



Improved $^{192,194,195,196}\text{Pt}(n, \gamma)$ and $^{192}\text{Ir}(n, \gamma)$ astrophysical reaction rates

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^{192}Pt is produced solely by the slow neutron capture (s) nucleosynthesis process and hence an accurate (n, γ) reaction rate for this nuclide would allow its use as an important calibration point near the termination of the s -process nucleosynthesis flow. For this reason, we have measured neutron capture and total cross sections for $^{192,194,195,196}\text{Pt}$ in the energy range from 10 eV to several hundred keV at the Oak Ridge Electron Linear Accelerator. Measurements on the other Pt isotopes were, in part, necessitated by the fact that only a relatively small ^{192}Pt sample of modest enrichment was available. Astrophysical $^{192,194,195,196}\text{Pt}(n, \gamma)$ reaction rates, accurate to approximately 3%–5%, were calculated from these data. No accurate reaction rates have been published previously for any of these isotopes. At s -process temperatures, previously recommended rates are larger (by as much as 35%) and have significantly different shapes as functions of temperature than our new rates. We used our new Pt results, together with $^{191,193}\text{Ir}(n, \gamma)$ data, to calibrate nuclear statistical model calculations and hence obtain an improved rate for the unmeasured s -process branching-point isotope ^{192}Ir .

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I. INTRODUCTION

The $^{192}\text{Pt}(n, \gamma)$ cross section is particularly important to nuclear astrophysics studies for at least two reasons, as illustrated in Fig. 1. First, ^{192}Pt is the heaviest nuclide produced solely by the s process (the so-called s -only isotopes) for which reliable neutron capture data do not exist across the range of temperatures ($kT = 6$ – 26 keV) needed by s -process models. Second, ^{192}Pt is partially bypassed during the s process by a branching at ^{192}Ir . Because this branching is expected to be practically independent of temperature and electron density, it is particularly sensitive to the neutron density during the s process. However, a meaningful analysis of this branching has not been possible because of the large uncertainty in the $^{192}\text{Pt}(n, \gamma)$ reaction rate.

The only previously reported measurement [1] of the $^{192}\text{Pt}(n, \gamma)$ reaction rate employed an activation technique, so the reaction rate was obtained only at one temperature ($kT = 25$ keV) and with a fairly large uncertainty (29%). Activation measurements are hampered by the comparatively long half-life of ^{193}Pt and by the facts that both the half-life and decay intensities for this isotope are fairly uncertain. Theoretical estimates of reaction rates for the Pt isotopes [2–6], using statistical models, differ by up to a factor of 3. The currently recommended $^{192}\text{Pt}(n, \gamma)$ rate [7–9], which is based on a semiempirical estimate, is a factor of 3 larger than that obtained in the activation measurement.

An experimental determination of the $^{192}\text{Pt}(n, \gamma)$ reaction rate using a time-of-flight technique has been stymied by the very small (0.79%) natural abundance of this nuclide. The necessary gram-size, high-enrichment sample has not been available. However, given improvements made in the neutron capture apparatus at the Oak Ridge Electron Linear

Accelerator (ORELA) (e.g., see Ref. [10]), measurements with the available, relatively small (700 mg of Pt) sample of modest enrichment (56.7%) seemed feasible. The small sample size required a fairly long measurement time and the use of the same material for both the neutron capture and transmission measurements (with the shape of the sample refabricated between measurements). The low enrichment required complementary measurements on $^{194,195,196}\text{Pt}$.

One benefit of having to make all of these measurements is that these data should be very useful for testing and improving the nuclear statistical model, which still must be relied upon to calculate reaction rates for nuclides such as ^{192}Ir that are beyond the reach of current measurement techniques. Hence, we used our new data, together with existing $^{191,193}\text{Ir}(n, \gamma)$ cross sections [11] to constrain statistical-model parameters and obtain an improved prediction for the $^{192}\text{Ir}(n, \gamma)$ rate at s -process temperatures.

II. EXPERIMENT AND DATA REDUCTION

The experimental apparatus has been described previously many times (e.g., see Ref. [12] and references contained therein), so only the salient features will be mentioned herein. The ORELA [13–15] was operated at a pulse rate of 525 Hz, a pulse width of 8 ns, and a power of 7–8 kW. Neutron energy was determined by time of flight. The samples were isotopically enriched metallic platinum. The isotopic compositions and sample thicknesses are given in Table I. With the exception of the ^{192}Pt and natural samples, enrichments were greater than 96%.

