixed invariant masses [39] in Fig. 4.

At all invariant masses, the effect of viscosity is to reduce  $v_2$  of dileptons. This can be seen by comparing the red (light gray; ideal) and blue (dark gray; viscous) curves in Fig. 4. Importantly, when several sources of dileptons contribute to the net dilepton yield, the final  $v_2$  is a weighted average of the different elliptic flows, with the weight being the dilepton yield. This makes the interpretation of both panels of Fig. 4 clear: In the low-mass region, where the HM thermal dileptons outshine those from the QGP, one observes the net  $v_2$  to follow more closely that of the HM. At higher invariant masses (M =1.5 GeV) where the QGP yield dominates that of the HM, the final thermal dilepton  $v_2$  is close to that of the dileptons from the QGP, even though  $v_2^{\text{HM}} > v_2^{\text{QGP}}$ . Therefore, monitoring the thermal dileptons as a function of their invariant mass should help to map out the transition from a HM-dominated regime to that of a QGP. Together with a model for the time-evolution of the colliding system, such measurements could turn into a measurement of the effective temperature of the different phases. In addition, as is more clearly observed for the QGP dilepton distribution, the viscous corrections reduce the peak of  $v_2$  by 45% and shifts it to higher momenta, mainly because of the momentum dependence of  $\delta n$ . The results shown here consistently include the effects of viscosity, of using a mediumdependent vector spectral density, and of using a 3 + 1Dhydrodynamics simulation.

The distribution of  $v_2$  as a function of invariant mass is given in Fig. 5. There, one can clearly see that the peaks related to the  $\rho - \omega$  complex and to the  $\phi$  are present in the  $v_2$  spectrum—also noticed in Ref. [39]. Unlike the invariant

mass distribution of the yield, the  $v_2$  distribution is actually sensitive to the presence of viscosity: It is decreased compared to its value in the inviscid case (see Fig. 5). One also notices that  $\rho - \omega$  complex is made slightly broader by the viscous dynamics, owing to the different temperature and flow profiles.

The study of thermal dileptons is challenging experimentally, as competing sources have to be removed. In the intermediate mass region, the most important of these sources is charm and beauty hadrons.<sup>3</sup> Charmed and beauty hadrons require precise c- and b-quark tagging before they can be removed. However, the physics of heavy flavor dileptons is interesting in and of itself, as it opens a "clean" window to study heavy quark energy loss and gain mechanisms. Thus, Sec. IV C of this paper is precisely dedicated to heavy quarks, more specifically to charmed quarks.

<sup>3</sup>One reaction producing dileptons that was not included here is  $4\pi \rightarrow e^+e^-$ . This channel was found to be subdominant at SPS energies [30], but will be considered in the future.



FIG. 5. (Color online) The thermal dilepton  $v_2$  as a function of M for both ideal hydrodynamics (top curve) and viscous hydrodynamics (bottom curve).



FIG. 6. (Color online) Left panel (a): Dilepton invariant mass yields compared with experimental data at 0-10% centrality: importance of Langevin dynamics. Right panel (b): Dilepton invariant mass yields compared with experimental data at 0-10% centrality: importance of thermal radiation. The experimental acceptance cuts are indicated on the figures.

#### C. Including the dilepton contribution from charm decays

In order to make comparisons with experimental results for dilepton yields for invariant masses up to—and including—the "intermediate mass region," the contribution from semileptonic decays of charm to dileptons must be included. As discussed in Sec. III, the dynamics of heavy quarks whose velocity  $\gamma v \lesssim 1$  is approximated accurately with a relativistic Langevin equation for its momentum. We use MARTINI [42,43] as an event generator for charm quarks in heavy-ion collisions: The momenta of pairs of charm quarks are sampled using PYTHIA8 and the geometry in the transverse plane is sampled with the Glauber model, the Langevin equation is solved using the same calculations with MUSIC (including shear viscosity) that determined the thermal dilepton rates, and finally the species of charmed hadrons, and their decays, are sampled.

The total contribution to dN/dM is shown in Fig. 6 (left panel), representing the comparison of all our results with preliminary data from the STAR Collaboration [44] for the dilepton yields in gold-gold collisions at RHIC in the 0–10% centrality class. Note that the STAR acceptance requires the electron candidates to have  $|\eta^{e}| < 1$  and  $p_{T}^{e} > 0.2$  GeV, and dileptons to have  $|y^{ee}| < 1$ . Many  $\omega$ ,  $\rho$ , and  $\phi$  mesons are produced in these collisions and decay into dileptons; the data from STAR includes thermal dileptons as well as dileptons from the decays of the many hadrons produced in heavy-ion collisions. For this reason, we include the "cocktail" yield, as evaluated by the experimental collaboration: An extrapolation of hadron yields decaying to dilepton yields. The solid green (gray) line represents the sum of the thermal rates, the cocktail, and the contribution of charm without evolution in the medium, while the solid purple (gray) line represents the sum of the thermal rates and the cocktail with the charm contribution after evolving according to relativistic Langevin dynamics. The energy exchange of charm quarks with the medium leads to a depletion in dN/dM at large M, and the charm contribution alone can differ by an order of magnitude at M = 2.1 GeV, depending on whether Langevin evolution is considered. The data have a slight preference for Langevin evolution, but the

size of the error precludes a stronger conclusion at this point. However, the inclusion—or not—of the possibility of charm energy variation will affect any determination of the cross sections using data for dilepton yields. At lower invariant masses, the STAR data seem compatible with this theoretical calculation. However, it is clear that acceptance-corrected data will make a much more compelling case for model compatibility.

