Description of the isotope chain ^{180–196}Pt within several solvable approaches

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Energies of the ground, β and γ bands as well as the associated B(E2) values are determined for each even-even isotope of the ^{180–196}Pt chain by the exact solutions of some differential equations which approximate the generalized Bohr-Mottelson Hamiltonian. The emerging approaches are called the sextic and spheroidal approach (SSA), the sextic and Mathieu approach (SMA), the infinite square well and spheroidal approach (ISWSA), and the infinite square well and Mathieu approach (ISWMA). While the first three methods were formulated in some earlier papers, ISWMA is an unedited approach of this work. Numerical results are compared with those obtained with the so-called X(5) and Z(5) models. A contour plot for the probability density as function of the intrinsic dynamic deformations is given for a few states of the three considered bands with the aim of evidencing the shape evolution along the isotope chain and pointing out possible shape coexistence.

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I. INTRODUCTION

Since the critical point symmetries [1-5] of the nuclear shape phase transitions were proposed, many experimental and theoretical efforts were made to find the nuclei described by the new symmetries. While at the beginning the X(5)[3] candidates were found in the mass region of $A \approx 150$ [6-8], recently a new region has been suggested for Os and Pt isotopes with $A \approx 180$ [9,10]. In Refs. [11,12] data for the isotopes ^{176,178,180,188,190,192}Os were analyzed with the sextic and spheroidal approach (SSA) [11], the Davidson and spheroidal approach (DSA) [12], and the infinite square well and spheroidal approach (ISWSA) [13], and the results were compared with those of the coherent state model (CSM) [14] and X(5). According to our analysis these isotopes present features for the U(5) \rightarrow SU(3) shape phase transition with the critical point reached for ¹⁷⁶Os and ¹⁸⁸Os. On the other hand, by applying the sextic and Mathieu approach (SMA) [15] to ^{188,190,192}Os, we see that the isotope ¹⁹²Os is a good candidate for the critical point of the phase transition between the prolate and the oblate shapes through the triaxial shape corresponding to $\gamma_0 = 30^{\circ}$.

Encouraged by the results for the Os isotopes, we consider the above-mentioned models also for the even-even ^{180–196}Pt isotopes. We aim not only at determining the energy spectra and the electric transition probabilities but also at showing the static deformation of each isotope in both the ground and excited states. Features like shape coexistence or a transition from the prolate to oblate shapes through a triaxial deformation are expected to show up. Keeping in mind that the SMA, the ISWMA, and the Z(5) [5] are suitable for the description of the triaxial nuclei lying close to $\gamma_0 = 30^\circ$, a comparison of their predictions represents a challenging task. ISWMA is the unedited model proposed in this paper.

Recently, in Ref. [10] it was shown that the isotope ¹⁸²Pt has some of the X(5) features. According to the interacting boson model-1 (IBM-1) [16] and the general collective model [17], this isotope manifests shape coexistence and is close to the critical point of the U(5) \rightarrow SU(3) shape phase transition.

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Evidence for shape coexistence was also presented for ^{176,178}Pt [18,19], ¹⁸⁴Pt [20], ¹⁸⁶Pt [21], and ¹⁸⁸Pt [22], which suggests that this behavior is a specific feature for Pt isotopes. Some investigations where the ground-state shape evolution in Pt isotope chain from the prolate towards the oblate shapes were performed in Refs. [23,24].

The objectives formulated above are achieved according to the following plan. In Sec. II, a short presentation of the formalisms used for the description of the Pt even-even isotopes is given. Numerical results and their comparison with the corresponding experimental data are discussed in Sec. III. The final conclusions are drawn in Sec. IV.

II. SHORT PRESENTATION OF THE MODELS

The formalisms X(5), Z(5), ISWSA, ISWMA, SSA, and SMA are derived by a set of approximations applied to the Bohr-Mottelson Hamiltonian [1],

$$H = -\frac{\hbar^2}{2B} \left[\frac{1}{\beta^4} \frac{\partial}{\partial \beta} \beta^4 \frac{\partial}{\partial \beta} + \frac{1}{\beta^2 \sin 3\gamma} \frac{\partial}{\partial \gamma} \sin 3\gamma \frac{\partial}{\partial \gamma} - \frac{1}{4\beta^2} \sum_{k=1}^3 \frac{\hat{Q}_k^2}{\sin^2 \left(\gamma - \frac{2\pi}{3}k\right)} \right] + C \frac{\beta^2}{2}, \quad (2.1)$$

amended with a potential [25,26]

$$V(\beta,\gamma) = V_1(\beta) + \frac{V_2(\gamma)}{\beta^2}.$$
 (2.2)

The form of the β and γ potential allows us to separate the β variable from the γ and the three Euler angles θ_1 , θ_2 , and θ_3 . Here, \hat{Q}_k 's denote the angular momentum components in the intrinsic reference frame. A full separation, however, may be achieved by expanding the rotor term in power series of γ around either of $\gamma_0 = 0$ or of $\gamma_0 = \pi/6$ and, moreover, by replacing the factor β^2 multiplying the γ -dependent term with

its average value, denoted hereafter by $\langle \beta^2 \rangle$. The resulting equations are

$$\left[-\frac{1}{\beta^4}\frac{d}{d\beta}\beta^4\frac{d}{d\beta} + \frac{\Lambda}{\beta^2} + v_1(\beta)\right]f(\beta) = \varepsilon_\beta f(\beta), \quad (2.3)$$
$$\left[-\frac{1}{\sin 3\gamma}\frac{d}{d\gamma}\sin 3\gamma\frac{d}{d\gamma} - W + v_2(\gamma)\right]\phi(\gamma) = \tilde{\varepsilon}_\gamma\phi(\gamma), \quad (2.4)$$

where the following notations are used:

$$v_1(\beta) = \frac{2B}{\hbar^2} V_1(\beta), \quad v_2(\gamma) = \frac{2B}{\hbar^2} V_2(\gamma),$$
$$\varepsilon_\beta = \frac{2B}{\hbar^2} E_\beta, \quad \tilde{\varepsilon}_\gamma = \langle \beta^2 \rangle \frac{2B}{\hbar^2} E_\gamma. \tag{2.5}$$

 Λ and W are the contributions coming from the rotor term and their expressions depend on the order of the γ series truncation.