Neutron capture measurements were made on flight path 7 at a source-to-sample distance of 40.12 m using a pair of C_6D_6 detectors and employed the pulse-height-weighting technique. A ^{10}B filter was used to remove overlap neutrons from preceding beam bursts, and a Pb filter was used to reduce γ -flash effects. These filters were placed in the beam

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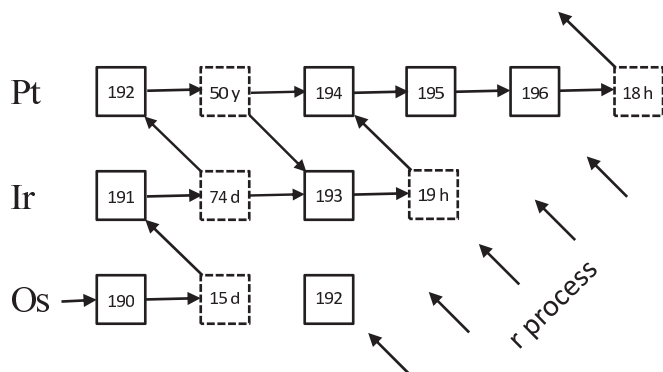


FIG. 1. Schematic diagram of the s -process path near Pt. Solid and dashed boxes represent stable and radioactive isotopes, respectively. Stable isotopes are labeled by their mass numbers whereas half-lives serve as labels for radioactive ones. Contributions to this mass region due to β decays following the r process, whose path runs through more neutron rich nuclides, are represented by diagonal arrows. Although ^{192}Pt is partially bypassed during the s process by a branching at ^{192}Ir , it is shielded against contributions from the r process by its stable isobar ^{192}Os ; hence, ^{192}Pt is an s -only isotope.

at a distance of 5 m from the neutron source. Cross-section normalization was made via the saturated resonance technique [16] using the 4.9-eV resonance in the $^{197}\text{Au}(n, \gamma)$ reaction.

A thin ^6Li -loaded glass scintillator [17], located 43 cm ahead of the sample in the neutron beam, was used to measure the energy dependence of the neutron flux. Separate sample-out background measurements were made, and measurements with a carbon sample were used to subtract the very small, smoothly varying background caused by sample-scattered neutrons.

Total neutron cross sections were measured via transmission on flight path 1 using a ^6Li -loaded glass scintillator at a source-to-detector distance of 79.827 m. Transmission measurements on a different sample were made at the same time, and hence under the same ORELA operating conditions, as the (n, γ) experiments (e.g., ^{196}Pt transmission was run at the same time as ^{194}Pt capture). The samples were cylindrical in shape, being between 7.9 to 16.8 mm in diameter. A ^{10}B filter was used to remove overlap neutrons from preceding beam bursts, and a Pb filter was used to reduce γ -flash effects. These filters were placed in the beam at a distance of 5 m from the neutron source. Separate runs for each sample were made at a pulse rate of 130 Hz to determine the residual background due to overlap neutrons from preceding

beam bursts. These runs were made at the same time as the $^{197}\text{Au}(n, \gamma)$ calibration measurements on flight path 7. The Pt samples were exchanged periodically with an empty sample holder, and with polyethylene and bismuth absorbers, which were used for determination of backgrounds.

\mathcal{R} -matrix analysis of the resolved resonance region will be described in a forthcoming paper. In the unresolved resonance region, the capture data were averaged over coarse energy bins and the relatively small corrections for multiple scattering and resonance self-shielding were calculated using the code SESH [18]. These data also were corrected for isotopic impurities in the samples using the current measurements.

III. AVERAGE CROSS SECTIONS AND ASTROPHYSICAL REACTION RATES

Cross sections averaged over coarse energy bins typically used in the calculation of astrophysical reaction rates are shown in Table II and Fig. 2. The cross sections in this figure have been multiplied by the square root of the energy at the center of each bin, effectively removing the underlying $1/v$ component, so that data over a wide range of energies for all four isotopes can be shown in the same graph. Open symbols in this figure represent average cross sections calculated directly from numerical integration of the data. Corrections for multiple scattering, self shielding, and isotopic impurities were calculated as described above. Solid symbols in Fig. 2 represent average cross sections calculated from the resonance parameters. Of course, in this case, corrections to the data are calculated on a resonance-by-resonance basis. Uncertainties common to both methods of calculating average cross section (e.g., due to normalization) are not included in this table or figure and therefore represent one-standard-deviation statistical uncertainties only. The good agreement between average cross sections obtained by the two techniques attests to the accuracy of the background subtraction and corrections we applied to our data. All four data sets show effects of neutron inelastic channels opening up, near 100 keV in ^{195}Pt and 300 keV in $^{192,194,196}\text{Pt}$.

SAMMY was used to calculate astrophysical reaction rates from the resonance parameters together with the averaged cross sections shown in Fig. 2. Resulting Maxwellian-averaged cross sections are given in Table III. As a check on the SAMMY calculations, reaction rates also were calculated following the technique of Ref. [20]. The two methods agree to within 0.5%. Statistical uncertainties in the reaction rates are negligible when compared to the overall normalization uncertainty. From

TABLE I. Isotopic compositions and thicknesses of samples.

Sample	Atomic percent						Sample thickness (10^{-3} atom/b)	
	^{190}Pt	^{192}Pt	^{194}Pt	^{195}Pt	^{196}Pt	^{198}Pt	Capture	Transmission
^{192}Pt	<0.5	56.97	26.16	11.23	4.70	0.90	0.463	4.53
^{194}Pt	<0.01	0.026	96.46	2.44	0.90	0.16	1.68	23.6
^{195}Pt	0.04	0.01	1.21	97.29	1.40	0.10	0.607	23.8
^{196}Pt	<0.01	0.04	0.67	1.69	97.25	0.28	1.71	14.4
Natural	0.01	0.79	32.9	33.8	25.3	7.20	–	7.57

TABLE II. Cross sections averaged over coarse energy bins for the $^{192,194,195,196}\text{Pt}(n, \gamma)$ reactions.