The right panel of Fig. 6 investigates the importance of thermal radiation to describe the STAR data. In the low invariant mass region, the cocktail systematically underestimates the data and including charmed hadrons (with Langevin dynamics) is not enough to raise the calculation to the level of the measurements: The inclusion of thermal radiation is crucial. For intermediate dilepton invariant masses, the situation is less clear, given STAR's current experimental uncertainties. However, the trend does suggest that thermal radiation from the QGP is present and must participate in the interpretation of the data.

The STAR Collaboration also has preliminary measurements of minimum bias  $v_2(M)$  of dileptons (albeit with still large error bars) over a large momentum range, and this also includes the dileptons produced by semileptonic decays of charmed mesons. A comparison with these data requires knowledge of the elliptic flow of the hadronic cocktail, which we leave for a future work. The theoretical results for this observable are shown in Fig. 7, not including the contribution of the cocktail. Including the charm contribution to  $v_2$  has two important effects: First, it reduces the  $v_2$  in the 0- to 1-GeV invariant mass range by about a factor of two, and it increases the  $v_2$  in the 1.5- to 2-GeV invariant mass range where the charm contribution dominates the dilepton yields. The flow of the charm contribution is smaller than the flow of the hadronic matter contribution and it is larger than the flow of the QGP contribution, but also bear in mind that the net elliptic flow is a weighted average of its individual components. Notably, the absolute magnitude of the final elliptic flow is small. Let it be made clear again: No efforts have been made here to search for conditions that will maximize this signal, such as going to a



FIG. 7. (Color online) Dilepton invariant mass  $v_2$  including thermal and charm contributions at 0–10% centrality.

higher centrality class, including fluctuating initial states, etc. This is left for a future systematic investigation of these effects.

Before leaving this section on results and moving to a conclusion, it is pertinent to recall that electromagnetic radiation samples the entire space-time history of the colliding system, not just the freeze-out stage. The validity for all times of the viscosity correction linear in the viscous pressure tensor [see Eq. (6)] to the thermal distribution functions can then be questioned. This investigation was performed in Ref. [5]; those results still hold and will not be repeated here. Suffice it to say that improved versions of  $\delta n$  will be explored in an upcoming work.

## V. CONCLUSION

In this paper, we have conducted a systematic study of viscosity effects on dilepton spectra in heavy-ion collisions (a) in the microscopic emission rates, (b) in the macroscopic evolution, and (c) in the semileptonic contribution. Viscosity affects the net thermal dilepton spectrum by first inducing a correction to the hadronic distribution functions. These corrections will mostly be seen in the part of the signal that is attributable to the QGP, as the shear pressure tensor  $\pi_{\mu\nu}$  is maximal in this phase. After describing the dilepton radiation in a hadronic ensemble gas and in a quark-gluon plasma and the viscous effects on the rates, those have been integrated with MUSIC in order to consistently investigate how the viscous dynamics affects the dilepton yield and elliptic flow. Note that viscosity will also affect the cooling rate of the hydrodynamic medium, which in turn will influence both the QGP and HM thermal dileptons. For essentially all conditions considered here, the effects of the viscous dynamics are numerically not large but are non-negligible. Moreover, and importantly, the viscous corrections are required to ensure theoretical consistency.

For the purpose of comparison with recent experimental data, the calculations presented in this work include a Langevin evolution of charmed quark distributions in a viscous hydrodynamics background. The dilepton signal originating from the charm decays was then added to that of thermal sources. These results compared well with preliminary data on Au-Au collisions from the STAR Collaboration at RHIC, suggesting that the data are consistent with the viscous corrections on both microscopic rates and macroscopic dynamics. As argued previously by many authors, the intermediate invariant mass region opens the possibility of measuring the energy shift of heavy quarks that interact with the hot and dense evolving medium, and the results shown here also support this assertion. Our calculations also suggest that it should be possible to access the QGP dilepton radiation in the intermediate mass region-from 1.2 to 2.5 GeV-provided that precise experimental tagging of heavy flavor exists. In that case, it may be experimentally possible to remove the lepton pairs originating from open charm and beauty decays, thus exposing direct radiation from the QGP. A simultaneous analysis of yield and  $v_2$  of the high-mass lepton pairs, coupled with a removal of nonphotonic electrons, would produce a clear picture of the early stages of the nuclear collision. As written earlier in this paper, future work will include a study of varying the initial states existing prior to the hydrodynamical evolution, as well as an exploration of the effects of the different QCD transport coefficients. In what concerns measurements, the program at RHIC together with dilepton measurements at the LHC will produce the necessary beacons of the QCD phase diagram.

#### ACKNOWLEDGMENTS

We are happy to acknowledge helpful discussions with G. Denicol, K. Dusling, I. Kozlov, M. Luzum, J.-F. Paquet, L. Ruan, and R. Vogt. This work was supported in part by the Natural Sciences and Engineering Research Council of Canada, U.S. DOE Contract No. DE-AC02-98CH10886, and U.S. NSF Grant No. PHY-1306359.