For the sake of fixing the notations and defining the main ingredients, in what follows the above-mentioned approaches will be briefly described. For details we advise the reader to consult Refs. [3,5,11–13,15,25,26]. In Eq. (2.3), the X(5), Z(5), ISWSA, and ISWMA models use a common potential in β , namely an infinite square well:

$$v_1(\beta) = \begin{cases} 0, & \beta \leqslant \beta_{\omega} \\ \infty, & \beta > \beta_{\omega}. \end{cases}$$
(2.6)

With such a choice Eq. (2.3) admits the Bessel functions of irrational order ν , as solutions:

$$f_{s,L}(\beta) = C_{s,L}\beta^{-\frac{3}{2}}J_{\nu}\left(\frac{x_{s,L}}{\beta_{\omega}}\beta\right), \quad s = 1, 2, 3, \dots$$
 (2.7)

 $C_{s,L}$ denotes the normalization factor, $x_{s,L}$ are the Bessel function zeros, while L is the total intrinsic angular momentum.

By contrast the SSA and SMA use in Eq. (2.3) a sextic oscillator plus a centrifugal barrier potential [27],

$$v_{1}^{\pm}(\beta) = (b^{2} - 4ac^{\pm})\beta^{2} + 2ab\beta^{4} + a^{2}\beta^{6} + u_{0}^{\pm},$$

$$c^{\pm} = \frac{L}{2} + \frac{5}{4} + M, \quad M = 0, 1, 2, \dots$$
(2.8)

Here, c^{\pm} has two different values, one for L even and other for L odd, while u_0^{\pm} are constants which are fixed such that the two potentials v_1^+ and v_1^- have the same minimum energy. Equation (2.3), with $\Lambda = L(L+1) - 2$ and the potential given by Eq. (2.8), is quasiexactly solved, the solutions being of the form

$$\varphi_{n_{\beta},L}^{(M)}(\beta) = N_{n_{\beta},L} P_{n_{\beta},L}^{(M)}(\beta^2) \beta^{L+1} e^{-\frac{a}{4}\beta^4 - \frac{b}{2}\beta^2},$$

$$n_{\beta} = 0, 1, 2, \dots, M,$$
 (2.9)

where $N_{n_{\beta},L}$ is the normalization factor, while $P_{n_{\beta},L}^{(M)}(\beta^2)$ are polynomials of order n_{β} in β^2 .

Concerning Eq. (2.4), X(5) and Z(5) chose an oscillator and a shifted oscillator potential, respectively:

$$v_2(\gamma) = c \frac{1}{2} (\gamma - \gamma_0)^2.$$
 (2.10)

Indeed, for X(5) $\gamma_0 = 0$ and the solutions of Eq. (2.4) are the generalized Laguerre polynomials, L_n^m :

$$\eta_{n_{\gamma},K}(\gamma) = C_{n,K} \gamma^{|K/2|} e^{-(3a)\gamma^2/2} L_n^{|K|} (3a\gamma^2),$$

$$n = \left(\frac{n_{\gamma} - |K|}{2}\right), \quad a = \frac{\sqrt{c}}{3}, \quad n_{\gamma} = 0, 1, 2, \dots$$
(2.11)

The quantum number K corresponds to the angular momentum projection on the intrinsic z axis. As for Z(5), $\gamma_0 = \pi/6$ and the corresponding equation (2.4) is obeyed by the Hermite polynomials H_n :

$$\eta_{\bar{n}_{\gamma}} = N_{\bar{n}_{\gamma}} H_{\bar{n}_{\gamma}} [b(\gamma - \pi/6)] e^{-b^2(\gamma - \pi/6)/2},$$

$$b = \left(\frac{c\langle\beta^2\rangle}{2}\right)^{1/4}, \quad \bar{n}_{\gamma} = 0, 1, 2, \dots$$
(2.12)

. 2 .

Both models, the X(5) and the Z(5), consider in Eq. (2.4) a zeroth order of approximation for the rotor term.

This is not the case for the ISWSA, ISWMA, SSA, and SMA models, where a second-order power expansion of both the rotor term and the periodic potential

$$v_2(\gamma) = u_1 \cos 3\gamma + u_2 \cos^2 3\gamma$$
 (2.13)

is used, which results in having the spheroidal $(S_{m,n})$ and Mathieu (\mathcal{M}_n) functions as solutions of the resulting differential equations, respectively:

$$\eta(\gamma) = S_{m,n}(\cos 3\gamma; c), \quad \eta(\gamma) = \frac{\mathcal{M}_n(3\gamma; q)}{\sqrt{|\sin 3\gamma|}}.$$
 (2.14)

The expressions of c and q are specified below.

The advantages of the Mathieu and spheroidal functions are that they are periodic, defined on a bound interval, and normalized to unity with the integration measure of $|\sin 3\gamma| d\gamma$, preserving in this way the hermiticity of the initial Hamiltonian. Note that the other approaches do not satisfy these conditions.

The total energy of the system is obtained by summing the eigenvalues of the β and γ equations:

$$\varepsilon = \varepsilon_{\beta} + \tilde{\varepsilon}_{\gamma}. \tag{2.15}$$

The excitation energies yielded by the formalisms used in the present paper are as follows:

X(5):
$$E(s,L,n_{\gamma},K) - E(1,0,0,0) = B_1(x_{s,L}^2 - x_{1,0}^2) + \delta_{K,2}X,$$

 $X = A_1 + 4C_1,$ (2.16)

with A_1 , B_1 , and C_1 arbitrary parameters. In our calcultions the parameter X is fitted.

Z(5):
$$E(s,L,n_{\gamma}=0,R) - E(1,0,0,0) = B_1(x_{s,L,R}^2 - x_{1,0,0}^2), \quad B_1 = \frac{1}{\beta_{\omega}^2} \frac{\hbar^2}{2B},$$
 (2.17)

ISWSA:
$$E(s, n_{\gamma}, m_{\gamma}, L, K) = B_1 x_{s,L}^2 + F \left[9\lambda_{m_{\gamma}, n_{\gamma}}(c) + \frac{u_1}{2} + \frac{11}{27}D - \frac{L(L+1)}{3} \right],$$

 $\lambda_{m_{\gamma}, n_{\gamma}} = \frac{1}{9} \left[\tilde{\varepsilon}_{\gamma} - \frac{u_1}{2} - \frac{11}{27}D + \frac{1}{3}L(L+1) \right], \quad c^2 = \frac{1}{9} \left(\frac{u_1}{2} + u_2 - \frac{2}{27}D \right),$
 $m_{\gamma} = \frac{K}{2}, \quad D = L(L+1) - K^2 - 2, \quad F = \frac{1}{\langle \beta^2 \rangle} \frac{\hbar^2}{2B},$
(2.18)