Energy range (keV)	Neutron capture cross section (mb)							
	^{192}Pt		^{194}Pt		^{195}Pt		^{196}Pt	
	Numerical Integration	Resonance	Numerical Integration	Resonance	Numerical Integration	Resonance	Numerical Integration	Resonance
3–5	1339 ± 30	1302 ± 29	694.3 ± 4.8	674.8 ± 4.7	2586 ± 22	2515 ± 21	481.2 ± 3.9	502.2 ± 4.1
5–7.5	1004 ± 22		607.0 ± 4.6	598.3 ± 4.5	1668 ± 20	1657 ± 20	285.4 ± 2.9	296.6 ± 3.0
7.5–10	871 ± 22		422.6 ± 4.7	426.1 ± 4.7	1292 ± 10		270.1 ± 3.2	275.1 ± 3.3
10–12.5	718 ± 22		414.4 ± 4.6	416.3 ± 4.6	1028 ± 10		189.8 ± 2.6	189.8 ± 2.6
12.5–15	606 ± 22		296.7 ± 3.8	301.9 ± 3.9	976 ± 10		184.5 ± 3.0	187.6 ± 3.0
15–20	547 ± 15		304.1 ± 2.8		804.4 ± 7.4		175.3 ± 1.9	
20–25	464 ± 15		254.5 ± 2.7		669.3 ± 7.1		153.7 ± 1.9	
25–30	424 ± 15		234.8 ± 2.8		622.8 ± 7.5		155.0 ± 2.0	
30–40	393 ± 11		228.8 ± 2.0		544.7 ± 5.5		129.4 ± 1.4	
40–50	352 ± 12		210.4 ± 2.3		477.9 ± 5.9		139.8 ± 1.7	
50–60	372 ± 12		207.2 ± 2.2		446.2 ± 5.5		129.1 ± 1.6	
60–80	322.7 ± 9.6		213.8 ± 1.9		420.5 ± 4.5		123.1 ± 1.3	
80–100	352.7 ± 8.9		218.3 ± 1.8		416.5 ± 4.2		122.9 ± 1.3	
100–120	362.8 ± 9.5		219.1 ± 1.9		337.3 ± 4.2		128.7 ± 1.4	
120–150	359.2 ± 8.0		208.1 ± 1.6		290.3 ± 3.4		117.9 ± 1.1	
150–175	319.6 ± 8.4		204.6 ± 1.7		248.9 ± 3.6		110.1 ± 1.2	
175–200	325.8 ± 8.3		194.6 ± 1.8		225.4 ± 3.6		100.8 ± 1.2	
200–225	323.9 ± 8.9		190.9 ± 1.8		193.6 ± 3.5		99.7 ± 1.2	
225–250	337.1 ± 8.8		184.9 ± 1.8		164.0 ± 3.4		95.9 ± 1.2	
250–300	334.8 ± 6.6		184.6 ± 1.3		128.3 ± 2.3		95.13 ± 0.90	
300–350	272.8 ± 5.8		164.1 ± 1.2		109.4 ± 2.1		87.22 ± 0.85	
350–400	227.1 ± 5.9		124.5 ± 1.2		96.1 ± 2.1		65.65 ± 0.78	
400–450	213.1 ± 5.7		109.1 ± 1.1		87.6 ± 2.1		52.45 ± 0.70	
450–500	214.5 ± 5.5		102.4 ± 1.1		79.5 ± 2.1		48.59 ± 0.65	

the uncertainties in the $^{197}\text{Au}(n, \gamma)$ and $^6\text{Li}(n, \alpha)$ cross sections, the statistical precision of the calibration measurements, the repeatability of the calibration runs, and uncertainties in the sample sizes and isotopes, we calculated normalization uncertainties of 3%–4% in the reaction rates.

We know of no measured $^{192,194,195,196}\text{Pt}(n, \gamma)$ reaction rates published in peer-reviewed journals. Our preliminary reaction rates reported in Ref. [21] are in agreement with the present rates to within the uncertainties. Although there have been numerous semiempirical and theoretical estimates, we know of only three other actual measurements of reaction rates to which our results can be compared, and only at a single temperature.

A measurement of the $^{192}\text{Pt}(n, \gamma)$ Maxwellian-averaged cross section (MACS) at $kT = 30$ keV, $\langle \sigma \rangle_{30} = 196 \pm 56$ mb, 2.5 times smaller than our measured rate, was reported in Ref. [1]. This measurement was made using a pseudo-Maxwellian neutron source at $kT = 25$ keV, which was then extrapolated to 30 keV. Activation measurements of the $^{192}\text{Pt}(n, \gamma)$ rate are hampered by the relatively long and somewhat uncertain half-life and the low energy of the decay of the product nuclide. Hence, it may not be too surprising that the most recent reaction-rate compilation [7] recommends a semiempirical estimate (590 ± 120 mb [22]) for this rate over the measurement of Ref. [1]. Note that $^{192,194,195}\text{Pt}(n, \gamma)$ reaction rates in the more recent KADoNiS compilation [8,9] are identical to those in Ref. [7].