#### APPENDIX A: VISCOUS CORRECTIONS TO QGP RATES

The ansatz for the form of the viscous correction that we chose to utilize for the QGP was previously explored in Ref. [20]. This ansatz originates from the continuity requirement of the the stress-energy tensor (or the Cooper-Frye formula) across the freeze-out surface. At freeze-out, the stress-energy tensor from the hydrodynamical simulation must be matched to the one from kinetic theory. That is,

$$T_{\text{ideal}}^{\mu\nu} + \pi^{\mu\nu} = \int \frac{d^3p}{(2\pi)^3 p^0} p^{\mu} p^{\nu} [n(p \cdot u) + \delta n(p \cdot u)]. \quad (A1)$$

Requiring that the stress-energy tensor be continuous during the entire hydrosimulation implies that the viscous correction to the equilibrium distribution function must be present in dileton production rates. For the extension to the thermal distribution, we use

$$n_{\text{total}}(p \cdot u) = n(p \cdot u) + \delta n(p \cdot u)$$
  
=  $n(p \cdot u) + \frac{C}{2T^2(\epsilon + P)} n(p \cdot u) [1 \pm n(p \cdot u)] p^{\alpha} p^{\beta} \pi_{\alpha\beta}$   
=  $n(p \cdot u) + \frac{C}{2} n(p \cdot u) [1 \pm n(p \cdot u)] \frac{p^{\alpha}}{T} \frac{p^{\beta}}{T} \frac{\pi_{\alpha\beta}}{\epsilon + P},$   
(A2)

where  $p^{\alpha}$  is the four-momentum of one of the incoming quarks,  $\epsilon + P$  is the local energy density and pressure respectively, T is the temperature, and  $\pi_{\alpha\beta}$  is the shear-stress tensor of the fluid. Substituting Eq. (A2) into Eq. (A1) yields

$$\pi^{\mu\nu} = \left[\frac{C}{2} \int \frac{d^3p}{(2\pi)^3 p^0} n(p \cdot u) [1 \pm n(p \cdot u)] p^{\mu} p^{\nu} \frac{p^{\alpha}}{T} \frac{p^{\beta}}{T}\right] \times \frac{\pi_{\alpha\beta}}{\epsilon + P},$$
(A3)

C is a proportionality constant that relates the hydrodynamical shear-stress tensor to its kinetic theory counterpart. In the context of a single component ensemble, C can be determined via [45]

$$\eta = \frac{C}{15T^3} \int \frac{d^3p}{(2\pi)^3 p^0} n(p \cdot u) [1 \pm n(p \cdot u)] [p^2 - (u \cdot p)^2]^2.$$
(A4)

One can solve for *C* in Eq. (A4) by expressing  $T^3$  in terms of entropy density:

$$s = \frac{4}{3}\frac{\epsilon}{T},\tag{A5}$$

$$\epsilon = \frac{T^4 g}{2\pi^2} \int_y^\infty \frac{x^3 \sqrt{1 - (y/x)^2} dx}{e^x \pm 1},$$
 (A6)

where  $\epsilon$  is the average energy density of a Fermi or Bose gas with distribution n,  $x = (p \cdot u)/T$ ,  $y = \sqrt{p^2}/T$ , g is the spin degeneracy factor, and  $p^2$  is the four-momentum squared. Finally solving for C is simplest in the rest frame of the fluid:

$$C = \frac{4a}{3\tilde{b}},$$
  

$$\tilde{a} = \frac{1}{2\pi^2} \int_y^\infty dx \frac{x^3 \sqrt{1 - (y/x)^2}}{e^x \mp 1},$$
 (A7)  

$$\tilde{b} = \frac{1}{30\pi^2} \int_y^\infty dx \frac{x^5 [1 - (y/x)^2]^{5/2}}{e^x \mp 1} \left\{ 1 \pm \frac{1}{e^x \mp 1} \right\}.$$

For the specific case of the QGP, in the approximation of a single-component fluid of massless quarks, *C* can be evaluated analytically and is  $C_q = \frac{7\pi^4}{675\zeta(5)} \approx 0.97$ . The modification of the distribution functions owing to

The modification of the distribution functions owing to viscosity have a nontrivial effect on the viscous rates of QGP dileptons. Since we will be including viscous effects on the hadronic dilepton rates, it is instructive to carefully explore the manner in which the simpler Born QGP rates get modified. Indeed, we use the same procedure for the HM case.

In the massless quark limit,

$$\frac{d^{4}R}{d^{4}q} = \int \frac{d^{3}p_{1}d^{3}p_{2}}{(2\pi)^{6}p_{1}^{0}p_{2}^{0}}n(p_{1}\cdot u)n(p_{2}\cdot u)\frac{q^{2}}{2}\sigma\delta^{4}(q-p_{1}-p_{2}) + \int \frac{d^{3}p_{1}d^{3}p_{2}}{(2\pi)^{6}p_{1}^{0}p_{2}^{0}}n(p_{1}\cdot u)n(p_{2}\cdot u)[1-n(p_{1}\cdot u)]\frac{q^{2}}{2}\sigma\delta^{4}(q-p_{1}-p_{2})C_{q}\frac{p_{1}^{\alpha}}{T}\frac{p_{1}^{\beta}}{T}\frac{\pi_{\alpha\beta}}{\epsilon+P}, \\
\frac{d^{4}R}{d^{4}q} = \frac{d^{4}R_{ideal}}{d^{4}q} + C_{q}\frac{J^{\alpha\beta}}{T^{2}}\frac{\pi_{\alpha\beta}}{\epsilon+P},$$
(A8)

