Observation of new microsecond isomers among fission products from in-flight fission of 345 MeV/nucleon ²³⁸U

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A search for isomeric γ decays among fission fragments from 345 MeV/nucleon ²³⁸U has been performed at the RIKEN Nishina Center RI Beam Factory. Fission fragments were selected and identified using the superconducting in-flight separator BigRIPS and were implanted in an aluminum stopper. Delayed γ rays were detected using three clover-type high-purity germanium detectors located at the focal plane within a time window of 20 μ s following the implantation. We identified a total of 54 microsecond isomers with half-lives of ~0.1–10 μ s, including the discovery of 18 new isomers in very neutron-rich nuclei: ⁵⁹Ti^m, ⁹⁰As^m, ⁹²Se^m, ⁹³Se^m, ⁹⁴Br^m, ⁹⁵Br^m, ⁹⁶Br^m, ⁹⁷Rb^m, ¹⁰⁸Nb^m, ¹⁰⁹Mo^m, ¹¹⁷Ru^m, ¹¹⁹Ru^m, ¹²⁰Rh^m, ¹²²Rh^m, ¹²¹Pd^m, ¹²⁴Pd^m, ¹²⁴Ag^m, and ¹²⁶Ag^m, and obtained a wealth of spectroscopic information such as half-lives, γ -ray energies, γ -ray relative intensities, and $\gamma\gamma$ coincidences over a wide range of neutron-rich exotic nuclei. Proposed level schemes are presented for ⁵⁹Ti^m, ⁸²Ga^m, ⁹²Bt^m, ⁹⁴Bt^m, ⁹⁵Bt^m, ⁹⁷Rb^m, ⁹⁸Rb^m, ¹⁰⁸Nb^m, ¹⁰⁸Zt^m, ¹⁰⁹Mo^m, ¹¹⁷Ru^m, ¹¹⁹Ru^m, ¹²⁰Rh^m, ¹²²Rh^m, ¹²²Rh^m, ¹²²Rh^m, ¹²²Rh^m, ¹²²Rh^m, ¹²²Rh^m, ¹²²Rh^m, ¹²¹Pd^m, ¹²⁴Ag^m, and ¹²⁵Ag^m, based on the obtained spectroscopic information and the systematics in neighboring nuclei. The nature of the nuclear isomerism is discussed in relation to the evolution of nuclear structure.

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I. INTRODUCTION

It is known that the study of microsecond isomers is a useful means to investigate the evolution of shell structure and nuclear shape far from stability. One example where this has been applied is in the region of neutron-rich exotic nuclei with atomic numbers $Z \sim 28-50$. Nuclear isomerism is sensitive to microscopic changes of nuclear structure. For instance, in the vicinity of doubly magic nuclei ⁷⁸Ni and ¹³²Sn, the so-called seniority isomers are generated by the stretched configuration of high-*j* orbitals such as $g_{9/2}$ and $h_{11/2}$ [1]. These isomers, such as ⁷⁶Ni^m [2], ⁷⁸Zn^m [3], and ¹³⁰Cd^m [4], provide a means to investigate the persistence of shell gaps far from stability. On the other hand, the so-called shape isomers [5], which are

generated due to shape coexistence [6], have been observed in the intermediate region between major shells. The well-known shape isomers are those appearing in the neutron-rich exotic nuclei with atomic numbers $Z \sim 38-42$ and neutron numbers $N \sim 58-61$ [6]. For instance, in the case of ${}^{98}\text{Sr}^m$ [7] and ${}^{100}\text{Zr}^m$ [8], the spherical second 0⁺ states appear at a low excitation energy above the prolate-deformed ground state, and exhibit isomeric transitions. Furthermore other isomers, such as the socalled *K* isomers [9,10] and the isomeric transitions between band heads of low-lying deformed states [11], have also been observed in the intermediate region between N = 50 and 82.

Systematic investigations of isomers in neutron-rich exotic nuclei between ⁷⁸Ni and ¹³²Sn are of interest, because not only shape coexistence but also shape transition phenomena have been predicted by several nuclear models [12–14]. Isomer spectroscopy of such exotic nuclei should allow one to test

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these nuclear models as well as to investigate the evolution of nuclear shape and shape coexistence. Moreover nuclei in this region are also important for the astrophysical r process [15,16], since they are predicted to be located on or close to the r-process path. Despite the interest, isomers in these neutron-rich exotic nuclei have not yet been studied in detail due to the difficulties in producing such very rare isotopes far from stability. The present work is the first attempt to investigate them comprehensively.

The production of radioactive isotopes (RI) using a projectile fragment separator [17] provides us with a unique opportunity to observe short-lived isomers with half-lives of a few microseconds, thanks to in-flight separation. The fast isotopic separation and identification of reaction products, which take place in several hundred nanoseconds, allow event-by-event detection of isomeric γ rays at the focal plane of the separator with small decay losses in flight. The γ decays are observed under low-background conditions after ion implantation. These features allow the search for and highly selective spectroscopy of new isomers over a wide range of exotic nuclei far from stability. This was first demonstrated by Grzywacz et al. in an experiment at GANIL in 1995 [18], and a number of the succeeding experiments [3,4,19-24] have provided a wealth of spectroscopic information on the nuclear structure of exotic nuclei. In these experiments not only projectile fragmentation of heavy-ion beams, but also in-flight fission of a uranium beam have been used as production reactions to populate isomers. In-flight fission is known to be an excellent mechanism for producing neutron-rich exotic nuclei, having large production cross sections over a wide range of atomic numbers, as demonstrated by experiments at GSI in which more than a hundred new isotopes were identified [25,26]. The features of in-flight fission also provide us with an opportunity to expand isomer spectroscopy towards the neutron drip line. In fact several new isomers in the region of mass number $A \sim 100$ have been recently discovered at NSCL by Folden et al. using in-flight fission of an 80 MeV/nucleon ²³⁸U beam [27]. However, in spite of the great potential of in-flight fission, its application to isomer spectroscopy has been limited to relatively narrow regions in the nuclear chart due to limited performance of fragment separator and primary uranium beam. To take full advantage of the production mechanism of in-flight fission, its reaction kinematics requires a fragment separator with large ion-optical acceptance, so that efficient collection of fission fragments can be achieved. Fission fragments are produced with much larger spreads in angle and momentum compared to projectile fragmentation. Increased intensity and energy of the primary uranium beam relative to past experiments provides significant benefits as well. These conditions are met in the new-generation in-flight RI beam facilities that have been or are being developed worldwide, allowing dramatic advances in isomer spectroscopy.

One of the new generation of devices, the superconducting in-flight separator BigRIPS [28,29], was recently developed at RIKEN RI Beam Factory (RIBF) [30]. It is a two-stage separator with large acceptance so that the features of in-flight fission can be exploited. The second stage of the BigRIPS separator is designed to have high momentum resolution which allows excellent particle identification for RI beams [31]. Such capabilities of the BigRIPS separator have been well demonstrated in an experiment to search for new isotopes using in-flight fission of a ²³⁸U beam at 345 MeV/nucleon, in which we discovered 45 new neutron-rich isotopes with atomic numbers ranging from 25 to 56 [32]. The experiment also allowed us to expand isomer spectroscopy towards the neutron drip line. While searching for new isotopes, we simultaneously observed delayed γ rays at the focal plane in order to search for new isomers as well as to perform isomer tagging [32]. As anticipated, a number of isomeric decays were observed, including those from new isomers in very neutron-rich exotic nuclei. We obtained new spectroscopic information on not only the new isomers but also on some known, previously reported, isomers. In this paper we report on the experimental results from our comprehensive isomer search among the fission fragments, which was performed using the BigRIPS separator at RIBF at RIKEN Nishina Center.

II. EXPERIMENT

The experiment was carried out using a ²³⁸U⁸⁶⁺ beam accelerated to 345 MeV/nucleon by the cascade operation of the RIBF accelerator complex. The beam intensity was approximately 0.2 particle nA (pnA) on target. The present data on new isomers were recorded during the same runs as the recent search for new isotopes [32]. The fission fragments produced by the in-flight fission of the uranium beam were analyzed and identified in the BigRIPS separator, transported to the focal plane of the ZeroDegree spectrometer [32] and implanted in an aluminum stopper. Delayed γ rays were detected using three clover-type germanium detectors within a time window of 20 μ s following the implantation (see Ref. [31] for the layout and configuration of the BigRIPS separator and ZeroDegree spectrometer). The details of the separator settings are summarized in Table I of Ref. [32]. As described in the reference we employed three different separator settings, each of which targeted new isotopes and new isomers in the regions with atomic numbers around 30, 40, and 50, respectively. In the following, these are referred to as the G1, G2, and G3 settings, respectively.

The details about the production, separation, and particle identification of fission fragments are given in Refs. [31] and [32]. Here we describe the experimental procedure for the measurement of isomer decays. The first stage of the BigRIPS separator was employed to collect and separate the fission fragments, while the second stage, combined with the following ZeroDegree spectrometer, served as a spectrometer for the particle identification of fragments. An achromatic aluminum energy degrader was used in the first stage of the BigRIPS separator to select the range of isotopes to be measured. Another energy degrader was used in the second stage in the case of the G3 setting, because further purification was needed. The angular acceptance of the BigRIPS separator was ± 40 mrad horizontally and ± 50 mrad vertically, while the momentum acceptance was set to $\pm 3\%$ by using slits at the dispersive focus located at the midpoint of the first stage. The particle identification was performed based on the ΔE -TOF- $B\rho$ method, in

which the energy loss (ΔE), time of flight (TOF), and magnetic rigidity ($B\rho$) were measured and used to deduce the atomic number (Z) and the mass-to-charge ratio (A/Q) of fragments. The TOF and $B\rho$ values were measured in the second stage of the BigRIPS separator, while the ΔE measurement was performed at the focal plane of the ZeroDegree spectrometer. Note that in the search for new isotopes the ΔE measured at the end of the BigRIPS separator was used for the G2 and G3 settings (see Ref. [32] for details about the detectors used).

The identified fragments were implanted into the aluminum stopper which was located in air at the focal plane of the ZeroDegree spectrometer. Prior to stopping, the ions were slowed down in an energy absorber consisting of a stack of aluminum plates. We chose the stopper thickness of 30 mm for the G1 setting, 10 mm for the G2 setting, and 10 mm for the G3 setting, based on the stopping-range calculation of ATIMA [33]. The thickness of the absorber stack was 15 mm, 13 mm, and 8 mm, respectively. The stopper was tilted by 26.6° both horizontally and vertically with respect to the beam axis. This allowed reduction of the overall γ -ray attenuation in the stopper, because we could use a thinner stopper plate as well as avoid large attenuation in some directions. The stopper had an effective area of $90 \times 90 \text{ mm}^2$. The beam profile on the stopper was monitored by using two sets of position-sensitive parallel plate avalanche counters (PPAC) [34]. The flight path between the production target and the stopper was 127.3 m long, and the flight time was typically 600-700 ns: e.g., 620 ns for ⁷⁸Zn in the G1 setting, 630 ns for ¹⁰⁸Zr in the G2 setting, and 670 ns for ¹³⁴Sn in the G3 setting.

The γ rays emitted from the stopped ions were detected using three clover-type high-purity germanium (Ge) detectors that were placed approximately 7 cm on the left, right, and lower sides of the stopper. The Ge detectors, each of which consists of four crystals, were operated in an add-back mode to construct γ -ray energy spectra. Lead collimators, 50 mm thick, were placed upstream of the Ge crystals to protect them from charged particles. The γ -ray detection was made based on the delayed γ coincidence technique. We recorded the time interval between the γ -ray emission and the ion implantation by using a multihit time-to-digital convertor (TDC) module, and the obtained time spectrum was used to deduce the half-life of isomeric state. The range of the multihit TDC was set to 20 μ s, and the time of implantation was taken from a plastic scintillation counter located at F3 in the BigRIPS separator [32]. We also recorded $\gamma\gamma$ coincidences among the Ge detectors. The time spectrum of $\gamma\gamma$ coincidence was constructed from the difference of the time intervals between detector hits, in order to identify true coincidence events. We investigated the $\gamma \gamma$ coincidences among all 12 individual crystals as well as among the three clover-type Ge detectors (operated in an add-back mode). The former provided $\gamma\gamma$ coincidence data for low energy γ rays.

The energy threshold and dynamic range of the Ge detectors were set to 50 and 4000 keV, respectively. The energy calibration and the measurement of absolute photopeak efficiency were made by using standard γ sources that were affixed on the front of the stopper. The measured overall efficiency of the three Ge detectors was 8.4% for 122-keV γ rays and 2.3% for 1408-keV γ rays in the case of the 30-mm-thick aluminum stopper (G1 setting), and 11.9% and 2.7% in the case of the 10-mm-thick aluminum stopper (G2 and G3 settings), respectively. The energy resolution (full width at half maximum) was 2.1 keV for 1-MeV γ rays. The online detection efficiency for isomeric γ rays was found to decrease to 70% of the above offline values in the worst case. This is due to pileups caused by prompt γ and x rays, which are emitted prior to the isomeric decay in the process of the ion implantation. In the offline analysis, we excluded γ -ray events with the time interval approximately five times larger than the half-life to reduce γ -ray backgrounds due to accidental coincidences between the γ -ray emission and the ion implantation. We also excluded those with a time interval <200 ns to reduce background events caused by the prompt γ and x rays. In investigating the $\gamma\gamma$ coincidences, the prompt events were rejected by gating on a two-dimensional plot of energy versus time.

Figures 1(a) and 1(b) show the Z versus A/Q particle identification (PID) plots obtained in the G2 setting, corresponding to PID plots without (a) and with (b) the delayed γ coincidence. Figure 1(a) shows clear identification of fission fragments including their charge states. Thanks to the excellent A/Q resolution, the peaks for fully stripped (Q = Z) and hydrogenlike (Q = Z - 1) ions are well separated from each other for each isotopic line. Figure 1(b) shows the observation of isomers in which isotope events having isomeric decays are clearly enhanced. Strongly observed isomers are labeled in an enlarged PID plot shown in Fig. 1(c). Here we selected γ -ray events with a time interval between 0.2 and 5 μ s to exclude the accidental coincidence events as well as the prompt events. The probability of the accidental coincidence was estimated from high implantation statistics isotopes that do not contain isomeric decays. The estimate is that on average only 2.5% of implants would generate a background coincidence.