There have been two previously reported [22,23] measurements of the $^{196}\text{Pt}(n, \gamma)$ reaction rate, both using the same technique as mentioned in the previous paragraph. Uncertainties in these previous results are about four times larger than ours. The most recent previous result (171 ± 22 mb at $kT = 30$ keV) [23] agrees with ours to within the quoted uncertainties, whereas the older result (197 ± 23 mb at $kT = 30$ keV) [22] is just outside the combined uncertainties. The currently recommended rate [9] is the weighted average of these two results together with the temperature dependence predicted by the statistical-model calculation of Ref. [6].

Our new $^{192,194,195,196}\text{Pt}(n, \gamma)$ reaction rates are compared to recommended rates from the most recent compilation [8,9] in Fig. 3. At s -process temperatures, our rates are far more accurate (uncertainties reduced by factors of 4–8) than the previously recommended rates. In general, previously recommended rates are larger (by as much as 35% at s -process temperatures) and have significantly different shapes, as functions of temperature, than our new rates.

IV. s -PROCESS NEUTRON DENSITY AND IMPROVED $^{192}\text{Ir}(n, \gamma)$ REACTION RATE

Although ^{192}Pt is an s -only isotope, it is partially bypassed during the s process by a branching at ^{192}Ir . Because this branching is expected to be practically independent of temperature and electron density, it is particularly sensitive to the

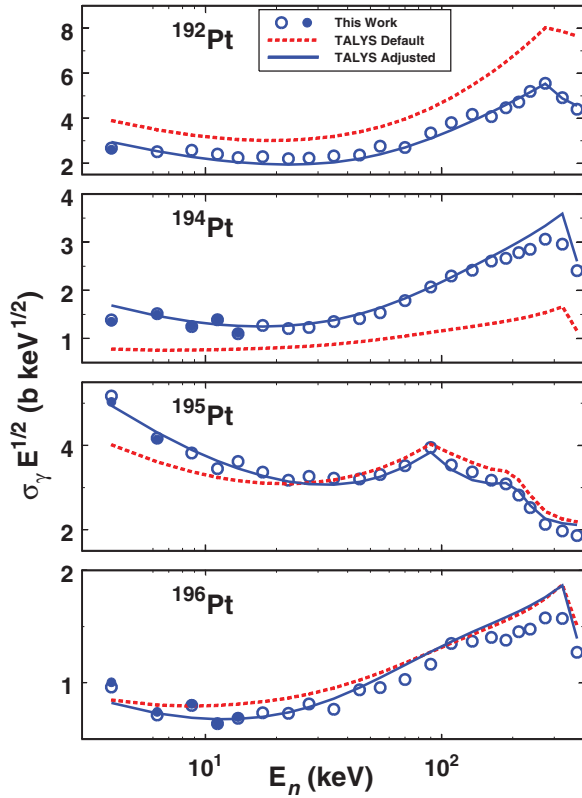


FIG. 2. (Color online) Reduced cross sections for $^{192,194,195,196}\text{Pt}(n, \gamma)$ averaged over coarse bins from 3 to 400 keV. Open symbols depict results obtained directly from the time-of-flight data whereas filled symbols show averaged cross sections calculated from the resonance parameters. Error bars, representing one-standard-deviation statistical uncertainties, are smaller than the symbols. Dashed red curves depict predictions, using default parameters, of the statistical-model code TALYS [19]. Solid blue curves depict TALYS calculations after calibration of model parameters using average resonance parameters. See text for details.

neutron density during the s process. A classical branching analysis [21] using our preliminary reaction rates resulted in a significantly lower neutron density [$n_n = (7^{+5}_{-2}) \times 10^7 \text{ cm}^{-3}$], inconsistent with the density [$n_n = (4.1 \pm 0.6) \times 10^8 \text{ cm}^{-3}$] resulting from analyses [24] of several other branchings. Because our new rates are consistent with our preliminary ones to within the uncertainties, we confirm this result. In contrast,

TABLE III. Maxwellian averaged cross sections for the $^{192,194,195,196}\text{Pt}(n, \gamma)$ reactions.

Thermal energy, kT (keV)	$\langle \sigma v \rangle / v_T$ (mb)			
	^{192}Pt	^{194}Pt	^{195}Pt	^{196}Pt
5	1261 ± 52	686 ± 21	2148 ± 65	408 ± 13
8	920 ± 38	508 ± 15	1497 ± 45	300.5 ± 9.4
10	801 ± 33	446 ± 13	1273 ± 38	263.7 ± 8.2
15	640 ± 26	361 ± 11	967 ± 29	214.8 ± 6.6
20	559 ± 23	320.3 ± 9.4	810 ± 24	190.8 ± 5.9
25	513 ± 21	297.2 ± 8.8	713 ± 21	176.7 ± 5.5
30	483 ± 20	282.5 ± 8.4	644 ± 19	167.4 ± 5.2