4~

where we decomposed the rate into its ideal and viscous contributions, ignoring all viscous corrections of order  $(\delta n)^2$ . Performing this integral is nontrivial. However, we know that the tensor  $J^{\alpha\beta}$  of viscous correction to the rate must solely depend on the momentum of the virtual photon  $q^{\alpha}$ , the flow  $u^{\alpha}$ , and the metric  $g^{\alpha\beta}$ . Hence,

$$J^{\alpha\beta} = b_0 g^{\alpha\beta} + b_1 u^{\alpha} u^{\beta} + b_2 q^{\alpha} q^{\beta} + b_3 (u^{\alpha} q^{\beta} + u^{\beta} q^{\alpha}) + b_4 (u^{\alpha} q^{\beta} - u^{\beta} q^{\alpha}).$$
(A9)

This is the most general form one can write down for  $J^{\alpha\beta}$ . However, since  $J^{\alpha\beta}$  is contracted with  $\pi^{\alpha\beta}$ —which must be a symmetric tensor (as part of  $T^{\alpha\beta}$ ); any antisymmetric piece of  $J^{\alpha\beta}$  must not contribute to this calculation as shown below. The coefficients  $b_0$  through  $b_4$  are obtained as

$$\begin{bmatrix} g^{\alpha\beta}J_{\alpha\beta} \\ u^{\alpha}u^{\beta}J_{\alpha\beta} \\ q^{\alpha}q^{\beta}J_{\alpha\beta} \\ (u^{\alpha}q^{\beta} + u^{\beta}q^{\alpha})J_{\alpha\beta} \\ (u^{\alpha}q^{\beta} - u^{\beta}q^{\alpha})J_{\alpha\beta} \end{bmatrix} = \begin{bmatrix} 4 & 1 & q^{2} & 2(u \cdot q) & 0 \\ 1 & 1 & (u \cdot q)^{2} & 2(u \cdot q) & 0 \\ q^{2} & (u \cdot q)^{2} & q^{4} & 2q^{2}(u \cdot q) & 0 \\ 2(u \cdot q) & 2(u \cdot q) & 2q^{2}(u \cdot q) & 2(q^{2} + (u \cdot q)^{2}) & 0 \\ 0 & 0 & 0 & 0 & 2q^{2} \end{bmatrix} \begin{bmatrix} b_{0} \\ b_{1} \\ b_{2} \\ b_{3} \\ b_{4} \end{bmatrix},$$
(A10)

whose solution is

$$\begin{bmatrix} b_0 \\ b_1 \\ b_2 \\ b_3 \\ b_4 \end{bmatrix} = \begin{bmatrix} \frac{1}{2} & -\frac{1}{2} \frac{q^2}{q^2 - (u \cdot q)^2} & -\frac{1}{2} \frac{1}{q^2 - (u \cdot q)^2} & \frac{1}{2} \frac{u \cdot q}{q^2 - (u \cdot q)^2} & 0 \\ -\frac{1}{2} \frac{q^2}{q^2 - (u \cdot q)^2} & \frac{3}{2} \left[ \frac{q^2}{q^2 - (u \cdot q)^2} \right]^2 & \frac{1}{2} \frac{q^2 + 2(u \cdot q)^2}{[q^2 - (u \cdot q)^2]^2} & -\frac{3}{2} \frac{q^2(u \cdot q)}{[q^2 - (u \cdot q)^2]^2} & 0 \\ -\frac{1}{2} \frac{1}{q^2 - (u \cdot q)^2} & \frac{1}{2} \frac{q^2 + 2(u \cdot q)^2}{[q^2 - (u \cdot q)^2]^2} & \frac{3}{2} \frac{1}{[q^2 - (u \cdot q)^2]^2} & -\frac{3}{2} \frac{u \cdot q}{[q^2 - (u \cdot q)^2]^2} & 0 \\ \frac{1}{2} \frac{u \cdot q}{q^2 - (u \cdot q)^2} & -\frac{3}{2} \frac{q^2(u \cdot q)}{[q^2 - (u \cdot q)^2]^2} & -\frac{3}{2} \frac{(u \cdot q)}{[q^2 - (u \cdot q)^2]^2} & 0 \\ 0 & 0 & 0 & 0 & \frac{1}{2q^2} \end{bmatrix} \begin{bmatrix} g^{\alpha\beta} J_{\alpha\beta} \\ u^{\alpha} u^{\beta} J_{\alpha\beta} \\ (u^{\alpha} q^{\beta} + u^{\beta} q^{\alpha}) J_{\alpha\beta} \\ (u^{\alpha} q^{\beta} - u^{\beta} q^{\alpha}) J_{\alpha\beta} \end{bmatrix}.$$