A delayed γ -ray energy spectrum has been obtained for each isotope by applying the time gate as well as a PID gate, and we have found a number of γ -ray peaks that we attribute to isomeric decays. If the statistics of the γ -ray peaks were poor, we made a significance test [35] to confidently identify them against the continuous background γ rays that were generated by accidental coincidences. The statistical probability of misidentification called the "p value" [35] was evaluated based on the background spectrum obtained from high implantation statistics isotopes that do not contain isomeric decays. The isomeric γ rays shown in the next section were those identified with a p value smaller than 0.1%, which corresponds to a confidence level better than 99.9%.

Background γ -ray peaks were observed in the delayed γ -ray energy spectra. We carefully analyzed those γ -ray peaks that appeared in multiple spectra, in order to find the origin of background γ rays. In addition to natural background γ rays such as 40 K, the following three origins were identified: (i) neutron-induced reactions such as 72 Ge(n, n') 72 Ge(691.55 keV, 0⁺, 444.2 ns) and 19 F($n, n'\gamma$) 19 F(197.14 keV, 5/2⁺, 89.3 ns) [36] that occur around the γ -ray detectors; (ii) β activities such as 24 Na and 26 Al produced from the stopper material in the process of the ion implantation; and (iii) isomeric γ decays from neutron-rich nuclei produced by secondary reactions of the implanted fragments.

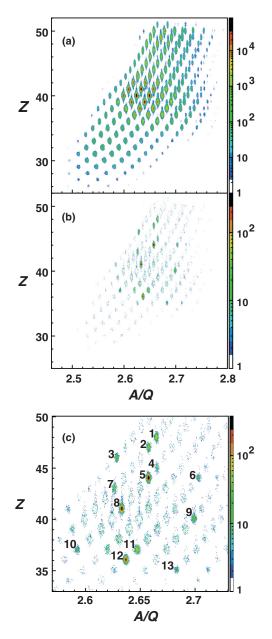


FIG. 1. (Color) Z versus A/Q PID plots of the fission fragments produced in the ²³⁸U + Be reaction at 345 MeV/nucleon. The data shown are those obtained with the G2 setting. (a) and (b) correspond to the PID plots without and with the delayed γ coincidence, respectively. (c) shows an enlarged plot of (b), in which the strongly observed isomers are labeled as 1: ¹²⁸Cd^m, 2: ¹²⁵Ag^m, 3: ¹²¹Pd^m, 4: ¹²⁰Rh^m, 5: ¹¹⁷Ru^m, 6: ¹¹⁹Ru^m, 7: ¹¹³Tc^m, 8: ¹⁰⁸Nb^m, 9: ¹⁰⁸Zr^m, 10: ⁹⁶Rb^m, 11: ⁹⁸Rb^m, 12: ⁹⁵Kr^m, and 13: ⁹⁴Br^m. The time gate used in (b) and (c) is from 0.2 to 5 μ s following ion implantation (see text).

The total number of the implanted fragments was 5.78×10^7 , 4.41×10^7 , and 8.46×10^7 in the G1, G2, and G3 settings, respectively, and the corresponding implantation rate is 529, 270, and 871 pps on average. The total number of the implanted fragments having isomeric decays was 8.97×10^6 , 1.03×10^7 , and 3.33×10^7 in the G1, G2, and G3 settings, respectively, and the corresponding implantation rate is 82.1, 63.1, and 343 pps on average. The net observation time in

the G1, G2, and G3 settings was 30.3, 45.3, and 27.0 h, respectively. The counting rates of the isomeric γ rays, including Compton-scattered events, are determined by the isomer ratio, the decay loss in flight, the internal conversion coefficients, and the γ -ray detection efficiency. The isomer ratio varies with the observed isomers, ranging from several percent to a few tens of percent.

III. RESULTS AND DISCUSSION

The microsecond isomers that we have observed in the present experiment are summarized in Tables I and II along with the BigRIPS settings, the number of implanted isotopes, the energy of isomeric γ rays (E_{γ}) , and the half-lives $(T_{1/2})$. The tables list 54 isomers in total, including the new isomers listed in Table I. The γ -ray peaks were fit using the code gf3 in the RADWARE software package [37]. The systematic uncertainty of γ -ray energies was estimated to be ± 0.5 keV from the accuracy of the energy calibration with standard γ sources. The statistical uncertainty of γ -ray energies was much smaller than the systematic one in most cases. The half-lives were deduced from fitting analysis of the time spectra of isomeric γ decays, in which the maximum likelihood method was used assuming exponential decay component(s) plus a constant background. In cases where more than one γ decay from a given isotope showed the same half-life, we added the time spectra to improve the half-life determination. The half-life values given in Tables I and II are those obtained from the latter time spectra. Some more details of the fitting analysis are given in their footnotes. The new isomers discovered in the present experiment were ${}^{59}\text{Ti}^m$, ${}^{90}\text{As}^m$, ${}^{92}\text{Se}^m$, ${}^{93}\text{Se}^m$, ${}^{94}\text{Br}^m$, ${}^{95}\text{Br}^m$, ${}^{96}\text{Br}^m$, ${}^{97}\text{Rb}^m$, ${}^{108}\text{Nb}^m$, ${}^{109}\text{Mo}^m$, ${}^{117}\text{Ru}^m$, ${}^{119}\text{Ru}^m$, ${}^{120}\text{Rh}^m$, ${}^{122}\text{Rh}^m$, ${}^{121}\text{Pd}^m$, ${}^{124}\text{Pd}^m$, ${}^{124}\text{Ag}^m$, and ${}^{126}\text{Ag}^m$. Figures 2 and 3 show their energy spectra and time spectra, respectively. Half-lives have been determined for these new isomers for the first time, except for ${}^{124}Pd^m$ and ${}^{126}Ag^m$, which were observed with $T_{1/2} > 20 \ \mu$ s. We classify ⁵⁹Ti^{*m*}, ¹¹⁷Ru^{*m*}, ¹²⁰Rh^{*m*}, ¹²¹Pd^{*m*}, and ¹²⁴Ag^{*m*} as new isomers due to incomplete or differing results in previous reports. In the cases of ¹¹⁷Ru^{*m*}, ¹²¹Pd^{*m*}, and ¹²⁴Ag^{*m*}, only the γ -ray energy spectra were presented without any analysis in a conference proceedings report [38] and details have not been published in a refereed paper. In the case of 120 Rh^m, the γ -ray energy spectrum published in Ref. [39] differs substantially from what was observed in the present work. For ⁵⁹Ti^{*m*} the γ -ray energy given in the previous report [40] differs significantly with our present data by as much as 8 keV. Figures 4 and 5 typically show the $\gamma \gamma$ coincidence spectra of ¹⁰⁸Nb^m and ¹¹⁹Ru^m, respectively. The¹⁰⁸Nb^m spectra represent the cases with good statistics and the ¹¹⁹Ru^m spectra those with typical statistics.

The observed isomers are shown in the nuclear chart of Fig. 6. We identified a number of new and known isomers in neutron-rich exotic nuclei between $N \sim 50$ and 82 from ⁷⁸Ni to ¹³²Sn. Isomeric states with $N \sim 58$ –61 have been known to be generated due to shape coexistence of spherical and prolate-deformed states [6]. In the present work several new isomers with Z < 38 were discovered, allowing us to

TABLE I. New isomers observed in the present work along with the corresponding BigRIPS settings, numbers of im	aplanted fragments,
half-lives $(T_{1/2})$, γ -ray energies (E_{γ}) , and γ -ray relative intensities (I_{γ}) .	

Isomers	BigRIPS setting ^a	Number of fragments	$T_{1/2} (\mu s)$	$E_{\gamma}(\text{keV})$	$I_{\gamma} \ (\%)^{b}$
⁵⁹ Ti ^m	G1	1.1×10^{4}	$0.587^{+0.057}_{-0.051}$	109.0	
$^{90}As^m$	G1	1.1×10^{4}	$0.20^{+0.12}_{-0.09}$	124.5	
92 Se ^m	G1	6.4×10^{4}	$10.3^{+5.5}_{-2.8}$	503.4	100(20)
			-2.8	538.8	89(18)
				897.8	76(18)
93 Se ^m	G1	3.7×10^{3}	$0.39_{-0.08}^{+0.12}$	208.3	85(26)
			0.00	469.9	100(43)
$^{94}\mathrm{Br}^m$	G1,G2	2.4×10^{5}	0.530(15)	88.0	2.6(20)
				92.1	44(3)
				115.5	6.2(13)
				165.7	12(2)
				179.0	15(2)
				202.6	100(4)
$^{95}\mathrm{Br}^m$	G1,G2	5.0×10^{4}	$6.7^{+1.1}_{-0.9}$	537.9	
$^{96}\mathrm{Br}^m$	G1,G2	4.1×10^{3}	$2.7^{+1.1}_{-0.7}$ c	311.5	
97 Rb ^m	G1,G2	1.4×10^{6}	$6.33^{+0.37}_{-0.34}$	77.1	
108 Nb ^m	G2	2.7×10^{6}	$0.109(2)^{e}$	64.3	1.6(4)
				77.6	100(2)
				89.0	56(1)
				102.2	19(1)
109 Mo ^m	G2	8.3×10^{3}	$0.194\substack{+0.076\\-0.049}$	69.7	
117 Ru ^m	G2	3.1×10^{5}	$2.487^{+0.058}_{-0.055}$	57.8	3.9(4)
				82.5	28(1)
				102.9	27(1)
				127.4	5.1(5)
				184.6	100(2)
119 Ru ^m	G2	1.6×10^{4}	$0.383\substack{+0.022\\-0.021}$	90.8	69(6)
				136.3	100(7)
120 Rh ^m	G2	1.7×10^{5}	$0.294\substack{+0.016\\-0.015}$	59.1	8.2(15)
				98.1	100(5)
122 Rh ^m	G2	8.6×10^{3}	$0.82\substack{+0.13 \\ -0.11}$	63.9	35(9)
			.0.005	207.1	100(17)
$^{121}\mathrm{Pd}^{m1\mathrm{g}}$	G2	2.8×10^{4}	$0.460^{+0.085}_{-0.092}$	135.5	
$^{121}\mathrm{Pd}^{m2\mathrm{h}}$			$0.463^{+0.083i}_{-0.094}$		
124 Pd ^m	G2	3.8×10^{4}	>20 ^j	$62.2(16)^{k}$	
$^{124}Ag^m$	G2,G3	5.7×10^{4}	$1.62^{+0.29}_{-0.24}$	75.5	32(7)
				155.6	100(12)
$^{126}Ag^m$	G2,G3	1.3×10^{5}	>20 ^j	254.8	

^aG1, G2, and G3 indicate the BigRIPS settings that were tuned for the $Z \sim 30$, $Z \sim 40$, and $Z \sim 50$ regions, respectively (see text). Experimental conditions are given in Ref. [32].

^bThe relative intensities were derived from photopeak counts which were corrected with the γ -ray detection efficiency. They are normalized to the intensity of the strongest one among the observed peaks. The errors shown are statistical only. We estimate that the γ -ray detection efficiency has systematic errors of 15%.

^cIn this case the constant-background component in the time spectrum was separately deduced by fitting the neighboring events on the high-energy side of the relevant γ -ray peak, because the long half-life relative to the TDC time range did not allow the fitting analysis with a decay component plus a constant background. The error shown also includes the fitting error of the constant-background component.

^dOnly the G2 data were used to deduce the half-life, because its signal-to-noise ratio was much better than that of the G1 data. This is due to the smaller γ -ray attenuation in the stopper (see text).

^eThe half-life was deduced by using the most intense γ ray at 77.6 keV.

^fThe half-life was deduced by using the most intense γ ray at 184.6 keV.

 ${}^{g}m1$ represents the isomer located at the first lowest excitation energy.

^hm2 represents the isomer located at the second lowest excitation energy.

ⁱThe half-life of the unobserved ¹²¹Pd^{m2} was deduced by assuming two cascading-decay components and a constant background for the time spectrum of the observed 135.5-keV γ ray (see text).

^jThe half-life was estimated as given here, because the isomeric γ -ray events were equally distributed in the 20- μ s range of the time spectrum. ^kThe statistical error is given in parentheses.

TABLE II. γ -ray energies and half-lives of previously known isomers observed in the present work. The BigRIPS settings and numbers of implanted fragments are also displayed. The previously reported γ -ray energies and half-lives are presented for comparison along with the references. New experimental information obtained in the present work is indicated by boldface with an underline.