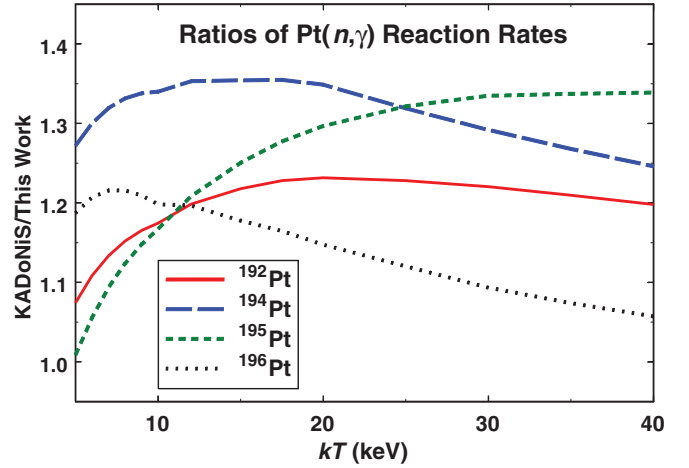


FIG. 3. (Color online) Ratios of astrophysical rates from Ref. [9] to our new rates for the $^{192,194,195,196}\text{Pt}(n, \gamma)$ reactions, depicted as solid red, long-dashed blue, short-dashed green, and dotted black curves, respectively.

just the opposite conclusion [$n_n = (4.3^{+3.4}_{-2.5}) \times 10^8 \text{ cm}^{-3}$] was reached in Ref. [1], based on a classical branching analysis using the only previous measurement of the $^{192}\text{Pt}(n, \gamma)$ reaction rate. The large reduction in extracted neutron density and its uncertainty from our new analysis are directly attributable to the substantially larger $^{192}\text{Pt}(n, \gamma)$ reaction rate and substantially reduced uncertainty, respectively, from our new measurements.

The large uncertainty in the above estimate of the neutron density from the ^{192}Ir branching is dominated by the assumed [21] factor of 2 uncertainty in the unmeasured $^{192}\text{Ir}(n, \gamma)$ reaction rate. Theoretical rates (which are based on nuclear statistical-model calculations) for this reaction listed in the latest compilation [9] vary by a factor of 2.7. In an attempt to obtain an improved rate for this reaction, we have used our new Pt data to constrain parameters in the nuclear statistical model. The code TALYS [19] was used for these calculations.

As shown in Fig. 2, average cross sections predicted by TALYS [19] using default parameters require normalizations ranging from 0.72 to 1.5 to yield reasonable agreement with our data. Also, the predicted cross section is significantly flatter as a function of energy than the measured one for $^{195}\text{Pt}(n, \gamma)$. As shown in Fig. 4, shapes of default TALYS cross sections also are significantly flatter than the $^{191,193}\text{Ir}(n, \gamma)$ data of Ref. [11]. There are substantial disagreements among the various $^{191,193}\text{Ir}(n, \gamma)$ measurements [11, 25–29]. We have chosen to use the data of Ref. [11] because they are the only sets to span the energy range needed to determine the reaction rate at s -process temperatures, and because other measurements by the same group with the same apparatus have shown, in general, to be reliable.

There are several model parameters which can be constrained by measured average resonance parameters. To this end, we report in Table IV D_0 , S_0 , and $\langle \Gamma_{\gamma 0} \rangle$ values resulting from \mathcal{R} -matrix analysis of the resolved resonance region. Because TALYS assumes that the Porter-Thomas distribution (PTD) [31] is valid (e.g., in calculating width fluctuation correction factors), D_0 and S_0 values in Table IV were

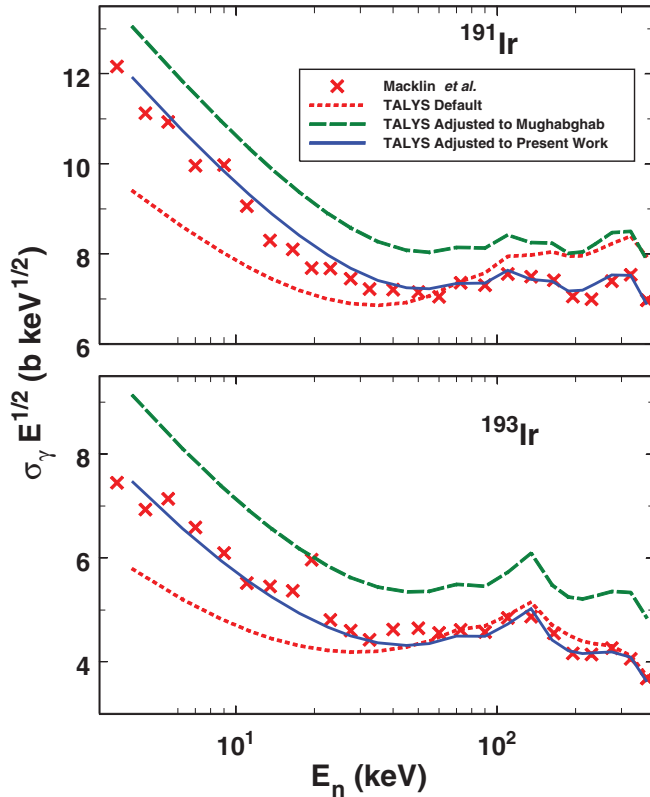


FIG. 4. (Color online) Reduced cross sections for ^{191,193}Ir(*n*, γ) averaged over coarse bins from 3 to 400 keV. Data are those of Macklin *et al.* [11] (X's). Dashed red curves depict predictions, using default parameters, of the statistical-model code TALYS [19]. Long-dashed green curves depict TALYS calculations after adjustment to average resonance parameters of Ref. [30]. Solid blue curves depict TALYS calculations after adjustment to local systematics. See text for details.