ISWMA:
$$E(s,n_{\gamma},L,R) = B_1 x_{s,L}^2 + F\left[9a_{n_{\gamma}}(L,R) + 18q(L,R) - \frac{3}{4}R^2 - \frac{5}{2}\right],$$

 $q = \frac{1}{36}\left(\frac{10}{9}L(L+1) - \frac{13}{12}R^2 + u_1 - \frac{9}{4}\right), \quad a_{n_{\gamma}} = \frac{1}{9}\left(\tilde{\varepsilon}_{\gamma} + \frac{3}{4}R^2 + \frac{5}{2}\right) - 2q,$
(2.19)

SSA:
$$E(n_{\beta}, n_{\gamma}, m_{\gamma}, L, K) = G[b(2L+3) + \lambda_{n_{\beta}}^{(M)} + u_0^{\pm}] + F\left[9\lambda_{m_{\gamma}, n_{\gamma}}(c) + \frac{u_1}{2} + \frac{11}{27}D - L(L+1)\right],$$

$$\lambda_{m_{\gamma},n_{\gamma}} = \frac{1}{9} \left[\tilde{\varepsilon}_{\gamma} - \frac{u_1}{2} - \frac{11}{27}D + \frac{1}{3}L(L+1) \right] + \frac{2L(L+1)}{27}, \quad G = \frac{\hbar^2}{2B},$$
(2.20)

SMA:
$$E(n_{\beta}, n_{\gamma}, L, R) = G[4bs(L) + \lambda_{n_{\beta}}^{(M)}(L) + u_{0}^{\pi}] + F\left[9a_{n_{\gamma}}(L, R) + 18q(L, R) - \frac{3}{4}R^{2} - \frac{5}{2}\right],$$
 (2.21)

where $\lambda_{n_{\beta}}^{(M)}(L)$ satisfy the equation

$$\left[-\left(\frac{d^2}{d\beta^2} + \frac{4s-1}{\beta}\frac{d}{d\beta}\right) + 2b\beta\frac{d}{d\beta} + 2a\beta^2\left(\beta\frac{d}{d\beta} - 2M\right)\right]P^{(M)}_{n_{\beta},L}(\beta^2) = \lambda^{(M)}_{n_{\beta}}P^{(M)}_{n_{\beta},L}(\beta^2).$$
(2.22)

The specific β and γ potentials of the six approaches used in the present paper are collected, for comparison, in Table I. The potentials in the β variable are to be amended by a centrifugal term due to the rotor component of the starting Hamiltonian.

The reduced *E*2 transition probabilities for ISWSA and SSA are determined with the reduced matrix element of the transition operator,

$$T_{2\mu}^{(E2)} = t_1 \beta \left[\cos \gamma D_{\mu 0}^2 + \frac{\sin \gamma}{\sqrt{2}} \left(D_{\mu 2}^2 + D_{\mu,-2}^2 \right) \right] + t_2 \sqrt{\frac{2}{7}} \beta^2 \left[-\cos 2\gamma D_{\mu 0}^2 + \frac{\sin 2\gamma}{\sqrt{2}} \left(D_{\mu 2}^2 + D_{\mu,-2}^2 \right) \right],$$
(2.23)

Approach	β potential	γ potential
X(5)	0, for $\beta \leqslant \beta_{\omega}$;	$\frac{c}{2}\gamma^2$.
	∞ , for $\beta > \beta_{\omega}$.	
Z(5)	0, for $\beta \leqslant \beta_{\omega}$;	$\frac{c}{2}(\gamma-\frac{\pi}{6})^2$.
	∞ , for $\beta > \beta_{\omega}$.	_ *
ISWSA	0, for $\beta \leqslant \beta_{\omega}$;	$u_1\cos 3\gamma + u_2\cos^2 3\gamma + \frac{9}{4\sin^2 3\gamma}.$
	∞ , for $\beta > \beta_{\omega}$.	· · · · · · · · · · · · · · · · · · ·
SSA	$(b^2 - 4ac^{\pm})\beta^2 + 2ab\beta^4 + a^2\beta^6 + u_0^{\pm},$	$u_1 \cos 3\gamma + u_2 \cos^2 3\gamma + \frac{9}{4 \sin^2 2\gamma}$.
	$c^{\pm} = \frac{L}{2} + \frac{5}{4} + m; \ m = 0, 1, 2, \dots$	+ SIII <i>3 y</i>
ISWMA	0 for $\beta \leq \beta_{\omega}$;	$-2q\cos 6\gamma;$
	∞ , for $\beta > \beta_{\omega}$.	$q = \frac{1}{2\epsilon} \left(\frac{10}{9} L(L+1) - \frac{13}{12} R^2 + \mu - \frac{9}{4} \right).$
SMA	$(b^2 - 4ac^{\pm})\beta^2 + 2ab\beta^4 + a^2\beta^6 + u_0^{\pm},$	$-2q\cos 6\gamma;$
	$c^{\pm} = \frac{L}{2} + \frac{5}{4} + m; \ m = 0, 1, 2, \dots$	$q = \frac{1}{36} \left(\frac{10}{9} L(L+1) - \frac{13}{12} R^2 + \mu - \frac{9}{4} \right).$

TABLE I. Here we list the β and γ potentials used by the approaches ISWSA, SSA, ISWMA, and SMA. For comparison the potentials characterizing X(5) and Z(5) are also given.

between the corresponding initial $|L_i M_i\rangle$ and final $|L_f M_f\rangle$ states, as described above:

$$B(E2; L_i \to L_f) = |\langle L_i || T_2^{(E2)} || L_f \rangle|^2.$$
 (2.24)

Here Rose's convention [28] was used for the reduced matrix elements. For SMA, ISWMA, and Z(5), in the expression of the transition operator (2.23) γ is substituted with $\gamma - 2\pi/3$. The argument is justified by the fact that $\gamma - 2\pi/3$ defines the axis 1 of the principal inertial ellipsoid. Indeed, the transformation from the laboratory to the intrinsic frame is a rotation defined by the matrix D_{MR}^L , where *M* and *R* are eigenvalues of the operator \hat{Q}_1 . The X(5) and Z(5) models keep only the zeroth-order approximation of the first γ term in the transition operator (2.23).