Isomers	Isomers BigRIPS Nu				Literature		
	setting ^a	of fragments			E_{γ} (keV)	$T_{1/2} (\mu s)$	Ref.
$^{43}S^m$	G1	4.6×10^{3}	320.9	$0.20\substack{+0.14 \\ -0.07}$	319	0.478(48)	[<mark>76</mark>]
${}^{50}\mathrm{K}^m$	G1	2.0×10^4	128.5;172.4	$0.138\substack{+0.050\\-0.041}$	44;70;101;127.4;171.4	0.125(40)	[77]
54Sc ^m	G1	2.8×10^4	110.7	$2.78^{+0.31}_{-0.26}$	110.5	2.77(2)	[20,44]
${}^{56}\mathrm{Sc}^m$	G1	8.0×10^{3}	140.6;188.2	$0.35^{+0.26}_{-0.12}$	47.7;140.5;187.8;587.2;727.1	0.290(30)	[44]
${}^{60}\mathrm{V}^{m1\mathrm{b}}$	G1	1.1×10^{5}	104.0	d	103.2	0.013(3)	[77]
${}^{60}\mathrm{V}^{m2\mathrm{c}}$	G1		99.7;104.0	$0.229\substack{+0.025\\-0.023}$	98.9;103.2	0.320(90)	[77]
$^{64}Mn^m$	G1	1.6×10^5	135.3	0.025	39.8;134.3	>100	[20,77]
					40.2;135.3	400(40)	[87]
67 Fe ^m	G1	2.2×10^{5}	366.8		366.4;387.7	64(14)	[88]
76 Ni ^m	G1	1.1×10^{4}	142.7;355.5;929.9;990.6	$0.409\substack{+0.058\\-0.050}$	144;354;930;992	$0.590\substack{+0.180\\-0.110}$	
$^{75}\mathrm{Cu}^{m1\mathrm{b}}$	G1	5.4×10^{5}	62.5	e	61.8	0.370(40)	[77]
$^{75}\mathrm{Cu}^{m2c}$	G1		66.8;62.5	$0.134\substack{+0.025\\-0.020}$	66.5;61.8	0.170(15)	[77]
78 Zn ^m	G1	1.3×10^{6}	145.7;730.0;890.5;909.1	$0.320\substack{+0.009\\-0.008}$	144.7;729.6;889.9;908.3	0.319(9)	[3]
82 Ga ^m	G1	1.1×10^{6}	141.0	$\underline{0.098}_{-0.009}^{+0.010}$	141	< 0.5	[27]
$^{92}\mathrm{Br}^m$	G1	1.7×10^{6}	98.6;106.3;139.7;154.9; 197.6 ;	$\underline{0.089}^{+0.007f}_{-0.008};$	99;106;139;155;169;239;	< 0.5	[27]
			239.5;259.5;301.9;778.8; 898.3 ;1039.8	$\underline{0.084}^{+0.010f}_{-0.009}$			
95 Kr ^m	G1.G2	1.7×10^{6}	82.6;114.3	1.582(22)	81.7;113.8	1.4(2)	[55]
⁹⁶ Rb ^m							
[~] Kb ^m	G2	6.4×10^{4}	59.9;90.3;93.3;117.2;123.2 ^h ; 149.7; 177.2;185.1;210.2;240.5;300.2 ^h ; 368.5 ^h ;461.4	1.72 ^{+0.16} g	38.0;59.3;89.5;92.8;116.8; 122.0;123.5;126.0;148.8; 166.1;177.6;185.4;209.9;240.3; 276.3;300.0;301.0;329.0;366.8; 369.2;402.4;405.5;461.6; 495.2;554.5	2.00(10)	[56]
98 Rb ^m	G1,G2	1.7×10^{6}	123.7;178.4	0.358(7)	116;124;178	$0.700\substack{+0.060\\-0.050}$	[27]
$97 {\rm Sr}^{m1b}$	G2	3.2×10^{3}	141.3;167.6	$0.52^{+0.16}_{-0.12}$	140.8;167.0	0.170(10)	[57]
97 Sr ^{m2c}	G2	0.2 / 10	1110,10110	-0.12	140.8;167.0;522.0	0.265(27)	[57]
					141.0;167.1;522.1	0.526(13)	[58]
100 Sr ^m	G2	1.6×10^6	129.6;288.1;1201.8	0.122(9)	129.3;204.0;288.2;997.5; 1201.0;1285.5;1370.7	0.085(7)	[10]
101 Y ^m	G2	1.7×10^{5}	128.0;203.5	$0.187\substack{+0.049\\-0.038}$	129;164;204;230;480	$0.860\substack{+0.090\\-0.080}$	[27]
108 Zr ^m	G2	1.5×10^{5}	174.4;277.9;347.2; 364.2 ; 426.3 ; 432.3 ;478.3; 485.3 ;603.9;	$0.536^{+0.026}_{-0.025}$	173.7;278.6;347.9;477.5;605.6	0.620(150)	
1063 71 10		1.2. 1.04	<u>642.3; 773.6; 795.9; 828.3</u>	0.55 ± 0.11		0.04(4)	
106 Nb ^m	G2	1.3×10^{4}	63.5;94.7;108.1;147.5;202.1;204.9	$0.66^{+0.11}_{-0.10}$	63.5;94.5;107.3;147.4;201.8;204.6		[78]
109 Nb ^m	G2	1.6×10^{6}	117.0;196.2;312.2	$0.114^{+0.008}_{-0.007}$	117;196.3;313.1	0.150(30)	[42]
$^{112}\mathrm{Tc}^{m}$	G2	7.2×10^{3}	93.1;259.2	$0.218\substack{+0.060\\-0.043}$	92;258	0.150(17)	[27,41]
113 Tc ^m	G2	6.7×10^{5}	114.4	$0.526\substack{+0.016\\-0.015}$	114	0.5(1)	[41]
$^{125}\mathrm{Ag}^{m}$	G2,G3	2.4×10^{5}	<u>103.2</u> ;669.8;684.0;714.5; 728.4; 765.1 ; 786.7	$0.498\substack{+0.021\\-0.020}$	672.1;686.3;716.6;731.2	0.473(111)	[75]
$^{125}\mathrm{Cd}^m$	G3	2.6×10^{3}	719.4;742.2		719.1;742.7	19(3)	[79]
$^{127}\mathrm{Cd}^m$	G2,G3	2.5×10^{5}	110.6;710.1;738.5;821.1;848.8	$11.0^{+9.2k,l}_{-3.5}$	110.4;711.2;738.8;821.2;848.9	17.5(3)	[80]
$^{128}\mathrm{Cd}^{m1b}$		6.5×10^{5}	439.2;644.9;783.9;1223.7	ee	440.0;645.8;784.6;1224.0	0.270(7)	[23]
$^{128}\mathrm{Cd}^{m2c}$			237.5;439.2;644.9;783.9;1223.7	d	237.9;440.0;645.8;784.6;1224.0	0.012(2)	[23]
$^{128}\text{Cd}^{m3j}$			68.8;237.5;439.2;536.8;	$3.76_{-0.37}^{+0.44}$	68.7;237.9;440.0;450.4;537.6;	3.56(6)	[23]
			644.9;783.9;1223.7		645.8;765.0;784.6;1224.0		
	G2,G3	9.1×10^{4}	128.0;138.0;538.2;1325.4	$0.248\substack{+0.021\\-0.019}$	128;138;539;1325	0.220(30)	[4]
$^{130}\mathrm{Cd}^m$							
$^{130}Cd^{m}$ $^{129}In^{m}$	G3	3.0×10^5	333.8;359.0;995.3;1354.6	$11.3^{+2.2k}_{-1.6}$	333.8;359.0;995.2;1354.1	8.7(7)	[81]
		$\begin{array}{l} 3.0\times10^5\\ 1.8\times10^6\end{array}$	333.8;359.0;995.3;1354.6 388.5	$\frac{11.3^{+2.2\mathbf{k}}_{-1.6}}{5.25^{+0.40\mathbf{k}}_{-0.35}}$	333.8;359.0;995.2;1354.1 388.8	8.7(7) 3.1(3)	[81] [82]

OBSERVATION OF NEW MICROSECOND ISOMERS AMONG ...

Isomers	BigRIPS	Number	E_{γ} (keV)	$T_{1/2} (\mu s)$	Literature		
	setting ^a	of fragments			E_{γ} (keV)	$T_{1/2} (\mu s)$	Ref.
130 Sn ^m	G3	2.5×10^4	96.7;391.6	$1.14_{-0.21}^{+0.27}$	96.54;391.39	1.61(15)	[84]
131 Sn ^m	G3	1.0×10^{6}	158.8;173.4;344.0;4273.8	$0.309_{-0.023}^{+0.024}$	158.57;173.28;344.40;	0.300(20)	[<mark>89</mark>]
					4102.3;4273.2;4446.0		
132 Sn ^m	G3	2.0×10^{7}	132.7;299.4;374.8;4041.2;4416.7	$2.088(17)^{m}$	64.4;132.5;299.6;310.7;375.1;	2.03(4)	[<mark>90</mark>]
					4041.1;4351.9;4416.2		
134 Sn ^m	G3	3.8×10^{6}	174.1;347.1;725.1	$0.086\substack{+0.008\\-0.007}$	174.0;347.8;725.6	0.080(15)	[85]
$^{136}\mathrm{Sb}^m$	G3	4.0×10^6	53.9;173.1	0.570(5)	51.4 ⁿ ;53.4;173.0	0.565(50)	[22,86]

TABLE II. (Continued.)

^aG1, G2, and G3 indicate the BigRIPS settings that were tuned for the $Z \sim 30$, $Z \sim 40$, and $Z \sim 50$ regions, respectively (see text). Each experimental condition is given in Ref. [32].

 ${}^{b}m1$ represents the isomer located at the first lowest excitation energy.

 $^{c}m2$ represents the isomer located at the second lowest excitation energy.

^dThe half-life was too short to be deduced in the present work.

^eThe half-life could not be deduced in the present work because two decay-component fitting was difficult due to the low statistics.

^fThe half-life was deduced to be 0.089 μ s for the 662-keV isomer using the γ rays at 106.3, 154.9, and 301.9 keV and 0.084 μ s for the 1138-keV isomer using those at 778.8, 898.3, and 1039.8 keV (see text).

^gThis value was deduced using the γ rays at 93.3, 117.2, 123.2, 210.2, 240.5, 300.2, 368.5, and 461.4 keV.

^hThis should be a doublet peak. See the previous report.

ⁱThe 196.2-keV γ ray was not used in deducing the half-life due to the mixture of the background γ ray at 197 keV (see text).

 ${}^{j}m3$ represents the isomer located at the third lowest excitation energy.

^kIn this case the constant-background component in the time spectrum was separately deduced by fitting the neighboring events on the high-energy side of the relevant γ -ray peak, because the long half-life relative to the TDC time range did not allow the fitting analysis with a decay component plus a constant background. The error shown also includes the fitting error of the constant-background component.

¹The half-life was deduced using the γ rays at 738.5 and 821.1 keV.

^mThis value was deduced using the most intense γ ray at 132.7 keV.

ⁿOnly conversion electrons were observed for the 51.4-keV transition in ${}^{136}Sb^m$.

investigate how shape coexistence evolves over the entire region. Several isomers were also observed in the region of $N \sim 65$ -70, in which a variety of nuclear shapes, including prolate, oblate, triaxial, and tetrahedral shapes [11,41-43], have been suggested. Furthermore we made the first broad exploration for the $N \sim 75$ region and discovered several new isomers. In addition to these observations, isomers in the lighter-mass region were also explored, some of which are of importance to study the shell evolution in the neutron-rich exotic nuclei between $N \sim 28$ and 50 [44-46].

Table I lists the measured γ -ray relative intensities (I_{γ}) for the new isomers. Those of known isomers are listed in Table III. In deducing the relative γ -ray intensities, we used the γ -ray detection efficiency that was evaluated by numerical simulation using the Monte Carlo code GEANT3 [47]. This allowed us to realistically take into account the γ -ray attenuation using the detailed geometrical layout of the stopper and detectors. The simulation well reproduced the measured γ -ray detection efficiency using standard γ sources that were affixed to the front surface of the stopper. We made the simulation for isomeric γ rays incorporating the realistic beam-spot profile and stopping-range distributions in the stopper, which were obtained based on the measured positions and $B\rho$ distributions of implanted fragments. The previously reported γ -ray relative intensities, such as those of ⁷⁸Zn^m (G1), ⁹⁵Kr^m (G1 and G2), ¹⁰⁰Sr^m (G2), ¹²⁷Cd^m (G2 and G3), ¹²⁸Cd^m (G2 and G3), ¹²⁹In^{*m*} (G3), ¹³¹Sn^{*m*} (G3), ¹³²Sn^{*m*} (G3), and ¹³⁴Sn^{*m*} (G3), are reproduced using the γ -ray detection efficiency thus obtained. We have also obtained some new results on the known isomers. They are indicated in Table II by boldface with an underline. New γ -ray peaks were identified for ${}^{92}\text{Br}^m$, ${}^{108}\text{Zr}^m$, and ${}^{125}\text{Ag}^m$. The half-lives of ${}^{82}\text{Ga}^m$ and ${}^{92}\text{Br}^m$ were determined for the first time. The precision of the half-lives was significantly improved for the following isomers: ${}^{95}\text{Kr}^m$, ${}^{98}\text{Rb}^m$, ${}^{108}\text{Zr}^m$, ${}^{109}\text{Nb}^m$, ${}^{113}\text{Tc}^m$, ${}^{125}\text{Ag}^m$, and ${}^{136}\text{Sb}^m$. Table II shows the comparison with the previous measurements. In most cases our half-life measurements are in agreement with those given in the literature.

The present experimental data allowed us to propose level schemes for some of the observed isomers. Level schemes were constructed based on the γ -ray relative intensities, the $\gamma\gamma$ coincidence information, and the correspondence of decay curves as well as the γ -ray energies and the half-lives. The energy sum, the cascade and side feeding, the intensity balance, and the multipolarity and reduced transition probability for isomeric γ -ray transitions were considered as well. Systematics of the level schemes and nuclear isomerism in neighboring nuclei as well as their theoretical studies were also compared, if available. In the case of odd-odd nuclei, we relied on the so-called Nordheim rule [48], in those cases where the configuration of single-particle orbitals could be inferred from the systematics of neighboring nuclei.

If the γ -ray energies were relatively low, the intensity balance analysis of cascading γ rays allowed us to deduce the multipolarities of transitions since the intensity correction for internal conversion is sensitive to the multipolarity [4,23,49].