corrected for the effects of missed resonances using the technique of Ref. [32] and assuming the PTD. The $\langle \Gamma_{\gamma 0} \rangle$ values in Table IV were determined from a maximum-likelihood analysis by assuming these widths were χ^2 distributed. The corresponding uncertainties are one-standard-deviation estimates from the maximum-likelihood analysis. For ¹⁹⁵Pt, separate $\langle \Gamma_{\gamma 0} \rangle$ values were determined for definite 0⁻ and 1⁻ resonance assignments.

To obtain an improved ¹⁹²Ir(*n*, γ) rate, first, we adjusted the level density (at the neutron separation energy) parameter $a(S_n)$ to obtain agreement with the measured D_0 values for ^{192,194,195,196}Pt (Table IV) and ^{191,192,193}Ir [30]. The resulting $a(S_n)$ and D_0 values are listed in Table V, from which it can be

TABLE IV. Average resonance parameters.

Nuclide	D_0 (eV)	$10^4 S_0$	$\langle \Gamma_{\gamma 0} \rangle$ (meV)
¹⁹² Pt	28.4 ± 1.2	2.06 ± 0.23	62.40 ± 0.95
¹⁹⁴ Pt	71.8 ± 2.9	2.01 ± 0.22	76.7 ± 1.1
¹⁹⁵ Pt	15.93 ± 0.41	1.94 ± 0.14	$109.9^{+2.9}_{-2.7}$ (0 ⁻), $127.3^{+1.5}_{-1.4}$ (1 ⁻)
¹⁹⁶ Pt	192 ± 12	1.95 ± 0.34	$85.9^{+1.8}_{-1.9}$

seen that the latter agree with the measured values in Table IV to within the uncertainties.

Next, we used the “gamgam” option in TALYS to input the measured $\langle \Gamma_{\gamma 0} \rangle$ values (Table IV and Ref. [30]). These quantities, together with the D_0 values, are used to normalize the γ -ray transmission coefficients in TALYS.

Next, we adjusted parameters of the neutron optical model potential (NOMP) to obtain agreement with the measured S_0 values (Table IV and Ref. [30]). The data for ^{192,194,195,196}Pt and ^{191,193}Ir indicate that $S_0 = 2.0$ to about 10% or better accuracy. Therefore, we used this value for ¹⁹²Ir, which is consistent with the much less certain value given in Ref. [30] ($S_0 = 3.7 \pm 1.8$). This step was particularly important for obtaining the correct energy dependence of the cross sections for ¹⁹⁵Pt and ^{191,193}Ir but did not have much effect for ^{192,194,196}Pt. In particular, a ratio $S_1/S_0 \approx 0.1$ was required for the former cases to obtain agreement with the cross-section shape versus energy. Sensitivity of the ¹⁹⁵Pt and ^{191,193}Ir(*n*, γ) cross section shape to the neutron strength functions appears to be a consequence of the smaller average level spacings for these nuclides compared to ^{192,194,196}Pt.

Default S_0 values in TALYS were about 70% of the measured ones. We could obtain both $S_0 \approx 2 \times 10^{-4}$ and $S_1/S_0 \approx 0.1$ by adjusting both the a_V and a_D parameters in TALYS. Default and adjusted values of varied TALYS parameters are given in Table V.

As shown in Figs. 2 and 4, adjusted TALYS cross sections are in good agreement with the ^{192,194,195,196}Pt data, but they are about 15%–25% larger than the ^{191,193}Ir data, indicating that the average resonance parameters in Ref. [30] are inconsistent with the average cross-section data of Ref. [11]. Therefore, it seems unlikely that the ¹⁹²Ir(*n*, γ) reaction rate calculated using average resonance parameters from Ref. [30] will be reliable.

Of the three needed parameters (S_0 , D_0 , and $\langle \Gamma_{\gamma 0} \rangle$), D_0 is likely the most problematical, especially given that there are relatively few known resonances for the Ir isotopes and because the recommended D_0 values in Ref. [30] appear to have undergone significant corrections for missed resonances. In contrast, S_0 and $\langle \Gamma_{\gamma 0} \rangle$ typically are much less sensitive to such effects, and the recommended values for the Ir isotopes for these two parameters are in line with expectations based on our Pt data. Therefore, constraining D_0 for ¹⁹²Ir seems to be key for obtaining more reliable predictions of the ¹⁹²Ir(*n*, γ) reaction rate.