III. NUMERICAL RESULTS

The formalisms presented in Sec. II were applied to some even-even isotopes of Pt: ^{180–196}Pt. It is commonly accepted that nuclear spectra can be classified by the values of the energy ratios:

$$R_{4_{g}^{+}/2_{g}^{+}} = \frac{E_{4_{g}^{+}} - E_{0_{g}^{+}}}{E_{2_{g}^{+}} - E_{0_{g}^{+}}}, \quad R_{0_{\beta}^{+}/2_{g}^{+}} = \frac{E_{0_{\beta}^{+}} - E_{0_{g}^{+}}}{E_{2_{g}^{+}} - E_{0_{g}^{+}}}.$$
 (3.1)

Moreover, it seems that nuclei satisfying a certain symmetry are characterized by almost constant ratios. The values of these ratios associated to the isotopes considered here are collected in Table II. As seen from there, the ratios $R_{4_g^+/2_g^+}$ for ^{180,182,184}Pt are close to that predicted by the X(5) approach. By contrast the other ratio $R_{0_{\beta}^+/2_g^+}$ indicates that these isotopes are far from the ideal picture of X(5). As a matter of fact this feature is consistent with the results of Ref. [29] that not all nuclear properties reach the critical point in a phase transition in the same isotope. We apply the approaches ISWSA and SSA to the mentioned isotopes in order to test their ability to account for these complementary features.

Concerning the description called Z(5) this is appropriate for ^{190,192,194,196}Pt, the statement being supported by the values of both ratios. Indeed, the detailed numerical analysis of Ref. [5] shows a good agreement between calculations and experimental data. In this context the application of the ISWMA and SMA to these isotopes will provide a sensible comparison of the formalisms on one hand and Z(5) on the other hand.

It is well known that the triaxial rigid rotor (TRR) predicts [30] a relation between the first three excited state energies:

$$\Delta E \equiv E_{3^+_{\gamma}} - E_{2^+_{\gamma}} - E_{2^+_g} = 0. \tag{3.2}$$

Due to this fact, the above equation is considered to be a signature for a triaxial deformation of $\gamma_0 = 30^\circ$. For the

isotope ¹⁹²Pt the above equation reads $|\Delta E| = 8$ keV, which means that the mentioned isotope is close to the ideal triaxial rigid rotor. Considering this isotope among the treated isotopes allows us to answer the question of whether these approaches are suitable for the description of the triaxial nuclei. The isotopes ^{186,188,190,192,194,196}Pt may be considered to be γ -unstable nuclei, having the ratio $R_{4a^+/2a^+}$ close to 2.5. A special case is that of ¹⁸⁶Pt, which has the head state of γ band higher in energy than the first β state, which results in claiming a γ stable picture. Most likely this nucleus exhibits the main features for the critical point of the phase transition from prolate to oblate shapes. Due to the specific structure of their potentials in the γ variable, ISWSA and SSA seem to be suitable to describe both the γ unstable and γ -stable deformed nuclei. Actually this argument justifies including ¹⁸⁶Pt and ¹⁸⁸Pt on the list of considered isotopes. In addition to the prolate-oblate transition along the Pt isotopic chain an alternative prolate-oblate transition has been considered in Ref. [31], with both transitions studied in Ref. [32].

Each approach involves a number of free parameters for energies as well as for B(E2) values. These are fixed by fitting some particular experimental data concerning either the excitation energies or the reduced transition probabilities. The results of the fitting procedure adopted are listed in Tables III and IV. As seen from these tables, the number of parameters used for fitting the spectra in X(5), ISWSA, and SSA are 2, 4, and 6, respectively, while in the fitting of B(E2) values 1, 2, and 2 parameters are used, respectively. Also, from Table III we notice that the number of parameters used for fitting the spectra in Z(5), ISWMA, and SMA are 1, 3, and 5, respectively, while in fitting the B(E2)s 1, 2, and 2 parameters are used, respectively.

Numerical results for the excitation energies of the ground, β , and γ bands, as well as for the quadrupole electric transitions between states of these bands, are compared with the corresponding experimental data in Tables V and VI, respectively. For each formalism the agreement between the calculation results and the corresponding experimental data is quantitatively appraised by the rms values of the deviations.

From Table V, one can see that spectra of the isotopes ¹⁸⁰Pt, ¹⁸²Pt, and ¹⁸⁴Pt are better described by SSA and ISWSA than by X(5). The best approach seems to be SSA. Moreover, the X(5) failure in explaining the data from the β band is removed by SSA, and that happens especially for ¹⁸²Pt. Concerning the γ band, all three formalisms, SSA, ISWSA, and X(5), encounter difficulties in explaining the bandhead energy. A possible explanation would be that the state 2^+_{γ} does not actually belong to the γ band. In this context we mention the fact that two alternative interpretations have been

TABLE II. Signatures of X(5), Z(5), and O(6) symmetries identified in the even-even isotopes $^{180-196}$ Pt. The two ratios are defined by Eq. (3.1).

	¹⁸⁰ Pt	¹⁸² Pt	¹⁸⁴ Pt	¹⁸⁶ Pt	¹⁸⁸ Pt	¹⁹⁰ Pt	¹⁹² Pt	¹⁹⁴ Pt	¹⁹⁶ Pt	X(5)	Z(5)	O(6)
$R_{4^{+}_{a}/2^{+}_{a}}$	2.69	2.71	2.67	2.55	2.52	2.49	2.48	2.47	2.46	2.90	2.35	2.50
$R_{0^+_\beta/2^+_g}$	3.12	3.23	3.02	2.46	3.00	3.11	3.78	3.86	3.19	5.65	3.91	

TABLE III. The parameters characterizing the X(5), ISWSA, and SSA approaches, determined by a fitting procedure, are listed for ^{180–188}Pt isotopes.

Nucl	B_1	(keV)	X (keV)	<i>F</i> (k	eV)	<i>u</i> 1	1	ι	<i>u</i> ₂	G (keV)	а	b	t_1	(W.u.) ^{1/}	2	<i>t</i> ₂ (W	$(.u.)^{1/2}$
	X(5)	ISWSA	X(5)	ISWSA	SSA	ISWSA	SSA	ISWSA	SSA	SSA	SSA	SSA	X(5)	ISWSA	SSA	ISWSA	SSA
¹⁸⁰ Pt	19.08	16.38	722.5	17.32	3.34	-0.15	-821.2	-104.6	-1000	1.04	1059	135	500.2	614.4	1750	0.0	0.0
¹⁸² Pt	18.02	16.39	720.7	11.33	5.33	-31.56	-1042	-163	-0.0007	0.81	1687	186	451.2	2200	6561	9062	89567
¹⁸⁴ Pt	17.28	16.83	739.7	3.35	6.25	-1000	-302.6	-1000	-262	0.62	3030	256	419.6	2422	7821	11331	122065
¹⁸⁶ Pt		16.25		16.82	3.08	-253.87	1471	6.75	-2326	0.85	1296	170		1728	5061	5978	58515
¹⁸⁸ Pt		21.5		41.99	14.55	-97.45	-466.2	81.07	165.8	1.45	1449	95		517.4	1717	0.0	0.0

studied in Ref. [42] with a related description appearing in Ref. [43]. A similar situation is met in ¹⁸⁶Pt. In ¹⁸⁸Pt, however, all three bands considered here are realistically described by SSA.