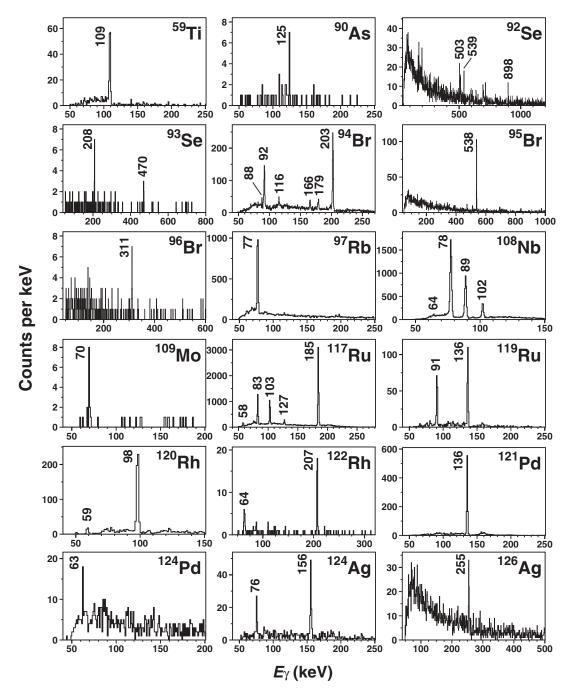


FIG. 2. Delayed γ -ray energy spectra of new isomers identified in the present work. Peaks are labeled by energies in keV. Unlabeled peaks such as those seen in ⁹²Se are the background γ rays described in text or unidentified γ rays. Note that the spectra of ⁹⁴Br and ¹⁰⁸Nb are given in 0.5 keV/channel.

This could not be applied in cases near the energy threshold where the γ -ray detection efficiency had the large uncertainty. The intensity correction was made using the total internal conversion coefficients calculated with the code BrIcc [49]. Furthermore the so-called recommended upper limits (RUL) analysis [50,51] was performed for the reduced transition probability to determine the upper limit of multipolarity for isomeric transitions.

Figures 7, 10, 11 and 15-18 show the level schemes that we propose for the new isomers: ⁵⁹Ti^{*m*} (Fig. 7), ⁹⁴Bt^{*m*} (Fig. 11),

⁹⁵Br^{*m*} (Fig. 11), ⁹⁷Rb^{*m*} (Fig. 11), ¹⁰⁸Nb^{*m*} (Fig. 15), ¹⁰⁹Mo^{*m*} (Fig. 15), ¹¹⁷Ru^{*m*} (Fig. 16), ¹¹⁹Ru^{*m*} (Fig. 16), ¹²⁰Rh^{*m*} (Fig. 17), ¹²²Rh^{*m*} (Fig. 17), ¹²¹Pd^{*m*} (Fig. 18), and ¹²⁴Ag^{*m*} (Fig. 17) and for the known isomers: ⁸²Ga^{*m*} (Fig. 7), ⁹²Br^{*m*} (Fig. 10), and ⁹⁸Rb^{*m*} (Fig. 11). These level schemes have been constructed for the first time. Figures 14 and 20 show our proposed revisions for the previously reported level schemes of ¹⁰⁸Zr^{*m*} (Fig. 14), and ¹²⁵Ag^{*m*} (Fig. 20). We could infer the nature of the nuclear isomerism for some isomers. Detailed discussions of the proposed level schemes are given in the following sections.

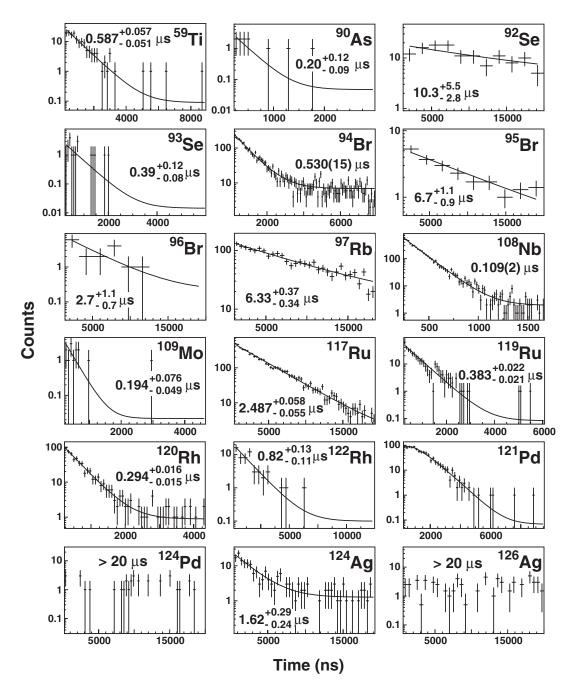


FIG. 3. Time spectra of new isomers identified in the present work. The solid line shows the fitted decay curve using an exponential decay component plus a constant background. The deduced half-lives are given in the figures except for ${}^{121}Pd^m$. Note that the time spectrum of ${}^{121}Pd^m$ is fit using two cascading-decay components plus a constant background. See text for the half-lives of ${}^{121}Pd^m$.

A. ⁵⁹Ti^m

The proposed level scheme for ${}^{59}\text{Ti}^m$ is shown in Fig. 7. A new isomeric γ ray at 109.0 keV with a half-life of $0.587^{+0.057}_{-0.051}$ μ s is observed. Based on the systematics in neighboring nuclei, we propose that the observed isomer in ${}^{59}\text{Ti}$ is an *E*2 isomer with spherical structure, which is generated due to a small transition energy. This type of isomer often appears in spherical nuclei. The reduced *E*2 transition probability *B*(*E*2) of the isomeric transition is $3.68^{+0.37}_{-0.34}$ W.u., consistent with spherical structure. The possibility of higher multipolarities is excluded, because the reduced transition probability significantly exceeds the RUL. The possibility of hindered *E* 1 or *M*1 transition is also excluded, because such an isomer, which usually appears as a shape isomer, has not been observed in the neighboring isotopes. Furthermore similar *E*2 isomers such as ${}^{54}Sc^m$ [20,44] and ${}^{56}Sc^m$ [44] appear in the vicinity of ${}^{59}\text{Ti}^m$, supporting our proposed origin.

The spin and parity (J^{π}) of the ground state in ⁵⁹Ti is tentatively assigned as $5/2^-$, based on the systematics of ⁵⁵Ti [52] and ⁵⁷Ti [44,53], which indicates that the $vf_{5/2}$ single-particle orbital is located above the $vp_{1/2}$ orbital in these neutron-rich Ti isotopes. Assuming such a shell ordering

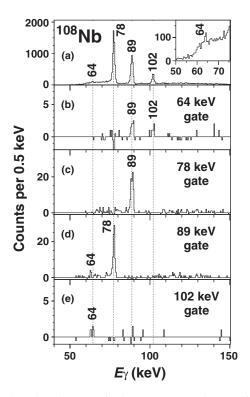


FIG. 4. Delayed γ -ray singles spectrum and $\gamma\gamma$ coincidence spectra of ¹⁰⁸Nb^m. The γ -ray spectra observed in coincidence with (b) the 64.3-, (c) 77.6-, (d) 89.0-, and (e) 102.2-keV γ rays are shown along with (a) the γ -ray singles spectrum. A partially enlarged γ -ray singles spectrum is shown in the inset of (a). Peaks are labeled by energies in keV.

and similarity to 57 Ti, the J^{π} of the 109.0-keV isomeric state in 59 Ti is tentatively assigned as $1/2^-$, with a configuration dominated by the $vp_{1/2}$ hole state. These assignments are consistent with the proposed E2 multipolarity for the isomeric transition. According to Ref. [44], the $1/2^-$ state in 57 Ti is located at 364 keV above the $5/2^-$ ground state. This implies

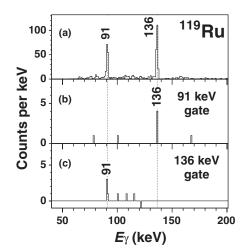


FIG. 5. Delayed γ -ray singles spectrum and $\gamma\gamma$ coincidence spectra of ¹¹⁹Ru^{*m*}. The γ -ray spectra observed in coincidence with (b) the 90.8- and (c) 136.3-keV γ rays are shown along with (a) the γ -ray singles spectrum. Peaks are labeled by energies in keV.

TABLE III. Deduced γ -ray relative intensities for previously reported isomers. New γ -ray peaks observed in the present work are indicated by boldface with an underline.

Isomers	E_{γ} (keV)	$I_{\gamma}(\%)^{\mathrm{a}}$	Isomers	E_{γ} (keV)	$I_{\gamma} \ (\%)^{a}$
$^{60}\mathrm{V}^{m1\mathrm{b}}$	104.0	60(8) ^d	$(^{109}\text{Nb}^{m})$	196.2	100(13) ^g
${}^{60}\mathrm{V}^{m2\mathrm{c}}$	99.7	$100(11)^{d}$		312.2	76(10)
75 Cu ^{$m1b$}	62.5	100(14) ^e	$^{108}{ m Zr}^{m}$	174.4	100(8)
$^{75}\mathrm{Cu}^{m2c}$	66.8	55(9) ^e		277.9	37(6)
$^{92}\mathrm{Br}^{m\mathbf{f}}$	98.6	100(5)		347.2	40(6)
	106.3	59(5)		364.2	25(6)
	139.7	10(6)		426.3	47(7)
	154.9	66(4)		432.3	22(6)
	<u>197.6</u>	2.1(36) ^g		478.3	33(7)
	239.5	9.6(21)		485.3	15(5)
	259.5	60(4)		603.9	50(9)
	301.9	12(3)		<u>642.3</u>	22(6)
	778.8	11(3)		773.6	21(7)
	<u>898.3</u>	6.1(19)		795.9	14(5)
	1039.8	20(3)		<u>828.3</u>	22(8)
98 Rb ^m	123.7	100(2)	$^{125}Ag^m$	<u>103.2</u>	70(5)
	178.4	45(2)		669.8	77(9)
$^{106}\text{Nb}^m$	63.5	40(13)		684.0	76(9)
	94.7	67(19)		714.5	100(11)
	108.1	40(15)		728.4	76(10)
	147.5	16(12)		<u>765.1</u>	14(4)
	202.1	66(19)		<u>786.7</u>	9.0(40)
	204.9	100(24)	$^{136}\mathrm{Sb}^m$	53.9	1.7(1)
¹⁰⁹ Nb ^{mh}	117.0	96(9)		173.1	100(1)

^aThe relative intensities were derived from photopeak counts which were corrected with the γ -ray detection efficiency. They are normalized to the intensity of the strongest one among the observed peaks. The errors shown are statistical only. We estimate that the γ -ray detection efficiency has systematic errors of 15%.

 ${}^{b}m1$ represents the isomer located at the first lowest excitation energy. ${}^{c}m2$ represents the isomer located at the second lowest excitation energy.

^dThe time window was set from 0.2 to 1.0 μ s.

^eThe time window was set from 0.2 to 2.1 μ s.

^fThe time window was set from 0.15 to 0.5 μ s.

^gThe contribution of the 197.1-keV γ ray originating from the ¹⁹F($n, n'\gamma$) reaction was subtracted based on the 197.1-keV peak observed in the neighboring isotopes: ⁹⁰Se for ⁹²Br^m and ¹⁰⁶Zr for ¹⁰⁹Nb^m, applying the same time window.

^hThe time window was set from 0.2 to 0.6 μ s.

that the subshell gap between the $vp_{1/2}$ and $vf_{5/2}$ orbitals gets smaller as the neutron number increases, thus generating the isomeric transition in ⁵⁹Ti.

B. ⁸²Ga^m

The proposed level scheme for ${}^{82}\text{Ga}^m$ is shown in Fig. 7 and the γ -ray energy spectrum and decay curve are shown in Fig. 8. A single isomeric γ ray at 141.0 keV is observed, confirming the previous measurement in Ref. [27] in which a γ ray at 141 keV was observed. The 197-keV peak is attributed to ${}^{19}\text{F}(n,n'\gamma){}^{19}\text{F}$. The half-life is measured to be $0.098{}^{+0.010}_{-0.009} \,\mu$ s for the first time. We propose that the observed isomer in ${}^{82}\text{Ga}$ is an E2 isomer with spherical structure,

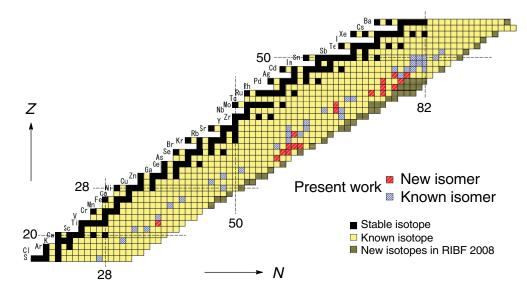


FIG. 6. (Color) Nuclear chart showing the observed isomers in the present work.

based on the systematics in neighboring nuclei. In this case the deduced B(E2) value is $4.09^{+0.43}_{-0.39}$ W.u., consistent with our interpretation. The possibility of other multipolarities is excluded similarly to the case of ${}^{59}\text{Ti}^m$. The ground state and the 141.0-keV isomeric state in ${}^{82}\text{Ga}^m$ are tentatively assigned as $J^{\pi} = 0^-$ and 2^- , respectively, according to the shell ordering in this region and the Nordheim rule for odd-odd spherical nuclei [48].

According to the β -decay spectroscopy of ⁸³Ge [45] and ⁸¹Ga [46], it can be assumed that the valence proton and neutron of the ground state of ⁸²Ga occupy the $\pi f_{5/2}$ and $\nu d_{5/2}$ orbitals, respectively. In such a configuration the Nordheim rule predicts that the 0⁻ coupling is energetically favorable for the ground state. Accordingly the isomeric state in ⁸²Ga has been assigned as $J^{\pi} = 2^{-}$. We propose that the configuration of the isomeric state is dominated by $[\pi f_{5/2} \otimes \nu s_{1/2}]2^{-}$, based on the ordering and spacing of the single-particle orbitals in the relevant neutron-rich region [45,46]. Furthermore it has been pointed out [45] that the $\nu s_{1/2}$ orbital in the odd-mass N = 51 isotones systematically goes down as the proton number decreases. The first $1/2^+$ state is located at 1031 keV

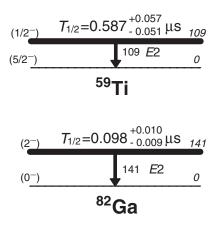


FIG. 7. Proposed level schemes of ⁵⁹Ti^m and ⁸²Ga^m.

for ⁸⁹Sr, 532 keV for ⁸⁷Kr, 462 keV for ⁸⁵Se, and 260 keV for ⁸³Ge. The present data suggest that the $\nu s_{1/2}$ orbital goes down further in ⁸²Ga, resulting in the occurrence of the *E*2 isomer in ⁸²Ga. Based on our observation, we cannot rule out the possibility that the $\nu s_{1/2}$ orbital is lower than the $\nu d_{5/2}$ orbital, in which case the J^{π} assignments to the ground and isomeric states would be reversed.