As noted above, D_0 and $\langle \Gamma_{\gamma 0} \rangle$ are used to normalize γ -ray transmission coefficients in TALYS. The equation used is

$$\frac{2\pi \langle \Gamma_{\gamma 0} \rangle}{D_0} = G_{\text{norm}} \sum_J \sum_{\Pi} \sum_{Xl} \sum_{I'=|J-l|}^{J+l} \sum_{\Pi'} \int_0^{S_n} dE_{\gamma} T_{Xl}(E_{\gamma}) \times \rho(S_n - E_{\gamma}, I', \Pi') f(X, \Pi', l), \quad (1)$$

where the J and Π sums are over compound nucleus states with spin J and parity Π that can be formed by s -wave incident neutrons, and I' and Π' denote the spin and parity of the final states that may be reached in the first step of the γ -ray cascade. The γ -ray transmission coefficient for type X (electric or magnetic) and multipolarity l for γ -ray energy E_{γ}

TABLE V. Default and adjusted TALYS parameters for level-density and neutron-optical models.

Nuclide	Default				Adjusted									
	$a(S_n)$ (MeV ⁻¹) ^a	D_0 (eV)	$10^4 S_0$	$10^4 S_1$	$a(S_n)$ (MeV ⁻¹)			D_0 (eV)			a_V ^b	a_D ^b	$10^4 S_0$	$10^4 S_1$
					LD1 ^c	LD2 ^c	LD3 ^c	LD1 ^c	LD2 ^c	LD3 ^c				
¹⁹¹ Ir	23.04	2.49	1.52	0.40	23.40	20.50	21.50	2.05	2.00	2.12	0.82	0.60	2.05	0.11
¹⁹² Ir	22.22	0.67	1.47	0.40	22.65	20.20	21.05	0.52	0.39	0.47	0.82	0.60	1.96	0.11
¹⁹³ Ir	21.41	7.01	1.43	0.40	21.90	19.90	20.60	5.35	5.26	5.65	0.80	0.60	2.04	0.10
¹⁹² Pt	24.00	21.97	1.38	0.42	23.49	19.69	21.01	28.39	28.44	28.47	0.80	0.68	2.04	0.13
¹⁹⁴ Pt	19.83	201.10	1.26	0.57	21.95	17.69	18.43	71.70	71.96	71.66	0.74	0.75	2.02	0.19
¹⁹⁵ Pt	20.38	17.99	1.26	0.42	20.60	17.97	18.38	15.87	15.89	15.91	0.78	0.75	2.03	0.15
¹⁹⁶ Pt	19.37	350.11	1.23	0.41	20.62	17.07	17.30	192.28	191.91	192.06	0.78	0.76	1.96	0.16

^aDefault TALYS level-density is model 1: constant temperature plus Fermi gas [33].

^bNumbers in these columns are factors by which default TALYS parameters were multiplied.

^cLD1 = TALYS level-density model 1, etc.

is denoted by $T_{Xl}(E_\gamma)$, the level density at excitation energy $S_n - E_\gamma$ is denoted by $\rho(S_n - E_\gamma, I', \Pi')$, and $f(X, \Pi', l)$ denotes the usual multipole selection rules. It is understood that the integral over dE_γ includes a summation over discrete states at lower excitation energies. The γ -ray transmission coefficient is related to the γ -ray strength function $f_{Xl}(E_\gamma)$ via

$$T_{Xl}(E_\gamma) = 2\pi f_{Xl}(E_\gamma) E_\gamma^{2l+1}. \quad (2)$$

Finally, G_{norm} is a normalization factor that ensures the equality of Eq. (1), and hence in practice the γ -ray strength functions are multiplied by this factor before they enter the nuclear reaction calculation in TALYS. Hence, it is possible to fit TALYS to a measured (n, γ) cross section by varying D_0 [by adjusting the level-density parameter $a(S_n)$]. Furthermore, the $a(S_n)$ values, adjusted to yield our D_0 values for Pt (Table V), are reasonably well fitted by a straight line as a function of mass number. Therefore, it seems likely that a reasonably reliable D_0 for ¹⁹²Ir could be obtained by using the average $a(S_n)$ value for ^{191,193}Ir, after they had been adjusted to yield agreement with the capture cross sections.

For these reasons, we used TALYS to fit the measured ^{191,193}Ir(n, γ) cross sections by adjusting the appropriate $a(S_n)$ value and used the average of the two $a(S_n)$ values to predict the ¹⁹²Ir(n, γ) cross section and reaction rate. We did this for all three level-density models in TALYS which allow $a(S_n)$ to be adjusted, and the values are given in Table V. The resulting D_0 values are about 2.0, 0.46, and 5.4 eV for ^{191,192,193}Ir, respectively. These adjusted ^{191,193}Ir D_0 values are intermediate to default TALYS values and those in the compilation of Ref. [30], whereas the adjusted D_0 values for ¹⁹²Ir are all smaller than the default and compilation ones.