The comparison of the numerical results yielded by SMA, ISWMA, and Z(5) with experimental data for the even-even isotopes $^{190-196}$ Pt, is made also in Table V, with the result in favor of SMA and ISWMA.

The electromagnetic transition probabilities, calculated with Eq. (2.24), are included in Table VI. By analyzing the rms values for each model, one may conclude that SSA and ISWSA describe the experimental data batter than X(5), while SMA and ISWMA describe the data better than Z(5). An explanation for this could be that X(5) and Z(5) use only the zeroth-order approximation of the harmonic part of the transition operator (2.23). Indeed, as shown in Table VI, for ¹⁸⁰Pt the results obtained by SSA and ISWSA using only the harmonic transition operator are almost identical with those of X(5). By contrast for 182,184 Pt, where the anharmonic contributions were included, the results of SSA and ISWSA are better than those of X(5). It is worth noticing that the rms associated to Z(5) for ¹⁹²Pt and ¹⁹⁶Pt are smaller than those provided by ISWMA. This situation might be caused by the fact the two approaches considered for the γ band have different descriptions. Indeed, in the framework of Z(5)the states of γ band are characterized by $n_{\gamma} = 0$, while the ISWMA γ states have $n_{\gamma} = 1$.

In Table VII we list the results for branching ratios of few states from the γ and β bands obtained by SSA, ISWSA, SMA, and ISWMA approaches, respectively. They are compared with the experimental data of Ref. [40]. For ^{190,192,194}Pt we list also the results yielded by the Z(5) formalism. The parameters determining the transition operator were fixed as follows: For

¹⁸⁸Pt and ¹⁹⁰Pt we kept t_1 as given in Tables III and IV, respectively, while t_2 was fixed by a least square procedure. The results for t_2 are also listed in Table VII. As for the rest of isotopes from the table, the parameters t_1 , t_2 are as listed in Tables III and IV.

Another objective of the present work is to determine the isotope shape in ground and excited states, within both the SSA and the SMA. Indeed, it is interesting to see how the shape changes when one passes from one isotope to another and moreover whether this picture is state dependent. We expect to visualize the shape phase transition and also possible shape coexistence. The static shape is defined by the values of the intrinsic variables β and γ for which the probability density (the probability in the volume unit of $d\beta d\gamma$),

$$P(\beta,\gamma) = |f(\beta)\phi(\gamma)|^2 \beta^4 |\sin 3\gamma|, \qquad (3.3)$$

reaches a maximum value. In Fig. 1, the contour plots are represented in the coordinates $(\beta \cos \gamma, \beta \sin \gamma)$. In order to save the space we chose two representatives for SSA, ¹⁸⁰Pt and ¹⁸⁸Pt, and one for SMA, ¹⁹⁰Pt. Indeed, the graphs corresponding to $^{182-186}$ Pt are similar to that of 180 Pt and those of ^{192–196}Pt resemble that of ¹⁹⁰Pt. We may ask ourselves why we make such plots once we know that the power expansion in γ was performed around $\gamma = 0^{\circ}$ and $\gamma = 30^{\circ}$. We notice that the density maxima are met not in the same point where the potential is minimum. The reason is that the density accounts also for the kinetic energy and moreover includes a factor defining the measure of the integration in the β and γ coordinates. These figures reflect the structure of the wave functions. Indeed, since the γ -dependent function depends on $\cos 3\gamma$ and the spheroidal functions are symmetric with respect to the space reflection transformation, the graphs exhibit the symmetry $\gamma \rightarrow \pi/3 - \gamma$. Concerning SMA, the

TABLE IV. The parameters characterizing the Z(5), ISWMA, and SMA approaches, determined by a fitting procedure, are listed for ^{190–196}Pt isotopes.

Nucl	B_1	(keV)	F (ke	eV)	<i>u</i>	l	G (keV)	а	b		t_1 (W.u.) ^{1/2}	2	<i>t</i> ₂ (W.u	l.) ^{1/2}
	Z(5)	ISWMA	ISWMA	SMA	ISWMA	SMA	ISWMA	SMA	SMA	Z(5)	ISWMA	SMA	ISWMA	SMA
¹⁹⁰ Pt	28.12	16.73	12.82	8.14	26.67	104.6	1.11	3014.12	84.00	27.49	28.14	96.38	0.00	0.0
¹⁹² Pt	29.45	17.84	13.98	7.87	9.49	121.8	2.95	616.5	22.98	23.94	24.51	55.10	102.4	1048
¹⁹⁴ Pt ¹⁹⁶ Pt	32.65 31.49	19.87 18.27	18.43 9.98	14.68 6.48	5.00 56.53	32.74 177.1	2.96 0.41	733.0 28322	33.05 250	18.76 20.77	16.94 19.79	43.42 130.2	137.6 172.9	968.6 7708