C. ⁹⁰As^m, ⁹²Se^m, ⁹³Se^m, and ⁹⁶Br^m

There is not enough information to propose level schemes for these new isomers, as we cannot rule out missing γ transitions due to low statistics. However, the discovery of these isomers provides us with future opportunities to

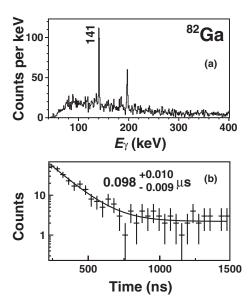


FIG. 8. (a) Delayed γ -ray energy spectrum and (b) time spectrum of ${}^{82}\text{Ga}^m$. The γ -ray peak at 197 keV is attributed to ${}^{19}\text{F}(n, n'\gamma){}^{19}\text{F}$ (see text). The deduced half-life is given in (b).

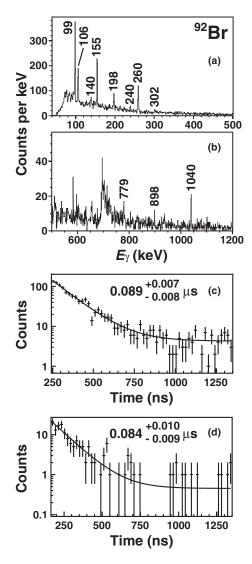


FIG. 9. (a), (b) Delayed γ -ray energy spectra and (c), (d) time spectra of ${}^{92}\text{Br}^m$. (a) and (b) show the low- and high-energy parts of the energy spectrum, respectively. Peaks are labeled by energies in keV. Unlabeled peaks in the energy spectra are the background γ rays described in text or unidentified γ rays. The time spectra were obtained using (c) the γ rays at 106.3, 154.9, and 301.9 keV for the 662-keV isomer and (d) those at 778.8, 898.3, and 1039.8 keV for the 1138-keV isomer. The deduced half-lives are given in the figures (see text).

investigate how the nuclear shape and shape coexistence evolve in the region of $N \sim 58-61$ [6] with decreasing Z number. Discussions on the evolution of nuclear shape in other neutron-rich nuclei in this region are given in the succeeding sections.

D. 9^2 Br^m

The γ -ray energy and time spectra of ${}^{92}\text{Br}^m$ are shown in Fig. 9. A total of eleven isomeric γ rays are observed, including two new γ rays at 197.6 and 898.3 keV. The halflife of ${}^{92}\text{Br}^m$ is determined for the first time. We find some disagreement with the previous measurement [27] as to the

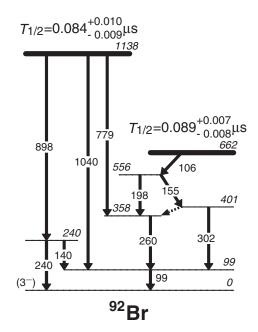


FIG. 10. Proposed level scheme of ${}^{92}\text{Br}^m$.

observed γ -ray peaks. Nine of the previously reported γ rays are confirmed, but those at 169 and 295 keV [27] are not identified in the present experiment. The level scheme of ${}^{92}\text{Br}^m$ was not proposed in Ref. [27], because they were unable to construct a plausible scheme using energy sum and intensity relations. In this previous experiment $\gamma\gamma$ coincidence was not measured.

Based on the energy sum analysis for the observed γ rays, we propose the level scheme shown in Fig. 10 for ${}^{92}\text{Br}^{m}$, in which two isometric states with similar half-lives are located at 662 and 1138 keV. The energy sum of the 197.6- and 259.5-keV transitions agrees well with that of the 154.9- and 301.9-keV transitions. The following energy sum relations are also found: 239.5 keV + 898.3 keV \approx 98.6 keV + 1039.8 keV \approx 98.6 keV + 259.5 keV + 778.8 keV \approx 1138 keV, 98.6 keV + 139.7 keV = 238.3 keV \approx 239.5 keV, and 259.5 keV + 778.8 keV = 1038.3 keV \approx 1039.8 keV. According to these relations, the observed γ rays are uniquely assigned to the level scheme. Furthermore, in the present experiment $\gamma\gamma$ coincidences are identified between the 98.6keV γ ray and those at 106.3, 154.9, 259.5, and 898.3 keV, between the 106.3-keV γ ray and those at 154.9 and 259.5 keV, and between the 259.5-keV γ ray and those at 154.9, 197.6, and 778.8 keV, consistent with the proposed level scheme. The intensity balance of γ rays seems fairly consistent as well. Note that we could not observe the 43-keV transition due to our energy threshold. The half-lives of the 662- and 1138-keV isomeric states are deduced to be $0.089^{+0.007}_{-0.008}$ and $0.084^{+0.010}_{-0.009} \ \mu$ s, respectively. Here we use the γ -ray peaks at 106.3, 154.9, and 301.9 keV for the 662-keV state and those at 778.8, 898.3, and 1039.8 keV for the 1138-keV state. The half-life of the 662-keV state was deduced assuming two decay components plus a constant background, because the background events of the 106.3-, 154.9-, and 301.9-keV γ rays contained Compton-scattered events originating from both

662- and 1138-keV states. We deduced the decay component of these background events by fitting the neighboring events on the high-energy side of the relevant γ -ray peak. The obtained decay parameter was fixed in the fitting analysis to deduce the half-life of the 662-keV state.

The RUL analysis restricts the multipolarity of the isomeric transition from the 662-keV state to E1, M1, or E2. The reduced transition probability reveals hindrance if the multipolarity is assigned as E1 or M1. The isomeric state at 1138 keV is likely due to high multipolarities of the transitions such as M2 or E3. In the neighboring odd-odd N = 57 isotone ⁹⁴Rb, two isomeric states with half-lives similar to those of ${}^{92}\text{Br}^m$ were recently identified at 1485.2 and 2074.8 keV, and it was proposed that they were spherical isomers caused by fully aligned configurations of the high-*j*, single-particle orbitals $\pi g_{9/2}$, $\nu g_{7/2}$, and $\nu h_{11/2}$ [54]. The ground state in ⁹⁴Rb is characterized by a configuration of $[\pi f_{5/2} \otimes \nu s_{1/2}]^{3-}$ and the 1485.2- and 2074.8-keV isomeric states are suggested to have spherical configurations of $[\pi g_{9/2} \otimes \nu g_{7/2}]^{8^+}$ and $[\pi g_{9/2} \otimes \nu h_{11/2}]^{10^-}$, respectively [54]. We speculate by analogy that the ground state in ⁹²Br has the same spherical configuration with $J^{\pi} = 3^{-}$ and the observed isomers in ⁹²Br are likely caused by a similar mechanism. Further experimental and theoretical studies are needed to clarify the nature of these isomers conclusively. Moreover, in the present experiment, an isomeric transition is also observed in ⁹⁰As for the first time, suggesting that similar isomerism may occur in this lighter N = 57 isotone.

E. 94 **Br**^{*m*}

A total of six new isomeric γ rays at 88.0, 92.1, 115.5, 165.7, 179.0, and 202.6 keV with a half-life of 0.530(15) μ s are identified in the γ -ray energy spectrum of ${}^{94}\text{Br}^m$. Our proposed partial level scheme is shown in Fig. 11. A $\gamma\gamma$ coincidence is observed between the 92.1- and 202.6-keV γ rays. A $\gamma\gamma$ coincidence between 115.5- and 179.0-keV γ rays is also observed, although the statistics are low. The energy sums of the cascading γ rays, which are 92.1 keV + 202.6 keV = 294.7 keV and 115.5 keV + 179.0 keV = 294.5 keV, agree well with each other. The 179.0-keV transition is placed in the upper part of the cascade because the 88.0-keV γ ray is observed. Inverse ordering of the cascading γ rays is also allowed, if we only rely on these results. The 165.7-keV γ ray cannot be assigned to the level scheme, although its decay curve is consistent with the half-life of ${}^{94}\text{Br}^m$.

In the neighboring N = 59 isotones, such as 95 Kr [55], 96 Rb [56], 97 Sr [57,58], 98 Y [59], and 99 Zr [60], the low-lying states systematically reveal spherical structure. *E*2 isomers with spherical structure were previously observed at the excitation energies below ~400 keV in the odd-odd isotone 98 Y as well as the even-odd isotones 95 Kr, 97 Sr, and 99 Zr. The present experiment confirms these isomers in 95 Kr and 97 Sr. We propose that the 295-keV state in 94 Br^m is an isomer with spherical structure, based on the systematics in these neighboring N = 59 isotones. The intensity balance analysis for the cascading 92.1- and 202.6-keV γ rays in 94 Br^m allows us to assign the multipolarity of the 92.1-keV transition to be *E*2, because it gives the most consistent

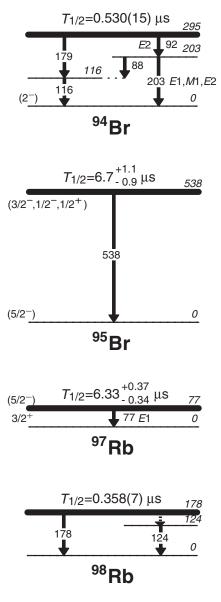


FIG. 11. Proposed level schemes of ${}^{94}\text{Br}^m$, ${}^{95}\text{Br}^m$, ${}^{97}\text{Rb}^m$, and ${}^{98}\text{Rb}^m$.

intensity balance after the correction of internal conversion. Here the multipolarity of the 202.6-keV γ ray is assumed to be *E*1, *M*1, or *E*2, because the higher multipolarities would give a longer half-life for the relevant state. Hence, taking into account the branching of γ rays, the *B*(*E*2) value of the 92.1-keV transition is estimated to be 2.49(24)–2.45(24) W.u., consistent with our proposal for ⁹⁴Br^m. Here multipolarities of *E*1 to *M*2 are considered for the 179.0-keV transition, based on the RUL. We speculate that the ground state in ⁹⁴Br has a spherical configuration of $[\pi f_{5/2} \otimes vs_{1/2}]2^-$ in analogy to the neighboring odd-odd N = 59 isotone ⁹⁶Rb [56]. Furthermore, in the present experiment, isomeric γ rays in ⁹³Se are also observed for the first time, implying the existence of a similar spherical isomer in this lighter N = 59 isotone.

In the N = 59 isotones ⁹⁶Rb [56], ⁹⁷Sr [57,58], ⁹⁸Y [59], and ⁹⁹Zr [60], prolate-deformed states were systematically identified above the low-lying spherical states, revealing shape

coexistence, and a band head of the deformed states, such as the [404]9/2⁺ states in ⁹⁷Sr [57,58] and ⁹⁹Zr [60], exhibits isomeric transitions. These deformed states have not been previously identified in ⁹⁵Kr [55]. In the present experiment we do not observe any isomeric γ rays from such deformed states in ⁹⁵Kr and ⁹⁴Br.

$\mathbf{F.}^{95}\mathbf{Br}^m$

The proposed level scheme for ⁹⁵Br^{*m*} is shown in Fig. 11. A single isomeric γ ray at 537.9 keV with a half-life of $6.7^{+1.1}_{-0.9} \mu s$ is observed for the first time. We propose that the 537.9-keV state in ⁹⁵Br is a prolate-deformed shape isomer located above the spherical ground state, based on the systematics of shape coexistence in the N = 60 isotones. These systematics will be discussed with ⁹⁷Rb^{*m*} in the next section. The reduced transition probability of the isomeric transition reveals very hindered values up to the *M*2 multipolarity, consistent with our proposed isomerism. The upper limit of possible multipolarity is *E*3, according to the RUL analysis. The N = 60 systematics allows us to tentatively assign the J^{π} of the ground state as $5/2^{-}$ and that of the 537.9-keV isomeric state as $[312]3/2^{-}$, $[310]1/2^{-}$, or $[440]1/2^{+}$. The details of the J^{π} assignments are also discussed in the next section.

G. ⁹⁷Rb^m

The proposed level scheme for ⁹⁷Rb^{*m*} is shown in Fig. 11. A single isomeric γ ray at 77.1 keV with a half-life of $6.33^{+0.37}_{-0.34}$ μ s is observed for the first time. The ground state of ⁹⁷Rb is known to exhibit a large static prolate deformation ($\varepsilon = 0.29$) characterized by the Nilsson orbit [431]3/2⁺ from spin and magnetic moments determined by laser spectroscopy [61]. The deformation is supported by the mass systematics in the relevant neutron-rich region [62] (see Fig. 1 in the reference). We propose that the 77.1-keV state in ⁹⁷Rb is a shape isomer, which has a spherical shape and decays to the deformed ground state, based on the systematics of shape coexistence in neighboring nuclei.

It is known that a sudden onset of prolate deformation occurs in the ground states of N = 60 isotones heavier than ⁹⁷Rb, which was observed as dramatic changes in the isotope shift and the two-neutron separation energy [62]. It is also known that shape coexistence occurs in the even-even N = 60isotones, in which a spherical second 0^+ state systematically appears above a prolate-deformed ground state at an excitation energy below 1 MeV [6]. Furthermore the spherical 0^+ state in these nuclei goes down with decreasing proton number [6]: e.g., 698 keV for 102 Mo [63], 331 keV for 100 Zr [8], 215 keV for 98 Sr [7], and the ground state of 96 Kr does not exhibit such deformation according to the mass systematics [62] and spectroscopic studies [64]. This systematic behavior of N = 60 nuclei imply that the energy levels of spherical and deformed states get closer as the proton number decreases and reverses between ⁹⁸Sr and ⁹⁶Kr. We speculate that the ⁹⁷Rb isotope is a transitional nucleus in which the spherical and deformed structures energetically compete with each other at a low excitation energy. It should be noted that similar shape

coexistence was suggested in the odd N = 60 isotone ⁹⁹Y, in which a spherical state is located at 599 keV above the deformed ground state [65]. We tentatively assign the J^{π} of the 77.1-keV state in ⁹⁷Rb as 5/2⁻, based on the J^{π} assignment of a spherical ground state in ⁹⁵Rb [66]. Accordingly the multipolarity of the 77.1-keV transition has been assigned as *E*1, which reveals a very hindered *B*(*E*1) value, consistent with our proposed isomerism. A little information is available for excited states in ⁹⁶Kr. Their investigations, including the search for deformed excited states in ⁹⁶Kr, are awaited to further clarify the evolution of shape coexistence in the N = 60isotones. In the present measurement we did not observe any isomeric transitions in ⁹⁶Kr although the statistics were as high as those of ⁹⁵Kr^m.