A simplification of the second-rank tensor  $J^{\alpha\beta}$  is made possible by using the identities  $u^{\alpha}\pi_{\alpha\beta} = g^{\alpha\beta}\pi_{\alpha\beta} = 0$ . Indeed,  $J^{\alpha\beta}$  is only proportional to  $q^{\alpha}q^{\beta}$  and the proportionality constant  $b_2$  is obtained via the projection operator

$$P_{\alpha\beta} = \frac{1}{2} \frac{g_{\alpha\beta}}{(u \cdot q)^2 - q^2} + \frac{1}{2} \left[ \frac{q^2 + 2(u \cdot q)^2}{[q^2 - (u \cdot q)^2]^2} \right] u_{\alpha} u_{\beta} + \frac{3}{2} \frac{q_{\alpha} q_{\beta}}{[q^2 - (u \cdot q)^2]^2} - \frac{3}{2} \left[ \frac{u \cdot q}{[q^2 - (u \cdot q)^2]^2} \right] (u_{\alpha} q_{\beta} + u_{\beta} q_{\alpha}).$$
(A11)

Since  $P_{\alpha\beta}J^{\alpha\beta}$  is a Lorentz invariant quantity, the most efficient way to compute it is in the rest frame of the fluid cell. Performing that computation yields

$$b_{2} = P_{\alpha\beta}J^{\alpha\beta} = \frac{1}{2|\mathbf{q}|^{5}}\int_{E_{-}}^{E_{+}} \frac{dE_{1}}{(2\pi)^{5}} \frac{q^{2}}{2} \sigma n(E_{1})n(q^{0} - E_{1})[1 - n(E_{1})]D, \quad D = \left[(3q_{0}^{2} - |\mathbf{q}|^{2})E_{1}^{2} - 3q^{0}E_{1}q^{2} + \frac{3}{4}q^{4}\right], \quad (A12)$$

where  $E_{\pm} = \frac{q^0 \pm |\mathbf{q}|}{2}$ . Finally, the Born rate with viscous corrections reads

$$\frac{d^4R}{d^4q} = \frac{q^2}{2} \frac{\sigma}{(2\pi)^5} \left[ \frac{1}{\exp(\beta q^0) - 1} \left\{ 1 - \frac{2}{\beta |\mathbf{q}|} \ln\left[\frac{n_-}{n_+}\right] \right\} + C_q \frac{q^\alpha}{T} \frac{q^\beta}{T} \frac{\pi_{\alpha\beta}}{\epsilon + P} \frac{1}{2|\mathbf{q}|^5} \int_{E_-}^{E_+} dE_1 n(E_1) n(q^0 - E_1) [1 - n(E_1)] D \right].$$
(A13)

### APPENDIX B: THE VECTOR MESON SELF-ENERGY AND ITS VISCOUS CORRECTION

Using the tools of the previous section, the goal of this section is to derive the viscous correction to the self-energy. To this end, we first outline the steps leading to the thermal self-energy and then we extend it to include viscous corrections.

#### 1. Thermal self-energy

To simplify the calculation, and without loss of generality, we choose the z axis such that the four-momentum of particle V is aligned with it, i.e.,  $p^{\mu} = (E, 0, 0, |\mathbf{p}|)$ . We further define the angle  $\theta$  between the z axis and the momentum  $k^{\mu} = (\omega, \mathbf{k})$  of particle a. Note that  $\theta$  is not the angle between  $p^{\mu}$  and  $k^{\mu}$ .

In the rest frame of particle a, it is possible to evaluate the angular part of the self-energy integral. From now on, prime (') is used to denote energy and momentum in V's rest frame and double prime (") is used to label a's rest frame. One can relate the energy in the two frames via

$$s = m_V^2 + m_a^2 + 2E''m_a = m_V^2 + m_a^2 + 2m_V\omega'.$$
(B1)

Hence,  $E'' = \frac{m_V}{m_a}\omega'$ . Furthermore, in V's rest frame,  $\omega = \frac{E\omega' + |\mathbf{p}||\mathbf{k}'|z'}{m_V}$ , where  $z' = \cos\theta'$ . Putting everything together,

$$\Pi_{Va}^{\mathrm{T}}(|\mathbf{p}|,T) = -4\pi \int \frac{d^{3}k}{(2\pi)^{3}\omega} n_{a}(\omega) \sqrt{s} f_{Va}^{\mathrm{c.m.}}(s) = -4\pi \int \frac{|\mathbf{k}'|^{2} d|\mathbf{k}' |dz'}{(2\pi)^{2}\omega'} n_{a} \left(\frac{E\omega' + |\mathbf{p}||\mathbf{k}'|z'}{m_{V}}\right) f_{Va}^{\mathrm{a's \, rest}} \left(\frac{m_{V}}{m_{a}}\omega'\right)$$
$$= -\frac{m_{V}}{\pi} \int_{m_{a}}^{\infty} |\mathbf{k}'| d\omega' f_{Va}^{\mathrm{a's \, rest}} \left(\frac{m_{V}}{m_{a}}\omega'\right) \int_{-1}^{1} dz' n_{a} \left(\frac{E\omega' + |\mathbf{p}||\mathbf{k}'|z'}{m_{V}}\right)$$
$$= -\frac{m_{V}m_{a}T}{\pi |\mathbf{p}|} \int_{m_{a}}^{\infty} d\omega' \ln \left[\frac{1 \pm \exp\left(-\omega_{+}/T\right)}{1 \pm \exp\left(-\omega_{-}/T\right)}\right] f_{Va}^{\mathrm{a's \, rest}} \left(\frac{m_{V}}{m_{a}}\omega'\right), \tag{B2}$$

where  $\omega_{\pm} = \frac{E\omega' \pm |\mathbf{p}||\mathbf{k}'|}{m_V}$ . This expression for the self-energy is evaluated on the mass shell of the vector meson V.