the		2(5)	318	747	266	868	548										244	810	530	358	280		584	325	405	473	245	:183	134	1961	084	274
⁸⁸ Pt a		/MA 2	4	54	466 I	227 1	01 2										22 1	288 1	45 2	l65 3	322 4		09	15	1 063	60 1	51 2	328 2	928 3	211 2	94 4	35
r ^{180–1} ions	¹⁹⁶ Pt	IS ISW	5 3	8	[4]	15 22	25 31										8 7	28 12	38 21	96 31	6 9 43		4	1 9	30 12	13 15	98 20	36 23	70 25	51 32	H 36	-
(5) for deviat		kp SM	56 25	77 74	26 141	53 221	44 312										35 94	62 142	213	299	396		89 72	15 95	93 128	10 154	07 200	228	50 287	315	382	62
nd X((5) E:	329 35	74 87	313 15	936 22	642 30										289 11	877 13	623	482	438		505 68	356 10	456 12	527 16	327 20	263	250 27	070	234	157
SA, a predi		'MA Z	4£	35	35 1	20	83 2										85 1	92 1	72 2	83 3	98 4		38	3 60	78 1	90 1	14 2	60 2	30 3	11 3	65 4	20
ISW of the	194 Pt	A ISW	3.	30	17 14	31 21	9 28										50 73	23 13	28 22	55 32	94 43		7 6	8	34 13	2 15	0 22	H6 23)4 31	5 32	5 36	-
SSA. lues c		xp SS	28 25	11 72	112 132	00 208	348 289										67 115	512 162	232	316	400		22 62	23 86	229 128	1499 149	200	22	30(300	390	66
by the ms va		Z(5) E	297 3	698 8	184 1	[747 2]	2383 28										163 12	693 1	2366	3140	1003		546 6	772 9	313 12	377 12	6603	2041	1663	6913	819	193
The r	L.	WMA	303	772	346]	010	759 2										705	254]	062	005	055 4		552	198	201	418	953 2	134 2	792 2	938 2	371 3	158
γ, yie	¹⁹² P	SA IS'	4	47	247 1	979 2	820 2										108	489 1	117 2	903 3	807 4		89		184 1	418 1	865 1	106 2	678 2	914 2	597 3	76
: <i>g</i> , <i>β</i> , s. [33		Exp S	317 2	785 6	365 12	2018 19	2729 28										195 1	439 1	5	5	3		612 6	921 8	201 1	482 1	869 18	2113 21	2591 20	5	35	
h i = n Ref		Z(5)	284	667	1130 1	1668 2	2276 2										1110 1	1617 1	2259	2999	3822	4724	521	737	1254 1	1315 1	2004 1	1949 2	2799 2	2644	3647	218
r ⁺ wit)t	WMA	282	721	1259	1885	2591										661	1173	1931	2815	3803	4885	581	812	1183	1391	1882	2062	2665	2816	3222	67
ates J a take	1 ⁰⁶¹	MA IS	225	545	206	872	620										332	260	875	607	426		548	348	159	369	808	600	559	742	391	67
and st al dat		Exp S	296	737 0	1288 1	1915 1	2628 2										921 8	1203 1	1	0	ŝ		598 (917 8	1128 1	1450 1	1733 1	6	0	0	ŝ	
$1 \gamma b_i$ iment isted.		V	183	545	1045	1667	2405	3256									849	1153	1716	2446	3314	4308	723	887	1089	1325	1594	1893	2223	2583	2971	89
β , and exper- also 1	$^{188}\mathrm{Pt}$	SSA	232	645	1170	1772	2429	3127									719	1193	1802	2493	3240	4028	681	860	1098	1316	1630	1868	2241	2489	2911	45
und, iding s, are		A Exp	266	671	1185	1783	2438	3105									66L	1115					909	936	1085		1636		2247			
ne gro espon n volt	ţ	ISWS ¹	123	362	685	1080	1543	2073	2667	3325	4045	4827	5671	6575	7540		642	856	1247	1744	2325	2982	917	1027	1161	1317	1492	1687	1899	2129	2377	155
, of th e corr lectro	¹⁸⁶ F	SSA	146	426	801	3 1250	3 1757	5 2315	5 2916	5 3556	l 4229	3 4933	7 5666	4 6424	3 7205		472	743	3 1134) 1604	2135	2718	849	970	1130	3 1290) 1505	1 1693	4 1954	2163	5 2462	107
i volts vith th kilo e		5) Exp	1 192	1 490	6 878	25 134.	53 1858	38 233	78 282	73 339:	21 405	22 4788	76 559	33 646	41 7408	15	2 472	0 798	91 122	32 1600	18	32	0 607	5 957	90 992	35 136	96 147(72 180	54 200	71 228(92 254:	_
ectror ared w its of		SA X(9 12	7 35	0 65	8 100	5 14	9 19	0 24	6 30	6 37	0 44	8 51	8 598	52 68-	17 T7	5 68	8	3 129	178	17 23	1 298	98 86	2 96	37 10	0 12	1 13	57 15'	9 17	5 19	86 219	5 15
cilo el compa i in un	$^{184}\mathrm{Pt}$	A ISW	1	3 34	9 65	6 101	8 142	5 192	9 247	4 306	7 371	1	6 513	7 598	2 685	0 776	1 66	2 87	8 126	5 172	3 23(8 293	7 85	2 96	0 108	4 123	8 139	7 156	6 175	4 196	1 218	15
s of k , are (given		kp SS.	53 13	36 39	98 74	31 117	07 165	04 218	27 274	82 334	69 396	93 461	67 527	97 595	86 665	35 736	32 58	44 82	34 119	00 165	217	273	t9 81,	40 93.	28 108	07 123	63 143	31 161	186	206	235	83
n unit ⊢196Pt χ and		((5) E	126 10	366 4.	584 79	069 12	515 17	021 22	585 27	205 32	881 38	613 44	400 51	241 58	136 66	75	712 49	339 84	347 12	859 18	450	111	347 6	956 92	087 10	237 13	405 14	589 17	790	005	236	164
for ¹⁹⁰ noted	1	WSA 3	121	353	999	047 1	492 1	999 2	568 2	195 3	882 3	627 4	430 5	290 6	208 7		. 47	800	246 1	734 1	303 2	943 3	349	55	084 1	234 1	402 1	587 1	789 1	008 2	243 2	56
gies, g aches ta, der	$^{182}\mathbf{p}$	SA IS	39	112	78 (216 1	710 1	251 1	830 2	442 3	083 3	749 4	437 5	143 6	867 7		37 (3 16,	185 1	652 1	180 2	755 2	805 ×	24	079 1	236 1	446 1	630 1	886 1	088 2	382 2	47
energ upproa tal dat		Exp S	155 1	420 4	775 7	1206 1	1698 1	2242 2	2832 2	3461 3	4094 4	4729 47	5403 5	5127 6	5905 6		500 5	856 7	1240 1	1650 1	2118 2	6	868	943 9	1034 1	1306 1	1438 1	1731 1		6	6	
Z(5) a		X(5)	133	387	724	1131	1604	2139	2736	3392	4108 4	4882	5715	6605	7552		753	993	1425	1967	2593	3292	856	971	1110	1269	1447	1642	1854	2082	2326	128
Exci and expe	Pt	SWSA	125	366	693	1093	1563	2100	2702	3369	4099	4892	5748	6663	7641		649	863	1258	1760	2348	3013	858	696	1105	1263	1440	1637	1853	2087	2338	108
LE V. WMA nding	180	SSA L	126	386	749	1194	1705	2273	2891	3552	4253	4989	5757	6555	7379		590	809	1173	1632	2164	2755	840	954	1101	1258	1464	1653	1909	2122	2421	58
TABI IA IS' respo		Exp	153	411	757	1182	1674	2229	2842	3505	4253	4985	5729	6551	7434		478	861	1248	1650			677	963	1049	1315		1727		2198		
SN. cor		+.,	$^{+8}_{-8}$	4%	6_g^+	8°+ 8°	10^+_{g}	12^+_8	$^{+}_{+8}$	16_g^+	18^+_s	20_s^+	22_{g}^{+}	24_{g}^{+}	26_{g}^{+}	28_{g}^{+}	0^{eta}_{eta}	$\frac{5}{\beta}$	$^{+6}_{+ \theta}$	6^{β}_{β}	8^{eta}_{eta}	10^+_{β}	4 4	$\widetilde{\omega}_+^{\times}$	$^{4}_{+}$	5^+	$\overset{^{^{^{^{^{^{^{}}}}}}}{\mathrm{0}}$	+~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~	$^{\star}_{\varkappa^+}$	$^+_{\gamma}$	10^+_{γ}	\times