Based on the above discussion of N = 60 systematics including ⁹⁷Rb^{*m*}, we propose the level scheme shown in Fig. 11 for ⁹⁵Br^{*m*}, with the spherical and deformed states reversed. The J^{π} assignment to the deformed state at 538 keV is made based on the Nilsson diagram in Fig. 7 of Ref. [67], assuming that the state has the same prolate-deformation parameter as ⁹⁷Rb ($\varepsilon = 0.29$).

H. 98 Rb^m

The proposed level scheme for ${}^{98}\text{Rb}^m$ is shown in Fig. 11 and the γ -ray energy spectrum and decay curve are shown in Fig. 12. Three isomeric γ rays at 116, 124, and 178 keV with a half-life of 700^{+60}_{-50} ns were previously reported [27] for ${}^{98}\text{Rb}^m$. In the present work only two isomeric γ rays at 123.7 and 178.4 keV with a half-life of 0.358(7) μ s are observed. The 116-keV γ ray is not confirmed, even though our photopeak counts of the 123.7-keV γ ray are approximately ten times more than those reported in Ref. [27]. The 116-keV γ ray should have been observed in this experiment if it originates

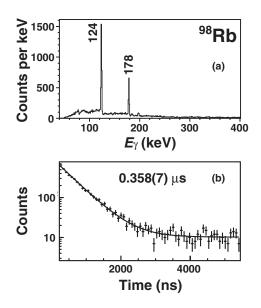


FIG. 12. (a) Delayed γ -ray energy spectrum and (b) time spectrum of ${}^{98}\text{Rb}^m$. Unlabeled peaks in the energy spectrum are the background γ rays described in text or unidentified γ rays. The deduced half-life is given in (b).

from the same isomeric state. We propose the level scheme shown in Fig. 11 for ${}^{98}\text{Rb}^m$, because a $\gamma\gamma$ coincidence is not observed between the 123.7- and 178.4-keV γ rays. The statistics of the present measurement can rule out that the γ rays are in a cascade.

In the neighboring N = 61 isotones, such as ⁹⁹Sr [68], ¹⁰⁰Y [69], ¹⁰¹Zr [70], and ¹⁰²Nb [69], the low-lying states systematically reveal prolate deformation. Based on the RUL analysis for reduced transition probability, we propose that the 178.4-keV state in ⁹⁸Rb is likely a shape isomer. The RUL analysis considering the branching of transitions restricts the multipolarities of the 178.4-keV transition and the unobserved 54.7-keV transition to E1, M1, or E2. The reduced transition probability reveals hindered values for all the possible multipolarities for the 178.4-keV transition, supporting our proposed isomerism. In the case of the unobserved 54.7-keV transition, the transition probability would have a significant hindrance except for an E2 multipolarity. The present data cannot rule out that the ordering of the unobserved 54.7-keV transition and the 123.7-keV transition could be reversed. However this does not affect the above discussion. In the present experiment, an isomeric γ ray is also observed in ⁹⁶Br for the first time, implying that similar isomerism occurs in this lighter N = 61isotone. Further investigations are needed to clarify the nuclear isomerism in the N = 61 isotones conclusively.

I. 108 Zr^m

The γ -ray energy and time spectra of ${}^{108}\text{Zr}^m$ are shown in Fig. 13 and our proposed level scheme for ${}^{108}\text{Zr}^m$ is shown in Fig. 14. A total of 13 isomeric γ rays at 174.4, 277.9, 347.2, 364.2, 426.3, 432.3, 478.3, 485.3, 603.9, 642.3, 773.6, 795.9, and 828.3 keV with a half-life of $0.536^{+0.026}_{-0.025}$ μ s are observed in this isomer. Among them the observation of five γ rays at 174.4, 277.9, 347.2, 478.3, and 603.9 keV was previously reported along with the half-life value of 620(150) ns [43]. The authors of Ref. [43] proposed a level scheme of rotational-like band up to 4⁺ consisting of the 174.4- and 347.2-keV transitions, based on the systematics of the 2⁺₁ and 4⁺₁ states in the lighter even-even Zr isotopes, although a $\gamma\gamma$ coincidence was not observed. In the present experiment we observe a $\gamma\gamma$ coincidence between the 174.4- and 347.2-keV γ rays, supporting their proposed level scheme.

We extended the level scheme based on the energy and intensity relation as well as the $\gamma\gamma$ coincidence information. All the observed γ rays are uniquely assigned to the level scheme in Fig. 14, including the ground-state rotational band which is extended up to the 8⁺ state. The energy sum analysis strongly supports the proposed level scheme, although there is some inconsistency in the relative intensity of the 426.3-keV γ ray. We tentatively assign the isomeric state of ¹⁰⁸Zr to the level at 2075 keV. The recent theoretical calculation [71] predicts that a two-quasineutron state with $J^{\pi} = 6^+$ appears at 1.997 MeV in ¹⁰⁸Zr. Our proposed isomeric state at 2075 keV is located close to this predicted high-*K* state, implying that the isomeric state in ¹⁰⁸Zr could be a *K* isomer. More details will be presented in our separate paper.

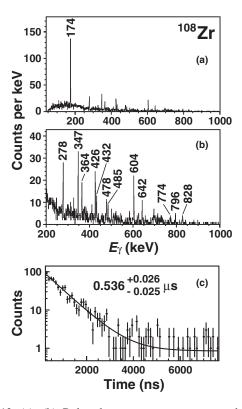


FIG. 13. (a), (b) Delayed γ -ray energy spectra and (c) time spectrum of ¹⁰⁸Zr^m. (b) shows a partially enlarged energy spectrum. Unlabeled peaks in the energy spectra are the background γ rays described in text or unidentified γ rays. The deduced half-life is given in the (c).

J. ¹⁰⁸Nb^m

The proposed level scheme for ¹⁰⁸Nb^m is shown in Fig. 15. Four new isomeric γ rays at 64.3, 77.6, 89.0, and 102.2 keV with a half-life of 0.109(2) μ s are observed. We clearly observe a $\gamma\gamma$ coincidence between the 77.6- and 89.0-keV γ rays as shown in Fig. 4. A $\gamma\gamma$ coincidence is also observed between the 64.3- and 102.2-keV γ rays. The energy sums of these cascading γ rays, which are 77.6 keV + 89.0 keV = 166.6 keV and 64.3 keV + 102.2 keV = 166.5 keV, agree well with each other. Furthermore a $\gamma\gamma$ coincidence between the 64.3- and 89.0-keV γ rays is observed. If we only rely on these results, reversed ordering of the cascading γ rays is also possible.

The ground state of ¹⁰⁸Nb is suggested to be a prolatedeformed state with $J^{\pi} = 2^+$ and characterized by a configuration of $\pi [422]5/2^+ \nu [411]1/2^+$, according to the recent β -decay spectroscopy of ¹⁰⁸Nb [72]. We propose to assign the multipolarities of the cascading 77.6-keV and 89.0-keV γ rays as *E*1 or *M*1, and *E*2, respectively, because such assignments give the most consistent intensity balance. If the ordering of these two γ rays is that shown in Fig. 15, the 89.0-keV isomeric *E*2 transition has a normal value in the Weisskopf estimate, suggesting that the 167-keV isomeric state in ¹⁰⁸Nb is due to the small transition energy between the deformed states in different bands.

If the ordering is reversed, the 77.6-keV isomeric transition becomes a hindered E1 or M1. In this case we propose that the 167 keV is an oblate-deformed shape isomer which decays

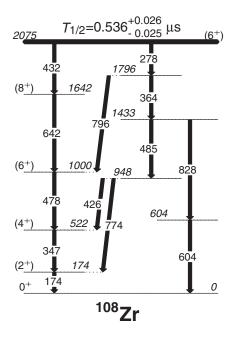


FIG. 14. Proposed level scheme of 108 Zr^{*m*}.

to the prolate-deformed states. In fact such shape coexistence has been recently suggested in the neighboring isomer ¹⁰⁹Nb^m [42]. Here the transition probability of the 89.0-keV *E*2 transition should have an enhanced value, because otherwise the corresponding state would have a long half-life. Further experimental studies are needed to propose the level scheme and the nuclear isomerism conclusively. The branching of the isomeric transitions was taken into account in deducing the transition probability.

K. ¹⁰⁹Nb^m, ¹¹²Tc^m, and ¹¹³Tc^m

It was recently suggested in Ref. [42] that the isomer ¹⁰⁹Nb^m is an oblate-deformed shape isomer located above prolatedeformed states. In the present experiment three isomeric γ rays at 117.0, 196.2, and 312.2 keV are observed, confirming all the previously reported γ rays in ¹⁰⁹Nb^m [42]. The present measurement has approximately ten times more statistics than the previous one, so that the precision of the half-life is

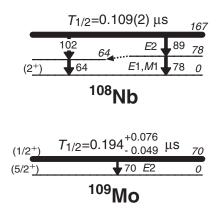


FIG. 15. Proposed level schemes of ¹⁰⁸Nb^m and ¹⁰⁹Mo^m.

improved significantly. Furthermore the improved statistics allow us to observe a $\gamma\gamma$ coincidence between 117.0- and 196.2-keV transitions for the first time, supporting the level scheme proposed in Ref. [42]. In the case of ¹¹²Tc^m and ¹¹³Tc^m, we confirm the previously reported isomeric γ rays: the 92- and 258-keV γ rays in ¹¹²Tc^m [41] and the 114-keV γ ray in ¹¹³Tc^m [41]. The isomeric states in ¹¹²Tc^m and ¹¹³Tc^m are suggested to be a triaxial-deformed shape isomer located above oblate-deformed states [41].

L. ¹⁰⁹Mo^m

The proposed level scheme for ¹⁰⁹Mo^{*m*} is shown in Fig. 15. A single γ -ray peak at 69.7 keV with a half-life of $0.194^{+0.076}_{-0.049}$ μ s is observed for the first time. The level scheme of ¹⁰⁹Mo was previously investigated by measuring prompt γ rays following spontaneous fission of ²⁴⁸Cm [73]. In that experiment the ground-state rotational band and the rotational band built on the 222-keV state were identified for ¹⁰⁹Mo, and the ground state of ¹⁰⁹Mo was suggested to be prolate deformed and characterized by the Nilsson orbit [402]5/2⁺ [73].

The RUL analysis limits the multipolarity of the isomeric transition up to E2. We propose that the isomeric state in ¹⁰⁹Mo is an *E*2 isomer whose transition takes place between the band heads of different deformed bands, in analogy to the neighboring isomer ${}^{107}Mo^m$ in which a similar isomeric transition was observed between the 65.4-keV state and the ground state [11]. According to Ref. [11], the ground state and the 65.4-keV state in ¹⁰⁷Mo^m are the band heads of the $5/2^+$ rotational band and the $1/2^+$ rotational band, respectively, and the 65.4-keV transition is an E2 transition with B(E2) = 5.9(5) W.u. The configurations of these states are characterized by $[413]5/2^+$ and $[411]1/2^+$, respectively. Furthermore these two bands were suggested to have similar prolate and triaxial deformation parameters [11]. In the case of ¹⁰⁹Mo^{*m*} the deduced B(E2) value is $10.5^{+4.2}_{-2.8}$ W.u., which is similar to that of 107 Mo^{*m*}, supporting our proposed isomerism. We tentatively assign the J^{π} of the isomeric state in ¹⁰⁹Mo as $1/2^+$, of which configuration may be characterized by $[411]1/2^+$. If the multipolarity is E1 or M1, the isomeric state in ¹⁰⁹Mo can be interpreted as a shape isomer since the transition probability reveals a hindered value. A shape isomer is suggested in the neighboring isomer 109 Nb^{*m*} [42].

M. ¹¹⁷Ru^m and ¹¹⁹Ru^m

The proposed level scheme for ¹¹⁷Ru^{*m*} is shown in Fig. 16. A total of five new isomeric γ rays at 57.8, 82.5, 102.9, 127.4, and 184.6 keV with a half-life of $2.487^{+0.058}_{-0.055} \mu$ s are observed in the case of ¹¹⁷Ru^{*m*}. A $\gamma\gamma$ coincidence is clearly observed between the 82.5- and 102.9-keV γ rays. A $\gamma\gamma$ coincidence between the 57.8- and 127.4-keV γ rays is also observed, although the statistics are low. The energy sum of these cascading γ rays agrees with each other as well as with the energy of the 184.6-keV γ -ray transition. Inverse ordering of the cascading γ rays is also possible.

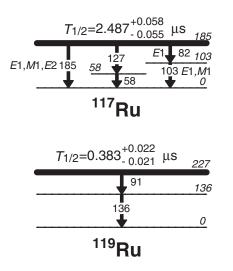


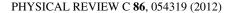
FIG. 16. Proposed level schemes of ¹¹⁷Ru^m and ¹¹⁹Ru^m.