## 2. Viscous corrections to the thermal self-energy

To calculate the viscous correction to the thermal self-energy, we proceed by including the  $\delta n$  correction to the thermal distribution function. Unlike the Bose distribution function present in the rates—which originates from the KMS relation and therefore is not related to the thermal distribution function of vector mesons—the distribution function present in the self-energy, Eq. (11), is indeed a distribution function of thermal particles, so the viscous correction to the thermal distribution in Eq. (A2) applies. Thus,

$$\delta \Pi_{Va}^{\mathrm{T}}(|\mathbf{p}|,T) = -4\pi \int \frac{d^3k}{(2\pi)^3 \omega} \delta n_a(k \cdot u) \sqrt{s} f_{Va}^{\mathrm{c.m.}}(s) = C_a \frac{K^{\alpha\beta}}{T^2} \frac{\pi_{\alpha\beta}}{\epsilon + P}.$$
(B3)

Note that  $C_a$  cannot be computed via Eq. (A7), since  $\delta \Pi_{Va}^T$  is describing a multicomponent mixture. Hence, a simplifying assumption is made:  $\forall a \ C_a = 1$ . Now we expand the tensor  $K^{\mu\nu}$  in the same manner as the QGP  $J^{\mu\nu}$  tensor encountered earlier:

$$K^{\mu\nu} = B_0 g^{\alpha\beta} + B_1 u^{\alpha} u^{\beta} + B_2 p^{\alpha} p^{\beta} + B_3 (u^{\alpha} p^{\beta} + u^{\beta} p^{\alpha}) + B_4 (u^{\alpha} p^{\beta} - u^{\beta} p^{\alpha}).$$
(B4)

Since the relation  $u^{\alpha}\pi_{\alpha\beta} = g^{\alpha\beta}\pi_{\alpha\beta} = 0$  still holds, we use the same projection operator as in Eq. (A11) to determine  $B_2$ . Thus,

$$B_{2,Va} = P_{\alpha\beta}K^{\alpha\beta} = -4\pi \int \frac{d^3k}{(2\pi)^3} n_a(u \cdot k) [1 \pm n_a(u \cdot k)] \frac{\sqrt{s}}{\omega} f_{Va}(s) \left[ \frac{1}{2} \frac{m_a^2}{(u \cdot p)^2 - p^2} + \frac{1}{2} \left[ \frac{p^2 + 2(u \cdot p)^2}{[p^2 - (u \cdot p)^2]^2} \right] (u \cdot k)^2 + \frac{3}{2} \frac{(p \cdot k)^2}{[p^2 - (u \cdot p)^2]^2} - 3 \frac{(u \cdot p)(u \cdot k)(p \cdot k)}{[p^2 - (u \cdot p)^2]^2} \right].$$
(B5)

Throughout this appendix, the upper (lower) sign refers to bosons (fermions). In the rest frame of the medium (using  $z = \cos \theta$  as before)

$$B_{2,Va} = -4\pi \int \frac{d^3k}{(2\pi)^3\omega} n_a (1\pm n_a) \sqrt{s} f_{Va} \left[ \frac{m_a^2}{2|\mathbf{p}|^2} + \left( \frac{3E^2}{2|\mathbf{p}|^4} - \frac{1}{2|\mathbf{p}|^2} \right) \omega^2 + \frac{3}{2} \frac{(E\omega - |\mathbf{p}||\mathbf{k}|z)^2}{|\mathbf{p}|^4} - \frac{3E\omega(E\omega - |\mathbf{p}||\mathbf{k}|z)}{|\mathbf{p}|^4} \right]$$
  
$$= -4\pi \int \frac{d^3k}{(2\pi)^3\omega} n_a (1\pm n_a) \sqrt{s} f_{Va} \left[ \frac{m_a^2}{2|\mathbf{p}|^2} + \frac{3|\mathbf{k}|^2 z^2 - \omega^2}{2|\mathbf{p}|^2} \right].$$
(B6)

By evaluating the integral in the rest frame of a, we obtain

$$B_{2,Va} = -4\pi m_a \int \frac{d^3 k'}{(2\pi)^3 \omega'} n_a \left(\frac{E\omega' + |\mathbf{p}||\mathbf{k}'|z'}{m_V}\right) \left[1 \pm n_a \left(\frac{E\omega' + |\mathbf{p}||\mathbf{k}'|z'}{m_V}\right)\right] f_{Va}^{a's rest} \left(\frac{m_V}{m_a}\omega'\right) \\ \times \left[\frac{m_a^2}{2|\mathbf{p}|^2} + \frac{3\left(\frac{E|\mathbf{k}'|z' + |\mathbf{p}|\omega'}{m_V}\right)^2 - \left(\frac{E\omega' + |\mathbf{p}||\mathbf{k}'|z'}{m_V}\right)}{2|\mathbf{p}|^2}\right] \\ = -\frac{m_a}{2\pi |\mathbf{p}|^2} \int_{m_a}^{\infty} d\omega' |\mathbf{k}'| f_{Va}^{a's rest} \left(\frac{m_V}{m_a}\omega'\right) \int_{-1}^{1} dz' n_b \left(\frac{E\omega' + |\mathbf{p}||\mathbf{k}'|z'}{m_V}\right) \left[1 \pm n_a \left(\frac{E\omega' + |\mathbf{p}||\mathbf{k}'|z'}{m_V}\right)\right] \\ \times \left[m_a^2 + (3|\mathbf{p}|^2 - E^2)\frac{\omega'^2}{m_V^2} + 4E|\mathbf{p}|\frac{\omega'|\mathbf{k}'|}{m_V^2}z' + (3E^2 - |\mathbf{p}|^2)\frac{|\mathbf{k}'|^2}{m_V^2}z'^2\right], \tag{B7}$$