The three ¹⁹²Ir(n, γ) cross sections calculated as described above agree to within 10%. Due to the normalization through Eq. (1), all four γ -ray strength function models yielded very similar predictions for a given level-density model. To further assess the uncertainty due to the level density, we varied $a(S_n)$ for each model by the average amount that our Pt $a(S_n)$ values (as functions of mass number, for each level-density model) deviated from linearity. This increased the maximum deviation of the various predicted ¹⁹²Ir(n, γ)

cross sections to 11.6%. Our recommended rate is the average of the largest and smallest rates predicted by TALYS following the above procedure, with base uncertainty calculated from the range of predictions. Additional uncertainties in this rate can arise from normalization [via Eq. (1)] to the measured average radiation width for ¹⁹²Ir [30], which can contribute a maximum uncertainty of 6%, and normalizations to the ^{191,193}Ir(n, γ) cross sections, which can add another 5% [7]. Therefore, we calculate that the overall uncertainty in the predicted ¹⁹²Ir(n, γ) reaction rate is about 22%. Our recommended MACS values for this reaction are given in Table VI.

In stellar models of the main s process, most of the neutron exposure occurs at temperatures near $kT = 8$ keV followed by a smaller exposure at about 23 keV. At these two temperatures, our recommended ¹⁹²Ir(n, γ) MACS values are 2.1 and 1.6, respectively, times larger than the rate [7] used in most previous s -process calculations. The larger ¹⁹²Ir(n, γ) MACS values we recommend would result in ¹⁹²Pt being more strongly bypassed during the s process. On the other hand, our new ¹⁹²Pt(n, γ) MACS value is smaller than the previously recommended rate [7], and this should result in less destruction of ¹⁹²Pt during the s process. New, realistic s -process calculations are needed to ascertain the net effect.

The previously recommended ¹⁹²Ir(n, γ) rate is based on calculations made with the statistical-model code NON-SMOKER [6,34], which had been normalized to data for nearby nuclides. The fact that the NON-SMOKER rate has a flatter temperature dependence suggests, based on our experience with TALYS as discussed above, that neutron strength functions (via

TABLE VI. Recommended ¹⁹²Ir(n, γ) MACS.

kT (keV)	MACS (mb)
5	8900 ± 1900
10	6000 ± 1300
15	4800 ± 1000
20	4080 ± 890
25	3590 ± 790
30	3220 ± 720
40	2670 ± 600

the NOMP) in NON-SMOKER need to be adjusted. This surmise is strengthened by the facts that NON-SMOKER predictions for the ^{195}Pt and $^{191,193}\text{Ir}(n, \gamma)$ cross sections also are flatter, as functions of energy, than the data, whereas predictions for $^{192,194,196}\text{Pt}$ are in better agreement with the data in this respect. Current consensus seems to be that (n, γ) cross sections are rather insensitive to the adopted NOMP (e.g., see Ref. [35] for a recent example). However, as we have demonstrated above, there are important exceptions. Therefore, it may be worthwhile to reexamine the sensitivity of theoretical rates to the NOMP, especially for nuclides predicted to have small average level spacings.

V. CONSTRAINTS ON PHOTON-STRENGTH-FUNCTION MODELS

Comparison of our $^{192,194,195,196}\text{Pt}(n, \gamma)$ data to TALYS calculations such as those described in the last section can help constrain γ -ray-strength-function models through the G_{norm} factor in Eq. (1); $G_{\text{norm}} > 1$ indicates that the γ -ray strength function (below S_n) is too small, and vice versa. However, the level-density model also enters this equation. Therefore, we calculated the average G_{norm} for 12 calculations, run for the four Pt isotopes with the three level-density models (which can be adjusted to measured D_0 values as described above), for each of the four γ -ray-strength-function models in TALYS. The resulting average G_{norm} values are 1.50 ± 0.80 , 0.39 ± 0.20 , 1.61 ± 0.70 , and 1.08 ± 0.46 for models 1–4, respectively, where listed uncertainties are standard deviations of the distributions. These results indicate that model 4 (Hartree-Fock-Bogolyubov [36]) yields results closest to the data, with nearly equal distribution of G_{norm} values below (7/12) and

above (5/12) 1.0. Model 2 (Brink-Axel Lorentzian [37,38]) gave the worst results, with all 12 G_{norm} values substantially smaller than 1.0, indicating that this γ -ray-strength-function model is consistently too large. Models 1 (Kopecky-Uhl generalized Lorentzian [39]) and 3 (Hartree-Fock BCS [36]) gave intermediate results with both having G_{norm} values larger than 1.0 (γ -ray-strength function too small) for 3/4 of the cases. Level-density models 1 (constant temperature plus Fermi gas [33]) and 2 (back-shifted Fermi gas [40]) consistently gave the largest and smallest, respectively, G_{norm} values.

VI. CONCLUSIONS

Our new $^{192,194,195,196}\text{Pt}(n, \gamma)$ and total cross section data represent very large improvements over previous work. Astrophysical (n, γ) reaction rates calculated from our data are substantially different from, and much more accurate than, currently recommended rates. We recommend a substantially larger $^{192}\text{Ir}(n, \gamma)$ reaction rate with steeper energy dependence, based in large part on calibration to local systematics deduced from our new Pt data. In addition, statistical-model calculations undertaken to assist this effort indicate that neutron strength functions may be more important for accurately predicting reaction rates for unmeasured nuclides than is routinely assumed.

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