the		Z(5)	32 51	70	84				25		24	33	24	0.34	10.5	0	4.6					0		0	9	Ξ
d with		SWMA	28 48	70	c 8				20		19	28	30.8	0.0033	13.6	1	8					0.69		0.41	6	13
npare	$^{196}\mathrm{Pt}$	SMA I	34 52	72	8/				23		22	29	15.5	0.18	9.8	21	0.02					0.32		0.19	5	6
t, are cor		Exp	$40.6^{+0.2}_{-0.2}\\60^{+0.9}_{-0.9}$	73^{+4}_{-73}	/8-78				5^{+5}_{-5}	ı	29^{+6}_{-29}	49^{+13}_{-13}	$2.8^{+1.5}_{-1.5}$	$0.0025^{+0.002}_{-0.002}$	$0.13\substack{+0.12\\-0.12}$	18^{+10}_{-10}	$0.26\substack{+0.23\\-0.23}$					$0.56\substack{+0.12\\-0.17}$		$0.48\substack{+0.14\\-0.14}$	16^{+5}_{-5}	
0-196F		Z(5)	26 41	57	60 LL						19		19.55	0		0		0	42			0	6			25
for ¹⁹	Pt	SWMA	20 34	49	80 80						14		21.43			1.9		1.75	87			1.21	21			22
1 Z(5)	194]	SSAI	37	51	10 70						15		$^{4}_{4}$ 9.13			39.9		$^{4}_{4}$ 1.29	71			$\frac{7}{7}$ 0.79	16			22
A, and		Exp	$49.2^{+0.1}_{-0.1}$ 85^{+5}_{-5}	67 ⁺²¹ 50 ⁺¹⁴	34^{+9}_{-9}						$21\substack{+4\\-4}$		$0.63_{-0.1}^{+0.1}$			$8.4^{+1.9}_{-1.9}$		$0.29^{+0.0}_{-0.0}$	89^{+11}_{-11}			$0.36^{+0.0}_{-0.0}$	14			
SWM		A Z(5)	42 68	94						92			Ũ					0		0	53	0				15
AA, IS	Pt	ISWM	41 71	103						85								3.42		7.13	38					17
nd SN	192	SSA	2 49 73	98						89								04 0.93		$_{07}^{07}$ 1.74	38					14
⁸⁸ Pt a		Exp	$57.2^{+1.}_{-1.}$ 89^{+5}_{-5}	70^{+30}_{-30}						102^{+10}_{-10}								$0.55^{+0.}_{-0.}$		$0.68^{+0.0}_{-0.0}$	38^{+10}_{-10}					
. 180-1		A Z(5)	56 89	123	148 166																					
els foi	⁹⁰ Pt	ISWM	56 95	138	191 191																					
mode	-	p SMA	3 3 86	119	144																					
I X(5)		A Ex	82 56 [±] 131	162	180 205	220																				
A, and	¹⁸⁸ Pt	SSA (5 82 5 136	171	226	249																				
SWS		A Exp	82^{+1}_{-1}					_																		
SA, I	Pt	A ISWS	2 162 2 228	1 248	252 (252	3 249	240																		40
the S	186]	cp SS/	$^{+8}_{-8}$ 162 $^{+13}_{-13}$ 232	+23 254 -23 254	-29 20(-26 250	$^{+26}_{-26}$ 252	+21 243	-36 25																		36
l with].		(5) E ₃	75 113 19 188	48 289	70 294 87 304	01 255	13 225	23 201 31																		36
mined 37–41		WSA X	179 . 236 1	235 1	205 1	188 2	173 2	147 2 147 2																		49
deteri 3,35,3	¹⁸⁴ Pt	SSA IS'	176 1 238 2	243	232	193	171	129																		43
lities [21,3]		Exp 3	127^{+5}_{-5} 210^{+8}_{-8}	26^{+12}_{-12}	810^{+10}_{-10}	$[83^{+17}_{-17}]$	65^{+17}_{-17}	80^{+5}_{-17}	ĥ																	
obabi Refs.		A X(5)	86 138	171	216	232	246																			80
ion pı from	² Pt	ISWS/	166 222	224	202	189	178																			52
ransit taken	18	o SSA	$^{+7}_{-7}$ 167	¹⁸ 232	22 57 204 204	18 185	10 164																			47
l E2 t data		5) Exj	6 108 9 188 ⁺	0 284 ⁺	-1 205 6 266 ⁺	6 158 ⁺	2 113-																			6
duced		/SA X(06 10 59 16	10 21	+1 24 55 26	85 28	01 30																			9
The re	$^{180}\mathrm{Pt}$	SA ISW	10 10 68 10	02 2	55 2 P	78 28	00 3																			36 3
VI. 7 ing ex		Exp S:	53^{+15}_{-15} 1 10^{+30}_{-30} 1	≥ 50 2	0 0	2	ŝ																			
ABLE	W.u.]	+.	1 - 1 - 1 - 1 - 1 - 1 - 1 - 1 - 1 - 1 -	+_∞ +	- ⁸ +8	10^+_{g}	12^{+}_{8}	16^{+}_{s}	8+ ⁸	. <u>+</u> .×	<u>ح ±</u> -	<u>+</u> ×	+,80	+ 8	+ 60	<u>+</u> >	<u>+</u> ,>	s^+	+ 8	+ .00	+	+ .00	+.*	+ 88	+ 00	n.)
T⁄- corres	B(E2) [$\mathbf{J}_i^+ \rightarrow \mathbf{J}_{j}'$	$2^+_{g} \rightarrow 0$ $4^+_{g} \rightarrow 2$	$6_{g}^{+} \rightarrow 4_{2}$	$\frac{8_g}{10_g^+} \downarrow c$	$12^+_{g} \rightarrow$	14^+_{s}	$16^+_{s^+} \downarrow \downarrow$ $18^+_{s^-} \downarrow$	$2_B^+ \rightarrow 0$	$3^+_{\gamma} \downarrow 2^+_{\gamma}$	\downarrow^{+}_{γ}	$6^+_{\mathcal{V}} ightarrow 4$	$0^+_{\beta} \rightarrow 2$	$2^+_{\beta} \rightarrow 0$	$2^+_{\beta} \rightarrow 4$	$0^+_{\beta} \rightarrow 2$	$2^+_{\beta} \rightarrow 2$	$2^+_{\gamma} ightarrow 0$	$2^+_{\gamma} \rightarrow 2$	$3^+_{\gamma} \downarrow 2$	$3^+_{\gamma} \rightarrow 4$	\downarrow^{+4}_{χ}	$^{++}_{\gamma} \rightarrow ^{+}_{4}$	$6^+_{\gamma} ightarrow 4$	$6^+_{\gamma} ightarrow 6$	rms (W.