We propose that the isomerism of 117 Ru^{*m*} is due to shape coexistence. The RUL analysis considering the branching of γ rays allows us to restrict the multipolarity of the 184.6-keV transition to E1, M1, or E2, each of which gives a hindered value for the transition probability, consistent with shape isomerism. The intensity balance analysis for the cascading 82.5- and 102.9-keV γ rays restricts the multipolarity of the 82.5-keV transition to E1 and that of the 102.9-keV transition to E1 or M1. The analysis also allows a combination of E2 and M2, but such possibility is excluded because the 102.9-keV state would have a long half-life. The 82.5-keV transition has a hindered transition probability if an E1 multipolarity is assumed. These results for the 184.6- and 82.5-keV transitions support our proposed isomerism of ¹¹⁷Ru^m. We note that inverse ordering of the 82.5- and 102.9-keV transitions does not affect the proposed isomerism.

In the case of ¹¹⁹Ru^{*m*}, two new isomeric γ rays at 90.8 and 136.3 keV with a half-life of $0.383^{+0.022}_{-0.021}$ µs are observed. We propose the level scheme shown in Fig. 16 for 119 Ru^{*m*}, based on an observed $\gamma\gamma$ coincidence between the two γ rays (see Fig. 5). The present data cannot rule out that the ordering of the 90.8- and 136.3-keV transition can be reversed. The RUL analysis allows us to restrict the multipolarity of isomeric transition to E1, M1, or E2, regardless of the ordering of these γ rays. In the case of E1 or M1, the transition probability reveals a hindered value, suggesting that ¹¹⁹Ru^m is a shape isomer. However, if it is an E2 transition, the transition probability is not hindered. Our intensity balance analysis indicates that the hindered E1 or M1 transition is more likely, although the statistics are low. Further experimental investigations are needed to determine the isomerism of ¹¹⁹Ru^{*m*} conclusively.

N. 120 Rh^{*m*}, 122 Rh^{*m*}, and 124 Ag^{*m*}

The following new isomeric γ rays are observed in the delayed γ -ray energy spectra: the 59.1- and 98.1-keV γ rays with a half-life of $0.294^{+0.016}_{-0.015} \ \mu s$ in $^{120}\text{Rh}^m$, the 63.9- and 207.1-keV γ rays with a half-life of $0.82^{+0.13}_{-0.11} \ \mu s$ in $^{122}\text{Rh}^m$,



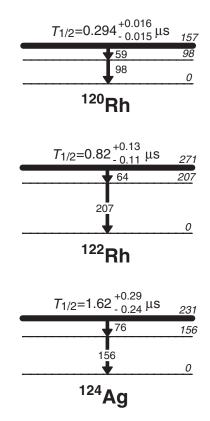


FIG. 17. Proposed level schemes of ¹²⁰Rh^m, ¹²²Rh^m, and ¹²⁴Ag^m.

and the 75.5- and 155.6-keV γ rays with a half-life of $1.62^{+0.29}_{-0.24}$ μ s in 124 Ag^m. We propose similar level schemes for these neighboring odd-odd isomers as shown in Fig. 17, because the γ -ray energy spectra resemble each other. In the case of 122 Rh^m, a $\gamma\gamma$ coincidence is observed between the 63.9- and 207.1-keV γ rays, although the statistics are low, supporting the proposed level scheme. In all cases the present data cannot rule out that the ordering of the transitions may be reversed. In the neutron-rich odd-odd Ag isotopes such as 122 Ag [74], β -decaying long-lived isomers appear systematically at low excitation energies. The nuclide 124 Ag may have such a long-lived isomer that is fed by the isomeric γ rays observed in the present experiment.

We do not have sufficient statistics to perform an intensity balance analysis for ¹²⁰Rh^m, ¹²²Rh^m, and ¹²⁴Ag^m. However, according to the RUL analysis for these isomers, E1, M1, or E2 are the most likely for the multipolarity of the isomeric transition, regardless of the ordering of the γ rays. The transition probability reveals hindered values except for in some instances the case of E2 multipolarity.

O. ¹²¹Pd^{m1,m2}

A new isomeric γ ray at 135.5 keV is observed in the delayed γ -ray spectrum of ¹²¹Pd. The time spectrum shows a bump at ~500 ns, not consistent with single-component decay (see Fig. 3). Above the 135.5-keV isomeric transition, there should be unobserved isomeric transitions with a γ -ray energy below our energy threshold or with a large internal conversion

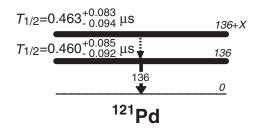


FIG. 18. Proposed level scheme of ¹²¹Pd^m.

coefficient, in which case the γ -ray intensity is too low to be observed in this experiment. If the ordering is reversed, the 135.5-keV transition should have a single decay component. Based on these results and also assuming two cascading-decay components, we propose the level scheme shown in Fig. 18 for $^{121}Pd^{m1,m2}$. The fitting analysis of the time spectrum allows us to determine the half-lives of the 135.5-keV transition $^{121}Pd^{m1}$ and the unobserved transition $^{121}Pd^{m2}$ as $0.460^{+0.083}_{-0.092} \ \mu$ s and $0.463^{+0.083}_{-0.094} \ \mu$ s, respectively. The population probabilities of these two isomers are determined in the fitting analysis. The RUL analysis allows us to restrict the multipolarity of the 135.5-keV transition to E1, M1, or E2. In the case of E1 or M1, the transition probability reveals a hindered value, while it is not hindered if the multipolarity is E2. Further experimental studies are needed to propose the level scheme conclusively.

P. ¹²⁵Ag^m

The delayed γ -ray energy and time spectra of ¹²⁵Ag^m are shown in Fig. 19 and our proposed level scheme for ${}^{125}Ag^m$ is shown in Fig. 20. Four isomeric γ rays at 672, 686, 717, and 731 keV with a half-life of 473(111) ns were previously reported for ${}^{125}Ag^m$ and a level scheme was proposed [75]. In the present experiment seven isomeric γ rays at 103.2, 669.8, 684.0, 714.5, 728.4, 765.1, and 786.7 keV with a half-life of 0.498^{+0.021}_{-0.020} μ s are identified. All the previously reported γ rays are confirmed, although the γ -ray energies systematically differ by 2-3 keV. Those at 103.2, 765.1, and 786.7 keV are new isomeric γ rays. The energy sum of the 684.0- and 714.5keV transitions agrees with that of the 669.8- and 728.4-keV transitions, supporting the existence of the 1399-keV state. The half-life deduced for the strongest γ -ray peak at 103.2 keV is in agreement with the one derived from the summed events of the previously reported γ rays. We place the 103.2-keV transition above the 1399-keV state as the isomeric transition, with the isomeric state located at 1502 keV. The energy sum of the 103.2-, 684.0-, and 714.5-keV transitions agrees with that of the 786.7- and 714.5-keV transitions, supporting the placement of the 103.2-keV transition as well as that of the 714.5-keV transition. The 765.1-keV γ ray cannot be assigned in the proposed level scheme.

The 1502-keV state in ¹²⁵Ag^{*m*} is likely an *E*2 isomer having spherical structure, based on the intensity balance analysis. The most consistent intensity balance is obtained between the 103.2-keV γ ray and the two γ rays in the cascades, if an *E*2 multipolarity is assumed for the 103.2-keV transition. The deduced *B*(*E*2) value is 1.08(12) W.u. for

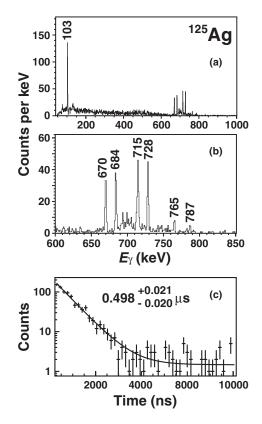


FIG. 19. (a), (b) Delayed γ -ray energy spectra and (c) time spectrum of ¹²⁵Ag^m. (b) shows a partially enlarged energy spectrum. Unlabeled peaks in the energy spectra are the background γ rays described in text or unidentified γ rays. The deduced half-life is given in (c).

this isomeric transition, supporting our proposed isomerism. Here the branching of the 103.2- and 786.7-keV transitions is taken into consideration. The J^{π} assignments to the energy levels in ¹²⁵Ag^m are tentatively made based on the theoretical

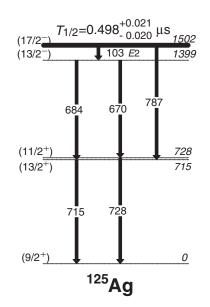


FIG. 20. Proposed level scheme of ${}^{125}Ag^m$.

predictions given in Ref. [75]. The ordering of the 669.8- and 728.4-keV transitions could be reversed.

Q. ¹²⁴Pd^m and ¹²⁶Ag^m

New isomeric γ -ray peaks are observed at 62.2 keV in 124 Pd^m and at 254.8 keV in 126 Ag^m. Their half-lives are estimated to be >20 μ s, because the isomeric γ -ray events are equally distributed within the 20- μ s range of the time spectra. The statistics are not sufficient to propose a level scheme.

R. Other isomers

All the previously reported isomeric γ rays in the following known isomers are identified in the present experiment: ⁴³S^m [76], 54 Sc^{*m*} [20,44], 60 V^{*m*1,*m*2} [77], 76 Ni^{*m*} [2], 75 Cu^{*m*1,*m*2} [77], ⁷⁸Zn^{*m*} [3], ⁹⁵Kr^{*m*} [55], ¹⁰⁶Nb^{*m*} [78], ¹²⁵Cd^{*m*} [79], ¹²⁷Cd^{*m*} [80], ¹²⁸Cd^{m1,m2} [23], ¹³⁰Cd^m [4], ¹²⁹In^m [81], ¹³⁰In^m [82], ¹³¹In^m [83], ¹³⁰Sn^m [84], ¹³⁴Sn^m [85], and ¹³⁶Sb^m [22,86] as listed in Table II, and overall we find good agreement with the previous measurements, including the relative γ -ray intensities and the level schemes. The previously reported $\gamma\gamma$ coincidences in ⁷⁸Zn^m [3], ⁹⁵Kr^m [55], ¹²⁹In^m [81], and ¹³⁴Sn^m [85] are confirmed. A $\gamma\gamma$ coincidence is observed between the 53.7and 173.0-keV γ rays in ¹³⁶Sb^{*m*}, supporting the level scheme proposed in Refs. [22,86]. The relative γ -ray intensities are deduced for ${}^{60}V^{m1,m2}$ [77], ${}^{75}Cu^{m1,m2}$ [77], and ${}^{106}Nb^m$ [78] for the first time. In the case of ${}^{60}V^{m1,m2}$, the observed intensity of the upper isomeric transition ${}^{60}V^{m2}$ is approximately two times higher than that of the lower transition ${}^{60}V^{m1}$ (see Table III). This implies the possibility of reversing the previously reported ordering of these two transitions [77].

Overall agreement with the previous measurements is also achieved in the case of the following known isomers: 50 K^m [77], 56 Sc^m [44], 64 Mn^m [20,77,87], 67 Fe^m [88], 96 Rb^m [56], 97 Sr^m [57,58], 100 Sr^m [10], 101 Y^m [27], 128 Cd^{m3} [23], 131 Sn^m [89], and 132 Sn^m [90], although some of the previously reported γ rays are not observed as seen in Table II. This disagreement on observed γ -ray peaks can be attributed to our limited statistics and energy threshold. We find disagreement with the previously reported half-life of 101 Y^m [27]: The present measurement is approximately five times shorter than the previously reported value.

IV. SUMMARY AND CONCLUSIONS

The present experiment includes the measurement of new and known isomers with $N \sim 58-61$ and $Z \sim 34-39$. It provides us with the opportunity to observe and elucidate the evolution of shape coexistence in the corresponding region of neutron-rich exotic nuclei. Previously it was shown that prolate-deformed states coexist with spherical states in this region [6]. In the present work we propose that the isomers ${}^{95}\text{Br}^m$, ${}^{97}\text{Rb}^m$, and ${}^{98}\text{Rb}^m$ are shape isomers generated due to such shape coexistence. Furthermore the energy difference between the prolate-deformed and spherical states systematically gets smaller as the proton number decreases along the PHYSICAL REVIEW C 86, 054319 (2012)

N = 60 line and their ordering could even be reversed. The isomeric state at 77.1 keV in ⁹⁷Rb^m is particularly interesting, because the 77.1-keV state and the ground state are located very close to each other, implying that the ⁹⁷Rb is likely a transitional nucleus. The ordering of the prolate-deformed and spherical states is reversed in our proposed level scheme of ⁹⁵Br^m. Further investigations including theoretical studies, particularly for the Se and Br isotopes, are needed to fully understand the evolution of shape coexistence and the nuclear isomerism in this region.

In the region of $N \sim 65-70$ and $Z \sim 40-43$, a variety of nuclear shapes and their coexistence are involved in the appearance of the isomers observed in the present experiment. The nuclear isomerism of 108 Zr^{*m*} was previously discussed in terms of a tetrahedral shape [43]. In the present work we propose the existence of a K isomer for 108 Zr^m instead, based on our expended level scheme and the recent theoretical prediction [71]. A low-lying oblate shape isomer was proposed in Ref. [42] in the case of 109 Nb^{*m*}, where the ground state is known to be of a prolate shape. Triaxial-deformed shape isomers were proposed above oblate-deformed states in the cases of ${}^{112}\text{Tc}^m$ and ${}^{113}\text{Tc}^m$ [41]. In the present work we propose two possibilities for the new isomer ${}^{108}Nb^m$: an E2 isomer caused by a small transition energy between deformed states in different bands or an oblate-deformed shape isomer located above a prolate-deformed state. We also propose that the new isomer ${}^{109}Mo^m$ is an E2 isomer resulting from the transition between deformed band heads, in analogy to the isomerism of neighboring ¹⁰⁷Mo^m [11]. These results indicate a rich variety of phenomena in this region and that more comprehensive information is needed to fully elucidate the nuclear structure.