where  $|\mathbf{k}|_{z} = \frac{E}{m_{V}} |\mathbf{k}'|_{z'} + \frac{|\mathbf{p}|}{m_{V}} \omega'$ . Performing the angular integral yields

$$B_{2,Va} = -\frac{m_V}{2\pi |\mathbf{p}|^2} \int_{m_a}^{\infty} d\omega' |\mathbf{k}'| f_{Va}^{a's \text{ rest}} \left(\frac{m_V}{m_a}\omega'\right) \times \left(\mathcal{A} + \mathcal{B} + \mathcal{C} + \mathcal{D} + \mathcal{E}\right),\tag{B8}$$

where

$$\mathcal{A} = \left(\frac{m_{V}T}{|\mathbf{p}||\mathbf{k}'|}\right) \left[m_{a}^{2} + \frac{(E|\mathbf{k}'| - |\mathbf{p}|\omega)^{2} - (E\omega' - |\mathbf{p}||\mathbf{k}'|)^{2}}{m_{V}^{2}}\right] [\exp(\omega_{-}/T) \mp 1]^{-1},$$

$$\mathcal{B} = -\left(\frac{m_{V}T}{|\mathbf{p}||\mathbf{k}'|}\right) \left[m_{a}^{2} + \frac{(E|\mathbf{k}'| + |\mathbf{p}|\omega)^{2} - (E\omega' + |\mathbf{p}||\mathbf{k}'|)^{2}}{m_{V}^{2}}\right] [\exp(\omega_{+}/T) \mp 1]^{-1},$$

$$\mathcal{C} = \pm 2\left(\frac{m_{V}T}{|\mathbf{p}||\mathbf{k}'|}\right)^{2} \left[(3E^{2} - |\mathbf{p}|^{2})\frac{|\mathbf{k}'|^{2}}{m_{V}^{2}} + 2\frac{E\omega'|\mathbf{p}||\mathbf{k}'|}{m_{V}^{2}}\right] \ln[1 \mp \exp(-\omega_{+}/T)],$$

$$\mathcal{D} = \pm 2\left(\frac{m_{V}T}{|\mathbf{p}||\mathbf{k}'|}\right)^{2} \left[(3E^{2} - |\mathbf{p}|^{2})\frac{|\mathbf{k}'|^{2}}{m_{V}^{2}} - 2\frac{E\omega'|\mathbf{p}||\mathbf{k}'|}{m_{V}^{2}}\right] \ln[1 \mp \exp(-\omega_{-}/T)],$$

$$\mathcal{E} = \mp 2\left(\frac{m_{V}T}{|\mathbf{p}||\mathbf{k}'|}\right)^{3} \left[(3E^{2} - |\mathbf{p}|^{2})\frac{|\mathbf{k}'|^{2}}{m_{V}^{2}}\right] \left\{\text{Li}_{2}[\pm \exp(-\omega_{+}/T)] - \text{Li}_{2}[\pm \exp(-\omega_{-}/T)]\right\},$$
(B9)

and  $Li_2$  is the dilogarithm function. Thus, the total self-energy is

$$\Pi_{V}^{\text{tot}}(M,|\mathbf{p}|,T) = \Pi_{V}^{\text{vac}}(M) + \sum_{a=N,\bar{N},\pi} \left\{ -\frac{m_{V}m_{a}T}{\pi|\mathbf{p}|} \int_{m_{a}}^{\infty} d\omega' \ln\left[\frac{1 \pm \exp\left(-\omega_{+}/T\right)}{1 \pm \exp\left(-\omega_{-}/T\right)}\right] f_{Va}^{a's \text{ rest}}\left(\frac{m_{V}}{m_{a}}\omega'\right) + C_{a}B_{2,Va}\frac{p_{V}^{\alpha}p_{V}^{\beta}}{T^{2}}\frac{\pi_{\alpha\beta}}{\epsilon + P} \right\}.$$
(B10)