FIG. 1. Probability densities for the states 0_g^+ , 10_g^+ , 0_β , 10_β , 2_γ^+ , 3_γ^+ , 9_γ^+ , and 10_γ^+ of ^{180,188}Pt and ¹⁹⁰Pt, calculated with SSA and SMA, respectively. The steps used in the contour plots are 30, 10, and 20 for ¹⁸⁰Pt, ¹⁸⁸Pt, and ¹⁹⁰Pt, respectively. Exceptions are 0_β^+ for ¹⁸⁸Pt and 8_β^+ for ¹⁹⁰Pt, where the steps are 12 and 25, respectively.

$\frac{B(E2;J^+ \rightarrow J^{\prime +})}{B(E2;I^+ \rightarrow I^{\prime +})}$		¹⁸⁸ Pt			1	⁹⁰ Pt			19	⁹² Pt	¹⁹⁴ Pt					
$\times 10^2$	Exp.	SMA	ISWSA	Exp.	SMA	ISWMA	Z(5)	Exp.	SMA	ISWMA	Z(5)	Exp.	SMA	ISWMA	Z(5)	
$\frac{2^+_{\gamma} \rightarrow 0^+_g}{2^+_{\gamma} \rightarrow 2^+_g}$	3.44	63	66	1.24	1.95	4.90	0	0.51	1.96	7.55	0	0.38	1.81	2.01	0	
$\frac{3_{\gamma}^{+} \rightarrow 2_{g}^{\$}}{3_{\gamma}^{+} \rightarrow 2_{\gamma}^{\$}}$	4.5	23	17	1.8	2.2	6.0	0	0.76	1.95	8.42	0	0.5	5.37	9.04	0	
$\frac{3^+_{\gamma} \rightarrow 4^+_g}{3^+_{\gamma} \rightarrow 2^+_{\gamma}}$		9.9	7.3	49	49	49	57	26	43	45	57		128	182	57	
$\frac{0_{\beta}^{\prime+} \rightarrow 2_{g}^{\prime+}}{0_{\beta}^{+} \rightarrow 2_{\beta}^{+}}$	≥11	23.2	17	11	14	31	19	3.8	8.1	31	19	7.9	10.5	31	19	
$\frac{2^{P}_{\beta} \rightarrow 0^{P}_{g}}{2^{+}_{\beta} \rightarrow 0^{+}_{\beta}}$	0.83	0.37	3.17	0.02	0.82	0.02	1.39	0.022	1.16	0.017	1.39		1.04	0.02	1.39	
$\frac{2^{P}_{\beta} \rightarrow 4^{P}_{g}}{2^{+}_{a} \rightarrow 0^{+}_{a}}$	19	66	49	4.2	44	68	42	≤2.8	28	68	42		35	68	42	
r.m.s.		35	32		16	27	16		13	30	22		3	14	7	
$t_2 (W.u.)^{\frac{1}{2}}$		-316.8	-184.9		2931	145.2										

TABLE VII. The branching ratios for some states of the ¹⁸⁸Pt and ^{190,192,194}Pt isotopes determined with SSA and ISWSA and SMA, ISWMA, and Z(5), respectively, are compared with the corresponding experimental data taken from Ref. [44].

mentioned symmetry is caused by the fact the potential in γ is function of $\cos^2 3\gamma$. Also, the node of the β function causes a doublet maxima with the same γ . For ¹⁸⁸Pt we notice equal density curves which surround two maxima of identical β . This situation is specific to the shape coexistence. It is worth mentioning that such transition is showing up despite the fact that for all isotopes $^{180-188}$ Pt we used a power expansion in γ around 0°. That means that the transition is caused not only by the potential shape but also by the structure coefficients involved in the associated differential equations. Actually, we calculated the spectroscopic properties of Pt isotopes with $A \ge 190$ also with a power expansion in γ around $\gamma = 0$. However, the results of SMA are characterized by smaller rms values for the deviations of the predictions from the experimental data. It is interesting to note that although we changed the description when we passed from ¹⁸⁸Pt to ¹⁹⁰Pt the probability density undergoes a smooth transition. The maxima surrounded by equidensity curves merge in one maximum at $\gamma = 30^{\circ}$ for ground and β band states, while for γ band states the doublets are well separated. How this picture is modified when additional degrees of freedom like octupole [45,46] or single particle [47,48] will be analyzed elsewhere.

Note that for ¹⁹⁰Pt the considered excited state in the β band is 8⁺ and not 10⁺ as happens for other isotopes. The reason is that, as seen from Table V, the highest spin state for which energies in nuclei with $A \ge 190$, calculated with SMA, is 8⁺.

IV. CONCLUSIONS

In the previous section we described some even-even isotopes of Pt by four solvable models emerging from the generalized Bohr Mottelson Hamiltonian. Indeed, for the isotopes with $180 \le A \le 188$ the approaches are those abbreviated by SSA and ISWSA, respectively, while for the rest of nuclei, $190 \le A \le 196$, the SMA and ISWMA are alternatively used. It is worth mentioning that the approach called ISWMA was used for the first time in the present paper. Since the first set exhibits some features of the X(5) "symmetry" we compared the results of our calculations with those obtained with the X(5) formalism, if they are available. As for the other isotopes the results were compared with the Z(5) results. One concludes that our results are slightly better than those obtained with X(5) and Z(5) methods regarding both the excitation energies and reduced *E*2 transition probabilities.

The wave function structure is nicely reflected in the contour plots for the probability density. It is suggested that due to the Hamiltonian symmetries the wave functions might be suitable for accounting for shape evolution as well as for possible shape coexistence.

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