The isomers in the region of $N \sim 75$ and $Z \sim 44-47$ were not explored in any detail prior to the present experiment. The present observation of the new isomers in ¹¹⁷Ru^m, 119 Ru^m, 120 Rh^m, 122 Ru^m, 121 Pd^m, 124 Pd^m, 124 Ag^m, and 126 Ag^m provide us with clues to understanding isomerism and the corresponding nuclear structure in this region. We propose ¹¹⁷Ru^m to be a shape isomer, based on the hindrance of its transition probability. The possible isomerism for ¹¹⁹Ru^m, 120 Rh^m, 122 Ru^m, 121 Pd^m, and 124 Ag^m is also discussed. The γ ray energy spectra of the observed isomers are similar to each other, implying the presence of a common mechanism. The occurrence of nuclear deformation in the region of $Z \sim 39-45$ and $N \sim 75$ was pointed out in Ref. [15] to be of astrophysical interest. The mass model based on the extended Thomas-Fermi plus Strutinsky integral (ETFSI) method indicates the occurrence of deformation in the relevant region [91]: The systematics of two-neutron separation energies exhibit humps around $N \sim 75$ in the isotopic chains of $Z \sim 39-45$ (see Fig. 3(c) in Ref. [91]). This behavior is very similar to what was observed around $N \sim 60$ in the isotopic chain of $Z \sim 38-42$ [62], where shape coexistence systematically appears. This could give rise to the shape isomers observed in the present work. We speculate that the isomers in the region of $N \sim 75$ could be generated in a similar way. The relevant neutron-rich region could be a new deformed region exhibiting shape coexistence. Further experimental and theoretical studies are needed to understand the nuclear structure and isomerism in this neutron-rich region.

Finally, isomers in a lighter-mass region with $N \sim 28$ – 50 are also observed and investigated in the present work, providing us with an opportunity to study shell evolution in the relevant neutron-rich exotic nuclei. For instance, the existence of the isomer ⁵⁹Ti^m suggests that the N = 34 subshell gap between the $\nu p_{1/2}$ and $\nu f_{5/2}$ orbitals gets smaller as the neutron number increases. Such shell evolution can be attributed to the attractive monopole interaction between the $\pi f_{7/2}$ and $\nu f_{5/2}$ orbitals [92].

In summary, we have conducted a search for new isomers using in-flight fission of a ²³⁸U beam at 345 MeV/nucleon, and a number of neutron-rich microsecond isomers were identified over a wide range of atomic numbers, including the discovery of 18 new isomers. A wealth of spectroscopic information was obtained, allowing us to present proposed level schemes for 17 neutron-rich exotic nuclei far from stability. In the present experiment we were able to explore nuclear isomerism over a wide range of neutron-rich exotic nuclei, allowing the investigation of the evolution of nuclear structure and shape coexistence. The experiment was made possible by the use of in-flight fission as well as the features of the BigRIPS separator which were fully exploited. These results demonstrate the potential of RIBF, which will allow isomer spectroscopy over a wide range of neutron-rich exotic nuclei far beyond the limits of conventional means.

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- [1] J. A. Pinston and J. Genevey, J. Phys. G 30, R57 (2004).
- [2] C. Mazzocchi *et al.*, Phys. Lett. B **622**, 45 (2005), and references therein.
- [3] J. M. Daugas et al., Phys. Lett. B 476, 213 (2000).
- [4] A. Jungclaus et al., Phys. Rev. Lett. 99, 132501 (2007).
- [5] P. Möller, A. J. Sierk, R. Bengtsson, H. Sagawa, and T. Ichikawa, At. Data Nucl. Data Tables 98, 149 (2012).
- [6] K. Heyde and J. Wood, Rev. Mod. Phys. 83, 1467 (2011).
- [7] F. Schussler, J. A. Pinston, E. Monnand, A. Moussa, G. Jung, E. Koglin, B. Pfeiffer, R. V. F. Janssens, and J. van Klinken, Nucl. Phys. A 339, 415 (1980).
- [8] F. K. Wohn, J. C. Hill, C. B. Howard, K. Sistemich, R. F. Petry, R. L. Gill, H. Mach, and A. Piotrowski, Phys. Rev. C 33, 677 (1986), and references therein.
- [9] R. A. Meyer et al., Nucl. Phys. A 439, 510 (1985).
- [10] B. Pfeiffer, G. Lhersonneau, H. Gabelmann, and K. L. Kratz, Z. Phys. A 353, 1 (1995).
- [11] J. Pinston et al., Phys. Rev. C 74, 064304 (2006).
- [12] J. Skalski, S. Mizutori, and W. Nazarewicz, Nucl. Phys. A 617, 282 (1997).
- [13] P. Moller, A. J. Sierk, R. Bengtsson, H. Sagawa, and T. Ichikawa, Phys. Rev. Lett. **103**, 212501 (2009).
- [14] J. P. Delaroche, M. Girod, J. Libert, H. Goutte, S. Hilaire, S. Peru, N. Pillet, and G. F. Bertsch, Phys. Rev. C 81, 014303 (2010).
- [15] H. Grawe, K. Langanke, and G. Martinez-Pinedo, Rep. Prog. Phys. 70, 1525 (2007).
- [16] B. Sun, F. Montes, L. S. Geng, H. Geissel, Y. A. Litvinov, and J. Meng, Phys. Rev. C 78, 025806 (2008).
- [17] J. P. Dufour *et al.*, Nucl. Instrum. Methods Phys. Res. A 248, 267 (1986).
- [18] R. Grzywacz et al., Phys. Lett. B 355, 439 (1995).
- [19] R. Grzywacz et al., Phys. Rev. C 55, 1126 (1997).

- [20] R. Grzywacz et al., Phys. Rev. Lett. 81, 766 (1998).
- [21] M. Pfützner et al., Phys. Lett. B 444, 32 (1998).
- [22] M. N. Mineva et al., Eur. Phys. J. A 11, 9 (2001).
- [23] L. Cáceres et al., Phys. Rev. C 79, 011301(R) (2009).
- [24] R. L. Lozeva et al., Phys. Rev. C 77, 064313 (2008).
- [25] M. Bernas et al., Phys. Lett. B 331, 19 (1994).
- [26] M. Bernas et al., Phys. Lett. B 415, 111 (1997).
- [27] C. M. Folden et al., Phys. Rev. C 79, 064318 (2009).
- [28] T. Kubo, Nucl. Instrum. Methods Phys. Res. B 204, 97 (2003).
- [29] T. Kubo et al., IEEE Trans. Appl. Supercond. 17, 1069 (2007).
- [30] Y. Yano, Nucl. Instrum. Methods Phys. Res. B 261, 1009 (2007).
- [31] T. Ohnishi et al., J. Phys. Soc. Jpn. 77, 083201 (2008).
- [32] T. Ohnishi et al., J. Phys. Soc. Jpn. 79, 073201 (2010).
- [33] J. Lindhard and A. H. Sorensen, Phys. Rev. A 53, 2443 (1996); ATIMA, http://www-linux.gsi.de/~weick/atima/.
- [34] H. Kumagai, A. Ozawa, N. Fukuda, K. Summerer, and I. Tanihata, Nucl. Instrum. Methods Phys. Res. A 470, 562 (2001).
- [35] C. Amsler et al., Phys. Lett. B 667, 1 (2008).
- [36] Table of Isotopes, 8th ed., Vol. I, edited by R. B. Firestone and V. S. Shirley (Wiley, New York, 1996).
- [37] D. C. Radford, RADWARE software package, available at http://radware.phy.ornl.gov/.
- [38] B. E. Tomlin, P. F. Mantica, and W. B. Walters, Eur. Phys. J. Special Topics 150, 183 (2007).
- [39] W. B. Walters et al., Phys. Rev. C 70, 034314 (2004).
- [40] L. Gaudefroy et al., Eur. Phys. J. A 23, 41 (2005).
- [41] A. M. Bruce et al., Phys. Rev. C 82, 044312 (2010).
- [42] H. Watanabe et al., Phys. Lett. B 696, 186 (2011).
- [43] T. Sumikama et al., Phys. Rev. Lett. 106, 202501 (2011).

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- [44] H. L. Crawford et al., Phys. Rev. C 82, 014311 (2010).
- [45] O. Perru *et al.*, Eur. Phys. J. A 28, 307 (2006).
- [46] D. Verney et al., Phys. Rev. C 76, 054312 (2007).
- [47] GEANT—Detector description and simulation tool, CERN Program Library Long Write-up W5013, CERN Geneva.
- [48] L. W. Nordheim, Phys. Rev. 78, 294 (1950).
- [49] T. Kibedi, T. Burrows, M. Trzhaskovskaya, P. Davidson, and C. Nestorjr, Nucl. Instrum. Methods Phys. Res. A 589, 202 (2008).
- [50] P. M. Endt, At. Data Nucl. Data Tables 23, 547 (1979).
- [51] P. M. Endt, At. Data Nucl. Data Tables 26, 47 (1981).
- [52] P. Maierbeck et al., Phys. Lett. B 675, 22 (2009).
- [53] S. Liddick et al., Phys. Rev. C 72, 054321 (2005).
- [54] I. Tsekhanovich et al., Phys. Rev. C 78, 011301(R) (2008).
- [55] J. Genevey, R. Guglielmini, R. Orlandi, J. A. Pinston, A. Scherillo, G. Simpson, I. Tsekhanovich, N. Warr, and J. Jolie, Phys. Rev. C 73, 037308 (2006).
- [56] J. A. Pinston, J. Genevey, R. Orlandi, A. Scherillo, G. S. Simpson, I. Tsekhanovich, W. Urban, H. Faust, and N. Warr, Phys. Rev. C 71, 064327 (2005).
- [57] J. K. Hwang et al., Phys. Rev. C 67, 054304 (2003).
- [58] A. Złomaniec, H. Faust, J. Genevey, J. A. Pinston, T. Rzaca-Urban, G. S. Simpson, I. Tsekhanovich, and W. Urban, Phys. Rev. C 72, 067302 (2005).
- [59] S. Brant, G. Lhersonneau, and K. Sistemich, Phys. Rev. C 69, 034327 (2004).
- [60] W. Urban et al., Eur. Phys. J. A 16, 11 (2003).
- [61] C. Thibault et al., Phys. Rev. C 23, 2720 (1981).
- [62] S. Naimi et al., Phys. Rev. Lett. 105, 032502 (2010).
- [63] D. De Frenne, Nuclear Data Sheets 110, 1745 (2009).
- [64] M. Albers *et al.*, Phys. Rev. Lett. **108**, 062701 (2012), and references therein.
- [65] R. F. Petry, H. Dejbakhsh, J. C. Hill, F. K. Wohn, M. Shmid, and R. L. Gill, Phys. Rev. C 31, 621 (1985).
- [66] G. Simpson et al., Phys. Rev. C 82, 024302 (2010).
- [67] G. Lhersonneau, B. Pfeiffer, R. Capote, J. M. Quesada, H. Gabelmann, K. L. Kratz, and the ISOLDE Collaboration, Phys. Rev. C 65, 024318 (2002).
- [68] G. Lhersonneau, B. Pfeiffer, H. Gabelmann, K. L. Kratz, and the ISOLDE Collaboration, Phys. Rev. C 63, 054302 (2001).

- [69] J. K. Hwang *et al.*, Phys. Rev. C 58, 3252 (1998).
- [70] G. Lhersonneau, H. Gabelmann, M. Liang, B. Pfeiffer, K. L. Kratz, H. Ohm, and the ISOLDE Collaboration, Phys. Rev. C 51, 1211 (1995).
- [71] Y. Shi, P. M. Walker, and F. R. Xu, Phys. Rev. C 85, 027307 (2012).
- [72] H. Penttilä et al., Phys. Rev. C 54, 2760 (1996).
- [73] W. Urban, C. Droste, T. Rzaca-Urban, A. Zlclsomaniec, J. L. Durell, A. G. Smith, B. J. Varley, and I. Ahmad, Phys. Rev. C 73, 037302 (2006).
- [74] K.-L. Kratz, B. Pfeiffer, F.-K. Thielemann, and W. Walters, Hyperfine Interact. 129, 185 (2000).
- [75] I. Stefanescu *et al.*, Eur. Phys. J. A **42**, 407 (2009).
- [76] F. Sarazin et al., Phys. Rev. Lett. 84, 5062 (2000).
- [77] J. M. Daugas et al., Phys. Rev. C 81, 034304 (2010).
- [78] J. Genevey, F. Ibrahim, J. A. Pinston, H. Faust, T. Friedrichs, M. Gross, and S. Oberstedt, Phys. Rev. C 59, 82 (1999).
- [79] G. S. Simpson, A. Scherillo, J. Genevey, R. Orlandi, J. A. Pinston, I. S. Tsekhanovich, N. Warr, A. Covello, and A. Gargano, J. Phys.: Conf. Ser. 267, 012031 (2011).
- [80] F. Naqvi et al., Phys. Rev. C 82, 034323 (2010).
- [81] J. Genevey et al., Phys. Rev. C 67, 054312 (2003).
- [82] A. Scherillo et al., Phys. Rev. C 70, 054318 (2004).
- [83] K. Andgren et al., Phys. Lett. B 672, 313 (2009).
- [84] B. Fogelberg, K. Heyde, and J. Sau, Nucl. Phys. A 352, 157 (1981).
- [85] A. Korgul et al., Eur. Phys. J. A 7, 167 (2000).
- [86] G. S. Simpson et al., Phys. Rev. C 76, 041303 (2007).
- [87] S. N. Liddick et al., Phys. Rev. C 84, 061305 (2011).
- [88] J. M. Daugas *et al.*, Phys. Rev. C **83**, 054312 (2011), and references therein.
- [89] B. Fogelberg et al., Phys. Rev. C 70, 034312 (2004).
- [90] B. Fogelberg, M. Hellström, D. Jerrestam, H. Mach, J. Blomqvist, A. Kerek, L. O. Norlin, and J. P. Omtvedt, Phys. Rev. Lett. 73, 2413 (1994).
- [91] J. M. Pearson, R. C. Nayak, and S. Goriely, Phys. Lett. B 387, 455 (1996).
- [92] T. Otsuka, T. Suzuki, M. Honma, Y. Utsuno, N. Tsunoda, K. Tsukiyama, and M. Hjorth-Jensen, Phys. Rev. Lett. 104, 012501 (2010).