- [1] C. Gale, S. Jeon, and B. Schenke, Int. J. Mod. Phys. A 28, 1340011 (2013).
- [2] R. Baier, P. Romatschke, D. T. Son, A. O. Starinets, and M. A. Stephanov, JHEP 04 (2008) 100.
- [3] B. Betz, D. Henkel, and D. Rischke, J. Phys. G 36, 064029 (2009).
- [4] C. Gale, S. Jeon, B. Schenke, P. Tribedy, and R. Venugopalan, Phys. Rev. Lett. 110, 012302 (2013).
- [5] M. Dion, J.-F. Paquet, B. Schenke, C. Young, S. Jeon, and C. Gale, Phys. Rev. C 84, 064901 (2011).
- [6] R. Chatterjee, H. Holopainen, T. Renk, and K. J. Eskola, Phys. Rev. C 83, 054908 (2011).
- [7] C. Shen, U. W. Heinz, J.-F. Paquet, I. Kozlov, and C. Gale, arXiv:1308.2111 [nucl-th].
- [8] T. Hirano, N. van der Kolk, and A. Bilandzic, Lect. Notes Phys. 785, 139 (2010).
- [9] U. W. Heinz, arXiv:0901.4355 [nucl-th].
- [10] W. Israel and J. Stewart, Annals Phys. 118, 341 (1979).
- [11] B. Schenke, S. Jeon, and C. Gale, Phys. Rev. C 82, 014903 (2010).
- [12] A. Kurganov and E. Tadmor, J. Comput. Phys. 160, 241 (2000).
- [13] K. Kajantie, J. I. Kapusta, L. D. McLerran, and A. Mekjian, Phys. Rev. D 34, 2746 (1986).
- [14] P. Petreczky, J. Phys. G 39, 093002 (2012).
- [15] E. Braaten, R. D. Pisarski, and T.-C. Yuan, Phys. Rev. Lett. 64, 2242 (1990).
- [16] P. Aurenche, F. Gelis, G. D. Moore, and H. Zaraket, JHEP 12 (2002) 006.
- [17] H.-T. Ding, A. Francis, O. Kaczmarek, F. Karsch, E. Laermann, and W. Soeldner, Phys. Rev. D 83, 034504 (2011).
- [18] R. Rapp, arXiv:1304.2309 [hep-ph].
- [19] K. Dusling, G. D. Moore, and D. Teaney, Phys. Rev. C 81, 034907 (2010).
- [20] K. Dusling and S. Lin, Nucl. Phys. A 809, 246 (2008).
- [21] G. J. Gounaris and J. J. Sakurai, Phys. Rev. Lett. 21, 244 (1968).
- [22] J. I. Kapusta and C. Gale, *Finite-Temperature Field Theory: Principles and Applications* (Cambridge University Press, Cambridge, UK, 2006).

- [23] R. Rapp, Phys. Rev. C 63, 054907 (2001).
- [24] R. Rapp and J. Wambach, Adv. Nucl. Phys. 25, 1 (2002);
   R. Rapp, J. Wambach, and H. van Hees, arXiv:0901.3289 [hep-ph].
- [25] V. L. Eletsky, M. Belkacem, P. J. Ellis, and J. I. Kapusta, Phys. Rev. C 64, 035202 (2001).
- [26] A. T. Martell and P. J. Ellis, Phys. Rev. C 69, 065206 (2004).
- [27] G. Vujanovic, J. Ruppert, and C. Gale, Phys. Rev. C 80, 044907 (2009).
- [28] S. Jeon and P. J. Ellis, Phys. Rev. D 58, 045013 (1998).
- [29] A. Monnai and T. Hirano, Phys. Rev. C 80, 054906 (2009).
- [30] J. Ruppert, C. Gale, T. Renk, P. Lichard, and J. I. Kapusta, Phys. Rev. Lett. 100, 162301 (2008).
- [31] J. Zhao (STAR Collaboration), J. Phys. G 38, 124134 (2011).
- [32] E. V. Shuryak, Phys. Rev. C 55, 961 (1997).
- [33] H. van Hees, V. Greco, and R. Rapp, Phys. Rev. C 73, 034913 (2006); S. Caron-Huot and G. D. Moore, Phys. Rev. Lett. 100, 052301 (2008); P. Gossiaux, J. Aichelin, and T. Gousset, Prog. Theor. Phys. Suppl. 193, 110 (2012); M. Younus, C. E. Coleman-Smith, S. A. Bass, and D. K. Srivastava, arXiv:1309.1276 [nucl-th].
- [34] G. D. Moore and D. Teaney, Phys. Rev. C 71, 064904 (2005).
- [35] M. Cacciari, M. Greco, and P. Nason, JHEP 05 (1998) 007.
- [36] S. Caron-Huot and G. D. Moore, JHEP 02 (2008) 081.
- [37] C. Peterson, D. Schlatter, I. Schmitt, and P. M. Zerwas, Phys. Rev. D 27, 105 (1983).
- [38] H. Song and U. W. Heinz, Phys. Rev. C 77, 064901 (2008).
- [39] R. Chatterjee, D. K. Srivastava, U. Heinz, and C. Gale, Phys. Rev. C 75, 054909 (2007).
- [40] J. Deng, Q. Wang, N. Xu, and P. Zhuang, Phys. Lett. B 701, 581 (2011).
- [41] P. Mohanty, V. Roy, S. Ghosh, S. K. Das, B. Mohanty *et al.*, Phys. Rev. C 85, 031903 (2012).
- [42] B. Schenke, C. Gale, and S. Jeon, Phys. Rev. C 80, 054913 (2009).
- [43] C. Young, B. Schenke, S. Jeon, and C. Gale, Phys. Rev. C 86, 034905 (2012).
- [44] L. Ruan (STAR Collaboration), arXiv:1207.7043 [nucl-ex].
- [45] D. Teaney, Phys. Rev. C 68, 034913 (2003).